

Measurement of Dielectric Loss in Silicon Nitride at Centimeter and Millimeter Wavelengths

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Abstract—This work presents a suite of measurement techniques for characterizing the dielectric loss tangent across a wide frequency range from ~ 1 GHz to 150 GHz using the same test chip. In the first method, we fit data from a microwave resonator at different temperatures to a model that captures the two-level system (TLS) response to extract and characterize both the real and imaginary components of the dielectric loss. The inverse of the internal quality factor is a second measure of the overall loss of the resonator, where TLS loss through the dielectric material is typically the dominant source. The third technique is a differential optical measurement at 150 GHz. The same antenna feeds two microstrip lines with different lengths that terminate in two microwave kinetic inductance detectors (MKIDs). The difference in the detector response is used to estimate the loss per unit length of the microstrip line. Our results suggest a larger loss for SiN_x at 150 GHz of $\tan \delta \sim 4 \times 10^{-3}$ compared to 2.0×10^{-3} and $\gtrsim 1 \times 10^{-3}$ measured at ~ 1 GHz using the other two methods. These measurement techniques can be applied to other dielectrics by adjusting the microstrip lengths to provide enough optical efficiency contrast and other mm/sub-mm frequency ranges by tuning the antenna and feedhorn accordingly.

Index Terms—Dielectrics loss, silicon nitride (SiN_x), millimeter wavelength, two-level system (TLS), quality factor

I. INTRODUCTION

Dielectric materials are used extensively in mm-wave superconducting detectors for astrophysical or cosmological surveys. Experiments with photometer arrays such as CMB-S4 use microstrip lines to transfer the mm-wave signal received at the antenna to the detectors, where dielectric loss limits the overall optical efficiency [1]. Dielectric loss is more critical for

on-chip spectroscopy, where it constrains the quality factor of the on-chip filter bank, which limits the combination of spectral resolving power and the optical efficiency [2]. The lumped-element MKID architecture typically utilize an interdigitated capacitor, but adapting to parallel-plate capacitors can significantly reduce the size of the resonators and increase their packing density. However, one disadvantage associated with parallel-plate capacitors is that more of the power is confined to the thin-film dielectric layer; thus, the loss mechanism in dielectrics will limit detector quality factor and increase the noise. In quantum information, the dielectric loss can cause decoherence [3]. The defining roles of low-loss dielectrics in multiple applications motivate this study.

A few low-loss dielectrics have been developed and are commonly in use, including SiO_2 [4], SiN_x [5], amorphous Si (a-Si) [6], and silicon carbide [7]. State-of-art loss tangent of $O(10^{-5})$ has been achieved for a-Si [6] and silicon carbide [7] at microwave frequencies. However, past measurements mostly focused on a few GHz, while applications for mm-wave astrophysical surveys deal with radiation at millimeter wavelengths. Though a few measurements at mm-wave have been reported, more data is needed to validate the dielectric performance across different deposition techniques at mm-wave and cross-compare with measurements at a few GHz. Given the discrepancy between cm and mm-wavelengths that a few groups observed for SiN_x , hydrogenated amorphous Si and SiC [8], [9], [10], more work is needed to study the discrepancy and elucidate other potential loss mechanisms at mm-wavelengths beyond the dielectrics.

This paper presents our recent effort to measure the dielectric loss from centimeter to millimeter wavelengths. We designed and fabricated a test structure that can probe dielectric loss using three methods: 1) fitting frequency vs. temperature response to a TLS model, 2) inverting the quality factor in the low power limit, and 3) measuring the differential loss for detectors coupled via different lengths of microstrip transmission lines. Compared to the Fabry-Pérot method in [10], our method doesn't require a coherent source that can sweep in frequency and only needs an internal blackbody to feed the antenna. The detector chip has a simple design but offers three different ways to measure the dielectric loss. In a previous paper [11], we presented a design where the signal from a paddle of the ortho-mode transducer (OMT) probe is split into two transmission lines with different lengths and coupled to two detectors. However, the detectors did not show a noticeable optical response. We then simplified the optical coupling and fed more power to each detector

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by differentiating the signal from a pair of OMT paddles and removing the power split. The new strategy showed the expected optical coupling level. Here, we report on progress made having resolved several technological issues since [11]: 1) developed a more robust electrical contact between the aluminum inductor and niobium capacitor, 2) solving the step coverage when the transmission line climbs up onto the capacitor pads, 3) fixed unwanted shorts from niobium etching residuals, and 4) improving parasitic leakage of the optical signal from neighbouring detectors.

