



Re-analysis of plasma temperature and density for stellar iron opacity experiments

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Summary: Plasma temperature and density analysis are refined for Sandia iron opacity experiments



Motivation: Fe opacity measured at solar interior temperature disagree with calculated opacity

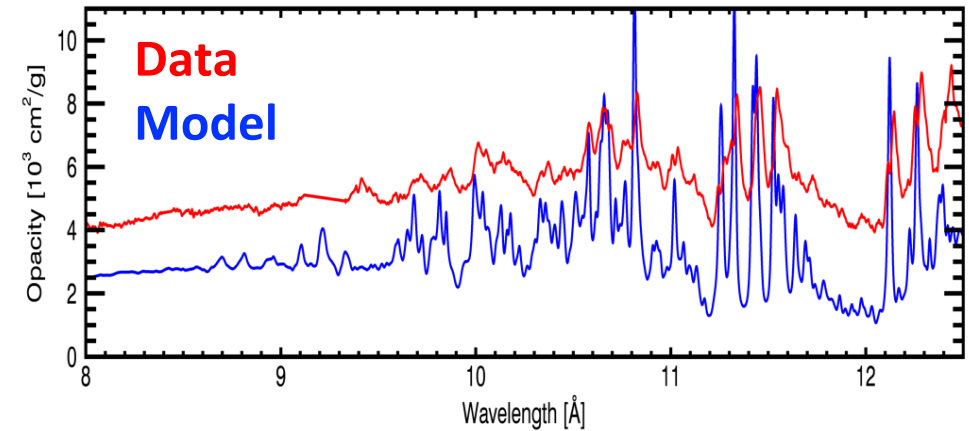
→ Are temperature (T_e) and density (n_e) accurate?

What we did: Refined temperature and density analysis

1. Line-shape calculations
2. Background determination
3. T_e , n_e analysis method

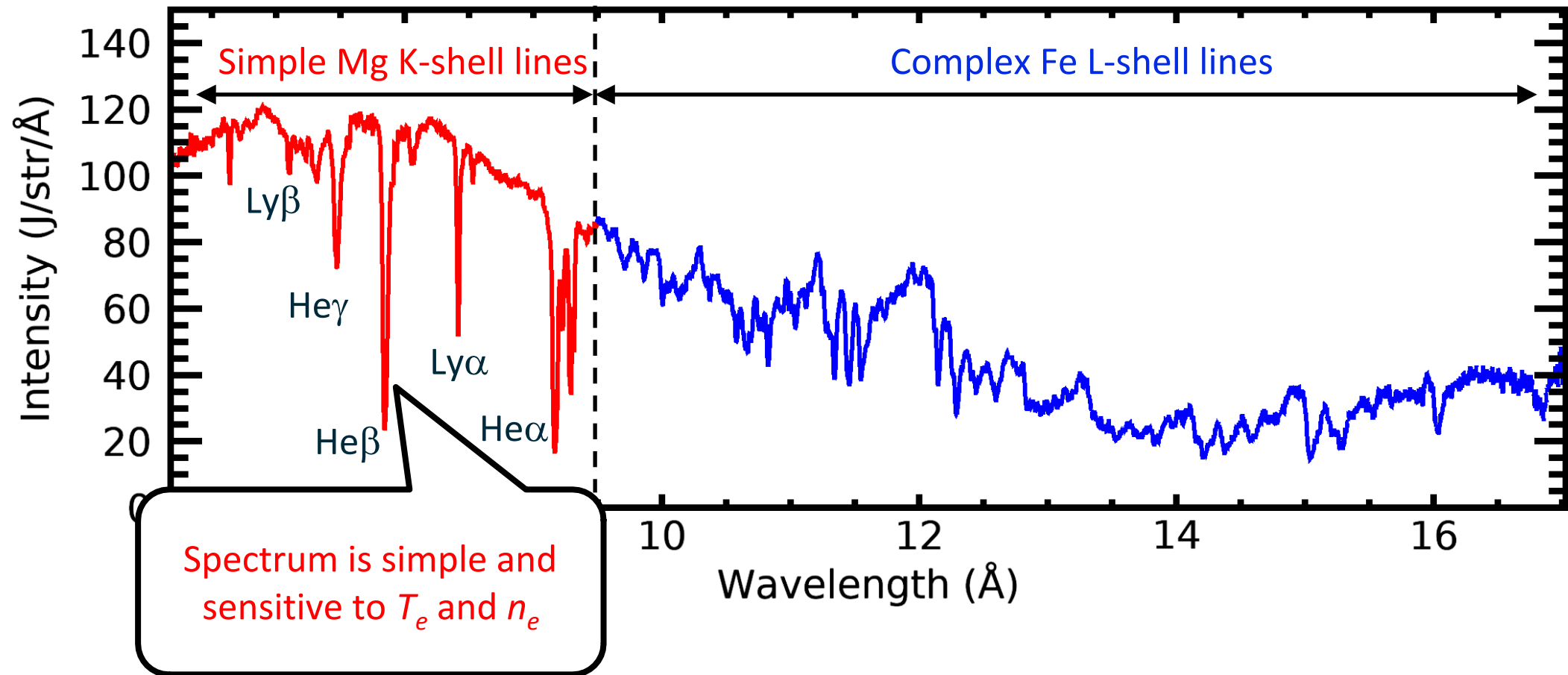
Result:

- New line shape increases T_e and n_e by 2% and 20%
- New analysis incorporates inconsistencies into uncertainties



Change in temperature and density is +1% and +26%, respectively, which are small to explain the previously reported Fe model-data discrepancy

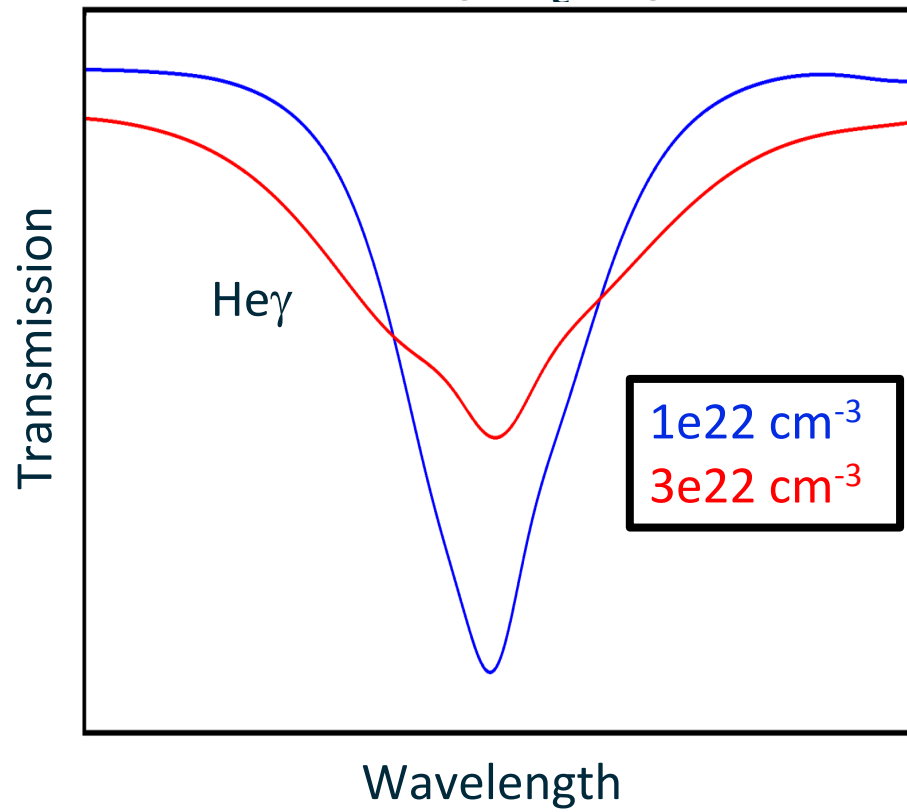
Temperature and density is analyzed by mixing Mg into the Fe opacity sample and analyzing its simple spectra



