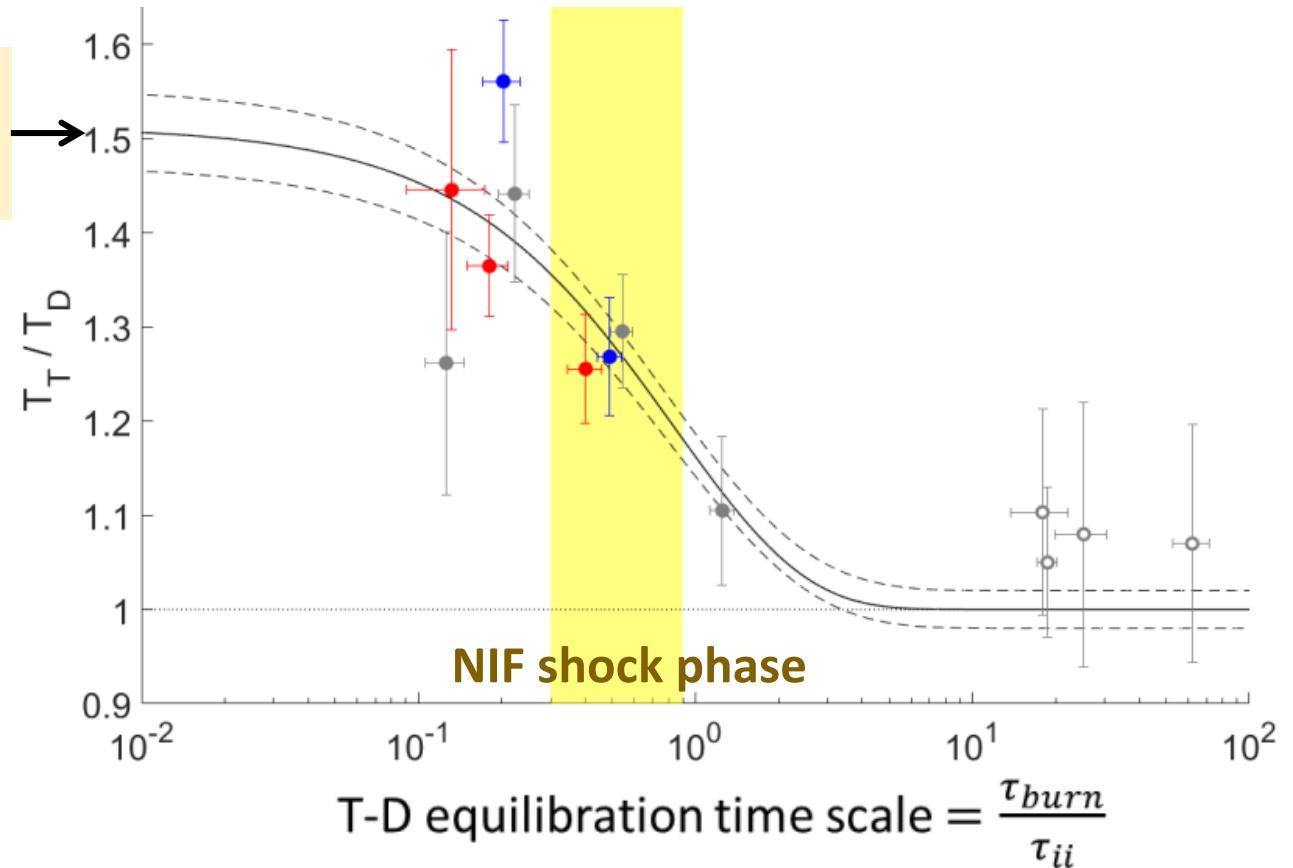


Thermal decoupling of deuterons and tritons during the shock-convergence phase in Inertial Confinement Fusion implosions

$$\left. \frac{T_T}{T_D} \right|_{t=0} = 1.5 = \frac{m_T}{m_D} \rightarrow$$



Neel Kabadi

10/2021

63rd Annual Meeting of the APS Division of Plasma Physics

11/9/2021

Kabadi et al., Phys. Rev. E **104**, L013201 (2021)

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L. Welch, B. Srinivasan
Virginia Tech



H. Sio, J. Sanchez
Lawrence Livermore National Laboratory



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University of Rome Sapienza



J. Eriksson
Uppsala University



W. Taitano, B. Keenan, S. Anderson, A. Simakov, L. Chacón
Los Alamos National Laboratory



Brian Appelbe
Imperial College London

The observed difference between T_{DTn} and T_{DDn} is explained by shock coupling proportional to mass, and equilibration consistent with hydro theory.

- This is the first direct lab-based experimental measurement of shocks coupling directly proportional to mass in multi-ion plasmas, verifying a long standing prediction.
- For NIF ignition implosions, during the shock convergence phase the D and T ions are out of thermal equilibrium which greatly impacts the ion viscosity.
- iFP simulations indicate that this result is robust, even in the highly kinetic regime.

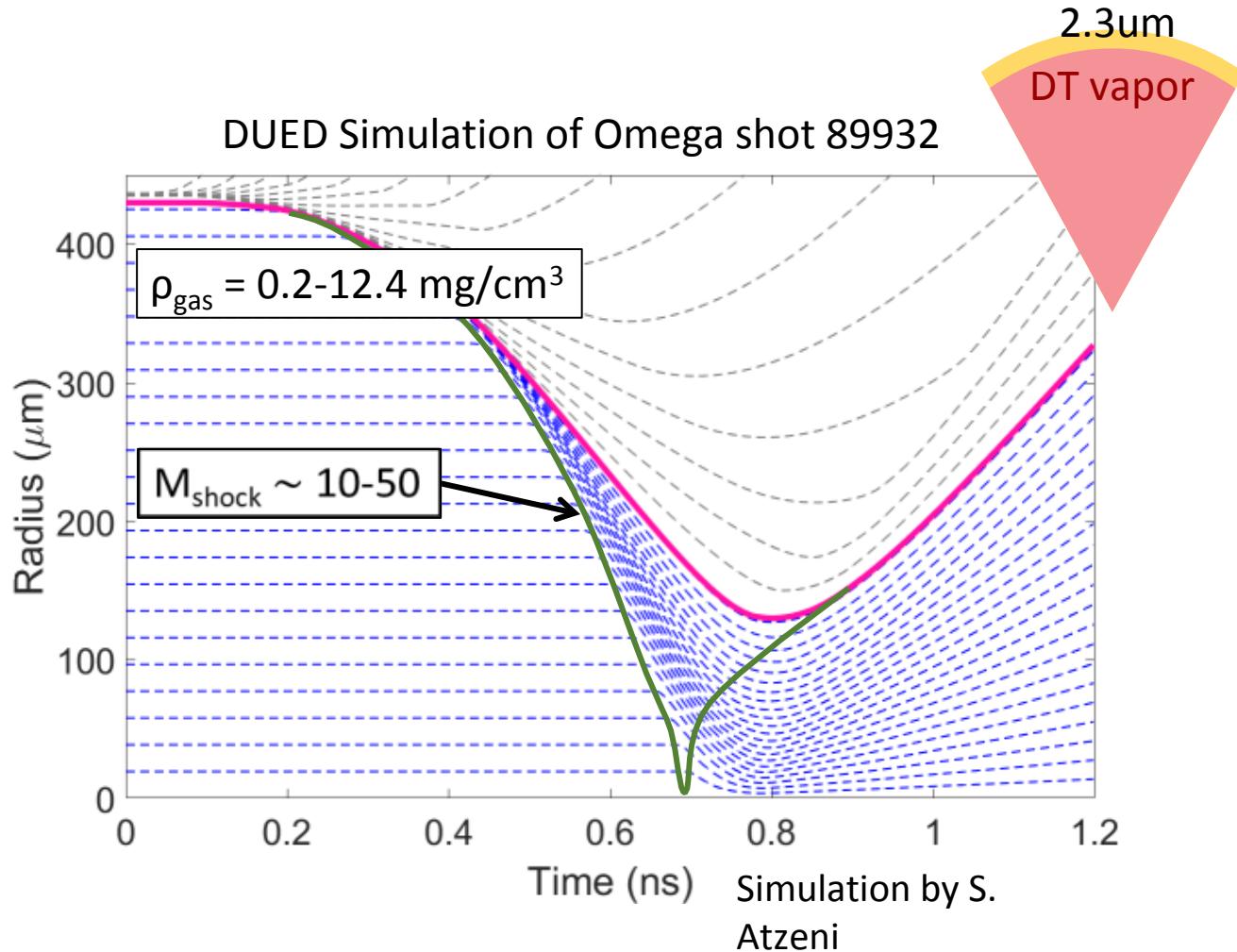
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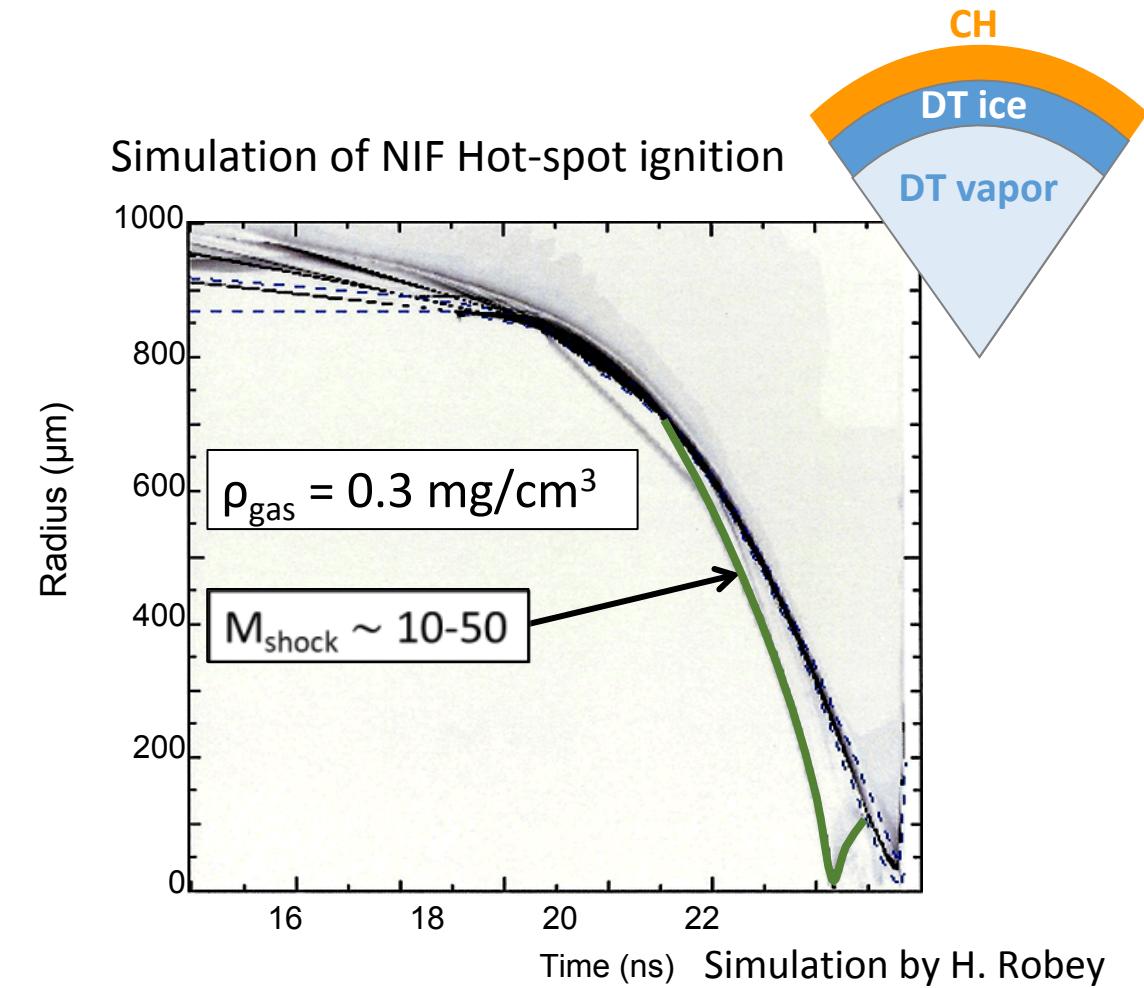
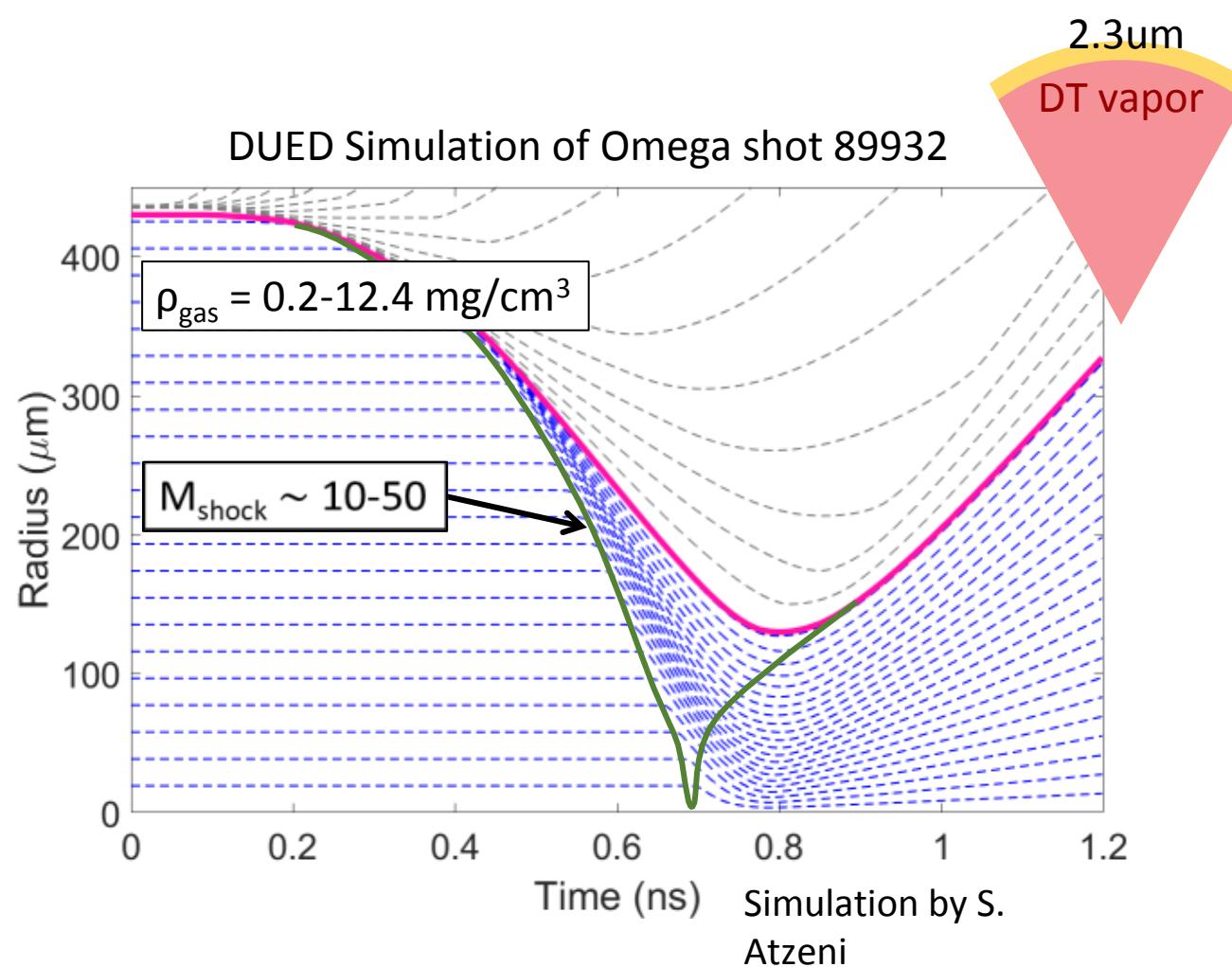
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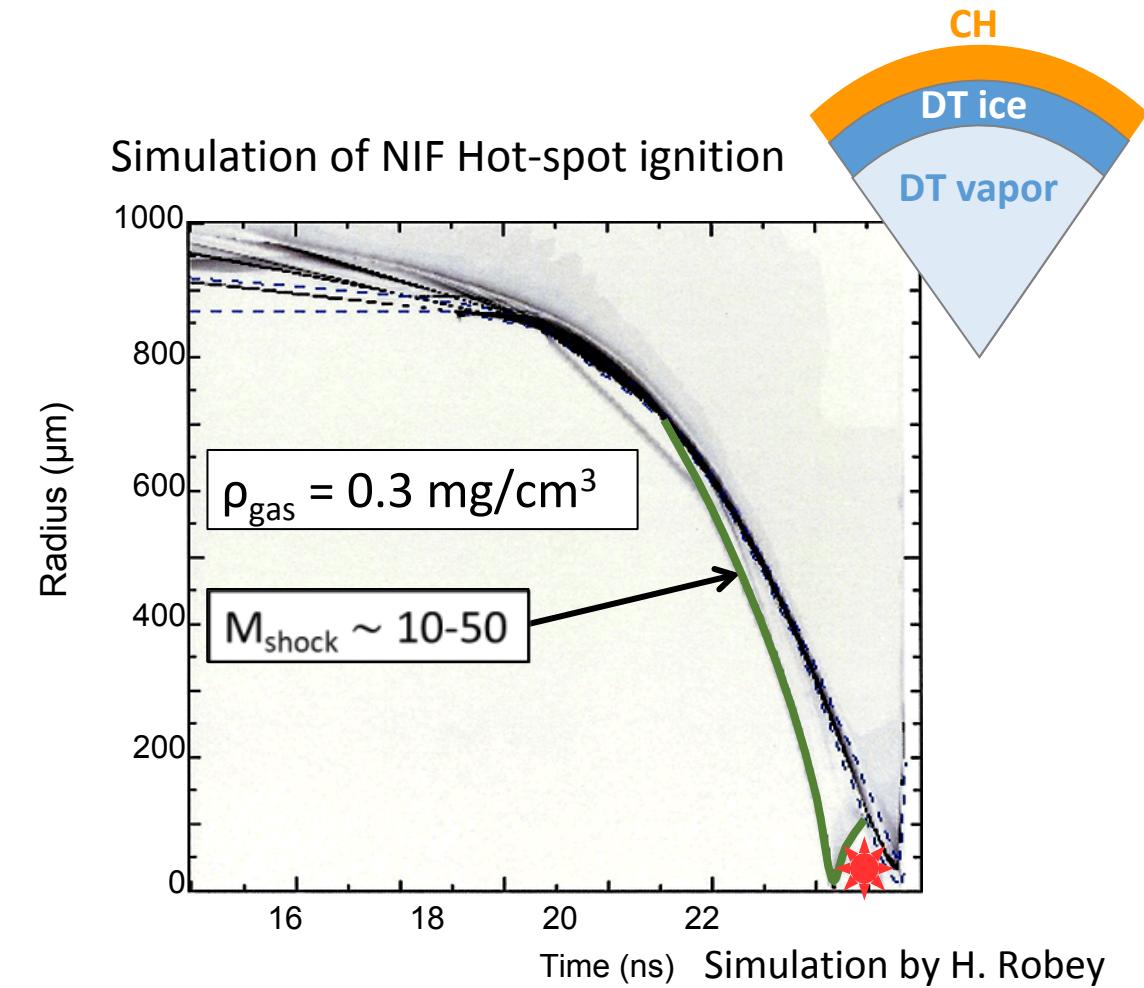
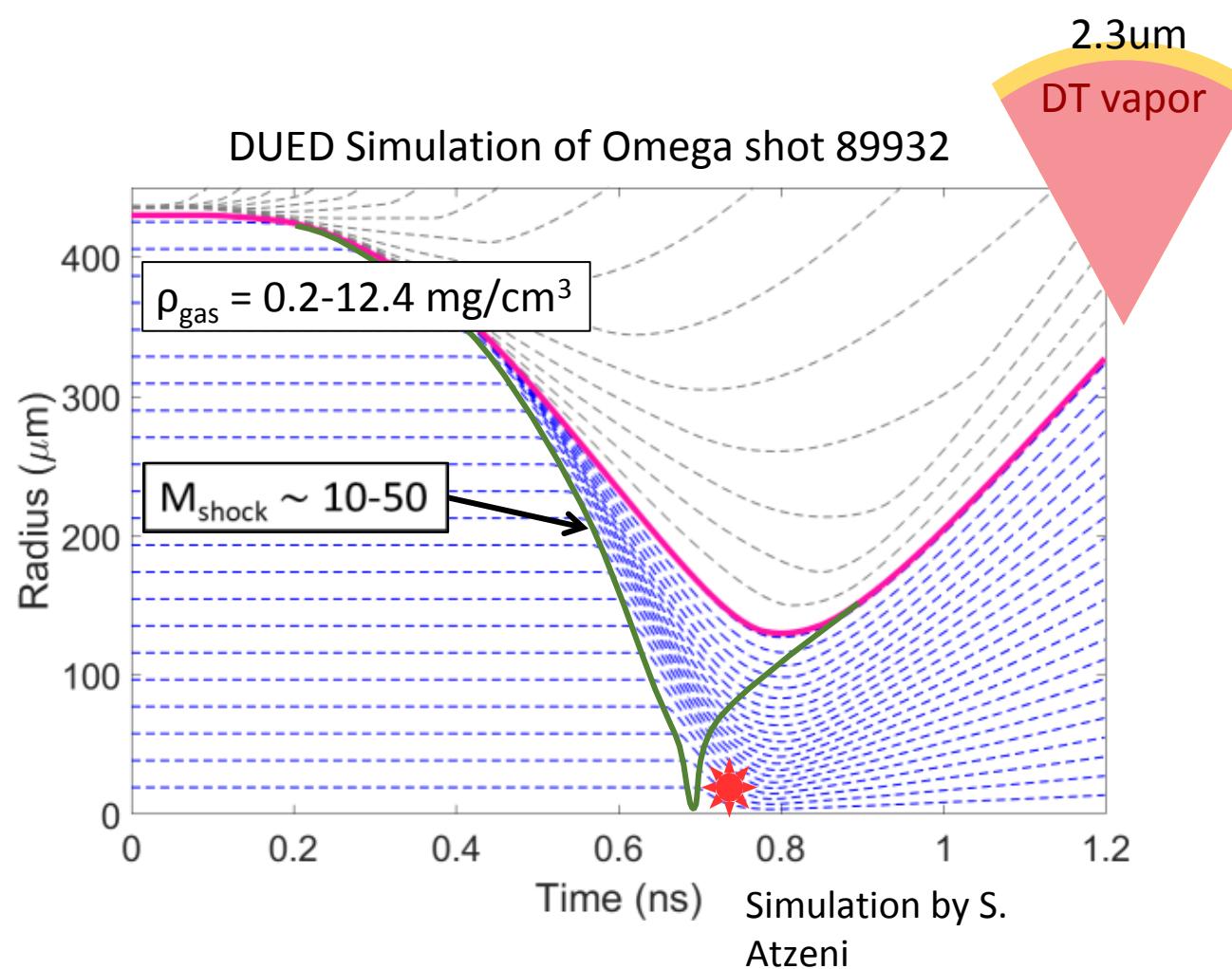
Thin shell shock driven implosions generate a shock convergence phase with similar conditions to ignition implosions



