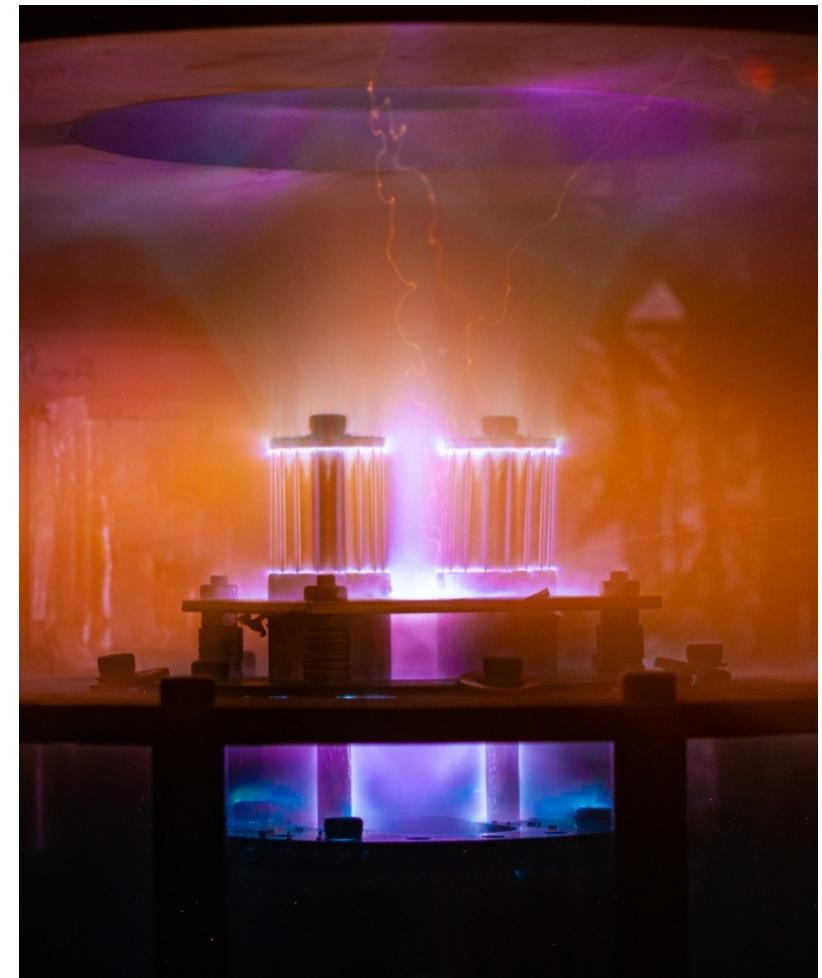


The New PUFFIN Generator at MIT

Jack Hare, jdhare@mit.edu



PSFC Plasma Science and Fusion Center

Sandia National Laboratories is a multimission laboratory managed and operated by National Technology & Engineering Solutions of Sandia, LLC, a wholly owned subsidiary of Honeywell International Inc., for the U.S. Department of Energy's National Nuclear Security Administration under contract DE-NA0003525.



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Sandia
National
Laboratories

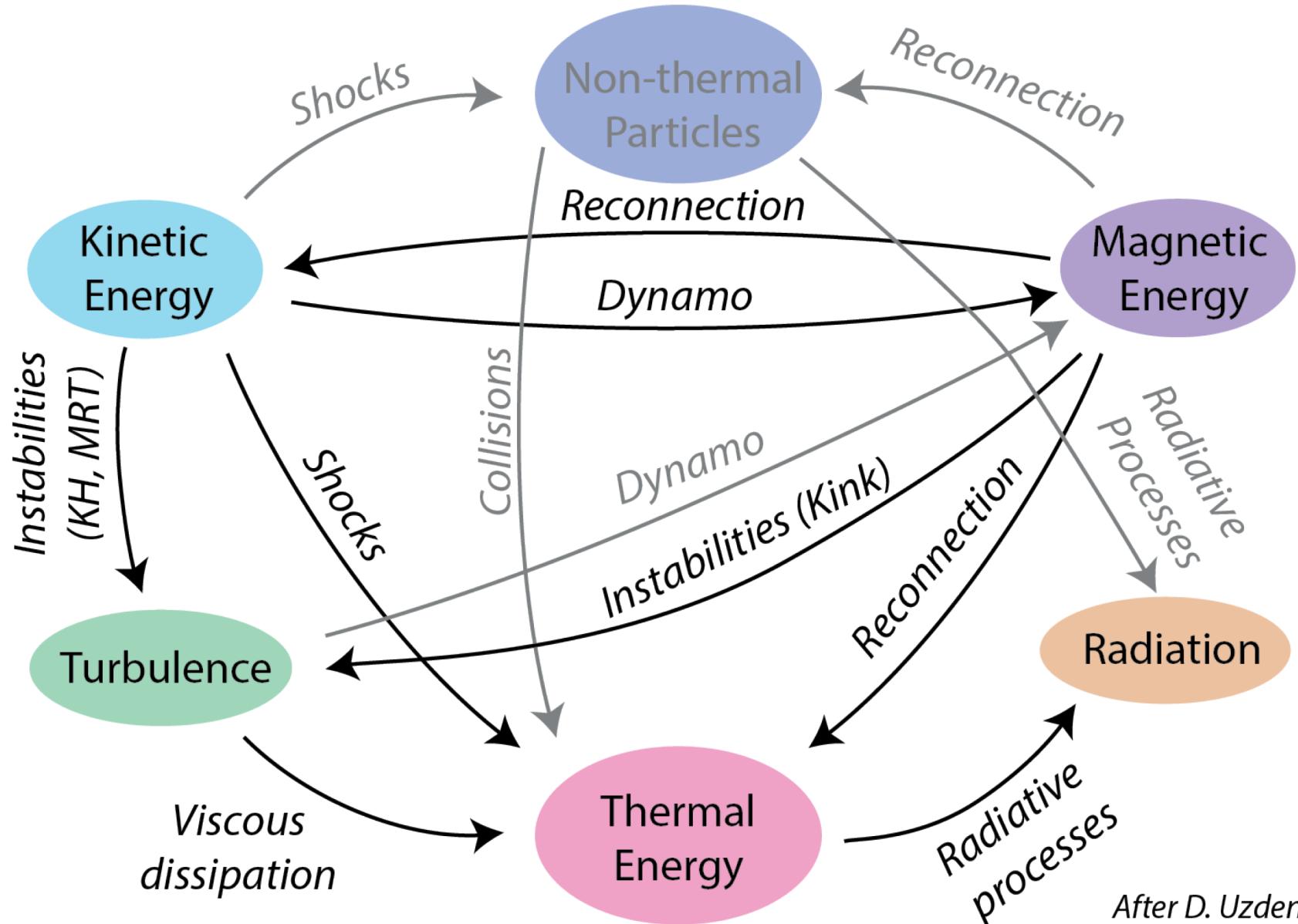


This work is supported
by the NSF and NNSA



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Research Interests: Energy Flows in Plasma



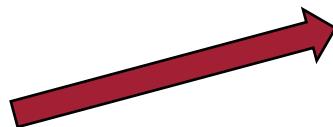


Research paths/talk outline

Magnetic
reconnection

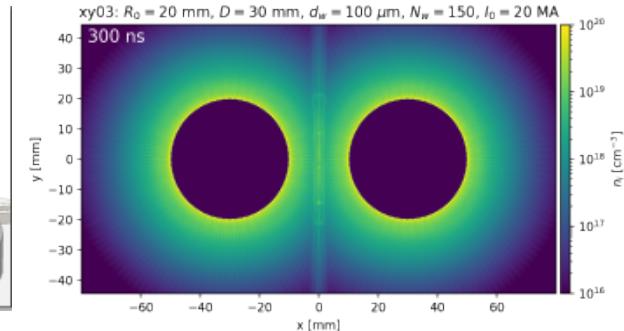
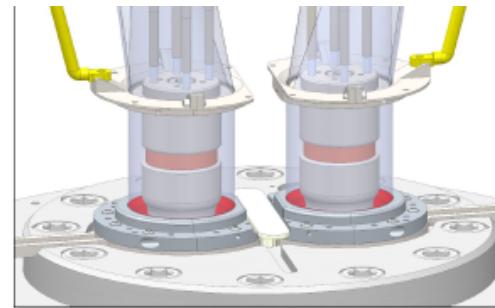


More current

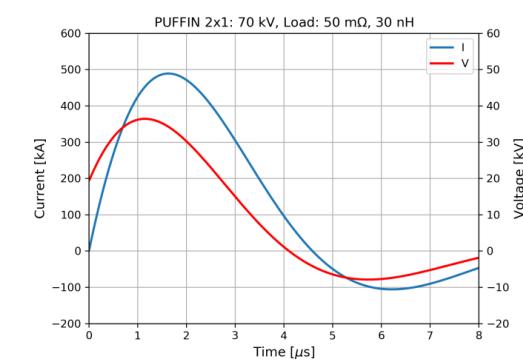
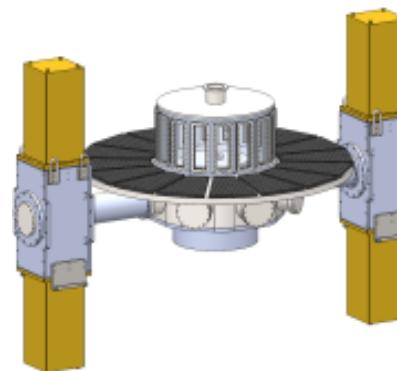