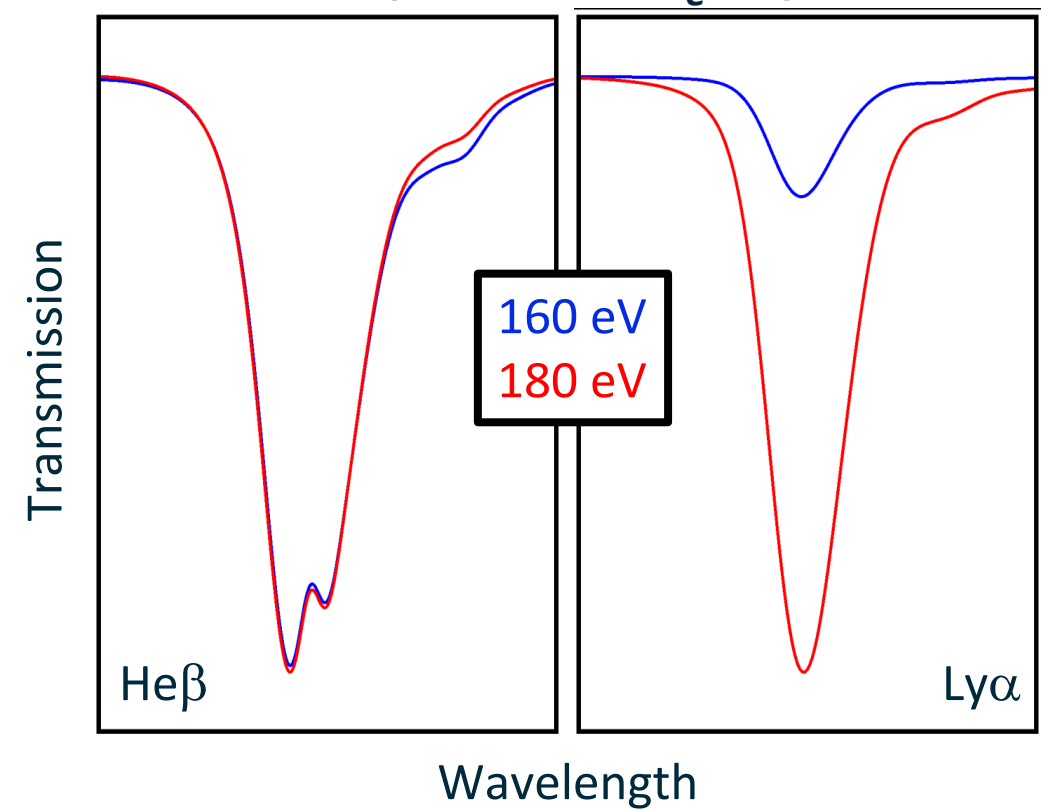
Electron density (n_e) and temperature (T_e) are inferred from Mg line broadenings and line ratios



Electron density (n_e) by line widths



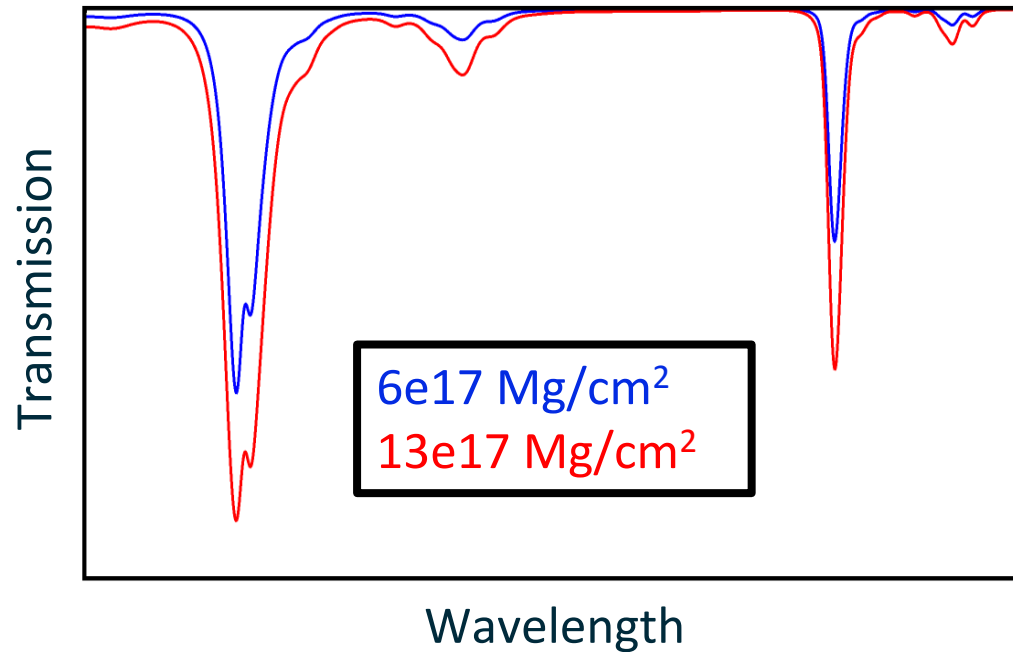
Electron temperature (T_e) by line ratios



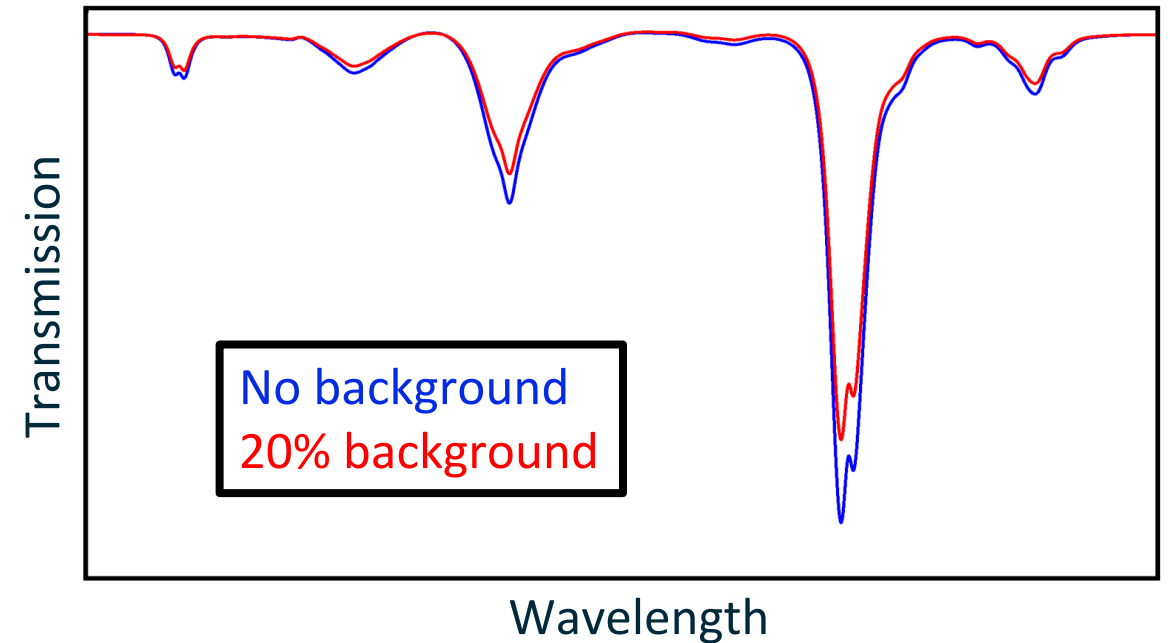
Mg areal density and residual background can be inferred from line depths and saturation on the strong lines



