


Article

Performance of the ICARUS Trigger System at the Booster and NuMI Neutrino Beams [†]

 Riccardo Triozzi on behalf of the ICARUS C ollaboration 

INFN Sezione di Padova, Università di Padova, 35122 Padova, Italy; riccardo.triozzi@pd.infn.it

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Abstract: The ICARUS-T600 liquid argon time projection chamber detector takes data at a shallow depth as the far detector of the Short Baseline Neutrino program at Fermilab, searching for sterile neutrinos with the Booster and Main Injector neutrino beams. The ICARUS trigger system exploits the temporal coincidence of the beams with scintillation light signals detected by 360 photo-multiplier tubes in limited TPC regions. The trigger efficiency measurement leverages cosmic rays collected without any scintillation light requirement, with timing from an external cosmic ray tagger system. The efficiency measured with stopping muons roughly saturates at $E_\mu \sim 300$ MeV, covering most of the expected energy range of charged-current neutrino interactions. For the latest ICARUS physics runs, special “adder” boards performing the analog sum of light signals were introduced as a complementary trigger to possibly recover low-energy neutrino interactions.

Keywords: neutrino physics; liquid argon; TPC; trigger


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1. The ICARUS Detector at Fermilab

The ICARUS-T600 liquid argon (LAr) time projection chamber (TPC) detector [1] currently takes data at a shallow depth as the far detector of the Short Baseline Neutrino (SBN) program at Fermilab [2], searching for sterile neutrinos with muon and electron neutrinos from the 0–2 GeV Booster Neutrino Beam (BNB) and the off-axis 0–5 GeV Neutrinos at the Main Injector (NuMI) beam. In recent years, several experiments reported evidence of anomalous neutrino oscillations (e.g., LSND [3] and MiniBooNE [4]) suggesting the existence of a fourth sterile neutrino state associated with a large $\Delta m^2 \sim 1$ eV² mass splitting, which would produce oscillations at short baselines. Several constraints on sterile neutrinos exist, including ICARUS measurements from data collected at the INFN Gran Sasso National Laboratories in Italy [5]. ICARUS is also pursuing a rich stand-alone program focused on neutrino oscillations, cross-sections, and going beyond the standard physics model.

The ICARUS detector, after operating deep underground at the Gran Sasso Laboratories until 2013, underwent an intense overhaul at the CERN and INFN laboratories [1] before starting collecting data at Fermilab in 2020. Refurbishing activities improved the cryogenics and purification systems, implemented new TPC read-out electronics, and upgraded the liquid argon light detection system with respect to the LNGS configuration [1,6].

This detector consists of two identical modules filled with 476 t of active liquid argon, each hosting two 1.5 m-drift LAr-TPCs separated by a shared cathode with $\sim 58\%$ transparency at a nominal 500 V/cm electric drift field (Figure 1). The TPC anodes consist of three readout wire planes with 3 mm pitch and 3 mm apart, oriented at 0 and

$\pm 60^\circ$, respectively, with respect to the BNB direction, where the first two Induction-1 and Induction-2 planes measure the ionization charge non-destructively, and the last Collection plane collects the charge.

Due to the shallow depth, ICARUS faces a challenging ~ 11 kHz cosmic ray rate. To reduce the hadron and photon cosmic ray component, a 3 m concrete overburden was deployed. In addition, the detector is surrounded by a segmented plastic scintillator Cosmic Ray Tagger (CRT) system, providing $\mathcal{O}(1\text{ ns})$ timing of crossing particles.

Scintillation light produced by ionizing particles in liquid argon [7,8] is detected by 360 8" photo-multiplier tubes (PMT) located behind the TPC wire planes [9], with 90 PMTs per TPC, providing a prompt signal used for the trigger system and for timing purposes.

The evolution of the trigger system through the initial physics runs, from Run1 (9–12 June 2022), through Run2 (20 December 2022–10 June 2023), to recent physics runs, is described in the following.