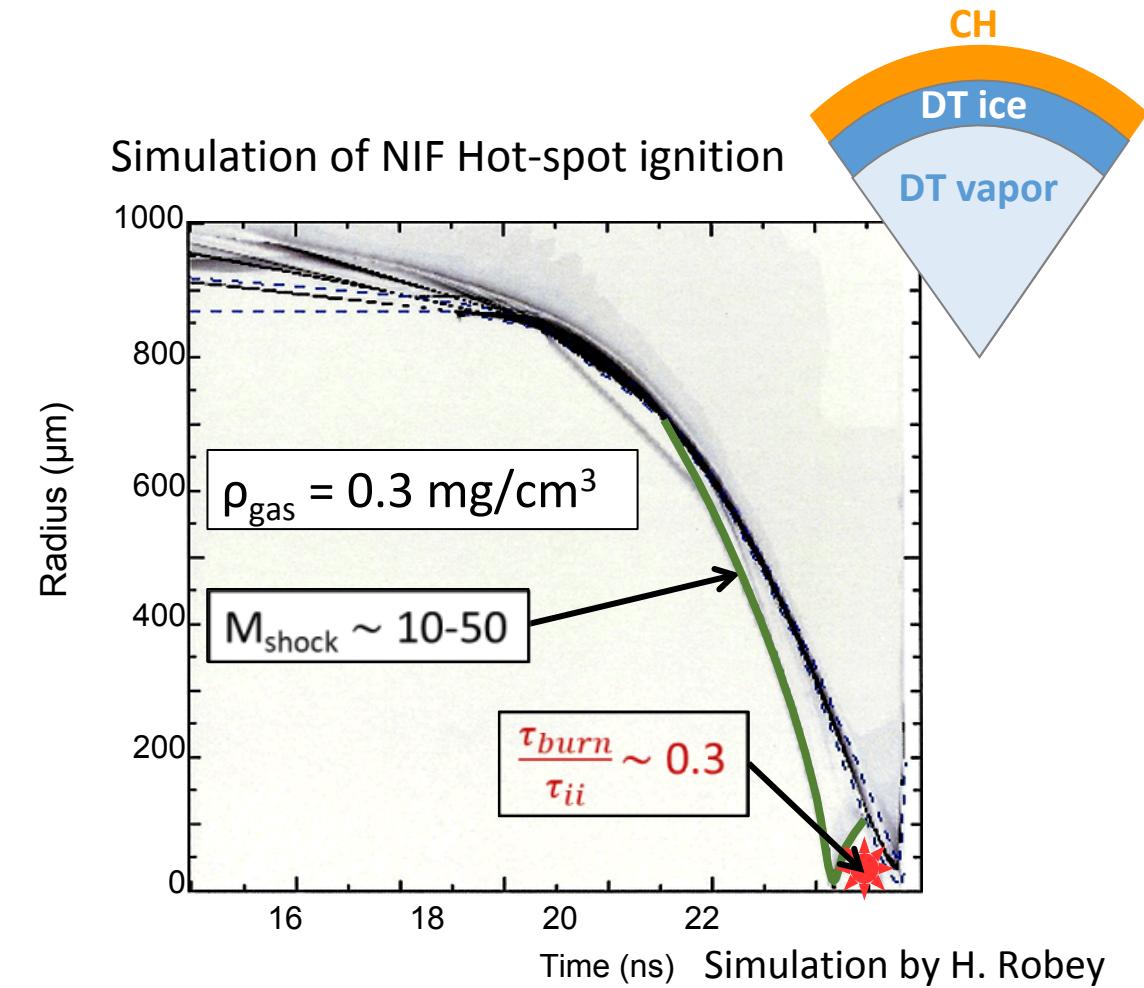
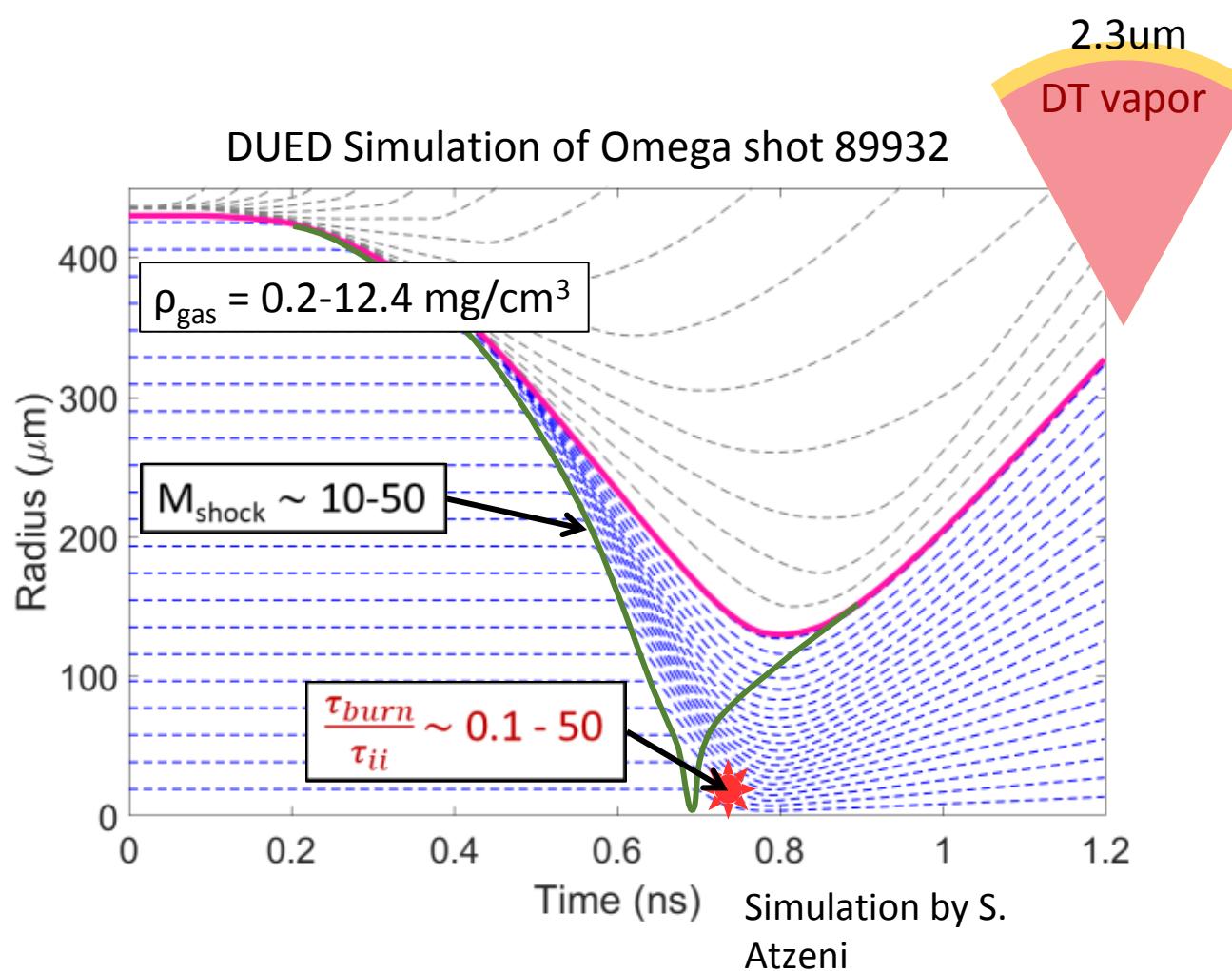
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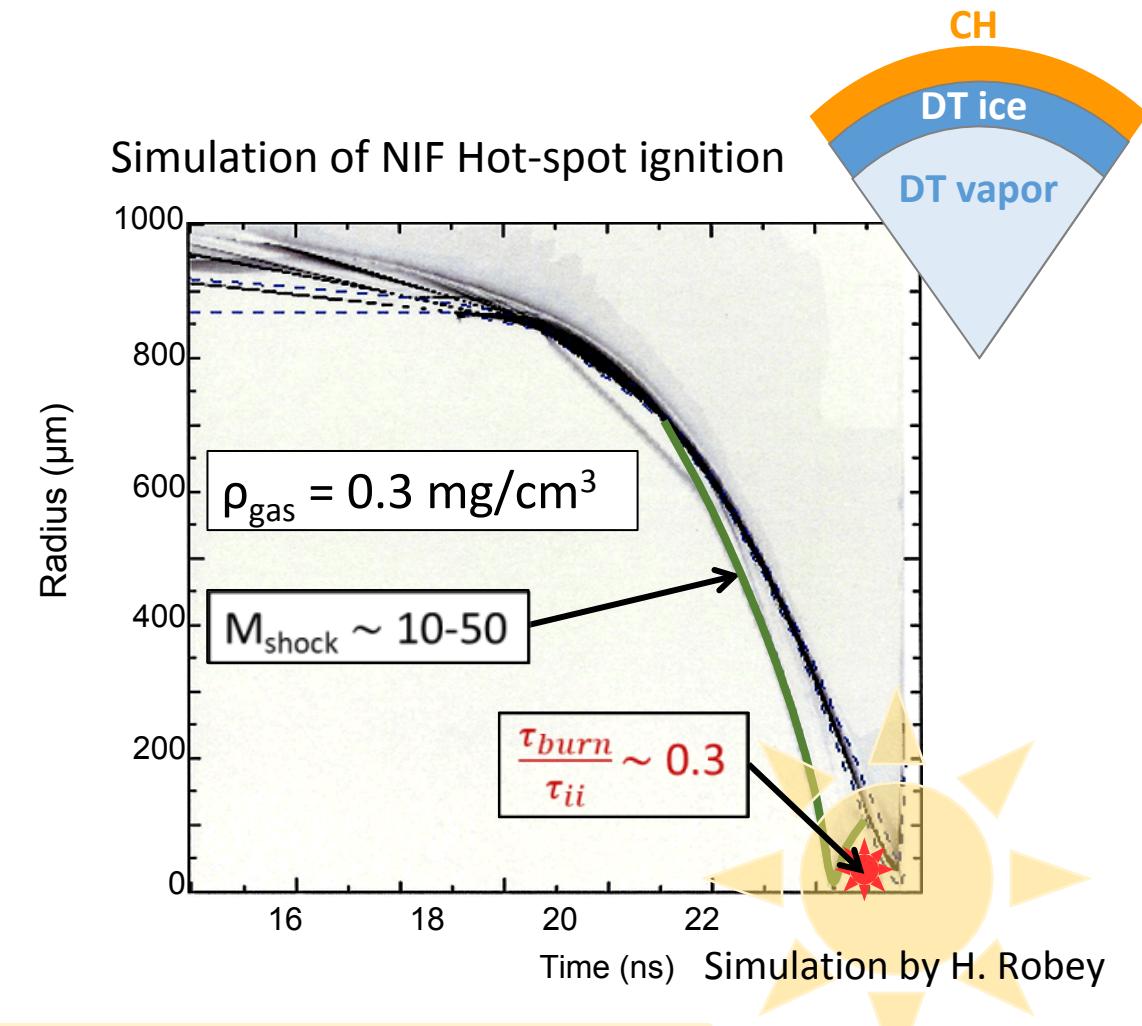
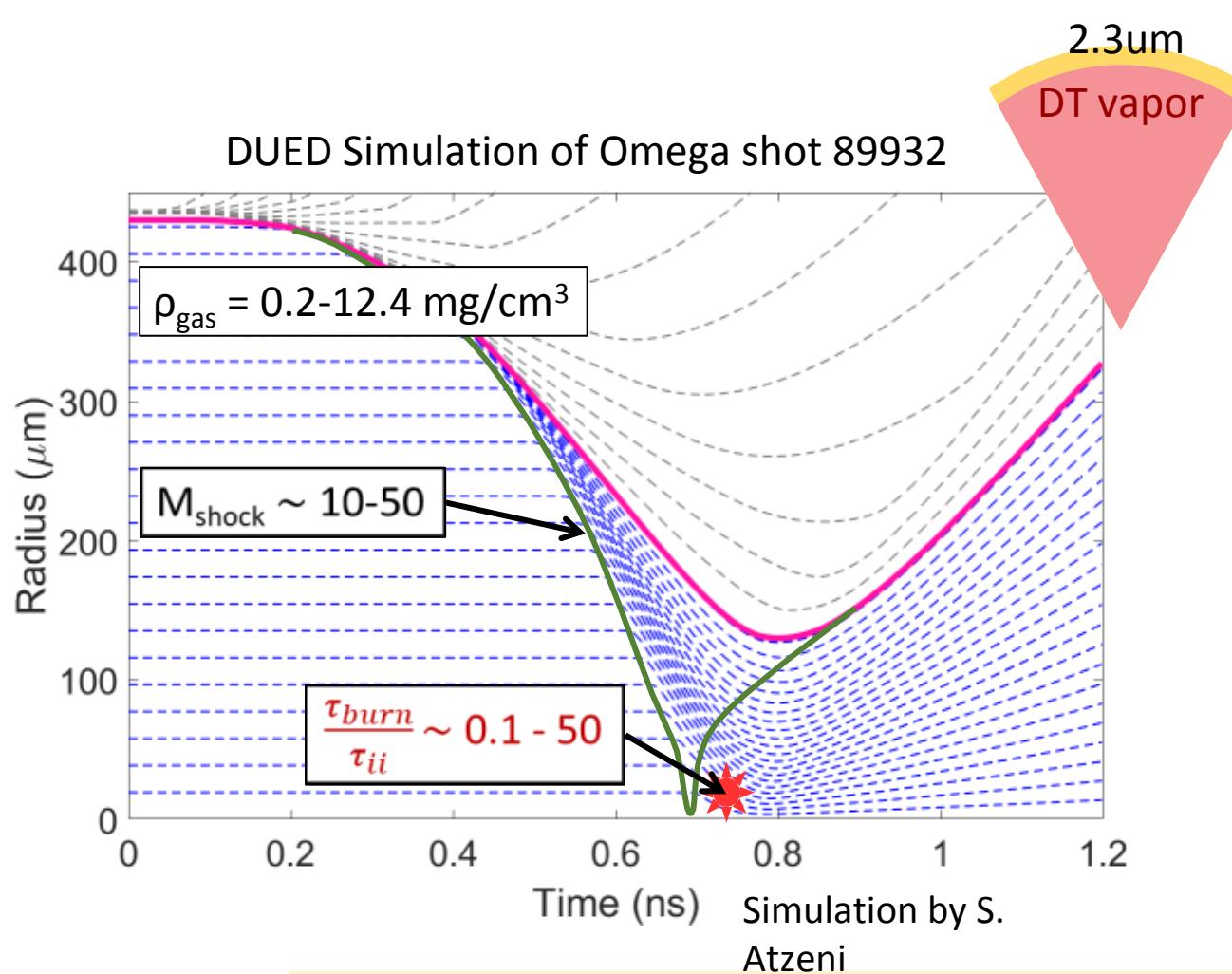
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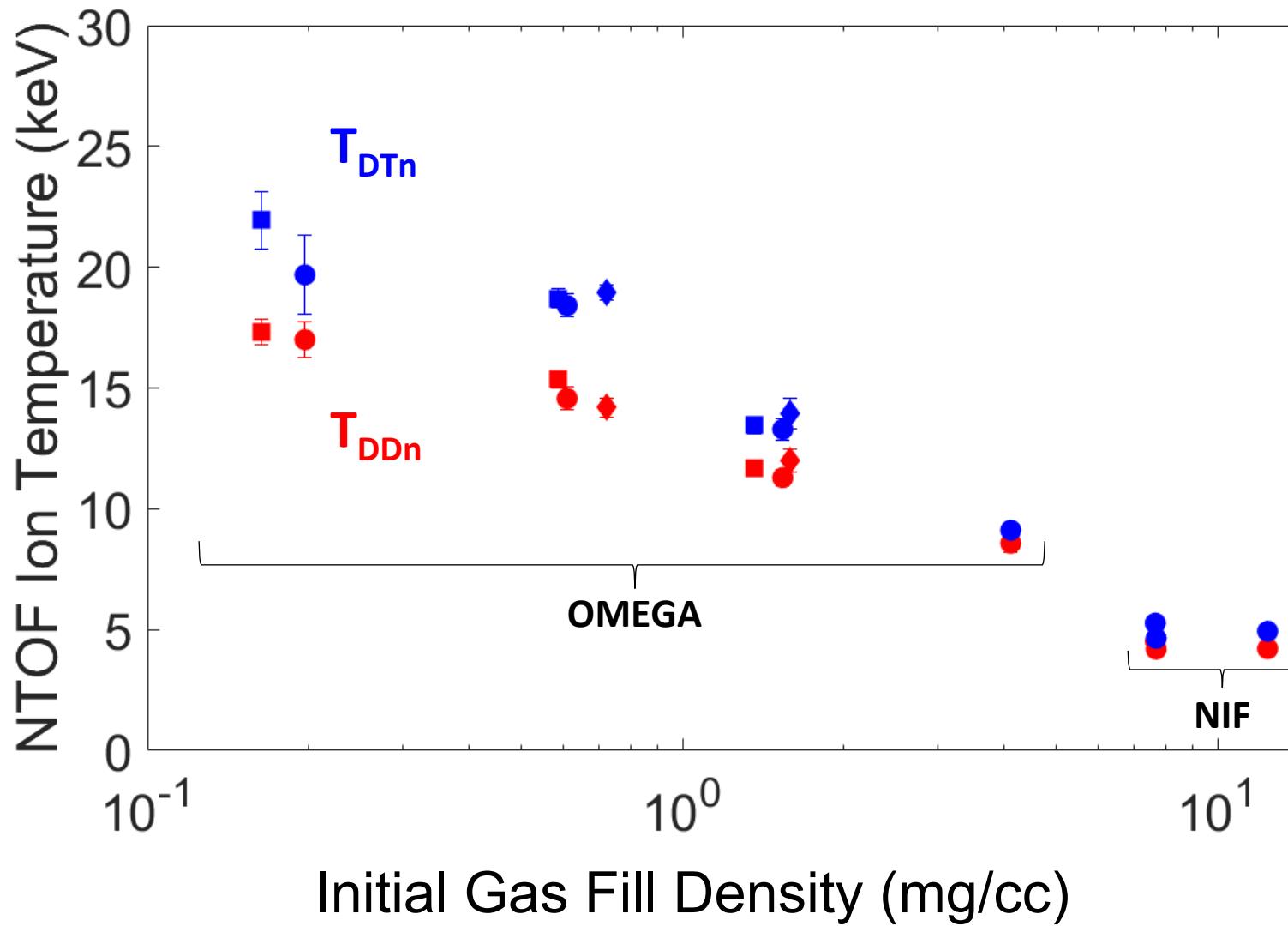


