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Three-dimensional Forecasting Models on a Desk-Top Computer
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INTRODUCTION

Recent advances in the desk-top computer capabilities, particularly those of an engineering workstation, are truly astonishing. A high performance workstation reportedly exceeded a supercomputer in certain scalar operations. Affordability and portability of a desk-top computer opened doors to many applications which were previously considered impossible.

The purpose of the present study is to demonstrate the feasibility of using an engineering workstation to operate a three-dimensional mesoscale modeling system to describe the transport and dispersion of atmospheric pollutants over complex terrain surrounding Vandenberg Air Force Base (VAFB), Vandenberg, California.

Our modeling system is composed of two numerical codes, HOTMAC (Higher Order Turbulence Model for Atmospheric Circulations) and RAP-TAD (Random Puff Transport And Diffusion). HOTMAC is a mesoscale forecast code that is able to model three-dimensional distributions of wind speed, wind direction, turbulence, temperature, and water vapor. RAP-TAD is a Lagrangian puff code based on the Monte Carlo statistical diffusion process. RAP-TAD can be used under extreme conditions where a conventional Gaussian plume model may fail.

MODEL EQUATIONS

HOTMAC

The basic equations for mean wind, temperature, mixing ratio of water vapor, and turbulence are similar to those used by Yamada and Bunker.¹

A terrain-following vertical coordinate system is used in order to increase the accuracy in the treatment of surface boundary condition:

$$z^* = H \frac{z - z_g}{H - z_g} \quad , \quad (1)$$

where z^* and z are transformed and Cartesian vertical coordinates, respectively; z_g is ground elevation; H is the material surface top of the model in the z^* coordinate and H is the corresponding height in the z -coordinate.

The governing equations, following the coordinate transformation, are:¹

$$\begin{aligned} \frac{DU}{Dt} = & f(V - V_g) + g \frac{\bar{H} - z^*}{\bar{H}} \left(1 - \frac{\langle \Theta_v \rangle}{\Theta_v} \right) \frac{\partial z_g}{\partial x} \\ & + \frac{\partial}{\partial x} \left(K_x \frac{\partial U}{\partial x} \right) + \frac{\partial}{\partial y} \left(K_{xy} \frac{\partial U}{\partial y} \right) \\ & + \frac{\bar{H}}{H - z_g} \frac{\partial}{\partial z^*} (-\bar{u}w) , \end{aligned} \quad (2)$$

$$\begin{aligned} \frac{DV}{Dt} = & -f(U - U_g) + g \frac{\bar{H} - z^*}{\bar{H}} \left(1 - \frac{\langle \Theta_v \rangle}{\Theta_v} \right) \frac{\partial z_g}{\partial y} \\ & + \frac{\partial}{\partial x} \left(K_{xy} \frac{\partial V}{\partial x} \right) + \frac{\partial}{\partial y} \left(K_y \frac{\partial V}{\partial y} \right) \\ & + \frac{\bar{H}}{H - z_g} \frac{\partial}{\partial z^*} (-\bar{v}w) , \end{aligned} \quad (3)$$

$$\frac{\partial U}{\partial x} + \frac{\partial V}{\partial y} + \frac{\partial W^*}{\partial z^*} - \frac{1}{H - z_g} \left(U \frac{\partial z_g}{\partial x} + V \frac{\partial z_g}{\partial y} \right) = 0 , \quad (4)$$

where

$$W^* \equiv \frac{\bar{H}}{H - z_g} W + \frac{z^* - \bar{H}}{H - z_g} \left(U \frac{\partial z_g}{\partial x} + V \frac{\partial z_g}{\partial y} \right) , \quad (5)$$

and

$$\frac{D(\cdot)}{Dt} \equiv \frac{\partial(\cdot)}{\partial t} + U \frac{\partial(\cdot)}{\partial x} + V \frac{\partial(\cdot)}{\partial y} + W^* \frac{\partial(\cdot)}{\partial z^*} . \quad (6)$$

In the above expressions, $\langle \cdot \rangle$ indicates an average over a horizontal surface. The second terms on the right-hand side of Equations (2) and (3) indicate the effects of ground slope. For simplicity, H is specified as

$$H = \bar{H} + z_{gmax} , \quad (7)$$

where $\bar{H} = 5000$ m and $z_{gmax} = 565.7$ m are used in this study. Expressions of the horizontal eddy viscosity coefficients, K_x , K_y , and K_{xy} are given in Yamada.²

