



Modeling the Effects of Radiation Reaction for Next Generation Pulsed Power Machine

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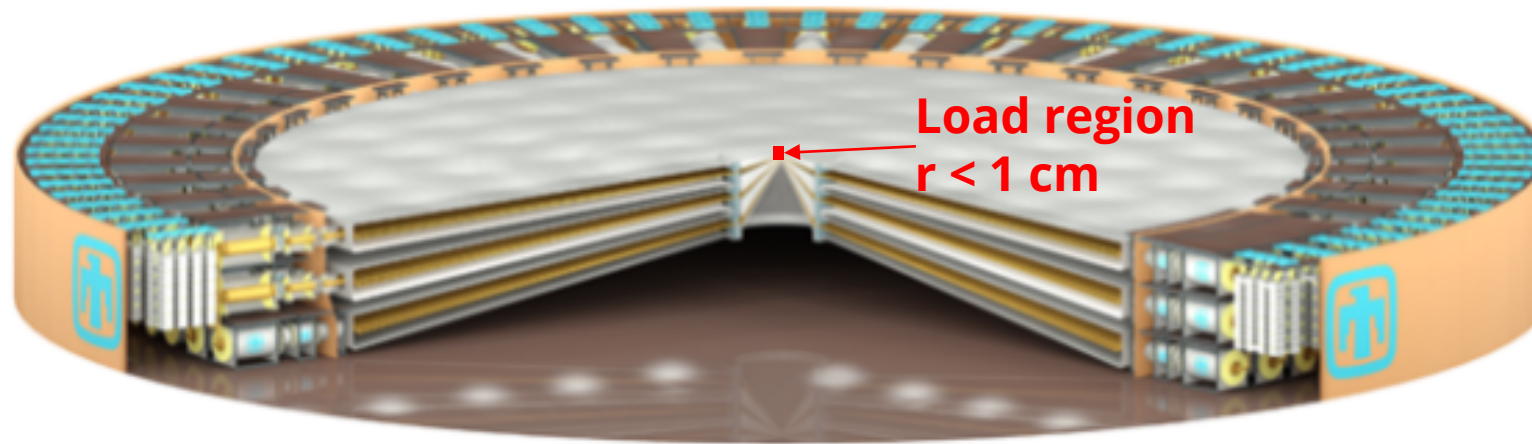
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Next Generation Pulsed Power (NGPP) Machine



- Sandia National Laboratories is designing a Next Generation Pulsed Power machine, which is capable of delivering >60 MA of current to a load.
- Present designs are investigating current pulse lengths in the range of 100 ns - 200 ns.
- Strong B-fields cause electrons to emit radiation and lose energy through a radiation reaction force. In this study, we examine how important radiation reaction effects are on power flow electrons outside of the load.
- At a radius of 1 mm: 60 MA \rightarrow 12,000 T.
- At a radius of 100 μm : 60 MA \rightarrow 120,000 T.

One design being
considered for
NGPP



What is the radiation reaction force?



- When accelerating charged particles, such as electrons, emit radiation as defined by the Larmor formula, the emission causes a radiation reaction force on the particle.
- From Ford and O'Connell, Phys. Lett. A, 174, pp. 182-184 (1993):

Lorentz Force Law:

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

Lorentz-Abraham-Dirac Equation:

not in existing power flow codes

$$\begin{aligned} \frac{d\mathbf{P}}{dt} &= \mathbf{F} + \mathbf{F}_{rad} \\ &= \mathbf{F} + \left[\tau \left(\gamma \frac{d\mathbf{F}}{dt} - \frac{\gamma^3}{c^2} \left(\frac{d\mathbf{v}}{dt} \times (\mathbf{v} \times \mathbf{F}) \right) \right) \right] \end{aligned}$$