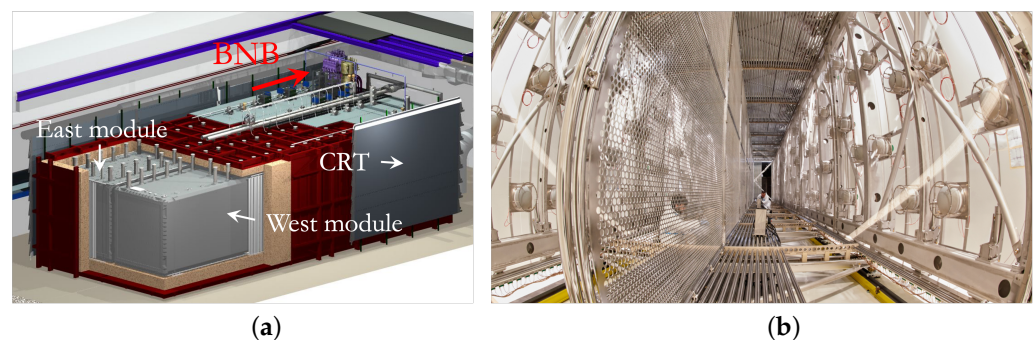


Figure 1. (a) Representation of the ICARUS modules inside the warm vessel, surrounded by the segmented CRT detector. A 3 m concrete overburden sits above the detector to reduce the cosmic ray background. (b) Picture of an ICARUS TPC during the overhauling activities, with the semi-transparent pierced cathode visible in the middle. On the right, the 8" PMTs are installed behind the anodic wire planes.

2. The ICARUS Trigger System

The ICARUS trigger system is online, fully implemented on hardware and FPGA programmable logic, and based on coincident light signals.

Light signals collected by the PMTs are continuously digitized and stored in 10 μs memory buffers by 500-MSa/s 16-channel digitizers [9], which also acquire reference timing signals from the accelerators. The 10 μs PMT acquisition window includes 3 μs of pre-trigger and 7 μs of post-trigger to accommodate for both the fast ($\tau \simeq 6\text{ ns}$) and slow ($\tau \simeq 1.6\text{ }\mu\text{s}$) components of scintillation light in liquid argon [7]. Each digitizer records signals from 15 PMTs in a 3 m longitudinal TPC slice, pairing them into 7 pairs (in logical OR) and 1 singlet low-voltage differential signal (LVDS) outputs activated by a light signal exceeding a digital 13-photoelectron threshold.

The digital LVDS signals in the west and east modules are fed into the trigger electronics, where two FPGAs ("West" and "East") evaluate at each clock cycle (25 ns) the LVDS patterns in the two modules according to a pre-defined trigger logic. The third "Global" FPGA combines the BNB and NuMI proton extraction signals with the West- and East-FPGA evaluation outcome, electing the first trigger of the two as the global trigger.

After the global trigger is issued, the Global-FPGA alerts the trigger electronics to start the readout of the TPC for 1.6 ms, of the CRT for 6 ms, and of all PMT signals for 26 μs around the trigger. The PMT acquisition window is enlarged from 10 μs to 26 μs for triggers in coincidence with the beam gate.

Considering the $\mathcal{O}(1\text{ GeV})$ energy scale of BNB and NuMI neutrinos, interactions are on average expected to be contained in 4 m longitudinally, motivating a trigger logic based

on limited TPC regions. To generate the “PMT-majority” global trigger, at least five LVDS signals are required in a 6 m longitudinal module slice in temporal coincidence with the BNB or NuMI beam gate windows (“majority-5”/“mj-5” setting). Each 6 m window contains 60 PMTs (30 per TPC, front-facing). For the initial Run1, three side-by-side slices were considered. The trigger was then improved in efficiency and uniformity in Run2 by introducing two additional overlapped windows (Figure 2).