This phase is difficult to study in ignition-relevant implosions in which the subsequent compression phase dominates the nuclear and X-ray observables.

Outline

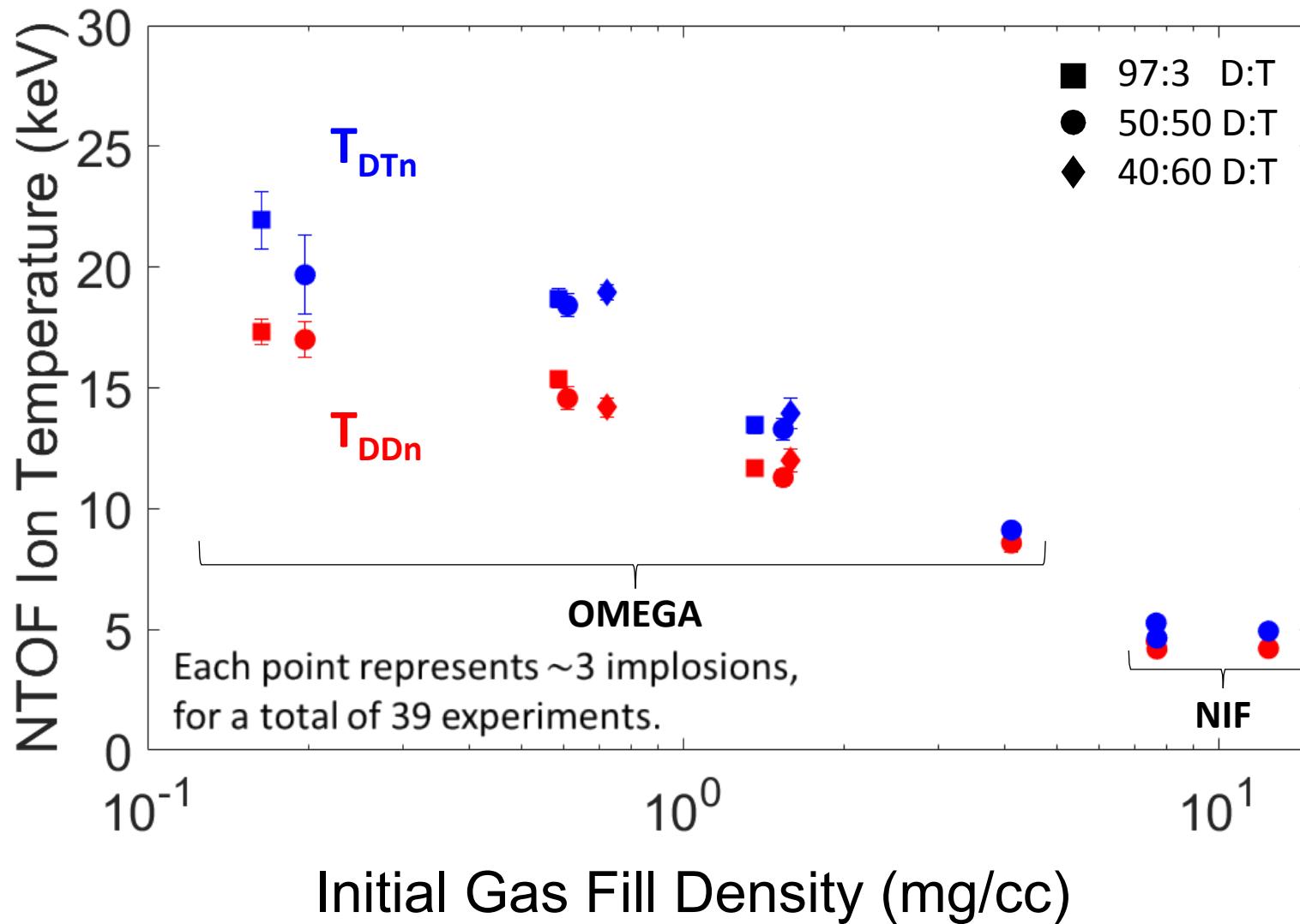
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The apparent D and T ion temperatures can be inferred from the measured DTn and DDn temperatures

$$T_{DTn} = \frac{m_T T_T + m_D T_D}{m_T + m_D}$$

$$\begin{array}{|c|c|c|} \hline & & \\ \hline & \equiv & \\ \hline & & \end{array}$$

$$T_T = T_{DTn} + \frac{m_D}{m_T} (T_{DTn} - T_{DDn})$$

For the NIF data, a higher order model taking into account temperature variation is used.

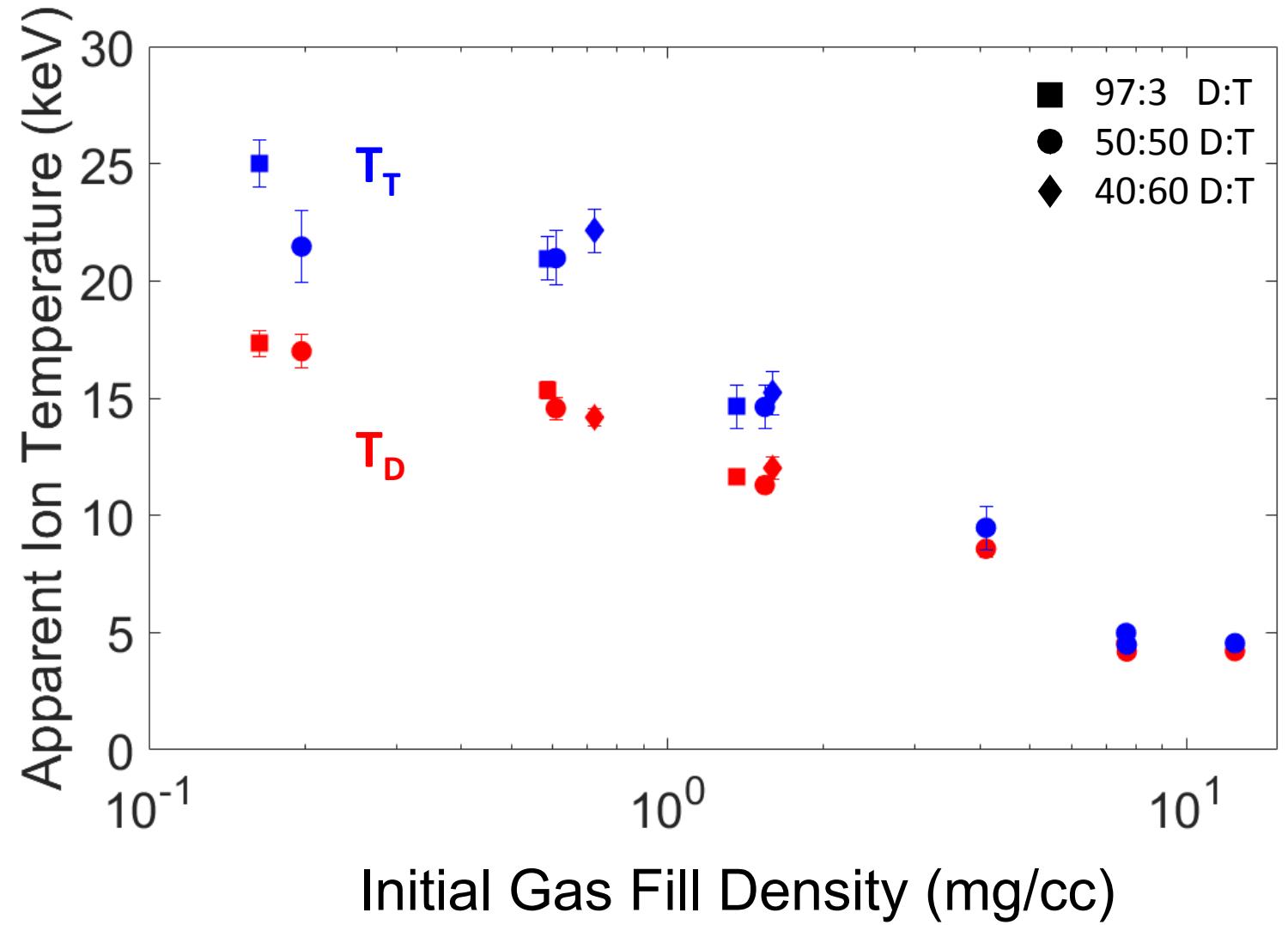
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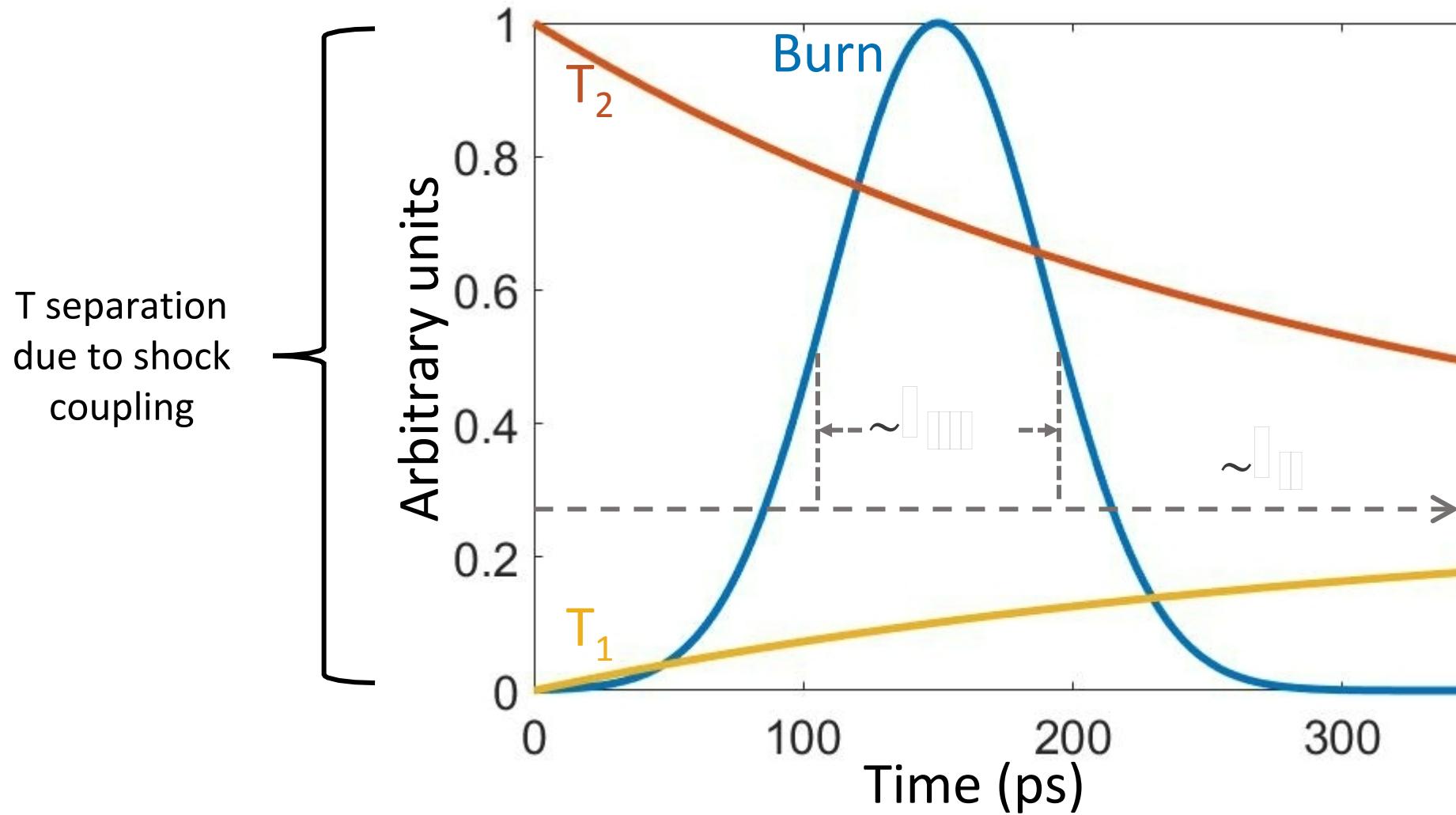
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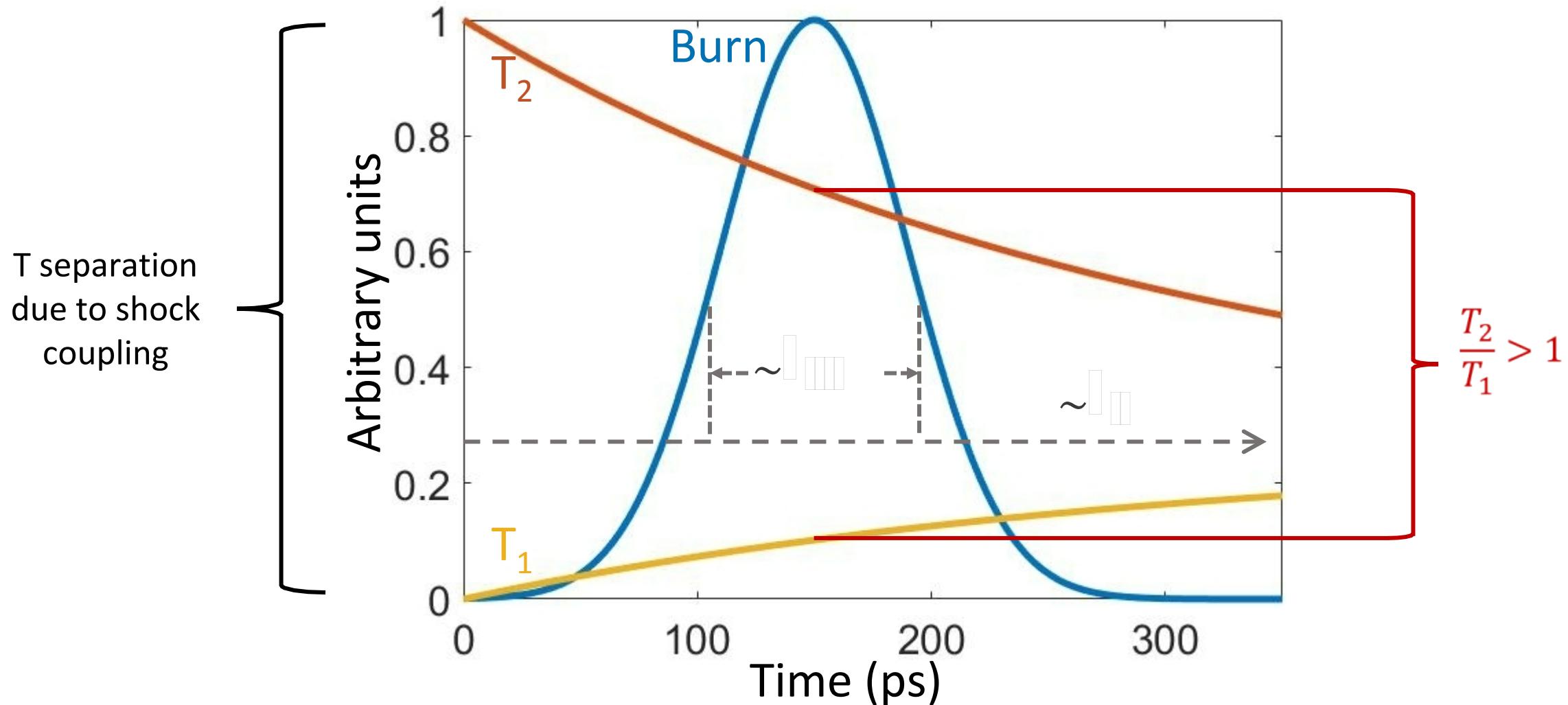
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$$\tau_{12} = 5.6 \times 10^{18} \frac{(m_1 T_2 + m_2 T_1)^{3/2}}{Z_1^2 Z_2^2 (m_1 m_2)^{1/2} n_2 \lambda_{12}} \text{ sec.} \quad \tau_{ii} = \left[\frac{1}{\tau_{12}} + \frac{1}{\tau_{21}} \right]^{-1}$$

