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Cumulative Fission Yields of Short-Lived Fission Products from ^{235}U and ^{239}Pu Measured by HPGe Gamma-Ray Spectroscopy

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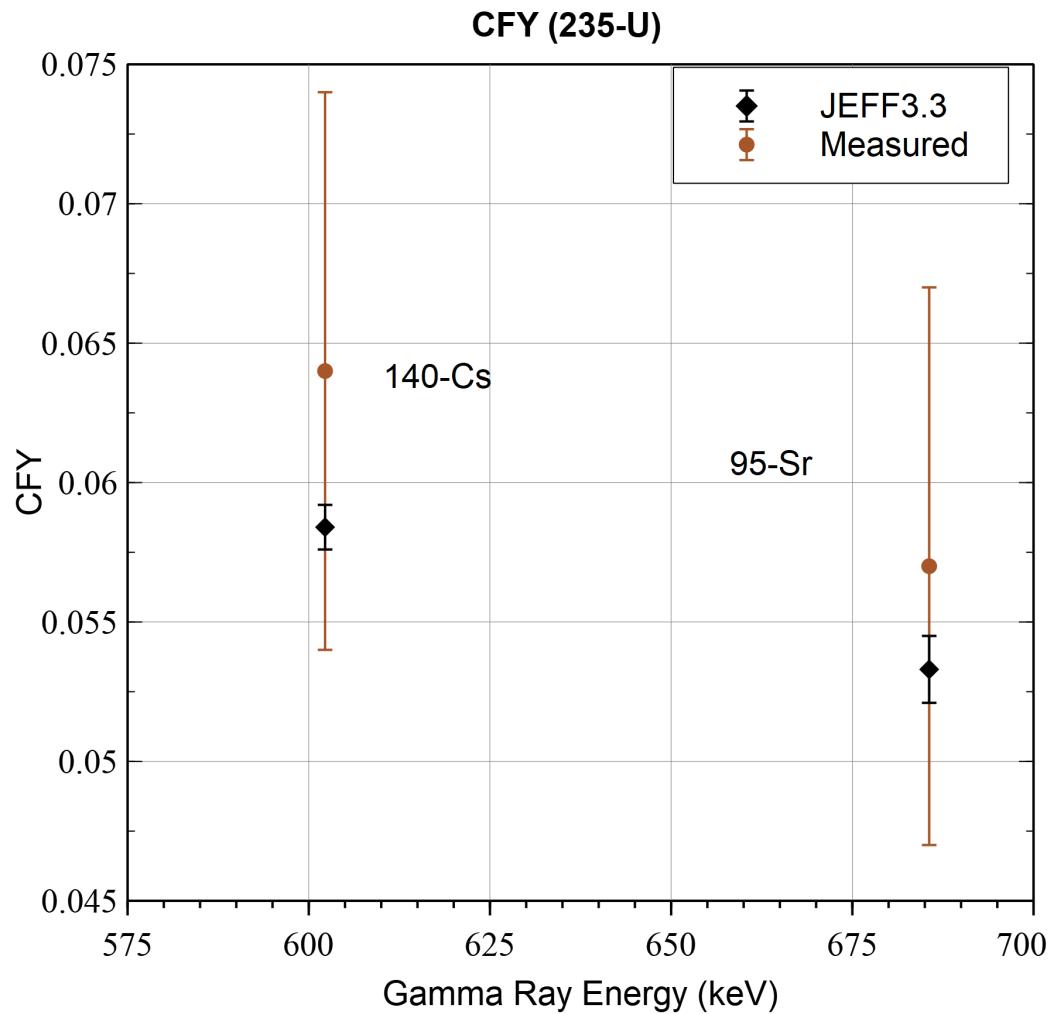
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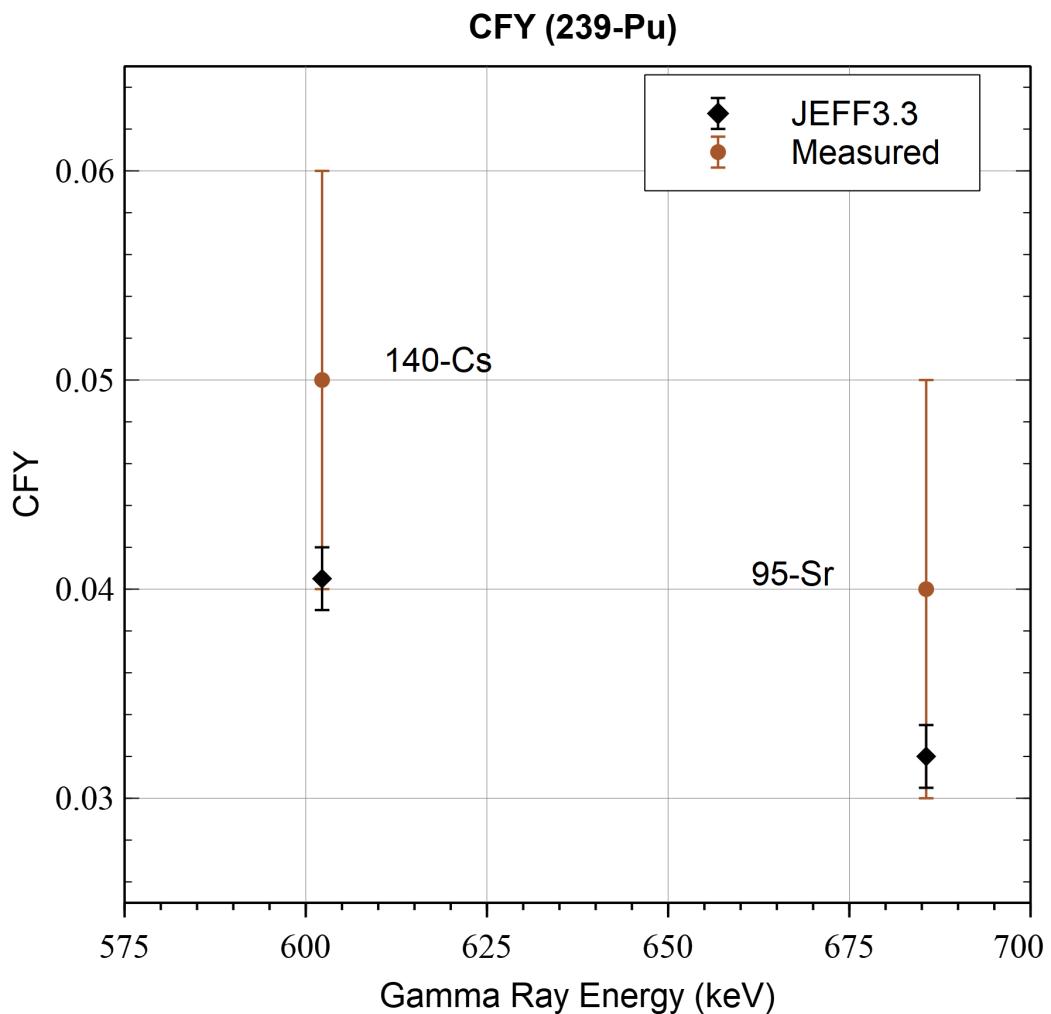
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¹ Graphical Abstract

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5 Highlights

6 **Cumulative Fission Yields of Short-Lived Fission Products from**
7 **^{235}U and ^{239}Pu Measured by HPGe Gamma-Ray Spectroscopy**

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9 • Research highlight 1 Preliminary study to show how gamma-ray spec-
10 troscopy of short-lived fission products from ^{235}U and ^{239}Pu may shed
11 light on the spectral deviation in the 5 to 7 MeV range in the reactor
12 neutrino spectra.