Mg areal density (ρL) by line depths



Background (b) by line saturation

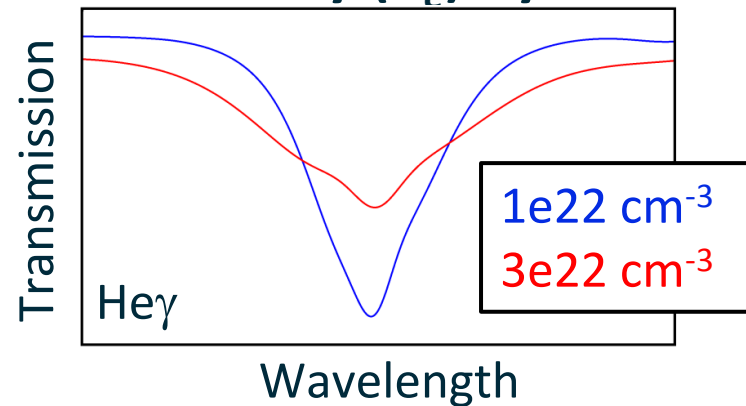


Tracer Mg spectroscopy can constrain n_e , T_e , ρL , and *background* simultaneously

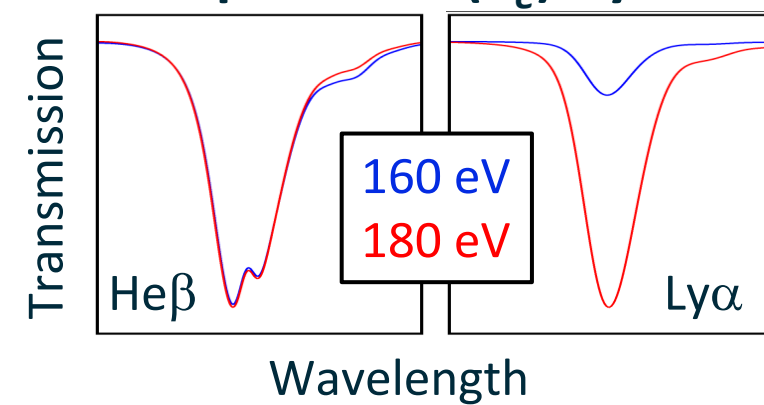


Mg is mixed into Fe opacity sample for plasma-diagnostic purposes

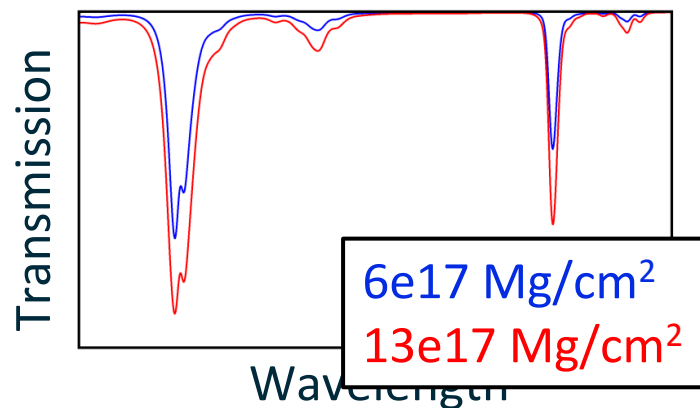
Electron density (n_e) by line widths



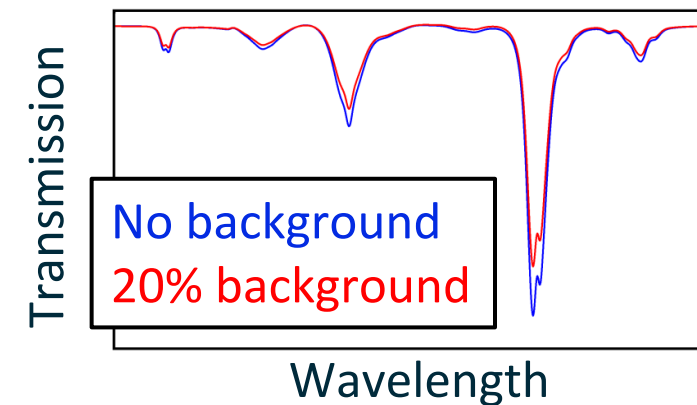
Electron temperature (T_e) by line ratios



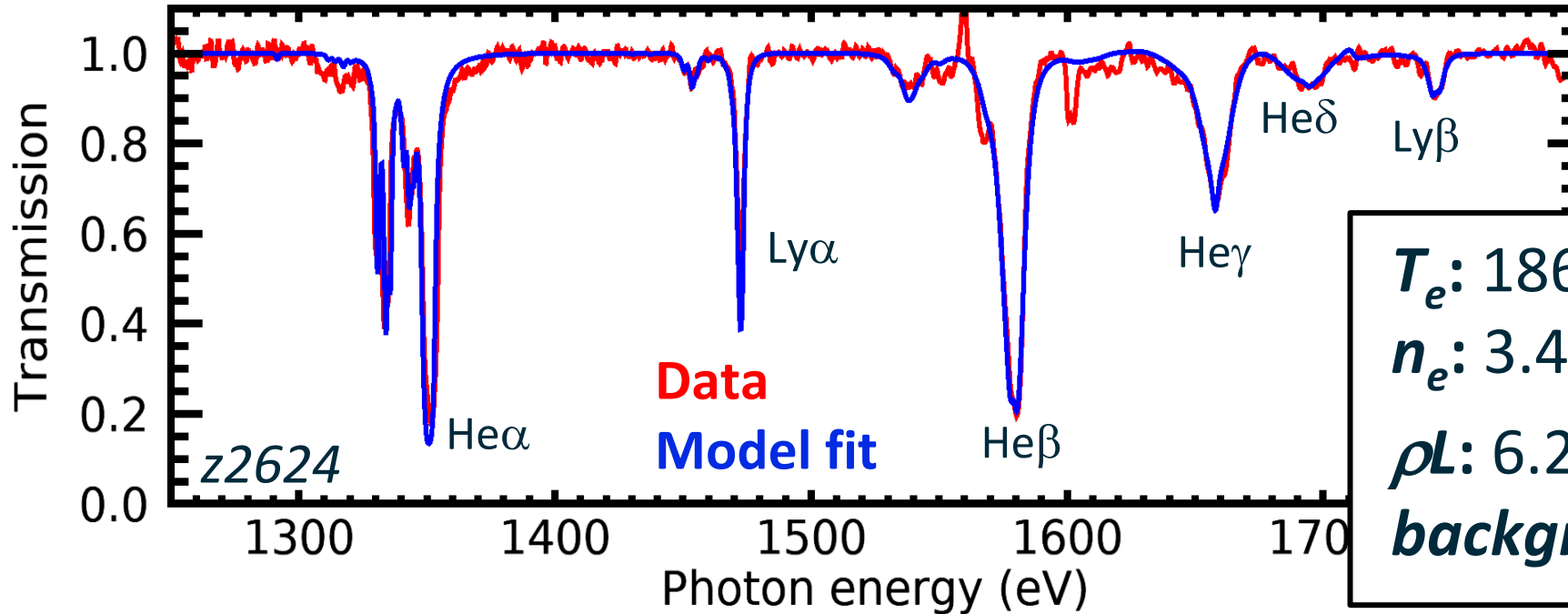
Mg areal density (ρL) by line depths



Background (b) by line saturation



n_e , T_e , ρL , and *background* are precisely determined by simultaneous fit to all the Mg lines



T_e : 186.07 eV ($\pm 0.4\%$)
 n_e : $3.49 \times 10^{22} \text{ cm}^{-3}$ ($\pm 1.7\%$)
 ρL : $6.27 \times 10^{-5} \text{ g/cm}^2$ ($\pm 2\%$)
background: 7.4 J/sr/\AA ($\pm 12\%$)

3 concerns:

1. Is line-shape calculation accurate? [1, 2]
2. Is background determination accurate?
3. Does the analysis capture **realistic uncertainties**?

