



# Molecular Dynamics of H Cottrell Atmosphere Formation in Al—Influence of Non-Dilute Concentration at Dislocation Cores



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## Hydrogen + Dislocations: How do they get along?



H is important fuel and proper storage is important, however it also leads to Hydrogen Embrittlement (HE)

To understand HE we must understand where H is going. Even with typically low H solubilities, enrichment near defects can result in HE.

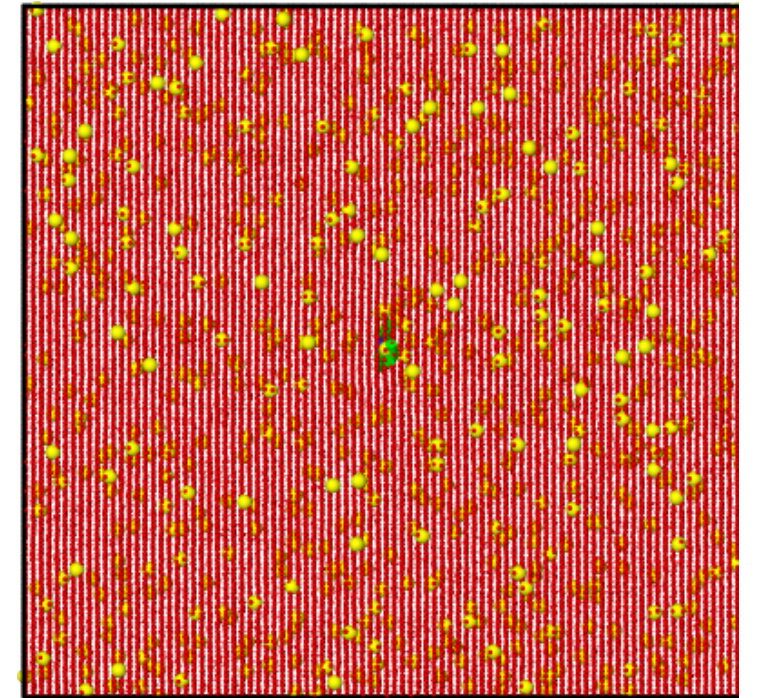
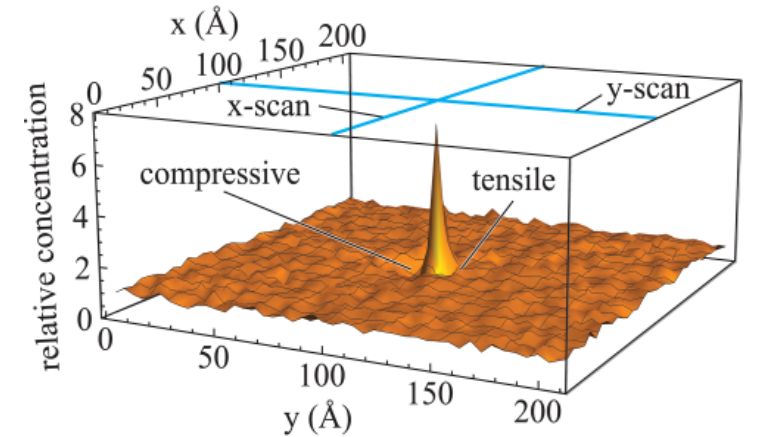
H-enhanced local plasticity(HELP) is a common theory used to explain the origin of HE that is concerned with Cottrell atmosphere.

To better understand HELP we focus on the formation of atmospheres in Al

# Outline



- MD simulation setup
- Transient formation of a H Cottrell atmosphere
- Compare and contrast to theoretical expressions
- Solute-Solute interactions at the core



## 4 Previous work

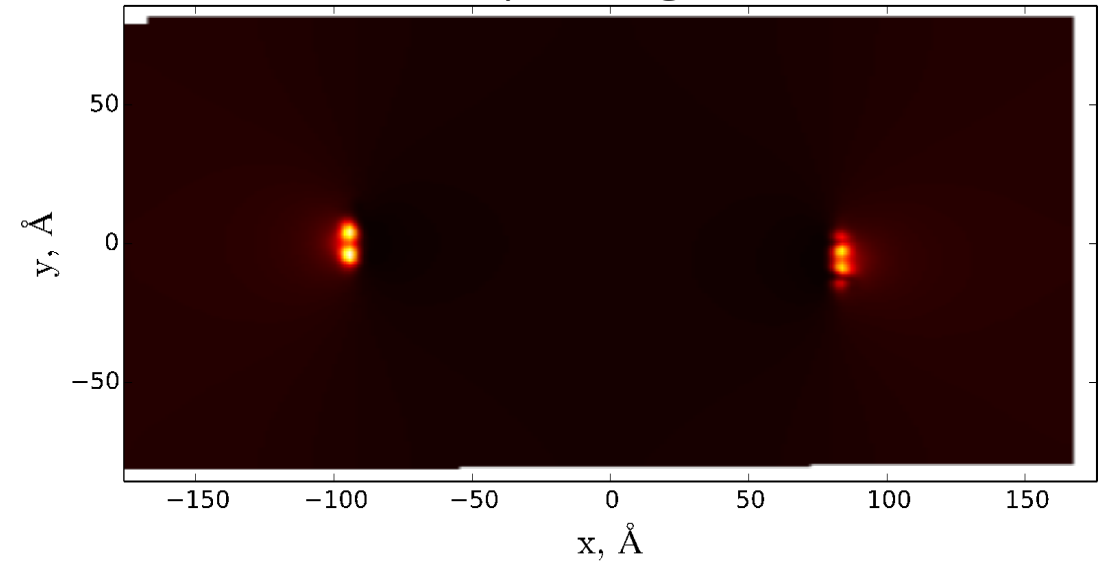
Molecular statics can calculate H insertion energy,  $E_{\text{ins}}$

Models and theories rely on  $E_{\text{ins}}$  to understand atmosphere behavior

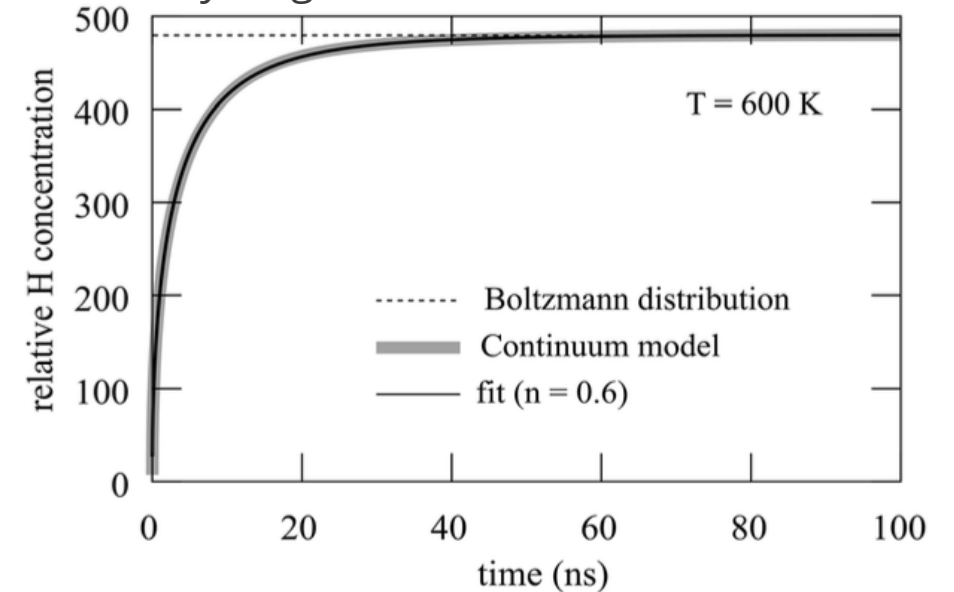
MD work with Cottrell atmospheres has been done but not on the transient formation

