

The New PUFFIN Generator at MIT

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Acknowledgements



MIT

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London

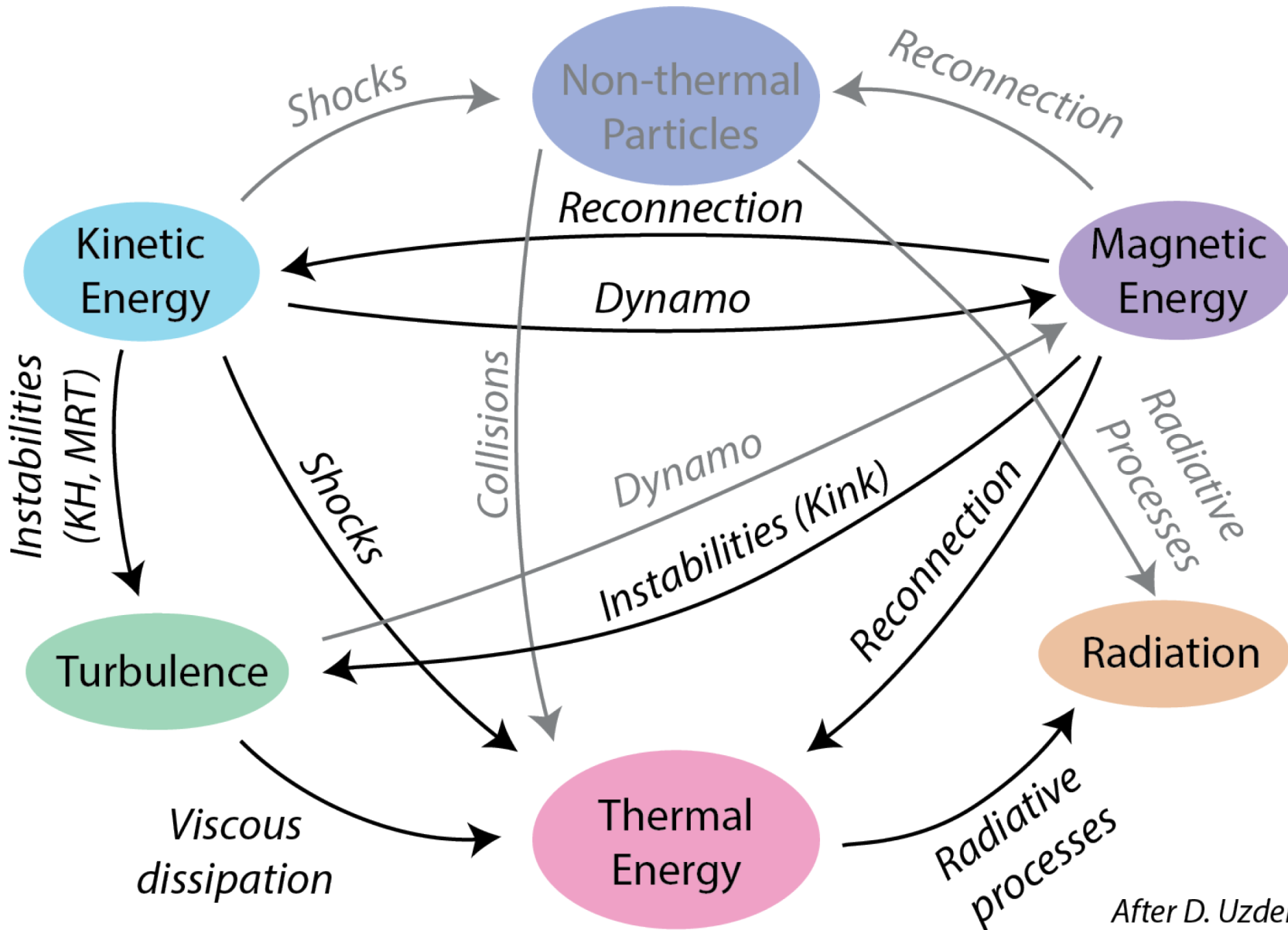


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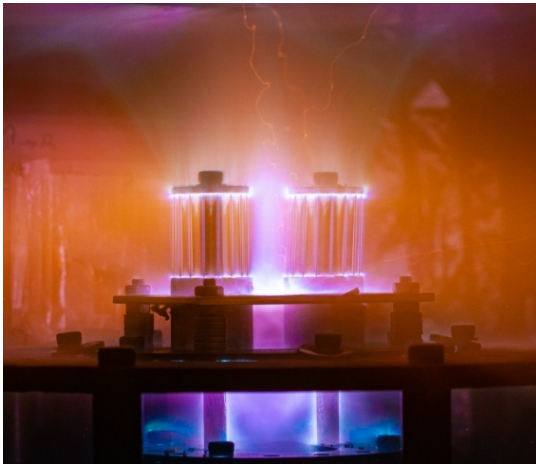
Research Interests: Energy Flows in Plasma



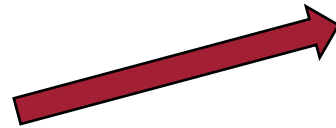
After D. Uzdensky



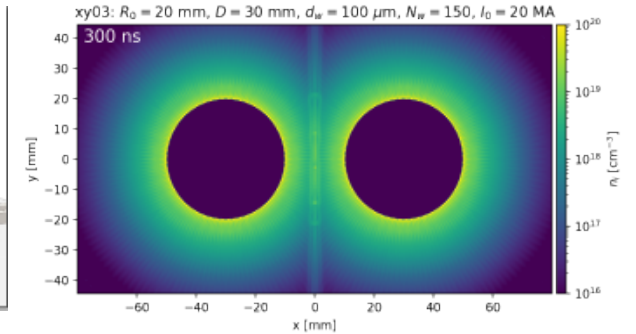
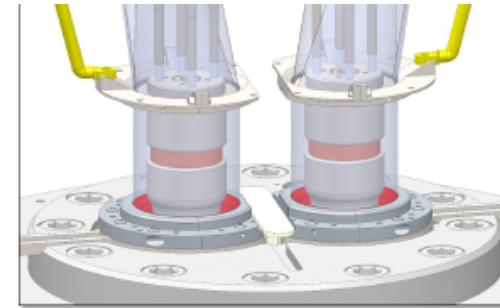
Magnetic
reconnection



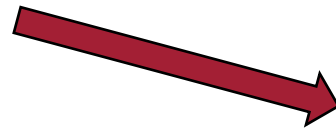
More current



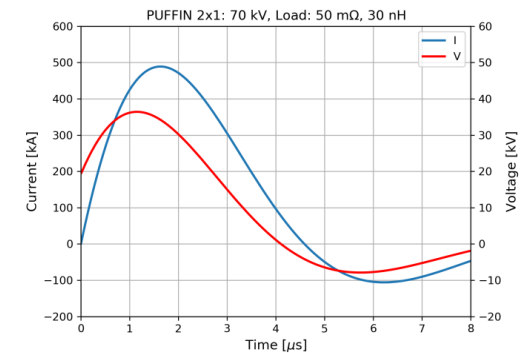
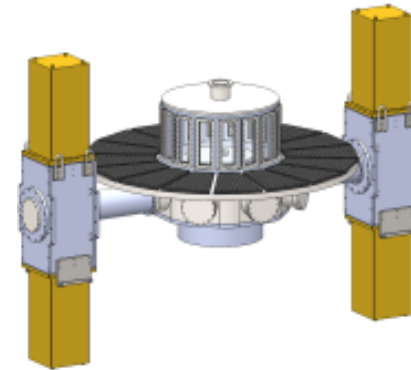
Radiatively cooled reconnection on Z



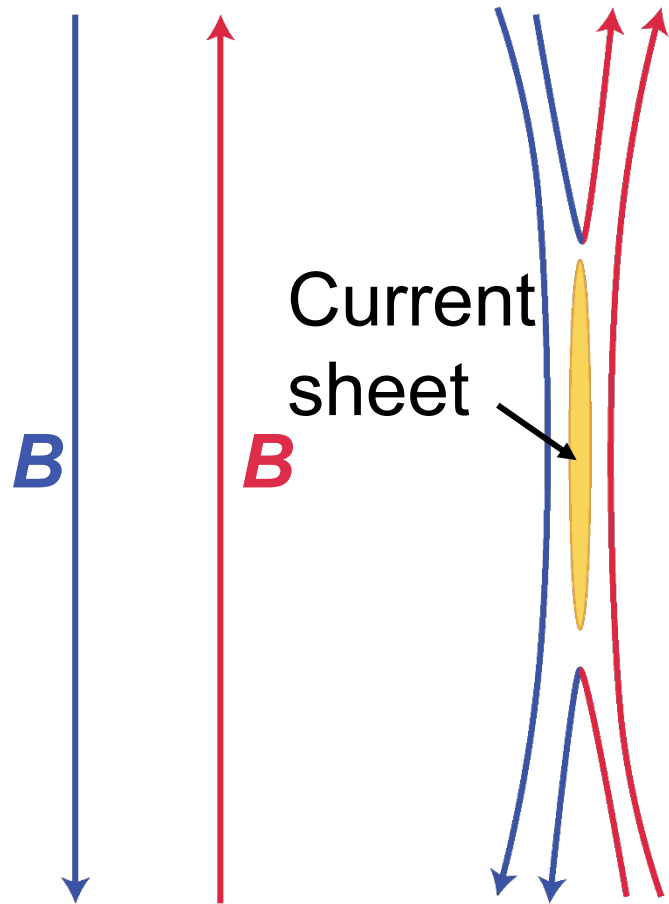
Longer current
pulse



PUFFIN: 2 μ s, 500 kA at MIT



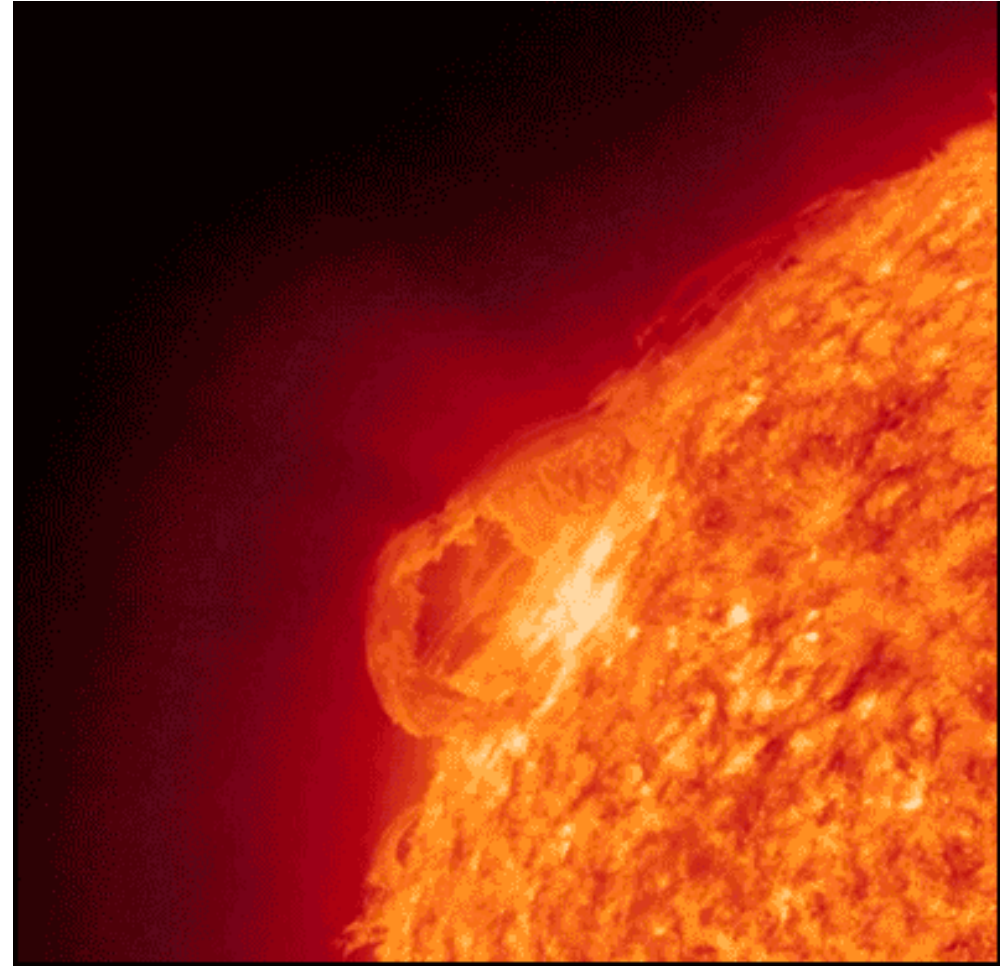
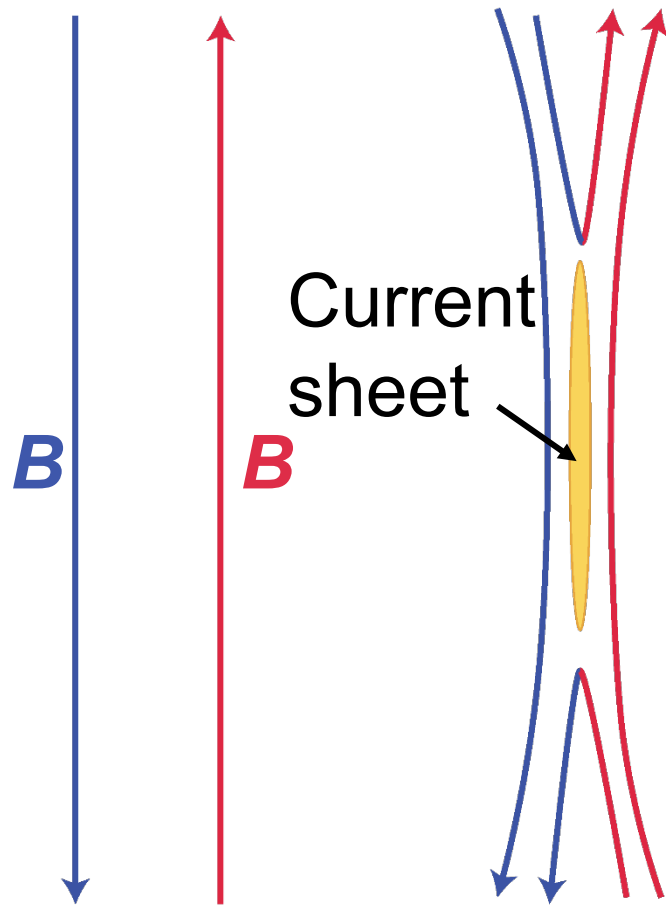
Magnetic Reconnection