The geostrophic winds U_g and V_g are computed from Yamada;³

$$fU_g \equiv fU_g(\bar{H}) \frac{\langle \Theta_v \rangle}{\langle \Theta_v(\bar{H}) \rangle} + g \frac{H - z_g}{\bar{H}} \int_{z^*}^{\bar{H}} \frac{1}{\langle \Theta_v \rangle} \\ \times \frac{\partial}{\partial y} \Delta \Theta_v dz' - \frac{g}{\bar{H}} \frac{\partial z_g}{\partial y} \int_{z^*}^{\bar{H}} \frac{\Delta \Theta}{\langle \Theta_v \rangle_v} dz' \quad (8)$$

and

$$fV_g \equiv fV_g(\bar{H}) \frac{\langle \Theta_v \rangle}{\langle \Theta_v(\bar{H}) \rangle} - g \frac{H - z_g}{\bar{H}} \int_{z^*}^{\bar{H}} \frac{1}{\langle \Theta_v \rangle} \frac{\partial \Delta \Theta_v}{\partial x} dz' \\ + \frac{g}{\bar{H}} \frac{\partial z_g}{\partial x} \int_{z^*}^{\bar{H}} \frac{\Delta \Theta_v}{\langle \Theta_v \rangle} dz' \quad , \quad (9)$$

where $\Delta \Theta_v \equiv \Theta_v - \langle \Theta_v \rangle$, and the abbreviated symbols $U_g(\bar{H}) \equiv U_g(x, y, \bar{H}, t)$; $V_g(\bar{H}) \equiv V_g(x, y, \bar{H}, t)$ etc. are used. The derivations of Equations (2)-(5), (8), and (9) are given in the Appendix of Yamada.³ A turbulence kinetic energy equation is given by

$$\frac{D}{Dt} \left(\frac{q^2}{2} \right) = \frac{\partial}{\partial x} \left[K_x \frac{\partial}{\partial x} \left(\frac{q^2}{2} \right) \right] + \frac{\partial}{\partial y} \left[K_y \frac{\partial}{\partial y} \left(\frac{q^2}{2} \right) \right] \\ + \left(\frac{\bar{H}}{H - z_g} \right)^2 \frac{\partial}{\partial z^*} \left[q l S_q \frac{\partial}{\partial z^*} \left(\frac{q^2}{2} \right) \right] - \frac{\bar{H}}{H - z_g} \\ \times \left(u \bar{w} \frac{\partial U}{\partial z^*} + v \bar{w} \frac{\partial V}{\partial z^*} \right) + \beta_g \overline{w \theta_v} - \frac{q^3}{B_1 \ell} \quad (10)$$

and a turbulence length scale ℓ is obtained from

$$\frac{D}{Dt} (q^2 \ell) = \frac{\partial}{\partial x} \left[K_x \frac{\partial}{\partial x} (q^2 \ell) \right] + \frac{\partial}{\partial y} \left[K_y \frac{\partial}{\partial y} (q^2 \ell) \right] \\ + \left(\frac{\bar{H}}{H - z_g} \right)^2 \frac{\partial}{\partial z^*} \left[q l S_\ell \frac{\partial}{\partial z^*} (q^2 \ell) \right]$$

$$\begin{aligned}
& + \ell F_1 \left[\frac{\bar{H}}{H - z_g} \left(-\overline{uw} \frac{\partial U}{\partial z^*} - \overline{vw} \frac{\partial V}{\partial z} \right) + \beta g \overline{w\theta_v} \right] \\
& - \frac{q^3}{B_1} \left[1 + F_2 \left(\frac{1}{kz} \right)^2 \right]
\end{aligned} \tag{11}$$

where $\overline{q^2} = \overline{u^2} + \overline{v^2} + \overline{w^2}$ is twice the turbulence kinetic energy; $\overline{w\theta_v}$, turbulence heat flux; θ_v , the fluctuation part of virtual potential temperature, and $(F_1, F_2, S_q, S_\ell, \text{ and } B_1) = (1.8, 1.33, 0.2, 0.2, \text{ and } 16.6)$, empirical constants determined from laboratory experiments (Mellor and Yamada⁴).