Understanding formation helps inform HELP

Boltzmann Heatmap for Edge Dislocation: Al-H



Peak Hydrogen Concentration Saturation





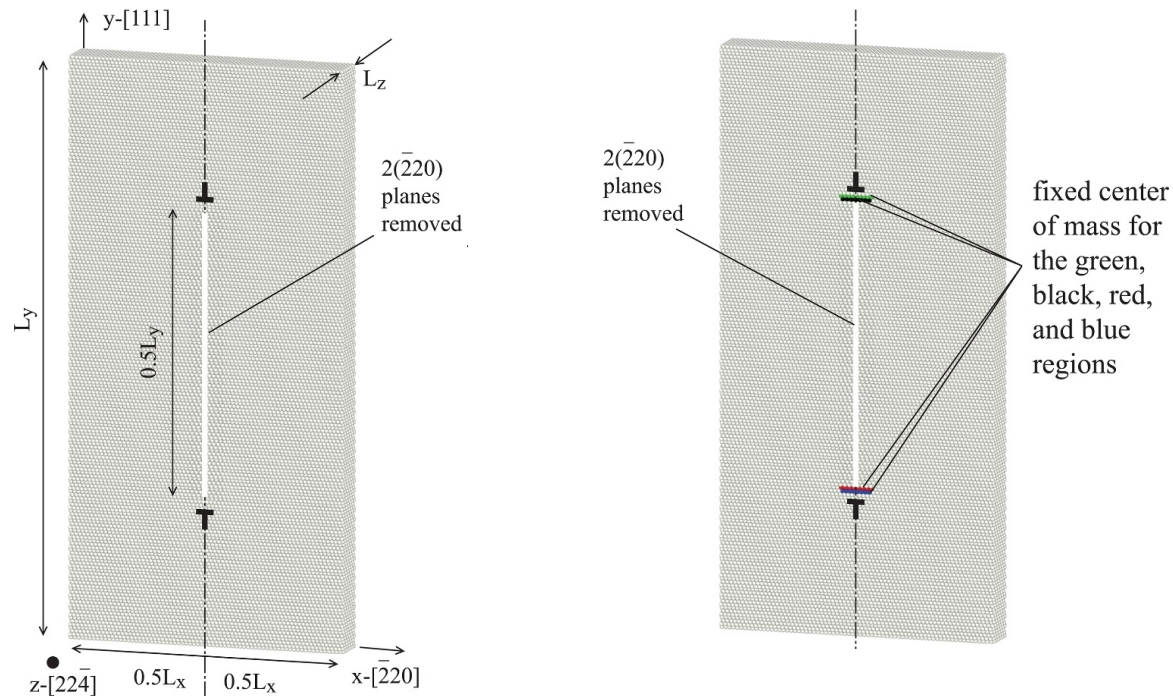
## 5

On MD time scales, 300K shows slow H diffusion

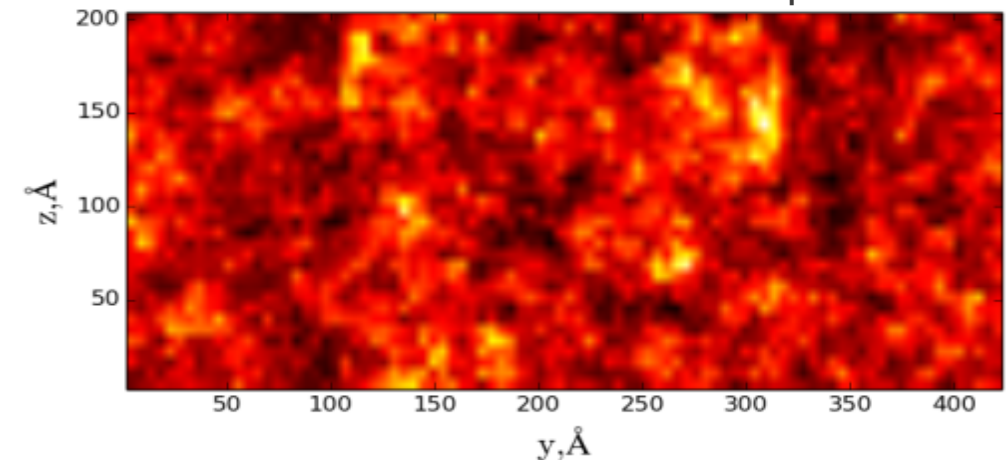
## Low Peierls stress materials edge bounces around

Dislocation needs to stay stationary.

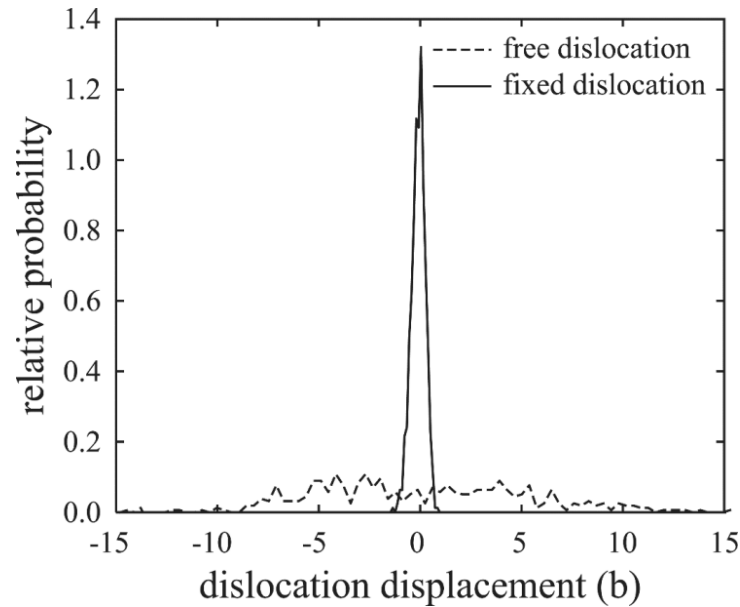
## Recenter the c.o.m. on partial planes to keep in place