T_e and n_e diagnostics are improved by refining (i) line-shape theory, (ii) background determination, and (iii) analysis method



Line-shape theory:

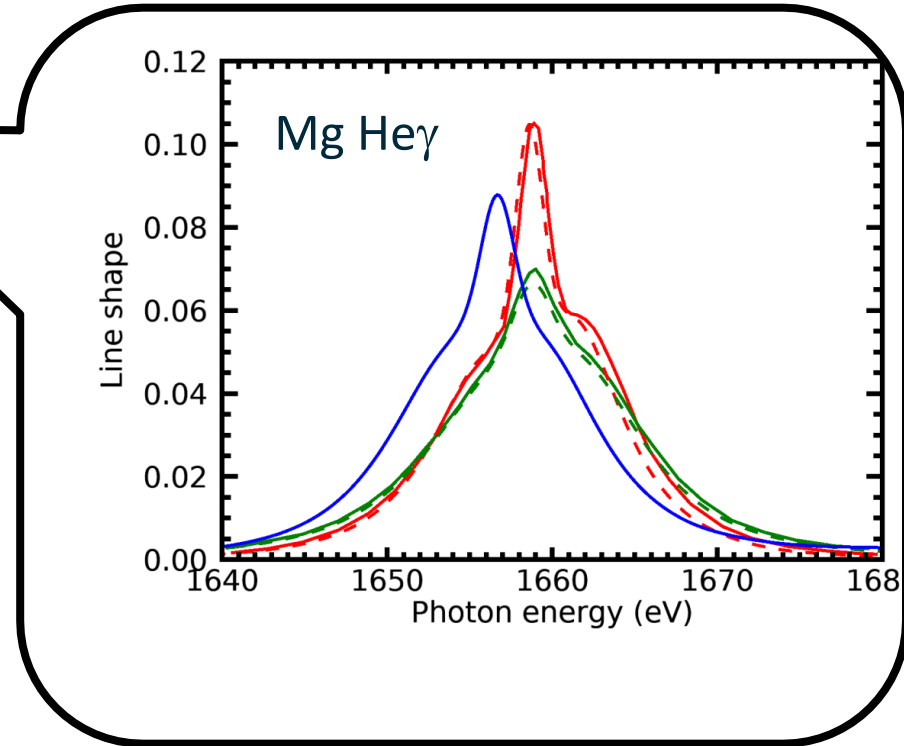
- Add missing physics called electron capture [1]
- Remove 3 common approximations [2]

Background determination [3]:

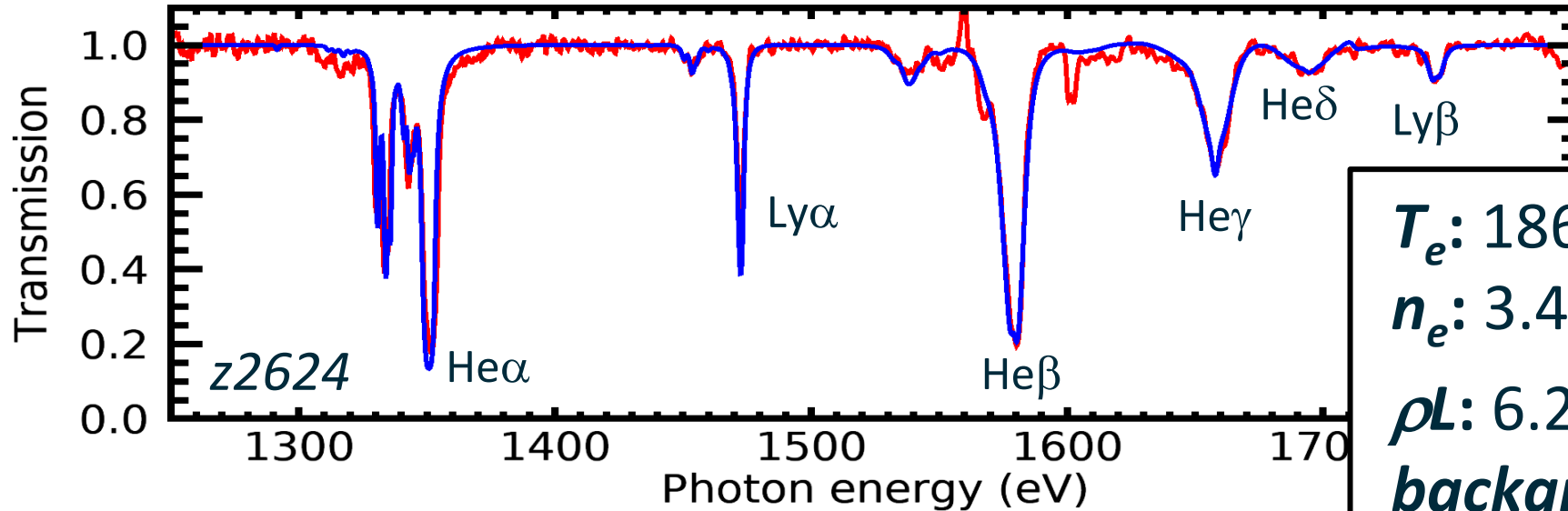
- Model-data-comparison method
- Dual-backlight-intensity method

T_e and n_e analysis method:

- Breakdown analysis



Simultaneous fits to all Mg lines produces too small uncertainties in T_e and n_e due to unrealistic assumptions



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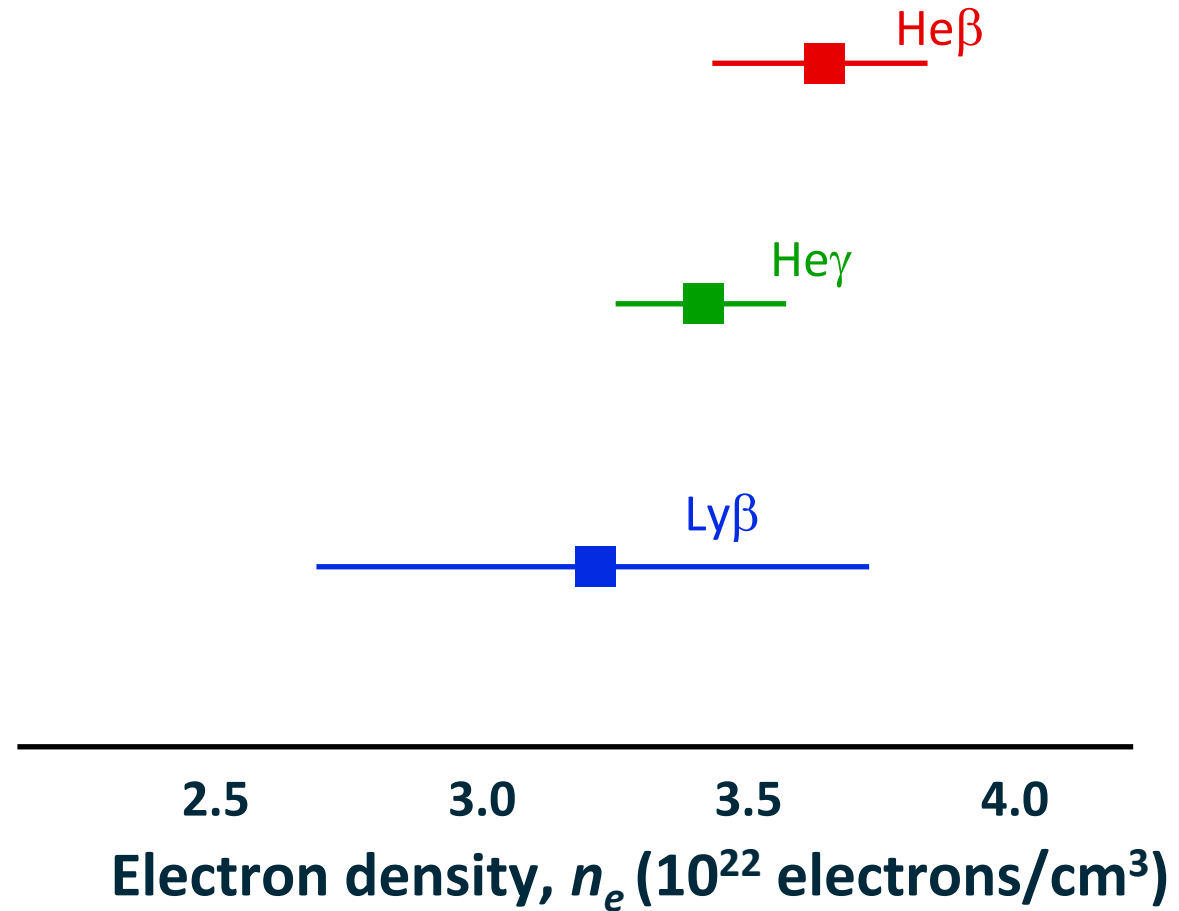
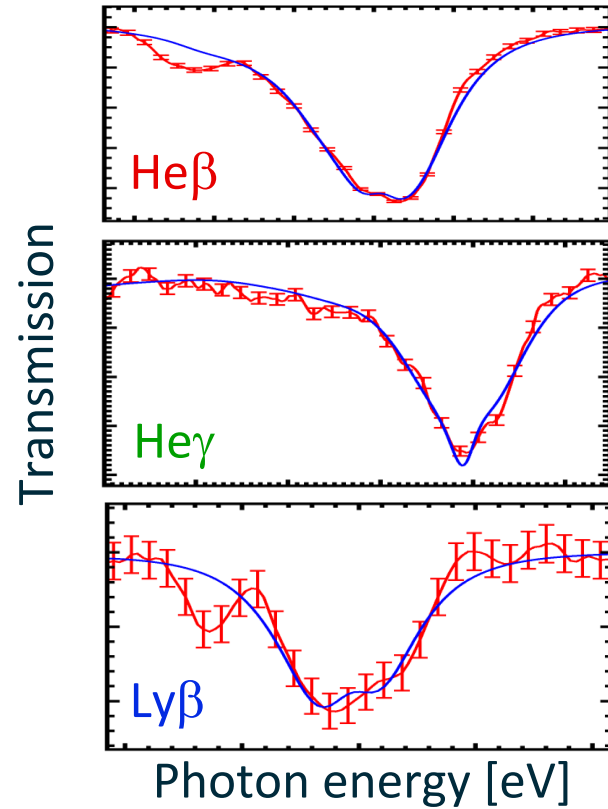
background: 7.4 J/sr/\AA ($\pm 12\%$)

