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# Water in Silicates: A Combined Shock & Spectroscopy Study

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## Can SiO<sub>2</sub> glass preserve volatiles during impacts?

- Water is abundant in planetary building blocks
- Deuterium to Hydrogen (D/H) ratio (Fig 1.) numerically represents the combination of sources of Earth's water
- One potential contribution to D/H ratio is water deposited via impactors (with  $v_{\text{impact}}$  on Earth ranging from 11 [km/s] to 53 [km/s])

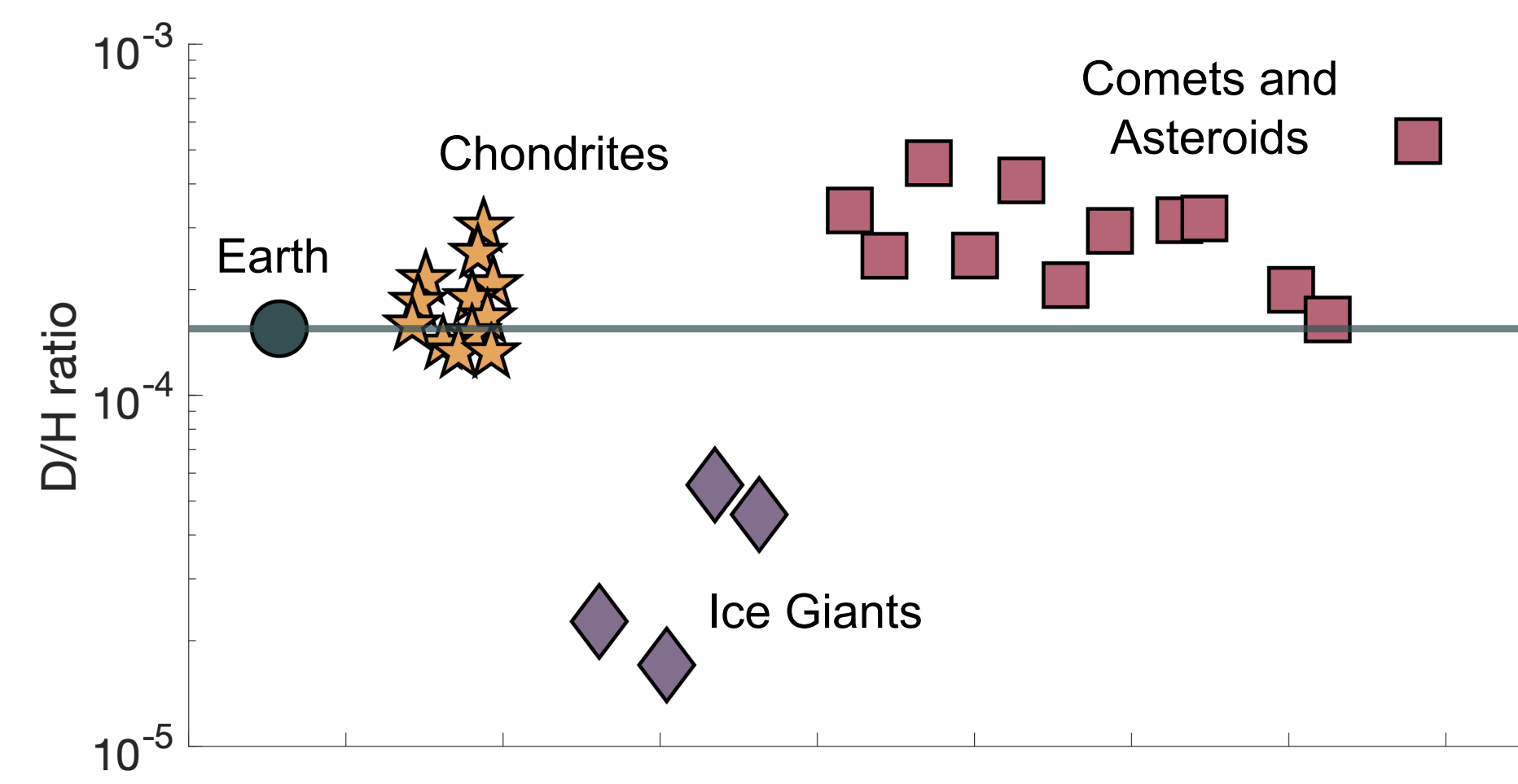


Fig. 1 Sources of water within our solar system

- Our goal is to experimentally constrain the contribution of impactors to Earth's D/H ratio by shocking hydrated SiO<sub>2</sub> glasses to see what happens to volatile species upon impact

## Driving questions

- Examine effects of hydration on physical properties of silicates
  - Does impact velocity change material response?
  - Does OH content change material response?
- What is the degree of devolatilization that occurred during shock?

## Shock quantifies material response properties

- Gas gun housed in the Dynamic Integrated Compression Experimental Facility
- Impact speeds 100, 200, 300 & 400 m/s
- Velocity interferometer system for any reflector (VISAR) collects velocity data from 4 channels (Figs 2 & 3)
- Capable of sample recovery post-impact