II. DEVICE DESIGN AND FABRICATION

Measurement devices optimised for mm-wave were fabricated using a similar process to that reported in [11] but with a few key differences. Devices start as a 150 mm silicon wafer with a commercially-grown 450 nm thermal SiO_2 and 2 μm low-pressure chemical-vapor deposited (LPCVD) SiN_x bilayer. These layers form the released membrane for the OMT probes. A 300 nm thick Nb ground plane is deposited and patterned by fluorine-based reactive ion etch (RIE). Next, a 500 nm-thick layer of SiN_x is deposited for the microstrip dielectric. The SiN_x was deposited by chemical vapor deposition using SiH_4 and N_2 precursors at a substrate temperature of 275°C. Following the microstrip dielectric, the microstrip wiring is fabricated from a 170 nm-thick Nb layer, along with a 30 nm NbN capping layer and patterned via liftoff. Next, a 50 nm Al layer is deposited and patterned into inductors via liftoff. Previously we had deposited the Al inductors before the Nb capacitors. However, due to issues with poor galvanic contact between the layers, we reversed the steps and added the NbN capping layer. The NbN limits oxidation of the Nb, making a galvanic contact to the Al more robust. Additionally, in previous devices, the Nb wiring layer was etched, which occasionally leads to shorting when the material is not removed from step sidewalls by the anisotropic RIE. Switching to using liftoff eliminated this issue. After the wiring layer processing, the OMT membrane is released by a silicon deep RIE from the backside of the wafer. We use the same etch step to separate individual chips from the wafer. The final step is an O_2 plasma used to remove a protective resist layer from the front side of the wafer.

Fig. 1 bottom and top show the chip design and a microscope image of a segment near a pixel. A coplanar waveguide (CPW) feedline is capacitively coupled to 20 MKIDs detectors, each containing a parallel-plate niobium capacitor and a meandered aluminum inductor line. We choose parallel-plate capacitors over interdigitated capacitors (IDCs) due to their higher energy filling factor F that is close to unity. Four of the 20 resonators are coupled to optical signals from the OMT via hybrid tee and different lengths of microstrip meander lines. At mm-wave, the dielectric loss can be extracted by comparing the optical response of detectors connected through varying lengths of microstrip transmission lines. In [11], we split the signal from one arm of the OMT into two detectors, which reduced the optical signal level to be comparable to systematics caused by temperature variation and noise scatter. To improve this, we increase the signal level by a factor of

four by removing the signal split and differentiating signals from two OMT fins with a hybrid tee. Other detectors should be dark, meaning they do not receive signals from the antenna via microstrip. The primary purpose of the dark detectors is to monitor systematics from temperature variation and optical leakage.