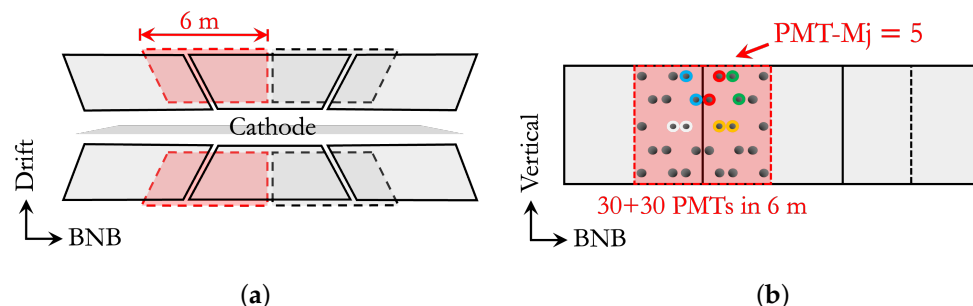


Figure 2. (a) View from the top of an ICARUS module, with two LAr-TPCs separated by a common cathode. For the trigger logic, light signals are evaluated in three side-by-side 6 m longitudinal module slices. For Run2, two additional slices (dashed lines) were introduced to improve efficiency and uniformity. (b) Side-view of the module. In the highlighted window, containing 60 PMTs (30 per TPC), an example of a mj-5 pattern is shown, with PMTs paired with LVDS channels.

A diagram of the FPGA implementation for the PMT-majority trigger is provided in Figure 3.

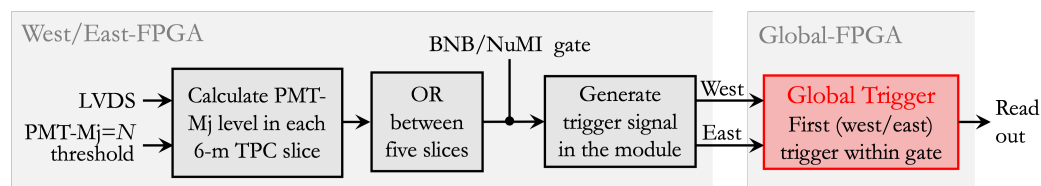


Figure 3. Diagram of the FPGA implementation for the PMT-majority trigger. At each clock cycle, the two West- and East-FPGAs evaluate the LVDS patterns in the corresponding module, calculating the multiplicity in each TPC slice, as defined in Figure 2. Each FPGA generates a trigger signal when the requirements are met within the BNB or NuMI gates. The Global-FPGA elects the first such trigger as the global trigger, activating the readout of the ICARUS sub-detectors.

Several triggers are routinely collected with both the BNB and NuMI beams. Triggers with the described PMT-majority logic are collected both “on-beam”, in coincidence with the beam spill extractions, and “off-beam”, where a gate signal is opened between beam spills to collect cosmic ray background statistics. Additionally, “minimum-bias” triggers can be acquired by just requiring that a gate is opened, without any request a priori on the scintillation light in the event. During data taking, minimum-bias triggers are collected every 200 (60) on-beam gates for BNB (NuMI) and every 20 off-beam gates. Minimum-bias data are used for measuring the trigger efficiency, for calibrations, and detector physics studies.

The ICARUS trigger was designed to select physical interactions within the BNB and NuMI beam spills and is observed to reject $\simeq 97\%$ of the spills that are either empty or containing negligible activity. Figure 4 shows the global trigger time with respect to the BNB (left) and NuMI (right) beam gate opening times for Run2 data ($\simeq 2 \times 10^{20}$ POT for BNB, $\simeq 2.7 \times 10^{20}$ POT for NuMI). The beam gate durations are slightly larger than the actual beam spills (respectively, 1.6 μ s and 9.6 μ s for BNB and NuMI). Beam-related activity (mainly neutrinos in the active volume, but also beam–halo interactions spilling into the detector) produces an evident excess in the on-beam data with respect to the off-beam data.

The relative fraction is higher for NuMI, which has higher beam intensity. The coarse NuMI beam structure consisting of 6 batches with $\simeq 100$ ns spacing is also observed. The peak in triggers at the gate opening is produced by energetic cosmic rays crossing the detector right before the beam gates, whose late scintillation light is enough to meet the PMT-majority trigger requirements, hence producing a global trigger.