Radiation Reaction Time-Constant for Electrons:

$$\tau = \frac{q^2}{6\pi\epsilon_0 mc^3} = 6.3 \times 10^{-24} s$$

Radiation Reaction in Crossed Uniform E and B Fields



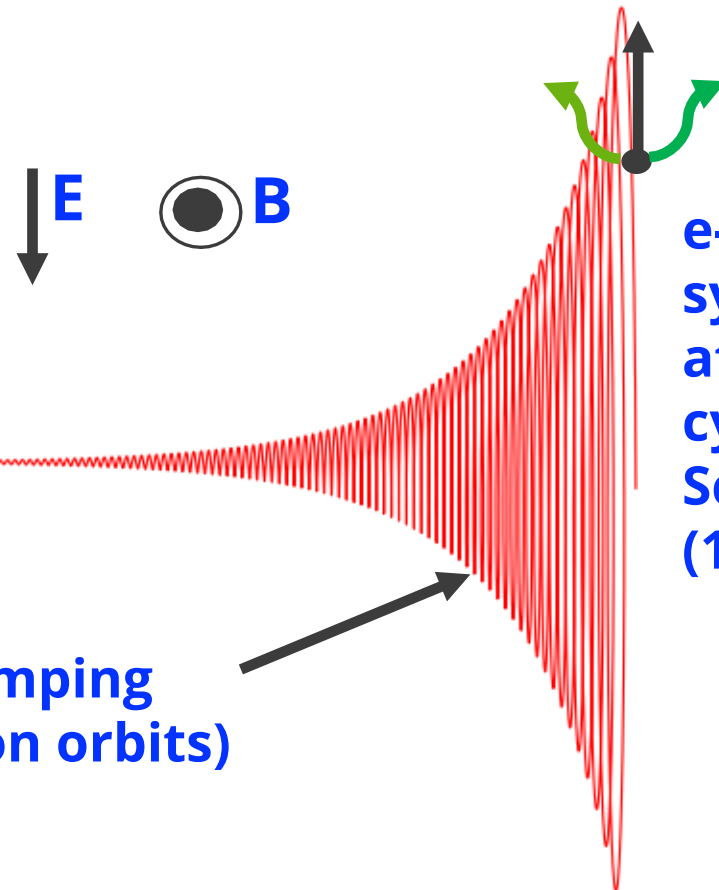
**Lorentz-Abraham-Dirac Equation:
In E x B Drift Frame ($P'_{||} = 0$)**

$$\frac{d\mathbf{P}'_{\perp}}{dt'} = -e\mathbf{v}'_{\perp} \times \mathbf{B}' - \frac{\gamma'\mathbf{P}'_{\perp}}{\tau_{damp}}$$

$$\tau_{damp} = \frac{m^2}{e^2 B'^2 \tau} \quad B' = B \sqrt{1 - \frac{E^2}{B^2 c^2}} \simeq B$$

**Undamped
Uniform E x B Drift**

**Damping of Cyclotron Motion
(for NGPP fields near a load damping
may occur over 10^7 - 10^9 cyclotron orbits)**



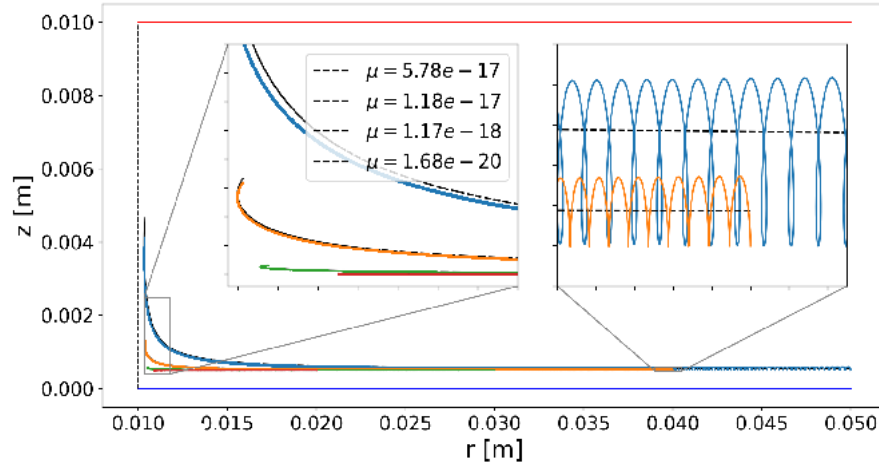
**e- emitting
synchrotron radiation
at harmonics of the
cyclotron frequency:
Schwinger, Phys. Rev.
(1949)**

