



# Revisiting iron opacity measurements at solar interior temperature

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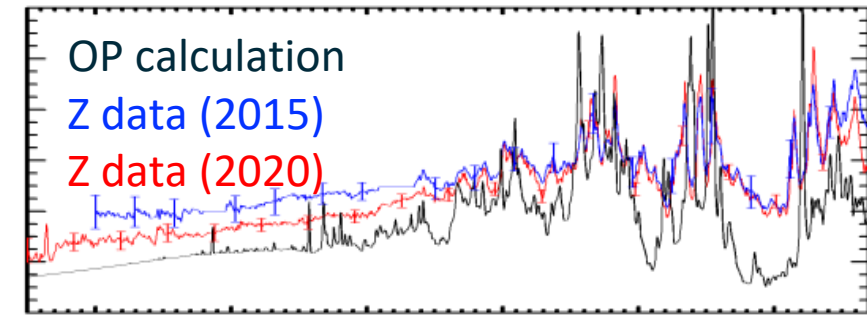


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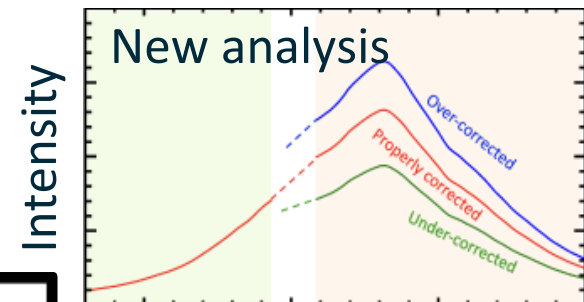
# Novel opacity-analysis methods enhance accuracy for Z opacity data



- There is significant disagreement between measured and modeled iron opacity → Is opacity analysis accurate?
- Large volume of backlight-only data enable accurate analysis  
Concern: Backlight-only data were collected over a decade. → Does backlight change over the decade?
- We developed new methods that do not rely on backlight-only data
- 7% transmission accuracy was confirmed through many synthetic-data tests



Wavelength



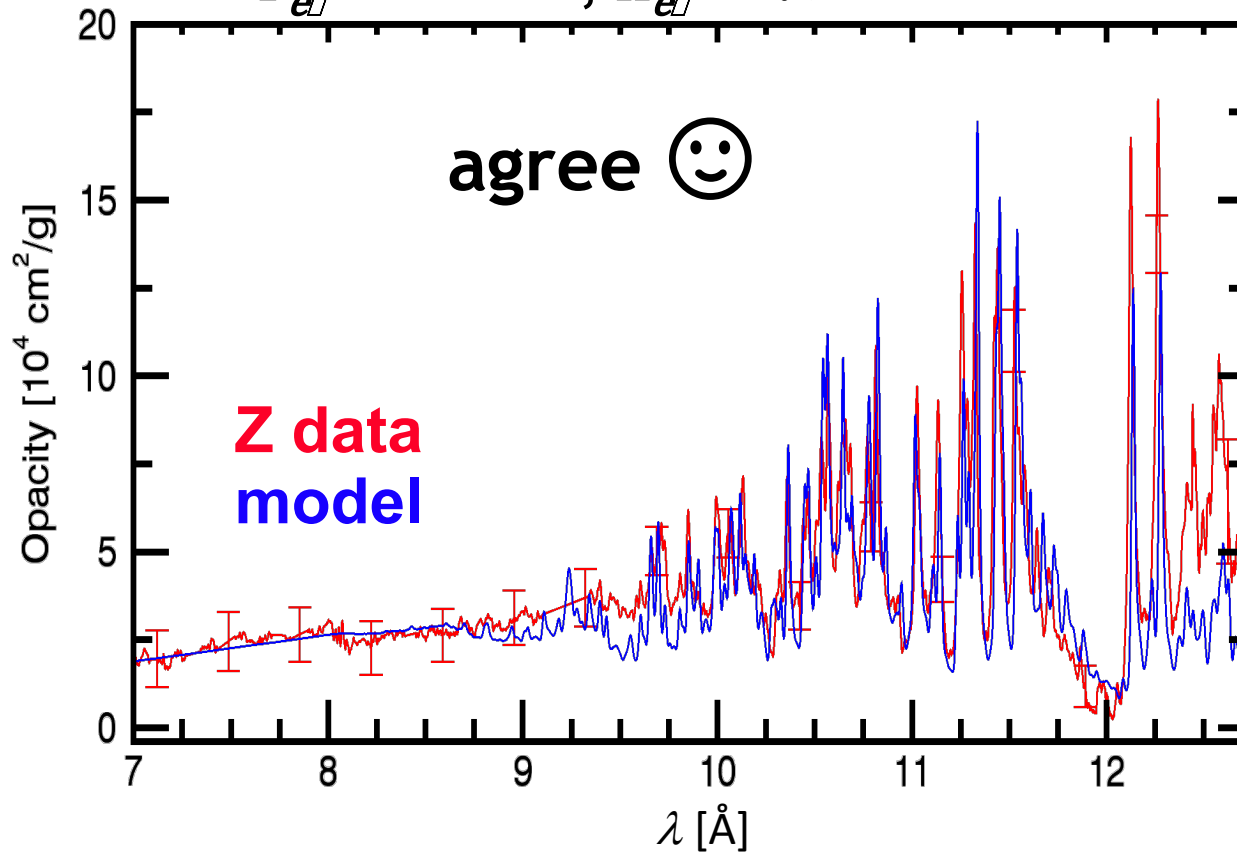
**We are working on re-analysis of all iron data from ~40 experiments and quantifying its impact on the solar problem**

We measured iron opacity at multiple conditions and found severe disagreement with models at solar temperatures



