

CONF 901057-22

UCRL-JC--105416

DE91 005432

DEC 27 1990

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**This paper was prepared for submittal to  
Capture Gamma-Ray Spectroscopy 7th Int. Sym.  
Asilomar, CA, October 14-16, 1990**

**October 14, 1990**

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INVITED PAPER AT THE "CAPTURE GAMMA-RAY SPECTROSCOPY" 7TH INTERNATIONAL SYMPOSIUM, ASILOMAR, CALIF., 10/14-19/1990

NUCLEAR SHELL MODEL CALCULATIONS WITH NON-LOCAL INTERACTIONS

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ABSTRACT

It is becoming clearer with time that non-locality of the nucleon-nucleon interaction can play a significant role in nuclear properties. In this talk we review evidence for such non-locality. Then, using a Gaussian interaction, we discuss the effect of non-locality on two body matrix elements in the nuclear shell model. Finally, we mention some applications. For example, non-locality leads to faster convergence of off-diagonal matrix elements.

I. WHAT IS NON-LOCALITY?

1. Definition

Let us begin with discussing what non-locality is. Basically, the Schroedinger equation for a particle in a conventional (local) potential has the well known form:

$$-(\hbar^2/2m)\psi(\vec{r}) + V(\vec{r})\psi(\vec{r}) = E\psi(\vec{r}). \quad (1)$$

which can also be written as:

$$-(\hbar^2/2m)\psi(\vec{r}) + \int V(\vec{r})\delta(\vec{r}-\vec{r}')\psi(\vec{r}')d^3r' = E\psi(\vec{r}). \quad (2)$$

For a non-local interaction, we have, instead:

$$-(\hbar^2/2m)\psi(\vec{r}) + \int V(\vec{r}, \vec{r}')\psi(\vec{r}')d^3r' = E\psi(\vec{r}). \quad (3)$$

i.e.  $\vec{r}$  and  $\vec{r}'$  don't have to be at the same point.

2. Relation to momentum dependence

Instead of writing the interaction as a non-local operator, we can express it in the form of a local, but momentum dependent interaction. Thus, for example, a non-local Gaussian interaction of the form:

$$V(\vec{r}, \vec{r}') = e^{-(\vec{r}+\vec{r}')^2/4a^2} \times e^{-(\vec{r}-\vec{r}')^2/4c^2} \quad (4)$$

is equivalent (apart from a multiplicative constant) to a momentum dependent Gaussian interaction of the form:

$$V(\vec{r}, \vec{p}) = e^{-r^2/a^2} \times e^{-p^2/c^2} \quad (5)$$

(Actually, this interaction has to be properly symmetrized w.r.t.  $r$  and  $p$ , so that it is Hermitean.) The connection between non-locality and momentum dependence is more complicated for a non-Gaussian interaction.

### 3. Relation to energy dependence

Instead of writing the interaction as momentum dependent, we can also try to represent it as local, but energy dependent. However, when we do this, we lose the orthogonality property of wavefunctions of different energy.

## II. WHY DO WE NEED NON-LOCALITY?

### 1. Theory

#### a. A simplified quark model for non-locality.

Some non-locality of the NN interaction at short distances is expected on account of the quark structure of nucleons, that is to say, due to the nucleon compositeness. To illustrate this point, consider a very crude model, in which each nucleon is replaced by a cluster consisting of two (not three) particles. The two particles have masses  $M$  and  $m$ , respectively. If there is a harmonic interaction between them,  $V = \frac{1}{2}\mu\omega^2r^2$ , where  $\mu$  is the reduced mass  $= Mm/(M+m)$  and  $r$  is the spacing between the particles, then the wavefunction of relative motion is a Gaussian proportional to  $e^{-\mu\omega r^2/2\hbar} = e^{-r^2/2b^2}$ . Now consider the interaction between two clusters whose centers of mass are separated by a distance  $R$ . We suppose that the interaction between the particles has a range very long compared to the cluster size  $b$ , and that it involves exchanging the light particles on the two clusters. Using standard cluster model methods, it can then be shown that the interaction between the clusters is just of the non-local form discussed above:

$$V(\vec{R}, \vec{R}') = e^{-(\vec{R}+\vec{R}')^2/4a^2} \times e^{-(\vec{R}-\vec{R}')^2/4c^2} \quad (6a)$$

with

$$a^2 = 2b^2/[1 + \frac{m}{M}]^2, \quad c^2 = 2b^2/[1 + \frac{M}{m}]^2 \quad (6b)$$