The internal heat energy equation is written in terms of $\Delta\Theta_v$ as discussed in Yamada and Bunker.⁵ We obtain a prognostic equation for $\Delta\Theta_v$ as

$$\begin{aligned}
\frac{D\Delta\Theta_v}{Dt} &= \frac{\partial}{\partial x} \left(K_x \frac{\partial \Delta\Theta_v}{\partial x} \right) + \frac{\partial}{\partial y} \left(K_y \frac{\partial \Delta\Theta_v}{\partial y} \right) \\
&+ \frac{\bar{H}}{H - z_g} \left[\frac{\partial}{\partial z^*} (-\overline{w\theta_v}) + \frac{1}{\rho C_p} \frac{\partial R_N}{\partial z^*} - W \frac{\partial \Theta_v}{\partial z^*} \right], \tag{12}
\end{aligned}$$

where $\partial < \Theta_v > / \partial x = \partial < \Theta_v > / \partial y = 0$ are used.

The longwave radiation flux $R_N / \rho C_p$ is computed according to Sasamori.⁶ A conservation equation for mixing ratio of water vapor is given by

$$\begin{aligned}
\frac{\partial Q_v}{\partial t} &= \frac{\partial}{\partial x} \left[K_x \frac{\partial Q_v}{\partial x} \right] + \frac{\partial}{\partial y} \left[K_y \frac{\partial Q_v}{\partial y} \right] \\
&+ \frac{\bar{H}}{H - z_g} \frac{\partial}{\partial z^*} (-\overline{wq_v}), \tag{13}
\end{aligned}$$

The turbulent fluxes in Equations (2), (3), (10), (11), (12), and (13) are obtained from simplified second-moment turbulence-closure equations (Yamada⁷):

$$(\overline{uw}, \overline{vw}) = -lq\tilde{S}_M \left(\frac{\partial U}{\partial z}, \frac{\partial V}{\partial z} \right) \tag{14a, b}$$

$$(\overline{w\theta_v}, \overline{wq_v}) = -\alpha lq\tilde{S}_M \left(\frac{\partial \Theta_v}{\partial z}, \frac{\partial Q_v}{\partial z} \right), \tag{15a, b}$$

where \tilde{S}_M and α are functions of the flux Richardson number, and $\alpha(\equiv K_H/K_M$ where K_H is an eddy diffusivity coefficient and K_M is an eddy viscosity coefficient) is the reciprocal of the turbulent Prandtl number. The expressions for \tilde{S}_M and α are obtained from the level 2 model (production = dissipation) of Mellor and Yamada⁸ where temporal and spatial derivatives in Eq. (9) are neglected. The readers are referred to Yamada⁹ for further discussions of the level 2 model equations. The final expressions for \tilde{S}_M and α are given in (Yamada⁷) and are not repeated here.

RAPTAD

A brief description of RAPTAD (Yamada and Bunker¹) is given herein. Locations of the center of each puff are computed from

$$x_i(t + \Delta t) = x_i(t) + U_{pi}\Delta t \quad (16)$$

where

$$U_{pi} = U_i + u_i \quad , \quad (17)$$

$$u_i(t + \Delta t) = au_i(t) + b\sigma u_i\zeta \quad (18)$$

$$a = \exp(-\Delta t/t_{Lx_i}) \quad , \quad (19)$$

and

$$b = (1 - a^2)^{1/2} \quad . \quad (20)$$

In the above expressions, U_{pi} is the particle velocity in the x_i direction; U_i , mean velocity; u_i , turbulence velocity; ζ , a random number from a Gaussian distribution with zero mean and unit variance, t_{Lx_i} , the Lagrangian integral time for the velocity u_i ; and σ_{u_i} , standard deviation of velocity fluctuation, u_i .

The Lagrangian time scales, $t_{Lx} = t_{Ly} = 10^4 s$ (Gifford¹⁰) and $t_{Lz} = 20s$, (Yamada and Bunker¹) are used in this study.