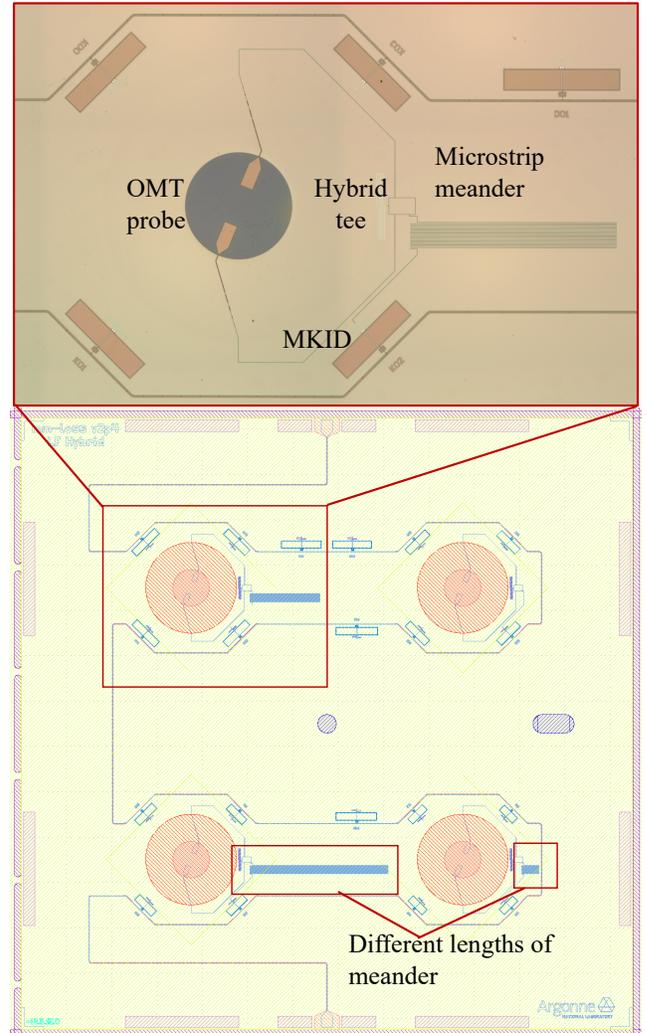


Fig. 1: Detector chip design (bottom) and the microscope image for one pixel (top). A CPW feedline couples to 20 MKIDs, out of which four receive optical from OMT probes through hybrid tees and different lengths of microstrip transmission lines. Other detectors do not directly connect to the antenna and are references for systematics and dark pickup.

III. DIELECTRIC LOSS FROM TLS MODEL AT 1-2 GHz

The two-level system (TLS) modeling in amorphous solids was first introduced by [12] and [13], where groups of atoms may tunnel between sites corresponding to different energy states. The dielectric modification by TLS can be derived assuming a log-uniform distribution of tunneling states, and the subsequent shift in resonant frequency for a resonator is given by [14]:

$$\frac{f(T) - f_0}{f_0} = \frac{F\delta_{\text{TLS}}^0}{\pi} \left[\text{Re}\Psi \left(\frac{1}{2} - \frac{\hbar f_0}{jk_B T} \right) - \log \frac{\hbar f_0}{k_B T} \right], \quad (1)$$

where $f(T)$ is the resonator frequency at temperature T , f_0 is the TLS-free resonance frequency, F is the energy filling factor in the amorphous dielectrics, δ_{TLS}^0 is the intrinsic TLS loss, Ψ is the digamma function, and k_B is the Boltzmann constant. Typically F and δ_{TLS}^0 are degenerate when fitting $f(T)$ vs. T to the model, but we eliminated the degeneracy in our system using parallel-plate capacitors with filling factors F close to one. An increase in stage temperature can break Cooper pairs and change the kinetic inductance, resulting in a downshift of the resonance frequency [15]. But this effect takes over after around 300 mK for aluminum and is beyond the range of our study.

We measured the resonance frequencies for our resonators using a standard vector network analyzer (VNA) setup with low-temperature attenuators and amplifiers and a room-temperature last-stage amplifier. A sample measurement was shown in Fig. 2 for a detector with SiN_x dielectric. We fit for the model in Eq. (1) to extract $F\delta_{\text{TLS}}^0$. The fit result of δ_{TLS}^0 is 2.2×10^{-3} for this resonator, which is representative of detectors on the same chip. Fig. 3 summarizes the distribution of $F\delta_{\text{TLS}}^0$ from two chips, which corresponds to the two groups in the histogram. $F\delta_{\text{TLS}}^0$ within each chip is uniform within $\pm 10^{-4}$ and the systematic shift between the two chips is also at the 10^{-4} level. To conclude, this measurement yields $\delta_{\text{TLS}}^0 = (2.0 \pm 0.2) \times 10^{-3}$ for SiN_x .

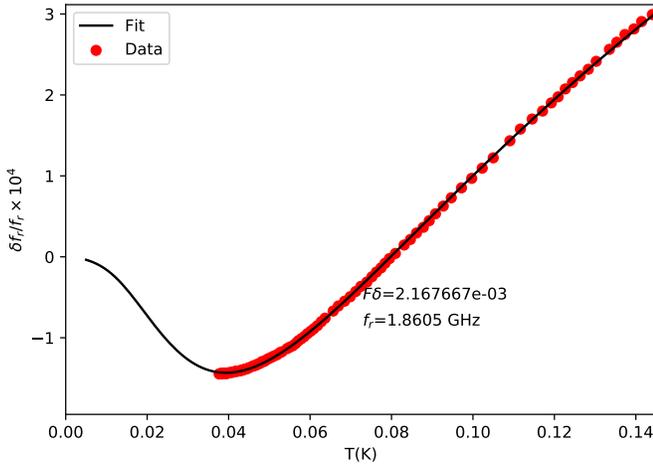


Fig. 2: Resonator frequency shift $f(T)$ vs. temperature T and the fit to a TLS model. The fit $F\delta_{\text{TLS}}^0$ is 2.2×10^{-3} for this resonator.