13 • Research highlight 2 Measured cumulative fission yields of ^{140}Cs and
14 ^{95}Sr from ^{235}U and ^{239}Pu are shown to be consistent with expectations
15 based on the JEFF3.3 fission yield library.

16 Cumulative Fission Yields of Short-Lived Fission
17 Products from ^{235}U and ^{239}Pu Measured by HPGe
18 Gamma-Ray Spectroscopy

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20 **Abstract**

21 In this study, we present a preliminary investigation focused on determining
22 cumulative fission yields for short-lived fission products. Our analysis
23 involves examining gamma spectra from the irradiated samples of ^{235}U and
24 ^{239}Pu using the High Flux Isotope Reactor. The motivation stems from the
25 observed discrepancy in the antineutrino energy spectrum within the range of
26 5 to 7 MeV. While several hypotheses have been proposed, a thorough anal-
27 ysis of fission yields provides an additional way of gaining insight into this
28 unexplained phenomenon. Our study suggests that the measured gamma
29 rays from ^{100}Nb , ^{140}Cs and ^{95}Sr are consistent with the expected values.
30 However, ^{93}Rb , ^{96}Y , ^{97}Y and ^{142}Cs cannot be quantified due to insufficient
31 statistics, interference from other gamma rays and the Compton scattering
32 background. Additionally, the calculated cumulative fission yields based on
33 the measured ^{140}Cs and ^{95}Sr are found to be consistent with the JEFF3.3
34 fission yield library. The present work shows that the potential of improv-
35 ing gamma-ray spectroscopy in the fission yields as a means to improve our
36 understanding of the neutrino spectrum.

37 *Keywords:* cumulative fission yield, gamma-ray spectroscopy, nuclear data
38 library

39 **1. Introduction**

40 Nuclear reactors are a large source of electron antineutrinos, making them
41 indispensable for investigating the properties of these particles. Approx-

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42 mately 6 antineutrinos are generated from a single fission event, and therefore,
43 a 1 GW thermal reactor emits about 10^{20} antineutrinos per second[1, 2].
44 In recent years, several reactor experiments were carried out to investigate
45 various properties of the reactor antineutrino flavor oscillations (Daya Bay[3],
46 RENO[4], STEREO[5], PROSPECT[6], NEOS[7], JUNO[8]) [2].

47 Experimental observations have revealed a spectral deviation in the 5 to
48 7 MeV range of antineutrinos when compared to the best available model.
49 Currently, this spectral feature remains unexplained [9]. While Hayes et
50 al. [10] have explored various potential sources contributing to this spectral
51 deviation, Dwyer and Langford [11] have pointed out that several obvious
52 systematic uncertainties such as absence of fission yields in the short-lived
53 isotopes have not been considered in the summation method. They suggest
54 that investigating fission yields could offer valuable insights into understanding
55 this spectral deviation. Sonzogni et al. [12] re-evaluated the thermal and
56 fast fission yields of ^{235}U in the ENDF/B database. Their analysis revealed
57 that the revision of thermal yields and decay probabilities for ^{86}Ge led to
58 about 10% variation in the calculated antineutrino spectrum in the 5 to 7
59 MeV energy range.

60 While measurement of fission yields from short-lived fission products re-
61 mains challenging [13], this study represents a feasibility study for measure-
62 ment of cumulative fission yields from short-lived fission products through
63 gamma-ray measurements. In our analysis, measured cumulative fission yield
64 is compared with the JEFF3.3 fission library to identify any disparities. The
65 JEFF3.3 database is selected for the comparison because it is the preferred
66 source of yields for antineutrino applications [9, 14]. Selection of the short-
67 lived fission products (Table 1) for our investigation is based on the list of
68 significant contributors at 5.5 MeV in the antineutrino spectrum [11, 12].

69 Unstable nuclides directly produced in fission undergo beta decay along
70 the isobar chain. Although these beta decays have very short half-lives,
71 they eventually yield nuclides with sufficiently long half-lives to allow mea-
72 surement of emitted gamma rays [15]. The direct fission yield of individual
73 nuclides in the primary fission event is referred to as Independent Fission
74 Yield (IFY), while the total yield of a nuclide including beta decay feeding is
75 termed Cumulative Fission Yield (CFY). These yields are expressed as per
76 fission event [12, 15].

Table 1

Decay data for 8 nuclides investigated in this study from the ENDF/B-VIII decay data sublibrary, including the decay chain gamma-ray with the strongest intensity selected for the present analysis. Uncertainty is given in the parenthesis.

Isotope	Half life (s)	Gamma Energy (keV)	Intensity
⁹³ Rb	5.84(2)	432.61(3)	0.202(14)
¹⁰⁰ Nb	1.4(2)	535.666(14)	0.46(6)
¹⁴⁰ Cs	63.7(3)	602.25(5)	0.53(3)
⁹⁵ Sr	23.90(14)	685.6	0.226
⁹² Rb	4.49(3)	814.98(3)	0.032(4)
⁹⁶ Y	5.34(5)	1750.4(2)	0.0235(24)
⁹⁷ Y	3.75(3)	3287.6(4)	0.181(19)
¹⁴² Cs	1.68(14)	359.598(14)	0.27(3)

77 2. Experiment

78 The ²³⁵U sample consists of 252.72 nanograms of natural uranium nitrate
 79 in an Inductively Coupled Plasma calibration solution. The ²³⁹Pu sample
 80 consists of 301.3 nanogram of National Institute of Standards and Technology
 81 (NIST) Certified Reference Material (CRM-137). The samples are irradiated
 82 using the PT-2 pneumatic tube of the High Flux Isotope Reactor (HFIR) at
 83 the Neutron Activation Analysis laboratory (NAA) of Oak Ridge national
 84 Laboratory. The measured thermal and epithermal neutron fluxes at the
 85 irradiation location [16] are 4.59×10^{13} n/cm²/sec and $1.96E \times 10^{11}$ n/cm²/sec
 86 respectively for ²³⁵U, and 4.43×10^{13} n/cm²/sec and 3.24×10^{11} n/cm²/sec
 87 respectively for ²³⁹Pu. The energy ranges for epithermal neutrons are 0.1 eV
 88 to 10 keV for ²³⁹Pu and 1 eV to 10 keV for ²³⁵U [17]. The neutron fluxes are
 89 measured using manganese and gold activation foils.

90 Each sample is irradiated for 30 seconds, and then transported to the
 91 detector chamber using the pneumatic tube transfer system [16] which in-
 92 troduces a 20-second delay prior to the gamma-ray measurement. Fig. 1
 93 shows the measured gamma-ray spectra of the irradiated ²³⁵U and ²³⁹Pu.

94 The gamma rays are measured with a 44% relative efficiency, ORTEC p-
 95 type coaxial HPGe detector with an aluminum end cap. Each sample is
 placed at 33 cm above the detector and measured for 30 seconds.