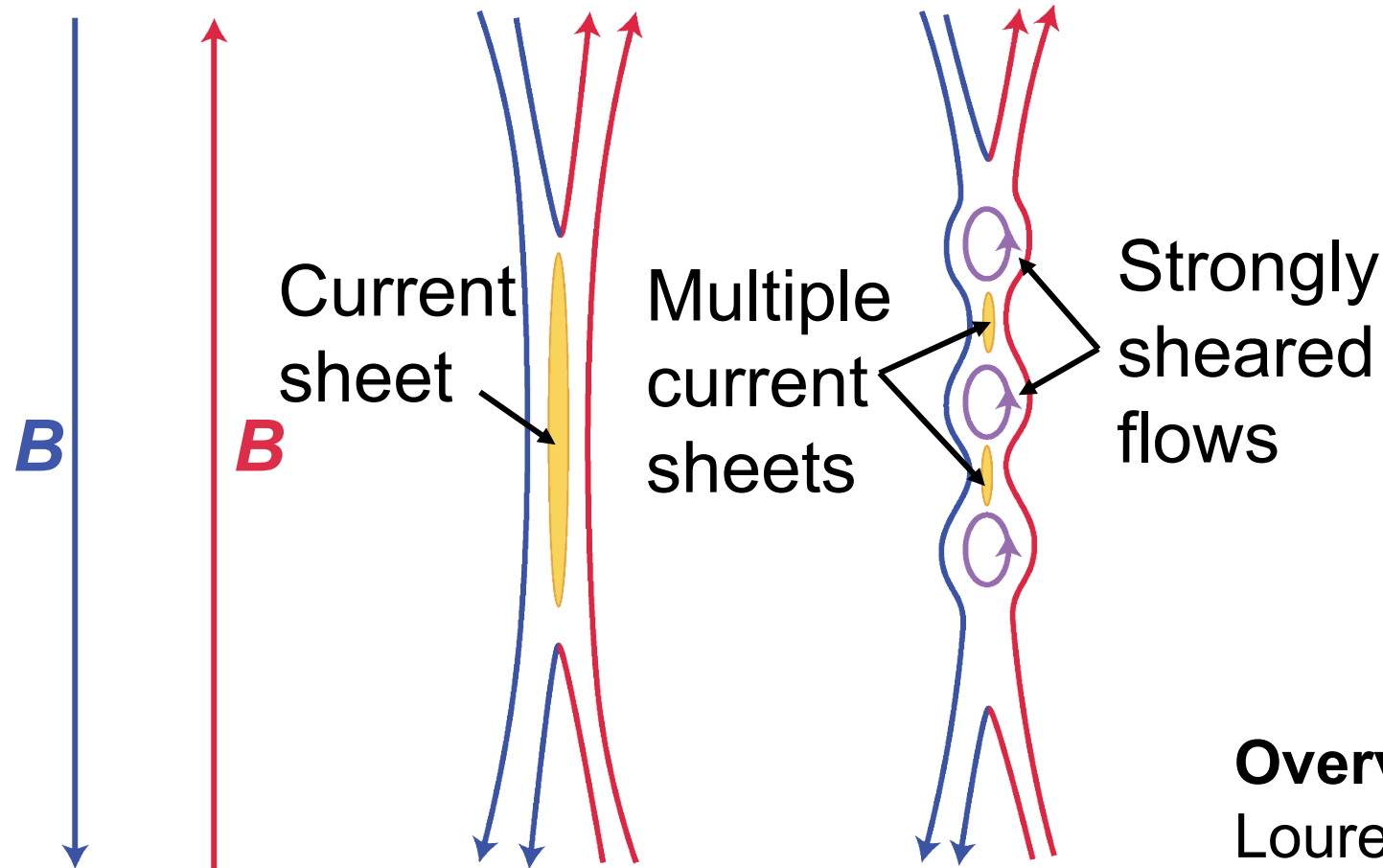
Magnetic Reconnection



Prediction: 1000 yrs. Reality: 10 minutes!

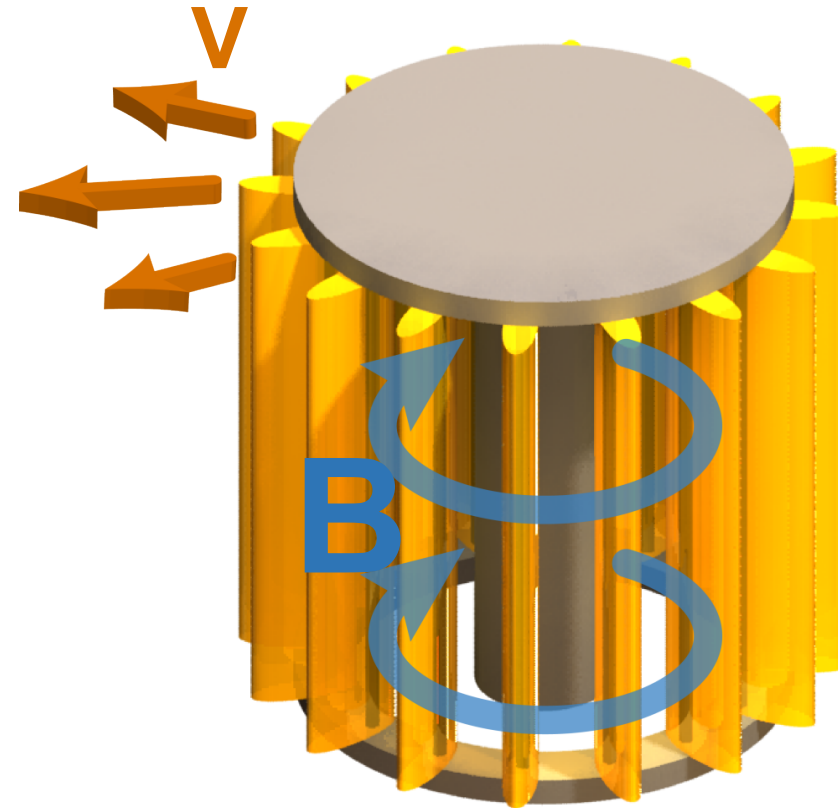
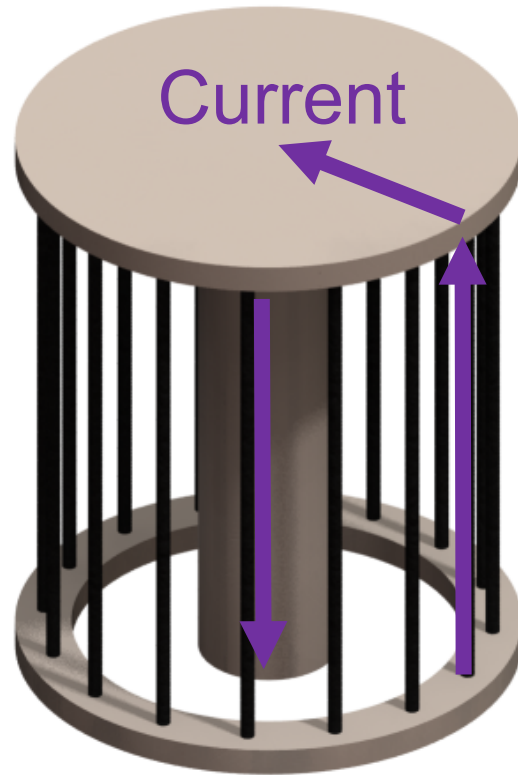


Plasmoids Lead to Fast Reconnection and Anomalous Heating



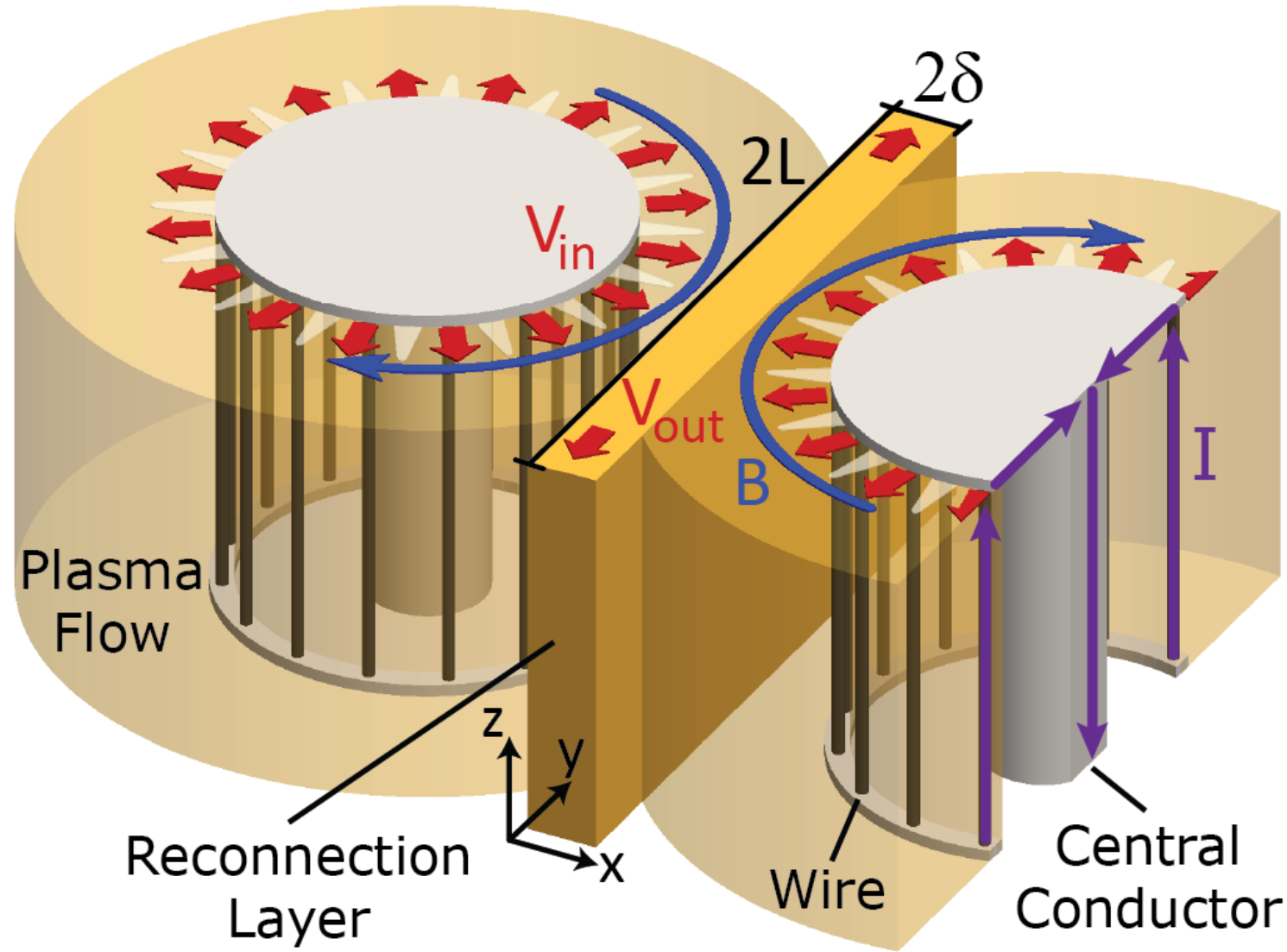
Overview of recent theory:
Loureiro, N. F., & Uzdensky, D.
A.(2015).
PPCF, 58, 014021