IV. DIELECTRIC LOSS FROM $1/Q_{\text{TLS}}$ AT 1-2 GHz

We can achieve quality factors at a few hundred thousand for resonators using interdigitated capacitors but only get quality factors at $O(10^3)$ level for resonators with parallel-plate capacitors and the same metal targets. This result suggests that most loss comes from TLS in SiN_x dielectric layer and that

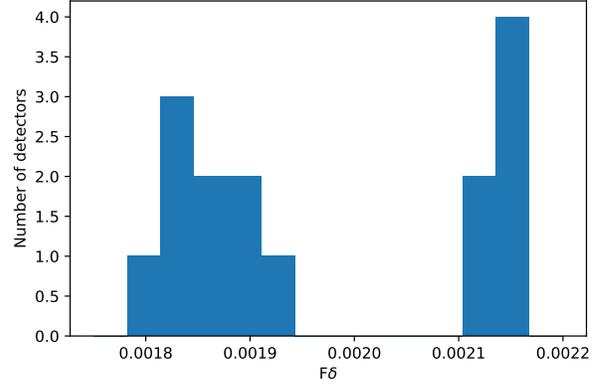


Fig. 3: Histogram showing the measured $F\delta_{\text{TLS}}^0$ for all resonators fitting to the TLS model. The two groups belong to two different detector chips. Albeit the small difference, the two chips both show $F\delta_{\text{TLS}}^0$ at the 2×10^{-3} level. The statistical error within each chip is within 10^{-4} .

the internal quality factor Q_i is limited by TLS loss. We can then use $1/Q_i$ to approximate dielectric loss.

However, one caveat for this measurement is that the driving power can saturate the TLS states and lead to underestimated $F\delta_{\text{TLS}}^0$. From [14], the resonator loss δ_{TLS} depends on power and temperature:

$$\delta_{\text{TLS}} = F\delta_{\text{TLS}}^0 \frac{\tanh\left(\frac{\hbar\omega}{2k_B T}\right)}{\sqrt{1 + \frac{\langle n \rangle}{n_c}}}, \quad (2)$$

where $\langle n \rangle$ is the average photon number in the resonator, and n_c is the TLS saturation photon number. $\langle n \rangle$ is proportional to power, so $1/Q_i \sim 1/Q_{\text{TLS}} = \delta_{\text{TLS}} \sim (1 + P/P_c)^{-1/2} F\delta_{\text{TLS}}^0$, where P is the power and P_c is the power for TLS saturation. We need the power sufficiently low to use $1/Q_i$ as a proxy for $F\delta_{\text{TLS}}^0$.

Fig. 4 shows the resonator Q_i vs. power at the feedline for all detectors we measured. We observe an increasing trend of Q_i vs. power due to TLS saturation. At low powers below -125 dBm, the VNA loses S/N. As a result, we cannot probe into the low-power limit and can only obtain a higher limit for Q_i (without TLS saturation) and a lower limit for $F\delta_{\text{TLS}}^0$. The histogram of resonator Q_i at -124 dBm is in Fig. 5 and centers around $Q_i \sim 1000 \pm 100$, corresponding to a constraint of $\delta_{\text{TLS}}^0 \gtrsim 1.0 \times 10^{-3}$. The results are consistent with those from fitting to the TLS model (Section III) in the same order of magnitude. We tried fitting the Q_i dependence on power to the $(1 + P/P_c)^{-1/2}$ model, but there is strong degeneracy between saturation power P_c and $F\delta_{\text{TLS}}^0$. To improve this method, we can switch to a lower-noise amplifier or increase the integration time during data collection. We also note that the high-power microwave Q_i of our microstrip with SiN_x is lower than that reported by other groups [16], [17]. The Q_i for resonators with aluminum inductors used in this paper still has a positive slope vs. power at -90 dBm, indicating their Q_i is still being limited by dielectric loss there. However, these resonators started to bifurcate at ~ -90 dBm. For all-Nb