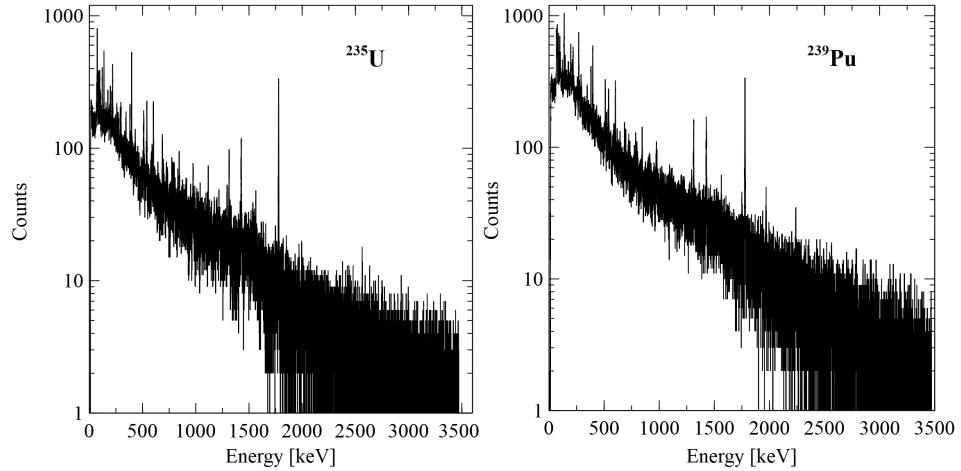


Fig. 1. Measured gamma-ray spectra from freshly irradiated ^{235}U and ^{239}Pu are plotted. See text for details.

96

97 3. Fission yields and expected gamma rays

98 The expected gamma-ray yield calculation starts by determining the num-
 99 ber of ^{235}U and ^{239}Pu nuclides initially present in the sample from the sample
 100 mass (m), Avogadro's number (N_A) and the molar weight (M). The number
 101 of nuclides (N_{fd}) directly produced from fission is given by Eq. (1).

$$N_{fd} = \text{IFY} \sigma_f \phi \frac{m N_A}{M} \quad (1)$$

102 The equation includes the IFY of a specific nuclide, the thermal neutron
 103 cross section (σ_f) and the thermal neutron flux (ϕ) [18]. The IFY are tabu-
 104 lated in the JEFF3.3 library, and the neutron cross section is based on the
 105 ENDF/B-VIII.0 neutron cross section standard sublibrary. The JEFF3.3
 106 and ENDF/B-VIII.0 fission yield libraries contain different IFY values for
 107 certain nuclides. This is demonstrated using the ^{140}Cs decay chain in Table
 108 2. In this example, JEFF3.3 does not have IFY for ^{140}Sb , so the IFY value

Table 2

Examples from the ^{140}Cs decay chain, showing the differing IFY of ^{235}U (top) and ^{239}Pu (bottom) from JEFF3.3 and ENDF/B-VIII.0 fission yield libraries. Uncertainty of each IFY is indicated in the parenthesis.

IFY	^{140}Sb	^{140}Te	^{140}I	^{140}Xe	^{140}Cs
JEFF3.3	No data	6.57E-08 (2.26E-08)	3.03E-04 (1.03E-04)	1.25E-02 (3.10E-03)	1.84E-02 (3.85E-03)
ENDF/B-VIII.0	2.82E-09 (1.81E-09)	9.04E-06 (5.78E-06)	1.11E-03 (7.13E-04)	2.59E-02 (1.04E-03)	3.05E-02 (1.83E-03)
IFY	^{140}Sb	^{140}Te	^{140}I	^{140}Xe	^{140}Cs
JEFF3.3	No data	2.33E-07 (8.06E-08)	4.77E-04 (1.63E-04)	1.83E-02 (4.06E-03)	2.18E-02 (4.52E-03)
ENDF/B-VIII.0	5.61E-11 (3.59E-11)	1.41E-06 (9.02E-07)	5.94E-04 (3.80E-04)	1.54E-02 (4.31E-04)	2.28E-02 (3.64E-03)

from ENDF/B-VIII.0 is used instead in our analysis. This suggests that N_{fd} will be different depending on the fission library used.

To determine the total number of nuclides (N_T) present at the end of irradiation, N_{fd} as well as each of its successors in the decay chain need to be determined from its own IFY and β -decayed. For the analysis presented in this work, only the β -decay path for the parent-daughter chains is used. The expected N_T is given in Eq. (2).

$$N_T = N_{fd} \frac{(1 - e^{-\lambda t})}{\lambda} + \sum_j Decay([N_{fd}]_j) \quad (2)$$

The first term gives the total number of a gamma-ray emitting nuclide produced directly from IFY during the irradiation. The second term describes the total number of a gamma-ray emitting nuclide resulting from the β -decay of j^{th} predecessor based on its own IFY.

In our study, the decay chains consist of 4 nuclides (^{93}Rb , ^{142}Cs) and 6 nuclides (^{140}Cs , ^{95}Sr , ^{92}Rb , ^{96}Y , ^{100}Nb , ^{97}Y). Each decay chain leading from IFY to the gamma-ray emitting nuclides measured in this study is described by a set of coupled linear differential equations describing the radioactive decays. These equations are reformulated as a set of matrices and solved using MATLAB [19]. The solution to each decay chain gives the number of a gamma-ray emitting nuclides produced from IFY and the decay of its predecessors during the 30-second irradiation. Fig. 2 summarizes the percent composition of the N_T based on the contribution from IFY and its successors. In general, only the parent or parent / grandparent of a given gamma-ray emitting nuclide remains at the end of irradiation. The nuclides further up in the decay chain have all decayed away due to their very short half-lives.

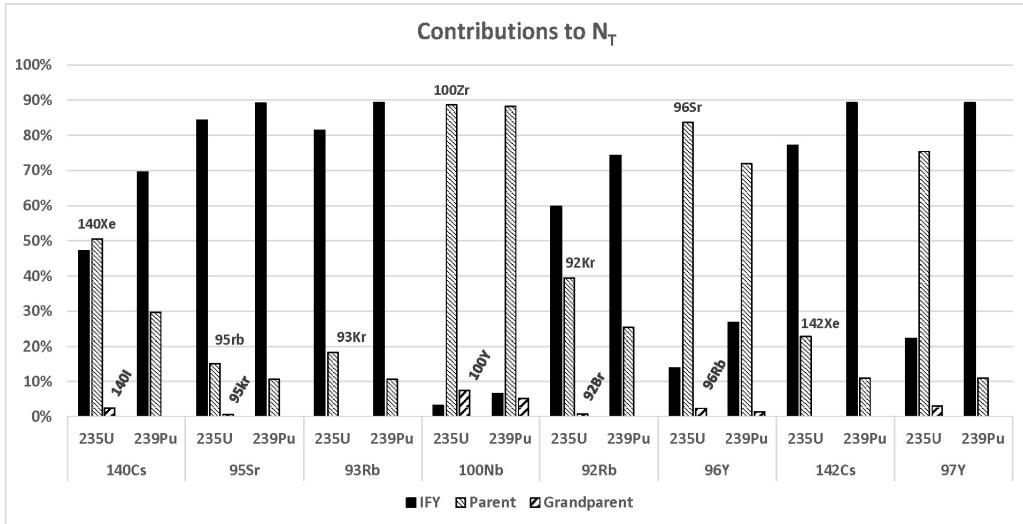


Fig. 2. The plot summarizes the percent contribution from the IFY and its successors to the composition of N_T in our study. For example in ^{235}U , 47% of N_T for ^{140}Cs is due to IFY of ^{140}Cs . The remaining contribution comes from ^{140}Xe (51%, parent) and ^{140}I (2%, grandparent). The predecessors further up the decay chain such as ^{140}Sb and ^{140}Te have all decayed away, and their contribution to ^{140}Cs is negligible. Nuclide labels for ^{235}U and ^{239}Pu are the same, and omitted in the ^{239}Pu portion of the chart for clarity.