Longer current
pulse

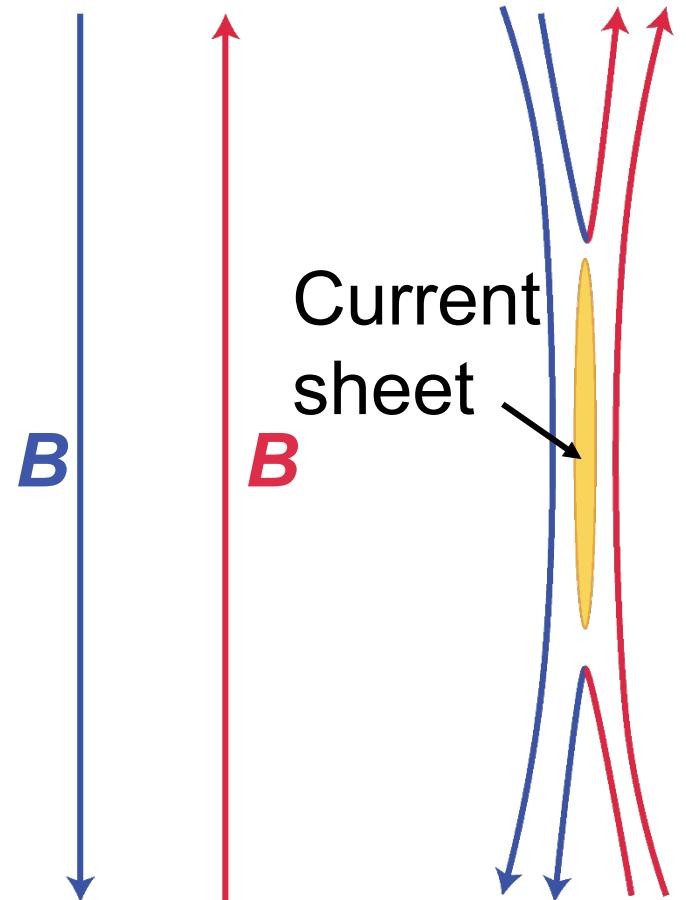
Radiatively cooled reconnection on Z



PUFFIN: 2 μ s, 500 kA at MIT



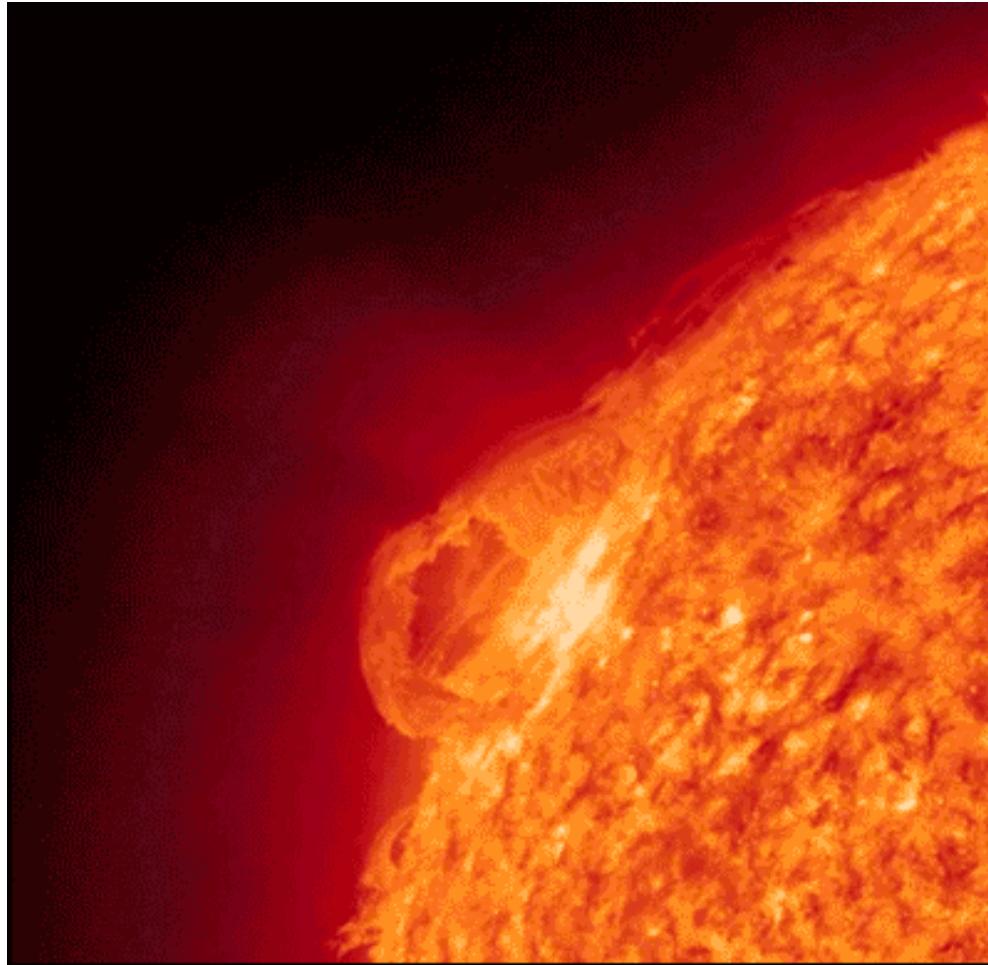
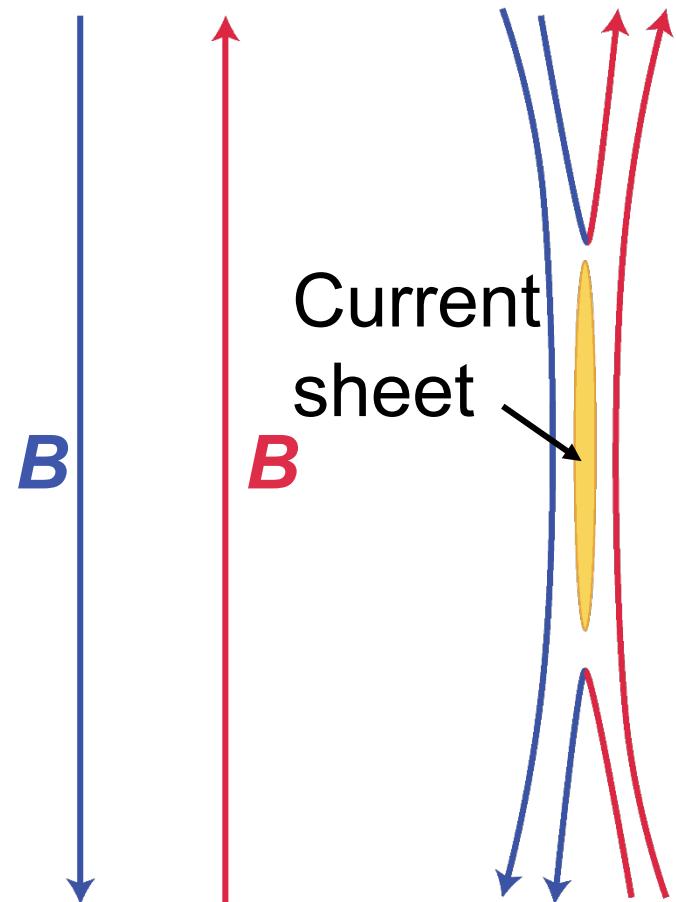
Magnetic Reconnection



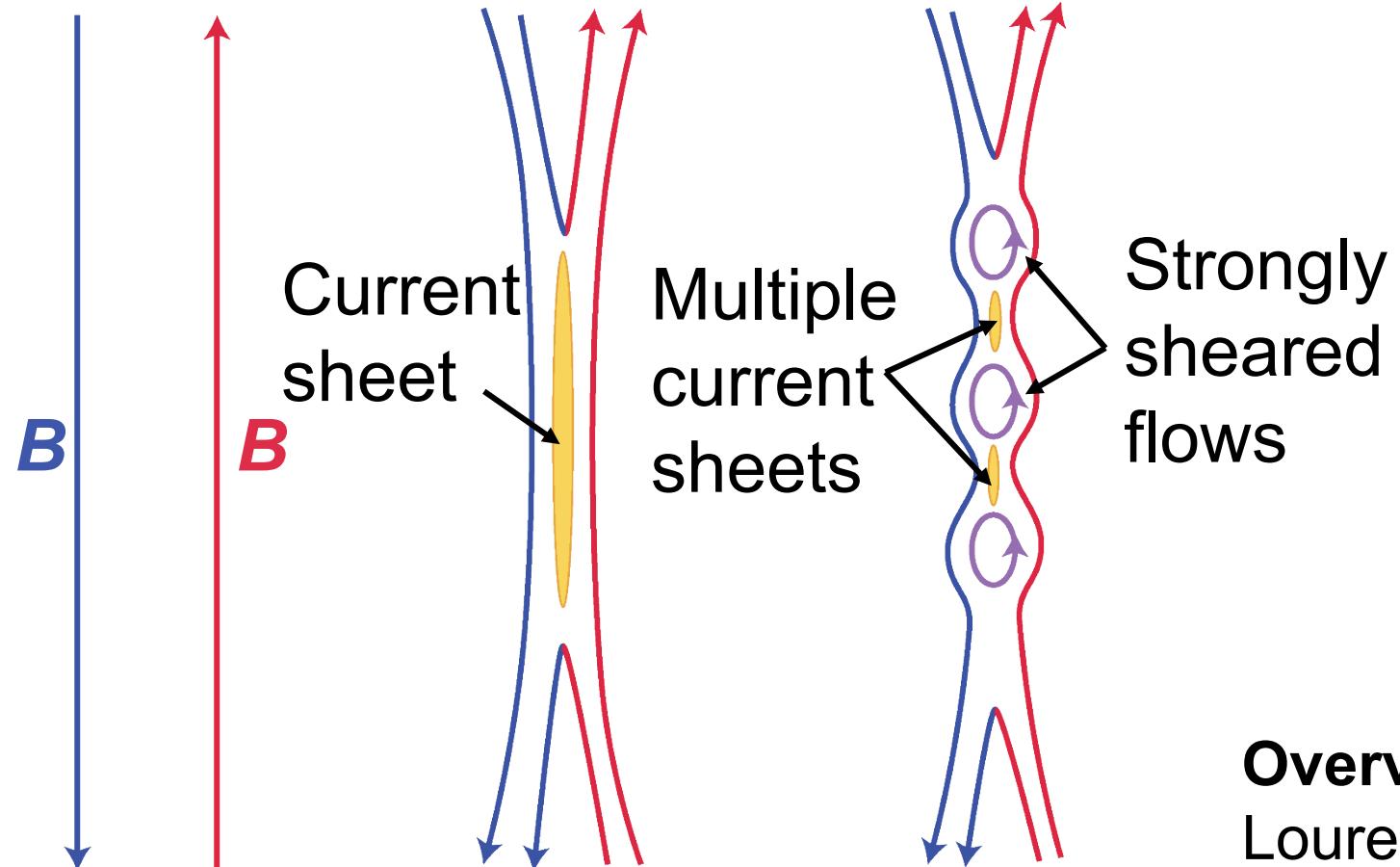


Magnetic Reconnection

Prediction: 1000 yrs. Reality: 10 minutes!

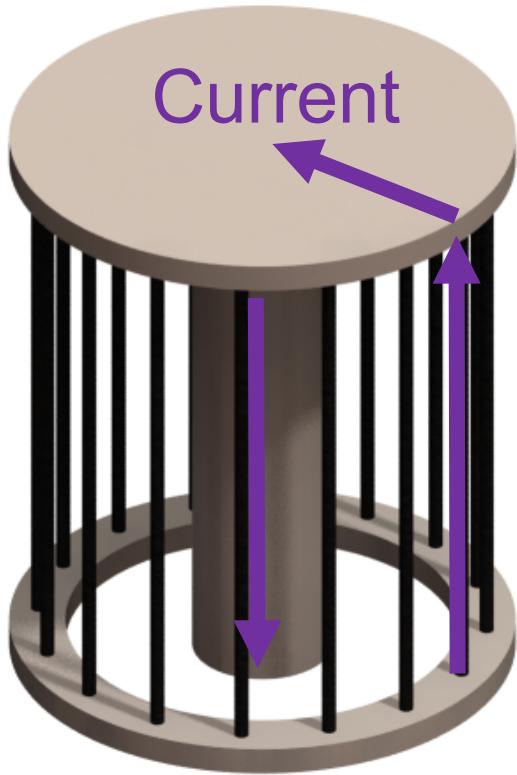


Plasmoids Lead to Fast Reconnection and Anomalous Heating

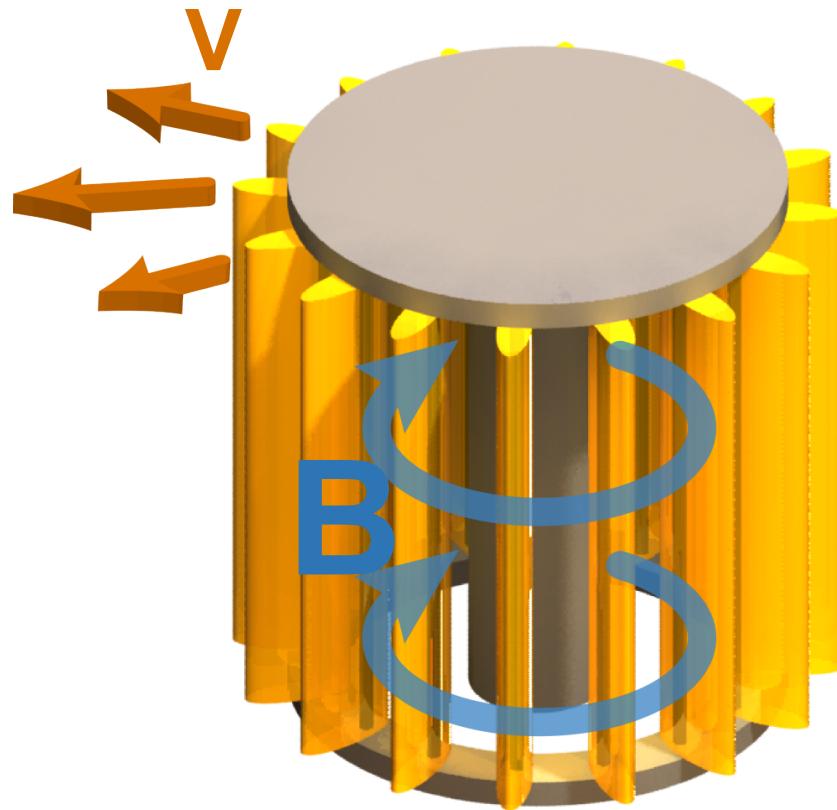


Overview of recent theory:
Loureiro, N. F., & Uzdensky, D. A. (2015).
PPCF, 58, 014021