Pulsed-power-driven Magnetic Reconnection

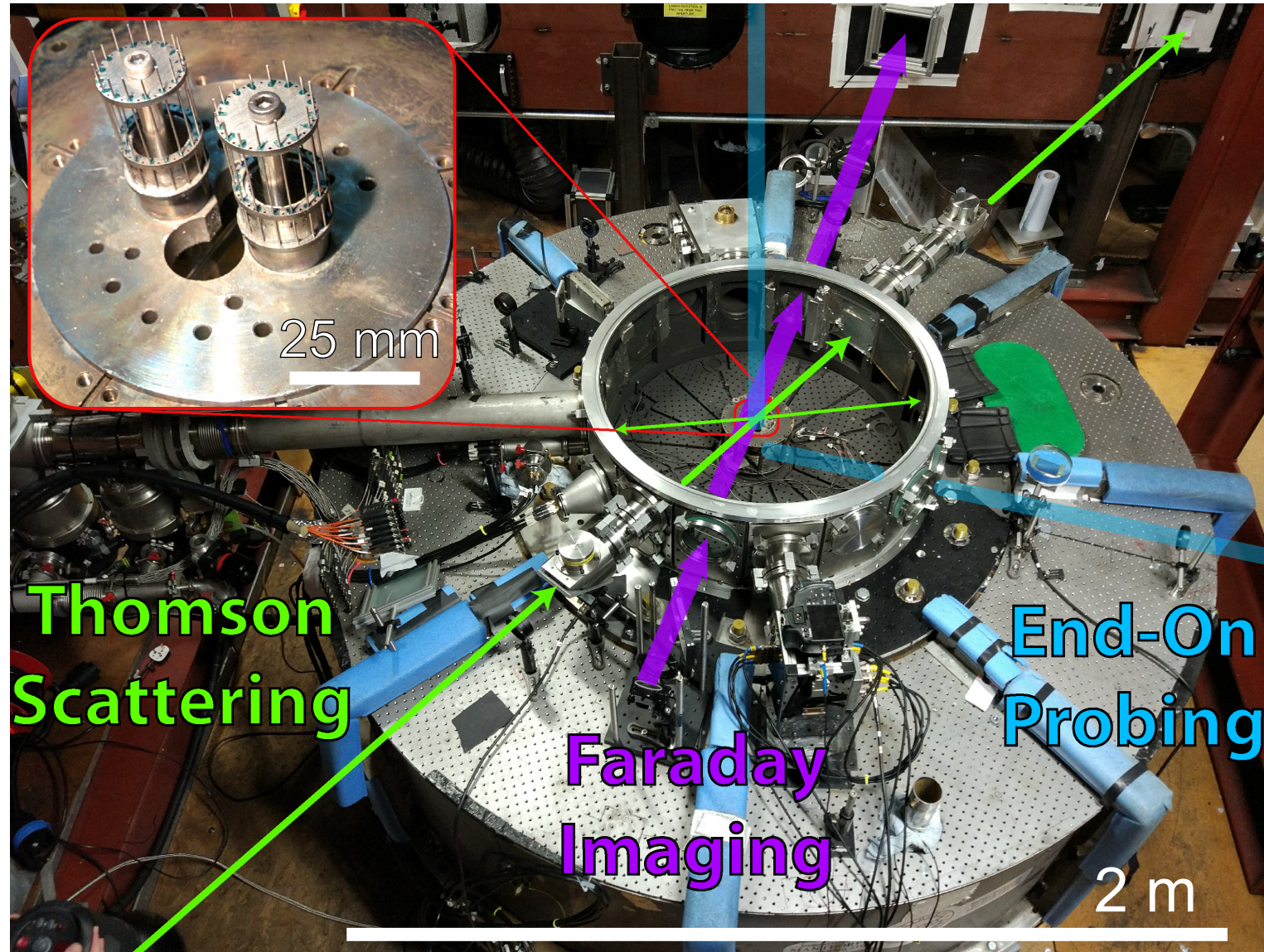


$I=1.4$ MA, 240 ns rise time

Magnetic Reconnection from Double Exploding Wire Arrays



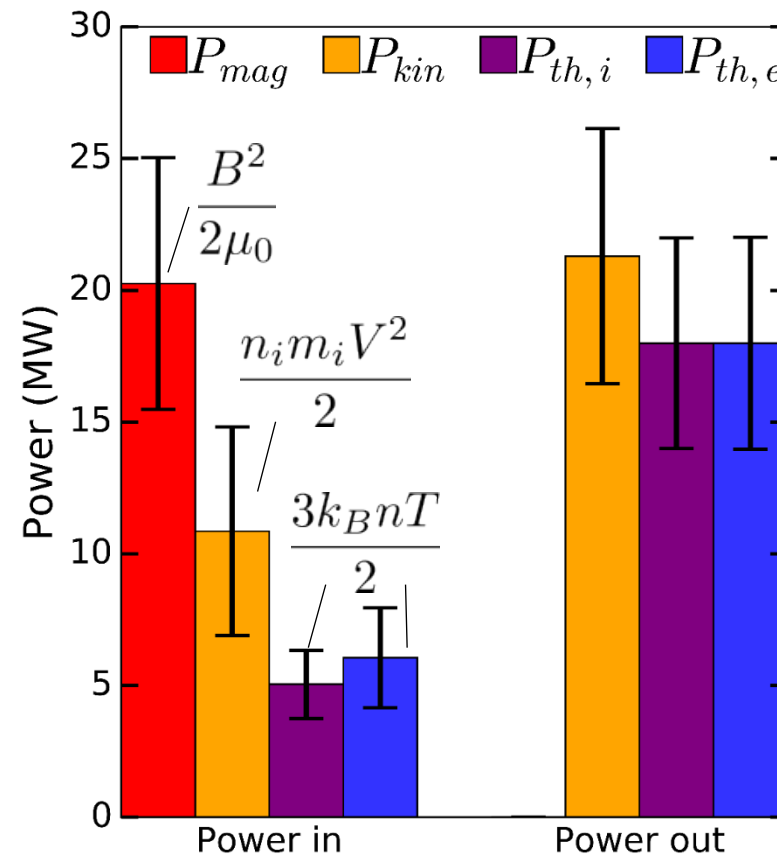
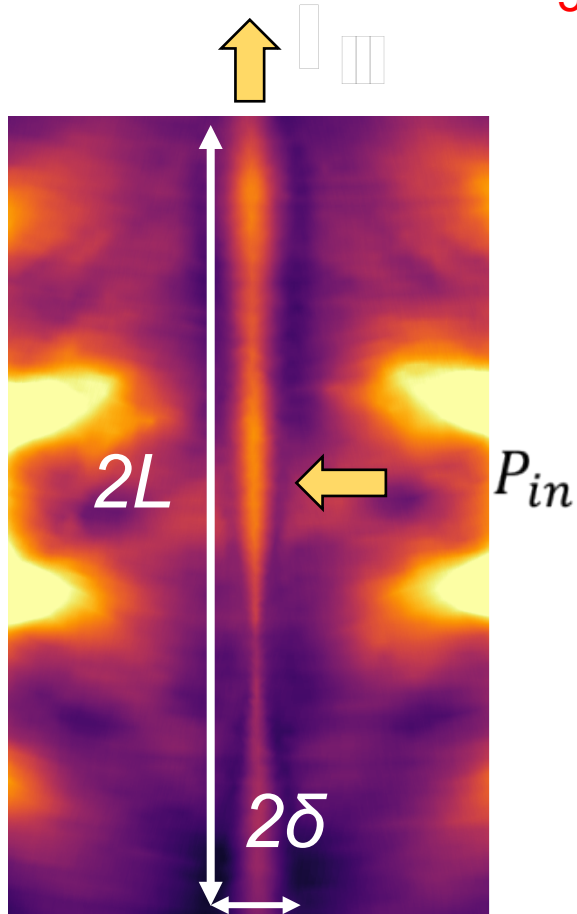
Overview of Diagnostic Suite on MAGPIE at Imperial College




Anomalous Heating in the Reconnection Layer



$$V_{in} L h \left(\underset{\sim 50\%}{E_{mag}} + \underset{\sim 25\%}{E_{kin}} + \underset{\sim 25\%}{E_{th,i}} + E_{th,e} \right) \approx V_{out} \delta h \left(\underset{\sim 40\%}{E_{kin}} + \underset{\sim 60\%}{E_{th,i}} + E_{th,e} \right)$$



Classical heating is too slow:

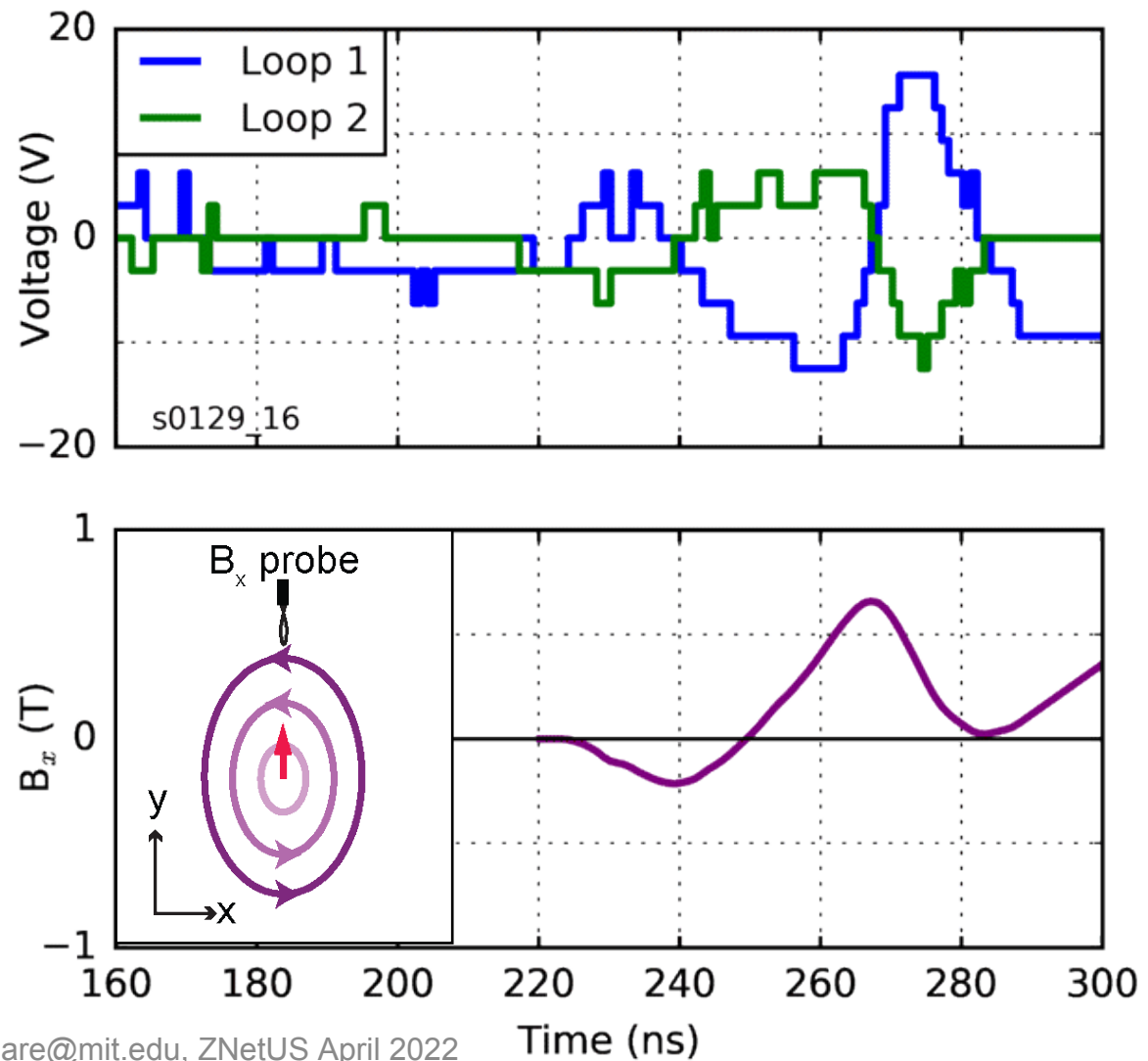
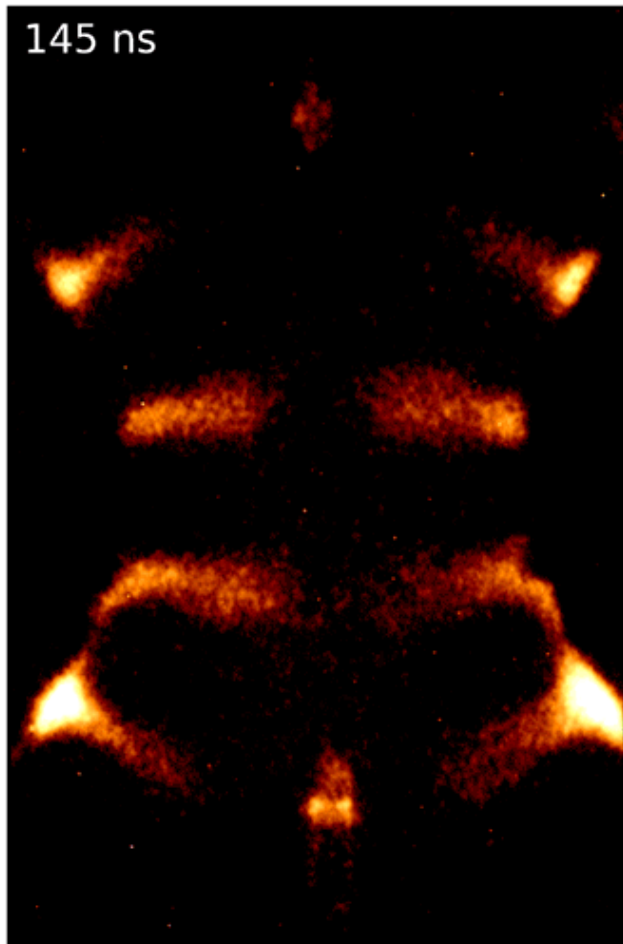
 ≈ 50 ns