Concentration χ at (X, Y, Z) is estimated by using the following expression:

$$\chi(X, Y, Z) = \frac{Q\Delta t}{(2\pi)^{3/2}} \sum_{k=1}^N \frac{1}{\sigma_{xk}\sigma_{yk}\sigma_{zk}} \exp\left(-\frac{1}{2} \frac{(x_k - X)^2}{\sigma_{xk}^2} - \frac{(y_k - Y)^2}{\sigma_{yk}^2} - \frac{(z_k - Z)^2}{\sigma_{zk}^2}\right)$$

$$\begin{aligned} & \times \exp \left(-\frac{1}{2} \frac{(y_k - Y)^2}{\sigma_{y_k}^2} \right) \left\{ \exp \left(-\frac{1}{2} \frac{(z_k - Z)^2}{\sigma_{z_k}^2} \right) \right. \\ & \left. + \exp \left(-\frac{1}{2} \frac{(z_k + Z - 2z_g)^2}{\sigma_{z_k}^2} \right) \right\}, \end{aligned} \quad (21)$$

where (x_k, y_k, z_k) is the location of k th particle; σ_{x_k} , σ_{y_k} , and σ_{z_k} are standard deviations of a Gaussian distribution; and z_g is the ground elevation. The variances are estimated based on Taylor's¹¹ homogeneous diffusion theory. For example, σ_y is obtained from

$$\begin{aligned} \sigma_y^2 &= 2\sigma_v^2 \int_0^t \int_0^\zeta R(\zeta) d\zeta dt \\ &= 2\sigma_v^2 t_{Ly} \left(t + t_{Ly} \exp \left(-\frac{t}{t_{Ly}} \right) - t_{Ly} \right), \end{aligned} \quad (22)$$

where a correlation function $R(\zeta) = \exp(-\zeta/t_{Ly})$ is used.

In a similar fashion, σ_x and σ_z are determined. The standard deviations σ_u , σ_v and σ_w at each particle location are obtained by interpolating grid values of a computation grid volume in which a particle is located.

SIMULATIONS AND DISCUSSIONS

The Mountain Iron (MI) diffusion experiments (Hinds and Nickola¹²) were conducted at Vandenberg Air Force Base during 1965 and 1966 to establish quantitative diffusion predictions for use as range safety tools in the "South Vandenberg" (SV) ballistic and space vehicle operations.

The experimental site, SV, is located along the California coast approximately 160 km west-northwest of Los Angeles. The coastline is oriented in approximately a north-south direction along the western side of SV, but changes abruptly at Point Arguello to a east-west direction. The coastline gradually changes to a north-south direction down to Point Conception and then changes again to a east-west direction. Santa Ynez Mountains form a east-west barrier along the coastline far south of SV.

The purpose of the present study is to demonstrate the feasibility of using an engineering workstation to operate three-dimensional models under realistic boundary conditions: land-sea contrast and complex topography. Thus, only qualitative discussions are given here, and more detailed study on model-observation comparisons is found in Yamada and Bunker.¹³

We selected the MI87 case which was characterized (Hunter¹⁴) as "Weak ambient and sea-breeze flow, with little distinction between wind vectors, but sharp cooling with fog rolling in."

Initial potential temperature profile was determined from the upper air soundings at 1300 lst. The potential temperature lapse rate was approximately $0.04^{\circ}\text{C}/\text{m}$ from the sea surface to 460 m msl and $0.044^{\circ}\text{C}/\text{m}$ above 460 m msl. Prevailing wind speed (2 m/s) and direction (225°C) were determined by examining five upper air soundings at 1300 lst.

The computational domain is $40 \times 48 \text{ km}^2$ with a horizontal grid spacing of 1 km. Integration started at 0500 lst, June 13, 1966 and continued for over 48 hours. The model plume was released continuously for 24 hrs starting at 2300 lst to examine the diurnal variation of the plume transport and diffusion processes.

Figure 1 shows the modeled horizontal wind vectors at 14 m above the ground at 0700 lst, June 14. Although the upper air wind direction was 225°C (southwesterly) downslope flows and land breeze developed in the surface layer due to cooling at the ground. The sunrise on June 14, 1966 was shortly before 0500 lst.

As the sun warms the ground, the air temperature in the surface layer first increases by convection. The air temperature over a sloped surface becomes higher than the temperature at the same elevation but away from the ground. The temperature difference results in a horizontal pressure gradient: a low pressure is normally located at a high elevation point. This pressure force produces upslope flows and sea-breezes that converge into ridges (Fig. 2).

The ground starts cooling shortly before sunset (around 2000 lst) and the air temperature close to the ground over a sloped surface is now lower than the air temperature at the same elevation but away from the surface. Thus, the direction of pressure gradient in the nighttime reverses from the daytime counterpart. However, the magnitude of the nighttime pressure force is normally much smaller than the daytime value, particularly in the summer season, for the following reasons.