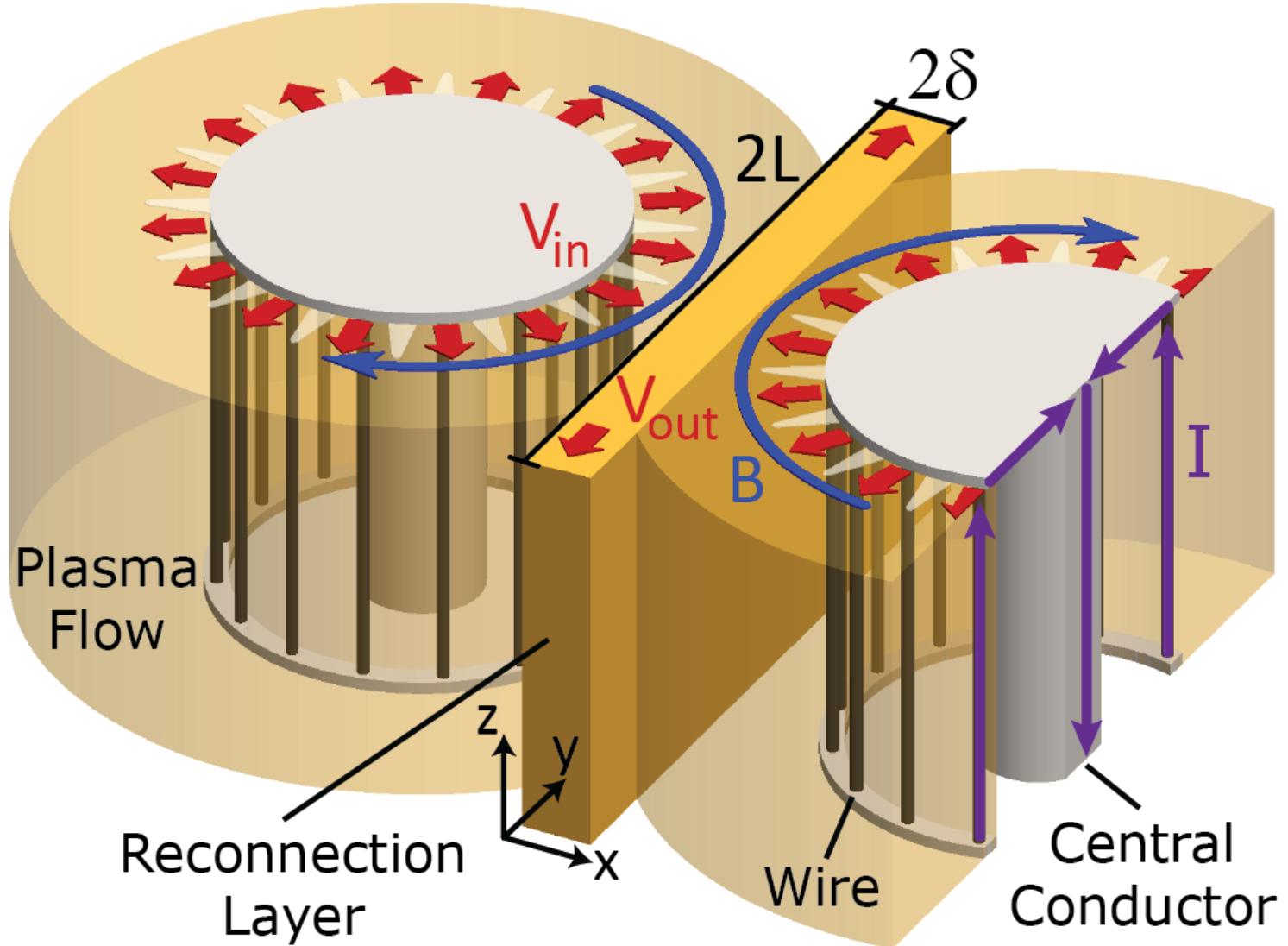
Pulsed-power-driven Magnetic Reconnection



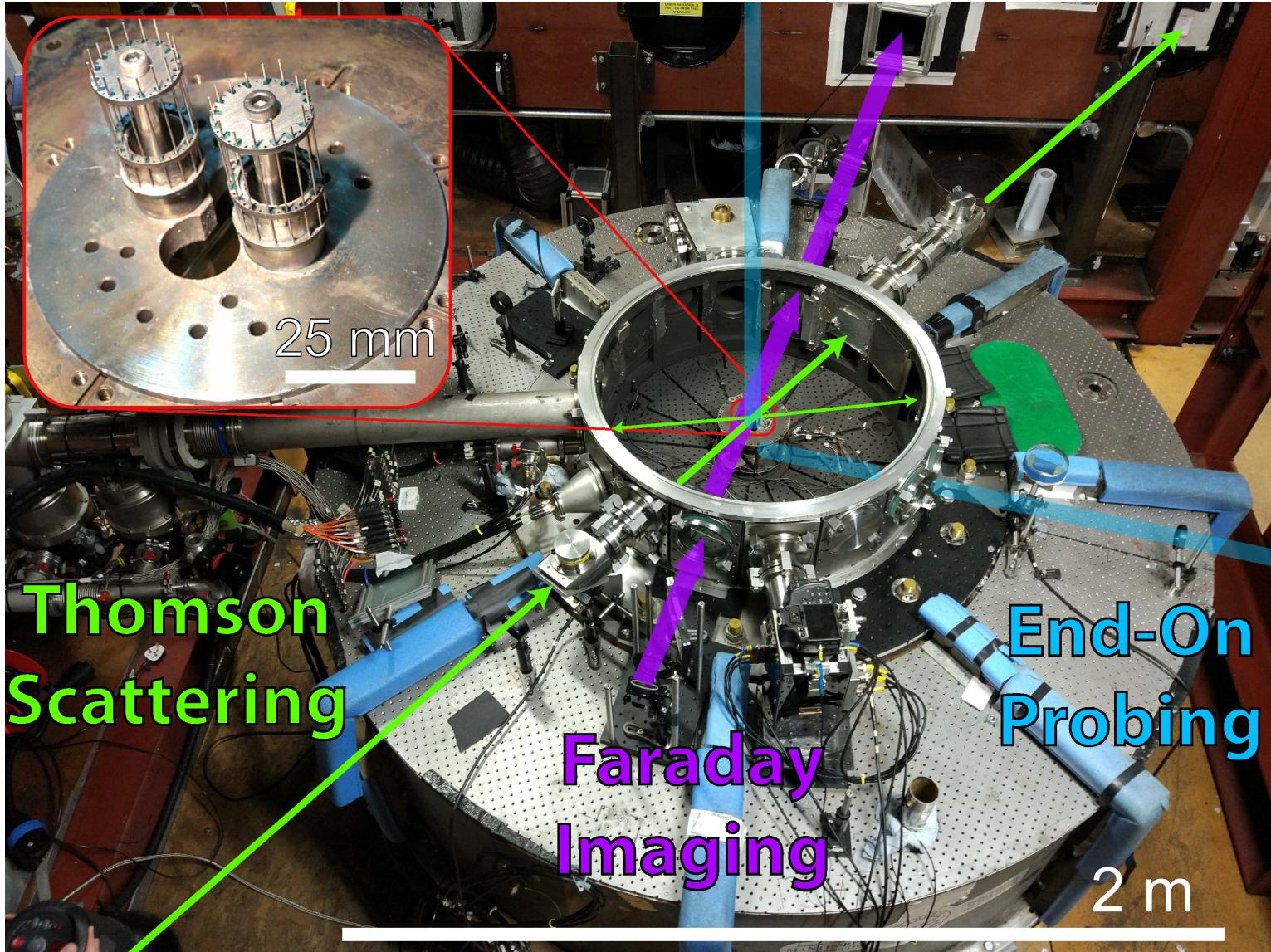
$I=1.4$ MA, 240 ns rise time



Magnetic Reconnection from Double Exploding Wire Arrays

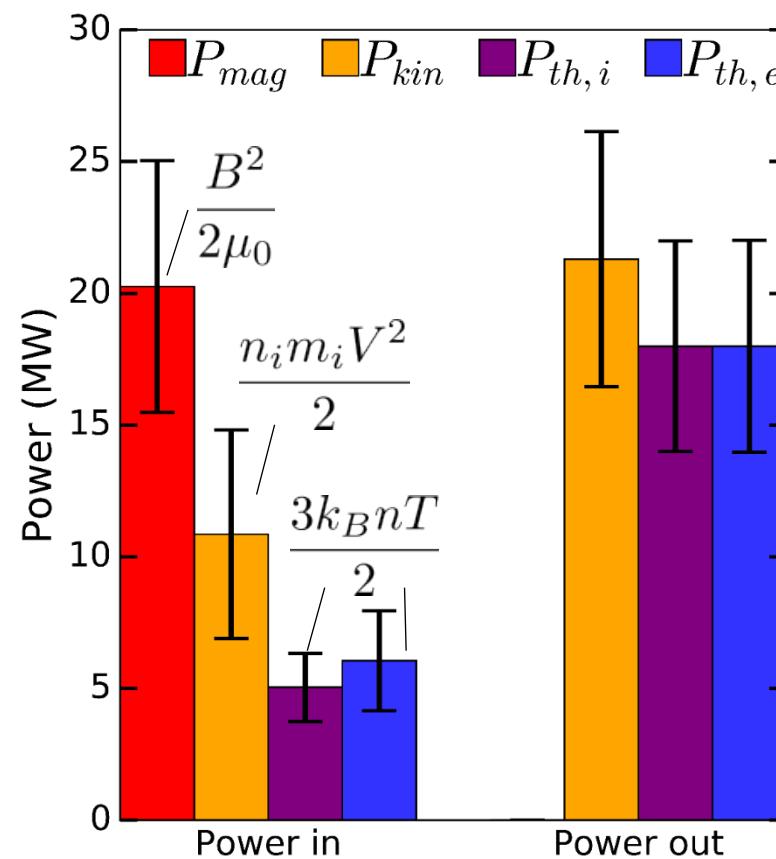
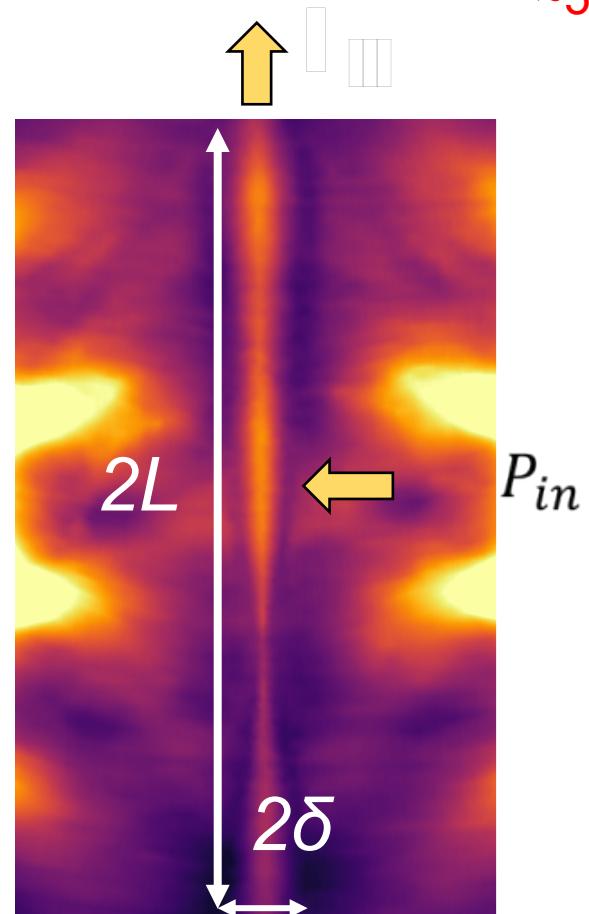