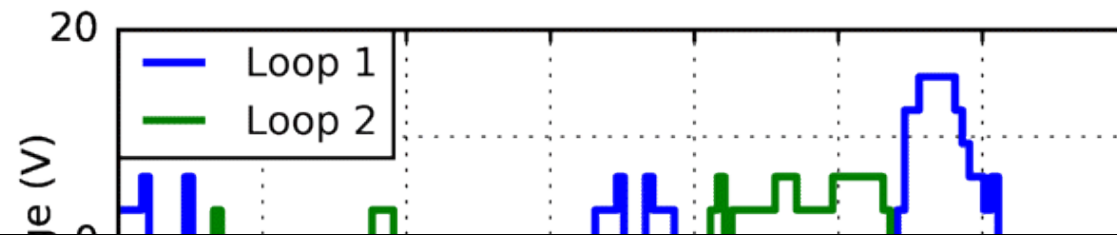
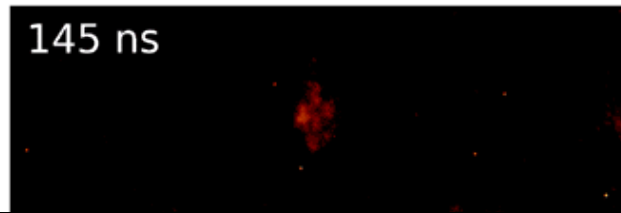
 ≈ 800 ns

 ≈ 350 ns

$$\tau_{exp} \ll \tau_{visc}, \tau_{res}$$

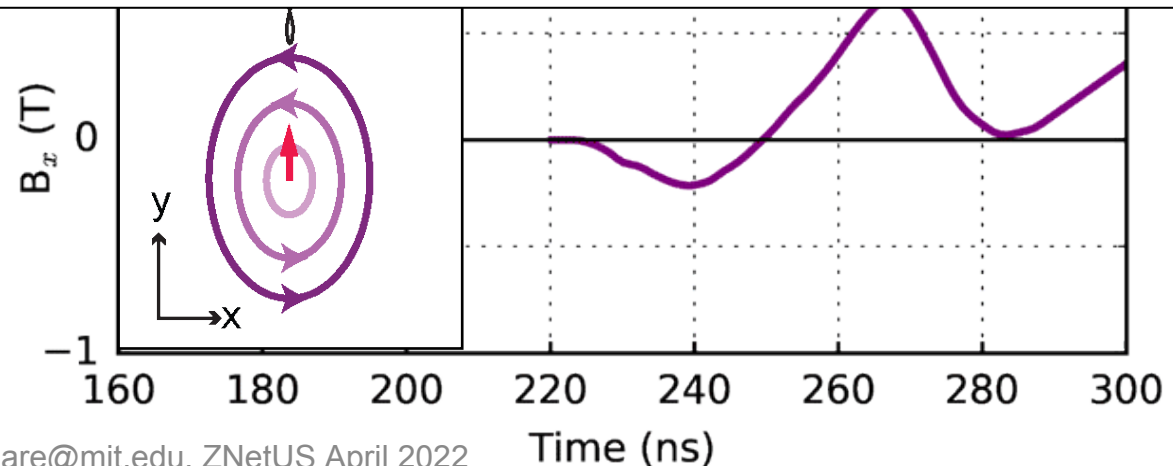
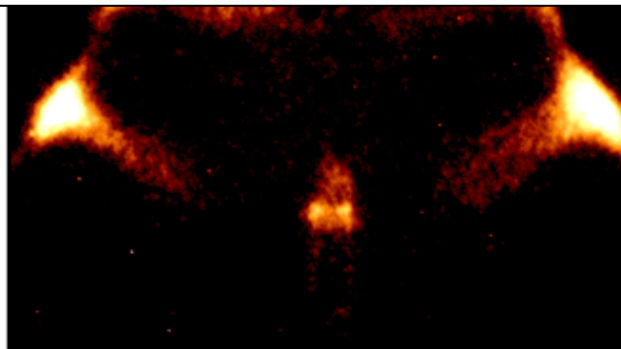
Magnetic Structure of Plasmoids





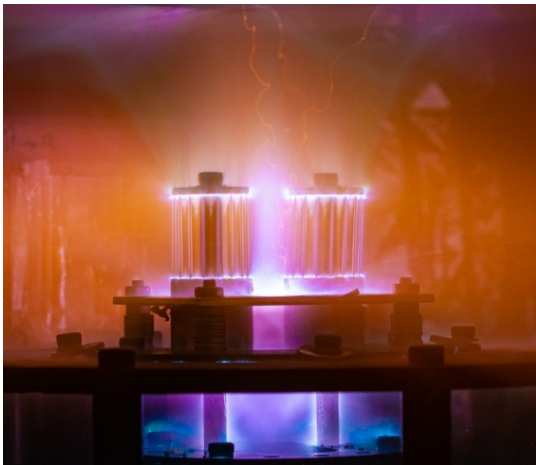
Open questions:

- How does reconnection layer dynamics scale with current?
- What is the quasi-steady behavior of reconnection?

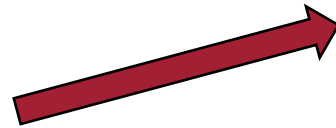




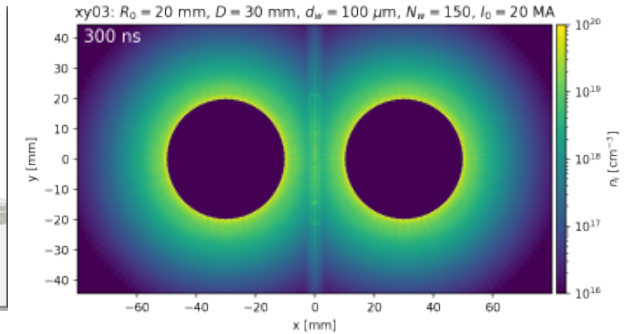
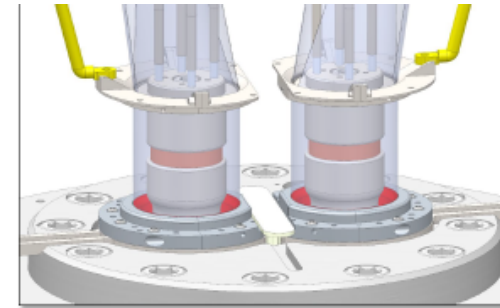
Magnetic
reconnection



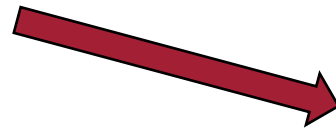
More current



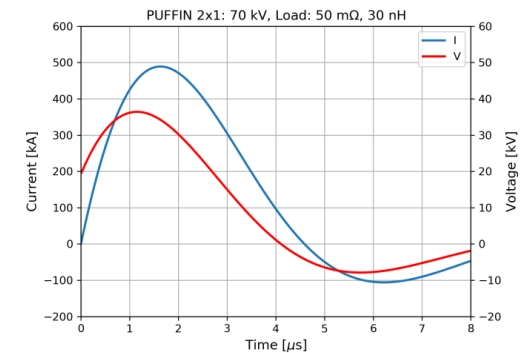
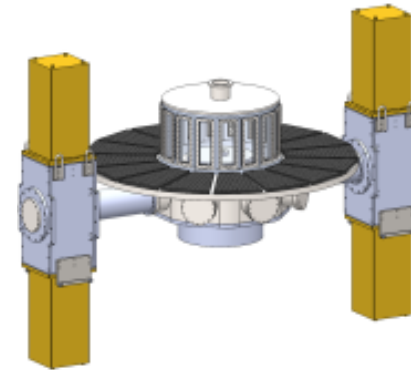
Radiatively cooled reconnection on Z



Longer current
pulse



PUFFIN: 2 μ s, 500 kA at MIT



Reconnection in Extreme Astrophysical Environments



Artist's impression of a black hole



M87 (EHT)



Crab Pulsar (Hubble/Chandra)



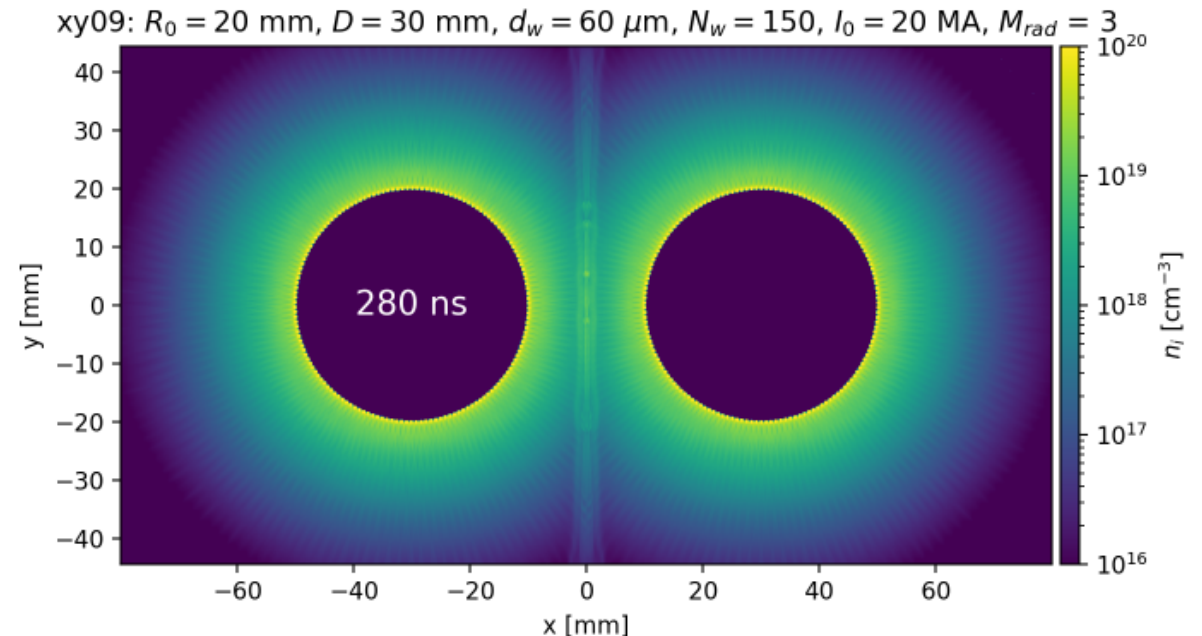
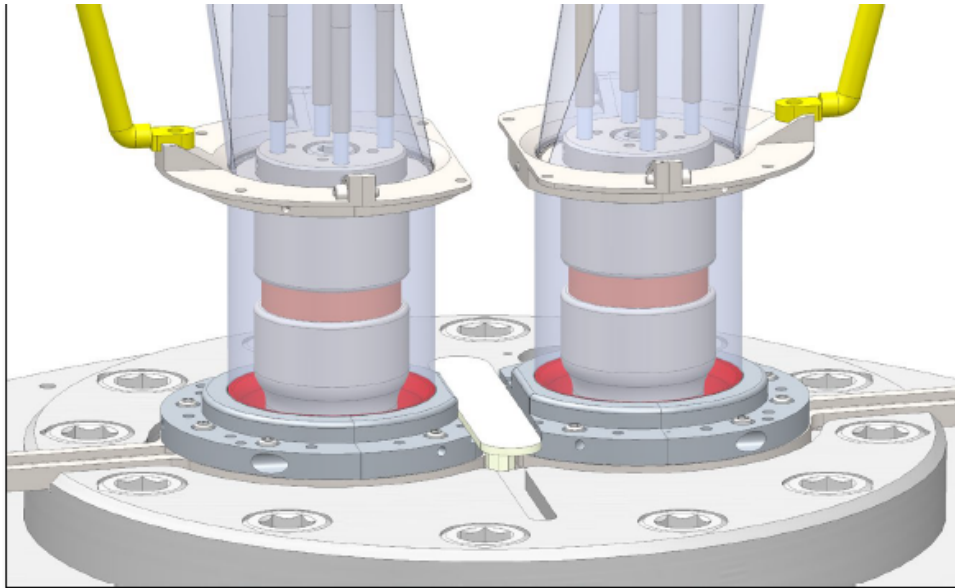
1. Cooling is a significant loss mechanism:
 - Modifies partition of magnetic energy between electrons, ions, kinetic
 - Leads to cooling instabilities, radiative collapse, termination of reconnection
2. X-rays: key observational signature in remote environments:
 - Where and when are X-rays produced – localized bursts?
 - How does this couple back to the reconnection process?