## H Concentration Map

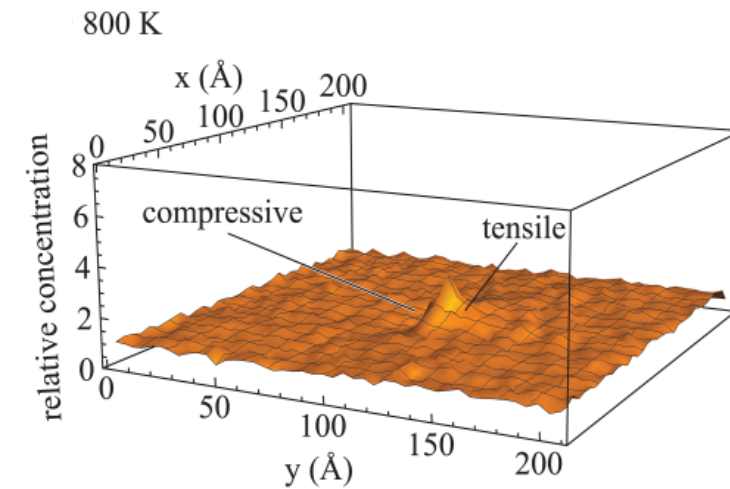
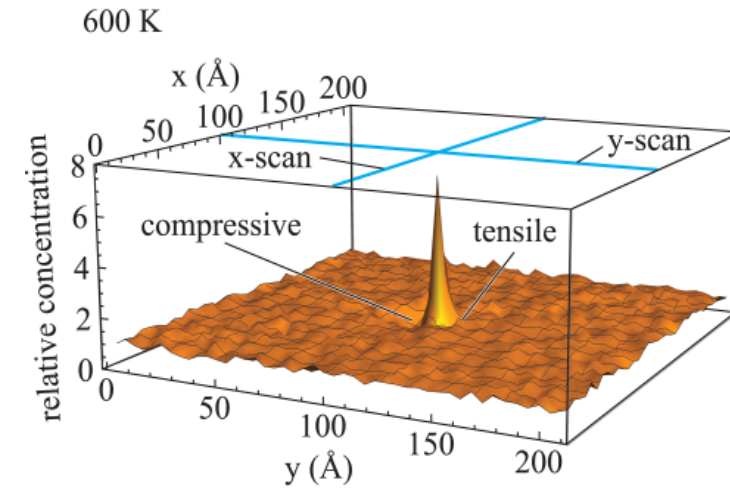
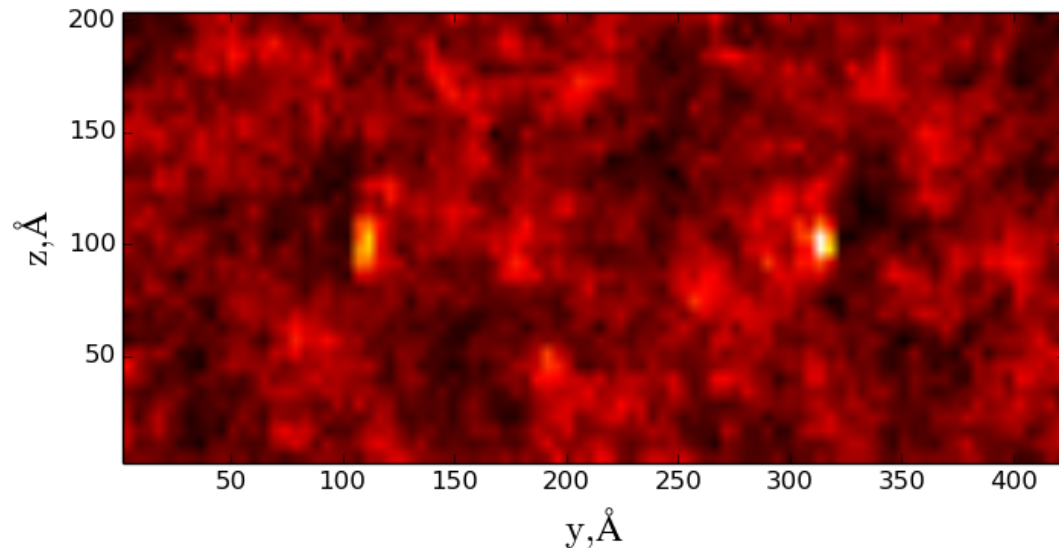