Analysis assumptions:

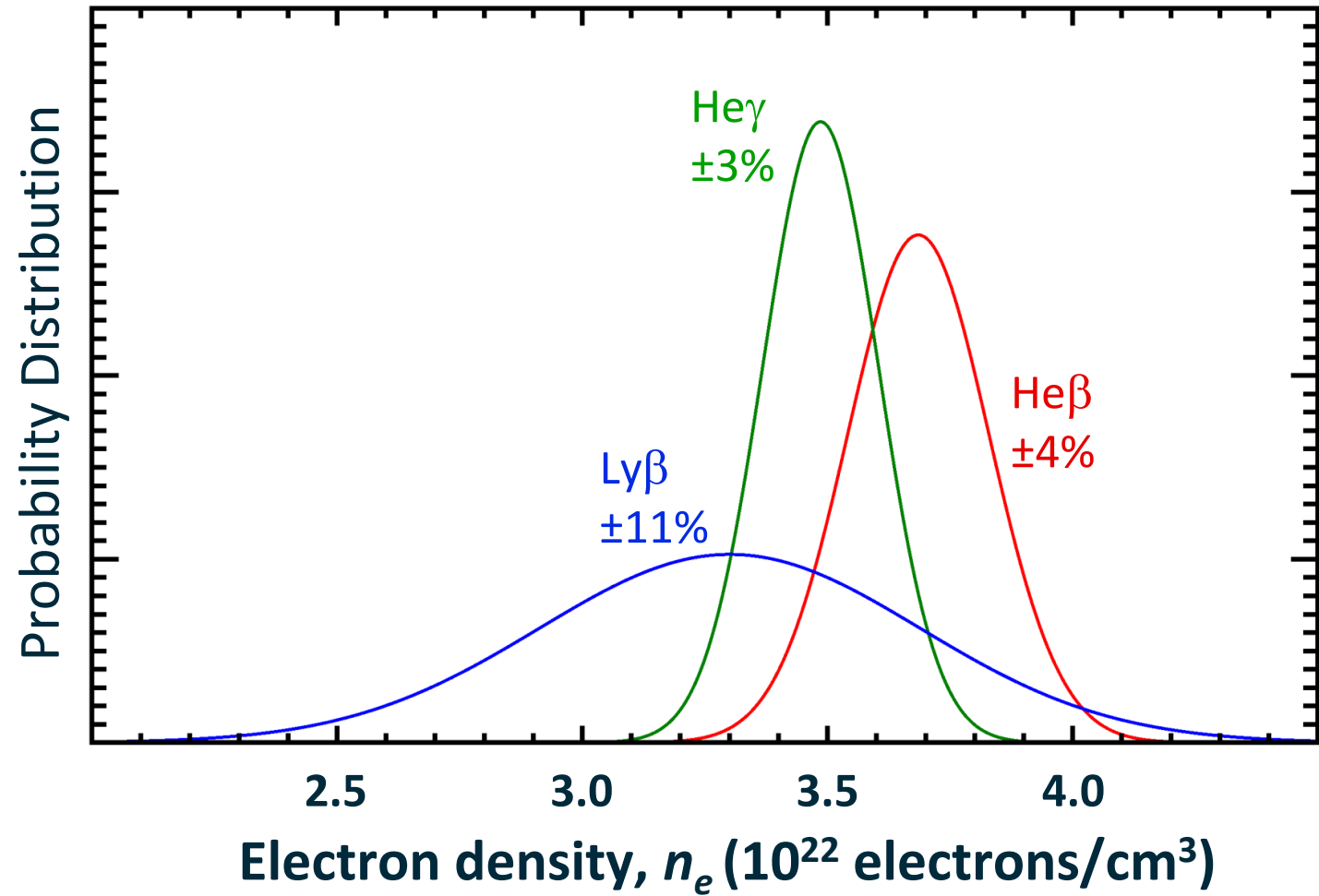
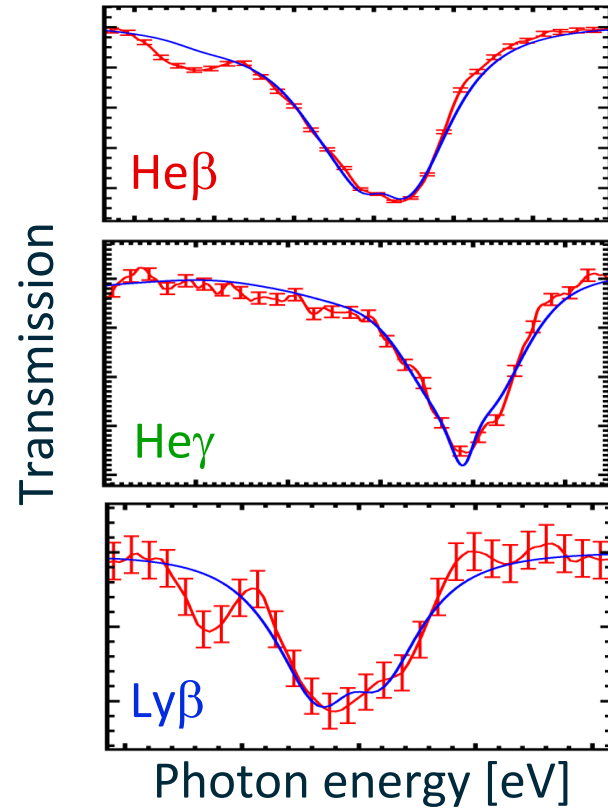
- **Spectral model** is perfect (atomic data, density effects, line shapes, etc)
- **Experimental** plasma is spatially/temporally uniform and LTE
- **Data processing is perfect**, and uncertainty on the data is independent and random