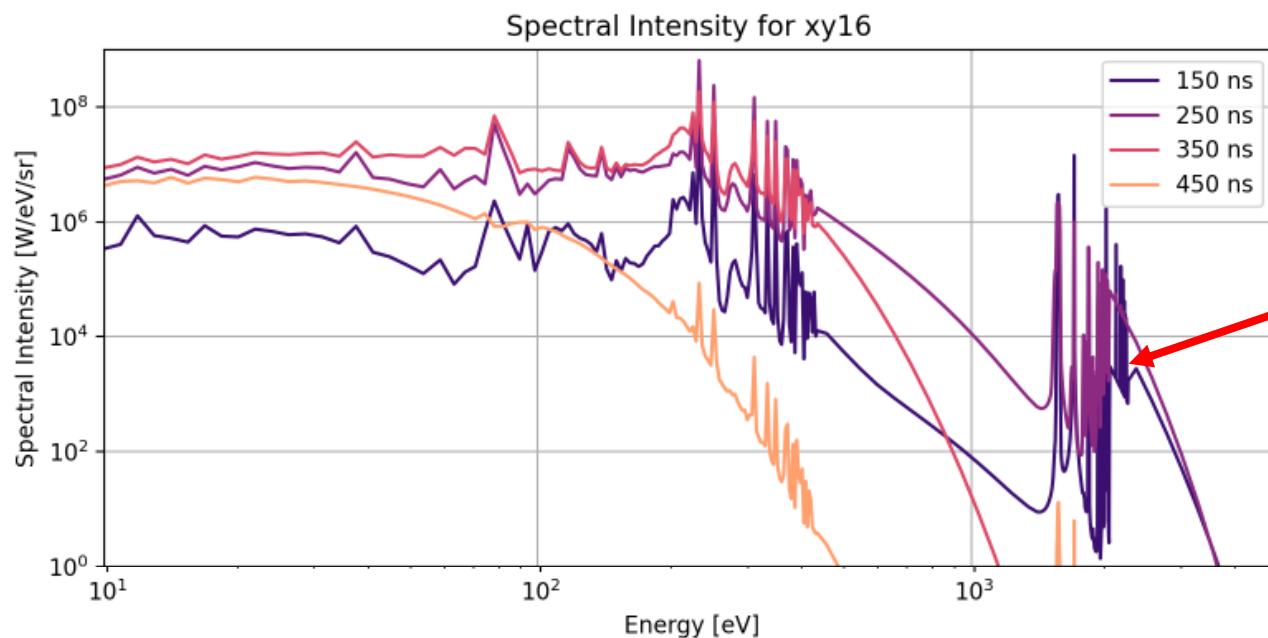
Z is the largest pulsed-power machine in the world

- 20-30 MA peak current compared to 1.4 MA on MAGPIE:
 - **Density** increase by $I^2 \sim 400$
 - **Magnetic energy** increase by $I^2 \sim 400$
 - **Cooling rate** increase by $I^4 \sim 160,000$

Unique capability: strongly radiatively cooled reconnection



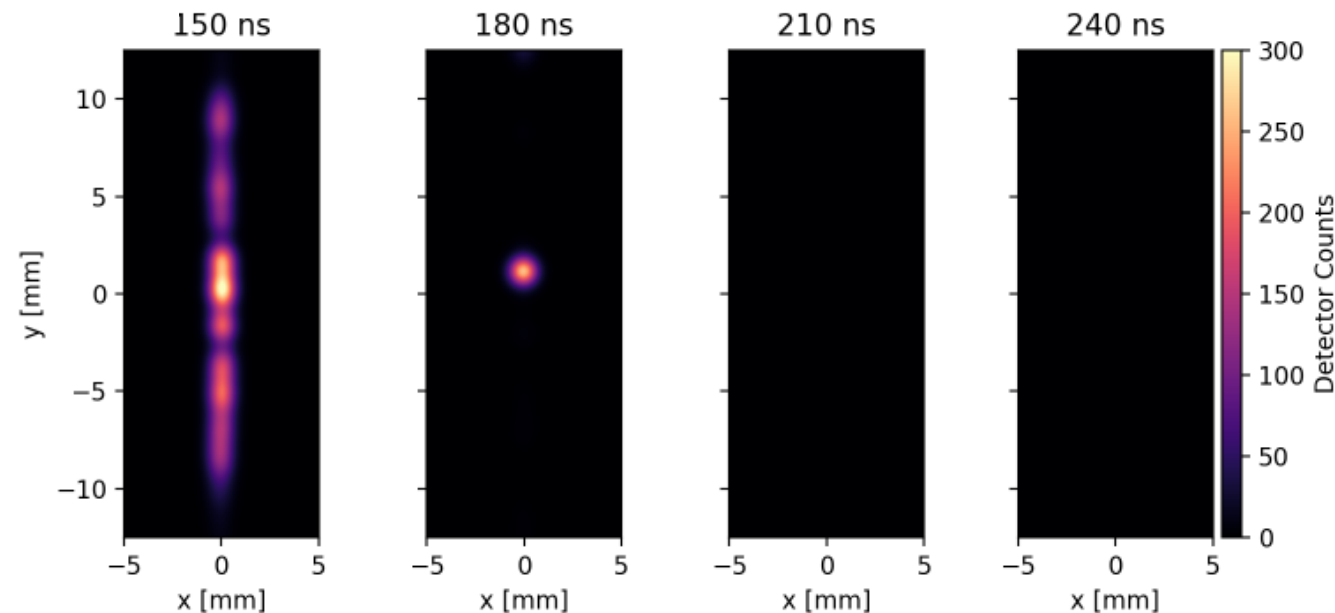
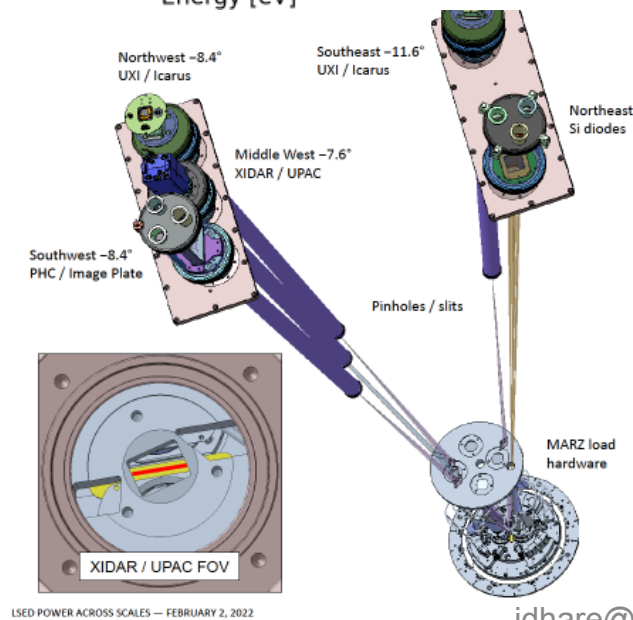
XP2: X-ray Post-Processor by Aidan Crilly & Jerry Chittenden



Al K-shell
disappears
after
collapse

ICARUS for 3DMARZ with 10umBe filter, 150 um pinhole

XP2 provides a
predictive
capability for X-
ray diagnostics

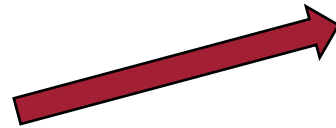




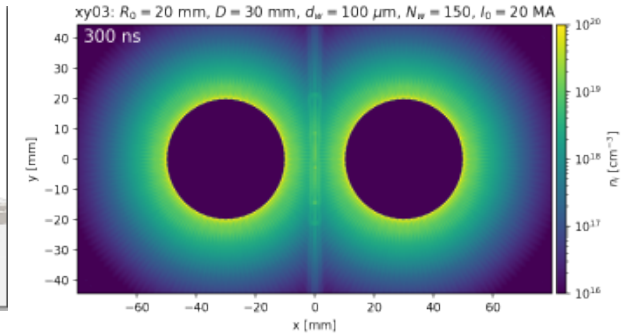
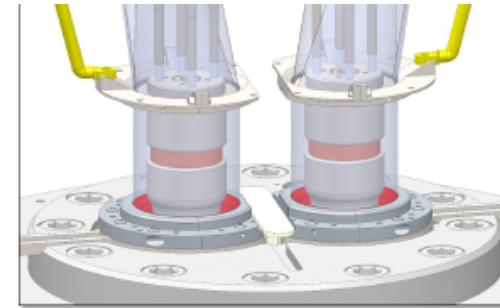
Magnetic
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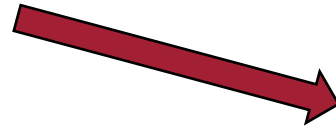
More current



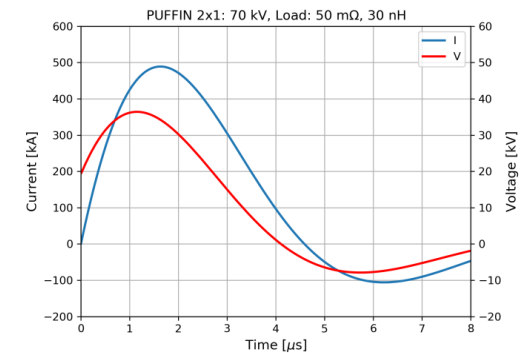
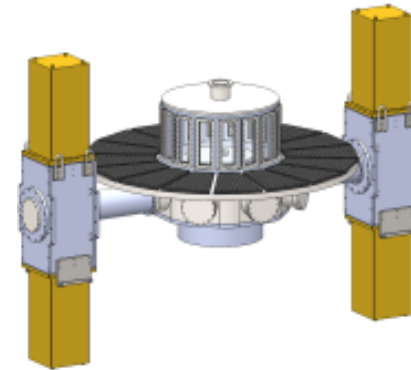
Radiatively cooled reconnection on Z