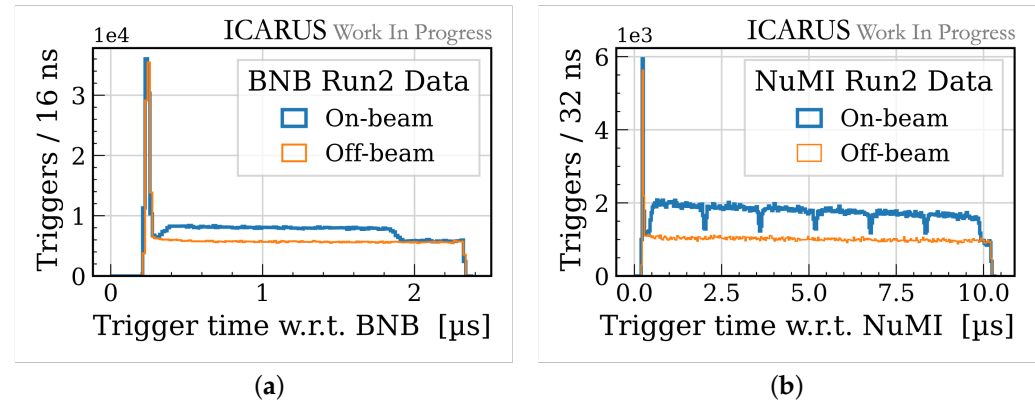


Figure 4. Profile of trigger times with respect to the BNB (a) and NuMI (b) beam gate opening times. Beam-related activity produces an excess of triggers in on-beam data with respect to off-beam data. The coarse NuMI beam structure, divided into 6 batches, is reconstructed at the trigger level.

2.1. Performance of the Trigger System

To study the performance of the trigger system, the trigger logic is emulated starting from the recorded PMT waveforms in off-beam minimum-bias data. To avoid biasing the trigger efficiency measurement, timing for cosmic ray particles is provided by the CRT, performing a spatial match between hits on the CRT scintillators and reconstructed tracks in the TPCs.

In particular, muons stopping inside the detector are selected (Figure 5), as they are similar in topology and energy to BNB muon neutrino charged-current interactions, and their energy can be precisely estimated from the residual range. To select the Bragg peak signature of stopping muons, a request is made on the median dE/dx local energy deposition at the end of the track to be larger than 4 MeV/cm. In addition, a calorimetric energy of at least 60 MeV is requested in the last 20 cm, based on the Bethe–Bloch prediction for muons stopping in liquid argon. More information on ICARUS calibrations for precise calorimetry can be found in [10].

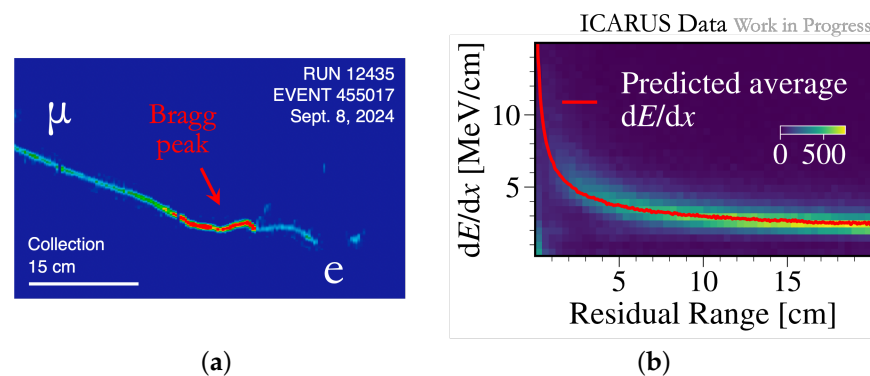


Figure 5. (a) Event display on the Collection plane of a stopping muon decaying into a Michel electron, showing the wires on the horizontal axis and the drift time in the vertical axis. The color scale indicates the collected charge, with red being higher. (b) Local energy deposition as a function of the residual range for the selected stopping muon tracks. Data are in good agreement with the Bethe–Bloch prediction, excluding the point at the lowest residual range, for which the local energy deposition estimation is unreliable given the uncertainty regarding the exact stopping position of the particle.