Two particular limits are of interest.

i. For  $m \ll M$ , we obtain  $a^2 = 2b^2$ ,  $c = 0$ , i.e. the interaction is local, with a

$$V(\vec{R}, \vec{R}') = e^{-\vec{R}^2/2b^2} \times \delta(\vec{R} - \vec{R}') \quad (7)$$

range governed by the overlap of wavefunctions between the two clusters. This case corresponds to the interaction between atoms. Since the atomic nucleus has a much larger mass than an electron, the interactions between atoms are nearly local.

ii. For  $m=M$ , which more closely corresponds to the quark model of nucleons, (except for containing 2 quarks instead of 3), we find  $a^2 = c^2 = b^2/2$ . For this case, the interaction is separable.

$$V(\vec{R}, \vec{R}') = e^{-(R^2 + R'^2)/b^2} \quad (8)$$

b. Relativistic Model

Alternatively, some non-locality also arises in relativistic nuclear models, due to relativistic retardation effects.

2. Experimental evidence: Deuteron radius

There are only a few well-known examples of non-local interactions known in nature. The best evidence (to date) for non-locality of the N-N interaction comes from the analysis of the deuteron charge radius<sup>1</sup>. With traditional (basically local) interactions, it is possible to fit N-N scattering phase shifts very well, but one obtains a deuteron radius about 1% larger than the empirical value. This is illustrated in Figure 1. The problem of the deuteron radius can be easily resolved, without destroying the fit to phase shifts, by making the potential non-local<sup>2</sup>. For example, there are simple unitary transformations on the wavefunction which change the short distance behavior, but keep intact both the orthogonality properties of non-degenerate pairs of wavefunction, and the wavefunction at large distances, which is what determines the phase shifts.

### III. NON-LOCALITY AND SHELL MODEL CALCULATIONS

#### 1. Two-body interaction matrix elements

It is well known that if we use harmonic oscillator single particle wavefunctions, then the two body matrix elements for Gaussian interactions can be calculated analytically. This also holds for a non-local Gaussian, in fact, the expressions are only slightly more involved. It is readily shown that in the short range limit; i.e., when both the range  $a$  and the non-locality range  $c$  are small compared to the oscillator length  $b$ , then the two body matrix elements reduce to those for a Skyrme interaction (without density dependence). It turns out that non-locality increases the even state matrix elements, but reduces the odd state matrix elements. If  $a = c$ , then only matrix elements with relative orbital angular momentum zero survive. In fact, in this case, the interaction is separable. Another case of interest occurs for a long range. Here the interaction effectively reduces to a quadrupole-quadrupole interaction. Finally, for  $ac = 2b^2$ , all two body matrix elements connecting different oscillator shells vanish identically. An interaction with these (somewhat unrealistic) values of the parameters gives rise to a mean field but no splitting between states belonging to the same irreducible representation of  $SU(3)$  but different  $L$  values.

We discuss here the Talmi integrals for non-local Gaussian interactions. Suppose we have the following non-local Gaussian interaction:

$$V = -V_0 \exp - \left[ \frac{(r+r')^2}{4a^2} + \frac{(r-r')^2}{4c^2} \right] \quad (9)$$

in even states, and  $\eta V$  in odd states.

Then the integrals involving relative motion can be calculated analytically. For non-local interactions, they are only slightly more complicated than for a local interaction.