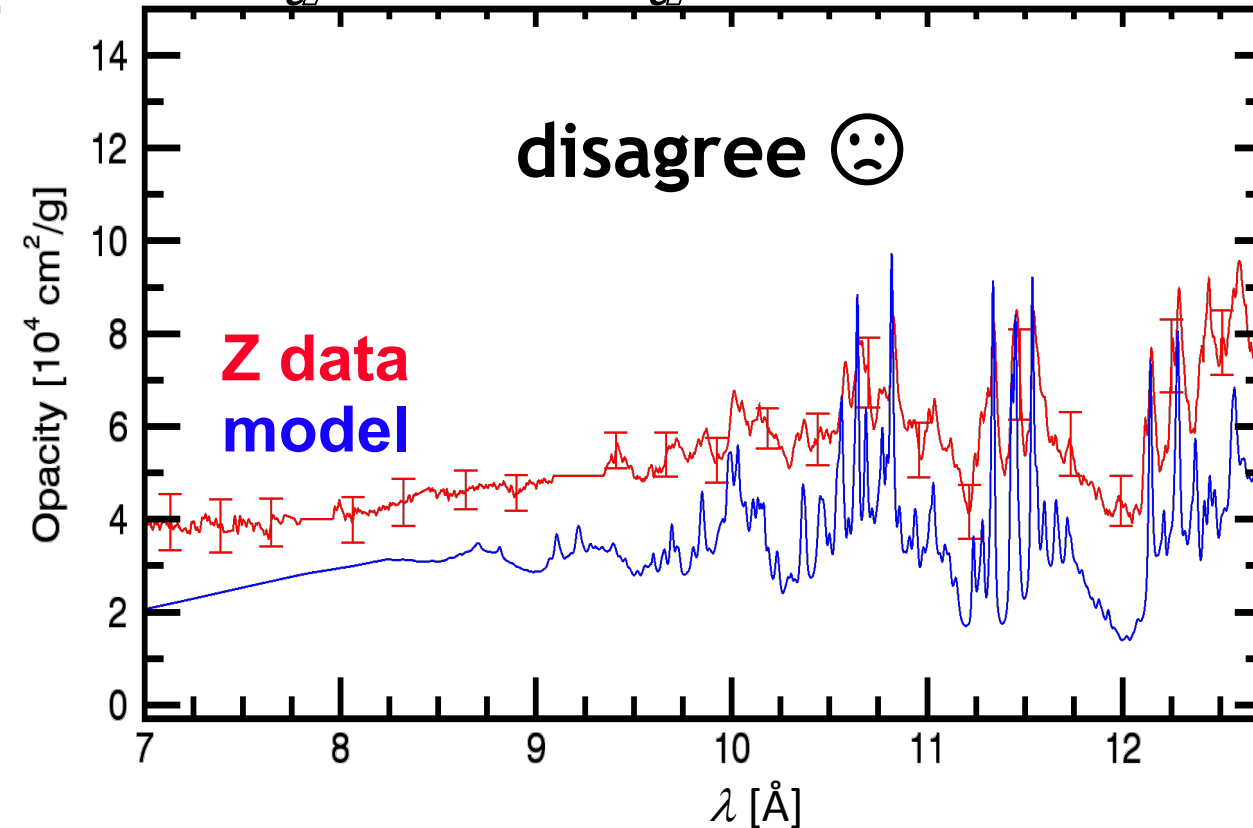
**Anchor 1**

$$T_e = 156 \text{ eV}, n_e = 6.9 \times 10^{21} \text{ cm}^{-3}$$



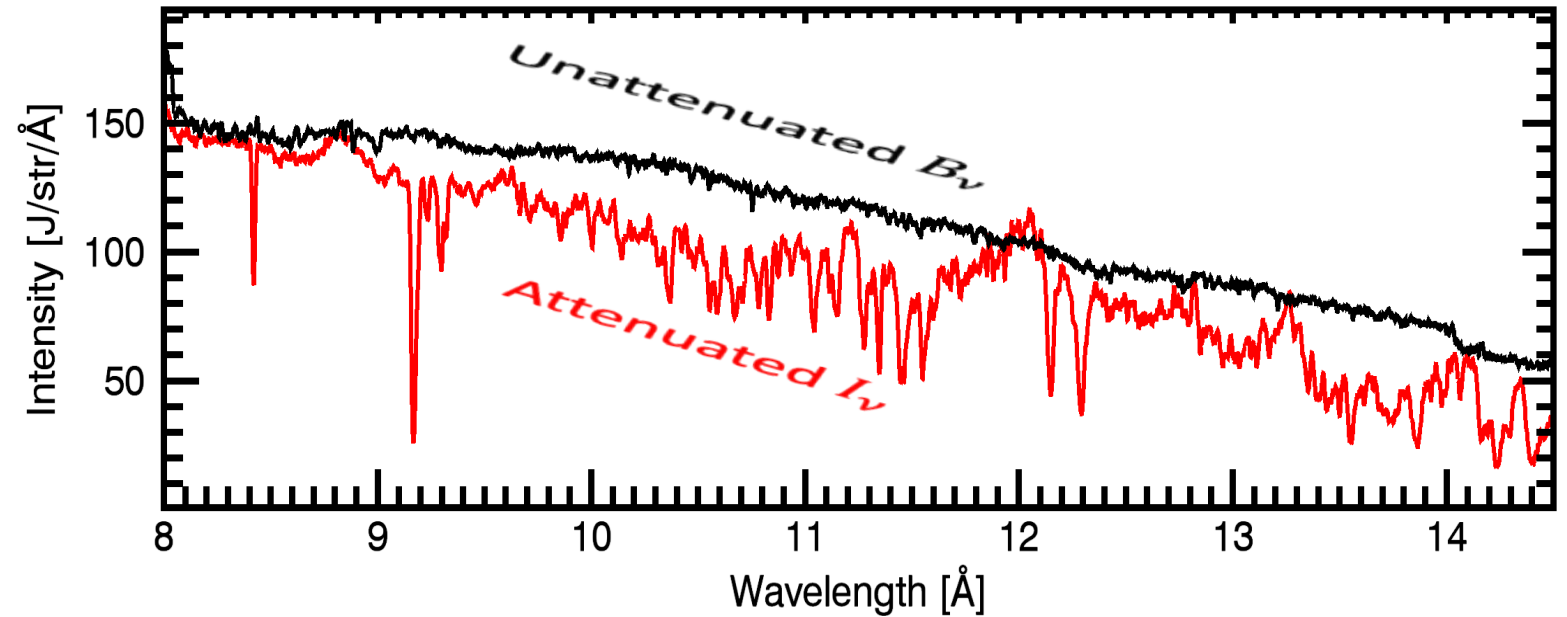
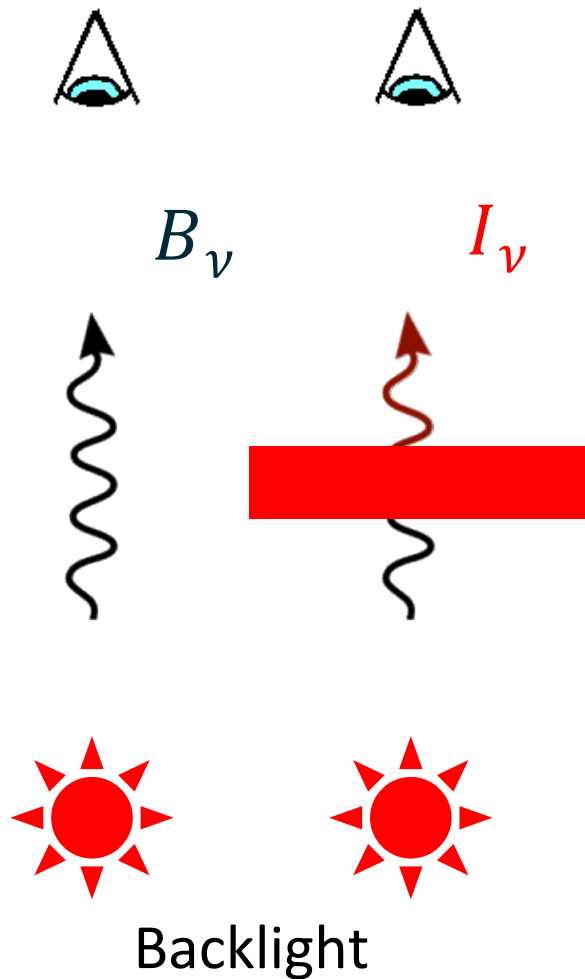
**Anchor 2 ~ solar temperature**