Overview of Diagnostic Suite on MAGPIE at Imperial College





Anomalous Heating in the Reconnection Layer



**Classical heating
is too slow:**

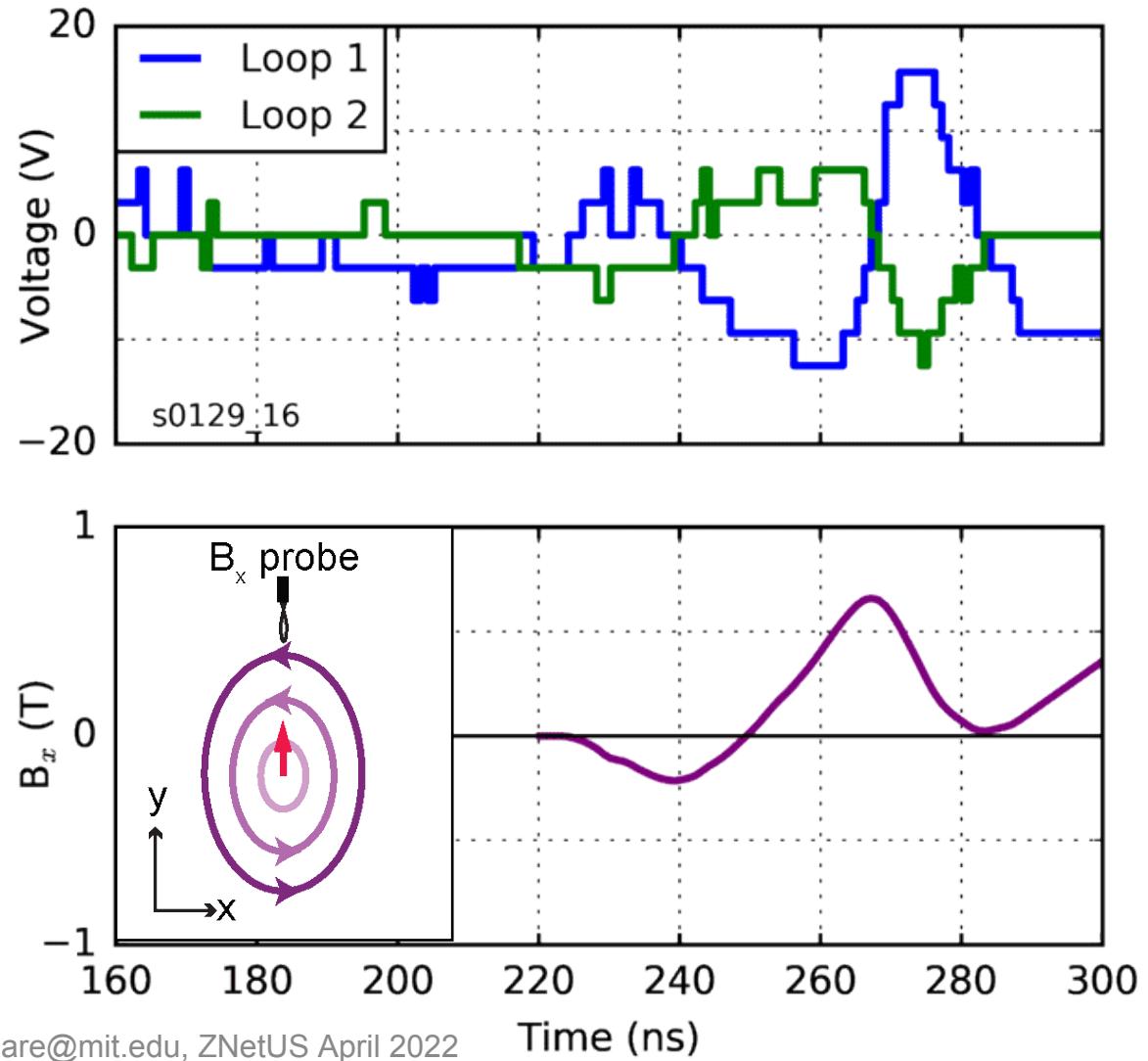
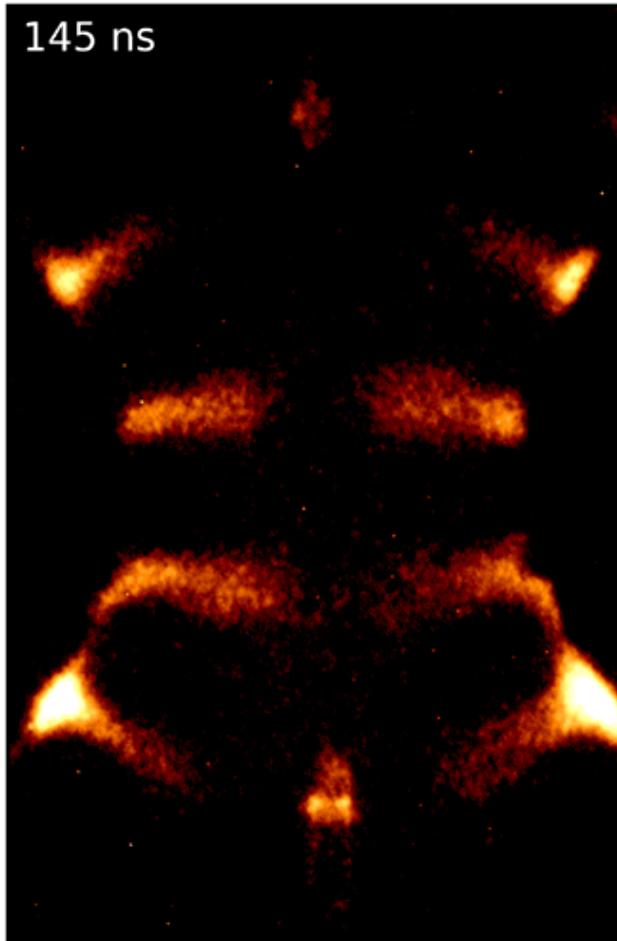
 $\approx 50\text{ ns}$

 $\approx 800\text{ ns}$

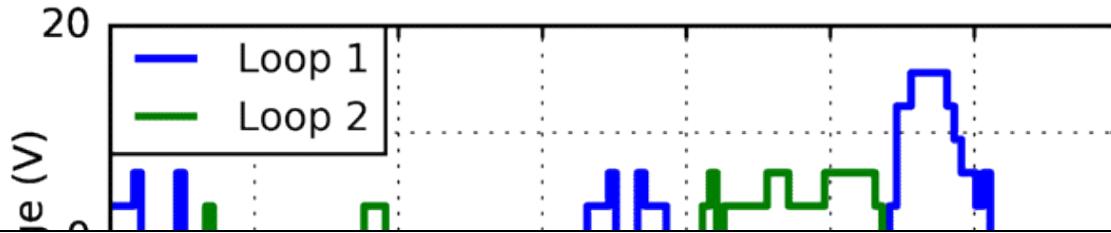
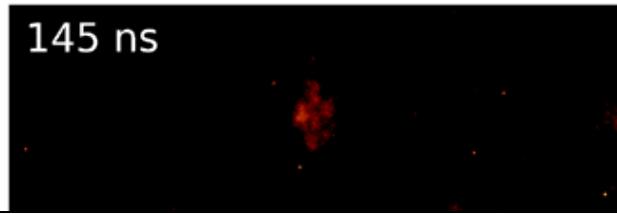
 $\approx 350\text{ ns}$

$\tau_{exp} \ll \tau_{visc}, \tau_{res}$

Magnetic Structure of Plasmoids

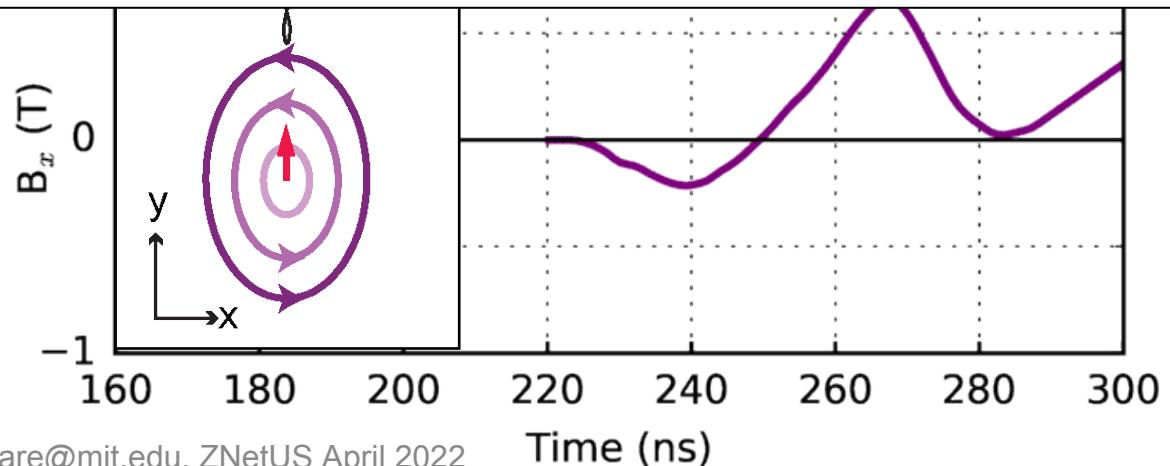
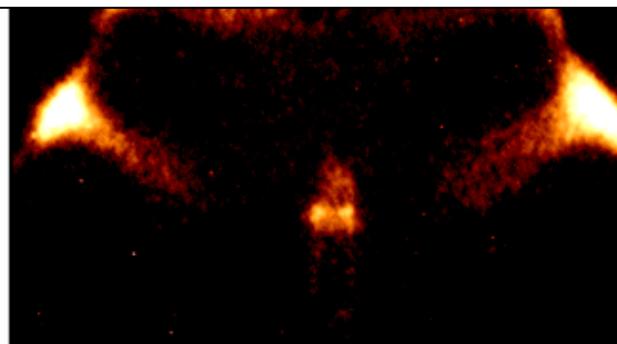


Magnetic Structure of Plasmoids



Open questions:

- How does reconnection layer dynamics scale with current?
- What is the quasi-steady behavior of reconnection?



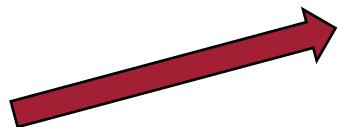


Research paths/talk outline

Magnetic
reconnection

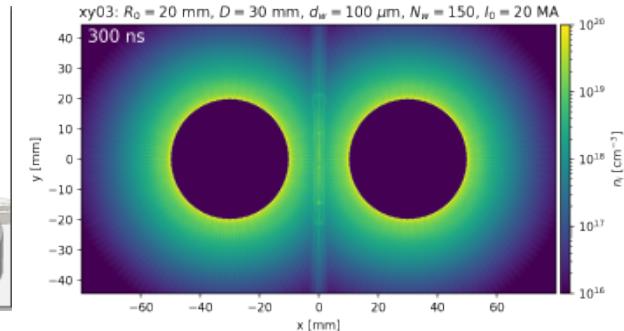
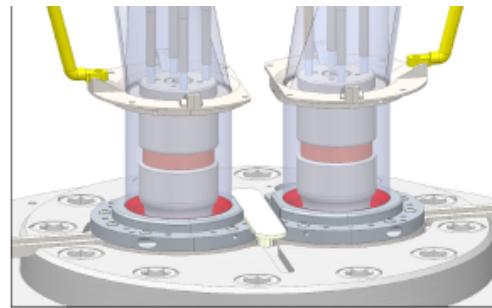


More current

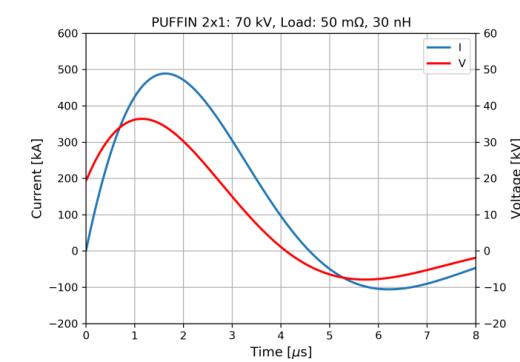
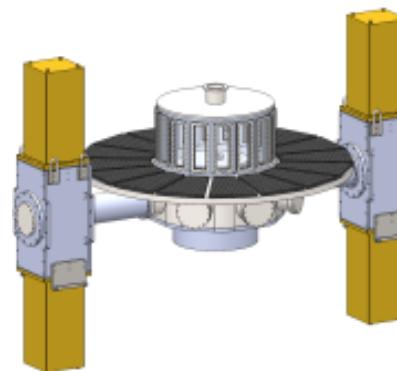


Longer current
pulse

Radiatively cooled reconnection on Z



PUFFIN: 2 μ s, 500 kA at MIT



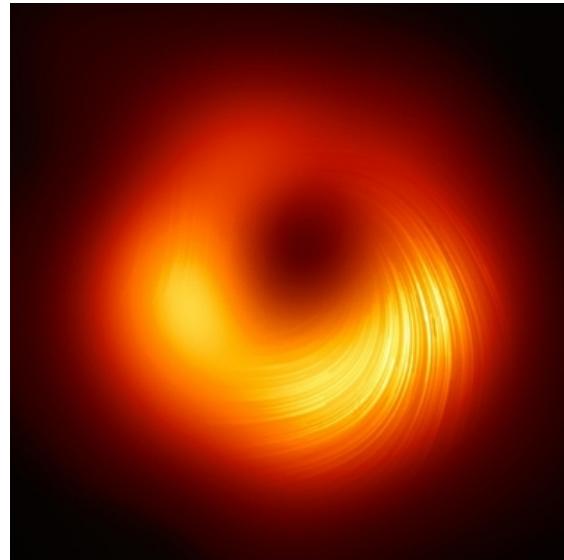
Reconnection in Extreme Astrophysical Environments



Artist's impression of a black hole



M87 (EHT)



Crab Pulsar (Hubble/Chandra)



1. Cooling is a significant loss mechanism:
 - Modifies partition of magnetic energy between electrons, ions, kinetic
 - Leads to cooling instabilities, radiative collapse, termination of reconnection
2. X-rays: key observational signature in remote environments:
 - Where and when are X-rays produced – localized bursts?
 - How does this couple back to the reconnection process?