We obtain:

$$I_{00} = V_0 \lambda^{3/2} (1-\mu)^{3/2} \quad (10a)$$

$$\text{where } \lambda = \frac{a^2}{a^2+2b^2}, \quad \mu = \frac{c^2}{c^2+2b^2} \quad (10b)$$

Here is an expression for another Talmi integral:

$$I_{11}/I_{00} = \eta(\lambda-\mu) \quad (11)$$

Note that if  $\lambda = \mu$ , i.e. if  $a = c$ , (local and non-local ranges are equal), then  $I_{11} = 0$ . Similarly, all other Talmi integrals with relative angular momentum larger than 0 vanish. This is not surprising, since this case corresponds to a separable interaction, which acts only in relative S-states. Thus, for example:

$$I_{22}/I_{00} = (\lambda-\mu)^2 \quad (12)$$

On the other hand, the Talmi integral  $I_{20}$  is finite even for a separable interaction.

$$I_{20}/I_{00} = (\lambda-\mu)^2 + \frac{3}{2}(1-\lambda-\mu)^2 \quad (13)$$

However, there is another interesting limit, that where  $\lambda + \mu = 1$ . This corresponds to the condition  $ac = 2b^2$ . For this case, we see that  $I_{20} = I_{22}$ , and, more generally,  $I_{N\ell}$  is independent of  $\ell$ . Furthermore, all off-diagonal matrix elements connecting different oscillator shells vanish.

For example,

$$I_{00-20}/I_{00} = \sqrt{\frac{3}{2}}(1-\lambda-\mu) \quad (14)$$

which equals 0 in this case.

It is well known that for an infinite range interaction, i.e.  $\lambda = 1$ ,  $\mu = 0$ , all off-diagonal matrix elements vanish. However, as is shown here, this result holds more generally, even for a finite range interaction, provided the range of the non-locality is chosen equal to  $c = 2b^2/a$ . In this limit, the interaction provides a mean field, but there is no pairing.

Finally, it should be noted that a finite range space exchange interaction,  $a = r_0$ ,  $c = 0$ ,  $\eta = -1$ , ( $\lambda = \lambda_0$ ,  $\mu = 0$ ) has the same Talmi integrals, and thus two-body matrix elements, as a zero range, but finite non-locality range  $a = 0$ ,  $c = r_0$ ,  $\eta = 1$ , ( $\lambda = 0$ ,  $\mu = \lambda_0$ ) interaction. This indicates that, to some extent the effect of non-locality can be simulated by a local but space exchange range interaction.

## 2. Connection with Skyrme Interaction

In this talk we will restrict our consideration to a density independent interaction. For a short range interaction, i.e. with both  $a$  and  $c \ll b$ , we obtain: for the two parameters  $\lambda$  and  $\mu$ :

$$\lambda = \frac{a^2}{a^2 + 2b^2} \rightarrow \frac{a^2}{2b^2} \left(1 - \frac{a^2}{2b^2}\right) \quad (15a)$$

$$\mu = \frac{c^2}{c^2 + 2b^2} \rightarrow \frac{c^2}{2b^2} \quad (15b)$$

The Skyrme parameters are defined as follows for a local short range interaction:

$$t_0 = \int V(r) d^3r = V_0 (\pi a^2)^{3/2} \quad (16a)$$

$$t_1 = -\frac{1}{3} \int V(r) r^2 d^3r = -t_0 a^2 / 2 \quad (16b)$$

$$t_2 = \frac{1}{3} \int V(r) r^2 d^3r = \eta t_0 a^2 / 2 \quad (16c)$$

$\eta$  is the ratio of odd state to even state interaction.

Then

$$I_{00} = V_0 \lambda^{3/2} (1-\mu)^{3/2} \rightarrow V_0 \left( \frac{\pi a^2}{2\pi b^2} \right)^{3/2} \times \left( 1 - \frac{3}{4} \frac{a^2}{b^2} \right) \quad (17)$$

$$I_{no} = \int \psi_n(r) V(r, r') \psi_0(r') d^3r d^3r' \quad (18)$$

For a short range but non-local interaction we must make the following changes:

in  $t_1$ , replace  $a^2$  by  $a^2 + c^2$ , and,

in  $t_2$ , replace  $a^2$  by  $a^2 - c^2$ .

Also,  $I_{00}$  is now given by:

$$I_{00} = V_0 \lambda^{3/2} (1-\mu)^{3/2} \rightarrow V_0 \left( \frac{\pi a^2}{2\pi b^2} \right)^{3/2} * \left( 1 - \frac{3}{4} \frac{a^2 + c^2}{b^2} \right) \quad (19)$$

Thus non-locality increases the apparent range of the even-state interaction, but decreases that of the odd-state interaction.