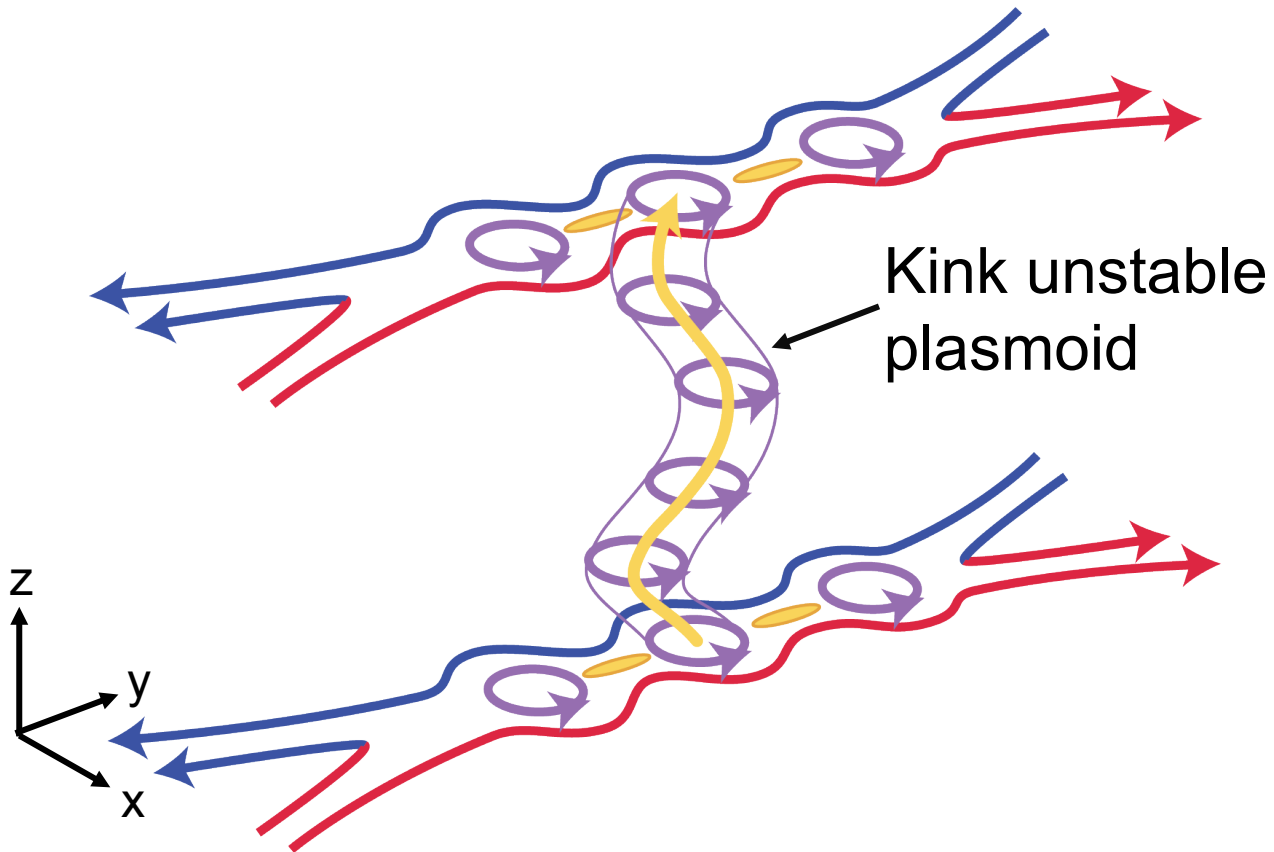
Longer current
pulse



PUFFIN: 2 μ s, 500 kA at MIT



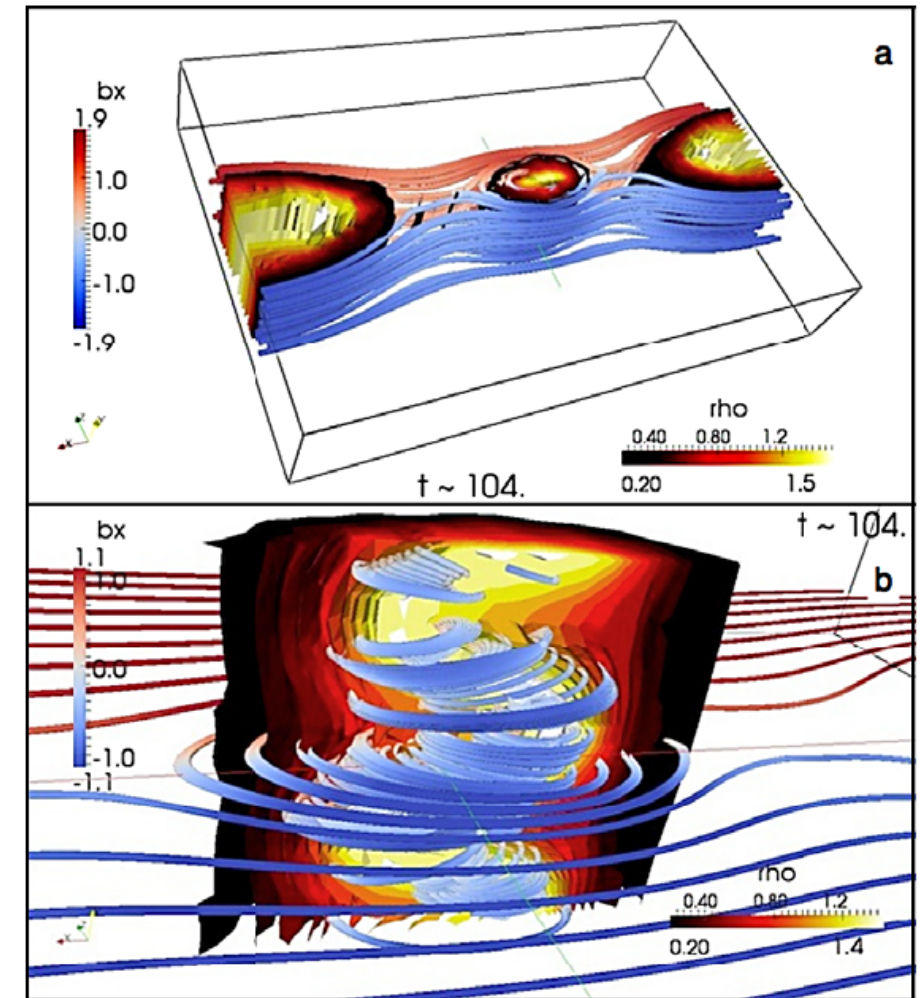
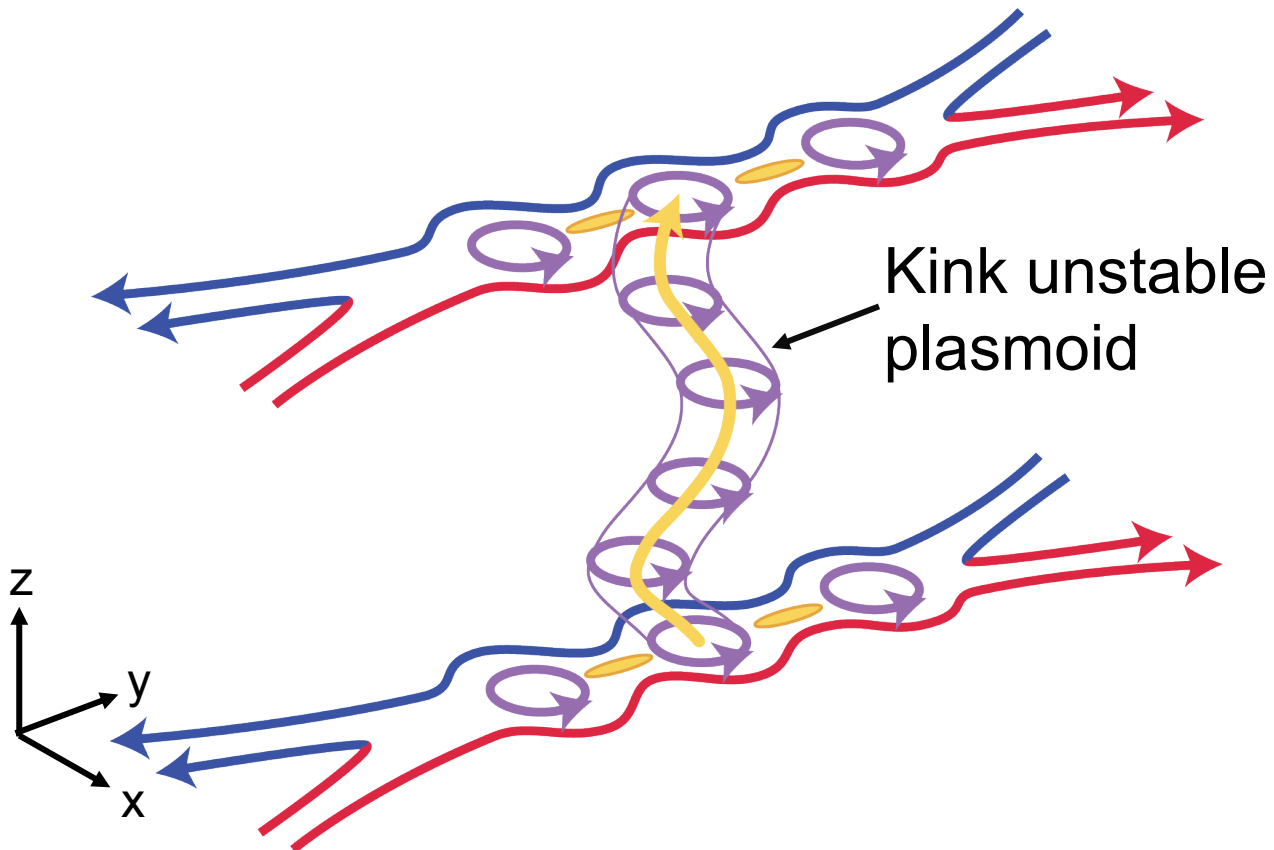
Interesting physics happens on long time scales



Kink Instability Leads to Turbulent Reconnection



Turbulent reconnection over a large volume,
as observed in astrophysics

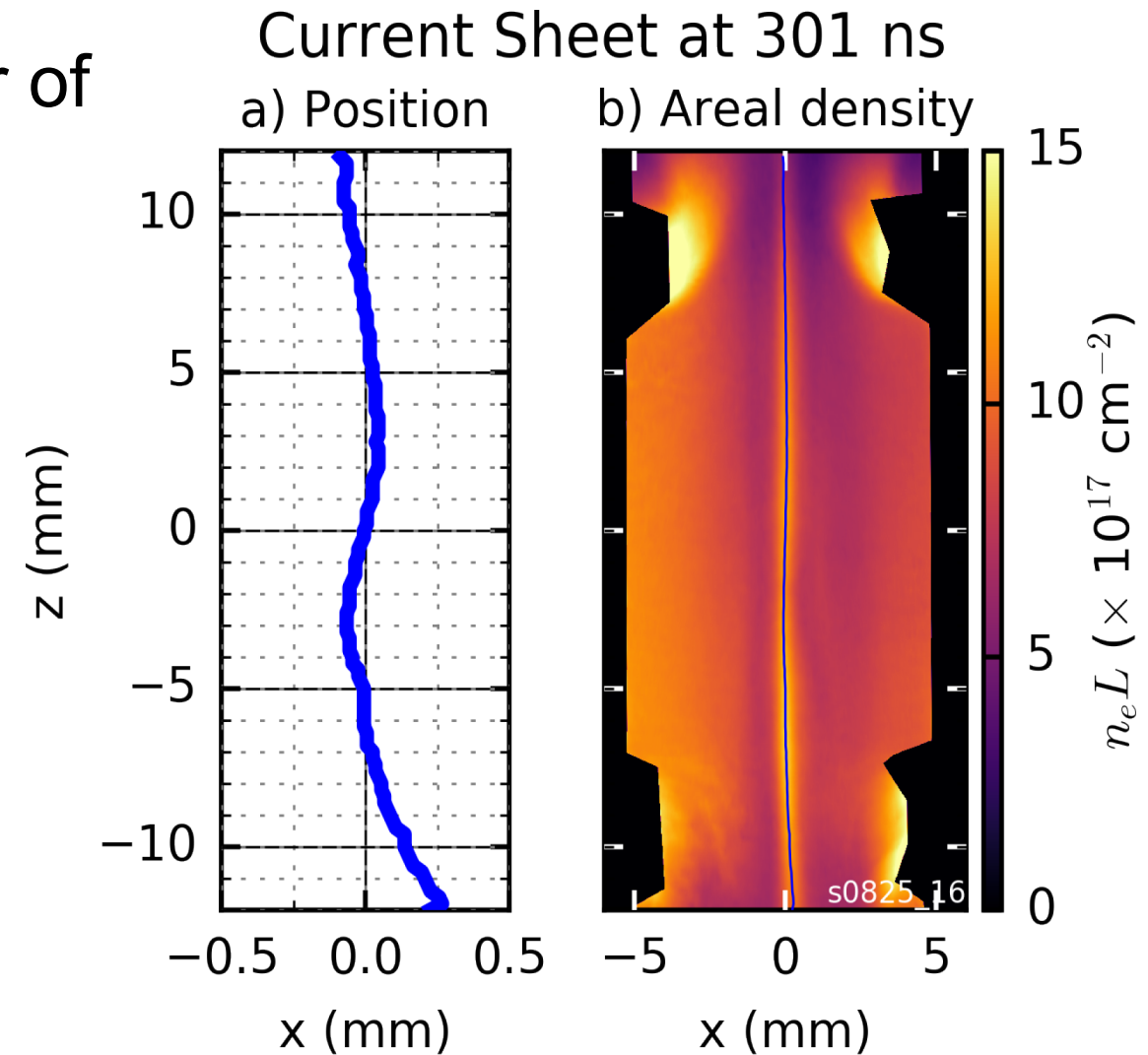
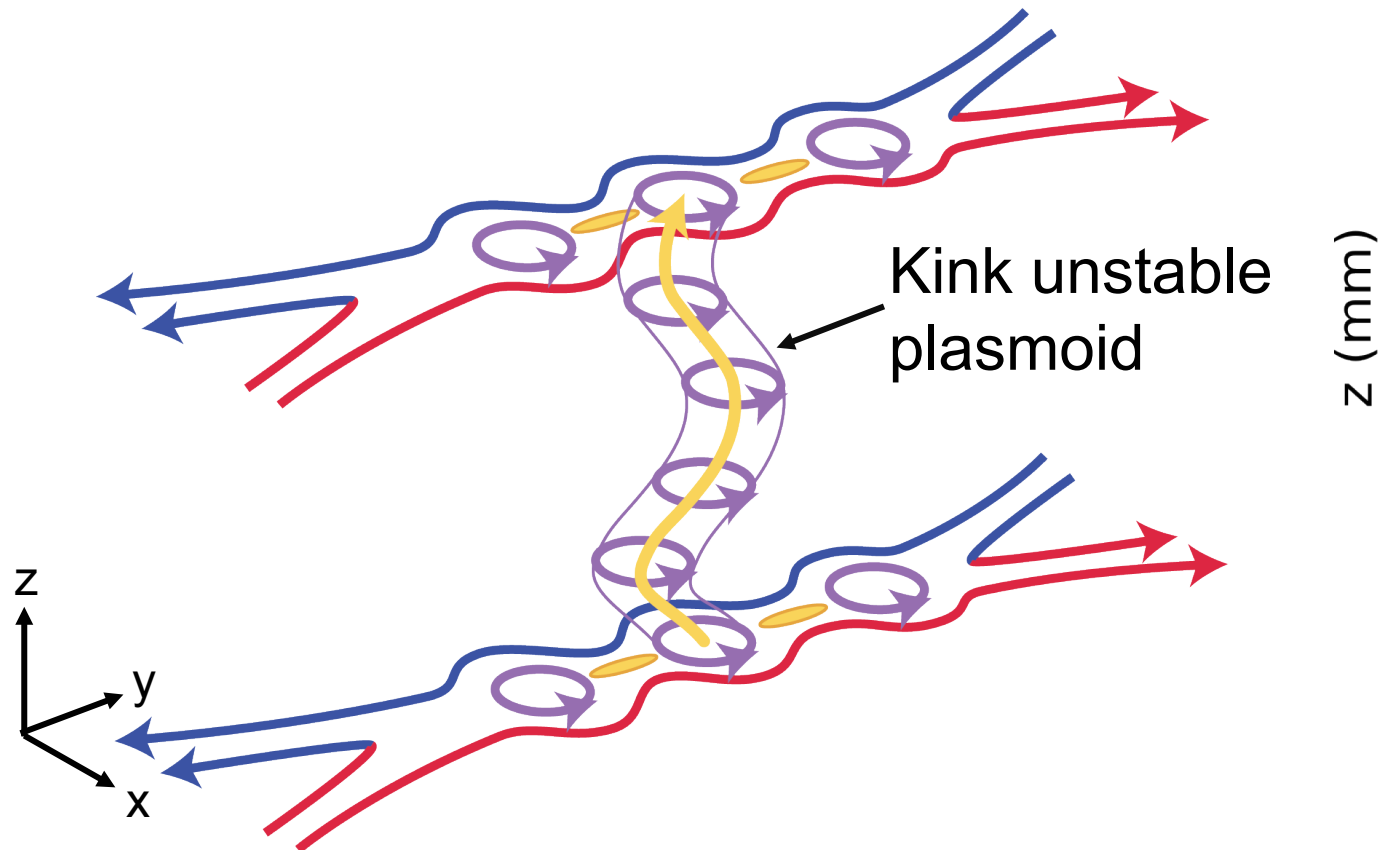


Lapenta, G., and L. Bettarini. *EPL* **93**, 6 (2011)