# MD mediated atmosphere formation



Dislocation remains in same spot

Shows characteristics of edge strain field



equilibrated for 2.5 ns

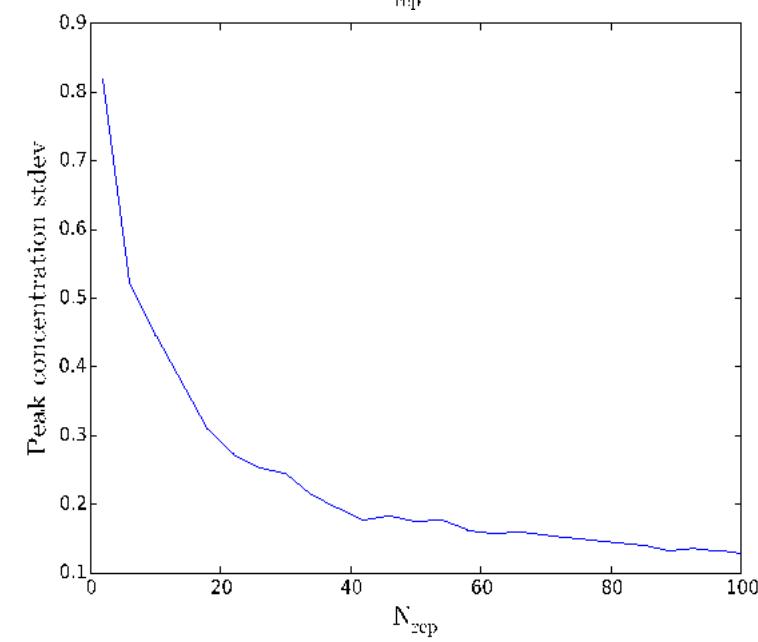
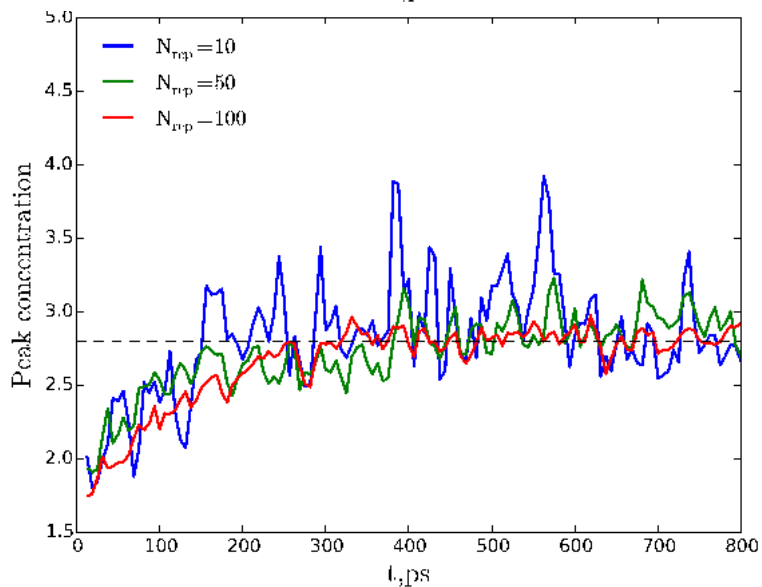
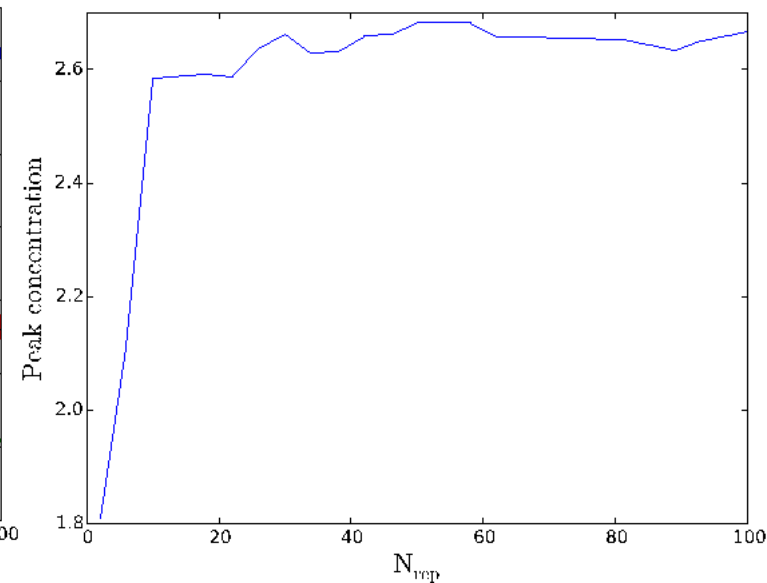
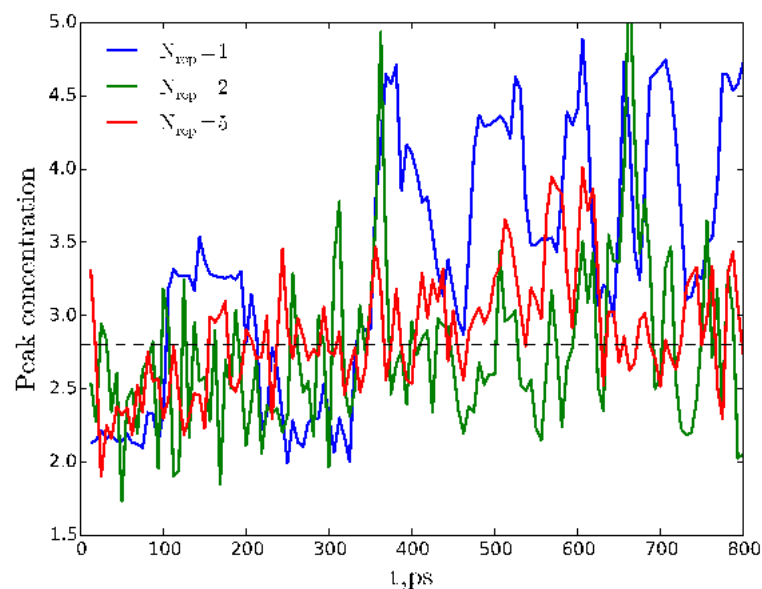
# Convergence Calculations



With more replicas the trend becomes apparent

With more replicas the curves smooths out

Alloying will increase the noise



# How does it match theory?

Fit to Cottrell theory shows good agreement

$$\hat{X}(t) = X_0 + (X_{\text{sat}} - X_0) \{1 - e^{-(t/t_{\text{cl}}^*)^{2/3}}\}$$

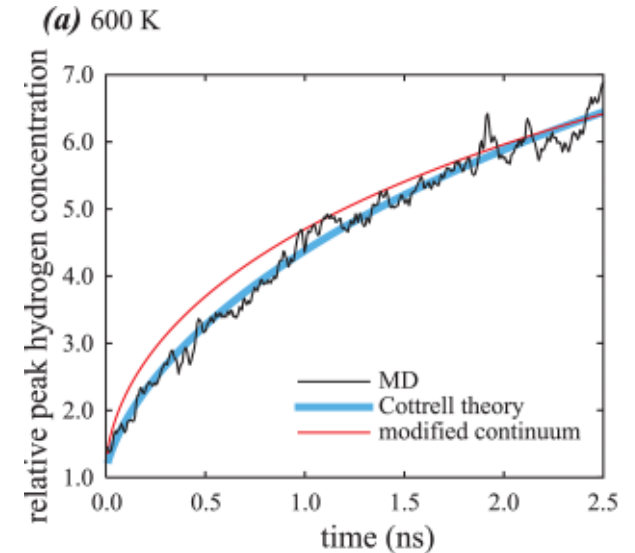
$$t_{\text{cl}}^* = \sqrt{2} b^2 k_B T / DW (X_{\text{sat}} / X_0)^{3/2}$$

$$\frac{t_{\text{cl}}^*(600\text{K})}{t_{\text{cl}}^*(800\text{K})} \sim 27$$

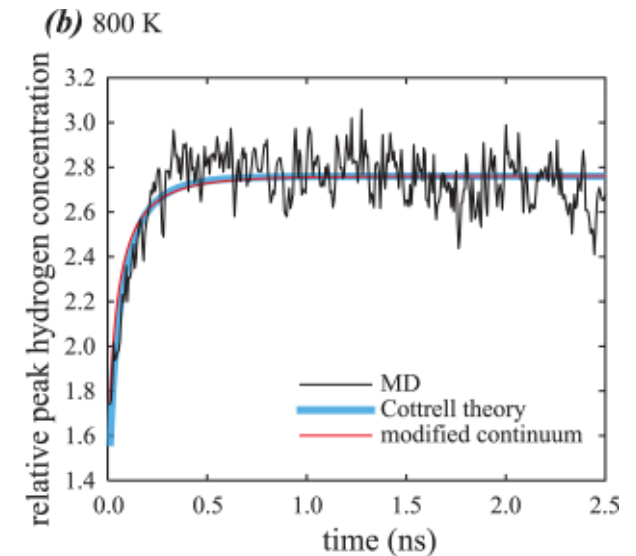
800 K plateaus ~0.25 ns → ~6.7 ns at 600K

Epperly and Sills theory accounts for finite atmosphere also shows good agreement

$X_0 = 1$   
 $X_{\text{sat}}(800\text{K}) \approx 2.8$   
 $D \approx 2.73 \text{ nm}^2/\text{ns}$   
 $W \approx 0.2 \text{ eV}$   
 $t_{\text{cl}}^*(800\text{K}) = 0.0684 \text{ ns}$



600K fit data with  $X_{\text{sat}}$   
as a free param  
 $X_{\text{sat}}(600\text{K}) = 8.3$   
 $t_{\text{cl}}^*(600\text{K}) = 1.830 \text{ ns}$



$\hat{X}(t)$  = peak concentration  
 $X_{\text{sat}}$  = saturation concentration  
 $X_0$  = background concentration  
 $t$  = time  
 $b$  = Burgers magnitude  
 $k_B$  = Boltz const  
 $T$  = temp  
 $D$  = diffusivity  
 $W = |\min_{\mathbf{x}} p_d(\mathbf{x}) \Delta V|$