132 The uncertainty in the N_T is determined from the IFY of a gamma-ray
 133 emitting nuclide and its predecessors, their decay constants, and thermal
 134 neutron cross sections as shown in Eq. (3) [20]. The IFYs make a greater
 135 contribution to the overall uncertainty than λ and σ . For example, the IFYs
 136 accounts for 99% of the uncertainty in ^{140}Cs and ^{95}Sr for both ^{235}U and ^{239}Pu
 137 while the contributions from λ and σ are 1%.

$$\delta(N_T) = N_T \sqrt{\sum \left(\left(\frac{\delta(\text{IFY})}{\text{IFY}} \right)^2 + \left(\frac{\delta(\lambda)}{\lambda} \right)^2 \right)_i + \left(\frac{\delta(\sigma_f)}{\sigma_f} \right)^2} \quad (3)$$

138 Finally, the expected gamma-ray yields (N_γ) during the subsequent 30-
 139 second measurement time are calculated from the N_T , the decay constant (λ),
 140 the absolute efficiency (ϵ) of the HPGe, and the gamma emission intensity
 141 (I_γ). The calculation includes the contributions from the predecessors con-
 142 tinuously decaying during the 20-second transportation delay. Except ^{140}Cs ,
 143 the half-lives of the measured nuclides are much shorter than the detector
 144 measurement time, therefore, it is necessary to decay-correct the peak counts.
 145 The ANSI standard for the correction factor is described in Ref.[21]. The
 146 uncertainty in the expected gamma-ray yield, $\delta(N_\gamma)$, is evaluated through a
 147 quadrature sum of uncertainties in N_T , ϵ , and I_γ . The dominant contributor
 148 to $\delta(N_\gamma)$ comes from the $\delta(N_T)$. For example, 99% of the overall $\delta(N_\gamma)$ in
 149 ^{140}Cs and ^{95}Sr for both ^{235}U and ^{239}Pu can be attributed to $\delta(N_T)$.

150 4. Measured gamma rays

151 The energy and full width at half maximum (FWHM) calibrations of the
 152 HPGe detector have been determined by analyzing known gamma-ray peaks.
 153 The absolute efficiency of the detector is estimated using the Geometry and
 154 Tracking (GEANT4)[22] simulation package. In the simulation, 17 gamma-
 155 ray energies are selected to cover the energy range from 50 keV to 3.5 MeV.
 156 Each gamma-ray simulation was performed using 1E+6 photons to determine
 157 the efficiency of the detector at each photon energy. The detector model in
 158 GEANT4 includes all the details of detector construction, including a 0.1
 159 cm thick aluminum window on the endcap of the detector and a 0.07 cm
 160 thick dead layer on the surface of the HPGe crystal. Dimensions of the
 161 HPGe were taken to be 6.5 cm in the diameter and 6.45 cm in the length
 162 based on published ORTEC documents[23]. According to the ANSI/IEEE

163 standard 325 [24], the relative efficiency of an HPGe is defined by Eq. (4).
 164 The absolute efficiency of HPGe at 1.33 MeV is measured with a source to
 165 detector distance of 25 cm. The relative efficiency is the ratio of this HPGe
 166 absolute efficiency to the absolute efficiency of a 3-inch by 3-inch Na(Tl)
 167 scintillator at 1.33 MeV measured at 25 cm (1.2E-3).

$$\text{Relative efficiency} = \frac{\text{Absolute efficiency}}{1.2 \times 10^{-3}} \quad (4)$$

168 To establish a benchmark, a GEANT4 simulation was performed for a
 169 point source placed at the standard distance of 25 cm from the detector.
 170 The absolute efficiency at 1333 keV was expected to be 5.3E-4 for a 44%
 171 relative efficient HPGe [25]. The simulated absolute efficiency was 5.9E-
 172 4(7.7E-5). Fig. 3 shows the simulated detector efficiency. The efficiency
 173 is fitted using the parametric equation given in the RADWARE software
 174 package [26]. Above 150 keV, efficiency is fitted with the parameters (D, E
 175 and F) in the form of:

$$\text{Efficiency} = e^{D+Ey+Fy^2} \quad (5)$$

176 where $y = \ln(E_\gamma/1000)$ and E_γ is a gamma-ray energy in keV.

177 Measured gamma-ray peaks are analyzed using two methods: non-linear
 178 fitting and a simple summation. The ANSI standard for the summation
 179 method is given in [21, 27, 28], and the detailed explanation is given in
 180 [29, 30]. Non-linear fit analysis is performed using a GF3M program from
 181 the RADWARE package [26] and an open-source software, GNUPLOT [31].
 182 (see Figs. 4 and 5)

183 The fitting algorithm offers a linear and a quadratic baseline fit. The
 184 linear baseline fit is chosen because it produces a less fitting error in the
 185 peak-count estimate than the quadratic fit. The linear baseline is given by
 186 $A + Bx$ where A and B are fitted parameters, and x is a channel number.
 187 Initial estimate of A and B are given by a straight line between the limits of
 188 the fit. In the Gaussian profile, the fitted parameters are the position, height
 189 of the peak and the peak width (FWHM). The initial values for position and
 190 height are determined based on a given peak. To sufficiently account for the
 191 linear baseline as an approximation to the Compton continuum, $3 \times \text{FWHM}$
 192 from the centroid is chosen as a fitting range for a well isolated Gaussian
 193 peak [24, 32].

194 Table 3 and Table 4 summarize the fit results for ^{235}U and ^{239}Pu . In
 195 general, the fitted peak energies are consistent with tabulated values for

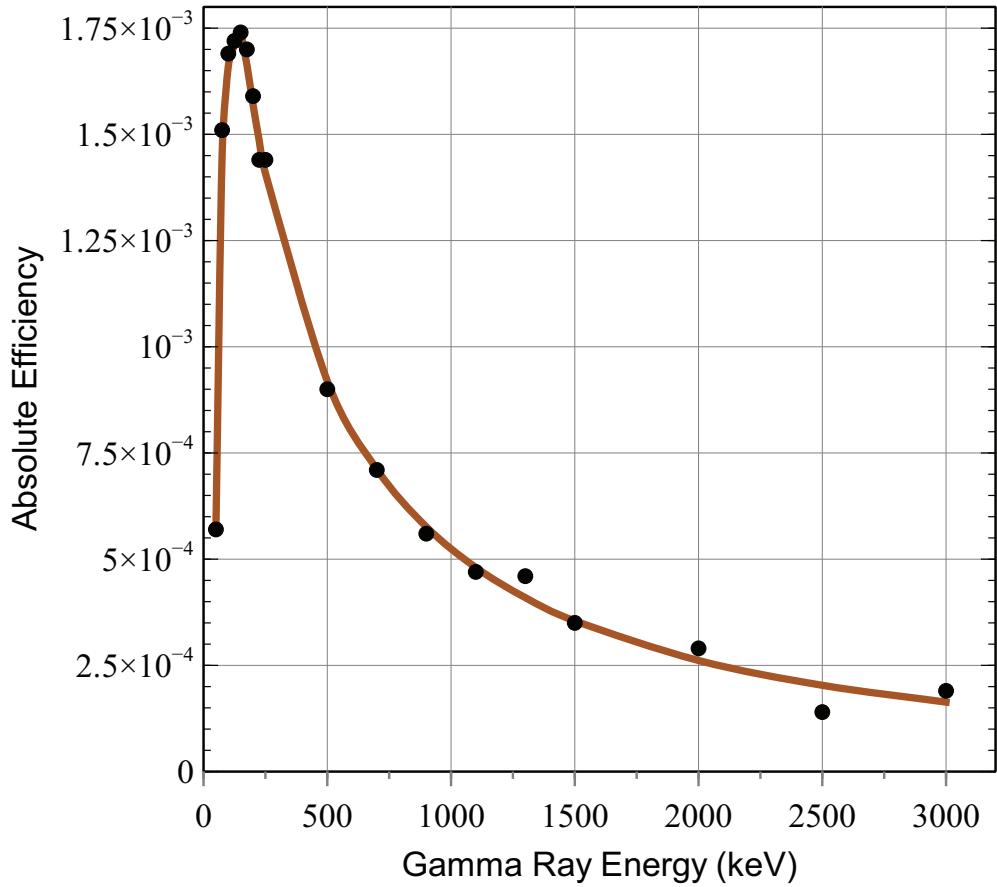


Fig. 3. GEANT4 efficiency for the ORTEC P-type 44% relative efficiency HPGe detector used here. The efficiency is fit using the parametric equation given in the RADWARE program.