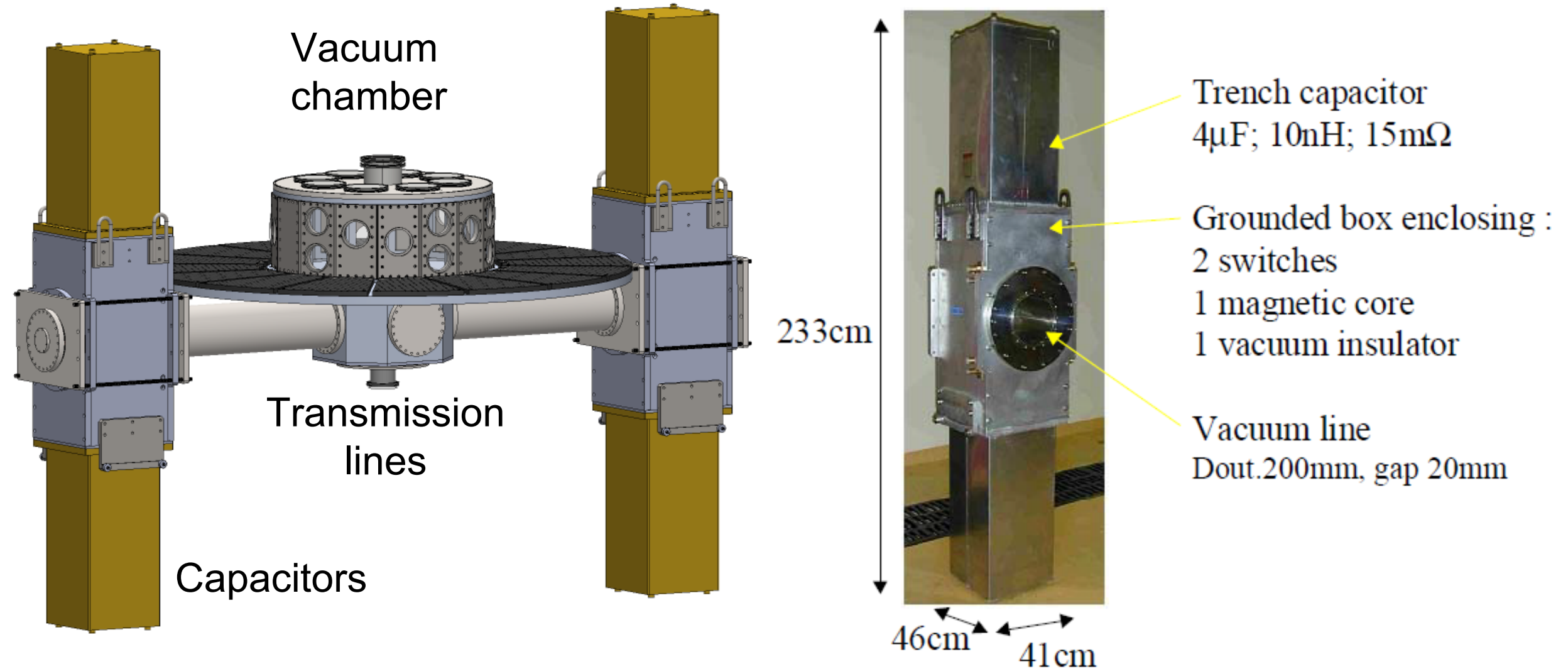


Open question:

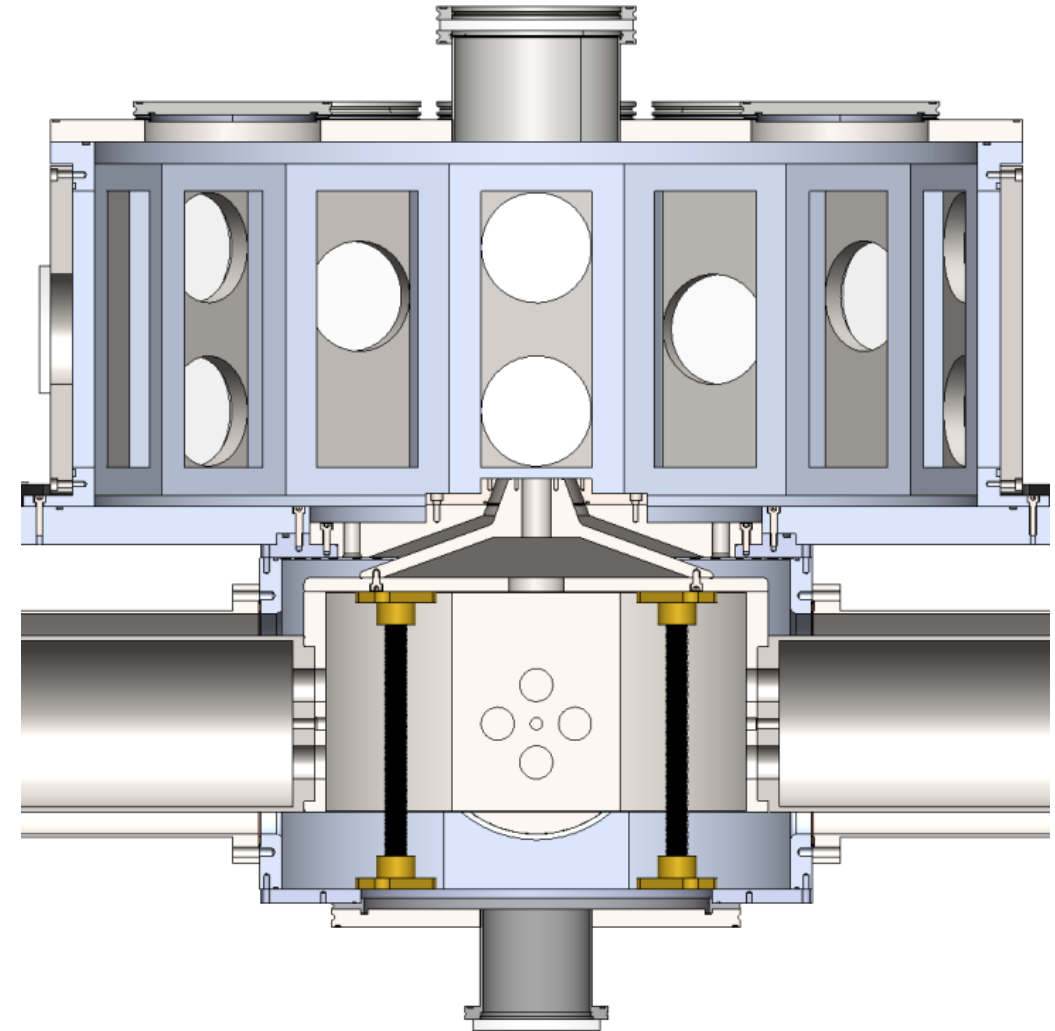
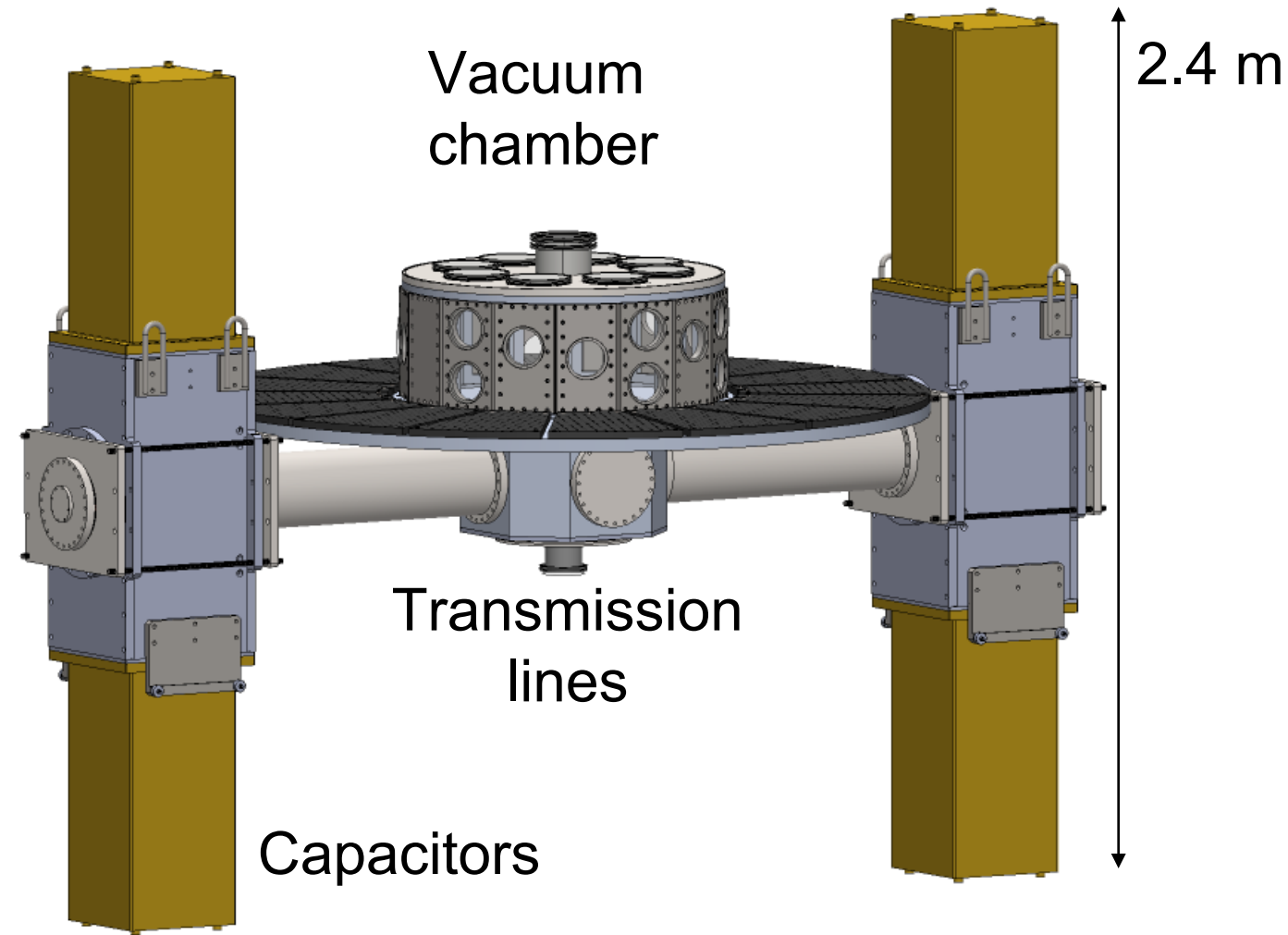
What is the long-term behaviour of this kink instability?



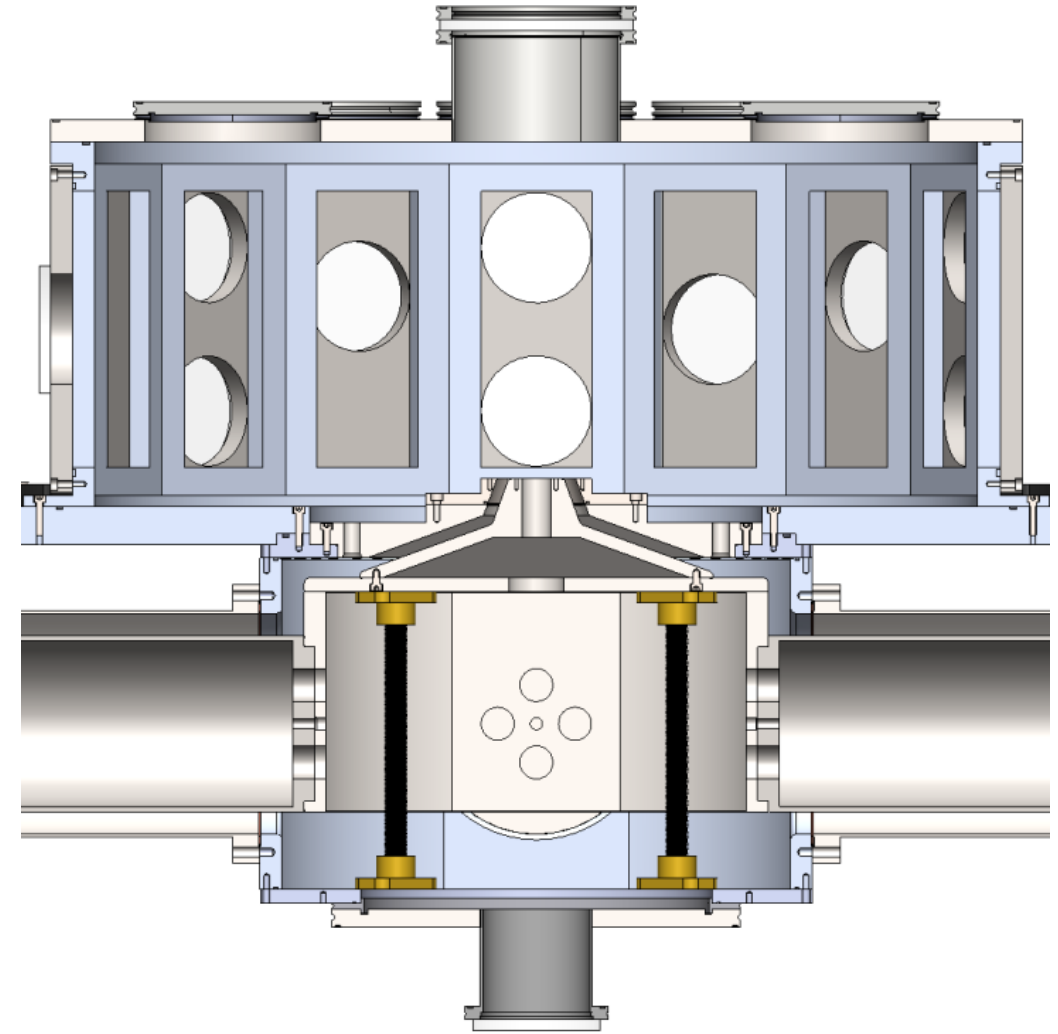
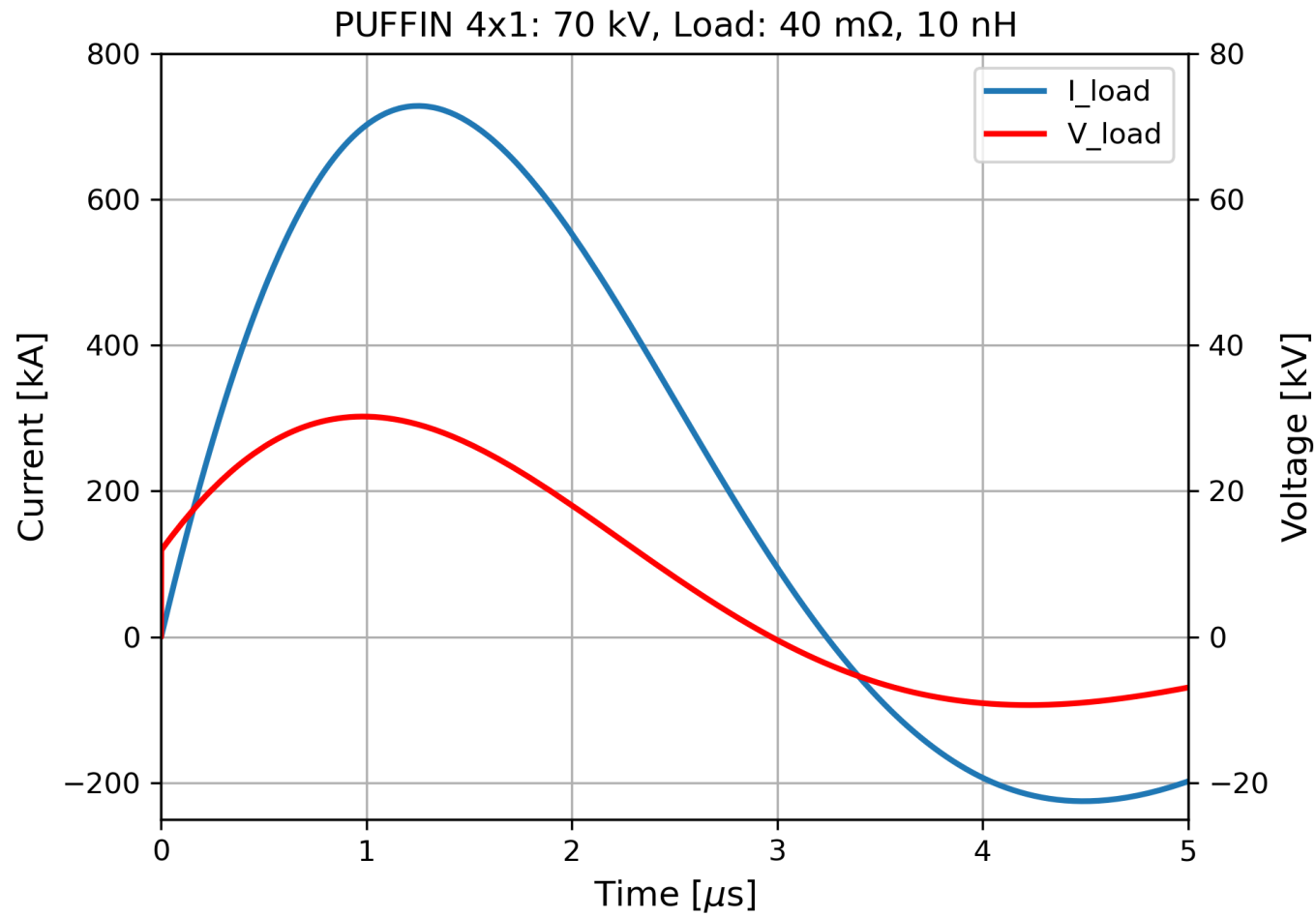
PUFFIN uses 2 LTD5 stages from SPHINX (CEA Gramat)