# Theoretical Overestimation of Peak Concentrations



The peaks estimated by MD  
are much lower than predicted

$E_{\text{ins}}$  must be less favorable  
near dislocation

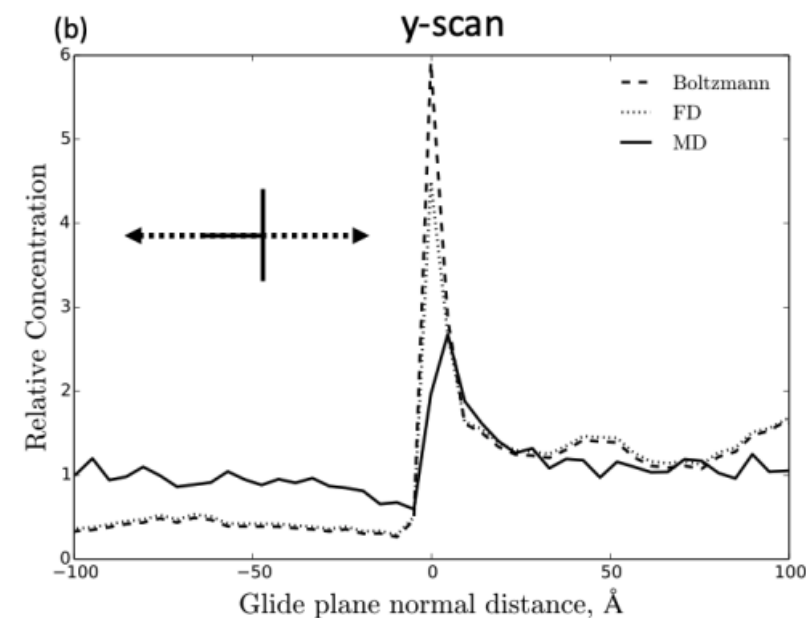
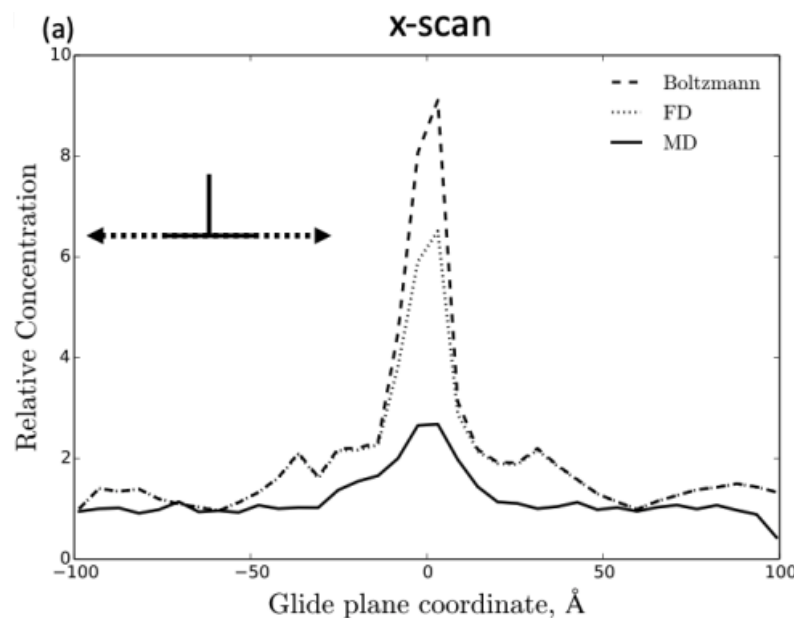
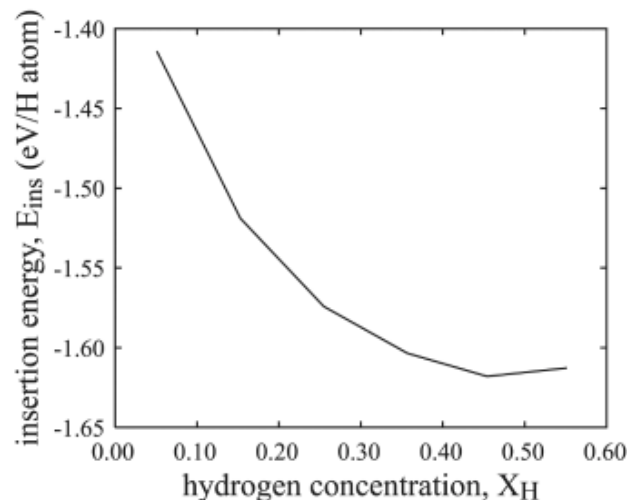
Boltzmann:

$$p_i = e^{-\frac{E_i}{k_b T}}$$

Fermi-Dirac:

$$n_i = \frac{1}{e^{-(E_i - \mu)/k_b T} + 1}$$

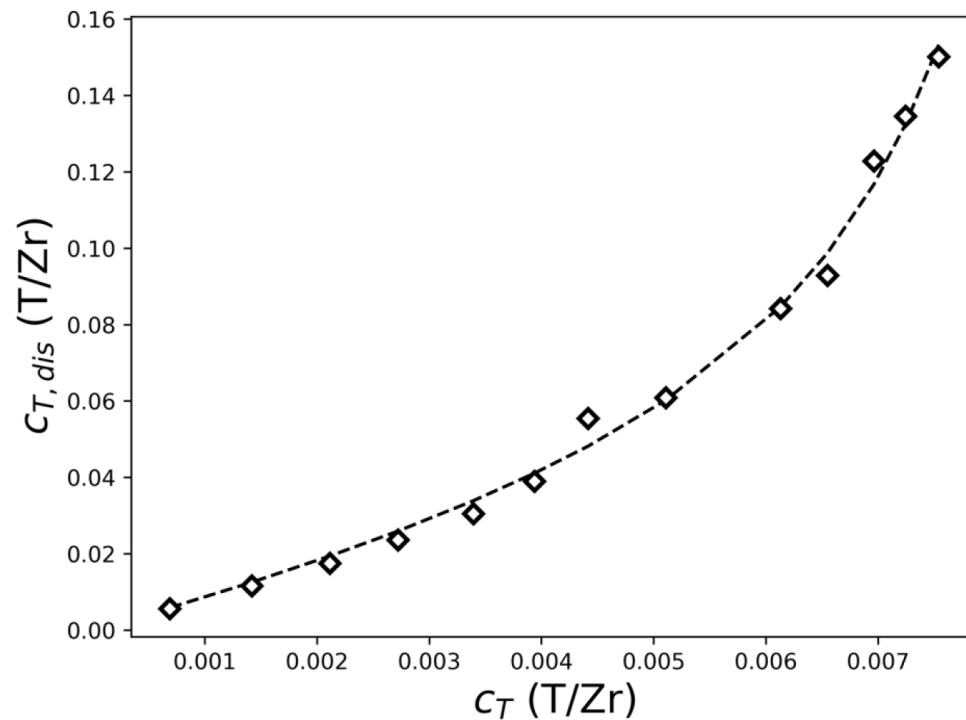
Bulk Al  $E_{\text{ins}}$  more  
attractive with more H