The majority-5 trigger efficiency measurement with stopping muons for Run2 is reported in Figure 6 as a function of the muon energy estimated from the range (>70 MeV for muons at least 20 cm long). The ICARUS trigger is close to full efficiency above roughly 300 MeV, with the plateau covering the spectrum of charged-current neutrino interactions from the BNB and the higher-energy NuMI beams. Related systematic uncertainties mainly come from the track reconstruction and selection and from the emulation algorithm, amounting to $\simeq 10\%$ below 200 MeV.

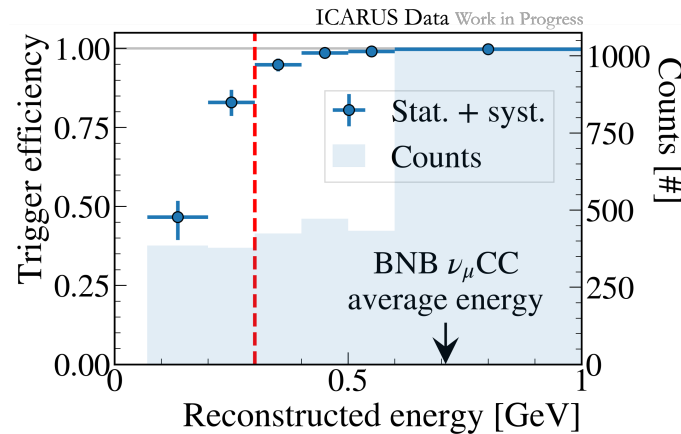


Figure 6. Trigger efficiency with stopping muons as a function of the muon energy estimated from the residual range for the Run2 settings (at least 5 PMT-pairs in one of five overlapped 6 m module slice). Statistical and systematic errors are included. The corresponding energy spectrum is reported in the light-blue histogram. For reference, the average simulated deposited energy in muon neutrino charged-current interactions is shown.

2.2. Cosmic Ray Rejection

ICARUS is operating at the Earth’s surface, facing a challenging $\simeq 11$ kHz cosmic ray background while searching for neutrinos from BNB and NuMI.

During the $\simeq 1$ ms drift of ionization electrons from beam events, tens of overlapping cosmic ray interactions are acquired. Such cosmic ray particles cross the detector before and after the beam spills, and can be rejected with timing. Scintillation light can provide a prompt signal to determine the time of interactions. Indeed, it would be challenging to acquire several ms of mostly baseline PMT data. To avoid saturating the data acquisition system, additional “out-of-time” triggers are acquired: in Run2, when a mj-9 trigger condition is met in a 2 ms time window around the beam spills, 10 μ s of PMT data are read out from the 180 PMTs in the corresponding module.

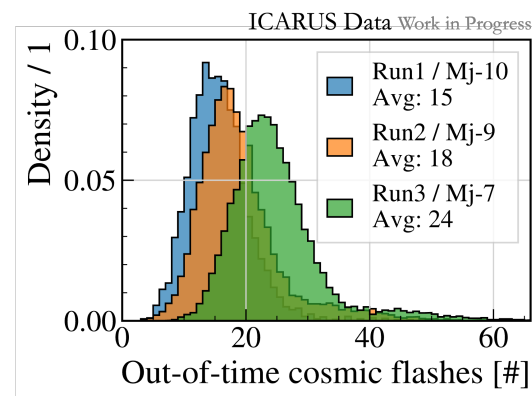


Figure 7. Number of out-of-time reconstructed light flashes per event in the three ICARUS run configurations. The substantial improvement in the collection of out-of-time triggers leads to an improved cosmic background rejection.

The out-of-time PMT-majority requirement was lowered from mj-10 in Run1, through mj-9 in Run2, down to a mj-7 for the latest Run3 (started in March 2024), hence improving the cosmic background rejection potential. Figure 7 shows the number of out-of-time reconstructed scintillation light flashes (i.e., clusters of PMTs fired in a module in a 1 μ s window) per event in the three ICARUS run configurations.