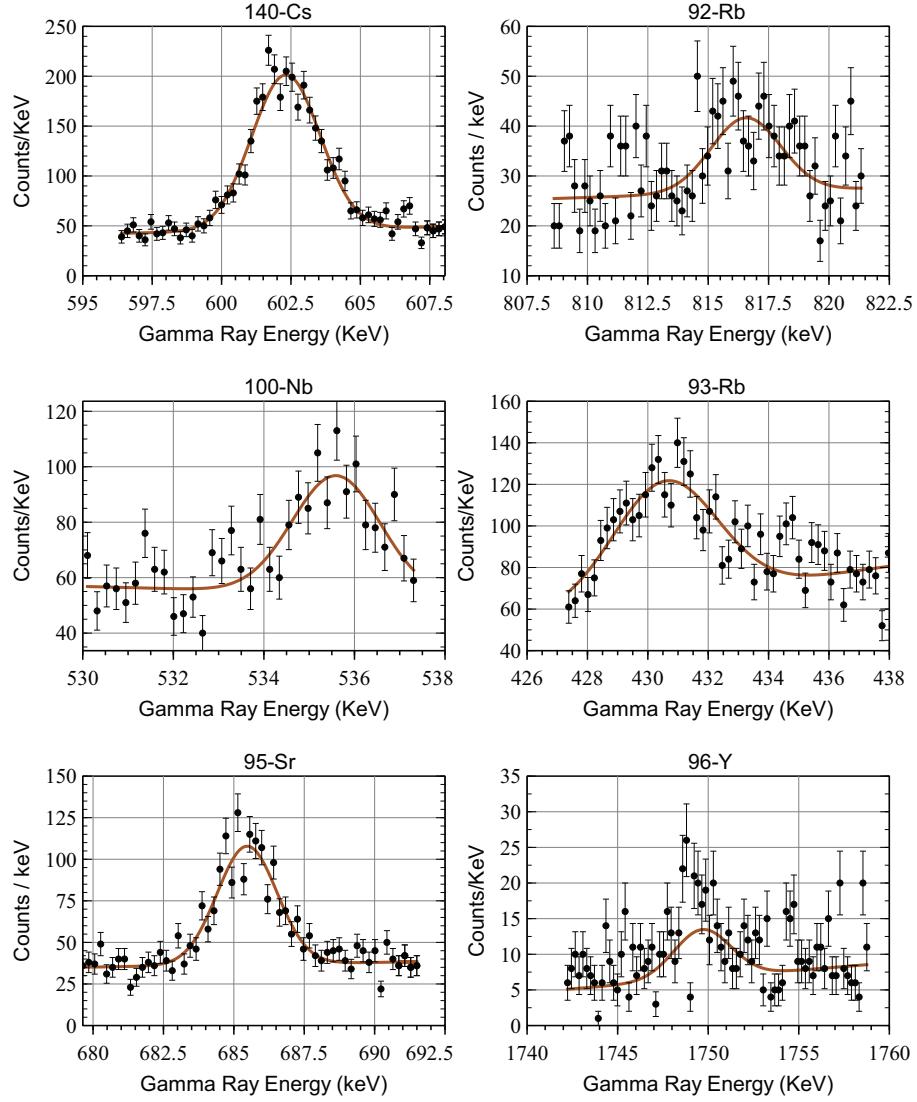


Fig. 4. Data and fits for the six gamma-ray peaks of interest from ^{235}U fission are shown. The measured spectra are shown by the dots with $1-\sigma$ uncertainties, and the fits by solid lines. Gamma ray peaks from ^{97}Y and ^{142}Cs are not detectable with the present statistics, and are omitted in this figure.

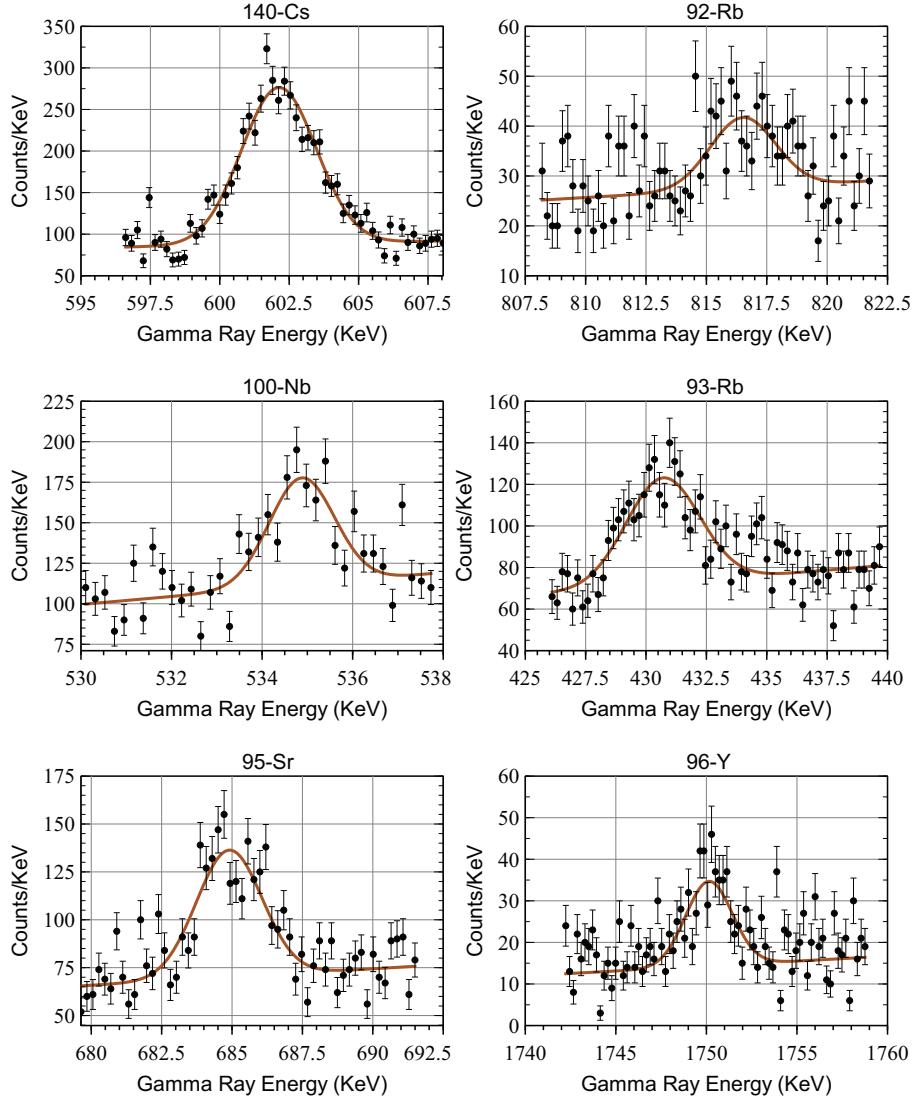


Fig. 5. Data and fits for the six gamma-ray peaks of interest from ^{239}Pu fission are shown. The measured spectra are shown by the dots with $1-\sigma$ uncertainties, and the fits by solid lines. Gamma ray peaks from ^{97}Y and ^{142}Cs are not detectable with the present statistics, and are omitted in this figure.