In the Skyrme approximation<sup>3</sup>, the density independent part of the interaction is written as:

$$V(p, r) = t_0 \delta(r) + \frac{1}{2} [p^2 \delta(r) + \delta(r) p^2] t_1 + p \cdot \delta(r) \cdot p t_2 \quad (20)$$

If  $t_2 = -t_1$ , i.e.  $\eta = 1$ , then

$$V(p, r) = t_0 \delta(r) + \frac{1}{2} [p^2 \delta(r) - 2p \cdot \delta(r) \cdot p + \delta(r) p^2] t_1 \\ = t_0 \delta(r) - \frac{1}{2} [\nabla^2 \delta(r)] t_1 \quad (21)$$

which is just a local interaction.

The Talmi integrals for relative angular momentum zero are given as follows:

$$I_{n0} \rightarrow \psi_{n0}^2(0) t_0 - \psi_{n0}(0) \nabla^2 \psi_{n0}(0) t_1 \quad (22)$$

Now

$$\nabla^2 \psi_{n0}(0) = -\frac{n+3/2}{b^2} \psi_{n0}(0) \quad (23)$$

Thus

$$I_{n0} = \psi_{n0}^2(0) \left( t_0 + \frac{n+3/2}{b^2} t_1 \right) = \psi_{n0}^2(0) t_0 \left( 1 - \frac{n+3/2}{b^2} (a^2 + c^2) \right) \quad (24)$$

Also,

$$I_{11} = [\nabla \psi_{11}(0)]^2 t_2 = \frac{1}{2} \eta t_0 \frac{a^2 - c^2}{b^2} \quad (25)$$

It is interesting to list here ratios of the Talmi integrals in the Skyrme approximation:

$$I_{20}/I_{00} = (\lambda - \mu)^2 + \frac{3}{2} (1 - \lambda - \mu)^2 \rightarrow \frac{3}{2} \times \left( 1 - \frac{a^2 + c^2}{b^2} \right) \quad (26a)$$

$$I_{22}/I_{00} = (\lambda - \mu)^2 \rightarrow O(a^4) \quad (26b)$$

$$I_{00-20}/I_{00} = \sqrt{\frac{3}{2}} (1 - \lambda - \mu) \rightarrow \sqrt{\frac{3}{2}} \left( 1 - \frac{1}{2} \frac{a^2 + c^2}{b^2} \right) \quad (26c)$$

$$I_{11}/I_{00} = \eta (\lambda - \mu) \rightarrow \frac{1}{2} \eta \frac{a^2 - c^2}{b^2} \quad (26d)$$

Finally, we discuss two possible applications of non-local interactions in the nuclear shell model.

### 3. Overlap integrals

Consider the filling of shells in the  $A = 100$  region. Here we have 40 to 45 protons, filling the  $p1/2$  and  $g9/2$  shells, while there are 55 to 60 neutrons which fill the  $g7/2$  and  $d5/2$  shells. There are, of course, interactions between protons and neutrons, and there is some empirical evidence from study of Zr and Mo isotopes that the interaction between the proton  $g9/2$  and neutron  $g7/2$  shells is especially large.<sup>4</sup> As the neutron  $g7/2$  shell fills, the proton  $g9/2$  shell drops in energy. The apparent special overlap between these shells is larger than calculated for a local zero or finite range interaction. It requires that the interaction is short range in phase space; i.e., the combined  $r$  and  $p$  space. We discuss here the overlap integrals for a simpler case, but one which has the essential physics, namely particles in the oscillator ( $s, d$ ) shell.

As can be seen from Table 1, for a local interaction, the overlap integrals are only slightly smaller between  $s$  and  $d$  than between two particles in the same orbits. However, non-locality enhances the difference, (though not by much unless the non-locality range is large). In the extreme case of a zero range interaction with infinite range non-locality, the overlap integral vanishes unless the two particles are in the same spatial state. It is like a zero range interaction in phase space; i.e., the combined  $p$  and  $r$  space.