$$\frac{dT_1}{dt} = \frac{1}{\tau_{12}} (T_2 - T_1). \quad \frac{dT_2}{dt} = \frac{1}{\tau_{21}} (T_1 - T_2).$$

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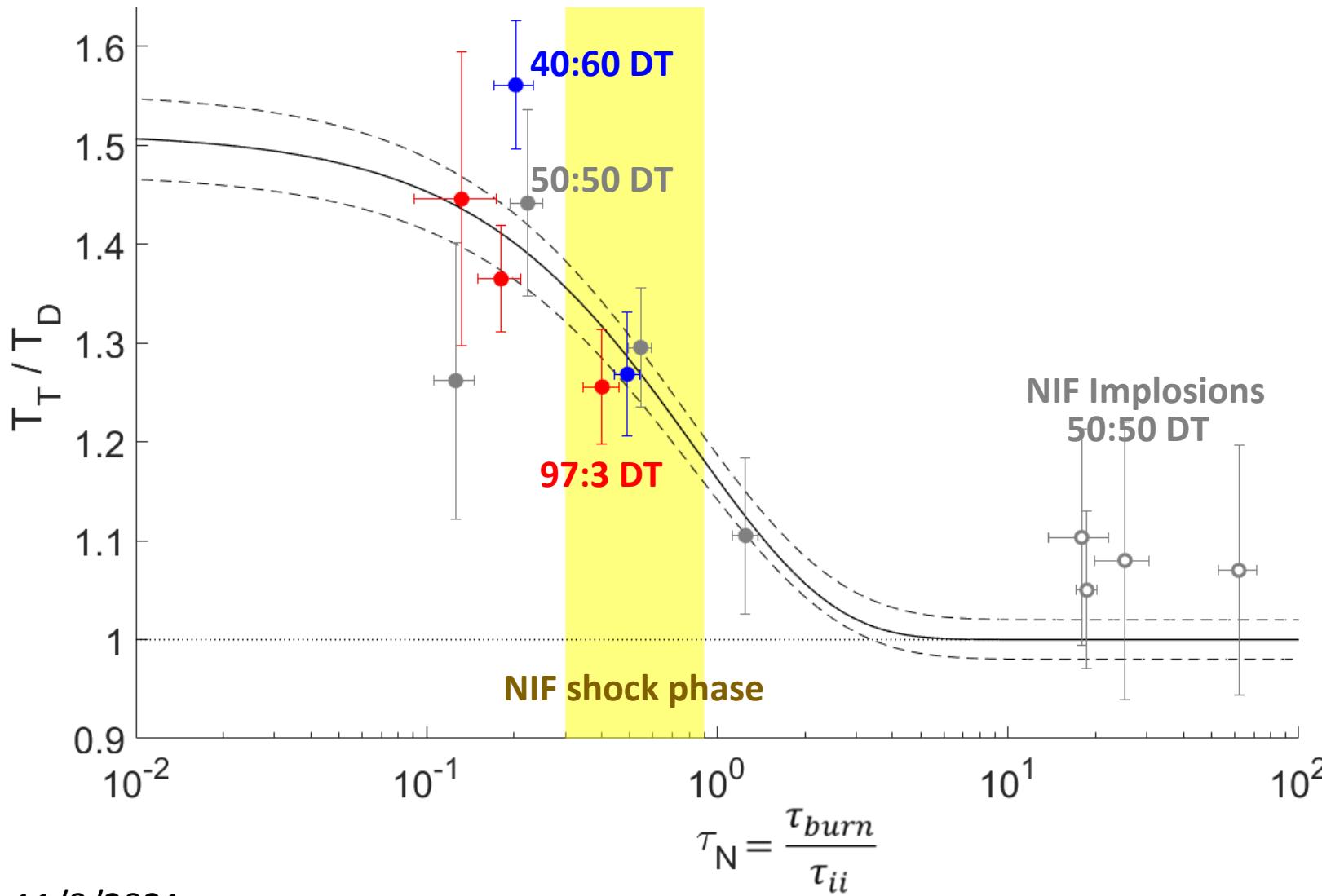
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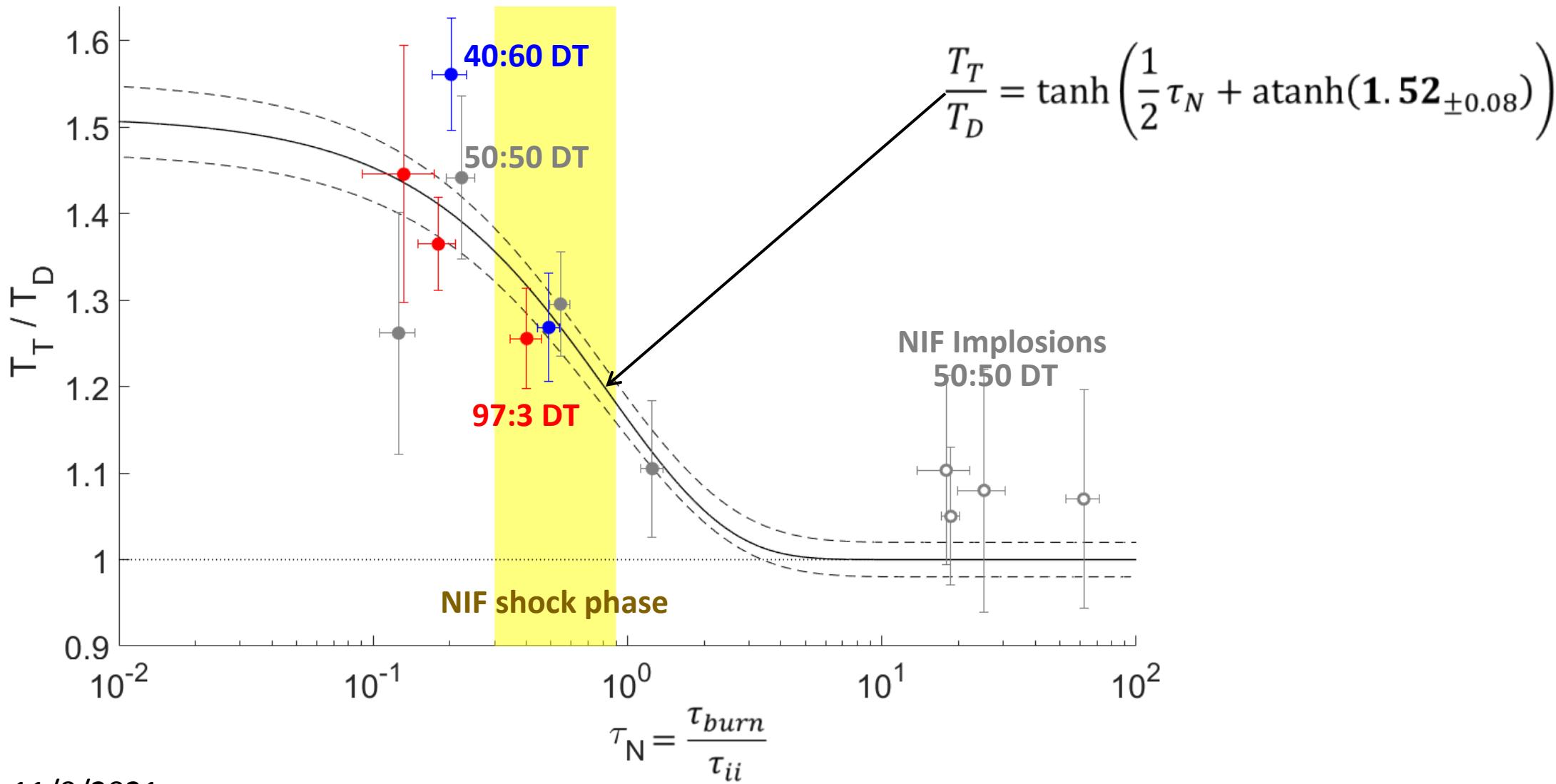
$$\frac{T_2}{T_1} = \tanh \left(\frac{1}{2} \frac{t}{\tau_{ii}} + \text{atanh}(R_0) \right), \quad R_0 = \left. \frac{T_2}{T_1} \right|_{t=0}$$

For each specific implosion we will average and take $\frac{t}{\tau_{ii}} \approx \tau_N = \frac{\tau_{burn}}{\tau_{ii}}$

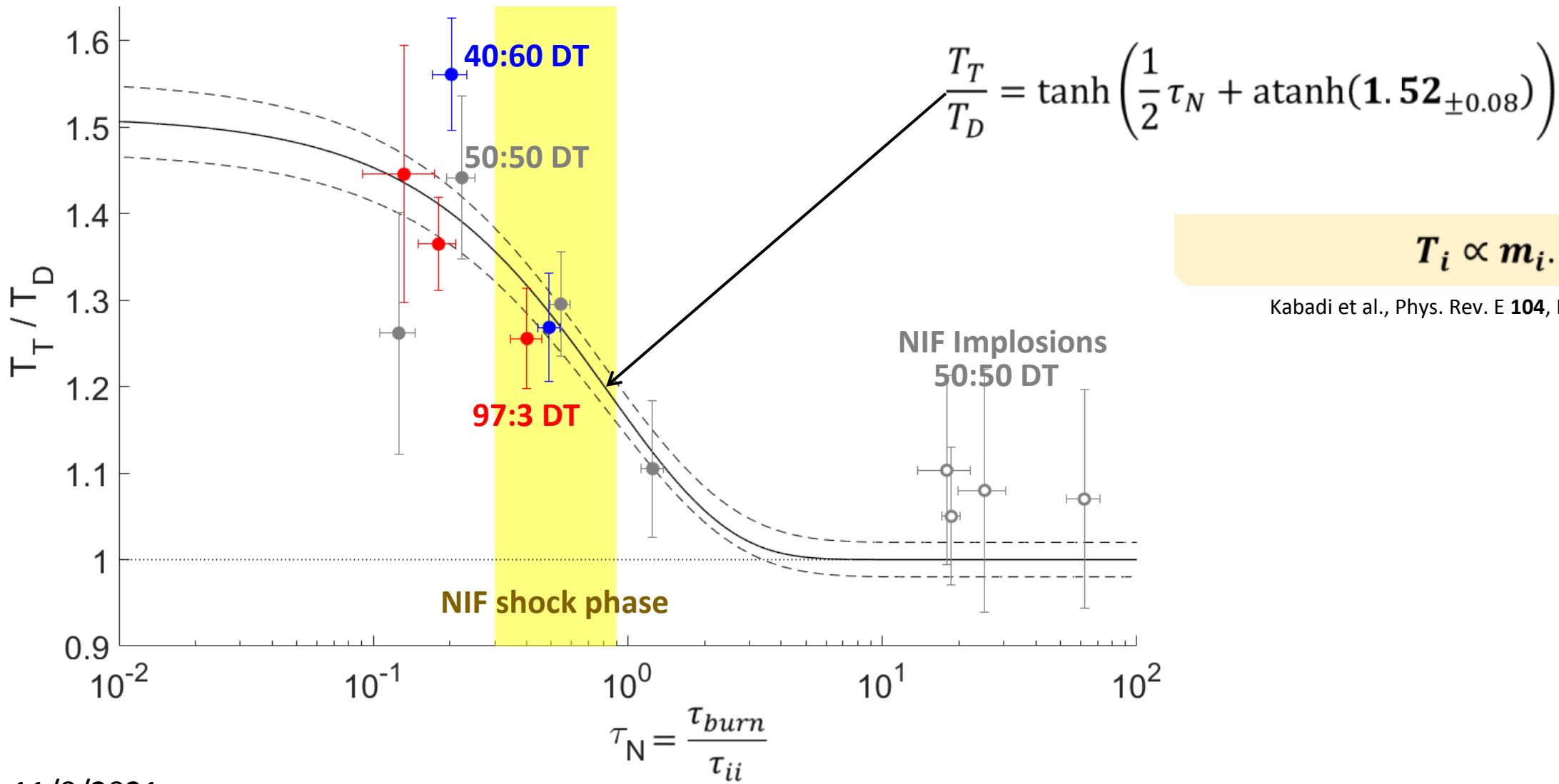
An initial temperature ratio of 1.5 is inferred from the data confirming long standing theory that shocks couple proportional to mass in multi-ion plasmas