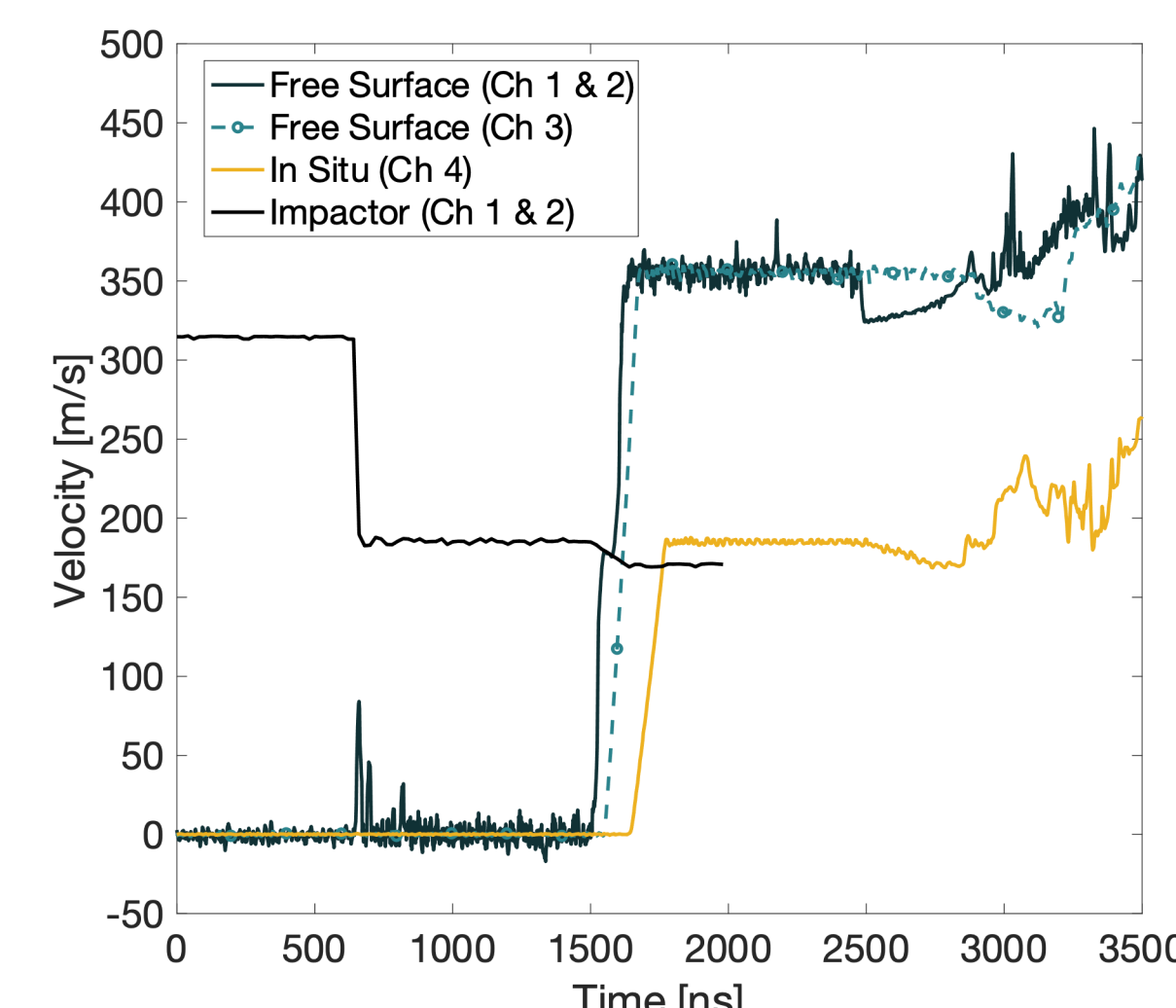
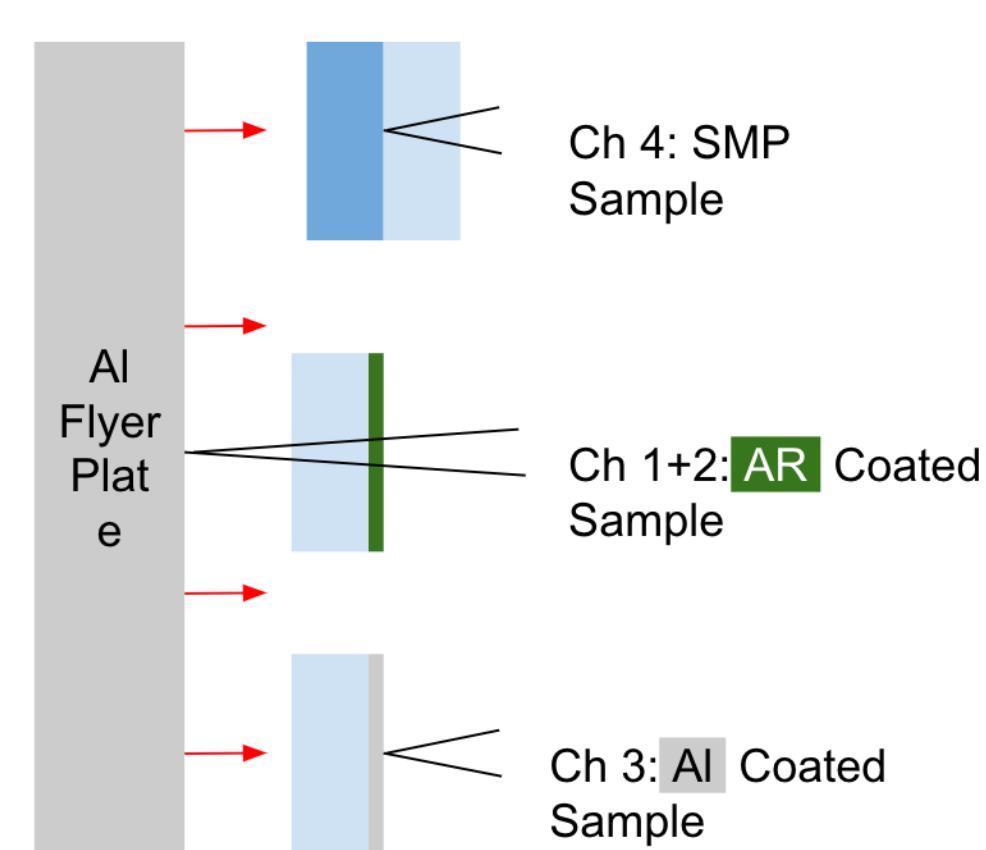


Fig. 2 (Left) Diagram of gas gun experimental setup. Fig 3 (Right) Example of velocities produced from VISAR on each channel