Drift – Kinetic Modeling Approach for Radial MITLs

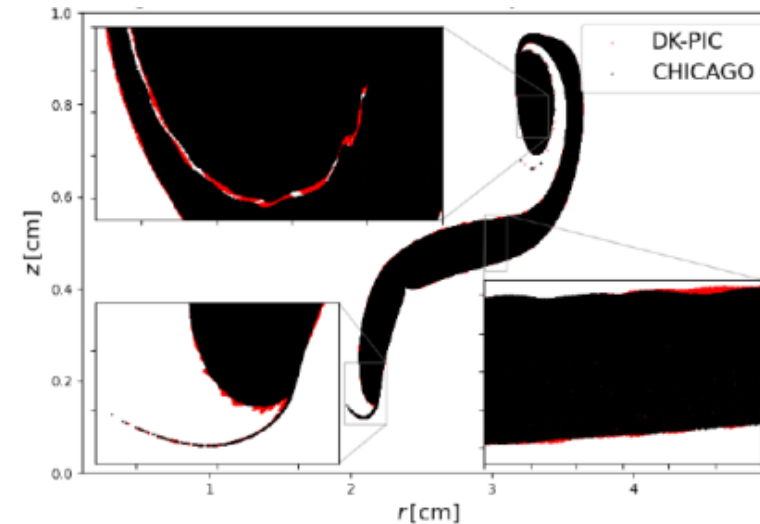


$$\mathbf{v}_{gc} = \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \frac{\mu}{q\gamma} \frac{\mathbf{B} \times \nabla B}{B^2} \quad \mu = \frac{p_{r,osc}^2 + p_{z,osc}^2}{2mB}$$

Guiding Center Drift Motion and Magnetic Moment



**Hess and Evstatiev,
IEEE Trans. Plasma Sci. (2021)
(vacuum fields with
magnetic moment conserved)**



**Evstatiev and Hess, TP11.00092
(vacuum/space-charge fields with
magnetic moment conserved)**

$$\mathbf{v}_{gc} = \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \frac{\mu(t)}{q\gamma} \frac{\mathbf{B} \times \nabla B}{B^2} \quad \frac{d\mu}{dt}\bigg|_{rad} = -\frac{2\mu}{\tau_{damp}} \sqrt{1 + 2\mu B/mc^2}$$