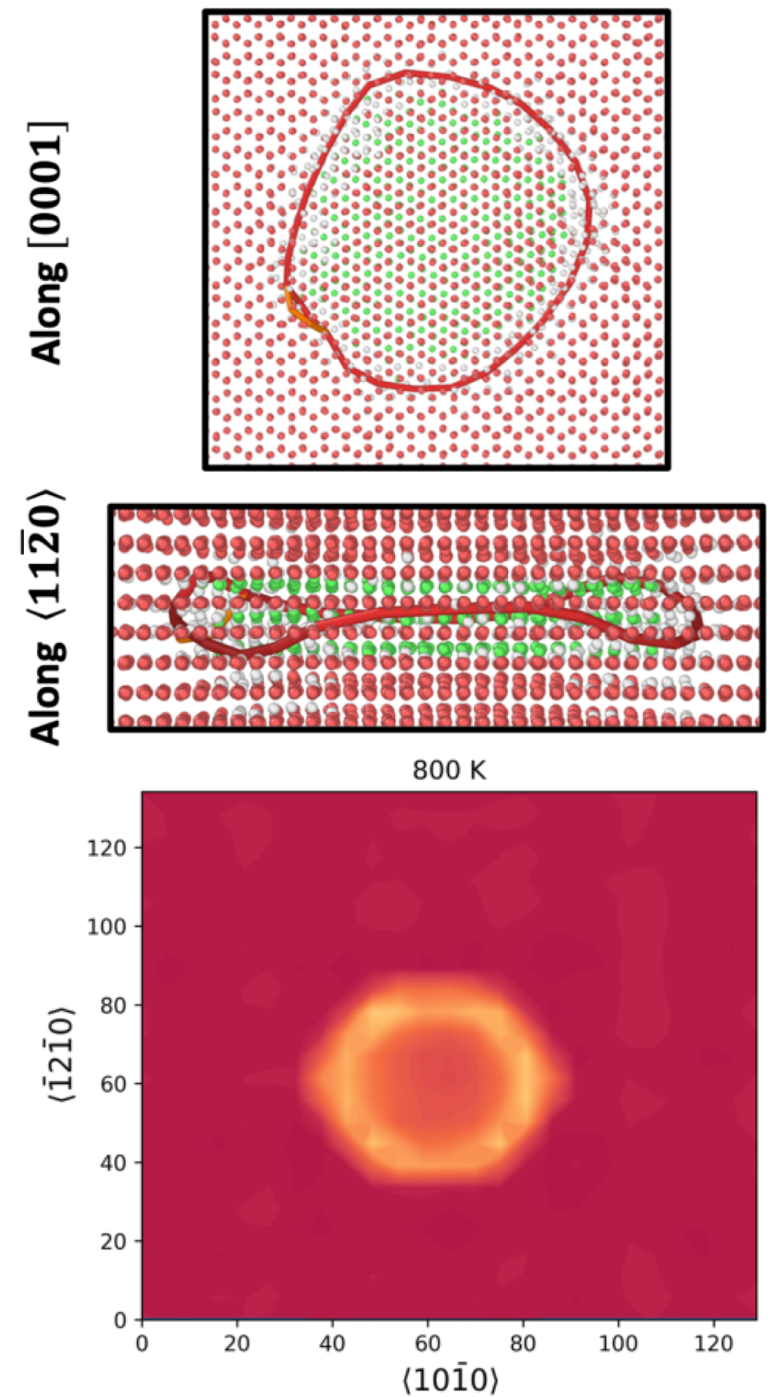
# Solute-Solute Interaction in Zr-T\*

T-T interactions in HCP Zr also show a solute-solute effect

With more T loading enrichment increases at vac-type dislocation loop.



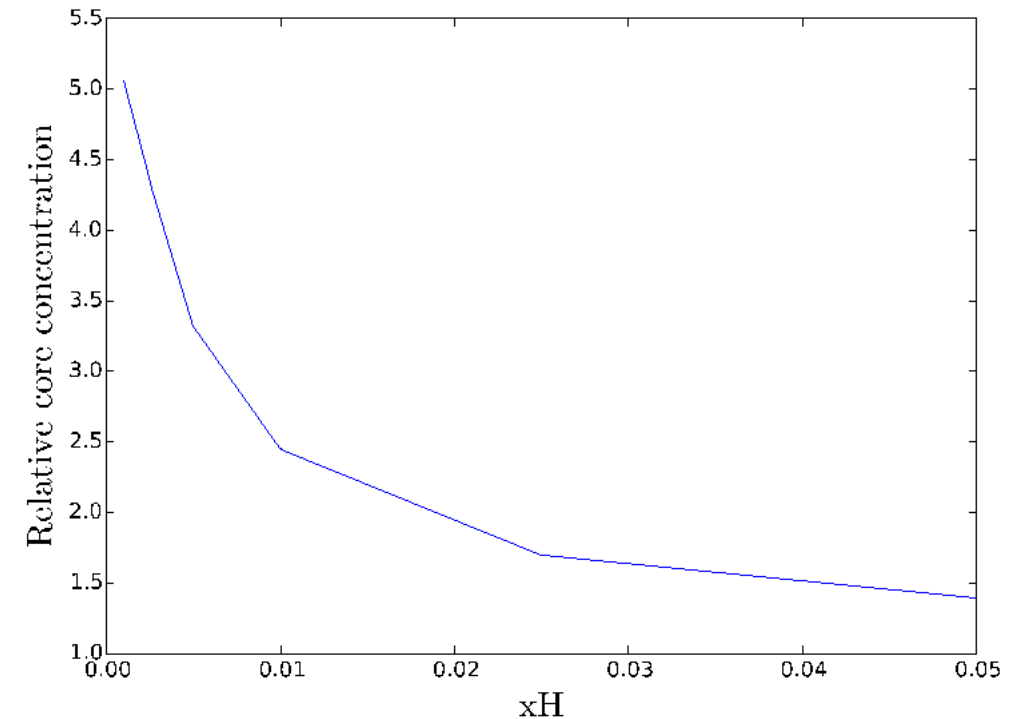
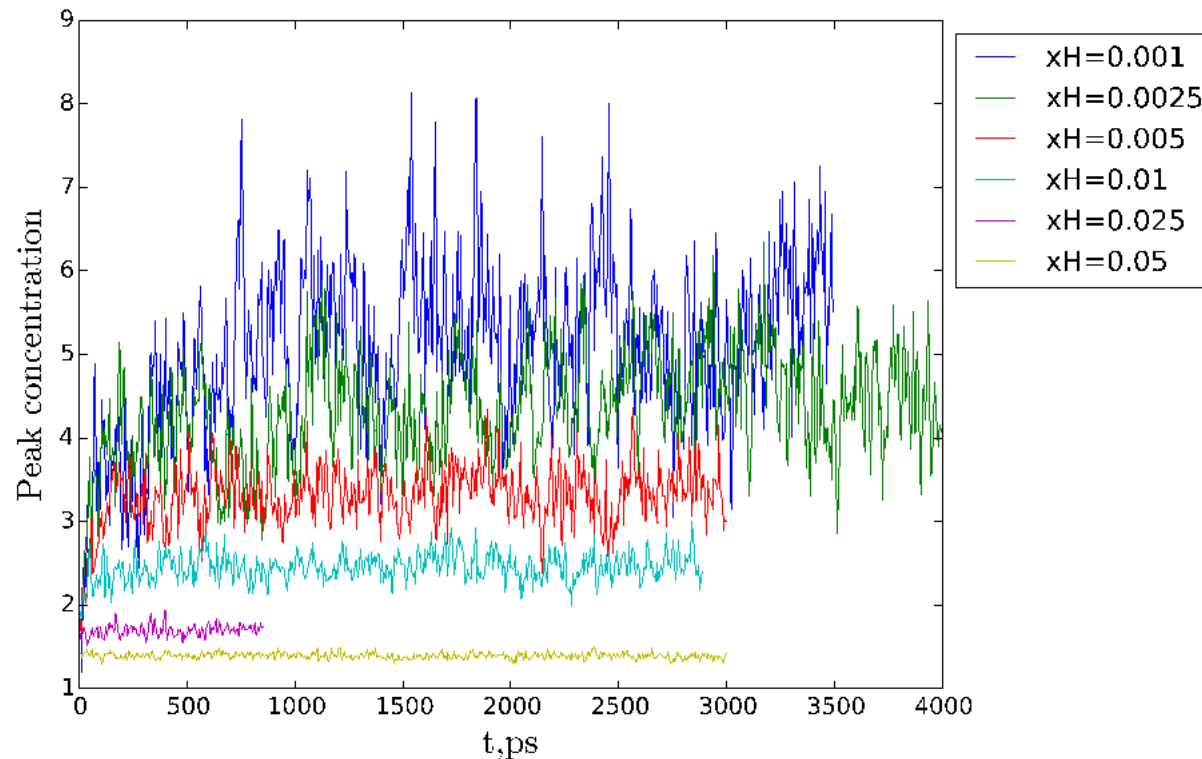
Fitting to equations reveals a strong T-T interaction.