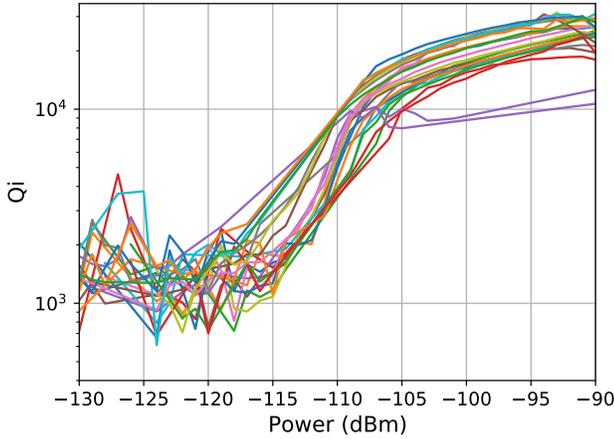


Fig. 4: Resonator quality factor vs. driving power. Each curve is a different detector. As the power increases, more TLS states become saturated, leading to a higher Q_i . At powers below -125 dBm, the VNA data becomes noisier, making it harder to constrain Q_i in the low power limit. Above -105 dBm, the Q_i is limited by other loss mechanisms, such as loss in the metal.

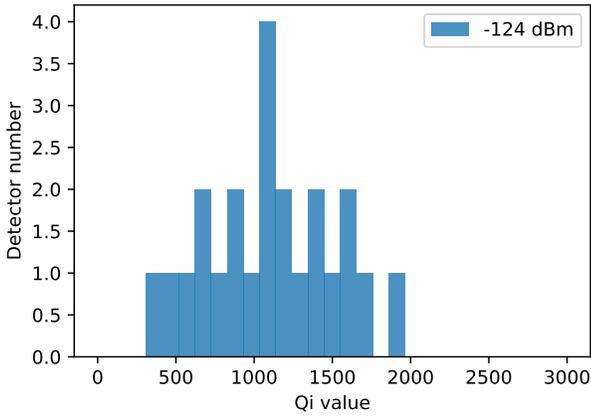


Fig. 5: Q_i distribution at -124 dBm feedline power.

resonators with higher bifurcation power, we noticed that the Q_i can reach 10^5 at -70 dBm. For this paper, we are only interested in the low-power limit of the Q_i set by the dielectric loss. The other factors that limit our high-power Q_i are orders of magnitude lower than the TLS loss.

V. DIFFERENTIAL OPTICAL EFFICIENCY AT 150 GHZ

The third method compares optical efficiency for detectors coupled through different lengths of microstrip lines, which requests an mm-wave optical coupling setup. A temperature-controlled cryogenic blackbody provides a method to sweep the incident power level, and is coupled to the detector sample box through low-pass metal-mesh filters and infrared blockers (Fig. 6).

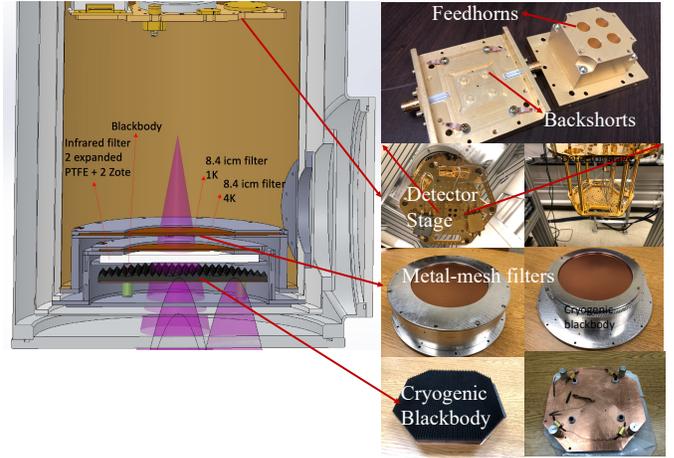


Fig. 6: Optical testing setup. The cryogenic blackbody is mounted on a copper plate with four heaters and thermometers for temperature control. An infrared filter attenuates the power above 1 THz, and two metal-mesh filters at 4 K and 1 K define a low-pass frequency window below 280 GHz. The detectors are mounted in a sample box with feedhorns and backshorts. We installed the optical components in a Bluefors dilution fridge shown on the left. The base temperature of the fridge is 8 mK. The optical window on the right of the CAD model came with the fridge and was not used for this paper.

The optical setup we built is compatible with both 150 and 220 GHz optical detector chips. Fig. 7 shows the detectors and their boxes. The 150 and 220 GHz chips were fabricated and diced from the same silicon wafer.