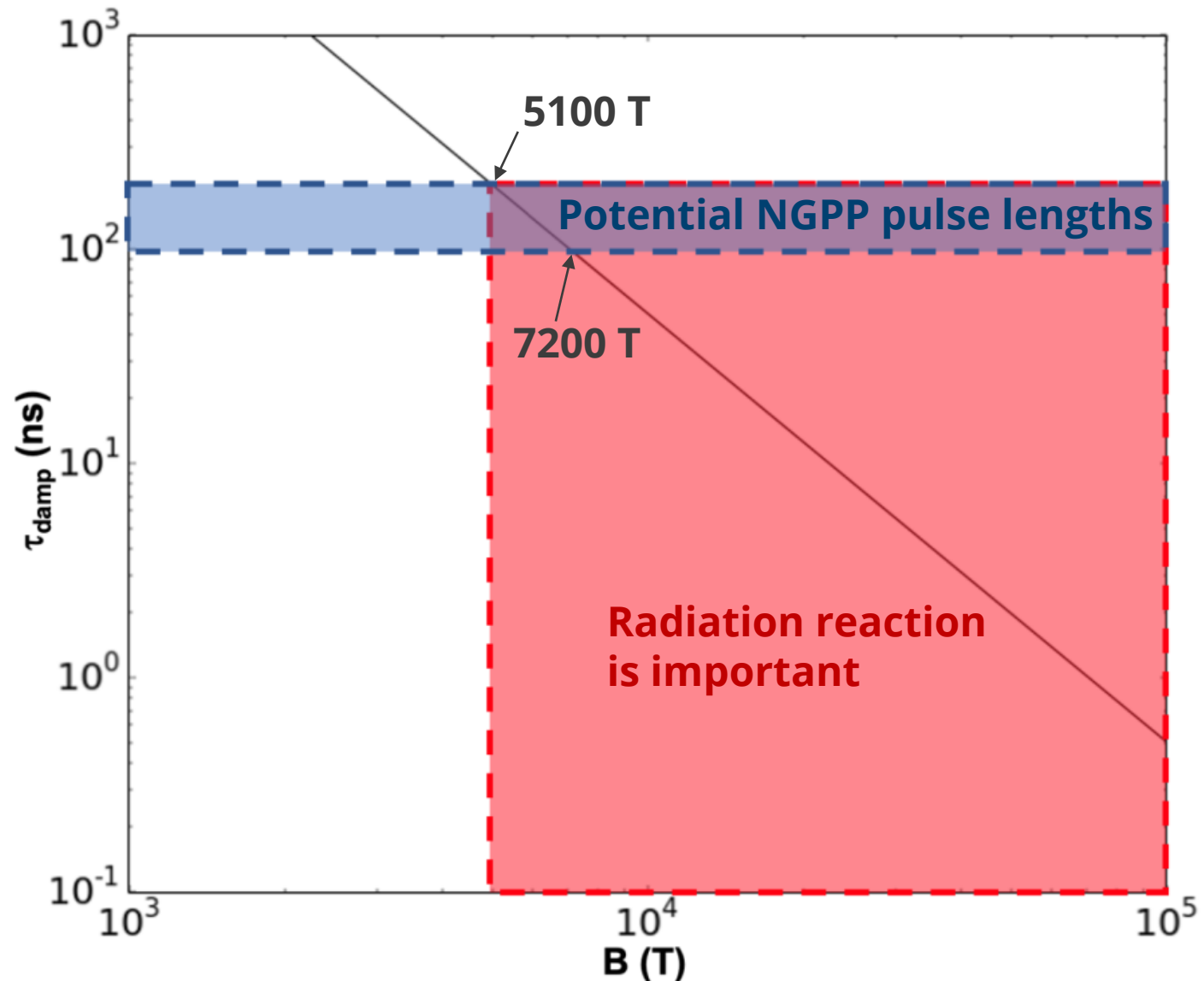
**Present work* (vacuum fields with
time-dependent magnetic moment
due to radiation reaction)**

***Hess and Evstatiev, SAND Report (2021) <https://www.osti.gov/servlets/purl/1814239>**

When does radiation reaction become important for NGPP?

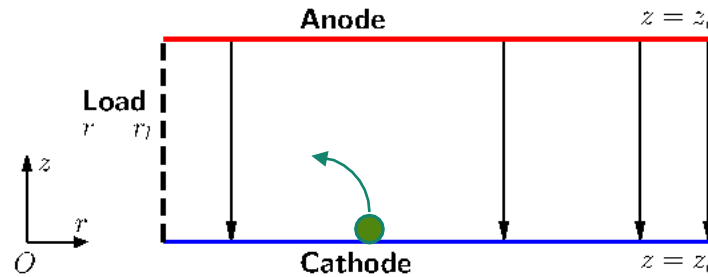


- One criterion for radiation reaction effects to become important on NGPP is the damping time becomes equal to (or less than) the NGPP pulse length.
- For a pulse length of 100 ns, radiation reaction becomes important $B > 7200$ T.
- For a pulse length of 200 ns, radiation reaction becomes important $B > 5100$ T.