The analysis returns unrealistically small T_e and n_e uncertainties even if these assumptions fail and different lines suggest inconsistent values

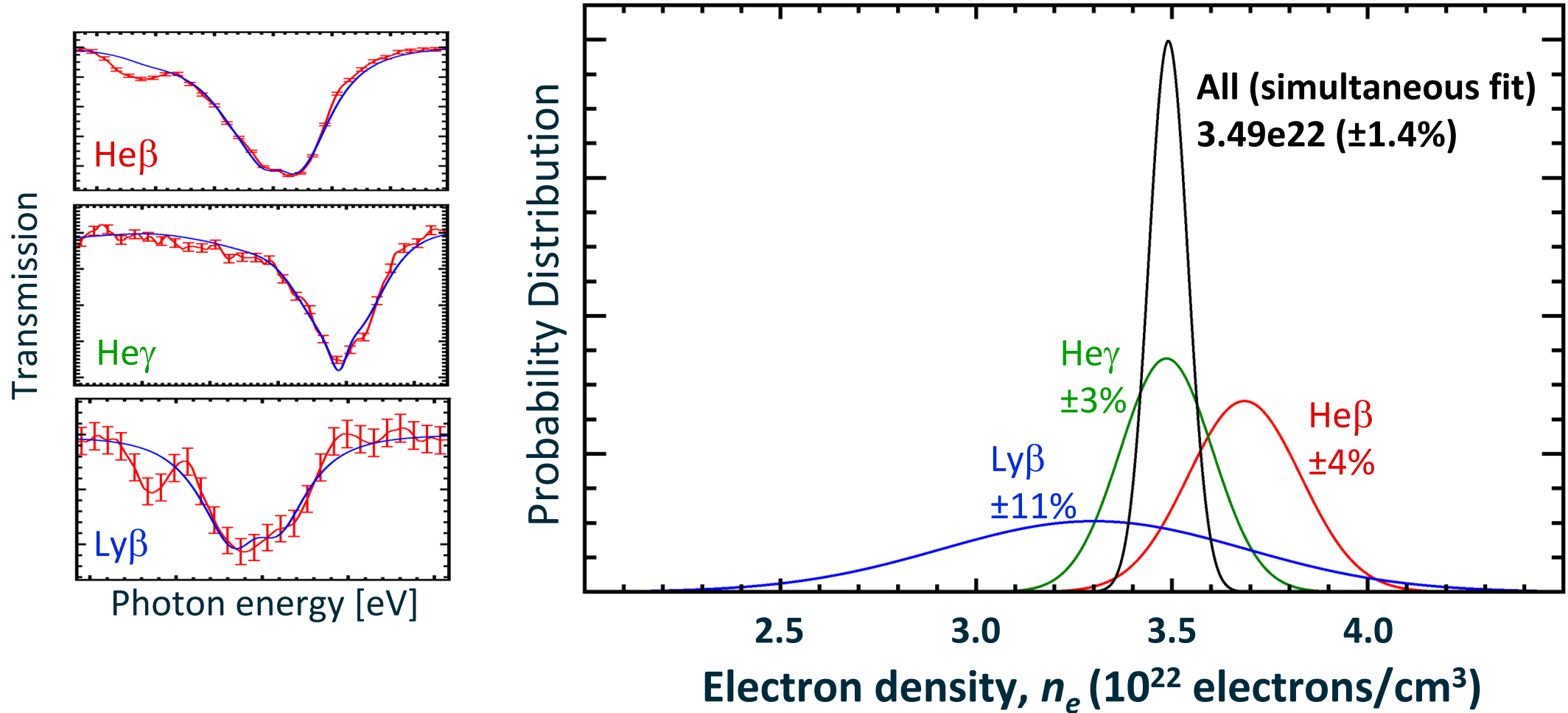
When we have multiple constraints on n_e , we can precisely determine n_e



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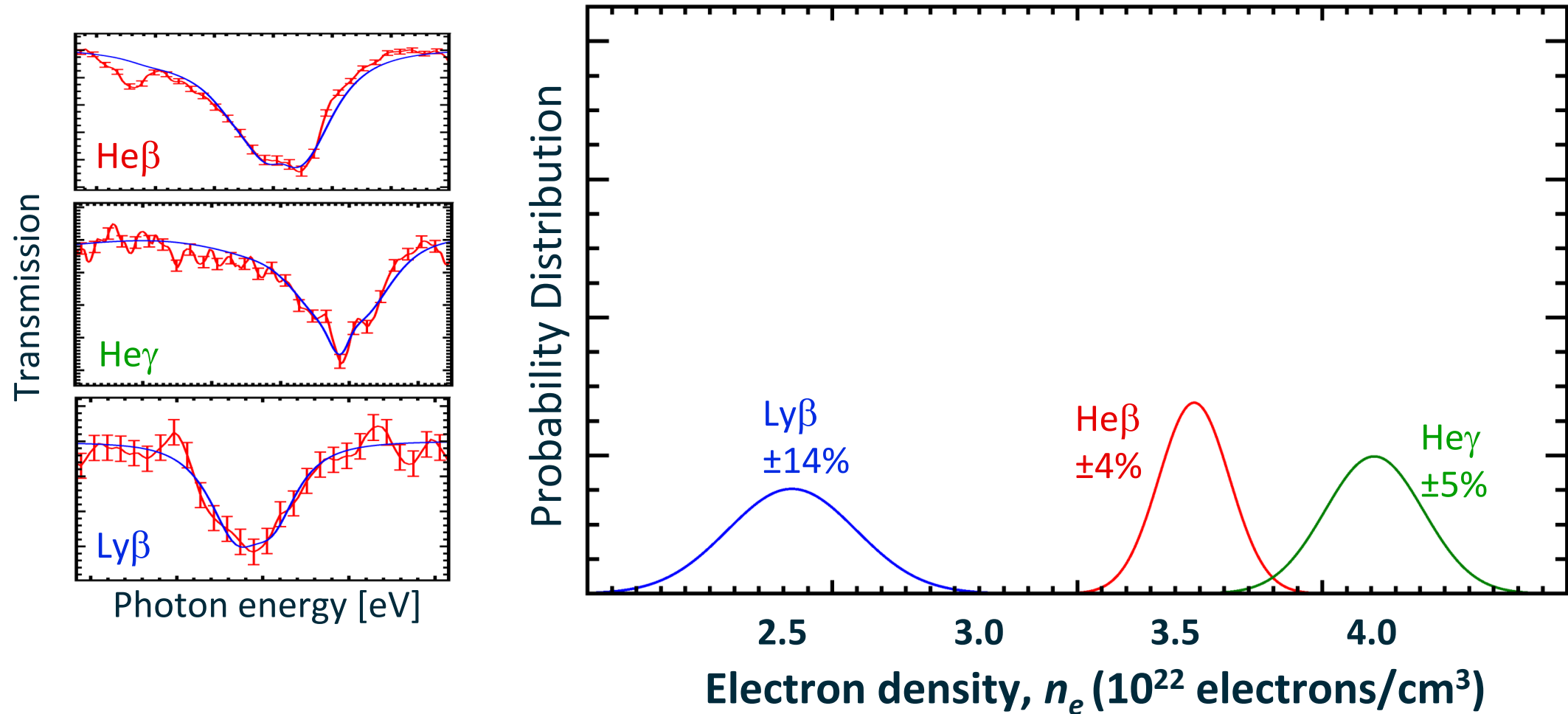