First, turbulence intensity in the surface layer during the nocturnal period is extremely small, comparing with the daytime value. Therefore, only the air relatively close to the ground is cooled by turbulent mixing. The depth of the nocturnal stable boundary layer is on the order of 100 m which is much smaller than the daytime convective boundary layer height of at least several hundred meters.

Second, the solar energy that reaches the ground in Summer is far larger than the net longwave radiation loss during the nighttime. In other words,

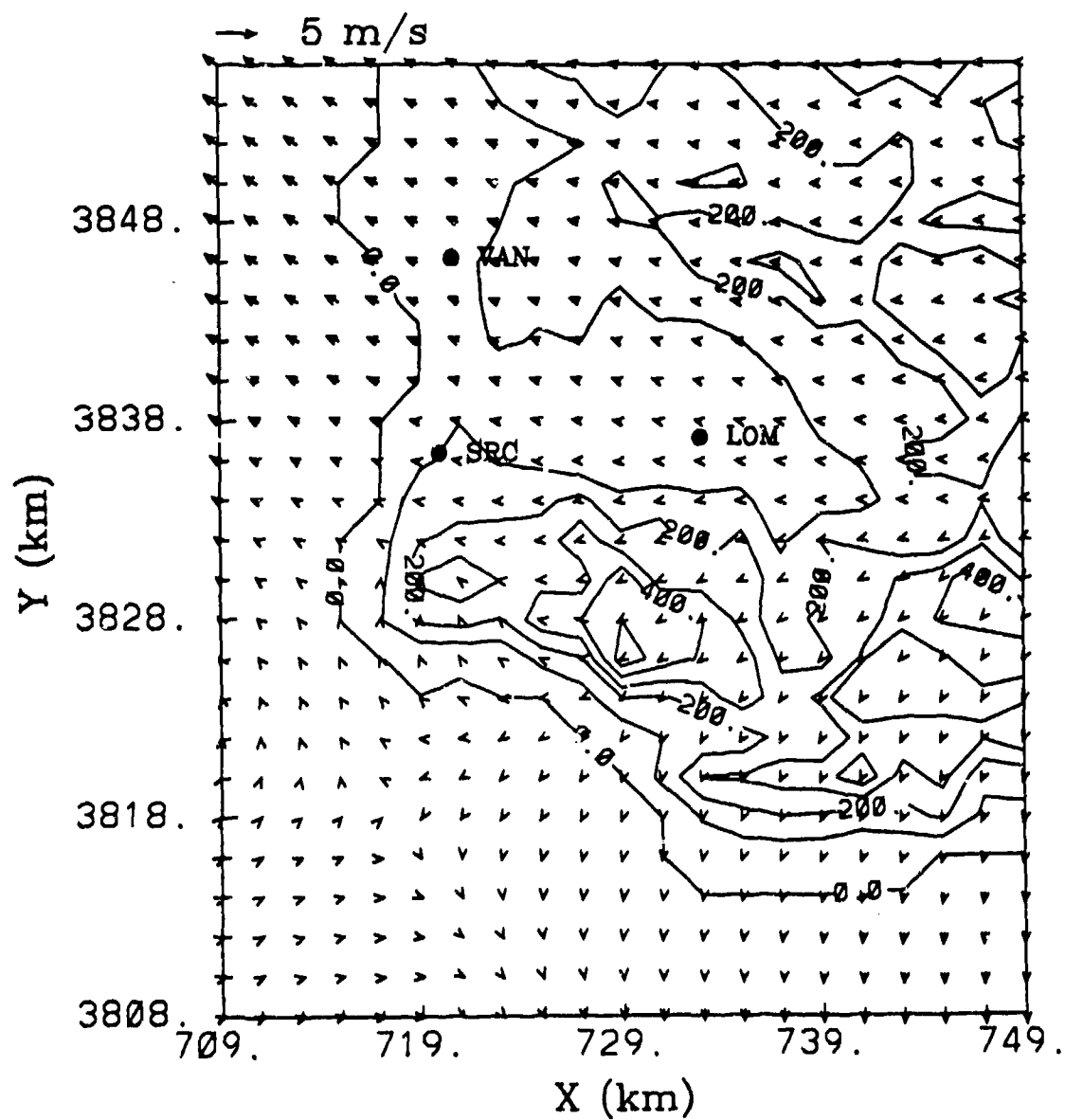


Figure 1. Modeled horizontal wind vectors at 14 m above the ground at 0700
1st, June 14, 1966. Terrain is contoured by solid lines with an increment of
100 m.