Table 3

Best-fit values and fitting statistics for the fission products from ^{235}U are summarized. Fitted energies are consistent with the calibrated energies. But fitted FWHMs show some divergence. The fit statistics (p-values) indicate acceptable fits for ^{100}Nb , ^{140}Cs and ^{95}Sr . (see Results section for detail).

^{235}U	χ^2/DOF	Centroid		FWHM			p-value
		Calibrated	Fitted	Calibrated	Fitted		
^{140}Cs	1.63	602.33	602.28(4)	1.93	2.87(8)	0.05	
^{95}Sr	1.43	685.57	685.43(6)	2.01	2.50(11)	0.03	
^{100}Nb	1.26	535.61	535.4(3)	1.87	2.4(10)	0.02	
^{93}Rb	1.21	432.67	430.38(15)	1.76	3.4(3)	< 0.01	
^{92}Rb	2.26	814.98	814.6(5)	2.12	1.1(12)	< 0.01	
^{96}Y	1.62	1750.50	1749.4(3)	2.73	1.9(6)	< 0.01	

196 both ^{235}U and ^{239}Pu . However, the p-values and FWHM fits suggest that
 197 a single Gaussian may be a poor model for some peaks, likely indicating
 198 interference from additional unidentified gamma rays. This could be clarified
 199 with improved counting statistics. As shown in Fig. (1), ^{239}Pu generates
 200 generally more gamma-ray activities than ^{235}U , suggesting more interference.
 201 This fact appears to be consistent with all p-values being lower for ^{239}Pu
 202 compared to ^{235}U .

203 5. Cumulative fission yields from measured gamma rays

204 The cumulative fission yield is calculated by evaluating the disintegration
 205 rate of nuclides relative to the total fission rate, and is given in Eq. (6)
 206 [33]. The disintegration rate of nuclides at the end of irradiation is derived
 207 from the measured gamma-ray counts, including appropriate adjustments
 208 for delay and count corrections. The overall fission rate is determined by the
 209 number of ^{235}U (^{239}Pu) in the sample, the thermal fission cross-section, and
 210 the thermal neutron flux.

$$CFY = \frac{\lambda N_p}{N_{fd} I_\gamma \epsilon (1 - e^{-\lambda t_{ir}})(e^{-\lambda t_d})(1 - e^{-\lambda t_m})} \quad (6)$$

Table 4

Best-fit values and fitting statistics for the fission products from ^{239}Pu are summarized. Similar to the case of ^{235}U , fitted energies are consistent with the calibrated energies while fitted FWHMs show some divergence. Unlike ^{235}U , the fit statistics (p-values) indicate poor fit quality for ^{100}Nb , ^{140}Cs and ^{95}Sr due to interference from other gamma rays. (see Results section for detail).

^{235}U	χ^2/DOF	Centroid		FWHM			p-value
		Calibrated	Fitted	Calibrated	Fitted		
^{140}Cs	1.90	602.33	602.03(5)	1.93	3.41(11)	< 0.01	
^{95}Sr	1.37	685.57	685.45(6)	2.01	2.41(11)	< 0.01	
^{100}Nb	2.74	535.61	534.75(10)	1.87	2.09(23)	< 0.01	
^{93}Rb	2.92	432.68	434(10)	1.76	5(18)	< 0.01	
^{92}Rb	1.89	814.96	815.4(3)	2.12	1.9(10)	< 0.01	
^{96}Y	1.9	1750.50	1750.19(13)	2.73	2.23(23)	< 0.01	

211 In Eq. (6), N_p is the measured gamma-ray peak count based on N_T , t_{ir} is the
 212 irradiation time (30 seconds), t_d is the transport delay time (20 seconds) and
 213 t_m is the detector count time (30 seconds). Because the gamma-ray peak
 214 counts for ^{93}Rb , ^{92}Rb and ^{96}Y are below the statistical limit of detection,
 215 proper N_p cannot be determined. The fission-produced ^{100}Nb at the end of
 216 irradiation has almost all decayed away during the 20-second transport delay
 217 due to its short half life (1.5 seconds), and is not detectable by the HPGe
 218 during the measurement time. The measured ^{100}Nb gamma rays are pri-
 219 marily due to the decay of ^{100}Zr (7.1 seconds). N_p from ^{140}Cs and ^{95}Sr are
 220 appropriate to use for calculating CFY. The uncertainty in the calculated
 221 CFY, $\delta(\text{CFY})$, is determined through a quadrature using N_p , λ , ϵ and I_γ .

222 6. Results

223 For each gamma-ray, the statistical significance is determined using the
 224 method of Refs. [29, 30, 27]. This method involves two statistical limits:
 225 Lc and Ld. The critical limit (Lc) is defined as the net count of a gamma
 226 ray peak above which a sample net count is statistically significant with the
 227 probability of false positive given by α . The detection limit (Ld) is defined

228 as the net count of a gamma ray peak above L_c that has a probability of
 229 false negative given by β . We adopt the usual convention of using $\alpha = \beta =$
 230 0.05 as the desired level of statistical significance. We note that the statistics
 231 in this method are based on a one-sided 95% confidence level so that the z
 232 statistic cutoff is 1.65, not 1.96. For the non-linear fitting method, L_c and
 233 L_d are given by Eq. (7) and (8) where $\sigma = \sqrt{B}$ and B = background count
 234 (no sample is present) respectively [29].

$$L_c = 2.33 \sigma \quad (7)$$

$$L_d = 2.71 + 4.65 \sigma \quad (8)$$

236 In Fig. (6) and Fig.(7), the measured gamma rays for ^{100}Nb , ^{140}Cs and
 237 ^{95}Sr are statistically significant, are above the minimum detection limit, and
 238 are fully consistent with the expected counts.

239 For ^{93}Rb , ^{92}Rb , and ^{96}Y , the obtained net counts from fitting significantly
 240 exceed the expected values. Using multiple Gaussian peaks does not improve
 241 the fitting results for ^{93}Rb and ^{92}Rb . In the case of ^{96}Y (1750.4 KeV),
 242 separation of ^{96m}Y (1750.06 KeV) from ^{96}Y cannot be achieved in the current
 243 data. Therefore, we estimate the gamma ray interference as follows.

244 For ^{93}Rb , 6 nuclides ($A = 90, 134, 138, 143, 144, 145$) produce a similar
 245 or larger order of magnitude of gamma-ray yield ($I_\gamma \times$ total fission yield) in
 246 the 432 keV region in our data [34, 35]. Based on the estimate of the gamma
 247 rays having a measurable effect, the proportion of gamma-ray yield of ^{93}Rb
 248 with respect to the 6 nuclides gives about 6% which is consistent with the
 249 expected net count of ^{93}Rb . In addition, ^{143}Ba (431.2 KeV, $I_\gamma = 0.0276$) is
 250 expected to produce approximately the same number of gamma ray counts
 251 in our data as ^{93}Rb (432.61 KeV, $I_\gamma = 0.202$), and is shown to be consistent
 252 with ^{93}Rb .

253 As for ^{92}Rb , 13 nuclides ($A = 82, 91, 92, 101, 132, 133, 132, 136, 137,$
 254 $139, 140, 144, 147$) produce a similar or larger order of magnitude of gamma-
 255 ray yields than ^{92}Rb in the 815 keV energy region [34, 35]. The proportion
 256 of measured peak counts from ^{92}Rb with respect to the thirteen nuclides is
 257 about 1.4% which is consistent with the expected net count of ^{92}Rb . For
 258 ^{239}Pu , 1% of the contribution in the 815 keV region is due to ^{92}Rb alone.