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S.-M. Thomas, et. al. American Mineralogist 93, 10 (2008), E. Stöpler, Contr. Mineral. And Petrol. 81, 1-17 (1982), C.-S. Zha, et. al. Phys. Rev. B 50, 13105 (1994), S.-M. Thomas, et. al Front. Earth Sci. 2 (2015) S.P. Marsh, LASL Shock Hugoniot Data (1980)]

## 1a. At increased $v_{\text{impact}}$ , material stiffness increases

**Lagrangian Sound Speed** quantifies material response during shock

$$c_L(u_p) = \frac{\Delta x}{\Delta t}$$

$$\delta c_L(u_p) = \sqrt{\left(\frac{\sigma x}{\Delta x}\right)^2 + \left(\frac{\sigma t_0}{z - t_0}\right)^2 + (\gamma)^2}$$

- Higher impact velocities, independent of OH presence, have stiffer response (higher  $c_L$ ) than lower impact velocities**
- At high impact velocities, hydrated samples have a stiffer response than dry samples
- Our results (Fig 5) are consistent with static compression experiments (Zha et. al., 1994)
- Error in 400 m/s sample likely due to lack of data points converting VISAR intensity to velocity

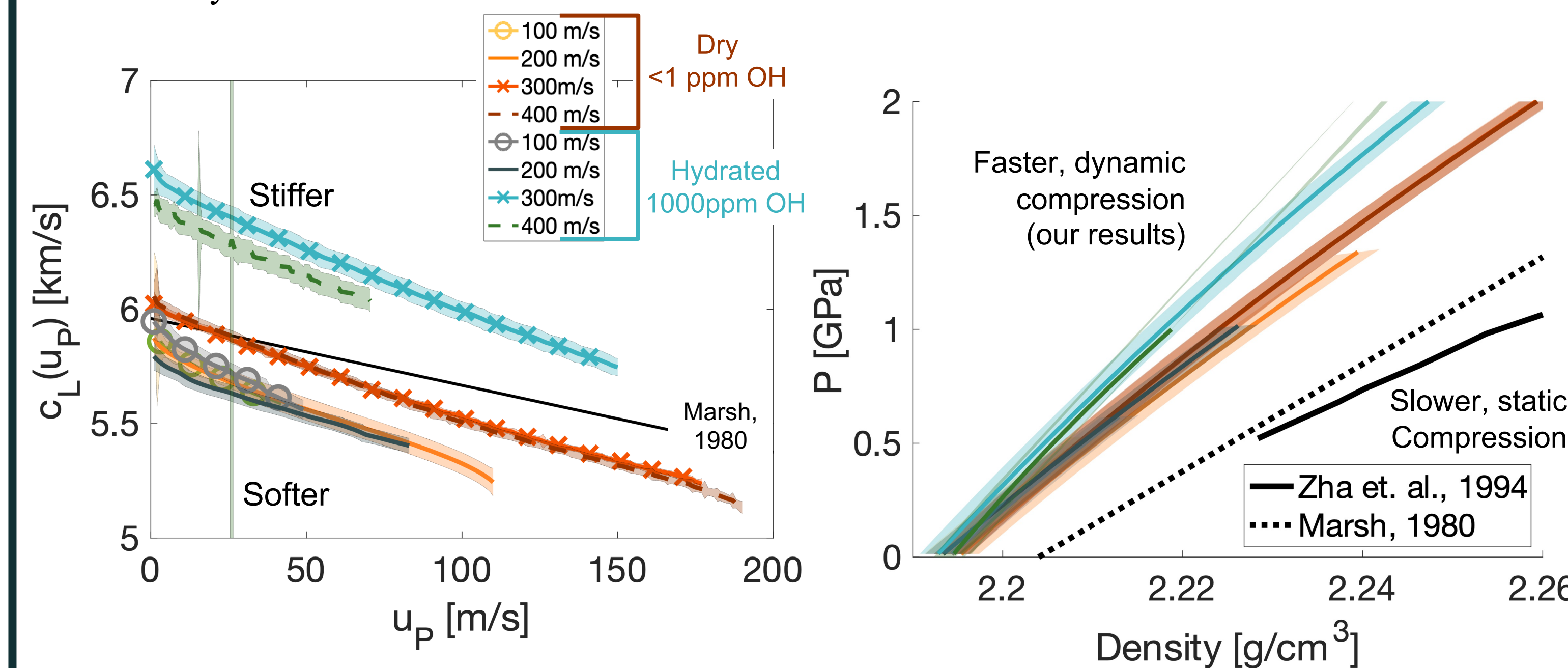


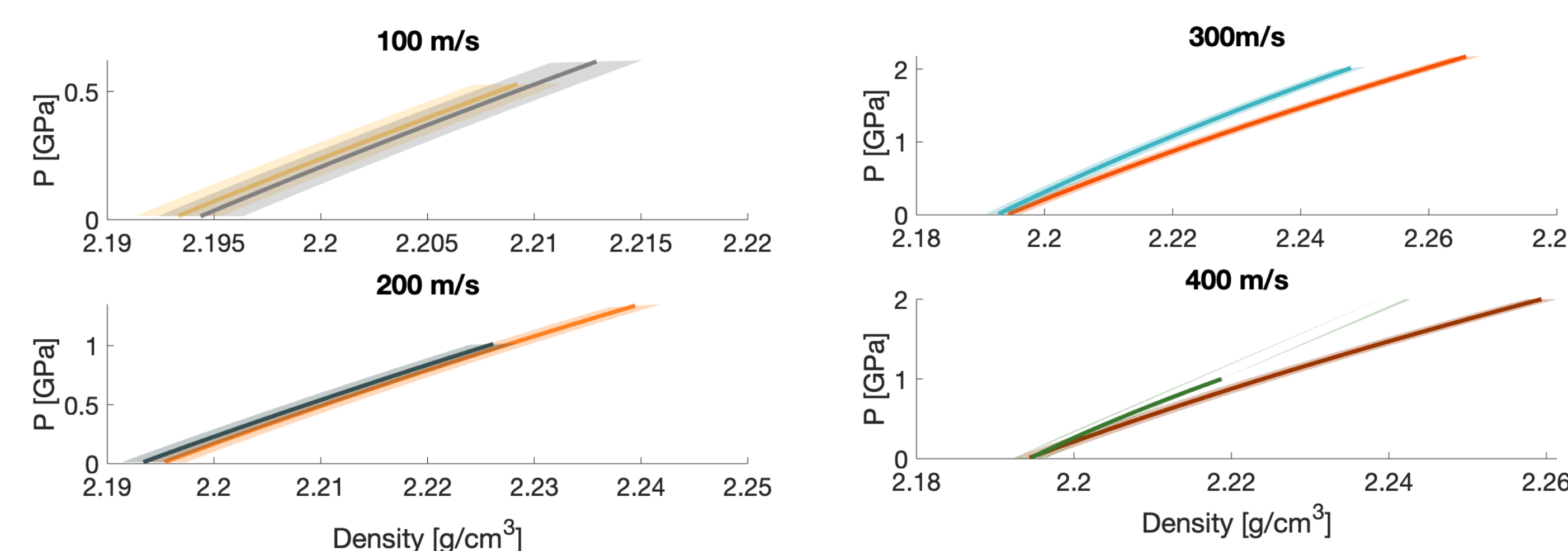
Fig. 4 (left) Lagrangian sound speeds of dry and hydrated glassed over increasing impact velocities. Fig 5. (right) Pressure density curve with both static (black line) and dynamic (colors) experimental results

## 1b. At increased $v_{\text{impact}}$ , OH affects material response

**Simplified form of Conservation Equations**

$$P = \rho_0 \int_0^u c_L du \quad \sigma P = \rho_0 \int_0^u \sigma c_L du \quad \rho = \rho_0 \left(1 - \int_0^u c_L^{-1} du\right)$$

- Compressional behavior is calculated from the sound velocity,  $c_L(u_p)$
- Dry and hydrated samples behave the **same at low** velocities (Fig 6)
- Dry and hydrated samples behave **differently at high** velocities (Fig 7)



Compositional variation pressure-density curve Fig 6 (left) low impact velocities. Fig 7 (right) high impact velocities.