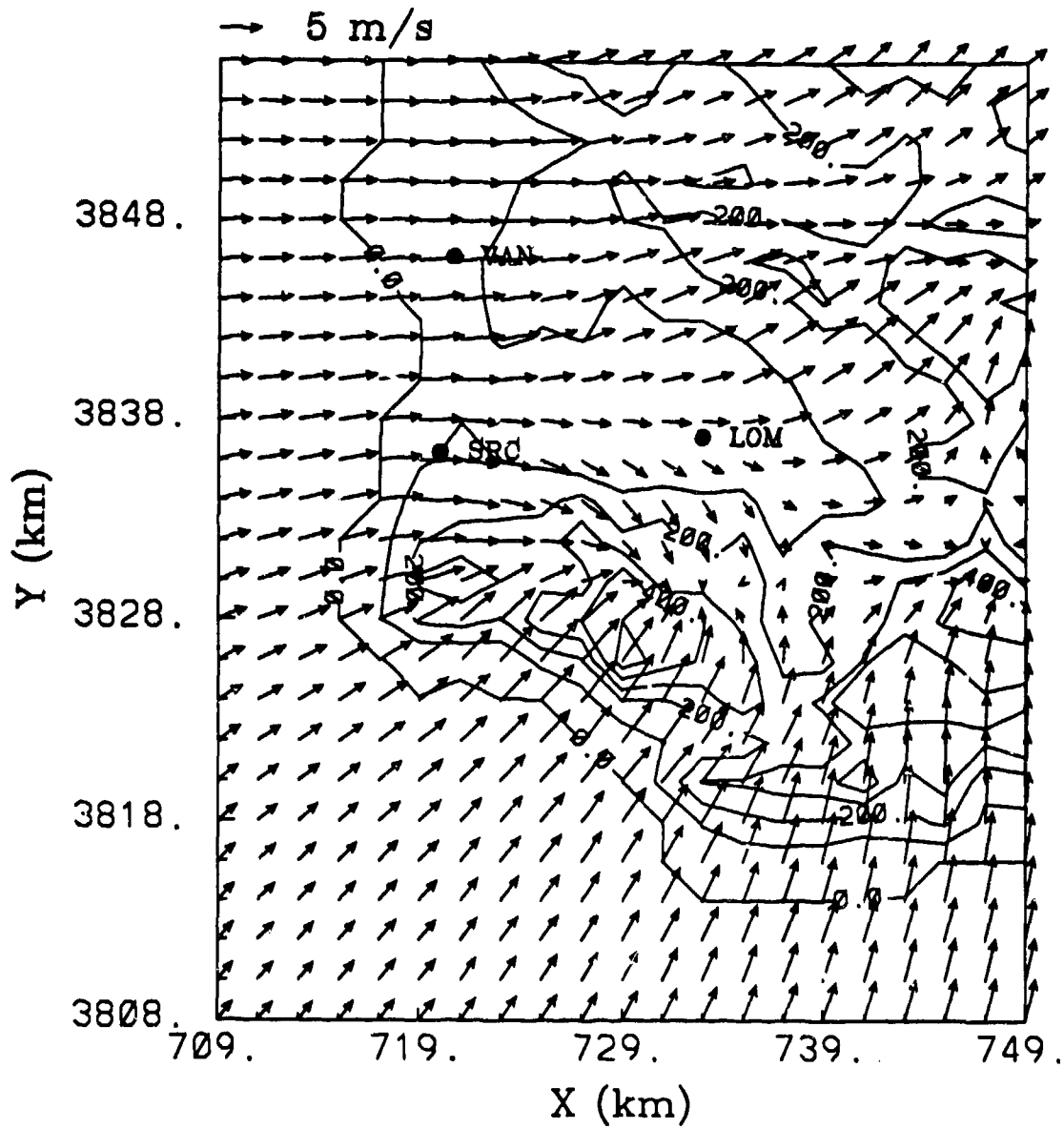


Figure 2. Modeled horizontal wind vectors at 14 m above the ground at 1400 lst, June 14, 1966. Terrain is contoured by solid lines with an increment of 100 m.