$$T_e = 182 \text{ eV}, n_e = 3.1 \times 10^{22} \text{ cm}^{-3}$$



**Key question: is the analysis accurate?**

# Sample opacity is inferred by measuring backlight with and without the sample



1. Determine transmission  $T_v$

$$T_v = \frac{I_v - \epsilon_v}{B_v - \epsilon_v}$$

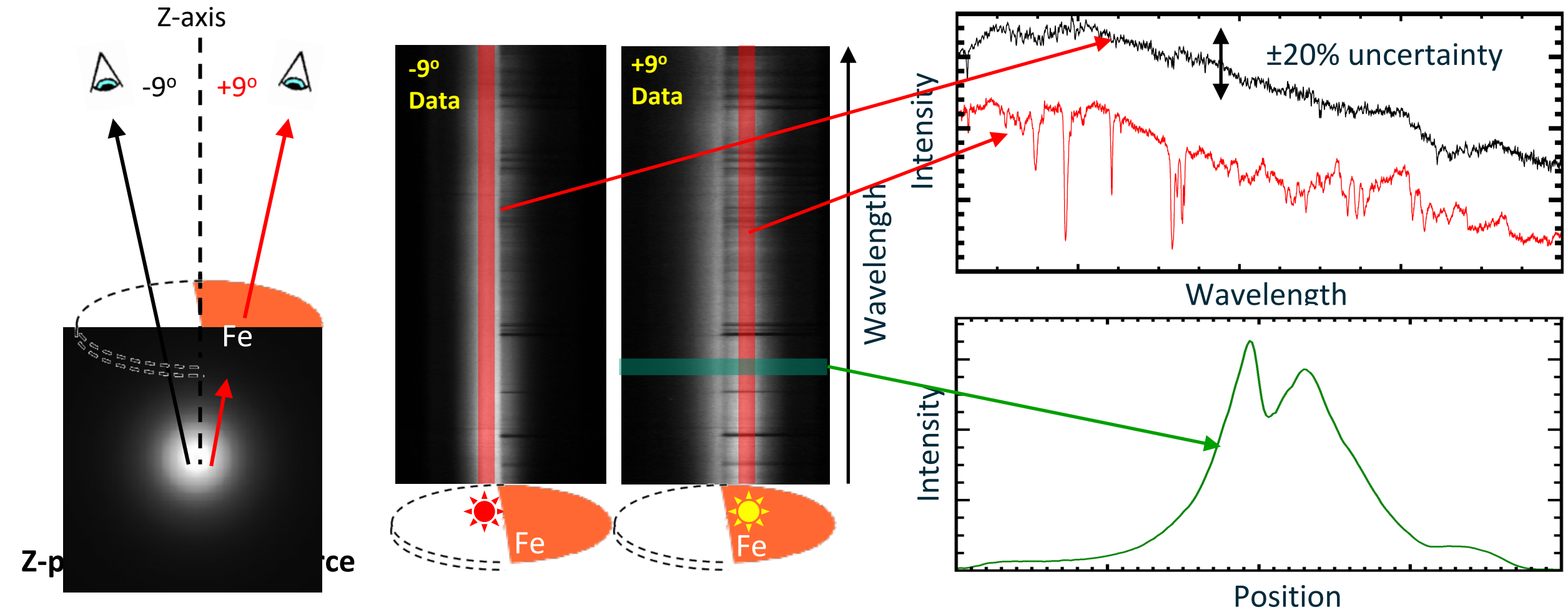
2. Convert  $T_v$  to opacity  $\kappa_v$

$$\kappa_v = -\ln T_v / \rho L$$

## Sources of uncertainty

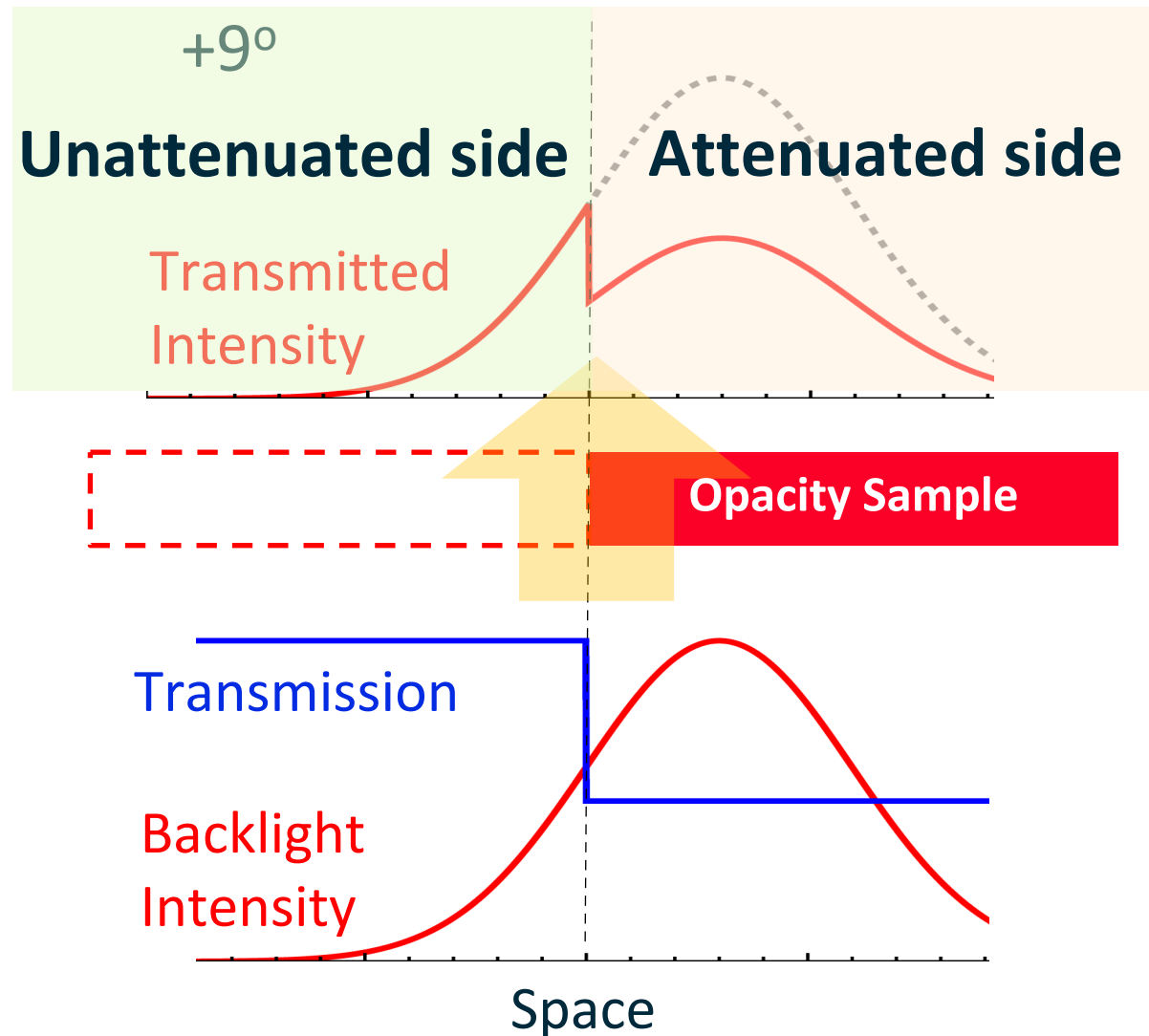
- Unattenuated  $B_v$
- Background  $\epsilon_v$
- Sample thickness  $\rho L$

Transmission spectra is determined by dividing attenuated by unattenuated spectra  $\rightarrow \pm 20\%$  uncertain