#### 4. Convergence of model space expansion

For a finite non-locality range, the two body matrix elements connecting different oscillator shells disappear faster than for a purely local interaction. This should lead to more rapid convergence of shell model energies as we increase the size of the model space. Table 2 indicates how non-locality increases the rate of convergence for the simple case of  ${}^4\text{He}$ , where, in lowest approximation, the nucleons are in the oscillator ground state. Such a faster convergence of the matrix elements can be readily understood. Intermediate states correspond to larger momenta. Now, a non-local interaction can be expressed as a momentum dependent one; i.e., it falls off faster with momentum, (not just momentum transfer, as with a local interaction) than a momentum dependent one. Since the intermediate states have larger momenta, the matrix elements involving such states will be reduced.

#### IV. SUMMARY

We have discussed the evidence for non-locality in the NN interaction, and the effect of a simple form of non-locality with Gaussians on nuclear shell model calculations. It remains to apply these ideas quantitatively to realistic NN interactions which fit scattering data, as well as the deuteron radius. Of course, this requires consideration of the density dependence in the NN interaction<sup>5</sup>.

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\* Work performed under the auspices of the U.S. Department of Energy by the Lawrence Livermore National Laboratory under Contract W-7405-Eng-48.

TABLE 1

Illustrating that non-locality leads to larger overlap between particles in the same orbit.

		$\lambda = \frac{a^2}{a^2+2b^2}$ ; $\mu = \frac{c^2}{c^2+2b^2}$		
$v_{ds}/v_{dd}$		$a \ 0$	$\sqrt{\frac{2}{3}}b$	
$c$	$\mu$	$\lambda \ 0$	0.25	
0	0	0.595	0.891	1.022
$\sqrt{\frac{2}{3}}b$	0.25	0.561	0.816	0.971
$\sqrt{2}b$	0.5	0.496	0.728	0.921
$\sqrt{6}b$	0.75	0.192	0.445	0.728
$\infty$	1.00	0.000	0.192	0.496
		$\sqrt{\frac{2}{3}}b$		$\sqrt{2}b$
$c$	$\mu$	$\lambda \ 0$	0.25	0.5
0	0	2.440	1.386	1.133
$\sqrt{\frac{2}{3}}b$	0.25	2.663	1.336	1.065
$\sqrt{2}b$	0.5	3.745	1.993	1.316
$\sqrt{6}b$	0.75	4.750	3.219	1.993
$\infty$	1.00	5.000	4.750	3.745
		$\sqrt{2}b$		$\sqrt{6}b$
$c$	$\mu$	$\lambda \ 0$	0.25	0.5
0	0	1.040	1.040	1.000
$\sqrt{\frac{2}{3}}b$	0.25	1.009	1.009	1.040
$\sqrt{2}b$	0.5	1.133	1.133	
$\sqrt{6}b$	0.75	1.386	1.386	
$\infty$	1.00	2.440	2.440	