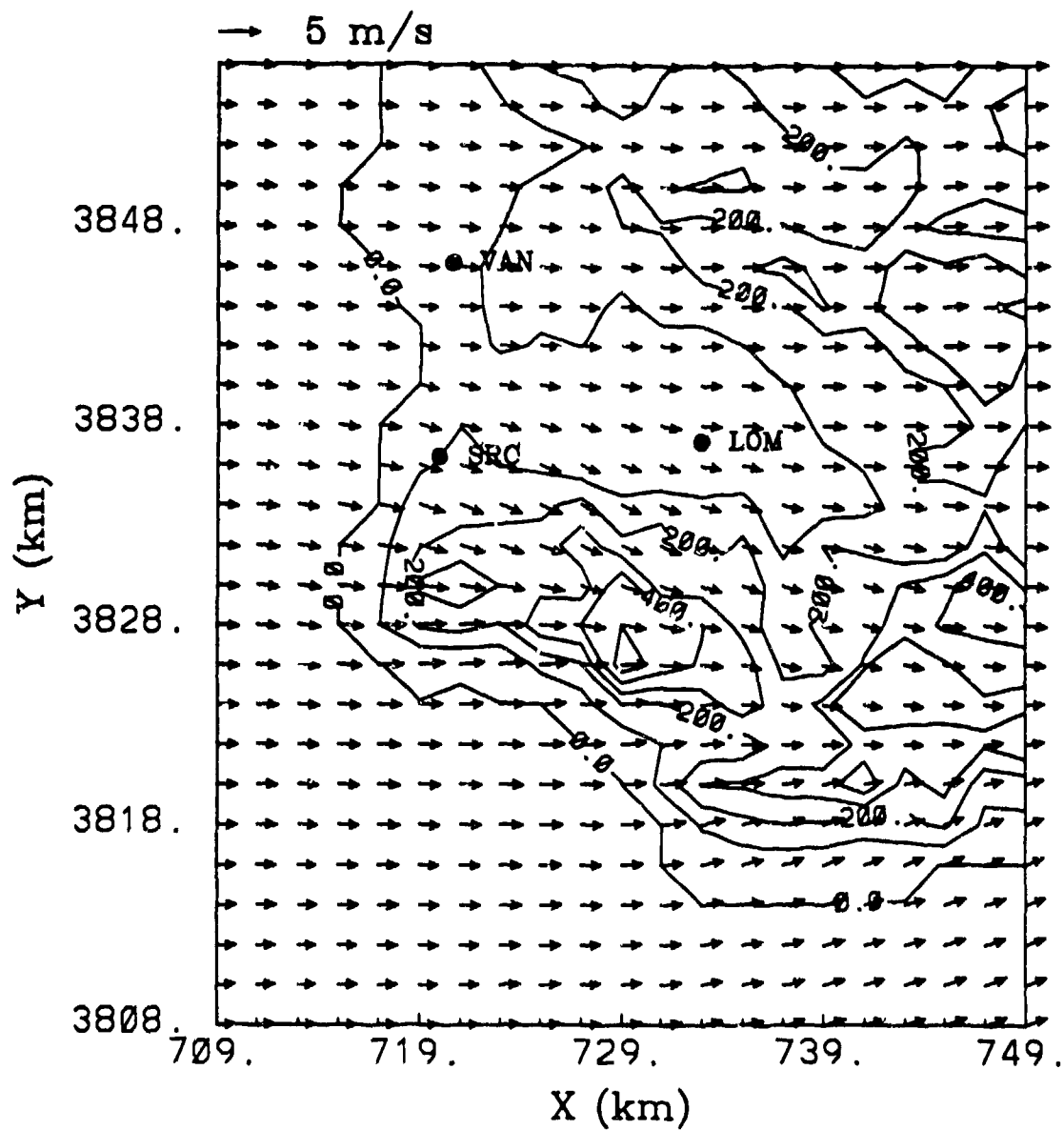


Figure 3. Modeled horizontal wind vectors at 14 m above the ground at 2300 1st, June 14, 1966. Terrain is contoured by solid lines with an increment of 100 m.

much more heat energy is input to the atmosphere during the day than the extraction of the heat energy from the atmosphere during the night.

The combination of the deeper boundary layer depth and greater heat energy available to heat the atmosphere during the daytime produced much larger pressure gradients than the nighttime counterpart. This is the reason why the modeled wind directions (Fig. 3) were not downslope although the upslope flow components encountered during the daytime (Fig. 2) almost completely diminished. The modeled wind direction later became downslope but the magnitude was very small.

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