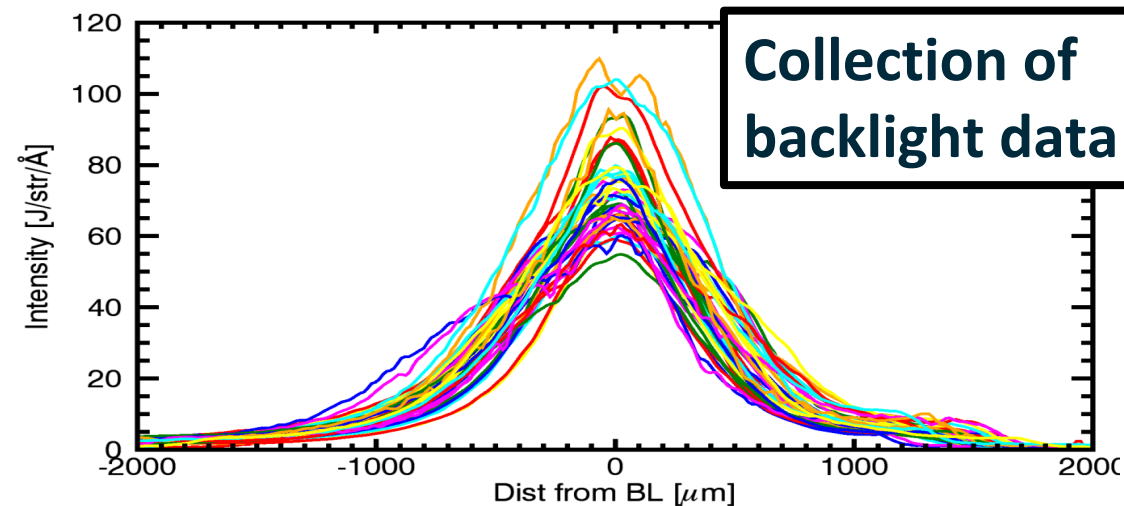
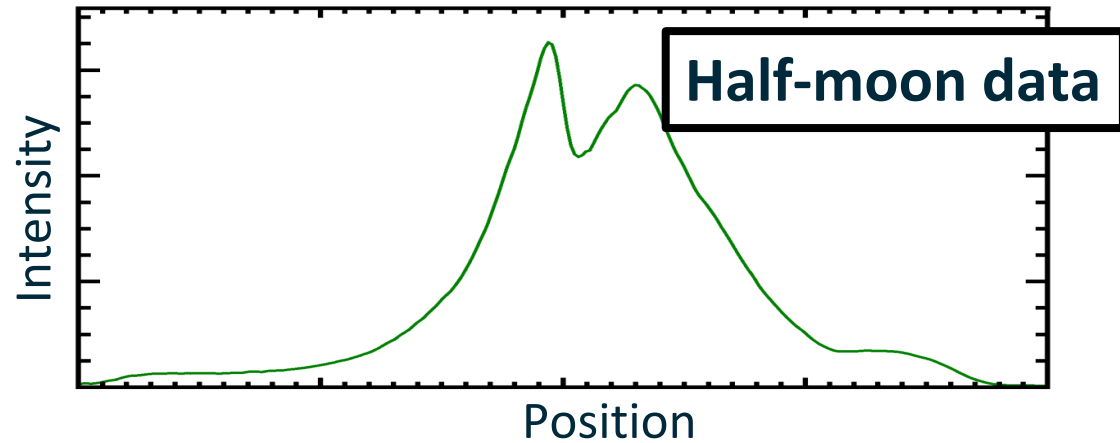
We use spatial shape to improve our accuracy of our transmission analysis

Spatial shape has unattenuated and attenuated side and provide essential clue on transmission

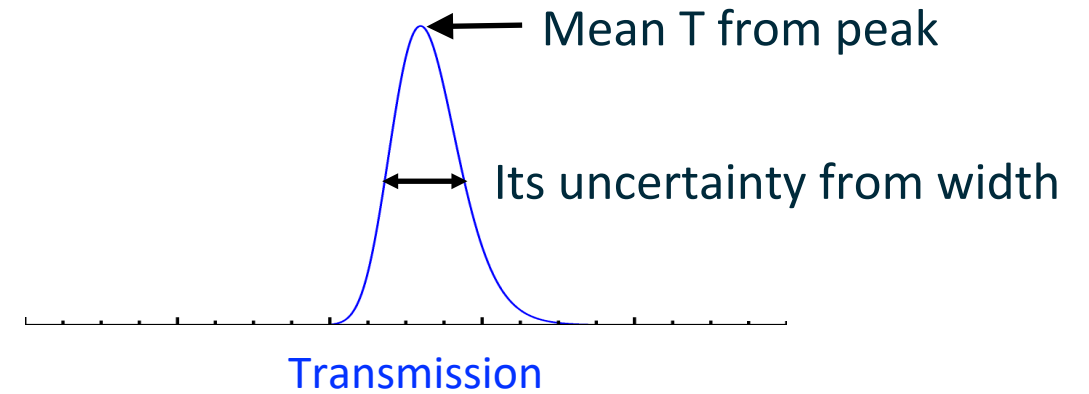




# We analyze measured half-moon sample aided by backlight statistics to improve transmission accuracy



Transmission Probability Distribution (TPD)

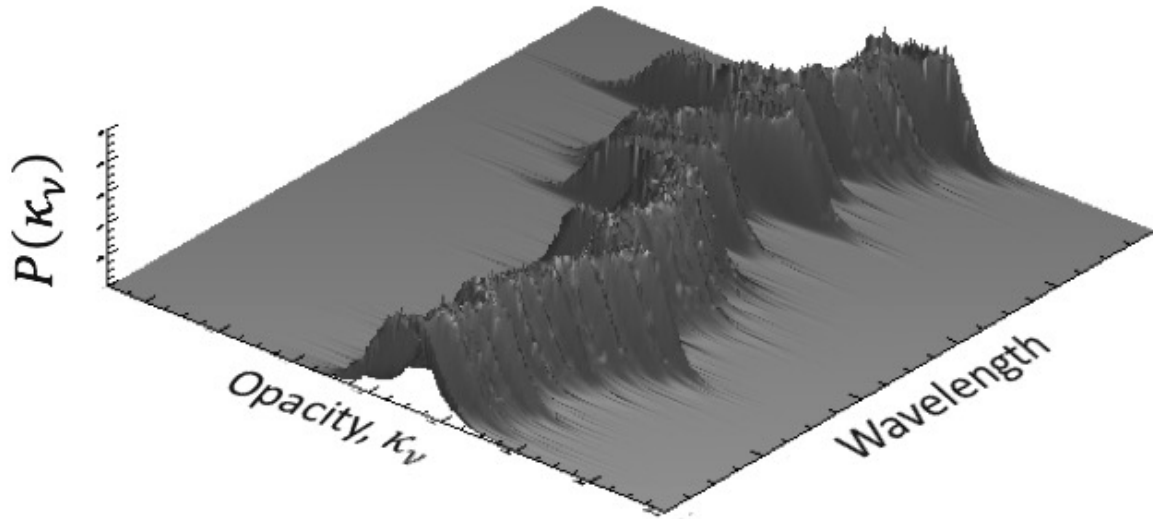


We developed multiple TPD methods that rely on different backlight statistics.