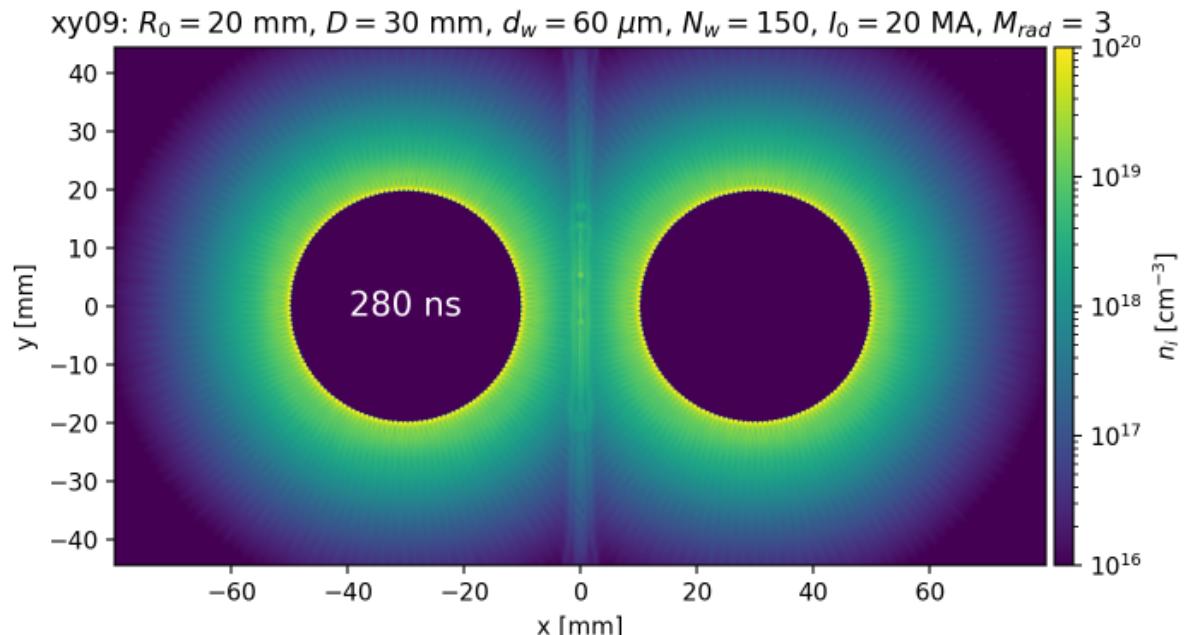
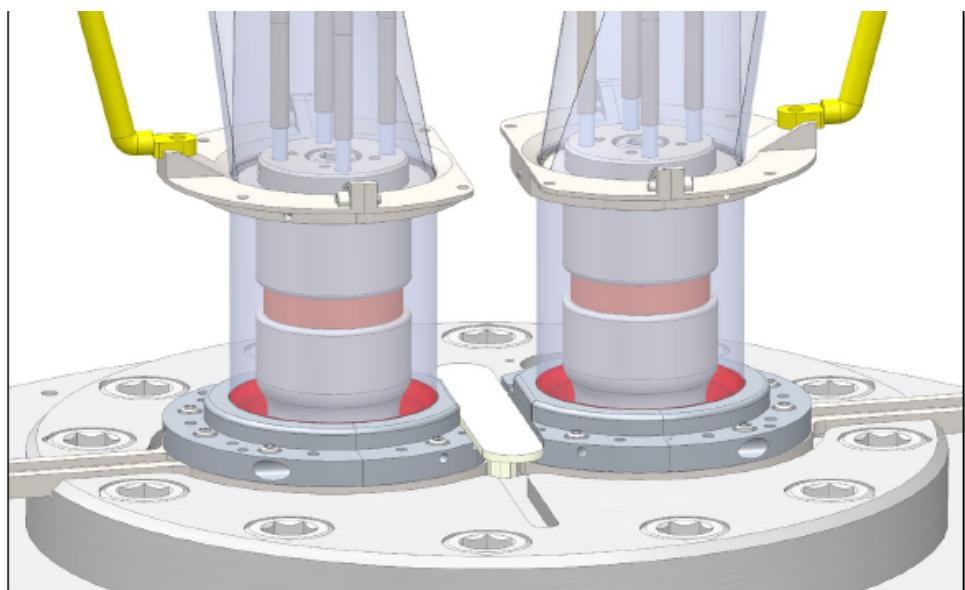


MARZ: Reconnection on Z

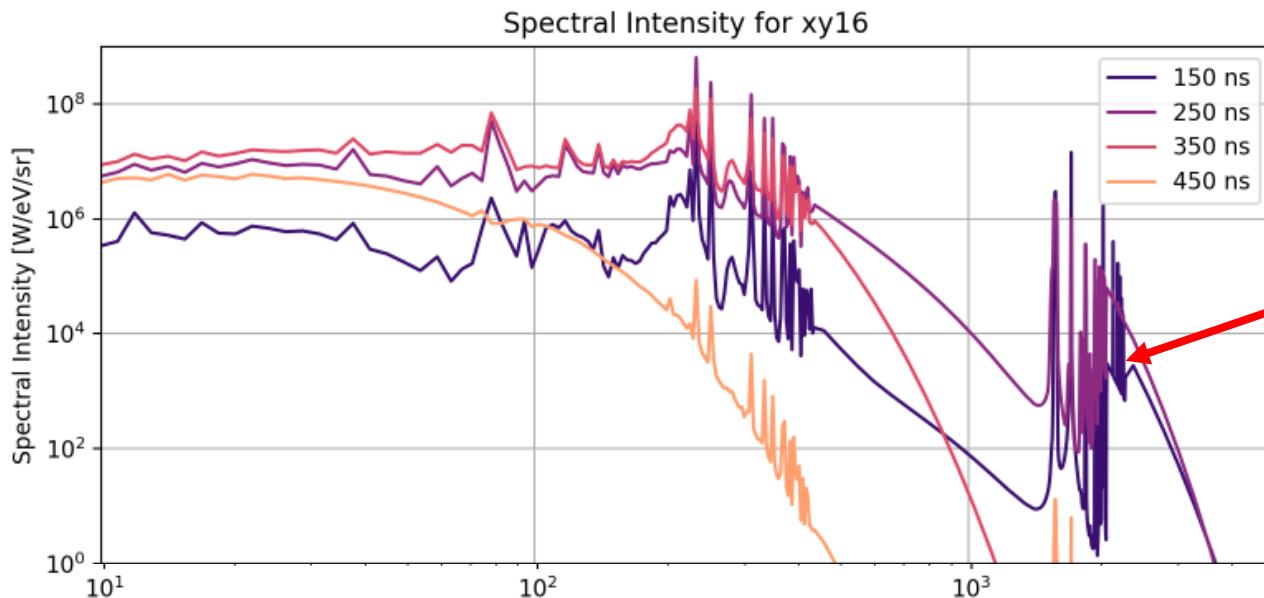
Z is the largest pulsed-power machine in the world

- 20-30 MA peak current compared to 1.4 MA on MAGPIE:
 - **Density** increase by $I^2 \sim 400$
 - **Magnetic energy** increase by $I^2 \sim 400$
 - **Cooling rate** increase by $I^4 \sim 160,000$

Unique capability: strongly radiatively cooled reconnection

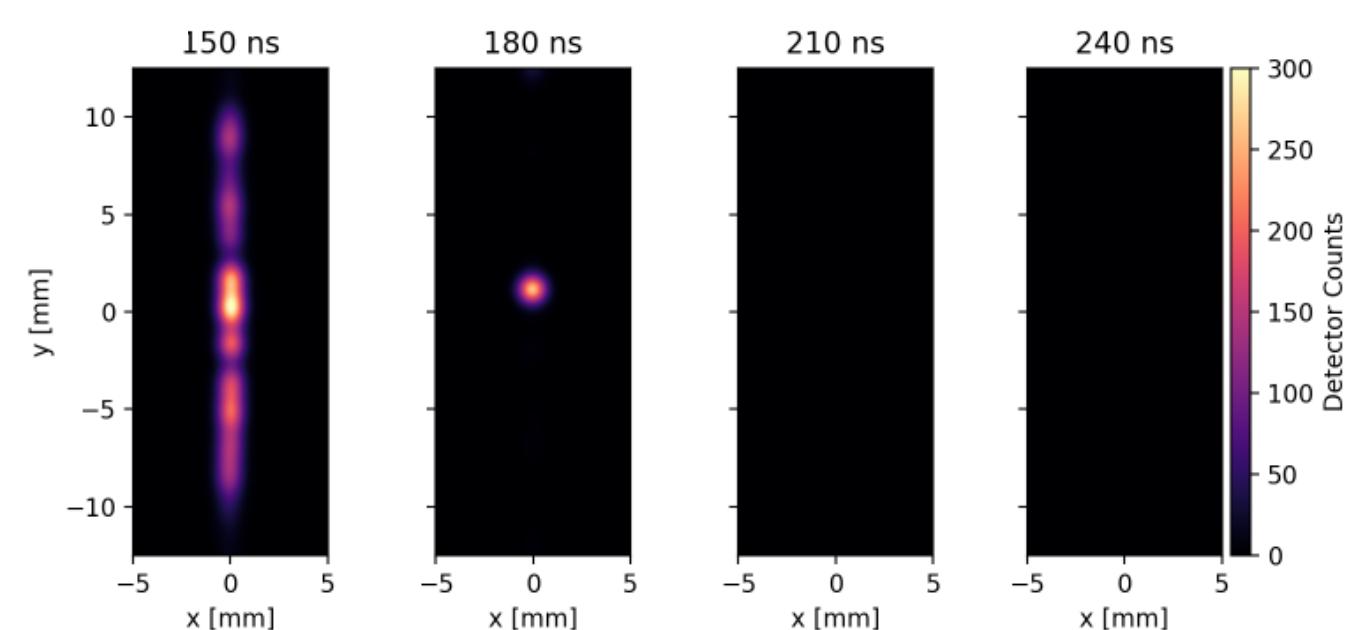
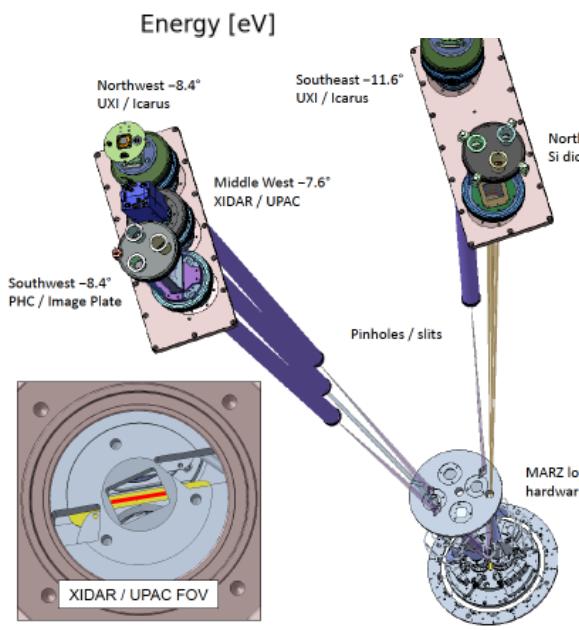


XP2: X-ray Post-Processor by Aidan Crilly & Jerry Chittenden



Al K-shell
disappears
after
collapse

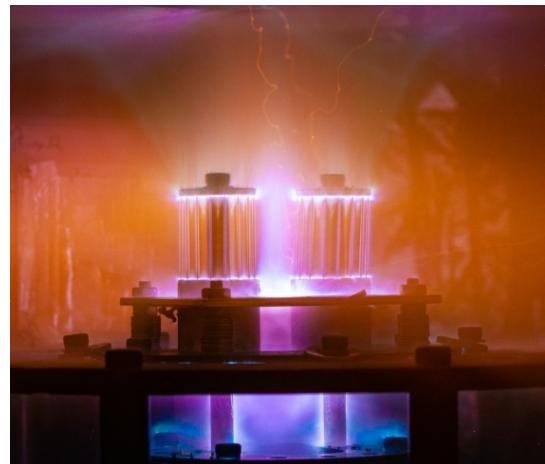
XP2 provides a
predictive
capability for X-
ray diagnostics



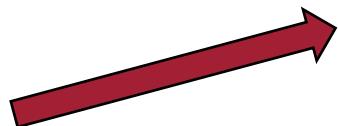


Research paths/talk outline

Magnetic
reconnection

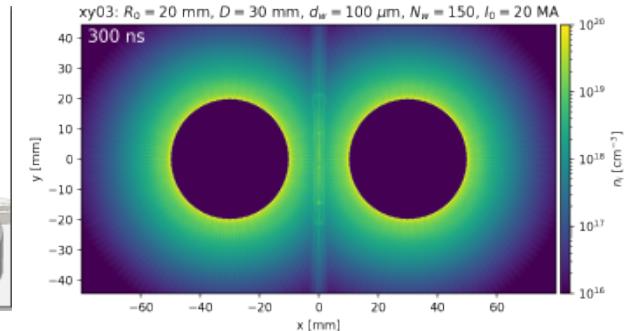
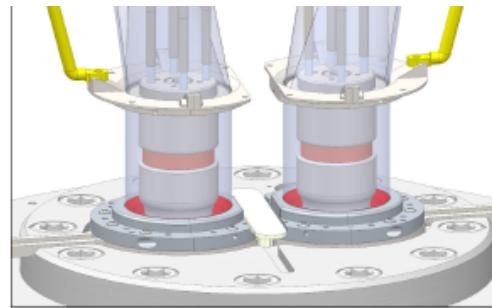