## Spectroscopy reveals post-impact OH content

- Bruker Tensor37 coupled to Bruker Hyperion microscope 15x objective and 10x eyepieces
- Transmission spectra recorded from 6000 to 4000 cm<sup>-1</sup> at ambient conditions (Fig 8)
- Double polished SiO<sub>2</sub> samples recovered from shock experiments
- Aperture size set to smallest sample and used for all experiments (image)

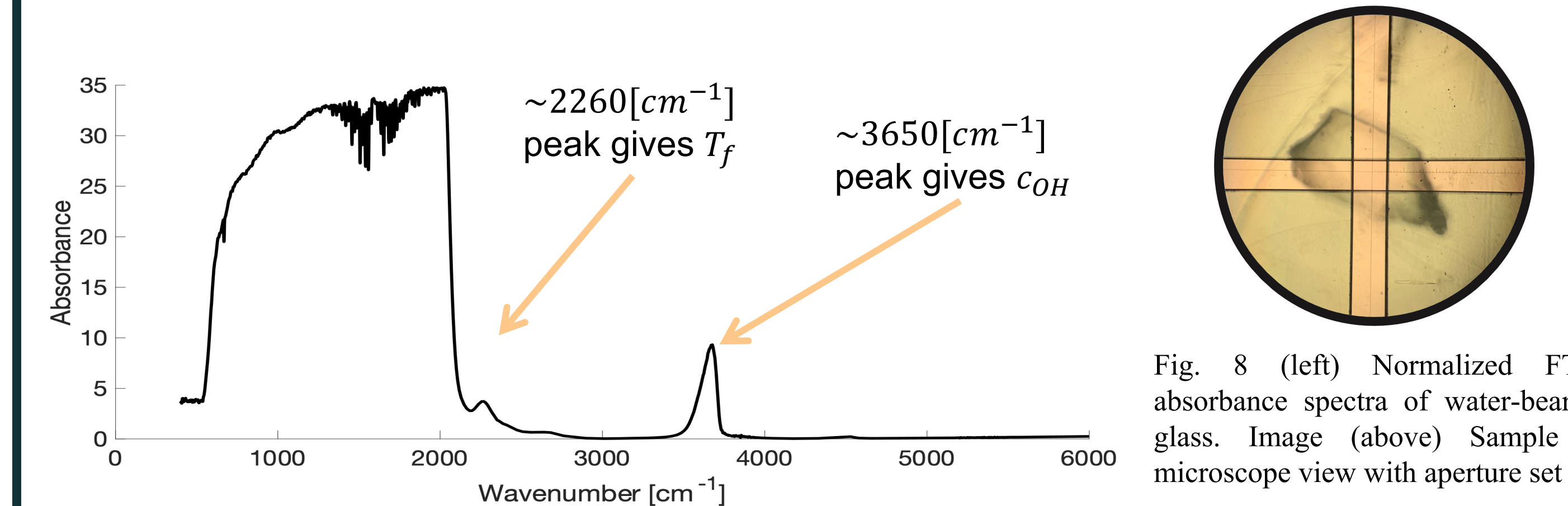


Fig. 8 (left) Normalized FTIR absorbance spectra of water-bearing glass. Image (above) Sample in microscope view with aperture set

## 2. OH varies in post-impact samples

**Beer-Lambert Law** relates water concentration to OH peak height (Stöpler, 1982)

$$\epsilon = \frac{1.8 \cdot A_i}{t \cdot \rho \cdot c_{H_2O}}$$

- Hydrated samples' post-impact  $\rho$  determined from fictive temperature

$$T_f \text{ (Fig. 9)} \rho \propto \frac{10^{-6}}{T_f} \propto \frac{10^4}{\lambda_{2260}}$$