259 ^{96m}Y has a larger gamma-ray intensity (0.88) than $^{96}\text{Y}(0.024)$, and larger
 260 IFY from $^{235}\text{U}(0.011)$ and $^{239}\text{Pu}(0.014)$ compared to ^{96}Y : $^{235}\text{U}(0.006)$ and
 261 $^{239}\text{Pu}(0.008)$. The fission analysis shows that 3% (1%) of the gamma ray
 262 yield is due to ^{96}Y alone for $^{235}\text{U}(^{239}\text{Pu})$. When fitted net counts are adjusted

263 appropriately to account for interference, they are shown to be consistent
264 with the expected net count. However, the results are below the statistical
265 significance and are inconclusive.

266 Due to their statistical insignificance of gamma-ray peak counts, the com-
267 putations of CFY for ^{93}Rb , ^{92}Rb , and ^{96}Y have been excluded. Similarly, the
268 calculation of CFY for ^{100}Nb is excluded because it is undetectable by the
269 HPGe detector during the measurement. CFY values for ^{140}Cs and ^{95}Sr have
270 been calculated and are shown in Fig. 8 (^{235}U) and Fig. 9 (^{239}Pu) alongside
271 the JEFF3.3 fission yield data for comparison. The primary sources of un-
272 certainty in $\delta(\text{CFY})$ come from the uncertainties associated with N_p and ϵ .
273 This indicates that enhancing the statistics for gamma-ray peak counts and
274 refining efficiency calibration can result in a better estimation of CFY.

275 Hayes et al. [1] draw attention to the possibility of a contribution from
276 epithermal neutron-induced fission. In this study, the highest neutron flux
277 is observed at 0.02 eV within the thermal neutron range. Consequently,
278 the recorded epithermal neutron flux accounts for just 0.4% (0.7%) of the
279 thermal neutron flux for ^{235}U (^{239}Pu). This implies that the number of
280 neutrons interacting with ^{235}U (^{239}Pu) in the thermal region is 234(137)
281 times greater than in the epithermal region. In our analysis, the contribution
282 from epithermal neutrons is deemed too minimal to significantly impact the
283 fission yield data.

284 7. Conclusion

285 Certain nuclides make substantial contributions to the antineutrino spec-
286 trum in the 5 to 7 MeV range. An examination of their fission yields provides
287 valuable insights for comprehending the unexplained spectral deviation in
288 this energy range. In this study, we determine the cumulative fission yields
289 of specific short-lived fission products from the irradiated ^{235}U and ^{239}Pu us-
290 ing the gamma-ray spectroscopy. The measured gamma rays for ^{100}Nb , ^{140}Cs
291 and ^{95}Sr are consistent with the expected. Statistics for the measured ^{93}Rb ,
292 ^{92}Rb and ^{96}Y are sub-optimal due to interference from other gamma rays
293 and Compton background. The gamma ray peaks from ^{97}Y and ^{142}Cs are
294 undetectable due to low fission yield, limited detector efficiency, and environ-
295 mental background. Because of its short half-life, we are unable to calculate
296 the CFY for ^{100}Nb . The calculated CFY for ^{140}Cs and ^{95}Sr are in agreement
297 in comparison with the JEFF3.3 database.

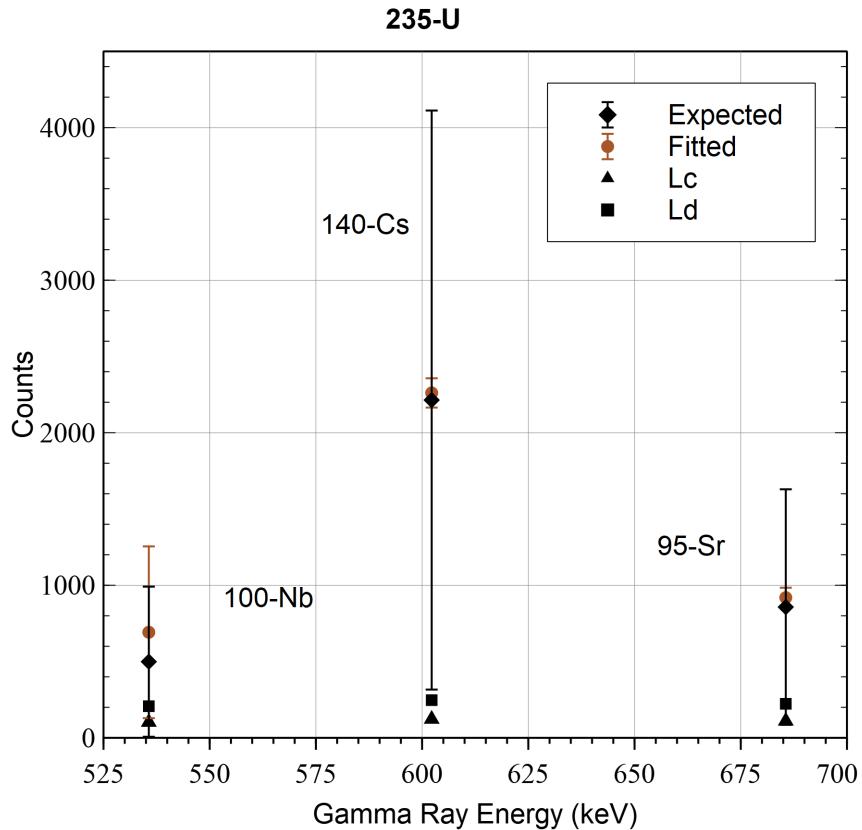


Fig. 6. Fitted and expected net counts and statistical limits and uncertainties for ^{235}U . The yields of ^{100}Nb , ^{140}Cs and ^{95}Sr are shown to be consistent with the expected values. ^{93}Rb , ^{92}Rb and ^{96}Y are below the statistical limit of detection, and are excluded from the plot for clarity.

298 A follow-up experiment offers the opportunity to improve various aspects
 299 of the current preliminary study. The primary source of uncertainty comes
 300 from IFY measurements, which need separate and dedicated experiments
 301 to refine their accuracy. To prevent the loss of nuclides, it is essential to
 302 reduce the 20-second transport delay. For instance, all fissioned ^{100}Nb nu-
 303 clides decay during the transport delay due to its short 1.5-second half-life.
 304 Addressing this challenge may require a substantial reconstruction of the
 305 sample transportation apparatus, incurring significant expenses. Additional
 306 improvements include the integration of multiple HPGe detectors to enhance
 307 gamma-ray detection through coincidence measurements, as well as installing