2.3. Adders

An independent trigger system based on the total amount of scintillation light rather than a multiplicity of fired PMTs is being developed to operate in parallel to the well-established PMT-majority trigger, possibly improving the event recognition efficiency at low energies.

Dedicated custom “adder” boards were developed, receiving signals from 15 PMTs in a 3 m TPC slice (corresponding to the same group of signals going into a digitizer) and splitting them into two 95% and 5% components. While the major signal component proceeds to the corresponding digitizer and is used for the PMT-majority trigger, the 5% components from the 15 PMTs are summed up analogically. The 24 adder signals from the 360 PMTs (12 adder signals per TPC) are proportional to the total deposited light in 3 m TPC slices. Adder signals are discriminated with a 60 mV threshold, chosen as a compromise between noise levels and maintaining sustainable trigger rates.

An adder global trigger is produced by requiring at least one adder signal over the threshold. For Run3, the adder trigger was introduced in logic-OR with the PMT-majority trigger, being able to distinguish between the two trigger sources at the hardware level (Figure 8). A global trigger rate of $\simeq 0.8$ Hz was obtained.

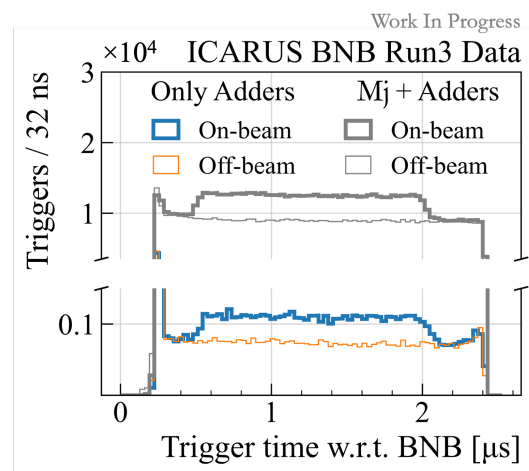


Figure 8. Profile of trigger times with respect to the BNB beam gate opening time for all triggers collected during Run3, where the PMT-majority is in logic-OR with the adder trigger. The contribution from the adder trigger is highlighted, with several beam-related triggers collected.

For out-of-time triggers, to avoid overcrowding the DAQ, only the four adders at the southernmost and northernmost corners (with respect to the BNB direction) in each module are considered, where the multiplicity-based logic is more inefficient due to border effects. A cosmic ray rate of $\simeq 11$ kHz is measured from the collected light, approaching the physical rate expected at the Earth’s surface.

Adders provide complementary information with respect to the multiplicity-based trigger logic, as interactions close to the PMTs can produce large scintillation light signals triggering PMTs, potentially not meeting the mj-5 requirement.

3. Conclusions

The ICARUS trigger, based on prompt scintillation light signals in liquid argon collected by 360 PMTs, is online and fully implemented on FPGA-based hardware. The main trigger is based on the multiplicity of PMT signals within limited TPC regions in coincidence with the proton extraction signals from the Booster and Main Injector accelerators.

The trigger efficiency, studied with stopping muons from cosmic ray minimum-bas data, is shown to saturate roughly at $E_\mu \sim 300$ MeV, with the plateau effectively covering the spectrum of BNB and NuMI charged-current interactions.

To improve the event recognition efficiency for low-energy events, a complementary trigger system was developed, based on the total amount of scintillation light rather than a multiplicity. The system, based on custom “adder” boards, was deployed and included with the main PMT-majority trigger for the latest physics runs, where several additional beam-induced interactions were collected.

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Abbreviations

The following abbreviations are used in this manuscript:

BNB	Booster Neutrino Beam
CRT	Cosmic Ray Tagger
LAr-TPC	Liquid Argon Time Projection Chamber
LVDS	Low-Voltage Differential Signal
NuMI	Neutrinos at the Main Injector
PMT	Photo-Multiplier Tube

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