

A UNIVERSAL FORMULA FOR THE QUASISTATIC SECOND-ORDER DENSITY

PERTURBATION BY A COLD MAGNETOPLASMA WAVE*

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Using the general expression for the ponderomotive Hamiltonian, we obtain the quasi-static quasi-neutral density change caused by the ponderomotive force of a cold magnetoplasma wave of arbitrary frequency and polarization:

$$\delta n(\underline{x}) = - \frac{|\tilde{E}(\underline{x})|^2 - |\tilde{B}(\underline{x})|^2}{4\pi(T_e + T_i)}.$$

This formula agrees with and extends previous results for unmagnetized and magnetized plasma.

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In studying the modulation of a finite-amplitude plasma wave, a number of authors have calculated the quasi-static quasi-neutral second-order density perturbation produced by the ponderomotive force of the modulation. With the representation

$$\phi(\underline{x}, t) = \phi(\underline{x}) \exp(-i\omega t) + \text{c.c.} \quad (1)$$

for a longitudinal magnetoplasma wave, the result¹

$$\delta n(\underline{x}) = - \frac{|\nabla \phi(\underline{x})|^2}{4\pi(T_e + T_i)} \quad (2)$$

has been obtained by Morales and Lee² for lower-hybrid waves, and by Shukla³ for electron magnetoplasma waves. The former authors remarked on the identity of formula (2) with the familiar expression for Langmuir wave modulation in unmagnetized plasma.

It is natural to inquire into the universality of formula (2). In this paper, we show that it does indeed apply to any longitudinal cold-plasma wave (for a single ion species⁴); i. e., the three solutions⁵ $\omega(\theta)$ of $\epsilon_L(\omega, \theta) = 0$, where $\epsilon_L \equiv \hat{k} \cdot \epsilon(\omega) \cdot \hat{k}$.

More importantly, we show that formula (2) can be simply generalized to apply to a cold plasma wave of any polarization, i. e., to a wave with non-zero $\nabla \times \underline{E}$. Here we use a local plane-wave⁶ representation

$$\underline{E}(\underline{x}, t) = \underline{\tilde{E}}(\underline{x}) \exp(i\underline{k} \cdot \underline{x} - i\omega t) + \text{c. c.}, \quad (3a)$$

$$\text{with} \quad \underline{\tilde{B}}(\underline{x}) = (c/\omega)\underline{k} \times \underline{\tilde{E}}. \quad (3b)$$

The generalization, derived below, is

$$\delta n(\underline{x}) = - \frac{|\underline{\tilde{E}}(\underline{x})|^2 - |\underline{\tilde{B}}(\underline{x})|^2}{4\pi(T_e + T_i)} \quad (4)$$

We note first that it reduces to (2) when $\underline{\tilde{B}} = 0$. Secondly, for the transverse unmagnetized case, where

$$|\underline{\tilde{B}}|^2 = (kc/\omega)^2 |\underline{\tilde{E}}|^2 = (1 - \omega_p^2/\omega^2) |\underline{\tilde{E}}|^2,$$

formula (4) becomes $\delta n/n = -(e^2/m\omega^2) |\underline{\tilde{E}}|^2/(T_e + T_i)$, the familiar result.⁷

Formula (4) can be used for any cold-magnetoplasma wave, e. g., lower hybrid in the electromagnetic region⁸, fast-magnetosonic-whistler⁹, Alfvén¹⁰, ordinary and extraordinary, etc., so long as (3b) is a valid approximation. (When it is not, use formula (10) below.)

Our derivation begins with the standard expression¹¹ for the quasi-static density perturbation, of species s , caused by the ponderomotive potential energy $\Psi_s(\underline{x})$ of an oscillation center¹² and by the

self-consistent electric potential $\phi(\underline{x})$:

$$\frac{\delta n_s(\underline{x})}{n_s^0} = - \frac{\psi_s(\underline{x}) + e_s \phi(\underline{x})}{T_s} \quad (5)$$

For two species (electrons and singly-charged ions), we impose quasi-neutrality ($\delta n_e = \delta n_i$, $n_e^0 = n_i^0$) to eliminate ϕ , and obtain the relation

$$\frac{\delta n(\underline{x})}{n^0} = - \frac{\psi_e(\underline{x}) + \psi_i(\underline{x})}{T_e + T_i} \quad (6)$$

Our expression for $\psi_s(\underline{x})$ is based on a useful relation¹³ for the ponderomotive Hamiltonian¹⁴ of an oscillation center. In the cold-species limit, Eq. (3) of Ref. (13) reduces to

$$n_s(\underline{x}) \psi_s(\underline{x}) = - (4\pi)^{-1} \underline{E}^*(\underline{x}) \cdot \chi_{\omega}^s(\underline{x}) \cdot \underline{E}(\underline{x}), \quad (7)$$

with the representation $\underline{E}(\underline{x}, t) \equiv \underline{E}(\underline{x}) \exp(-i\omega t) + \text{c. c.}$, where χ_{ω} is the well-known¹⁵ cold-species susceptibility. (We note that χ is proportional to density, so that ψ is density-independent; but the dependence of χ on possibly nonuniform magnetic field $\underline{B}_0(\underline{x})$ appears in ψ .)

Inserting (7) into (6), we have

$$\delta n(\underline{x}) = \frac{\underline{E}^*(\underline{x}) \cdot (\chi_{\omega}^e + \chi_{\omega}^i) \cdot \underline{E}(\underline{x})}{4\pi(T_e + T_i)} \quad (8)$$

Now we use the field equation

$$(\chi_{\omega}^e + \chi_{\omega}^i) \cdot \underline{E}(\underline{x}) = -\underline{E}(\underline{x}) + (ic/\omega)\nabla \times \underline{B}(\underline{x}), \quad (9)$$

where $\underline{B}(\underline{x}) = (c/i\omega)\nabla \times \underline{E}(\underline{x})$, to obtain

$$\delta n(\underline{x}) = - \frac{|\underline{E}(\underline{x})|^2 - |\underline{B}(\underline{x})|^2 - (c/\omega)\text{Im} \nabla \cdot \underline{E}^*(\underline{x}) \times \underline{B}(\underline{x})}{4\pi(T_e + T_i)} \quad (10)$$

Finally, for a local plane wave, with $\underline{E}(\underline{x}) = \tilde{\underline{E}}(\underline{x}) \exp i\mathbf{k} \cdot \underline{x}$ and (3b), one may drop the complex Poynting term in (10), as higher order in $kV \ll \tilde{E}$; the result is then Eq. (4).

Two points should be kept in mind in applying (4): second-order magnetic perturbations may be of significance¹⁶; and the quasi-static assumption may be invalid.¹⁷

Footnotes and References

1. The numerical factor in the denominator is sometimes given incorrectly as 8π . If, instead of (1), one uses $\phi(\underline{x}, t) = \text{Re } \phi(\underline{x}) \exp(-i\omega t)$, the factor should be 16π .
2. G. J. Morales and Y. C. Lee, Phys. Rev. Lett. 35, 930 (1975).
3. P. K. Shukla, J. Plasma Phys. 18, 249 (1977).
4. For more than one ion species, formulas (2) and (4) generalize to less beautiful forms.
5. We note that for the lowest-frequency solution (ion-cyclotron wave), the cold plasma model may be invalid.
6. More correctly, $\exp i \underline{k} \cdot \underline{x} \rightarrow \exp i \Theta(\underline{x})$, with $\underline{k}(\underline{x}) = \nabla \Theta$.
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