Radial MITL Field Equations:

$$E = -\frac{\mu_0 I}{2\pi} \ln \frac{r}{r_l} e_z \quad B = -\frac{\mu_0 I(t) \theta}{2\pi r} \quad I(t) = I_{peak} \sin^2 \frac{\pi t}{2t_{pulse}}$$



Radial MITL with fixed load driven by time-dependent current / electron emitted at cathode

Drift Equations:

$$\mathbf{v}_{gc} = \frac{\mathbf{E} \times \mathbf{B}}{B^2} + \frac{\mu(t)}{q\gamma} \frac{\mathbf{B} \times \nabla B}{B^2} \quad \mu|_{t_0} = \frac{mE^2}{2B^3}$$

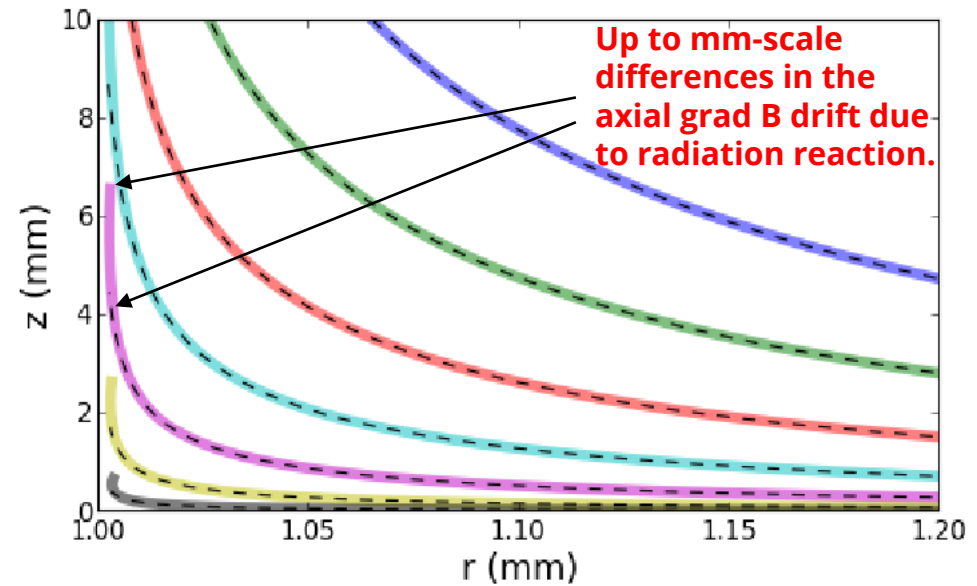
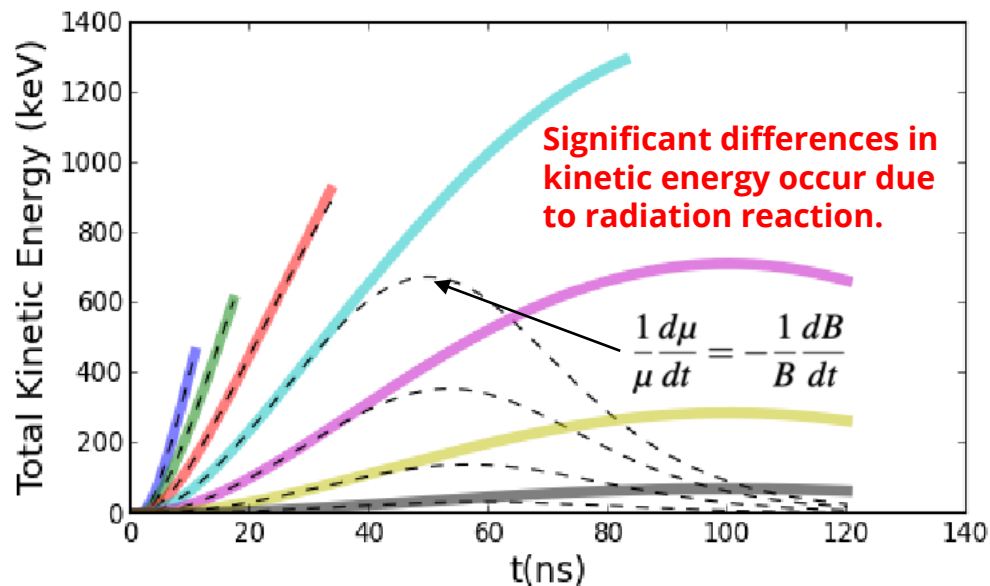
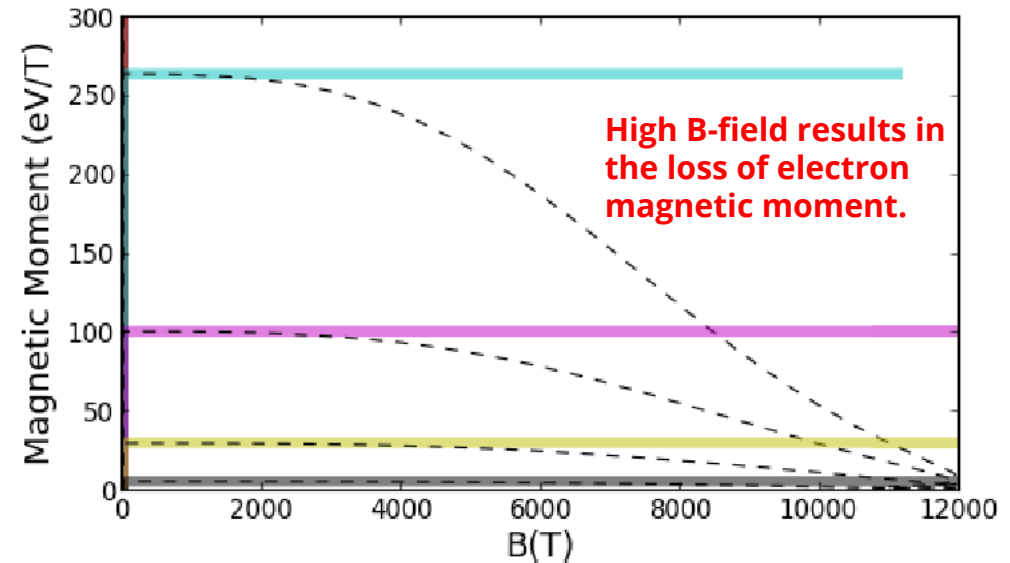
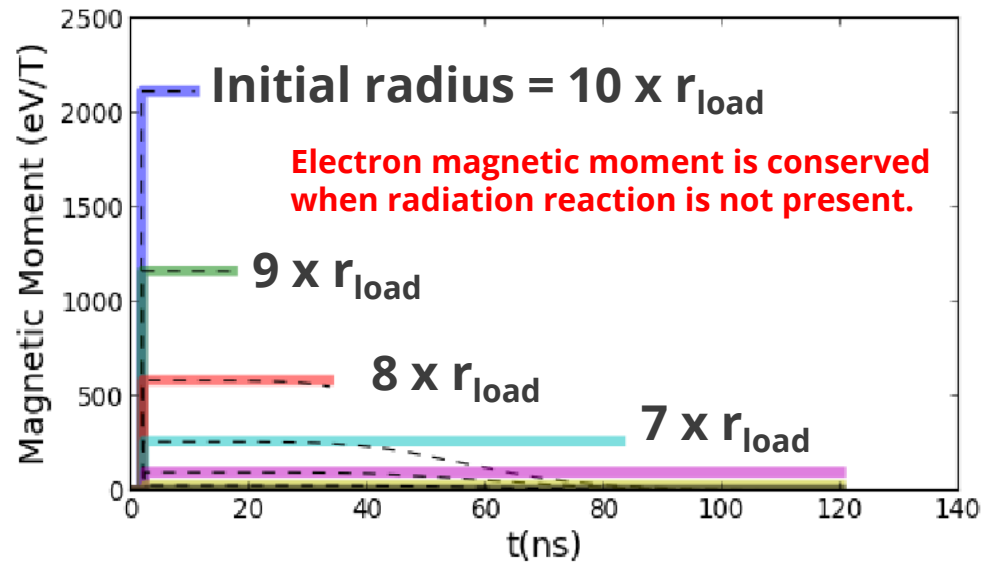
Magnetic Moment / Energy Equations:

$$p_{cyc} = \sqrt{2\mu m B} \quad K_{cyc} = (\sqrt{1 + 2\mu B/mc^2} - 1)mc^2$$

$$\frac{d\mu}{dt}|_{rad} = \frac{2\mu}{p_{cyc}} \frac{dp_{cyc}}{dt}|_{rad} = -\frac{2\mu}{\tau_{damp}} \sqrt{1 + 2\mu B/mc^2}$$

$$K_{tot} = K_{cyc} + \frac{1}{2}mv_{gc}^2$$

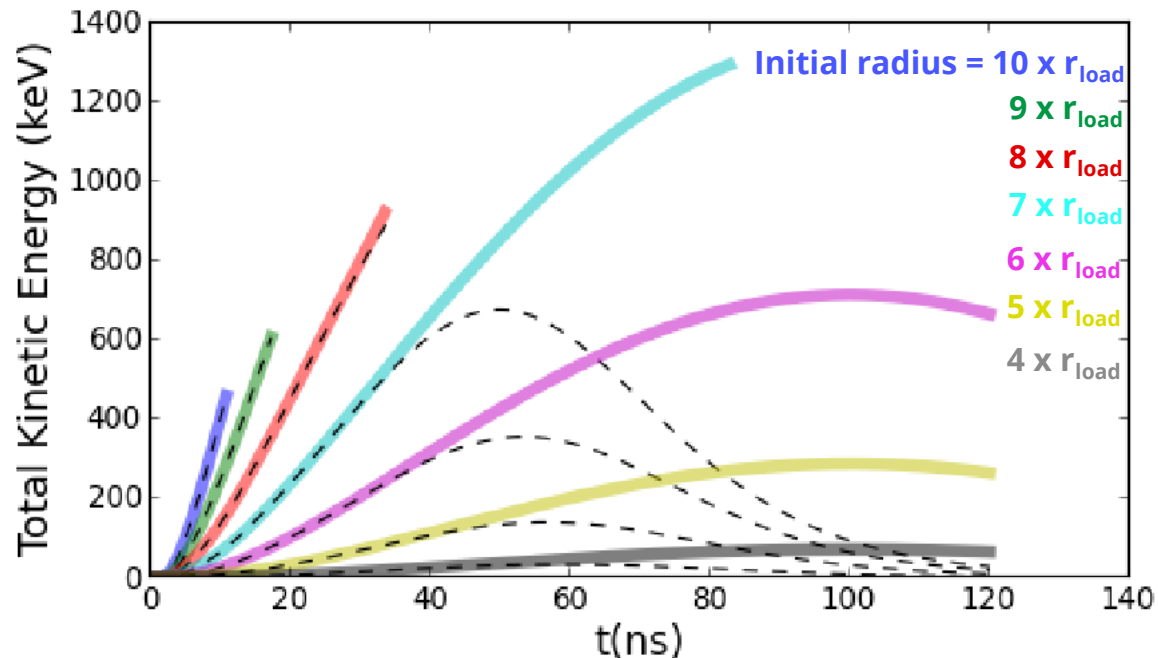
Results for Radial MITL with $I_{\text{peak}} = 60 \text{ MA}$, $t_{\text{pulse}} = 100 \text{ ns}$, $r_{\text{load}} = 1 \text{ mm}$ (solid no radiation reaction , dashed with radiation reaction)