More current

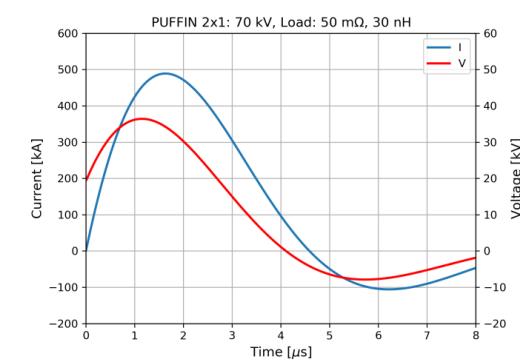
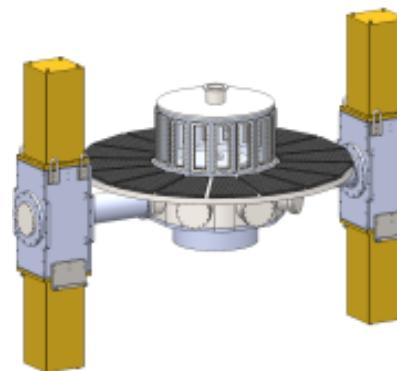


Longer current
pulse

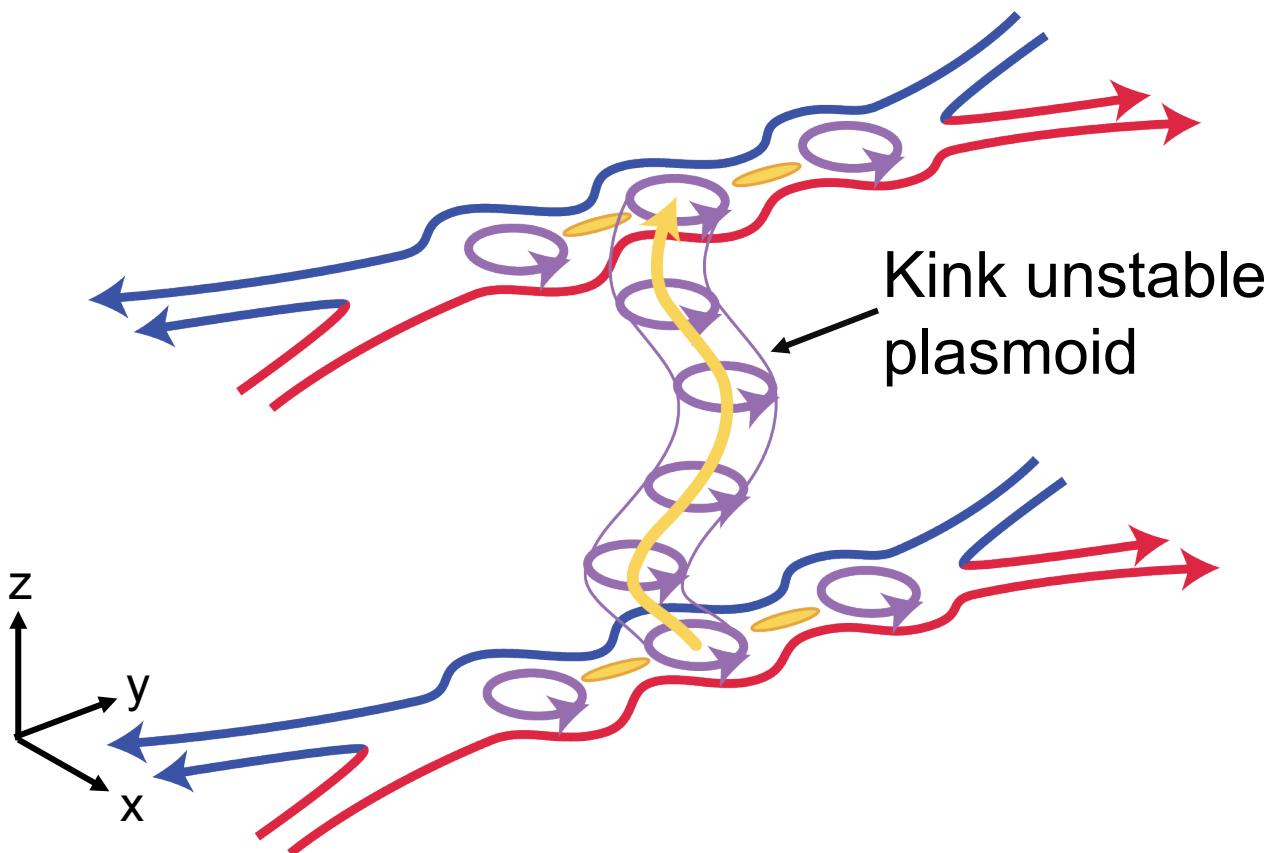
Radiatively cooled reconnection on Z



PUFFIN: 2 μ s, 500 kA at MIT



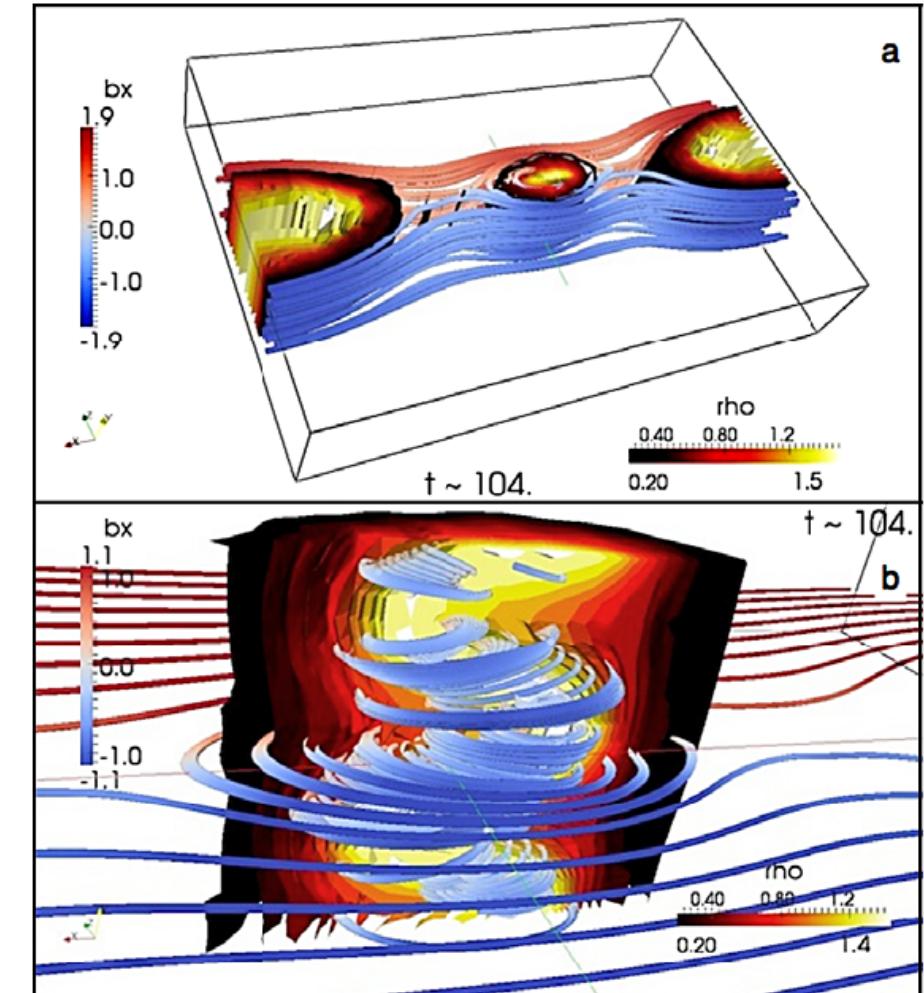
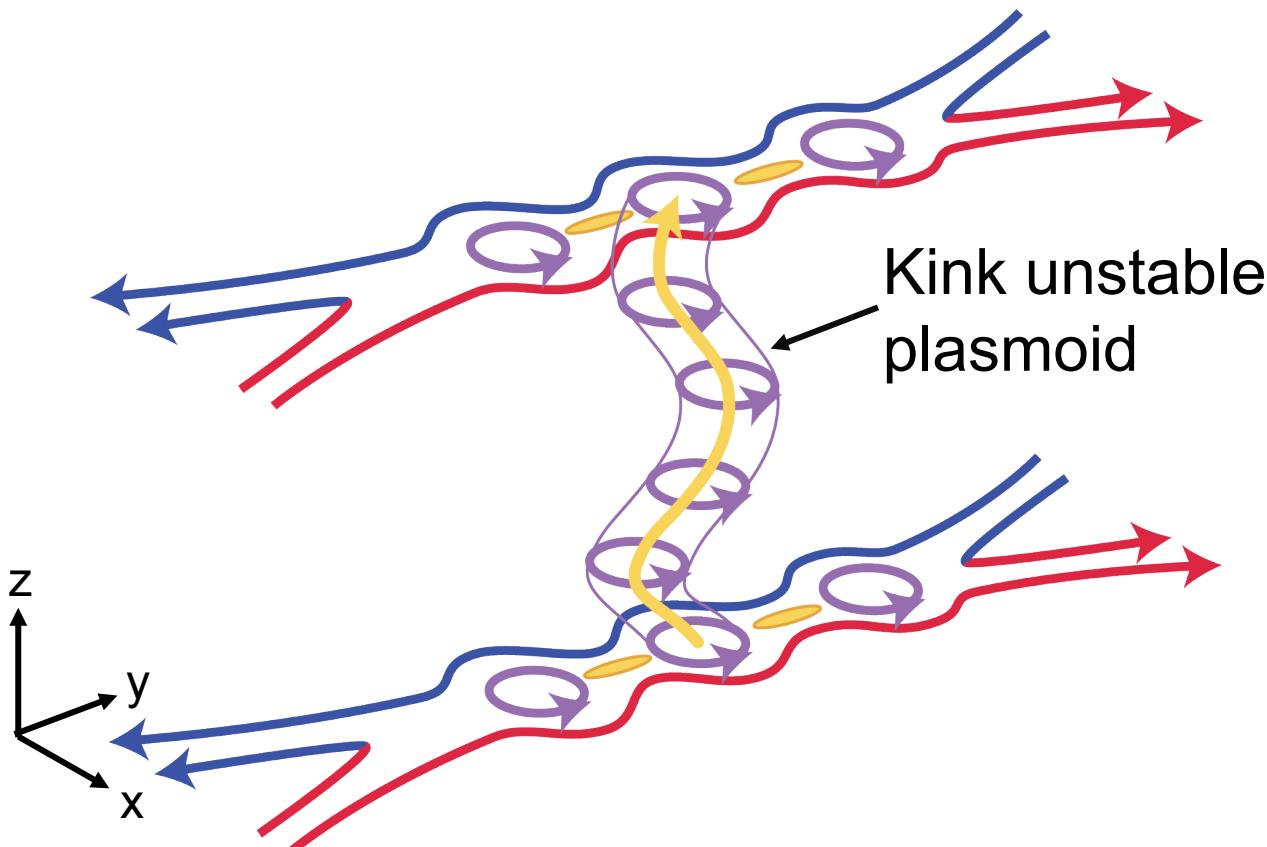
Interesting physics happens on long time scales



Kink Instability Leads to Turbulent Reconnection



Turbulent reconnection over a large volume,
as observed in astrophysics



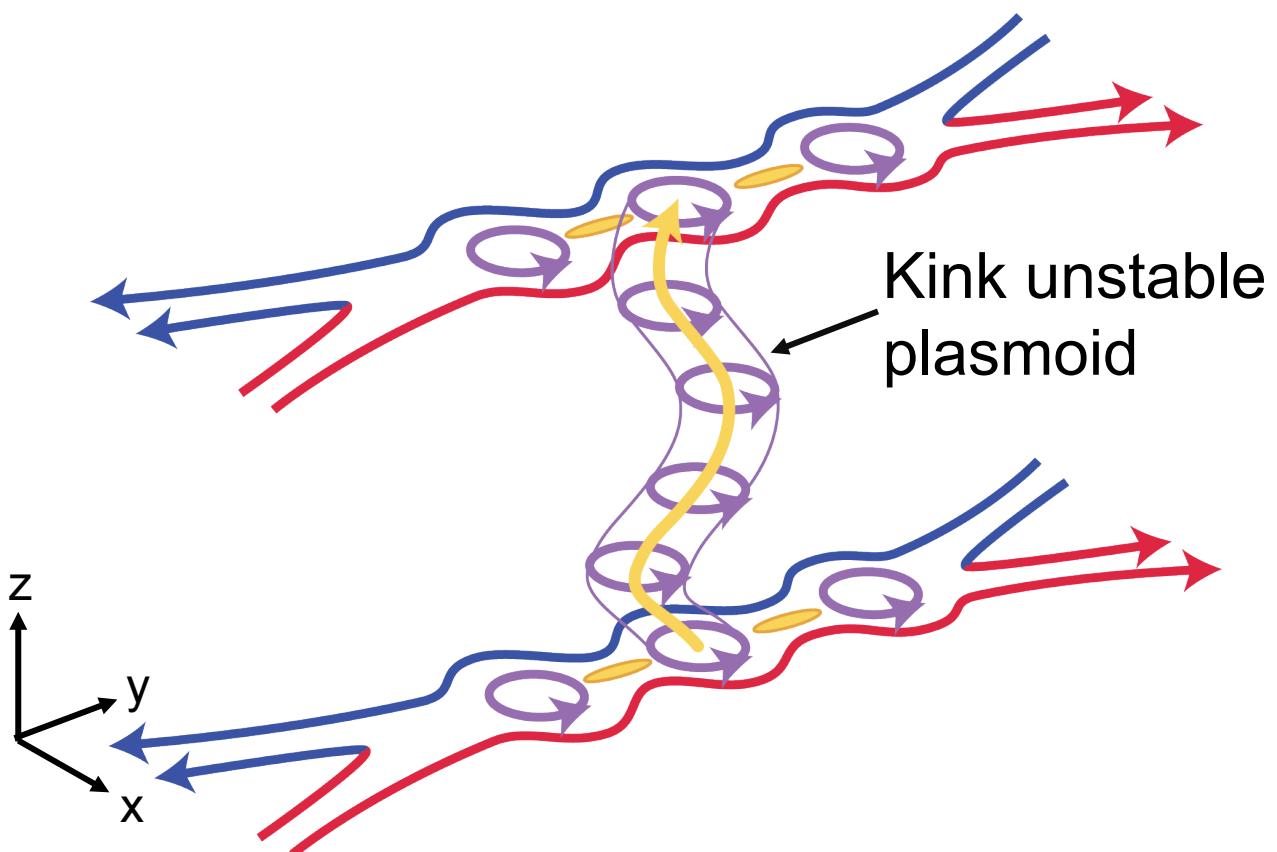
Lapenta, G., and L. Bettarini. *EPL* 93, 6 (2011)