- Peak height of the OH band varies between samples of the same initial composition
- Samples impacted at 300 m/s show little to no OH peak

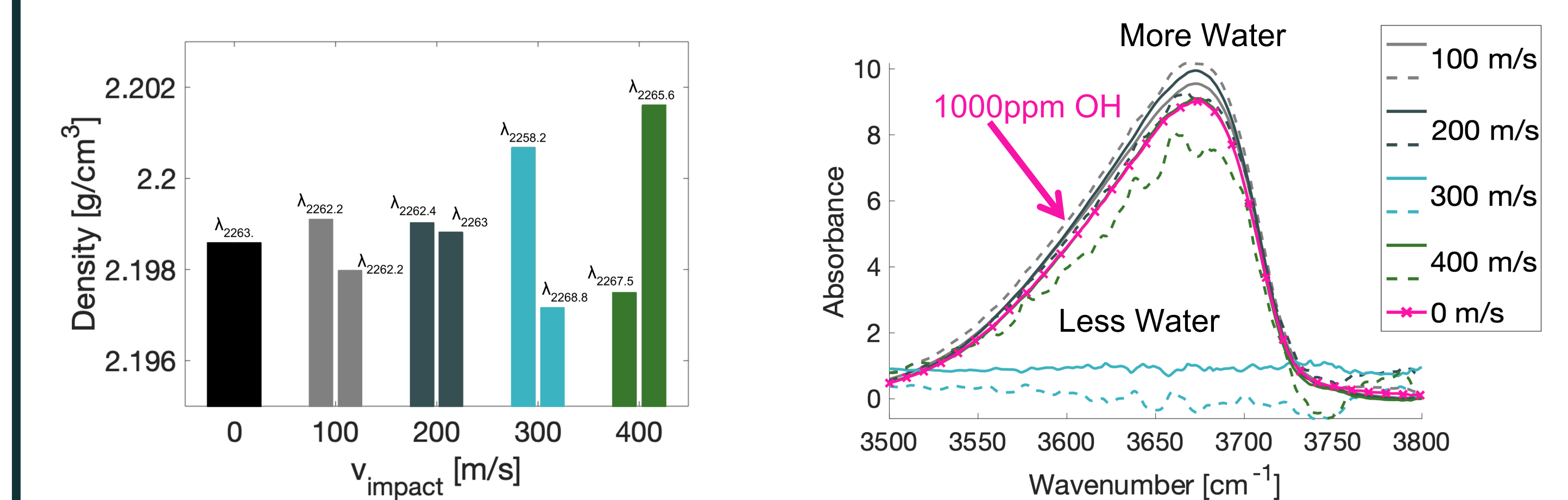


Fig 9 (left) Post-impact hydrated sample densities with  $\lambda_{2260}$  peak location labeled. Fig. 10 (right) OH band at ~3650 [cm<sup>-1</sup>] background-subtracted, normalized FTIR absorbance spectra of initially water-bearing glasses

## SiO<sub>2</sub> glass preserves OH during impacts!

- Does impact velocity change mat response?
    - Independent of composition, SiO<sub>2</sub> glass has a stiffer response at high  $v_{\text{impact}}$  (Fig 4)
  - Does OH content change mat resp?
    - Compressional behavior of samples diverges at higher  $v_{\text{impact}}$  between dry and hydrated samples (Figs 7 and 8)
- What is the degree of devolatilization that occurred during shock?
    - Initial results suggest OH survives impacts within SiO<sub>2</sub> glasses (Fig 10)
    - Peak height of OH band varies between samples of the same initial composition
    - More work is needed to quantify  $c_{OH}$
    - Vary aperture size of FTIR to see if degree of devolatilization varies across the sample geometry
  - Future plans to increase sample complexity to chemicals more representative of solar system compositions (basalt)