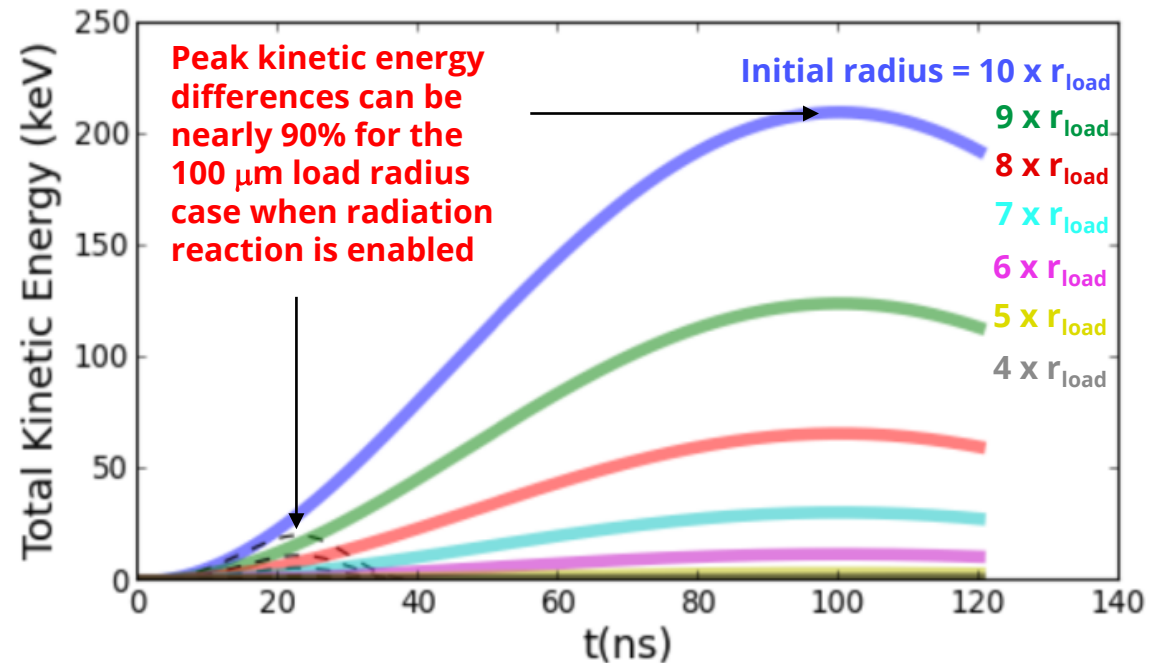
Decreasing the Load Radius Enhances Kinetic Energy Reduction Due to Radiation Reaction

- By decreasing the load radius to 100 μm (and keeping other MITL parameters fixed), the B-field increases by a factor of 10x near the load compared to the previous load radius = 1 mm case.
- The damping time $\propto 1/B^2$ decreases by 100x causing enhanced radiation reaction effects.

Load Radius of 1 mm



Load Radius of 100 μm



- We simulated single particle electron motion and kinetic energy loss due to radiation reaction in a high-current radial MITL near a load for NGPP design parameters using a guiding center drift approach.
 - The key piece of physics is the introduction of a time-dependent magnetic moment which is damped due to radiation reaction.
- Electrons in B-fields $>10,000$ T (which will occur in the MITL near a load ≤ 1 mm in 60+ MA drivers) can experience significant gyrokinetic energy loss/prevention of kinetic energy gain due to radiation reaction.
- Radiation reaction also affects axial grad B drift motion in the radial MITL by damping the electron's magnetic moment.
- Future work will include coupling radiation reaction physics with space-charge fields for modeling electron dynamics in radial MITLs.