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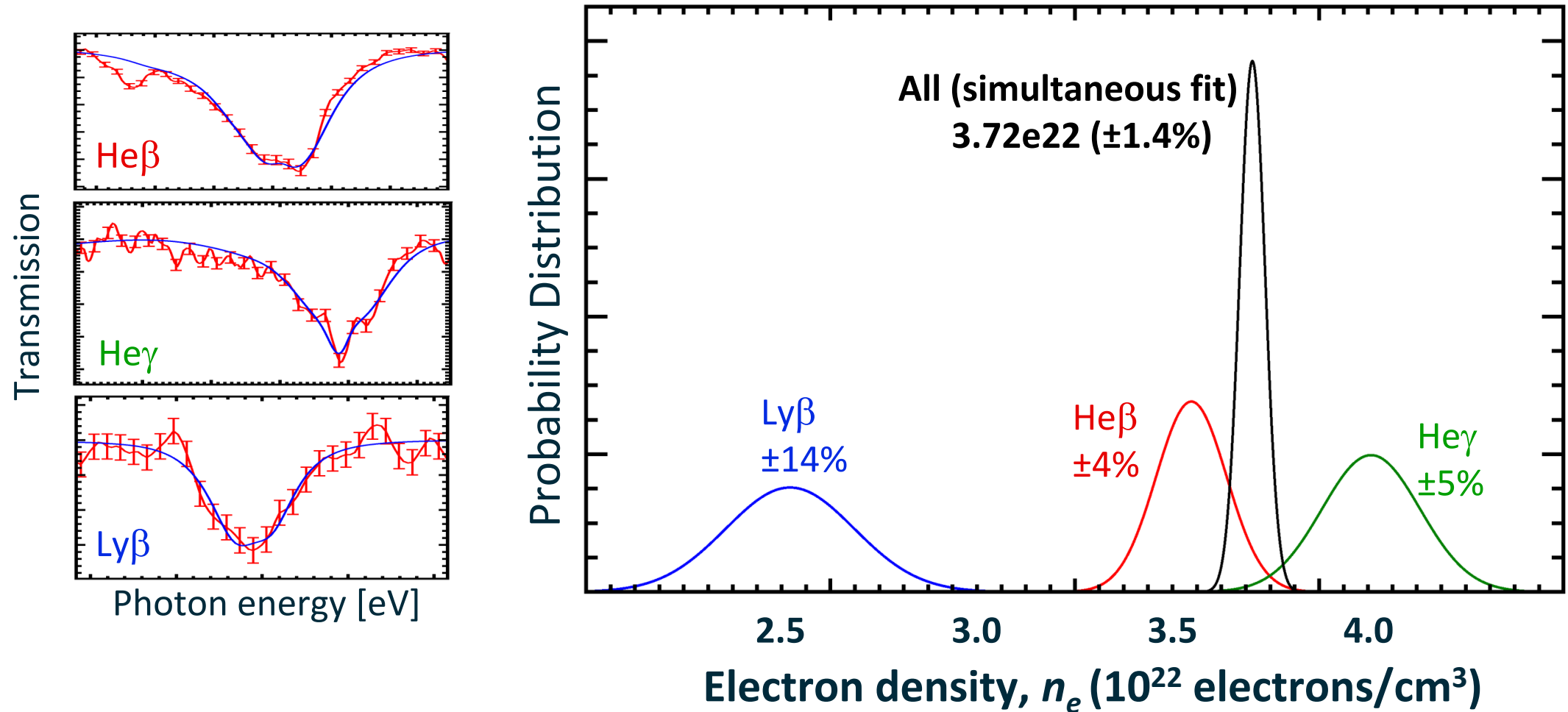
We can determine n_e very precisely by analyzing all ne sensitive lines simultaneously

Fits to the whole spectrum give equally precise results even when different lines disagree with each other



- Inconsistencies are most likely caused by failure of the analysis assumptions
- We want to incorporate the inconsistencies into uncertainty instead of giving precise

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New

Focus on one parameter at a time and take the average and standard deviation from all line-ratios, -broadening, and depths analyses

Step1: Preliminary: Simultaneous fit to all lines

Step2: Background: Determine and subtract background from the data [1]

Step3: n_e : Analyze Mg He β , He γ , Ly β line shapes [2,3]

Step4: ρ_L : Analyze Mg He β and He γ line depths

Step5: T_e : Analyze 11 temperature sensitive line ratios

More
Correlation
Less

* Errors due to preceding steps are propagated to the first order

New

Focus on one parameter at a time and take the average and standard deviation from all line-ratios, -broadening, and depths analyses

Step1: Preliminary: Simultaneous fit to all lines

Pros:

- More realistic uncertainties
- Be aware of inconsistencies
→ Clues for problems
- Clear on what physics we rely on for each parameter

Step2: Background

Step3: n_e : n_e

Step4: ρ_L : ρ_L

Step5: T_e : T_e

Cons:

- Tedious
- Parameter correlation is approximated to the 1st order

data [1]

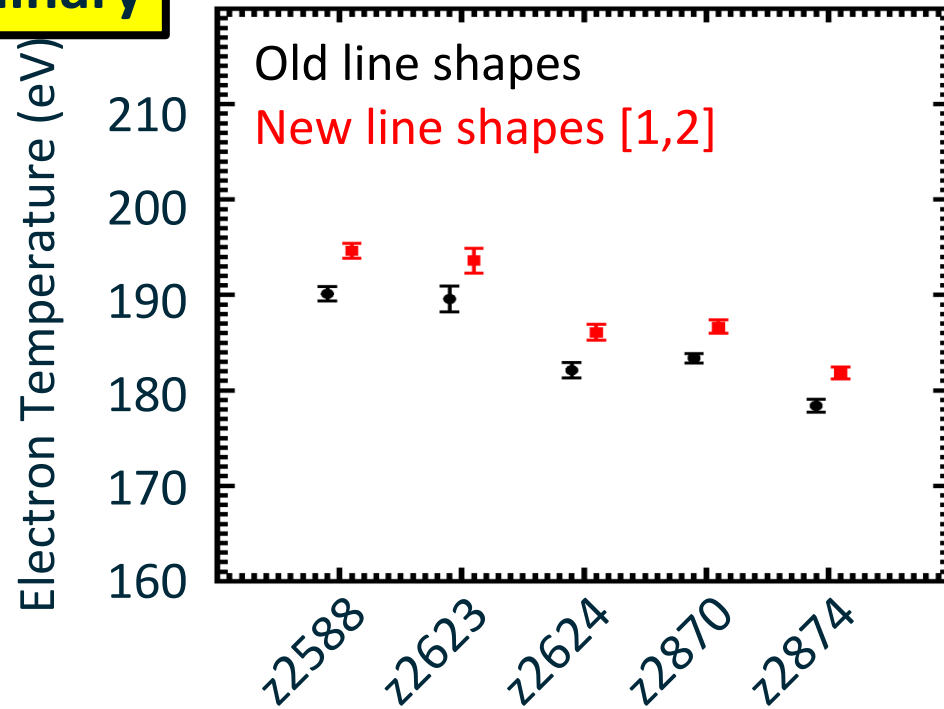
More Correlation Less

* Errors due to preceding steps are propagated to the first order

New line-shapes increased inferred T_e and n_e by 2% and 20%, respectively



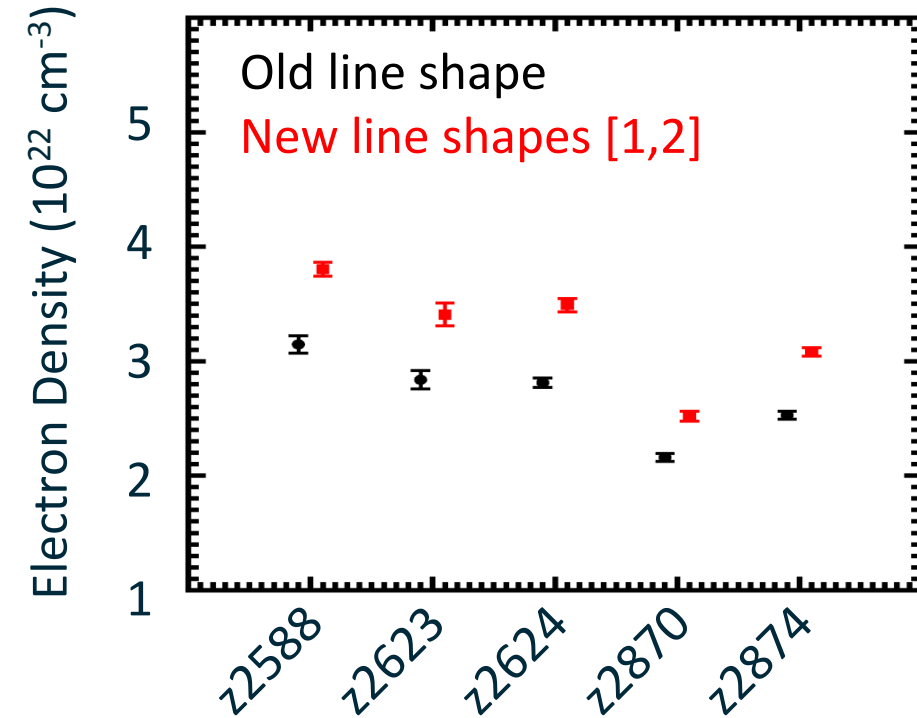
Preliminary



185 ± 5 eV ($\pm 3\%$)