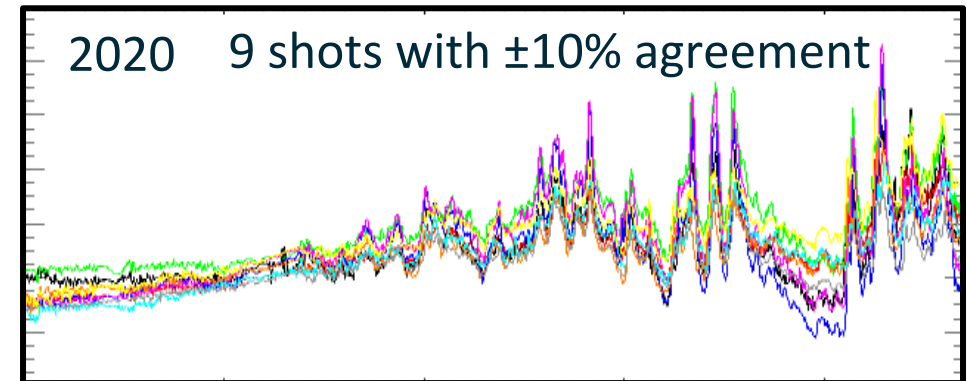
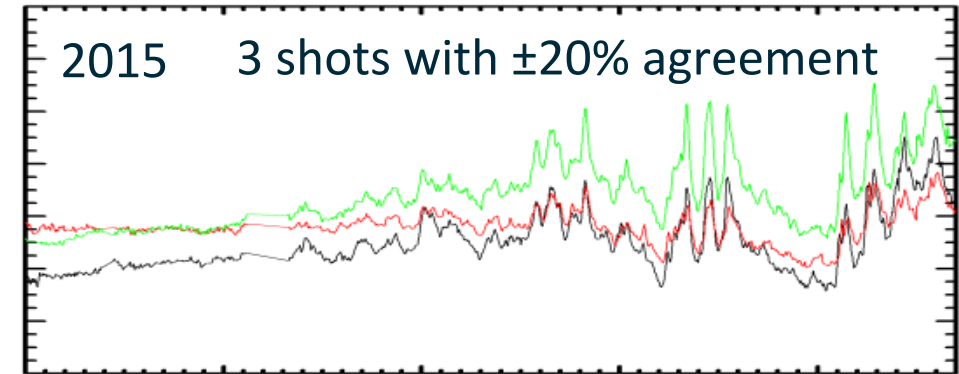
# TPDs\* are converted to opacity probability distribution; This method significantly improved analysis accuracy



Asymmetric non-Gaussian opacity PDF\*



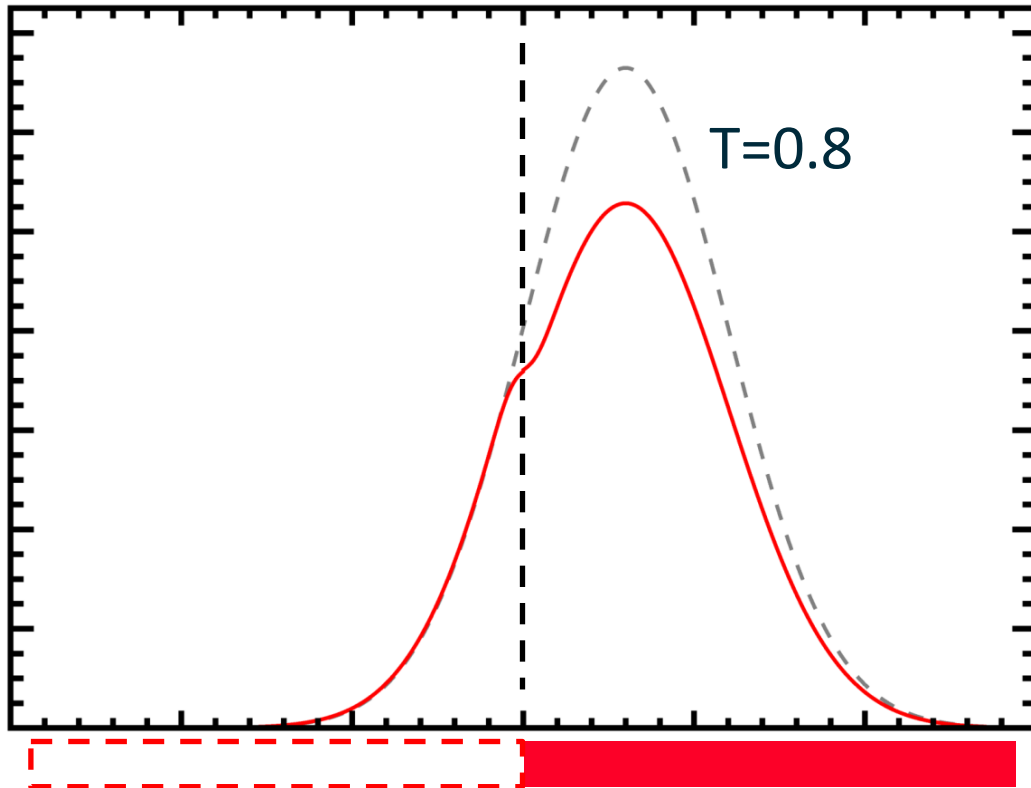
- Large volume of backlight-only data statistics
- Monte Carlo for robust errors propagations
  - Backlight intensity,  $B_v$
  - Background,  $\epsilon_v$
  - Sample areal density,  $\rho L$



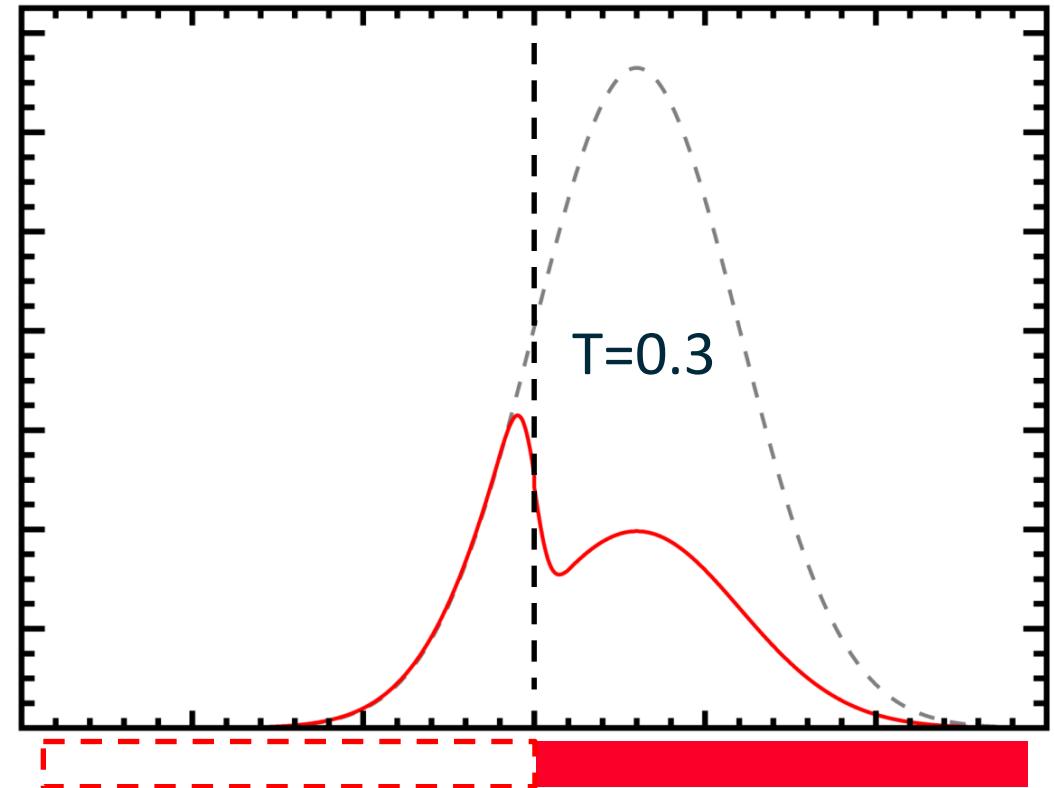
**Concern: Backlight-only data were collected over more than a decade**  
**→ Can we assess opacity independently of the backlight data?**