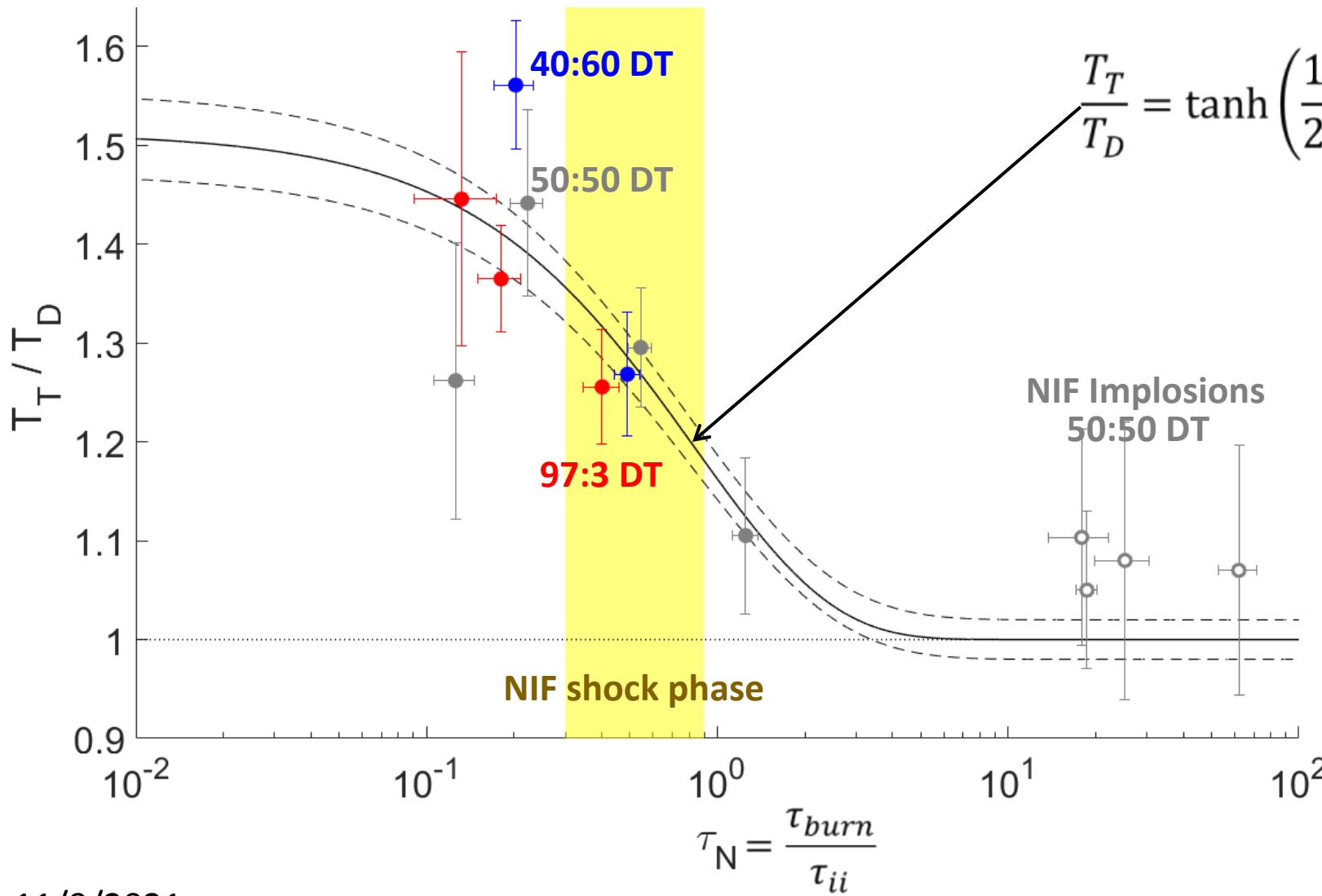
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$$\frac{T_T}{T_D} = \tanh\left(\frac{1}{2}\tau_N + \operatorname{atanh}(1.52_{\pm 0.08})\right)$$

This is the first lab-based measurement of this mass dependence corroborating recent astrophysical findings published in **Nature Astronomy***

Miceli et. al. *Nature Astronomy* 3, no. 3 (March 2019): 236–41.

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With multiple ion temperatures, calculation of the ion viscosity with average ion properties is incorrect

$$\mu = 1.01\text{E-15} \frac{T_i^{5/2} A^{1/2}}{Z^4 \ln \Lambda} \text{ g cm}^{-1} \text{ s}^{-1}$$

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Fill density mg/cc	τ_N	T ratio	T_{DTn} keV	μ average ion $\text{g cm}^{-1} \text{ s}^{-1}$	μ multi ion $\text{g cm}^{-1} \text{ s}^{-1}$	% error
4	1.1	1.1	8	3.1E-14	3.7E-14	-16%

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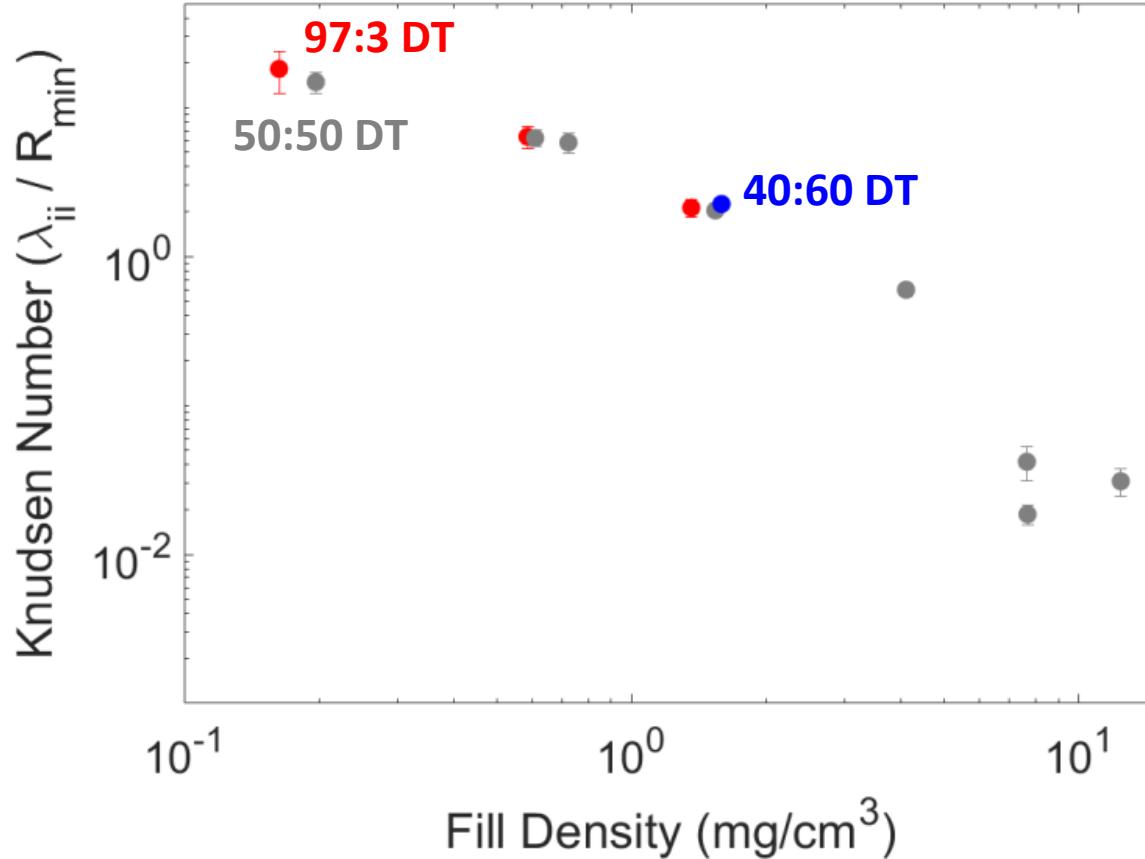
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0.2	0.1	1.45	21	1.6E-13	2.8E-13	-42%

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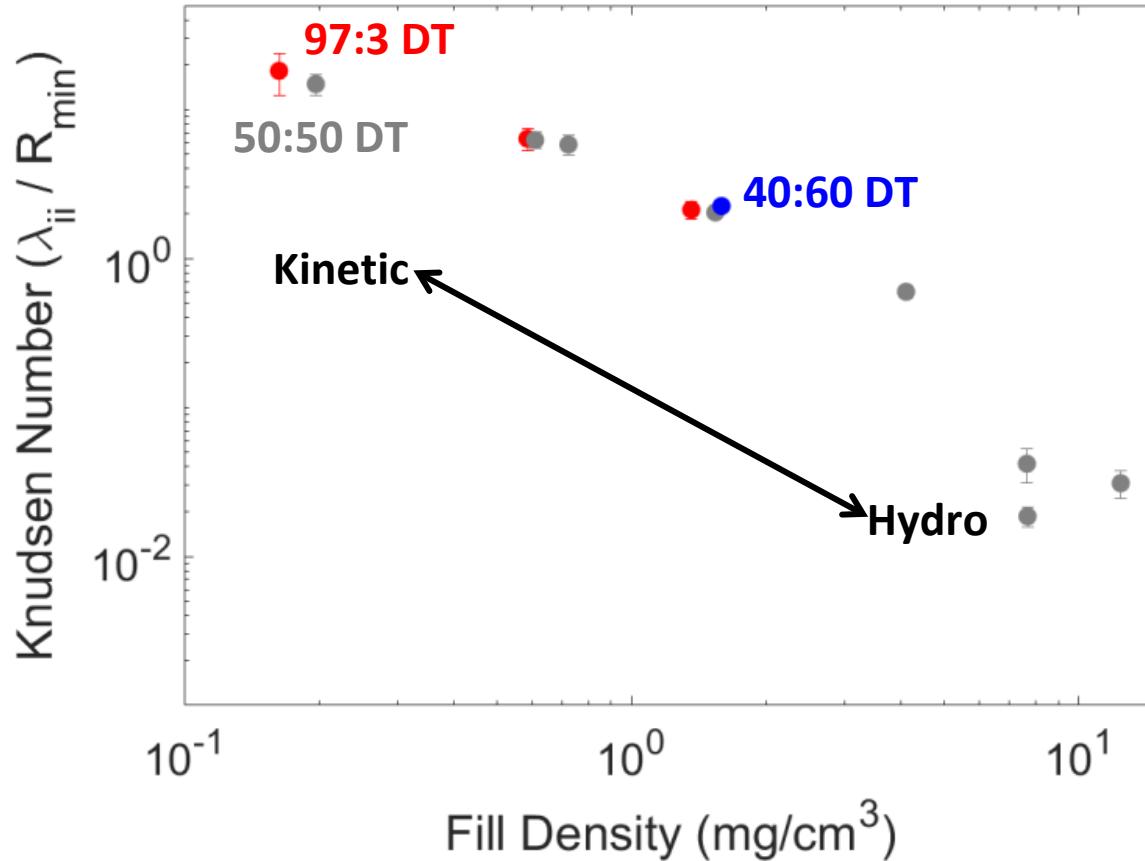
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Many of these implosions are in a regime where non-Maxwellian distribution functions are likely



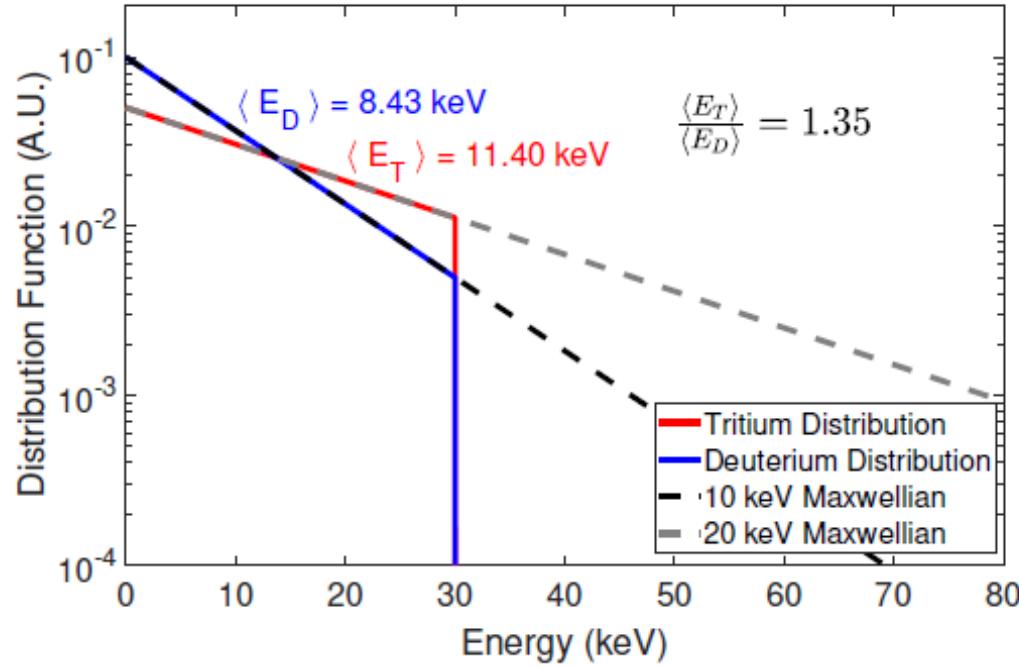
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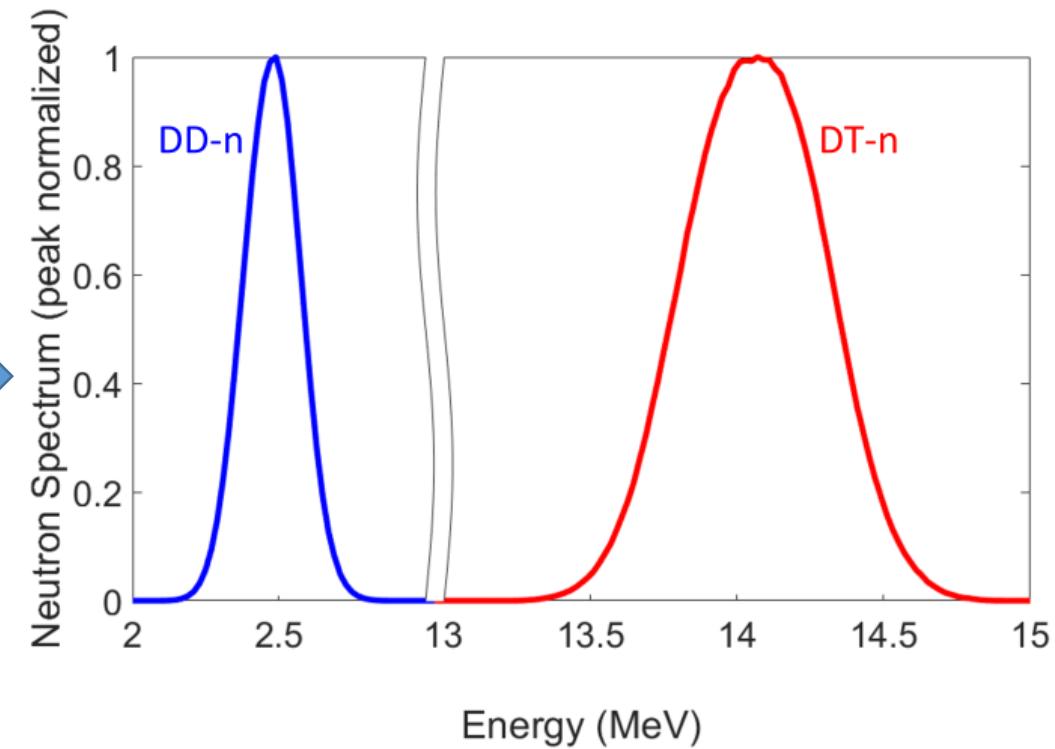
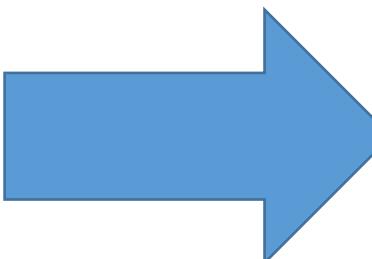
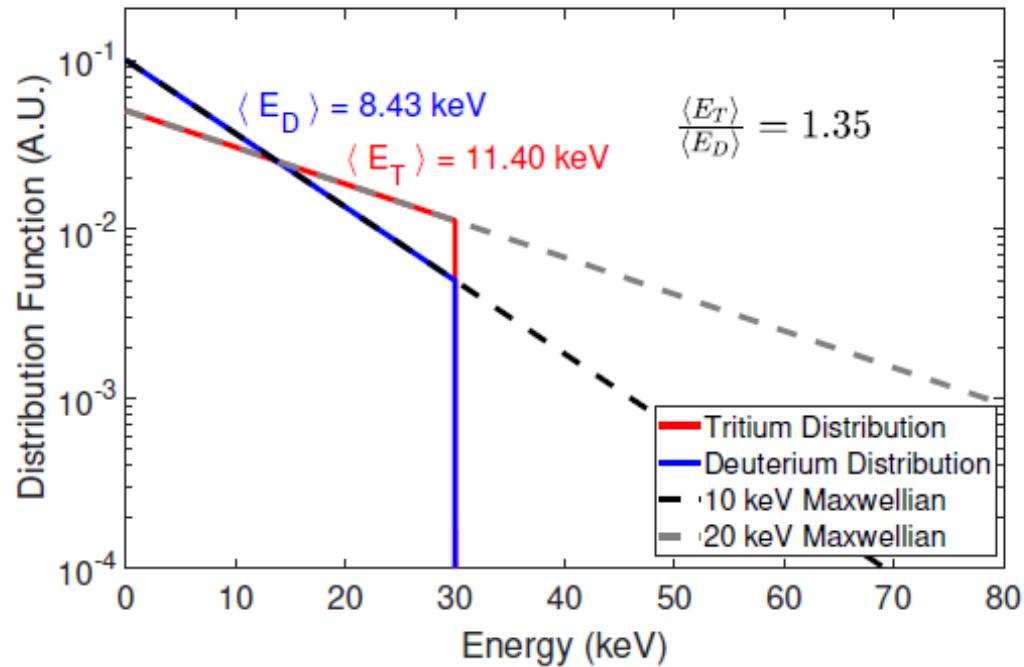


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To test the impact of non-Maxwellian distribution functions on the inferred temperature ratio, a series of Monte-Carlo calculations were performed