# Solute-Solute Interaction in Ni-H



- This phenomenon is observed in systems outside of Al-H
- Ni-H systems with variable loading of H show different enrichment at the core

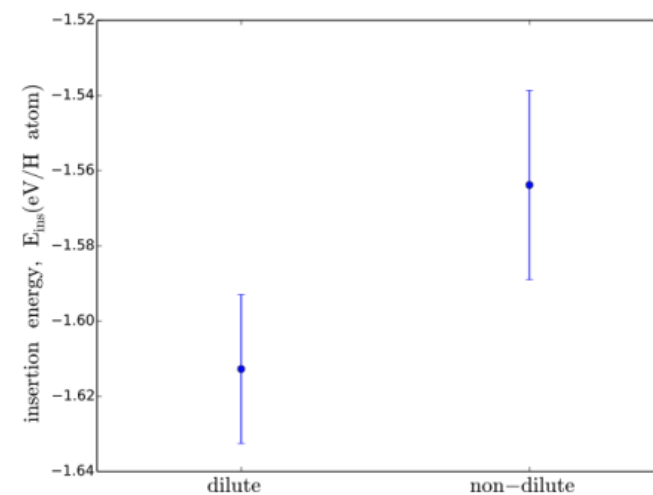
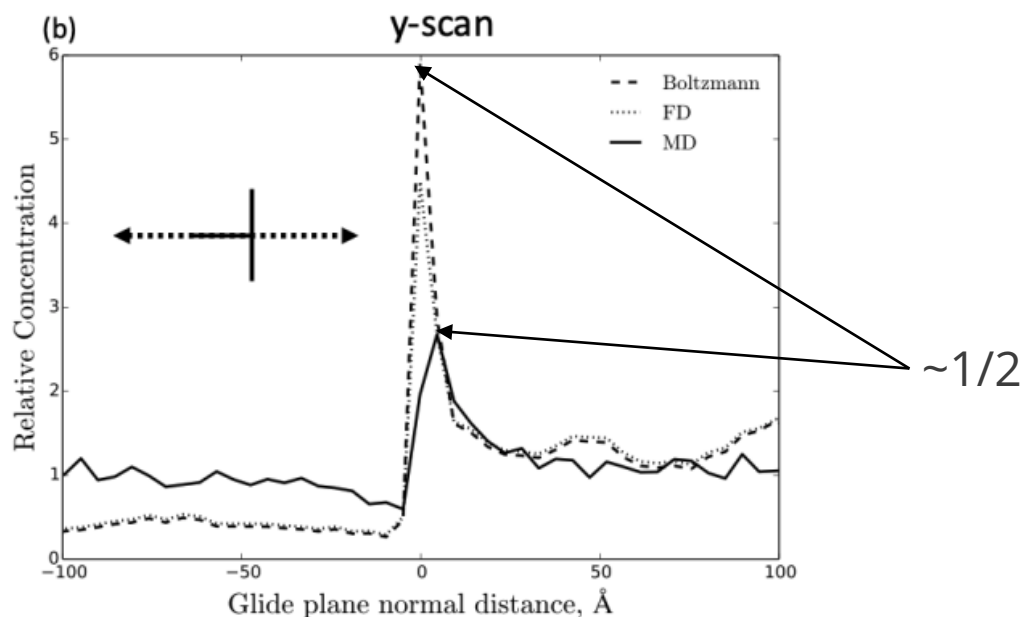
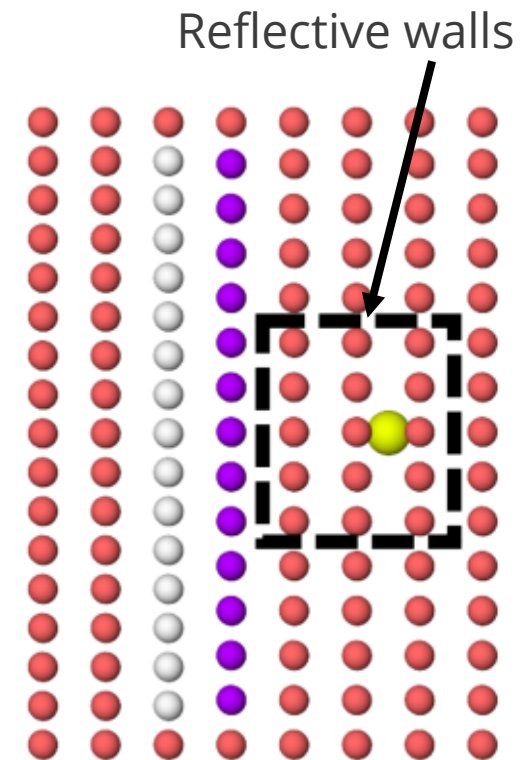