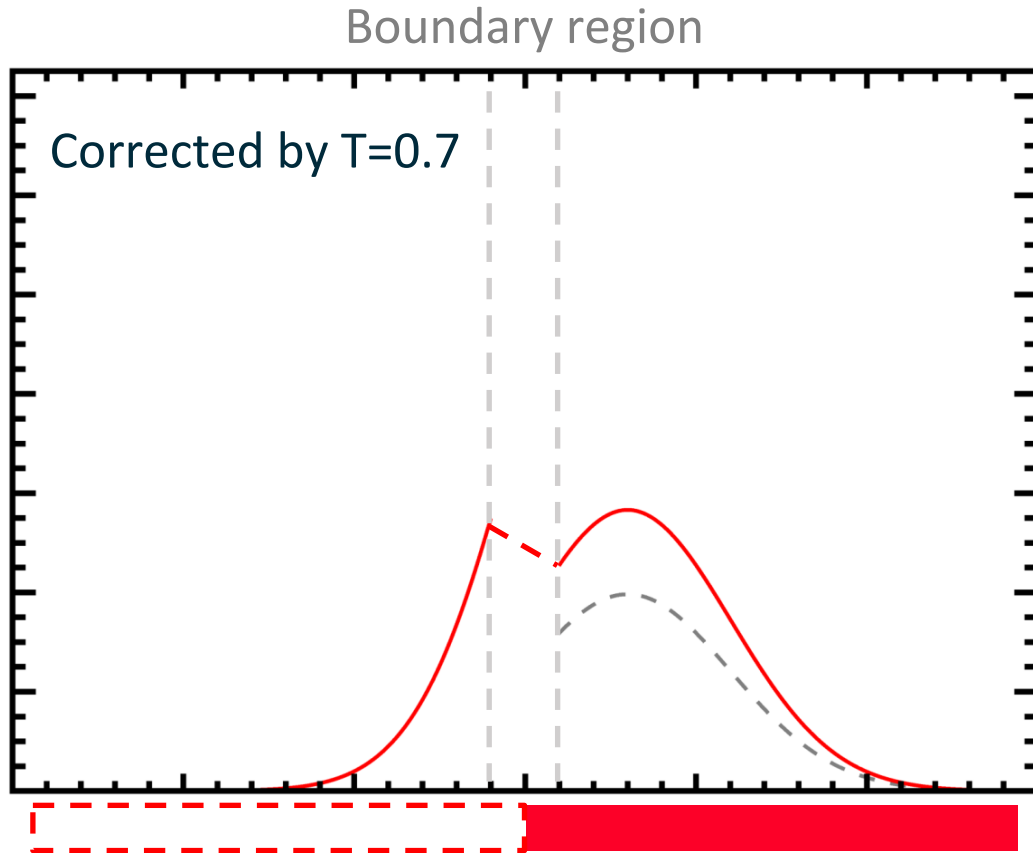
# Which half-moon spatial profile corresponds to lower sample transmission?