Hints of a kink instability on MAGPIE

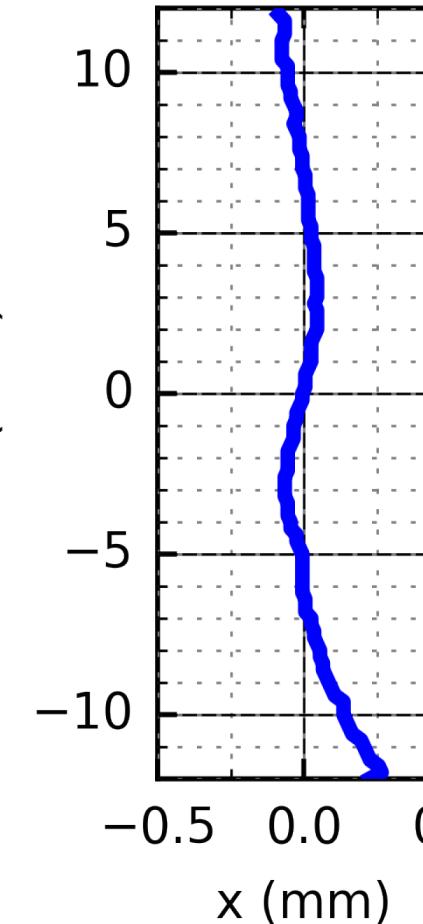
Open question:

What is the long-term behaviour of this kink instability?