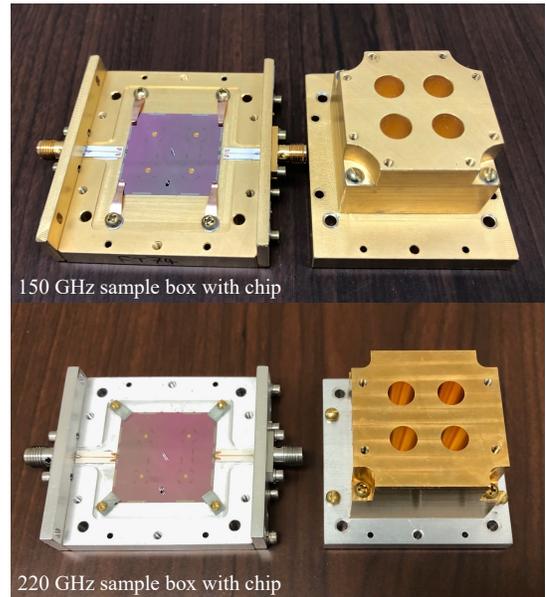


Fig. 7: 150 and 220 GHz sample boxes with detector chips installed. Left of each subfigure is the box bottom with the detector chip. Right is the box lid with feedhorns.

To measure the optical response of the detectors, we heat the blackbody to different temperatures T_{BB} and record the resonator frequency shift df/f_0 vs. T_{BB} . Here df is the frequency

shift, and f_0 is the resonant frequency. We took measurements for the 150 GHz device and show the results in Fig. 8, where we identified three optical detectors with decreasing optical response corresponding to microstrip meanders lengths of 6.4, 21.3, and 68.3 mm. The detector with the longest microstrip length of 128.3 mm has low optical response comparable to other dark detectors, suggesting its optical power is too weak from too much attenuation. The frequencies for the dark detectors also shifted slightly due to the chip heating. We keep the detector stage at a constant $T_{\text{stage}} = 247$ mK where $df/dT_{\text{stage}} \sim 0$ to minimize the detector response to base temperature variation. However, as the blackbody radiates more power to the device, the temperature gradient between the chip and the stage PID thermometer can increase as a function of blackbody temperature, causing detector frequency shifts. We successfully reduced this effect by fine-tuning the stage PID temperature such that df/dT_{stage} is closer to zero. But since the device temperature (T_{chip}) is not a constant and depends on the blackbody temperature T_{BB} , df/dT_{stage} cannot always stay zero at varying T_{BB} even if we PID control T_{stage} to a constant. We will always have a small amount of T_{chip} response, which can be monitored and corrected for using dark detectors.

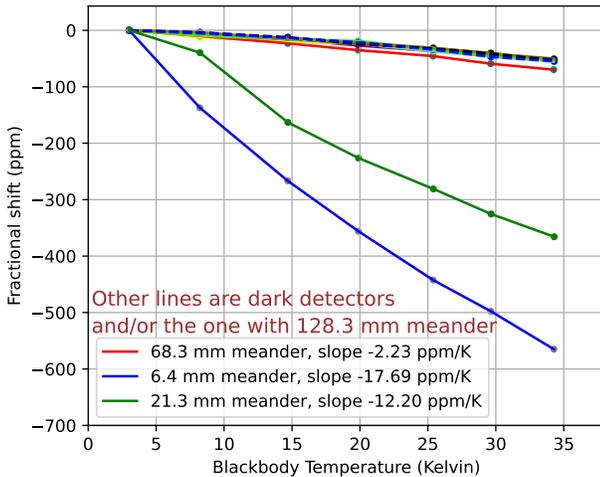


Fig. 8: Frequency shift vs. cold load temperature. We have four optical detectors coupled to 6.4, 21.3, 68.3, and 128.3mm meanders. The detector with the longest length (128.3mm) did not show a clear optical response due to too much attenuation. Dark detectors showed some response to cold load temperature due to chip heating.

The next step is to extract the dielectric loss from the measured optical responses. We use

$$\tan \delta = \frac{c}{2\pi L \sqrt{\xi_r} f_{\text{opt}}} \ln \left(\frac{\eta_{\text{long}}}{\eta_{\text{short}}} \right) \quad (3)$$

to calculate the loss through unit microstrip line length [8]. Here c is the speed of light, L is the length difference, f_{opt} is the optical band central frequency (150 or 220 GHz depending on the chip style), $\eta_{\text{short}}(\eta_{\text{long}})$ is the optical efficiency for the

short (long) meander, and permittivity $\xi_r = 7.0$. The fractional frequency shift can be decomposed into

$$df/f_0 = P_{\text{opt}} \frac{df/f_0}{dn_{\text{qp}}} \frac{dn_{\text{qp}}}{P_{\text{opt}}} = P_{\text{opt}} \frac{df/f_0}{dn_{\text{qp}}} \frac{\eta \tau_{\text{qp}}}{\Delta V_L}, \quad (4)$$