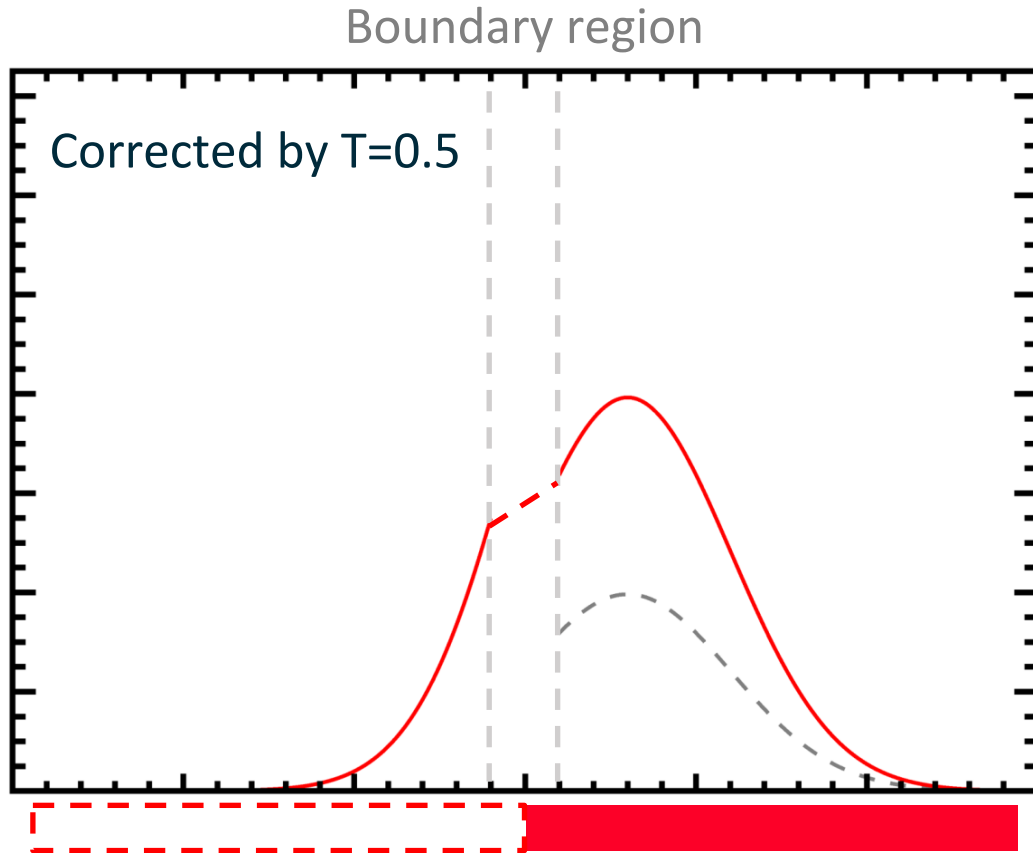
More drastic drop over the boundary



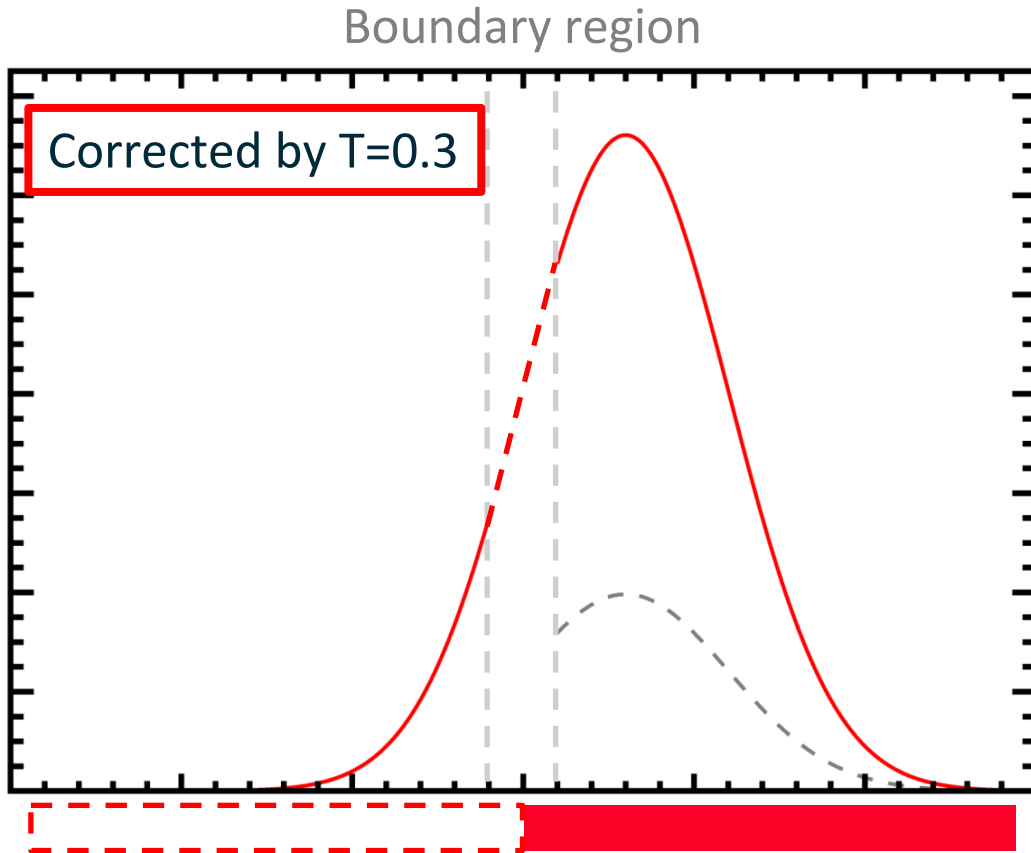
We can search for the transmission that makes the T-corrected profile nice and smooth