TABLE 2

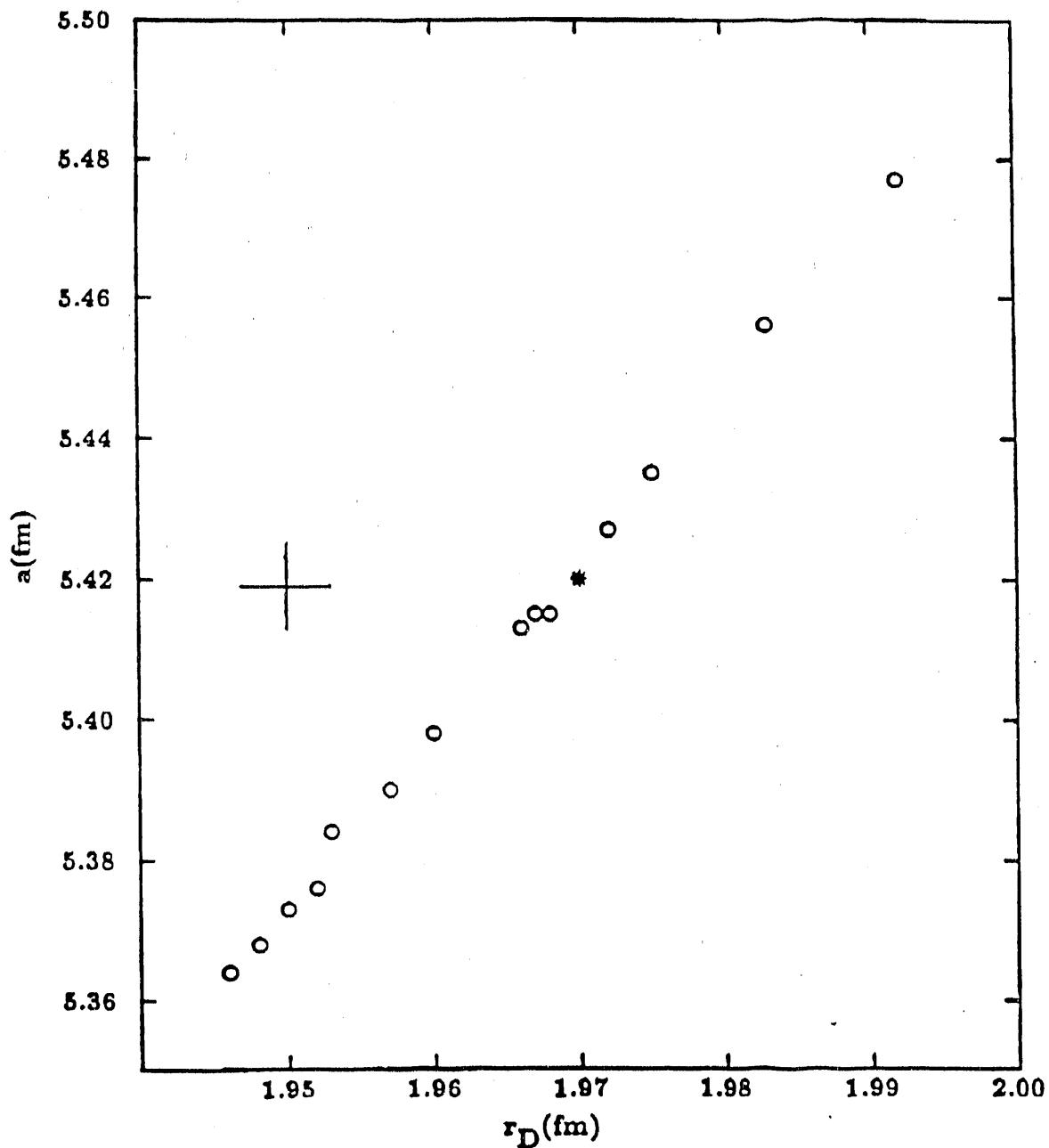
Illustrating that non-locality speeds up the convergence of the model space expansion.

$a = \sqrt{\frac{2}{3}}b$	$\lambda = \frac{a^2}{a^2+2b^2} = \frac{1}{4}$	$\mu = \frac{c^2}{c^2+2b^2}$	0	$\frac{1}{4}$	$\frac{1}{2}$	$\frac{3}{4}$
$I_{20}/I_{00}$	$= \frac{3}{2} (1 - \lambda - \mu)^2 + (\lambda - \mu)^2$		0.91	0.38	0.16	0.25
$I_{00-20}/I_{00}$	$= \sqrt{\frac{3}{2}} (1 - \lambda - \mu)$		0.92	0.61	0.31	0.00
$I_{00-40}/I_{00}$	$= \sqrt{\frac{15}{8}} (1 - \lambda - \mu)^2$		0.77	0.34	0.08	0.00
$I_{00-60}/I_{00}$	$= \sqrt{\frac{35}{16}} (1 - \lambda - \mu)^3$		0.62	0.18	0.02	0.00

FIGURE 1

Calculated triplet scattering length  $a$  and deuteron radius  $r_D$  for realistic NN potentials, and comparison with empirical value. From Ref. 1.

Calculated  $a$  and  $r_D$  for realistic potentials (Klarsfeld et al.)



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