where P_{opt} is the optical power, n_{qp} is the quasiparticle density, η is the optical efficiency including line loss and pair-breaking efficiency, τ_{qp} is the quasiparticle lifetime, Δ is the bandgap, and V_L is the inductor volume [18]. We designed our inductor volume V_L to be the same for all detectors. τ_{qp} and Δ are the same for the same wafer. P_{opt} should be the same for the same feedhorn, antenna, and blackbody illumination that covers the entire beam of the detectors. We measured $df/(f_0 dn_{\text{qp}})$ from resonant frequency vs. operating temperature. n_{qp} depends on the operating temperature $n_{\text{qp}} = 2N_0 \sqrt{2\pi k_B T \Delta} \exp(-\Delta/k_B T)$ [19], where N_0 is the single spin density of states at the Fermi level. The difference for $df/(f_0 dn_{\text{qp}})$ measured within 0.7% for different detectors. Since all other terms are approximately the same, the df/f_0 vs. optical power (scaled with blackbody temperature T_{BB}) slopes can be directly divided to get the optical efficiency ratio, or $\eta_{\text{long}}/\eta_{\text{short}}$ in Eq. 3. We get $\tan \delta \sim 0.003$ from 21.3 mm/ 6.4 mm detector responses and 0.005 from 68.3 mm/ 6.4 mm. On average $\tan \delta \sim (4 \pm 1) \times 10^{-3}$. We note that our loss tangent is larger than reported by [20]. Our group is currently developing SiN_x dielectrics using several deposition methods: chemical vapor deposition (CVD), ion beam-assisted sputtering (IBAS), and PECVD. We have measured microwave loss tangents as low as 2×10^{-4} [21] in silicon-rich SiN_x films deposited by PECVD. A future paper will present additional details on the dielectric optimization for mm-wave loss tangent; the focus of this paper is on our measurement methodology.

VI. MEASUREMENT SYSTEMATICS

The systematics for our optical measurements include the chip heating effect discussed in Section IV, optical leakage, and optical coupling variation among different pixels. We can control chip heating by choosing an appropriate stage PID temperature where the df/dT_{stage} responsivity is low and using the dark detectors to estimate a correction. One concern about our optical coupling is that different antennas will have slightly different coupling efficiency if they are misaligned under the feedhorns. A series of simulations suggest that the optical efficiency variation should be within ten percent if we can align the antennas and feedhorns within 0.1 mm. Optical leakage mitigation is a more challenging task [22]. Below we discuss the leakage measurements and a method for reducing the optical leakage with absorber chips.

We observed dark detector response at 10% compared to optical detectors before any leakage mitigation. Fig. 9 shows the response level for the dark detectors. We measured the optical response for detectors on the same device with different numbers of feedhorns covered over the metallic tape. The responses of all dark detectors were ~ 1 ppm/K with all feedhorns open and uniformly reduced in half after we covered two of the four feedhorns. The amount of reduction is independent of the detector location on the device. When we

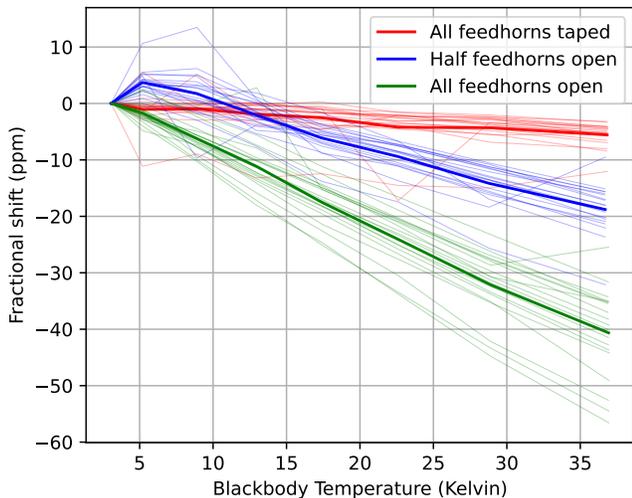


Fig. 9: Fractional frequency shift vs. blackbody temperature for dark detectors with different numbers of feedhorns taped. The response levels for most dark detectors are proportional to the number of open feedhorns. The thick line is the average over all detectors (thinner lines).