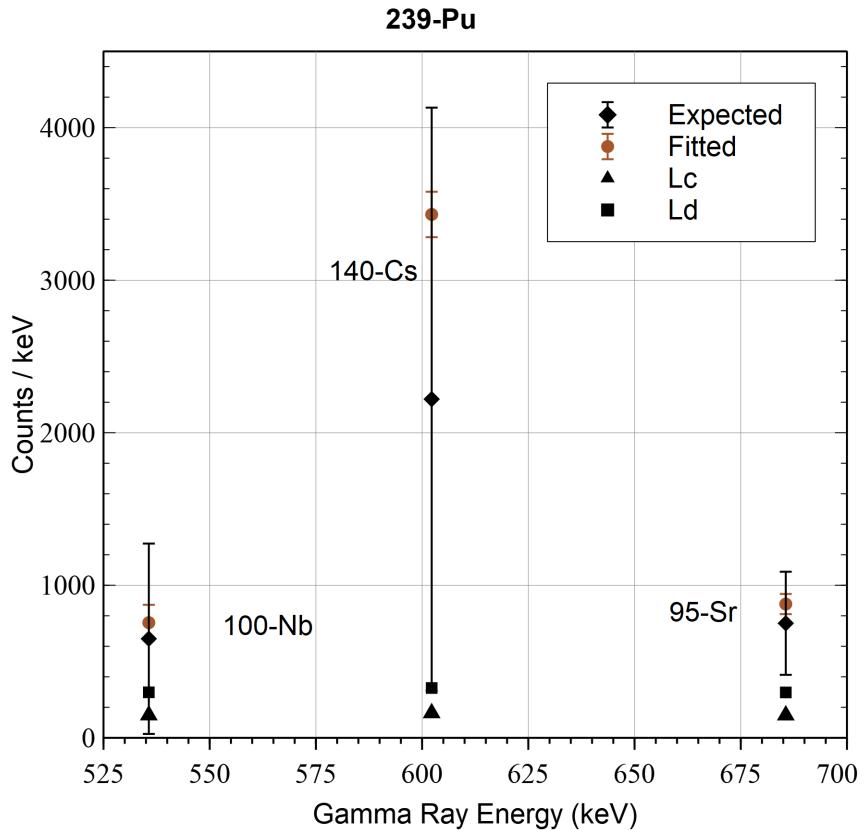


Fig. 7. Fitted and expected net counts and statistical limits and uncertainties for ^{239}Pu . The yields of ^{100}Nb and ^{95}Sr are plotted and shown to be consistent with the expected values. The fitted ^{140}Cs is about 35% larger than the expected value, suggesting a possible problem with the JEFF3.3 fission yield library. ^{93}Rb , ^{92}Rb and ^{96}Y are below the statistical limit of detection, and are excluded from the plot for clarity.

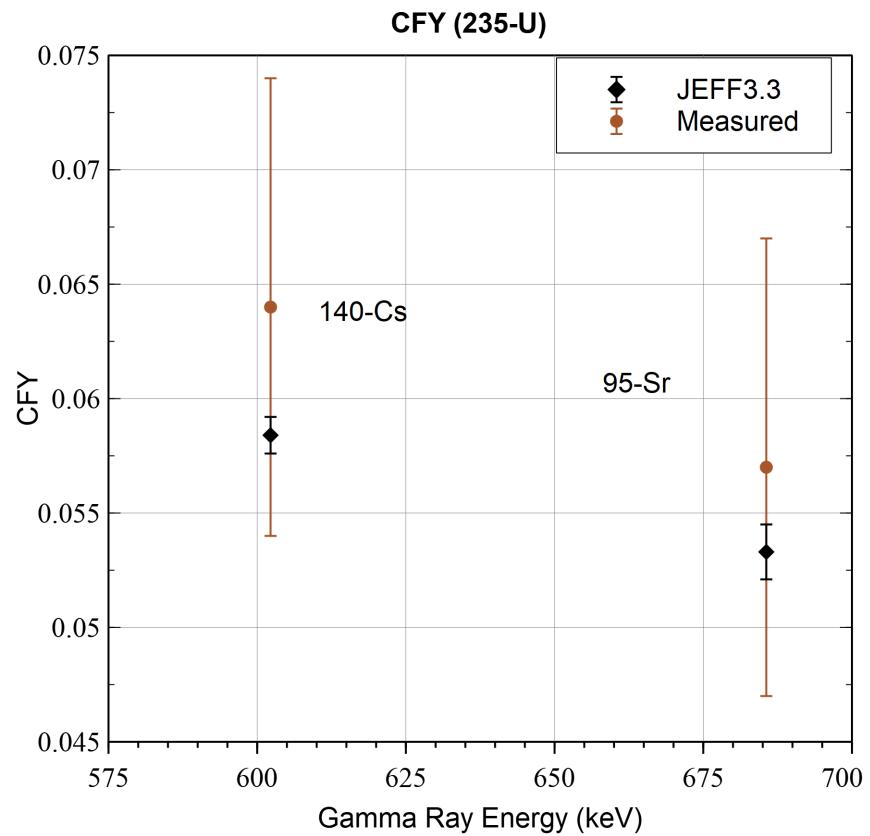


Fig. 8. Comparison of measured CFY and JEFF3.3 library CFY from ^{235}U . Within the uncertainty, measured CFY are consistent with JEFF3.3 CFY.

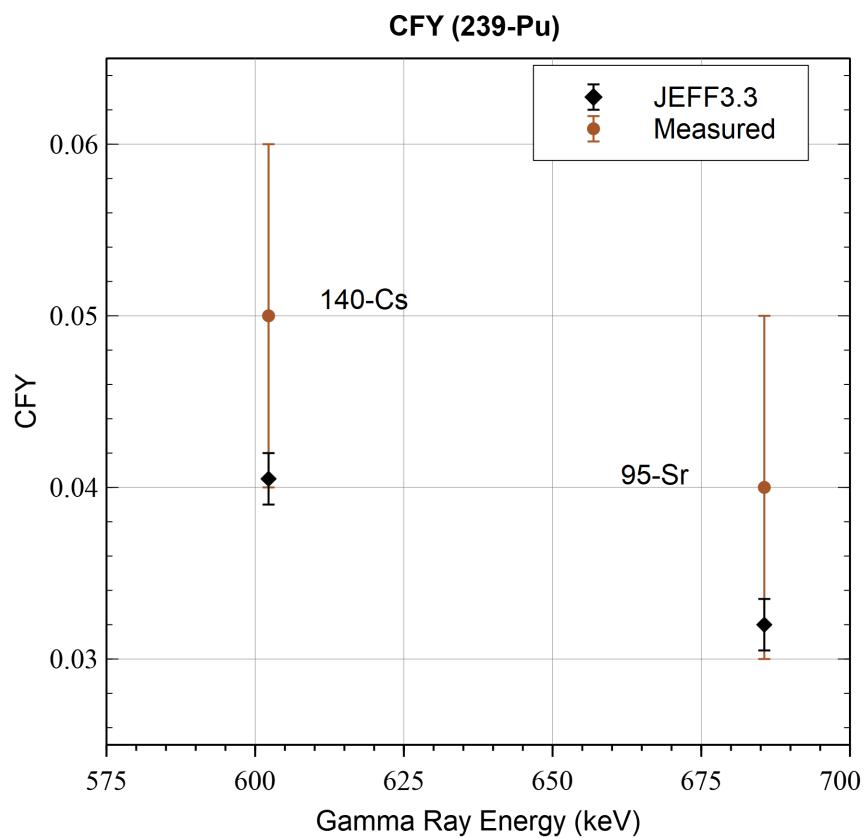


Fig. 9. Comparison of measured CFY and JEFF3.3 library CFY from ^{239}Pu . Within the uncertainty, measured CFY are consistent with JEFF3.3 CFY except ^{96}Y . Measured CFY for ^{96}Y is about 5 factors larger than CFY from JEFF3.3.

308 better shielding to minimize background interference. The fission analysis
309 could be improved by incorporating additional decay paths, such as beta-
310 delayed-neutron emissions and isomers. But this would significantly increase
311 the complexity of the analysis, necessitating the use of more advanced soft-
312 ware tools. Although CFY measurements show consistency, a subsequent
313 study can further enhance their accuracy by accounting for correction fac-
314 tors such as beam fluctuations and gamma-ray attenuation.

315 8. Declaration of competing interest

316 The authors declare that they have no known competing financial inter-
317 ests or personal relationships that could have appeared to influence the work
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