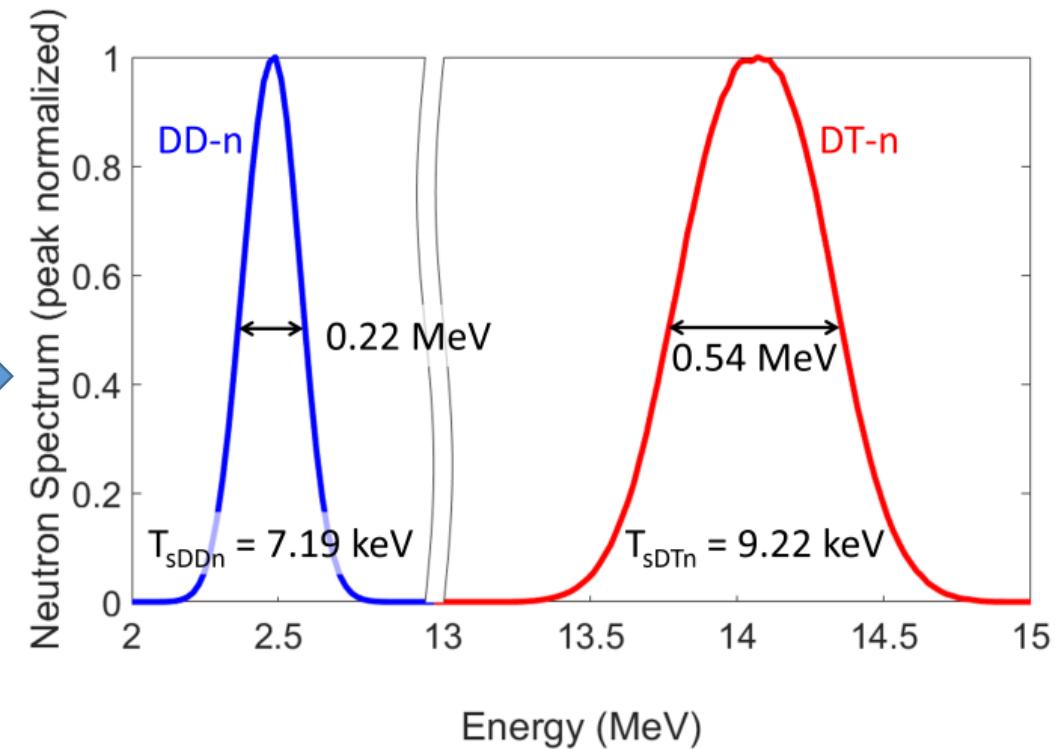
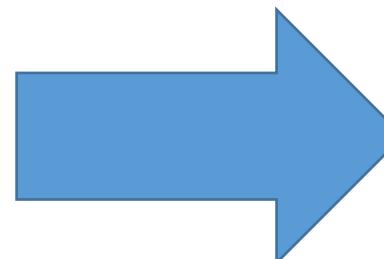
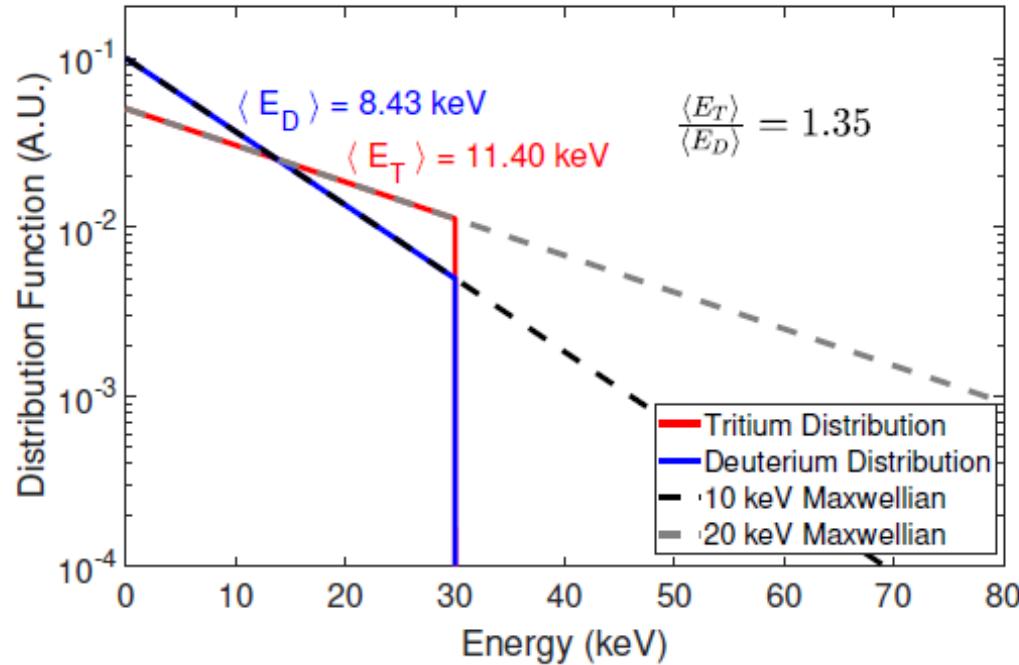


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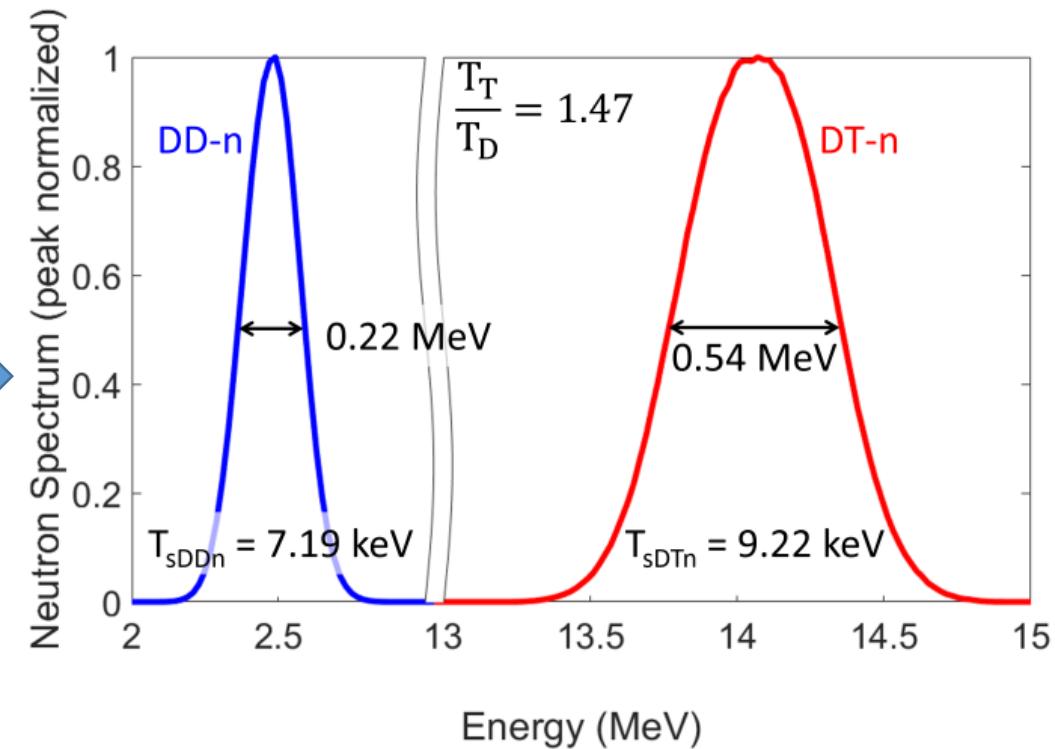
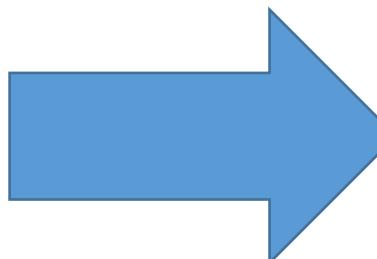
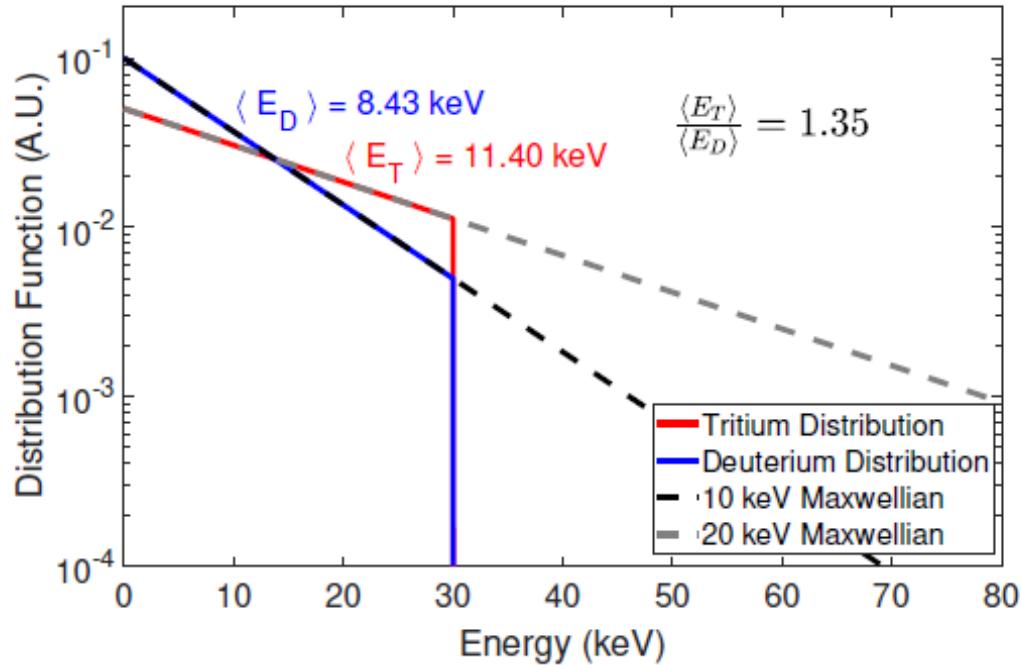
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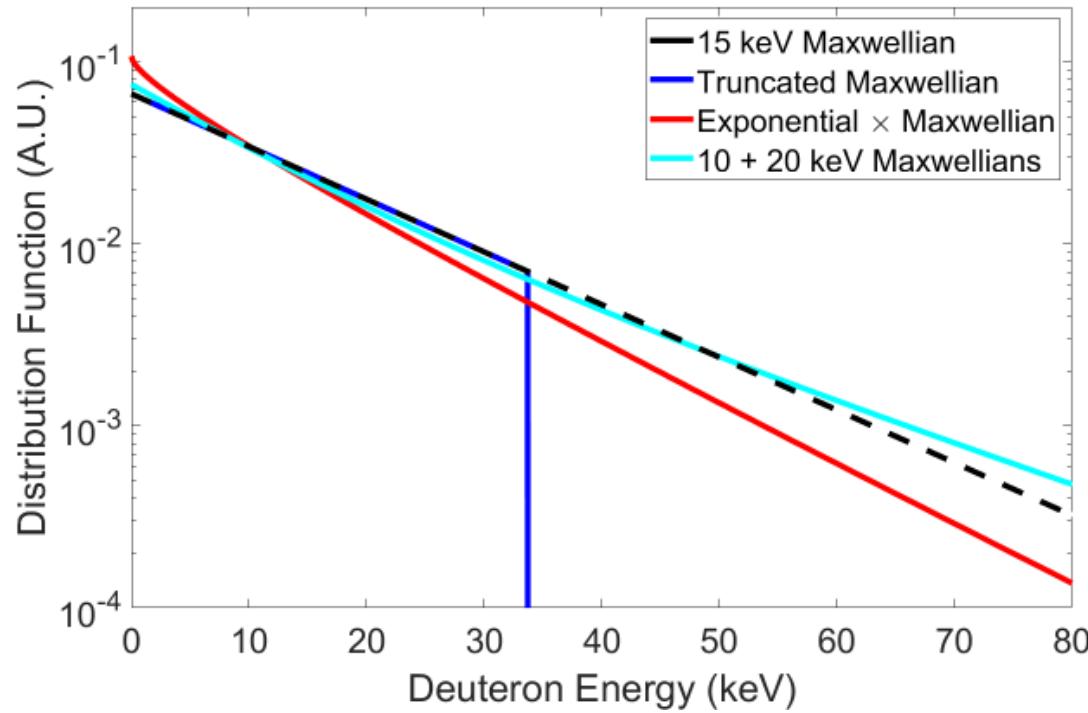
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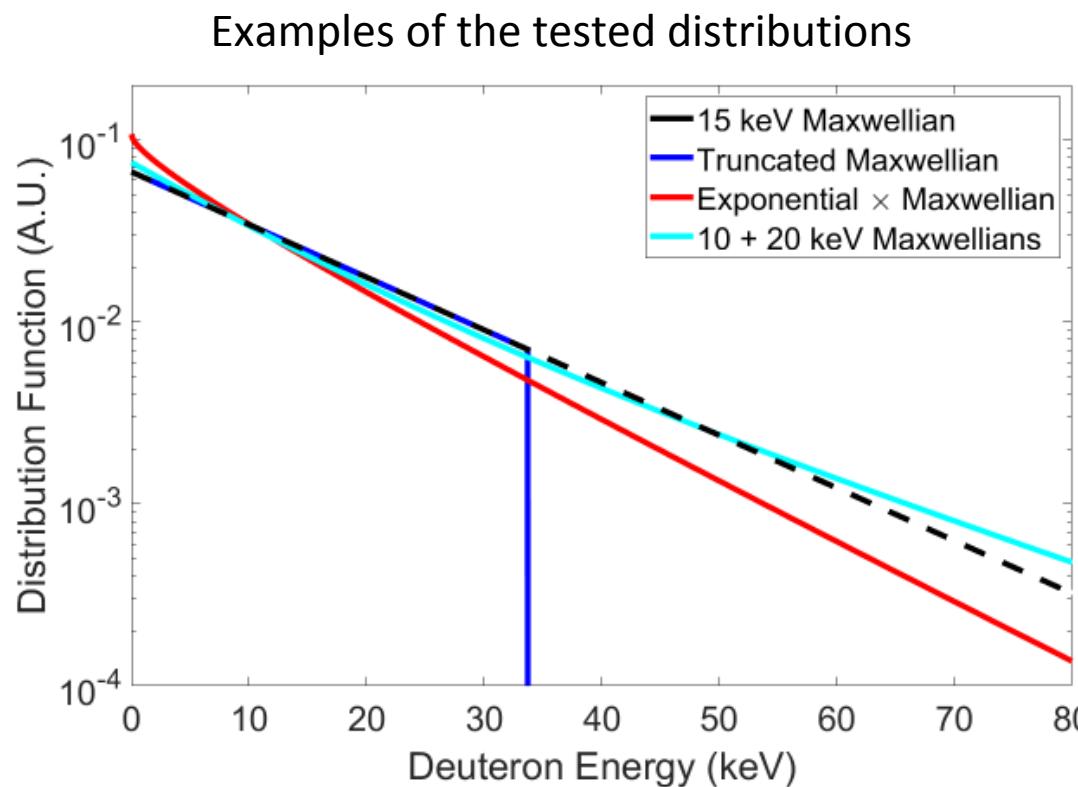
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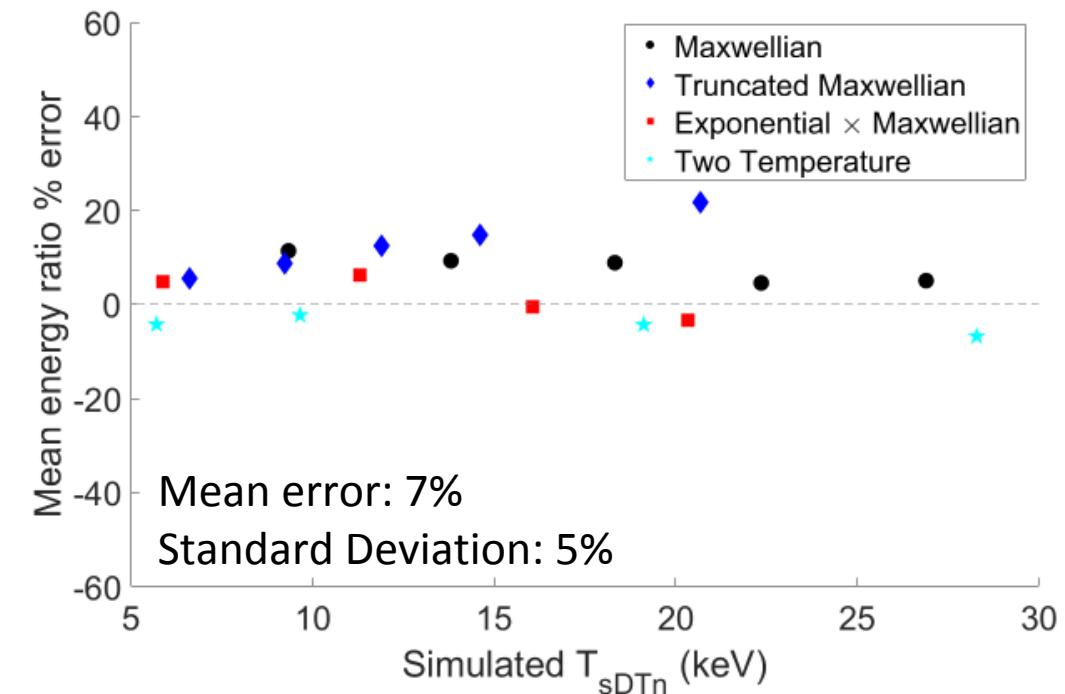
Examples of the tested distributions