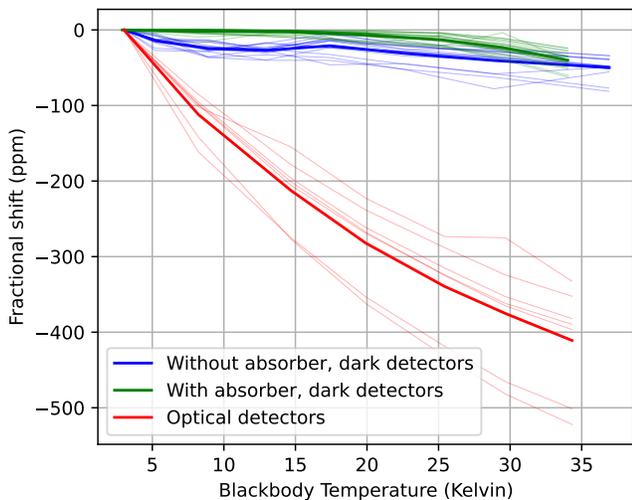


Fig. 10: Fractional frequency shift vs. blackbody temperature for the dark detectors without absorber chips, dark detectors with absorber chips, and optical detectors. The dark pickup is at $\sim 10\%$ of the response level for an optical detector and can be mitigated by installing absorber chips. The thick line is the average over all detectors (thinner lines).

covered all feedhorns, the detector responses were all close to zero, with a small slope likely caused by chip heating. The measurements are consistent with a model where optical leakage through the bottom of the feedhorn randomly scatter within the sample box and can terminate at any inductors. The leaked photons can break Cooper pairs in the inductor and shift the resonant frequency. The first data point of the blue line in Fig. 9 was an outlier caused by stage temperature variation when we started the temperature PID control.

To address this parasitic optical leakage, we designed and fabricated dedicated absorbing chips and placed them both above and below the detector chip to absorb spurious photons. Each absorber chip comprises a piece of silicon with a tuned absorbing structure that is patterned into an aluminium thin film. Fig. 10 shows the responses for optical detectors and dark detectors with and without the absorber chips. The dark response was about 1 ppm/K before adding the absorber chips and was reduced to below measurement noise level afterward. The nonlinear shift for the green curve at $T_{BB} > 25$ K is likely caused by chip heating. A picture of the top absorber device is in Fig. 11.

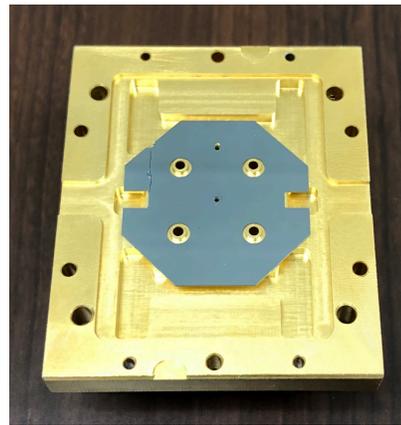


Fig. 11: An absorber chip is mounted in the box lid for terminating spurious photons. An additional absorber chip is underneath the detector chip and is not shown in this picture. The side with absorbing patterns should face the detector chip for both absorber chips. The other side of the absorber chip has a metal layer that reflects the radiation to enhance the absorption.

VII. CONCLUSIONS

This paper presents dielectric loss measurements for SiN_x using three different methods covering millimeter to centimeter wavelengths with the same test structure. At ~ 1 -2 GHz, we measured SiN_x loss to be 2×10^{-3} using the TLS fitting method and $\gtrsim 1.0 \times 10^{-3}$ by inverting Q_i at low excitation power. At 150 GHz, we measured the loss to be $\sim 4 \times 10^{-3}$ using the differential optical efficiency method. The loss at 150 GHz is a factor of ~ 2 larger compared to the loss at ~ 1 -2 GHz, which is consistent with NIST measurements [8]. Next, we will measure dielectric loss for a few more materials, such as SiO_x , silicon-rich SiN_x , and amorphous silicon. Early results showed a promising loss tangent of 2×10^{-4} at 1 GHz for silicon-rich SiN_x . We still need to understand why the loss tangent at 150 GHz is larger than at ~ 1 GHz. One possibility is other loss mechanisms beyond TLS at mm-wave, which can be tested using samples with lower loss, such as silicon-rich SiN_x . We have made 220 GHz samples and boxes and plan to extend the measurements to 220 GHz.

VIII. ACKNOWLEDGEMENTS

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