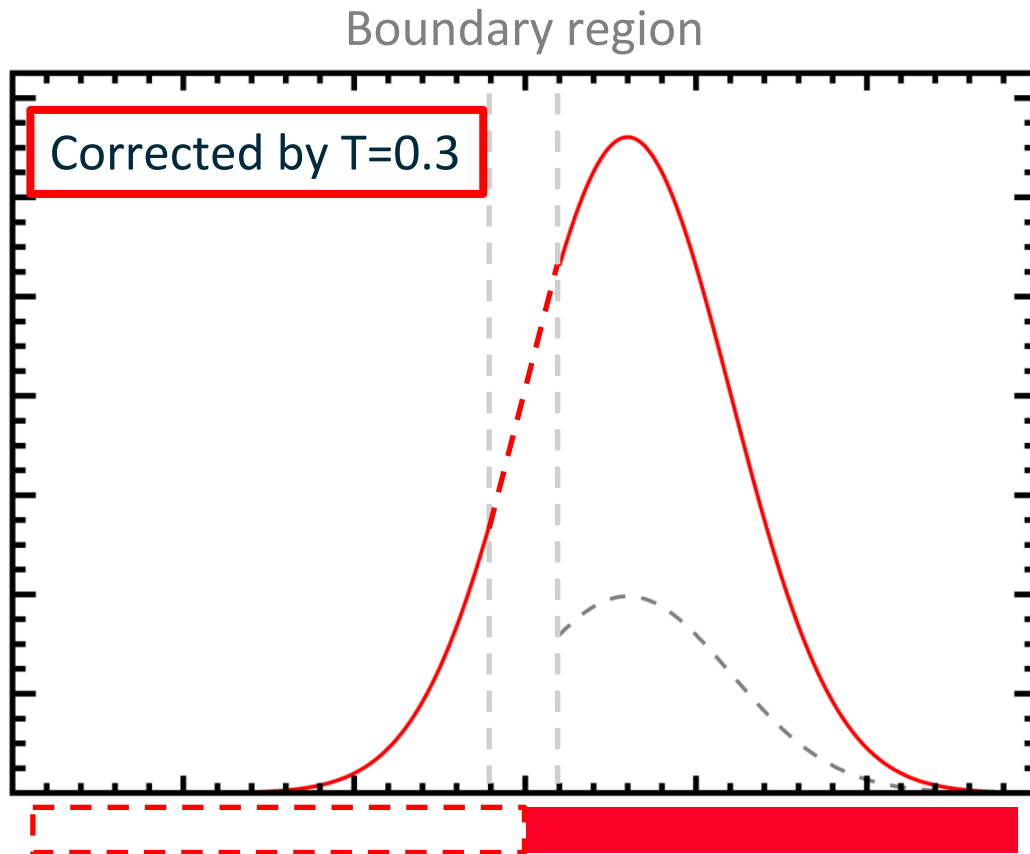
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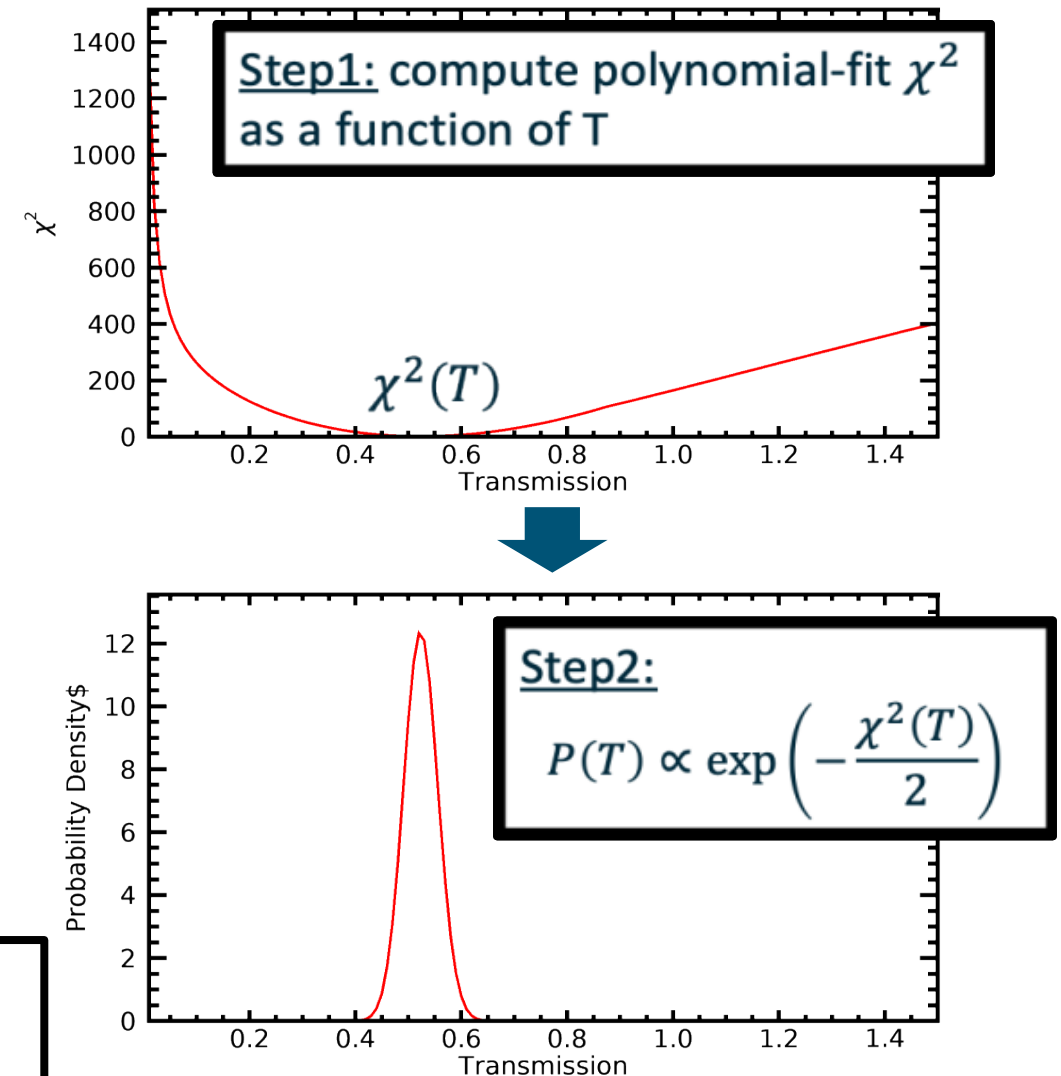
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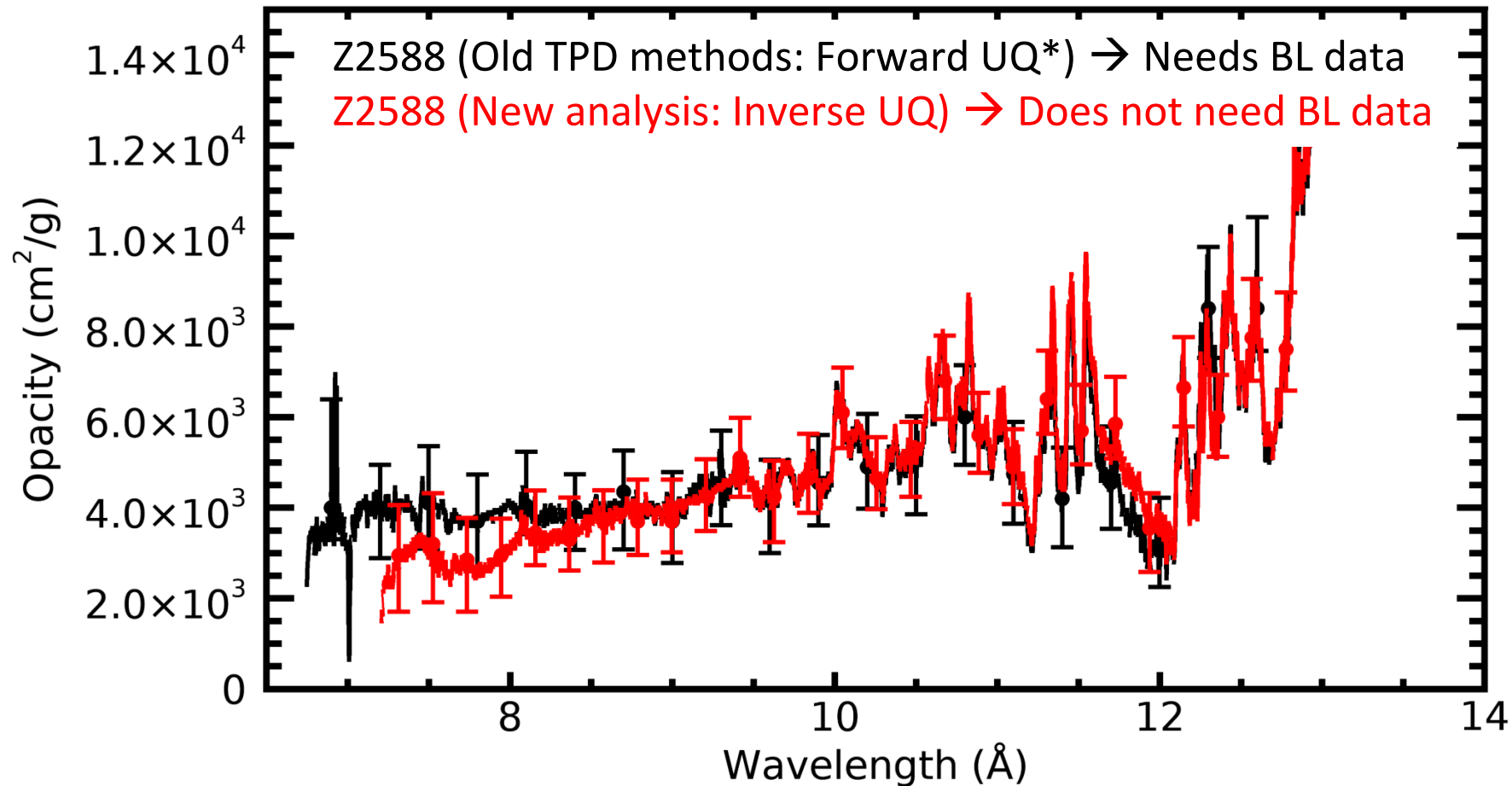


We have tested this method with many synthetic half-moon data created from backlight-only data



# The new method is applied on a few experiments

## The preliminary results are encouraging



**We are working on applying both techniques to all iron data**  
**Reanalysis involves ~40 iron experiments and >1,000 spectral images**

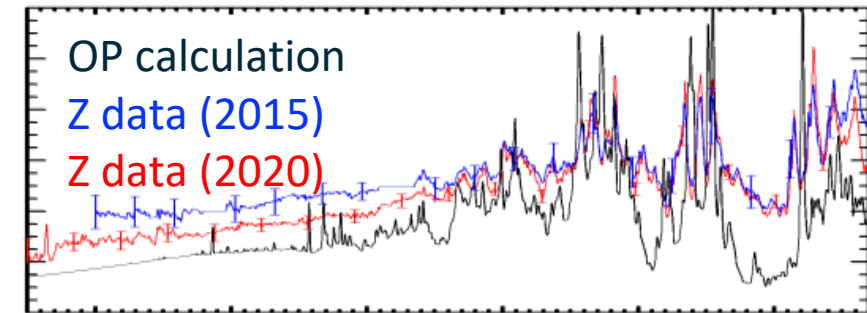
\*UQ = Uncertainty quantification



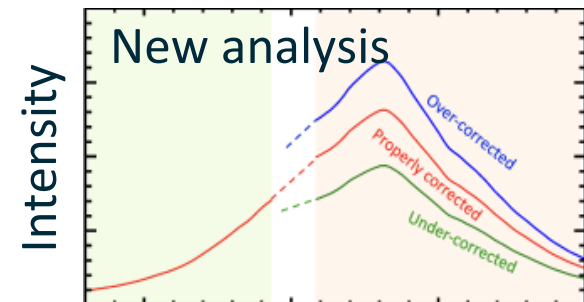
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