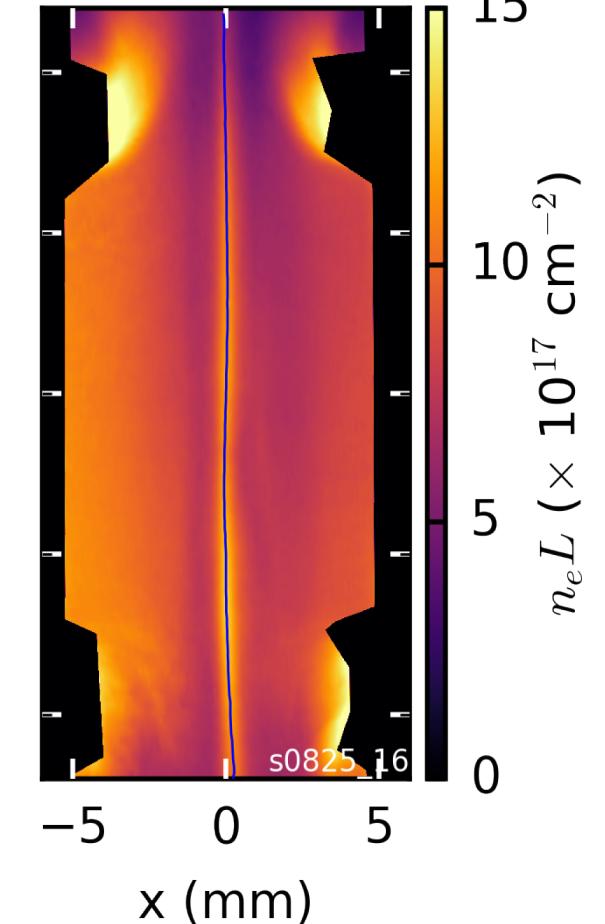


Current Sheet at 301 ns

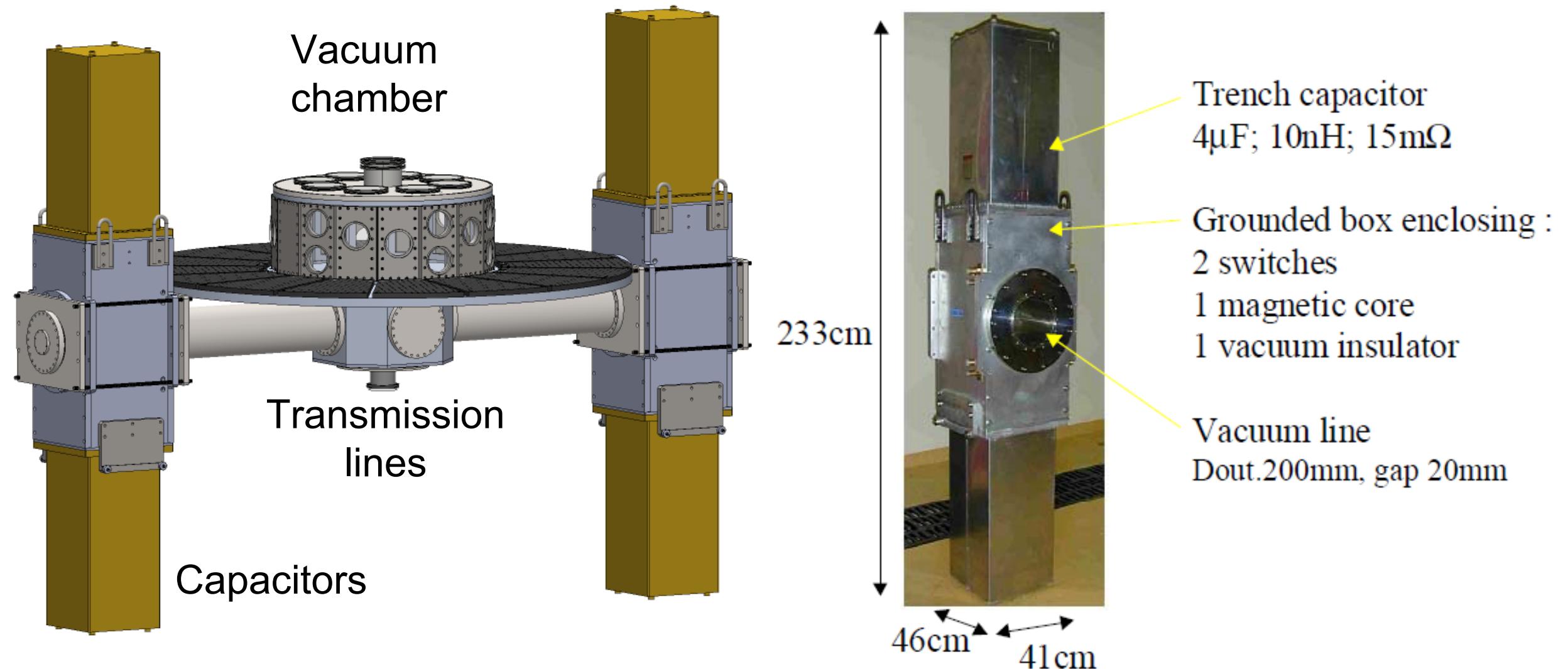
a) Position



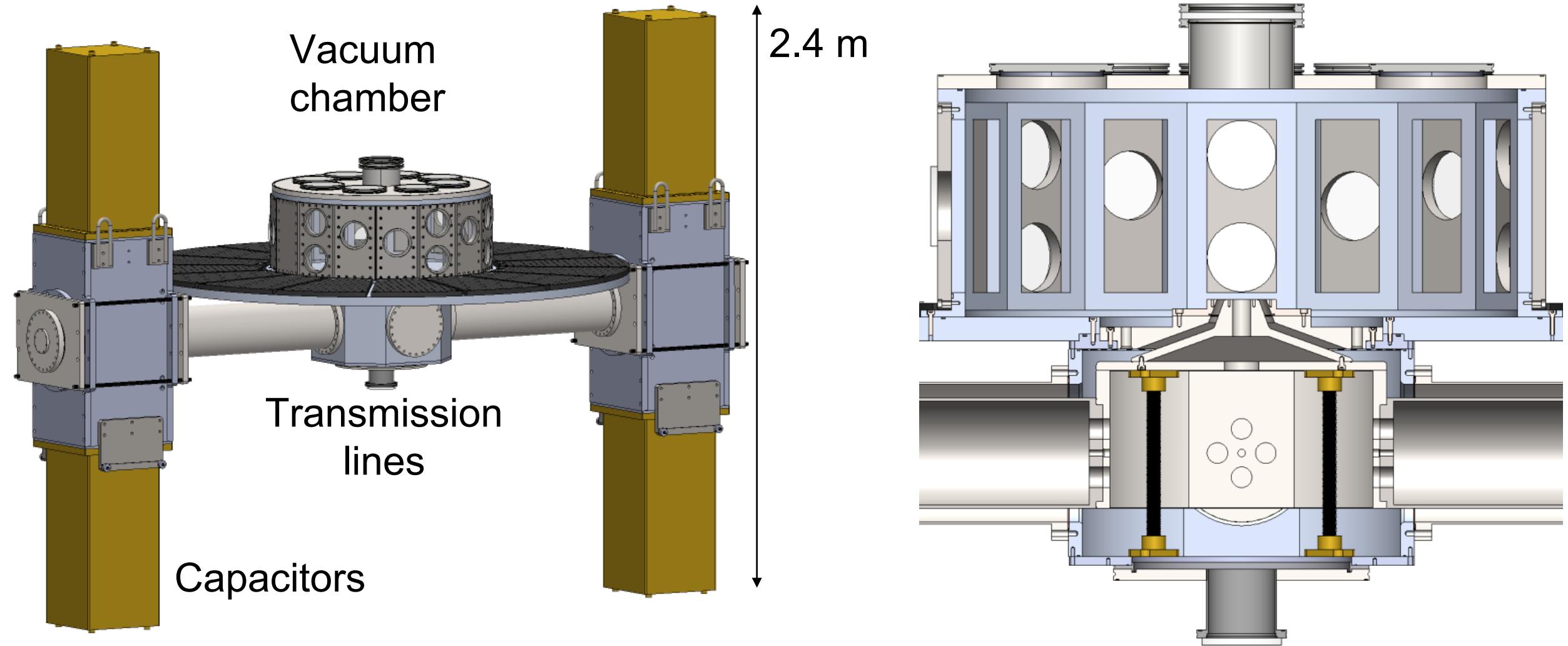
b) Areal density



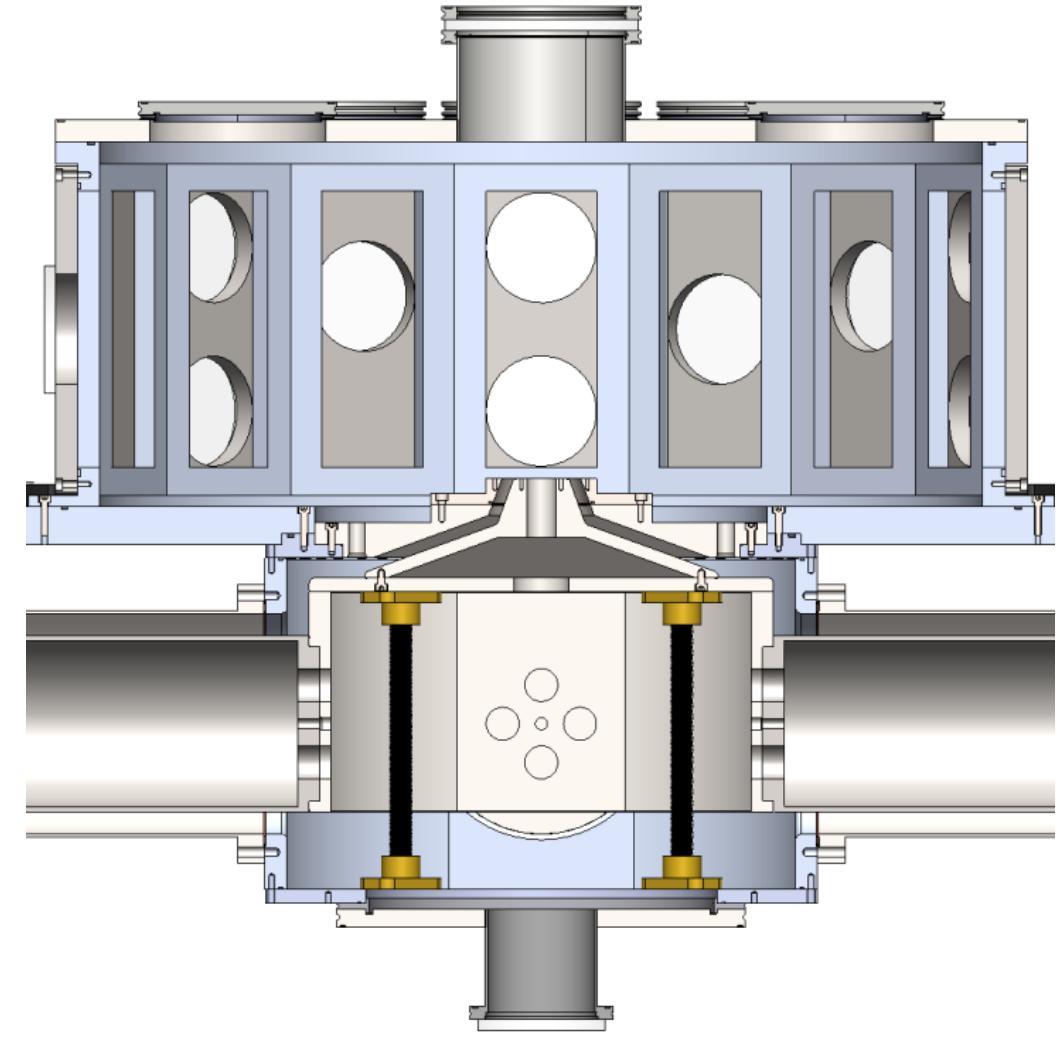
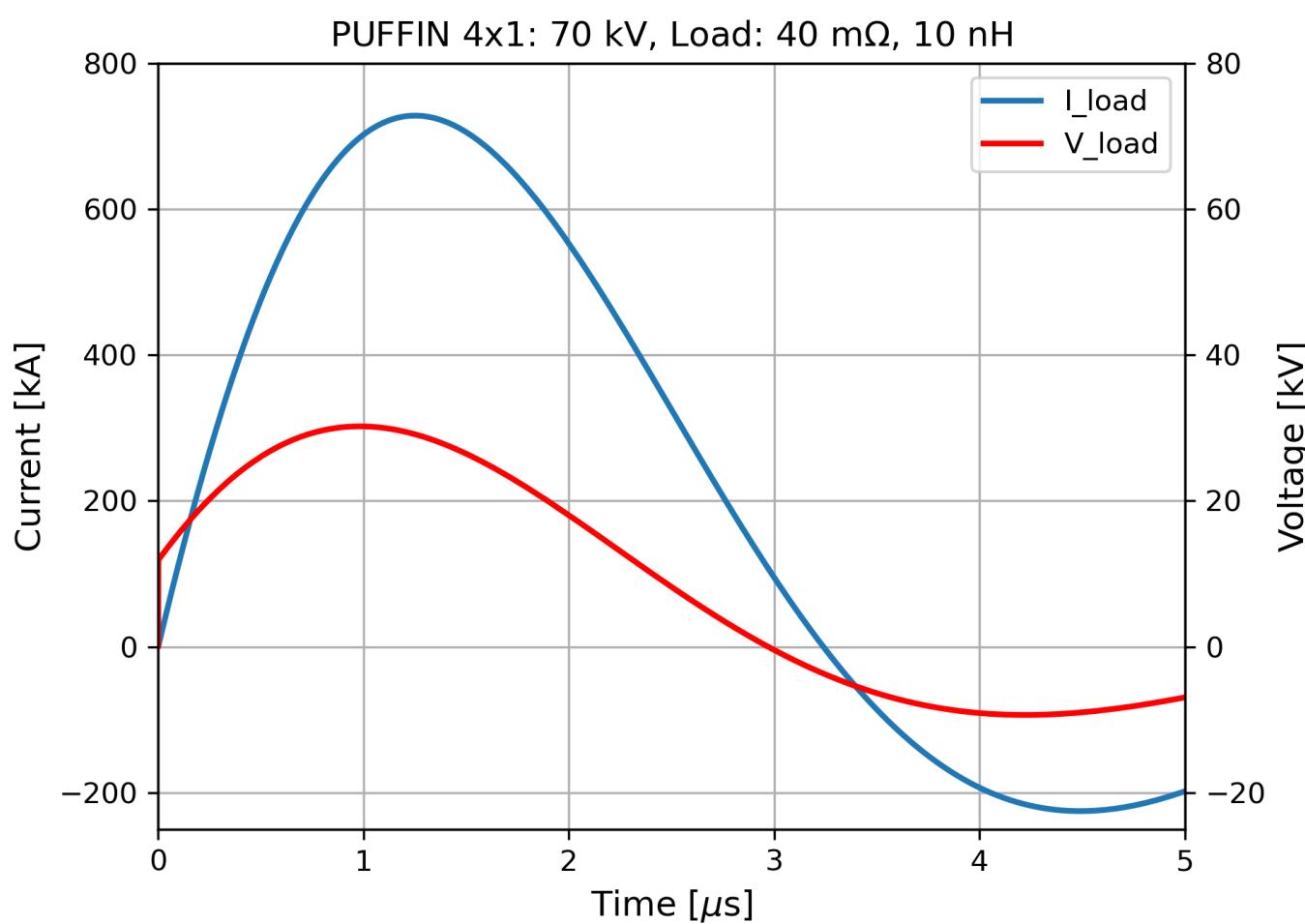
PUFFIN uses 2 LTD5 stages from SPHINX (CEA Gramat)



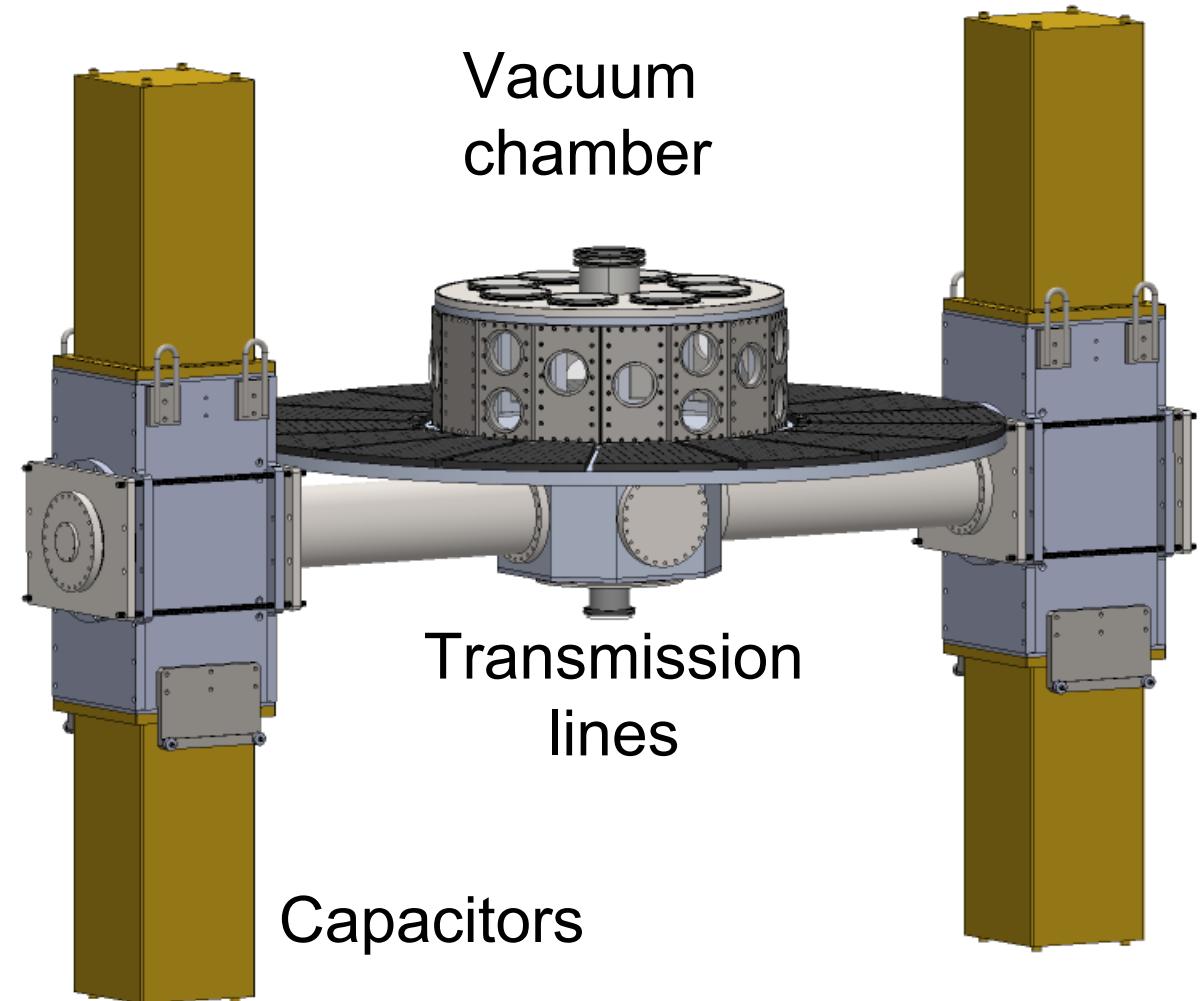
PUFFIN 2x1 will drive around 700 kA with a 1.6 μ s rise time



PUFFIN 2x1 will drive \sim 700 kA with a \sim 1.5 μ s rise time



PUFFIN: a long-pulse driver for fundamental physics



Timeline:

- LTD stages arrive May 2022
- Lab renovation finished December 2022
- Design and construction of power feed expected early 2023
- First experiments in 2023

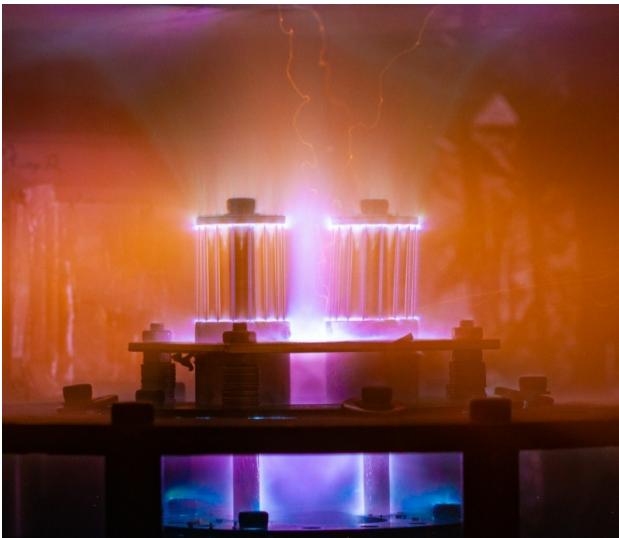
Diagnostics:

- Laser imaging shadowgraphy, schlieren, interferometry and imaging refractometry
- XUV and optical framing cameras
- Thomson scattering (spatially resolved, ion and electron features)

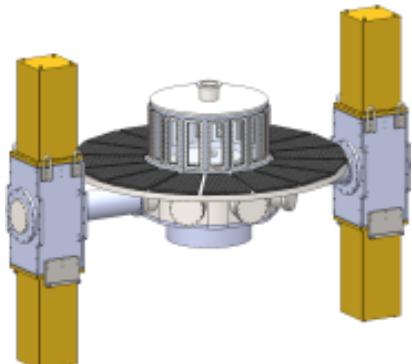


Summary

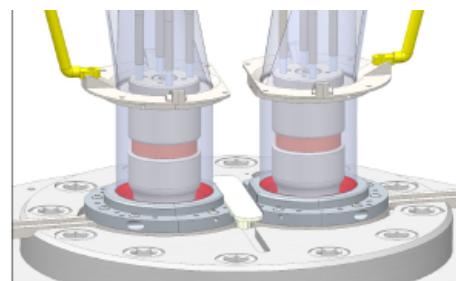
Magnetic reconnection



PUFFIN: $1.6 \mu\text{s}$,
700 kA at MIT



Radiatively cooled
reconnection on Z



Unifying theme: energy flows in plasmas