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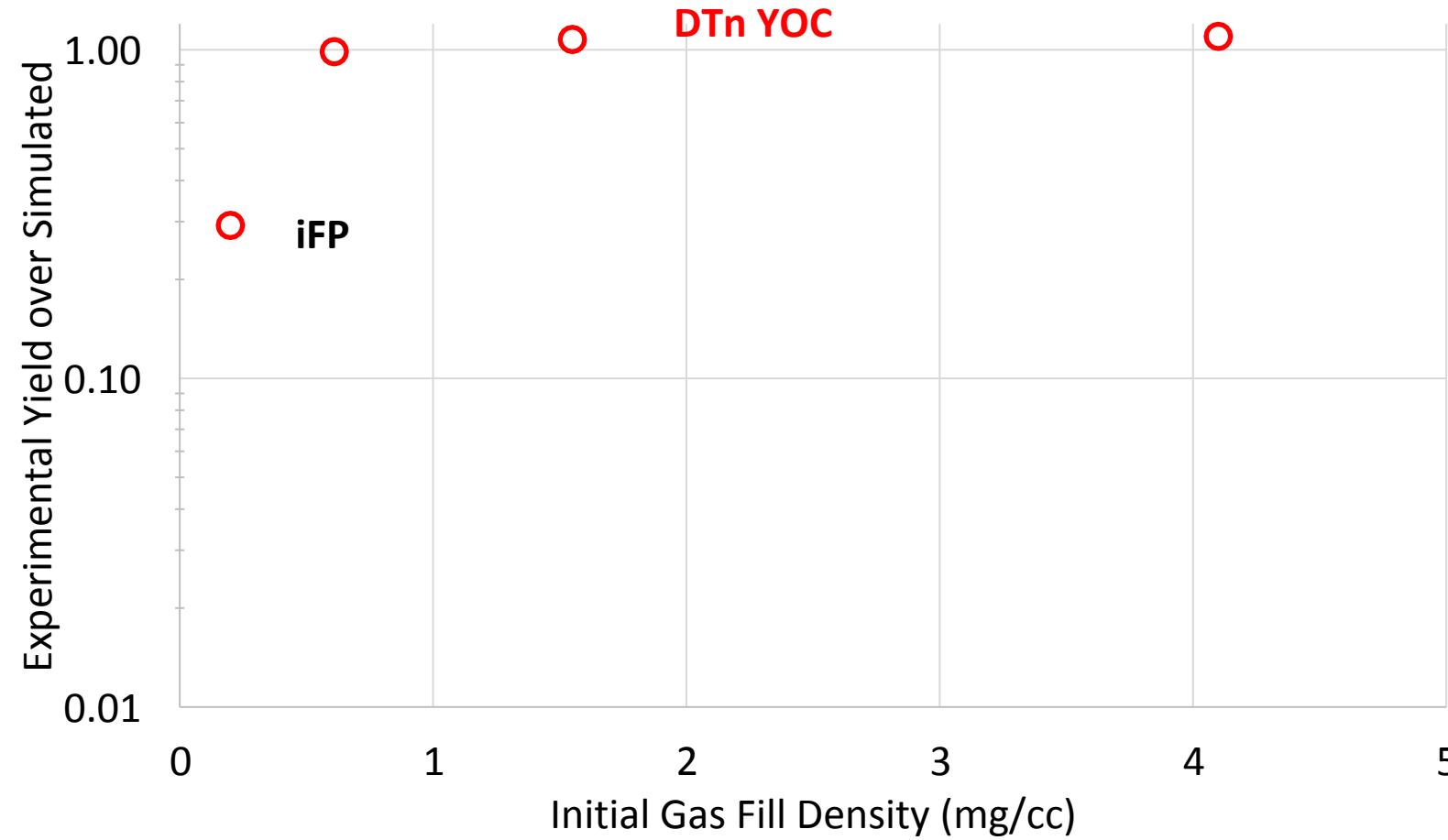


There is large error on the absolute temperature, but when taking a ratio of apparent D and T ion temperatures the error is reduced.