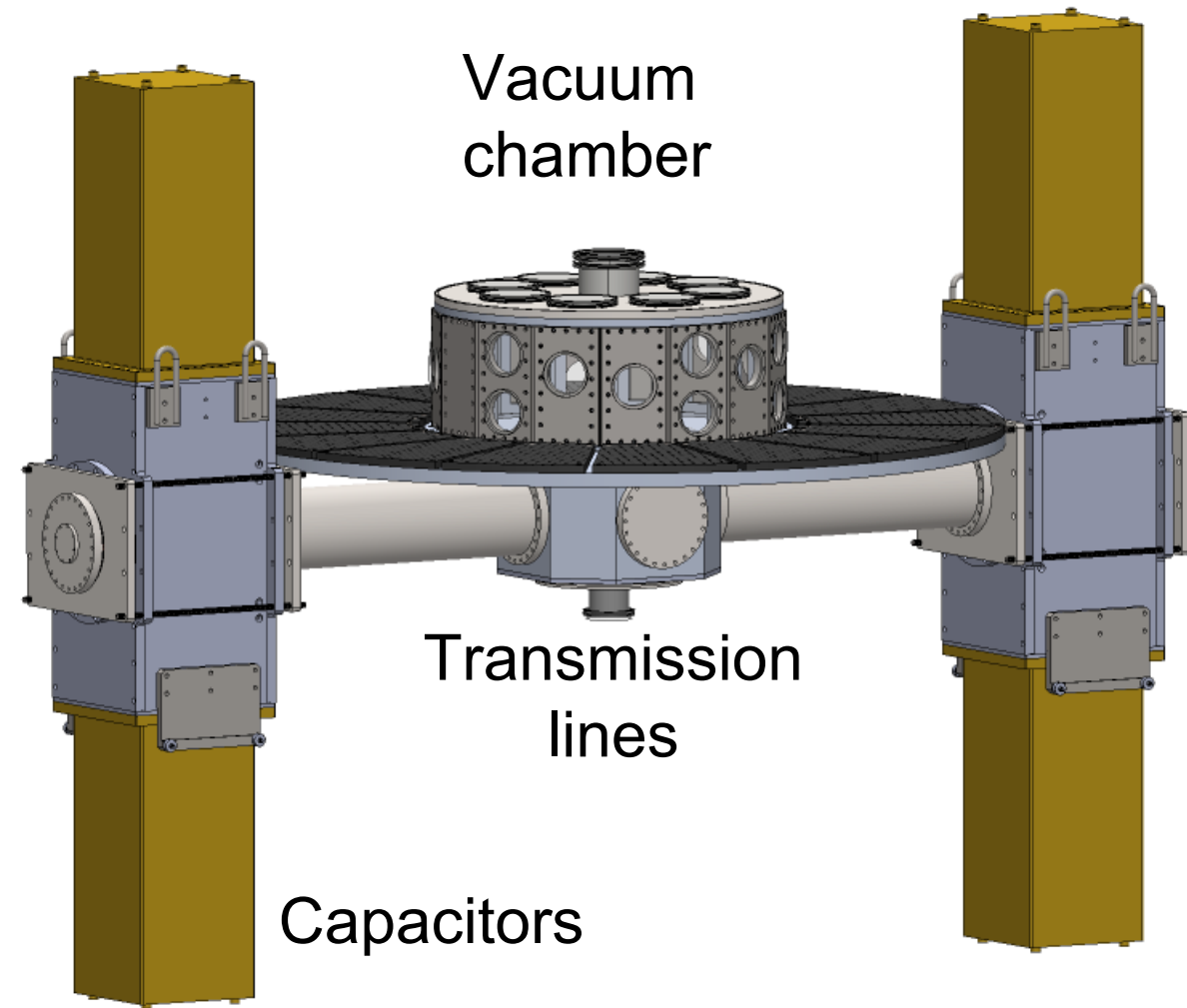
PUFFIN 2x1 will drive around 700 kA with a 1.6 μ s rise time



PUFFIN 2x1 will drive ~ 700 kA with a ~ 1.5 μ s rise time



PUFFIN: a long-pulse driver for fundamental physics



Timeline:

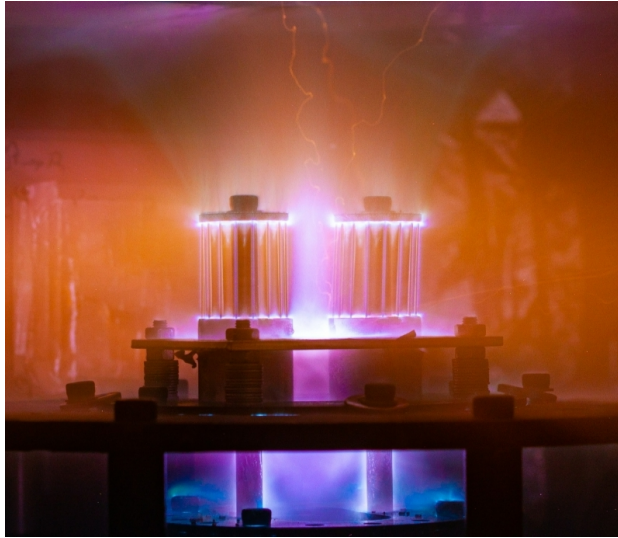
- LTD stages arrive May 2022
- Lab renovation finished December 2022
- Design and construction of power feed expected early 2023
- First experiments in 2023

Diagnostics:

- Laser imaging shadowgraphy, schlieren, interferometry and imaging refractometry
- XUV and optical framing cameras
- Thomson scattering (spatially resolved, ion and electron features)

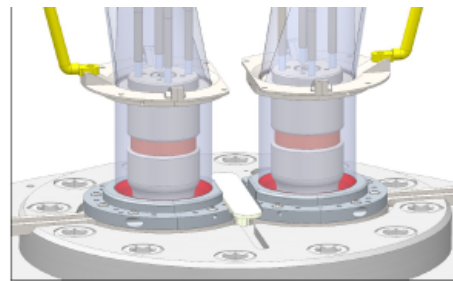
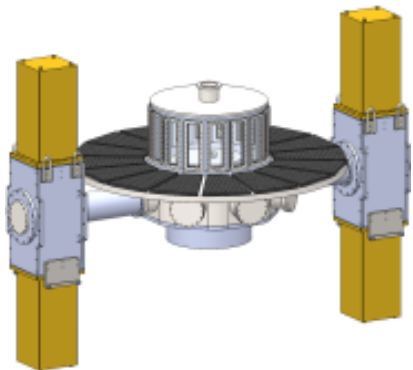


Magnetic reconnection

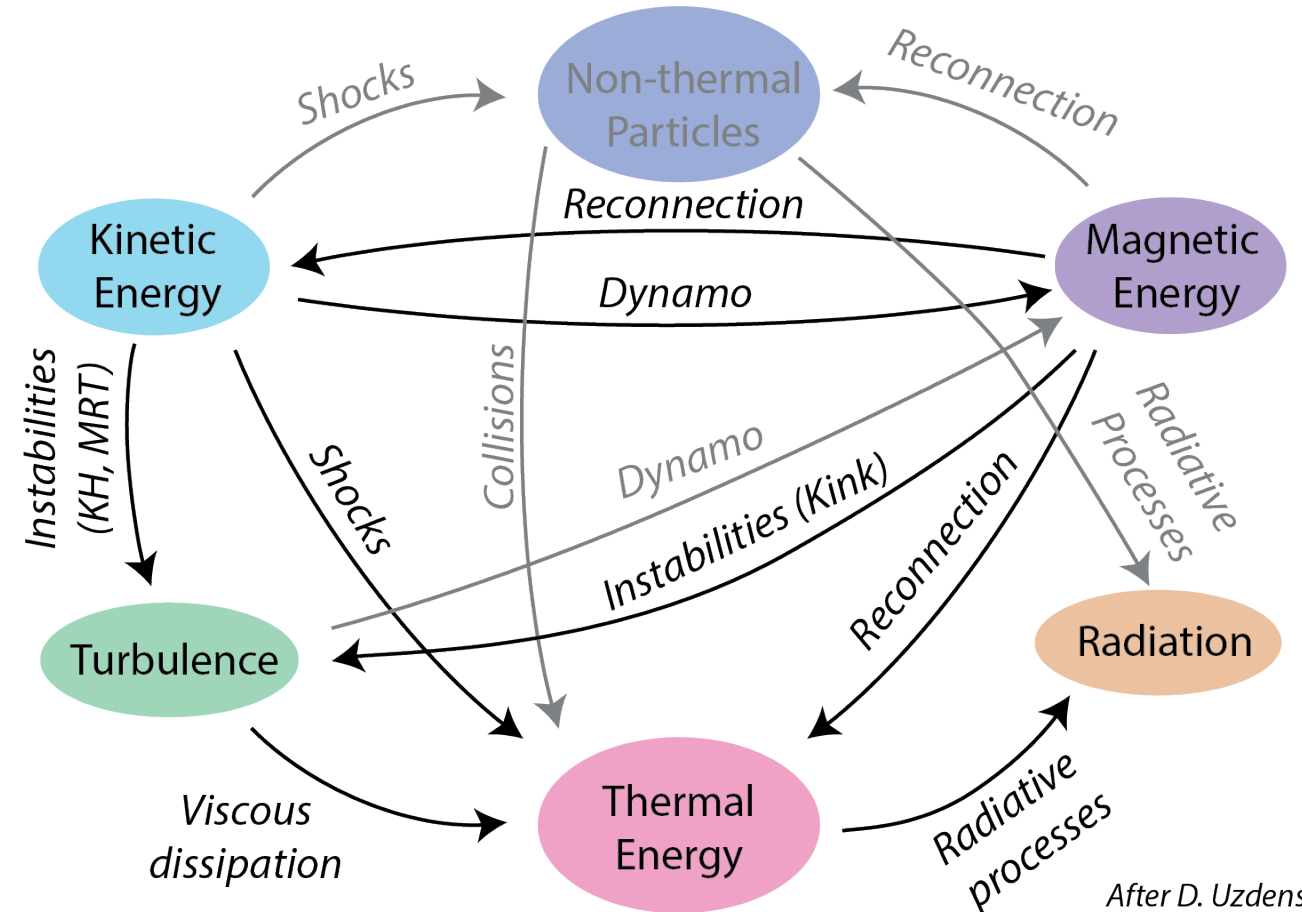


PUFFIN: $1.6 \mu\text{s}$,
700 kA at MIT

Radiatively cooled
reconnection on Z



Unifying theme: energy flows in plasmas



After D. Uzdensky