189 ± 5 eV ($\pm 3\%$)



$(2.7 \pm 0.4) \times 10^{22} \text{ cm}^{-3}$ ($\pm 14\%$)

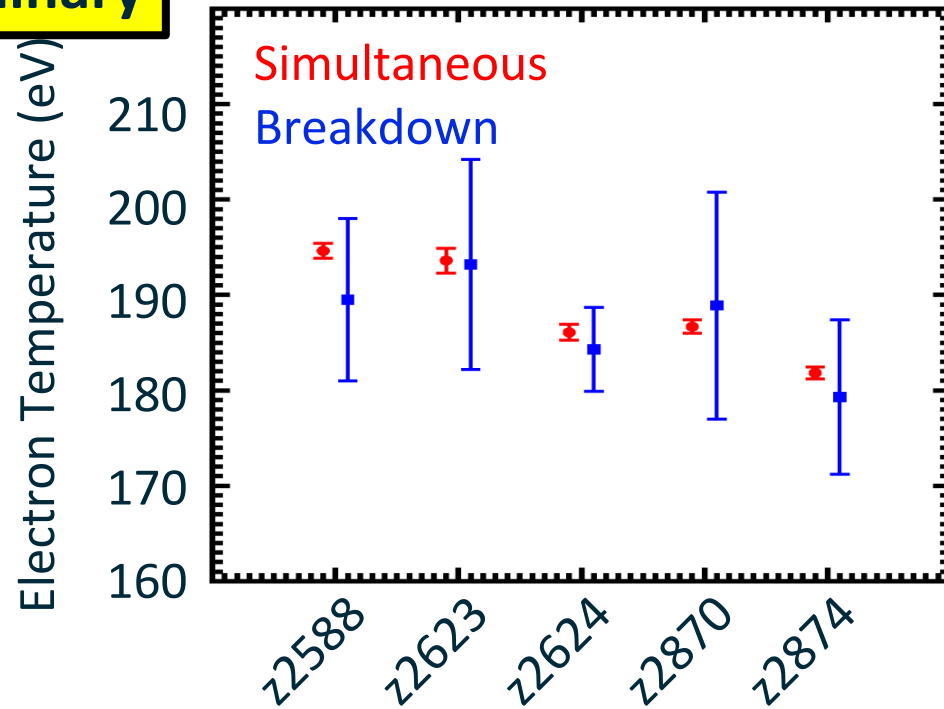


$(3.3 \pm 0.5) \times 10^{22} \text{ cm}^{-3}$ ($\pm 15\%$)

Breakdown analysis produced more realistic individual uncertainties with little impact on the average values



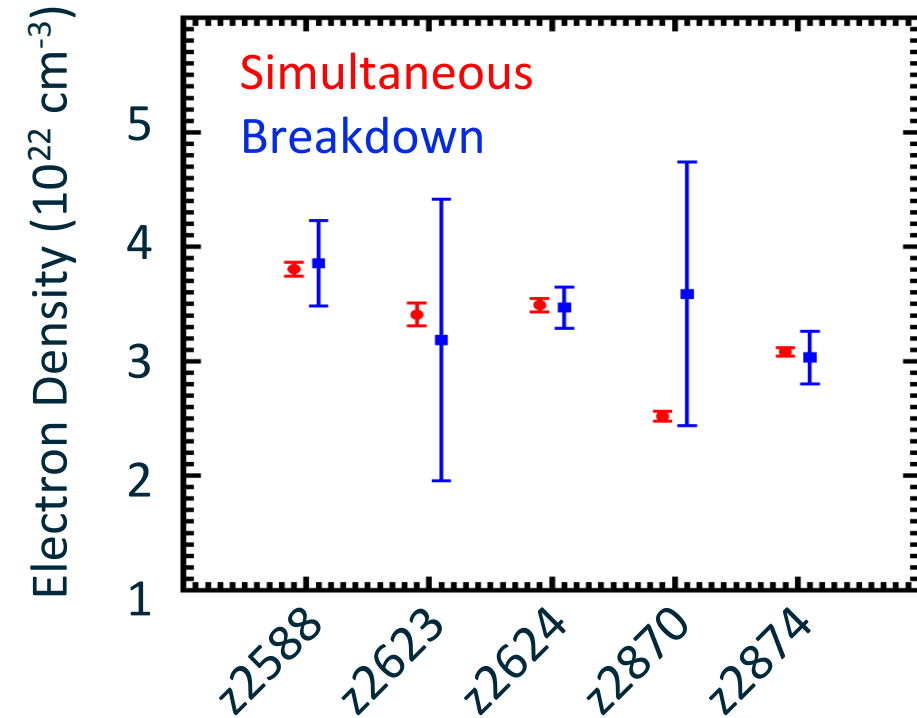
Preliminary



$189 \pm 5 \text{ eV } (\pm 3\%)$



$187 \pm 5 \text{ eV } (\pm 3\%)$



$(3.3 \pm 0.5) \times 10^{22} \text{ cm}^{-3} (\pm 15\%)$



$(3.4 \pm 0.3) \times 10^{22} \text{ cm}^{-3} (\pm 10\%)$

Summary: Plasma temperature and density analysis are refined for Sandia iron opacity experiments



Motivation: Fe opacity measured at solar interior temperature disagree with calculated opacity

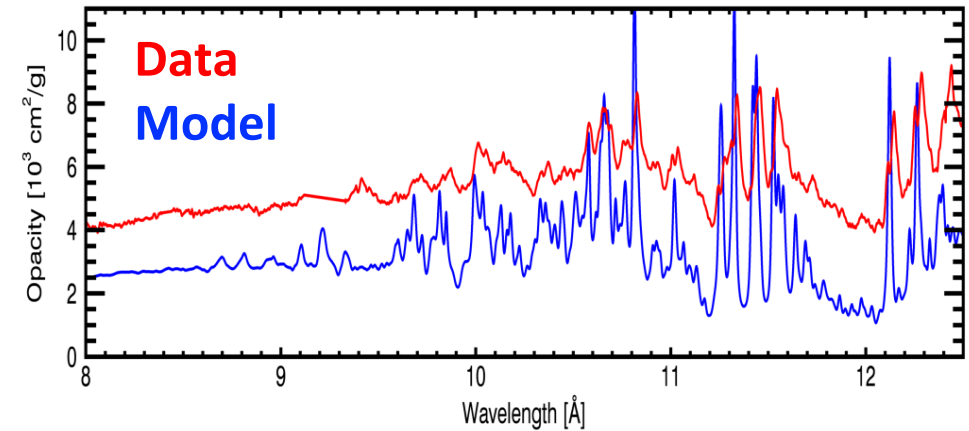
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- Clarify the final number is standard deviation
- Make the summary shorter
- Show the latest comparison
- Raise WD as another application