# Measuring Solute-Solute Interactions

To measure the effects of having nearby H, use reflective walls

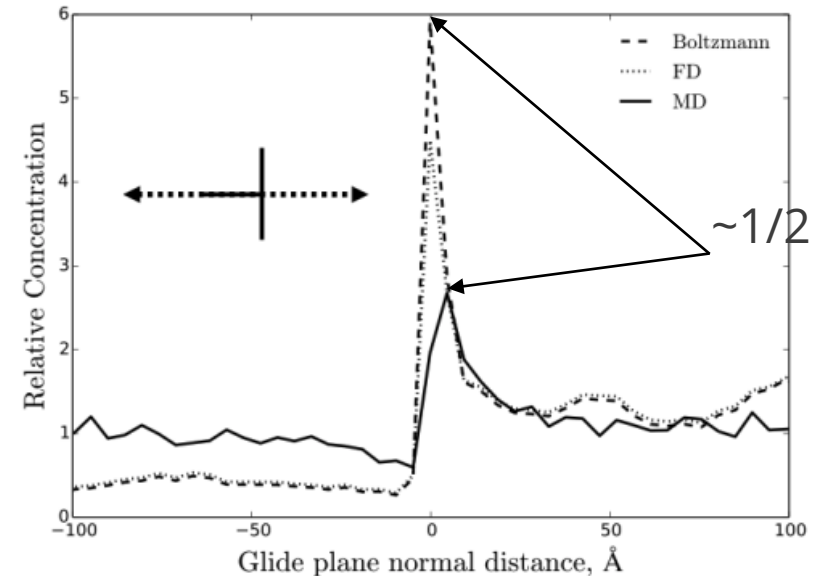
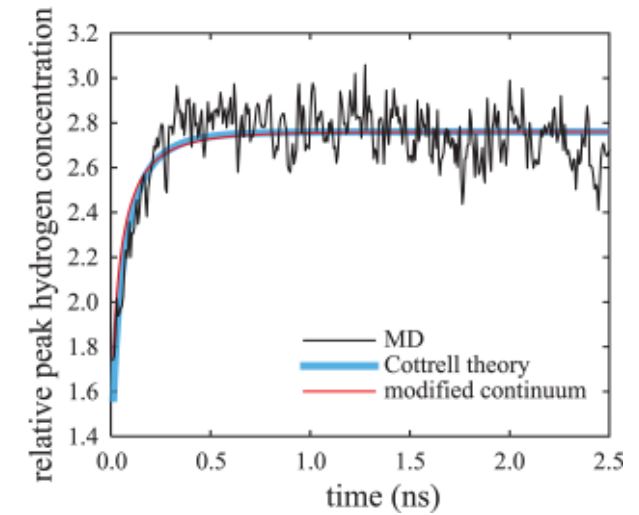
Time-average simulation with 1 (dilute), or 2(xH~.01) H in the box

A change of 50meV leads to the enrichment at the dislocation of ~1/2 the dilute enrichment



# Conclusions

- Simulated the transient formation of hydrogen Cottrell atmospheres
- Compared to theoretical expressions
- Calculated the solute-solute interactions near a dislocation
- Solute-Solute Interactions should be considered





# Acknowledgements\*

\*Due to work from home, hair/facial hair may have changed



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Richard Skelton (SNL)



Chris San Marchi (SNL)



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**Thanks for your time,  
questions?**



# Extra Slides



# Running average to Instantaneous value



$$\overline{X_H}(x, y, t) = \frac{1}{t} \int_0^t X_H(x, y, t) dt$$

$$X_H(x, y, t) = \frac{\partial [t \cdot \overline{X_H}(x, y, t)]}{\partial t}$$

$$X_{H,exp}(x, y, t) = C_H X_H(x, y, t)$$

$$\frac{\partial f_i}{\partial t} = \begin{cases} d_{i,1,0}, i = 2 \\ d_{i,1,-1}, i = 3 \\ \frac{15d_{i,1,-1} - 6d_{i,2,-2} + d_{i,3,-3}}{10}, 4 \leq i \leq N - 2 \\ \frac{4d_{i,1,-1} - d_{i,2,-2}}{3}, i = N - 1 \\ d_{i,1,-1}, i = N \\ d_{i,0,-1}, i = N + 1 \end{cases}$$

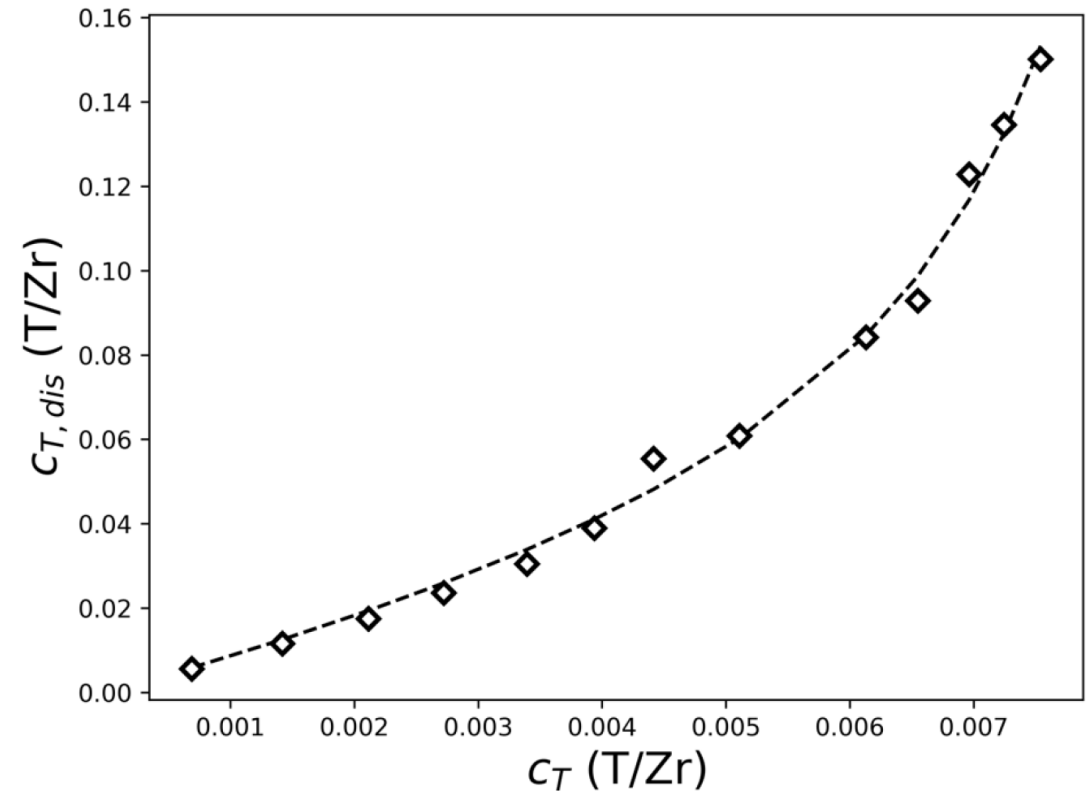
# Zr-T Solute-Solute Interactions



$$E(c_{T,dis}) = E_0 + c_{T,dis}E_{TT}$$

$$c_{T,dis}/(1 - c_{T,dis}) = c_T/(1 - c_T) \exp \left[ -\frac{E_0 + c_{T,dis}E_{TT}}{k_B T} \right]$$

$$E_0 = -0.127 \pm 0.002 \text{ eV and } E_{TT} = -0.363 \pm 0.024 \text{ eV}$$





# Example of out-of-date Pictures