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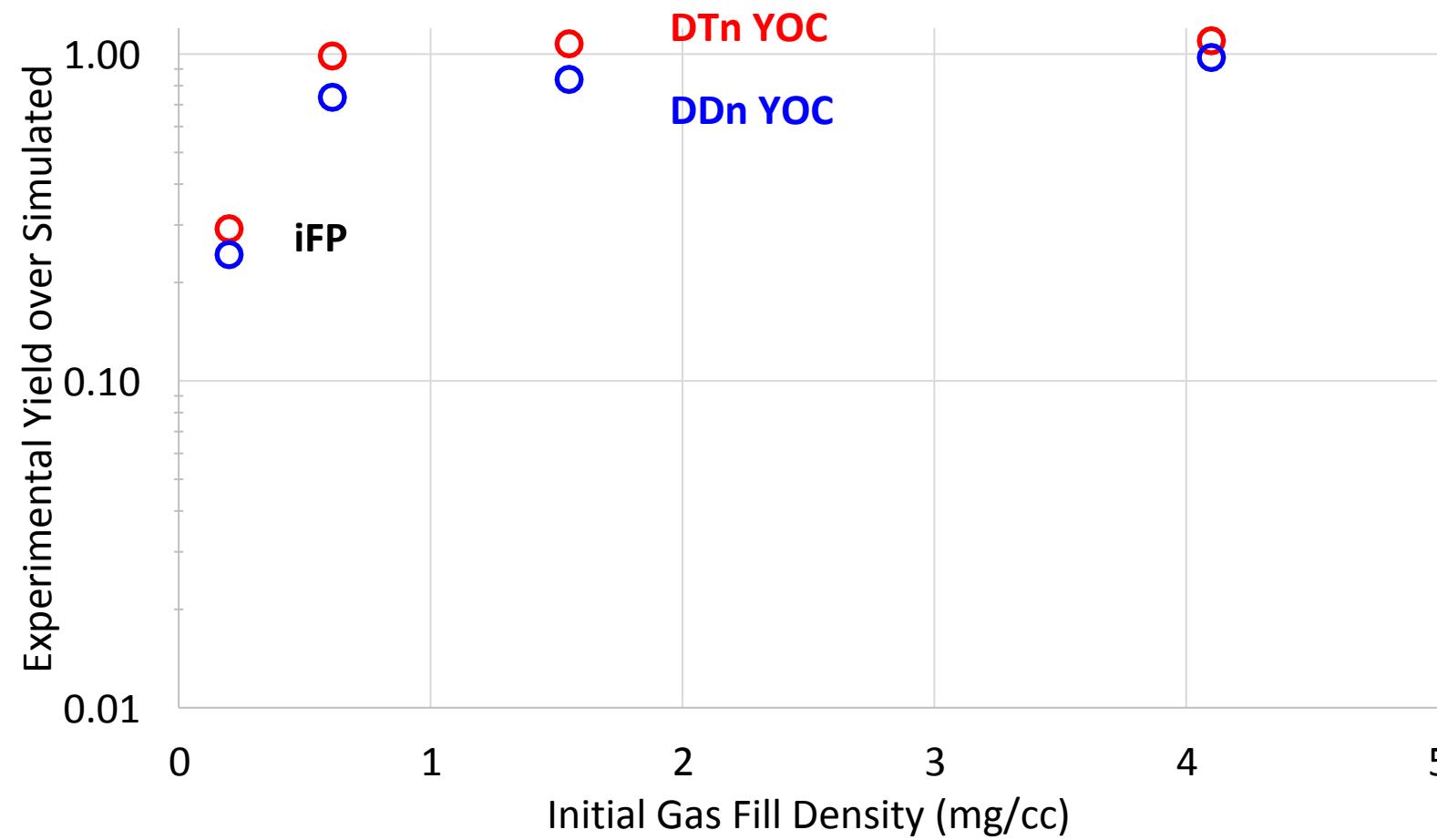
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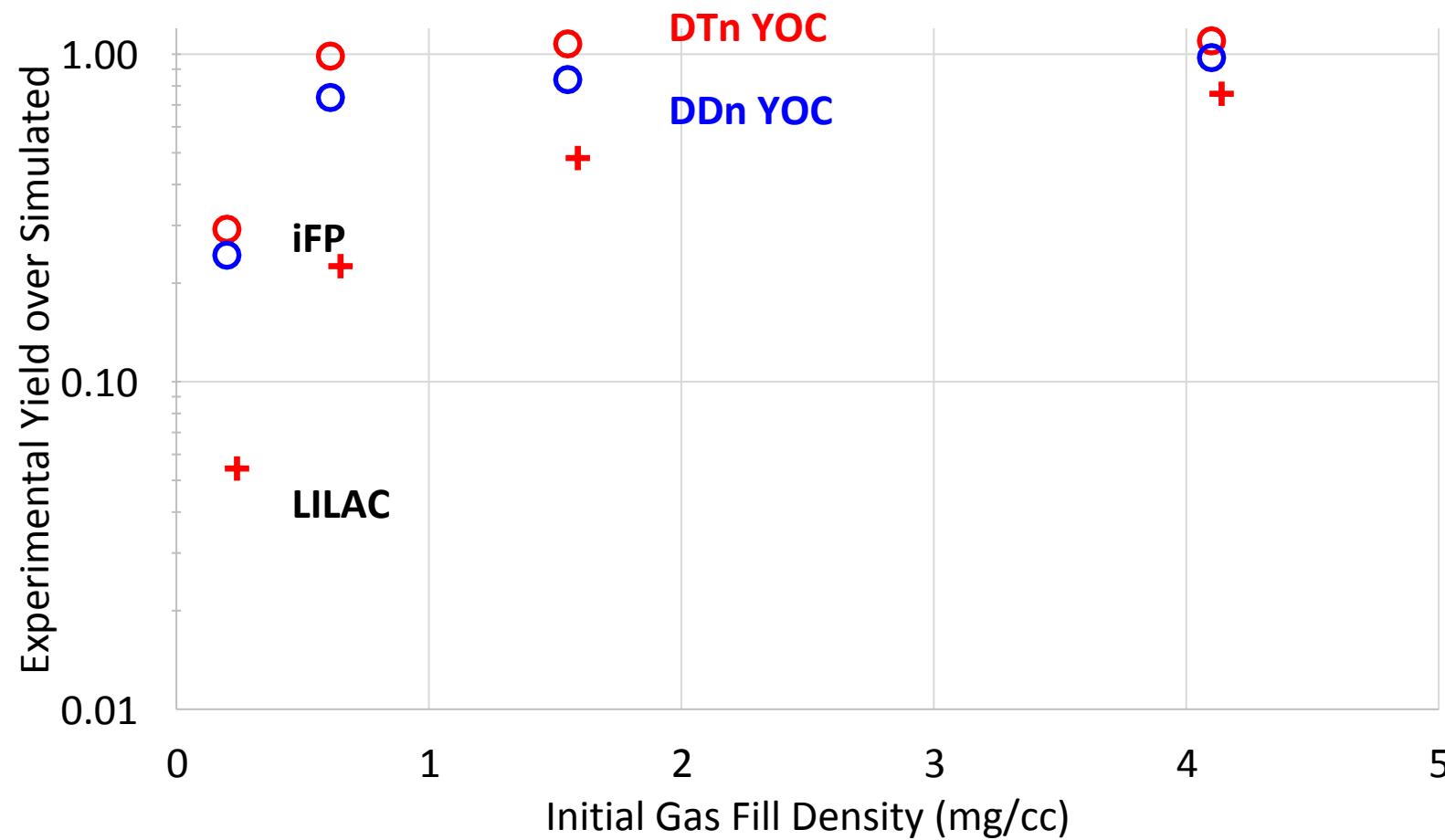
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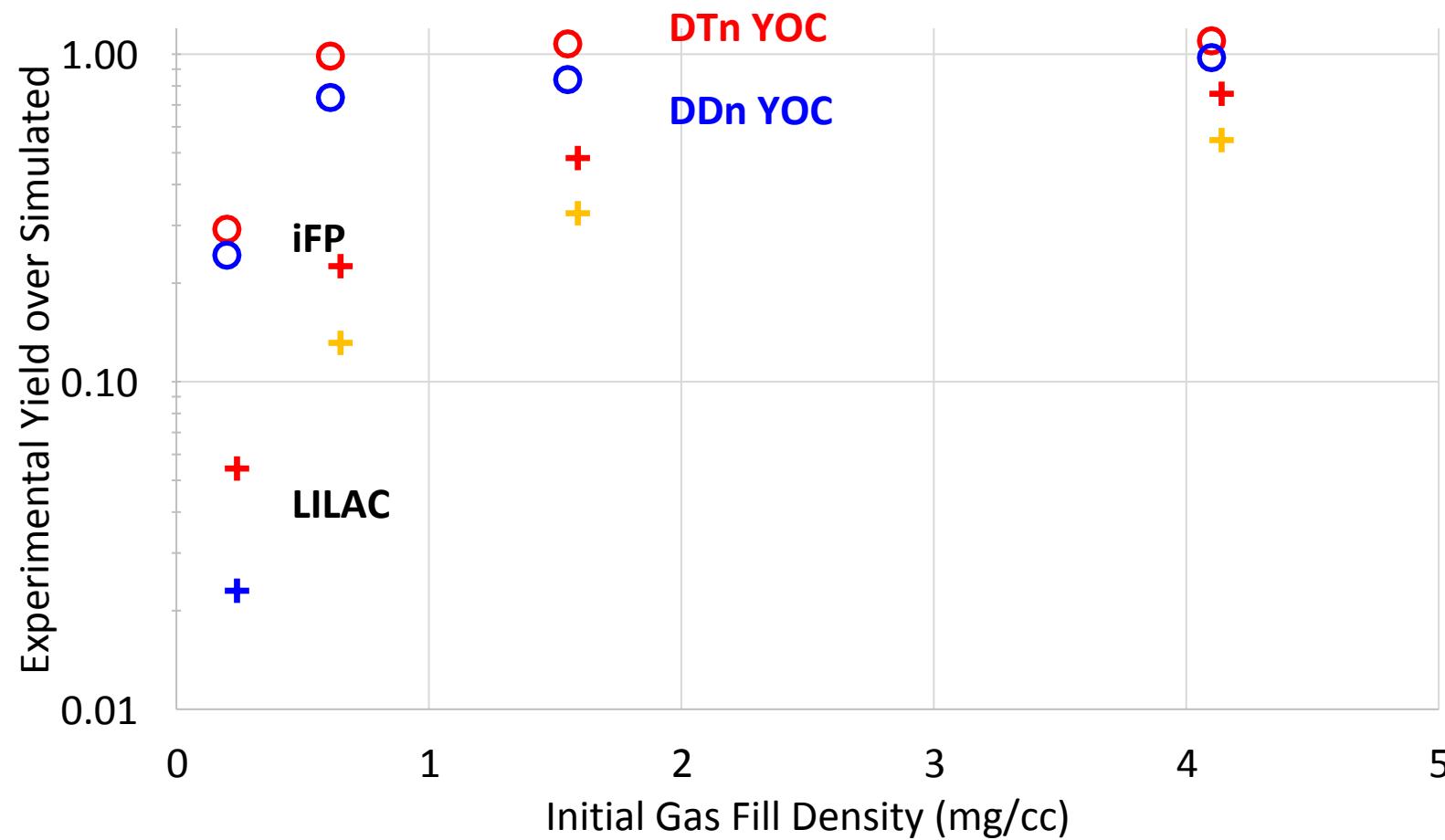
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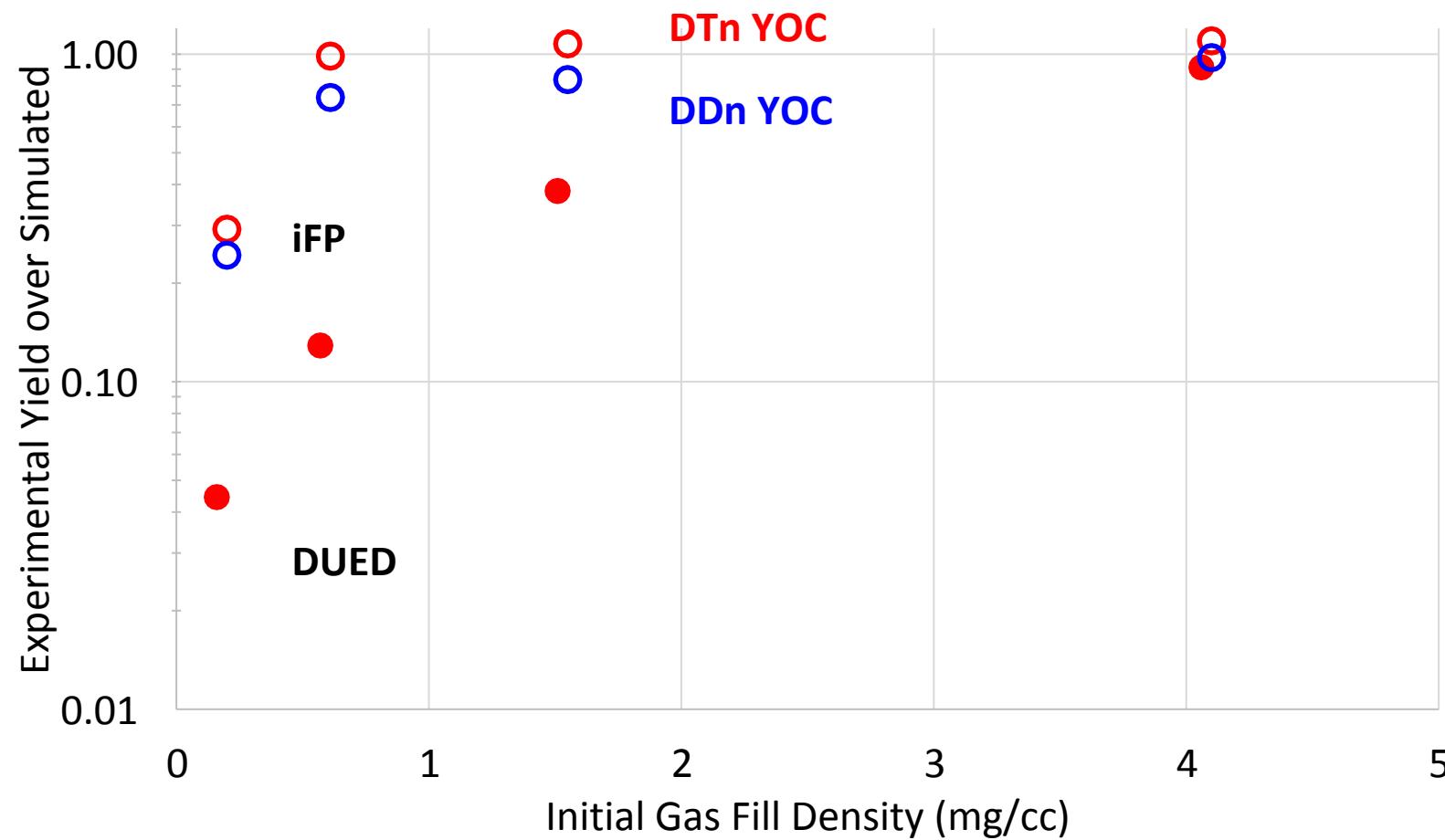
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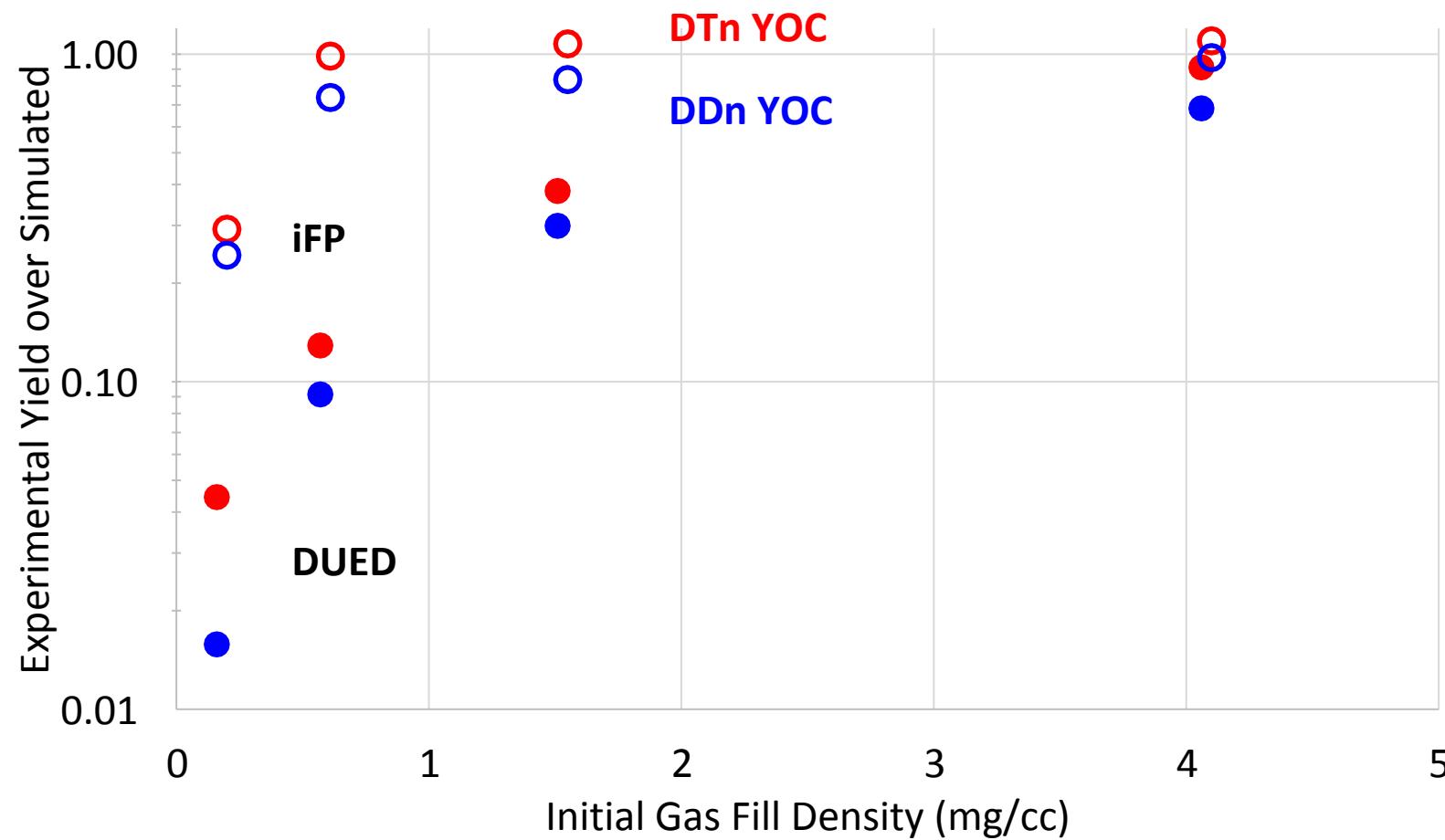
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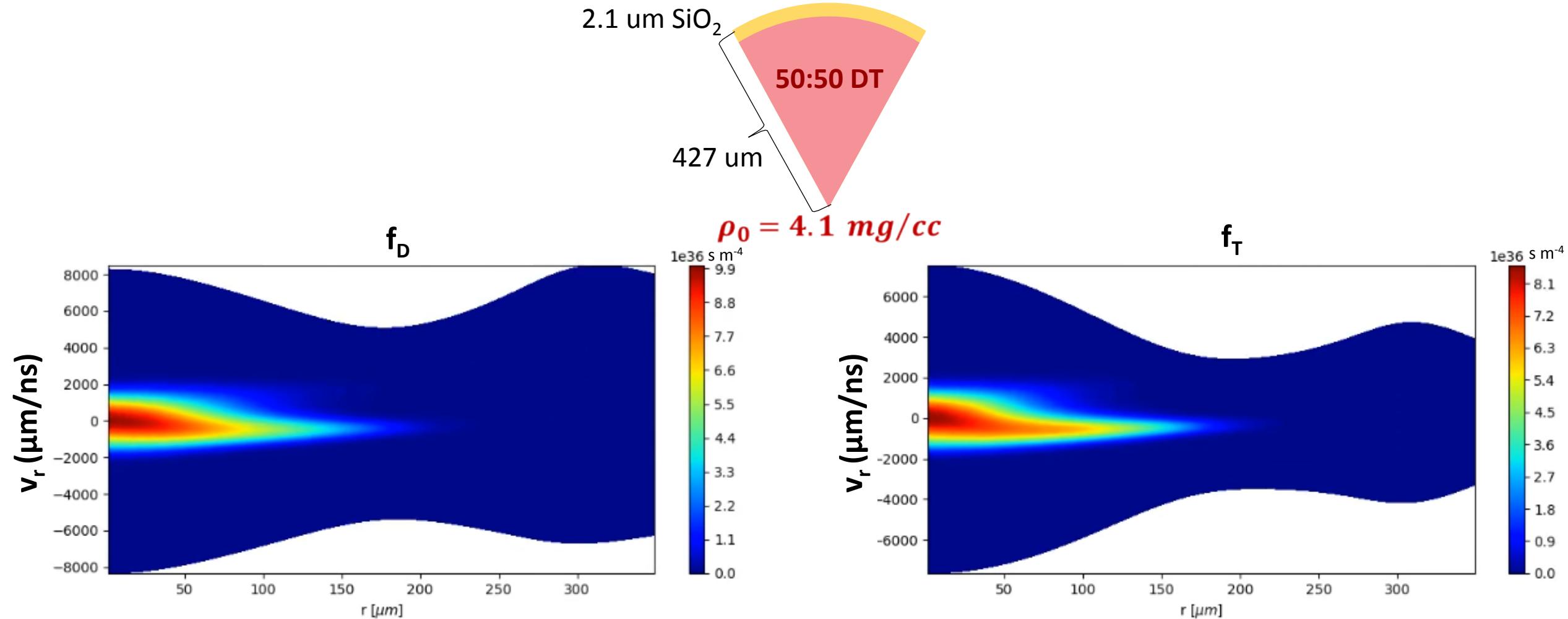
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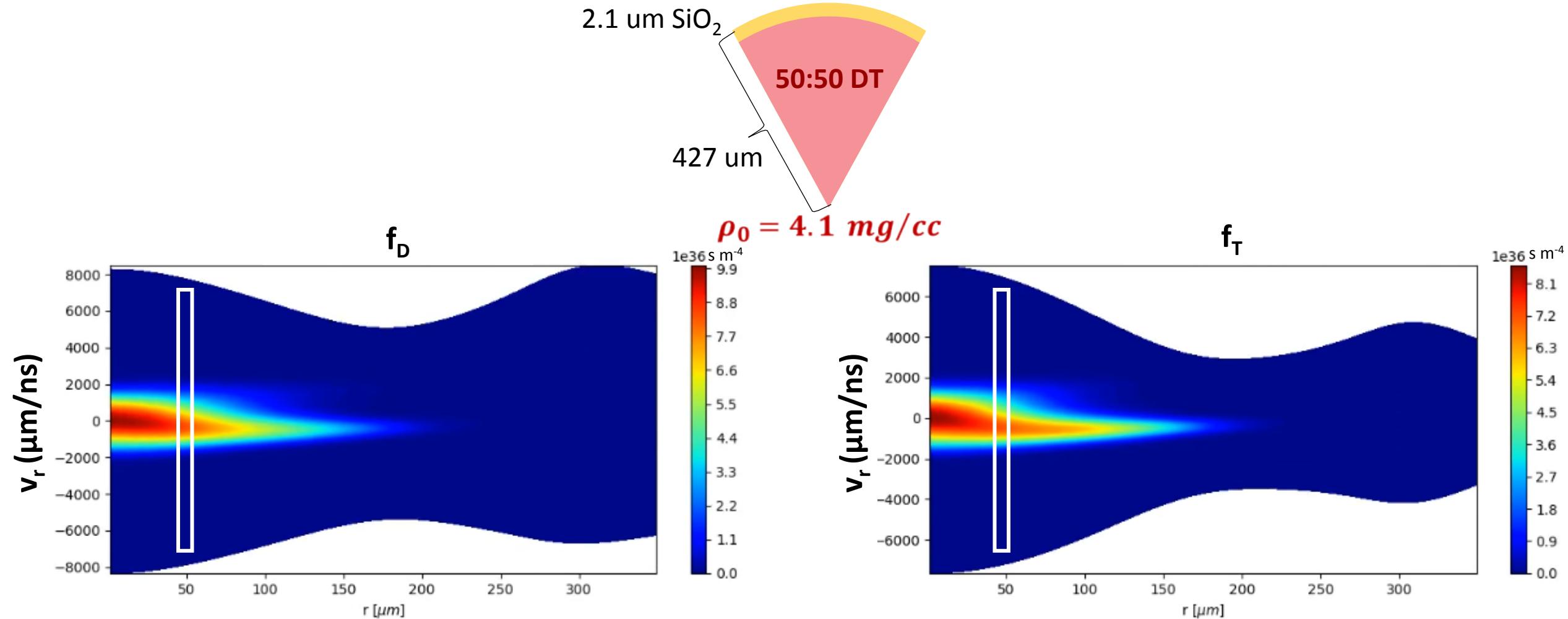
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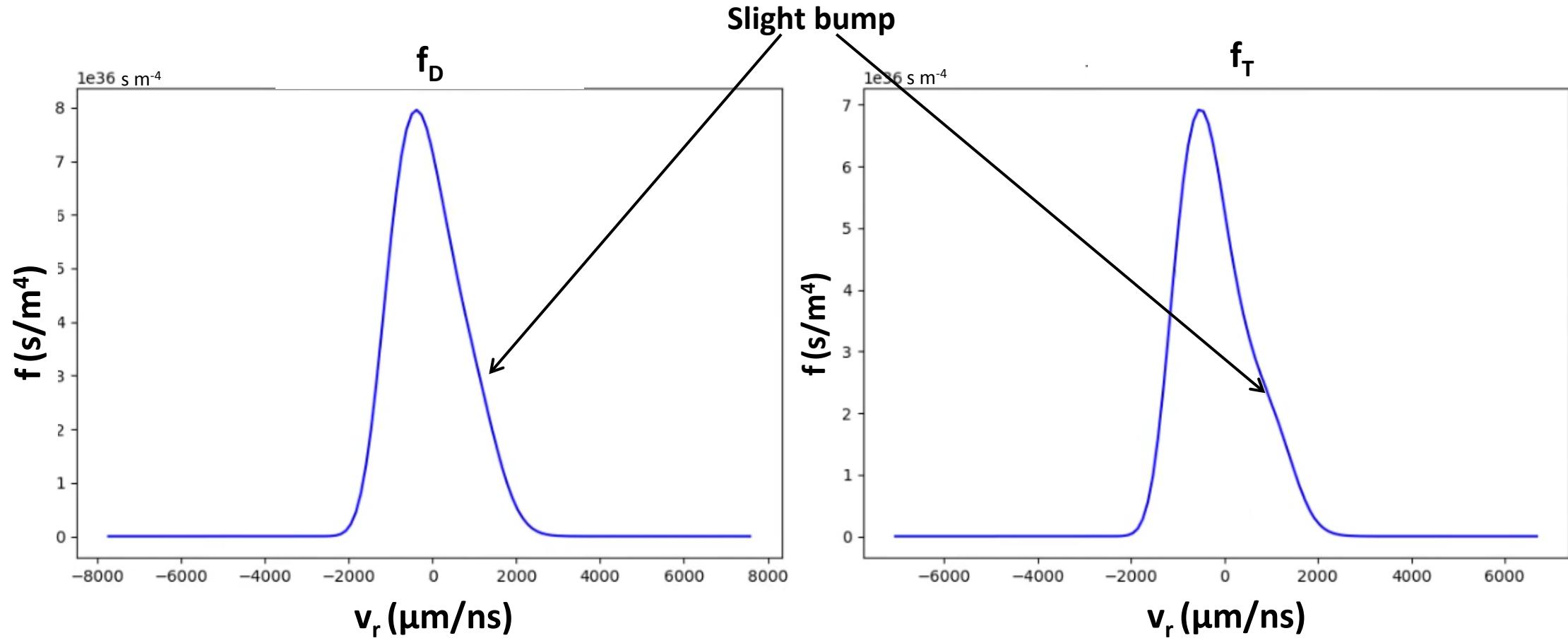
These are the bang time distribution functions integrated over the orthogonal velocity coordinate.

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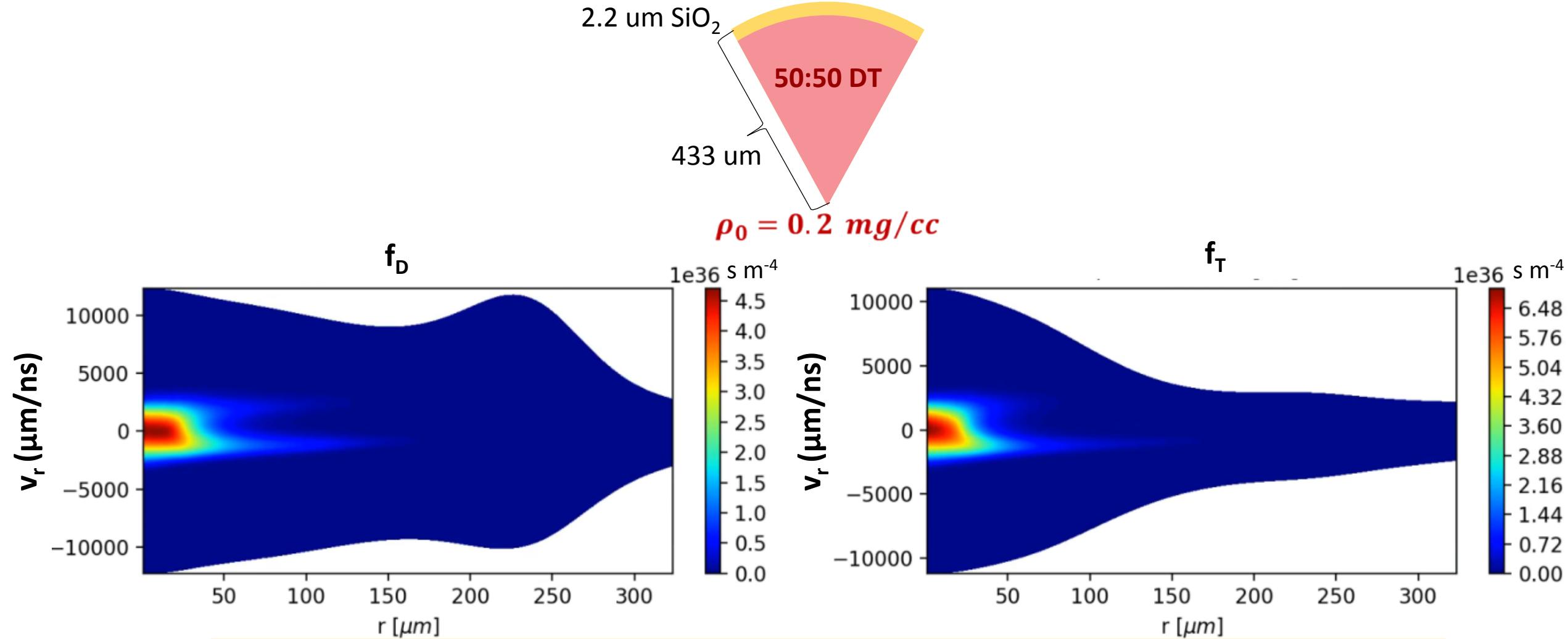
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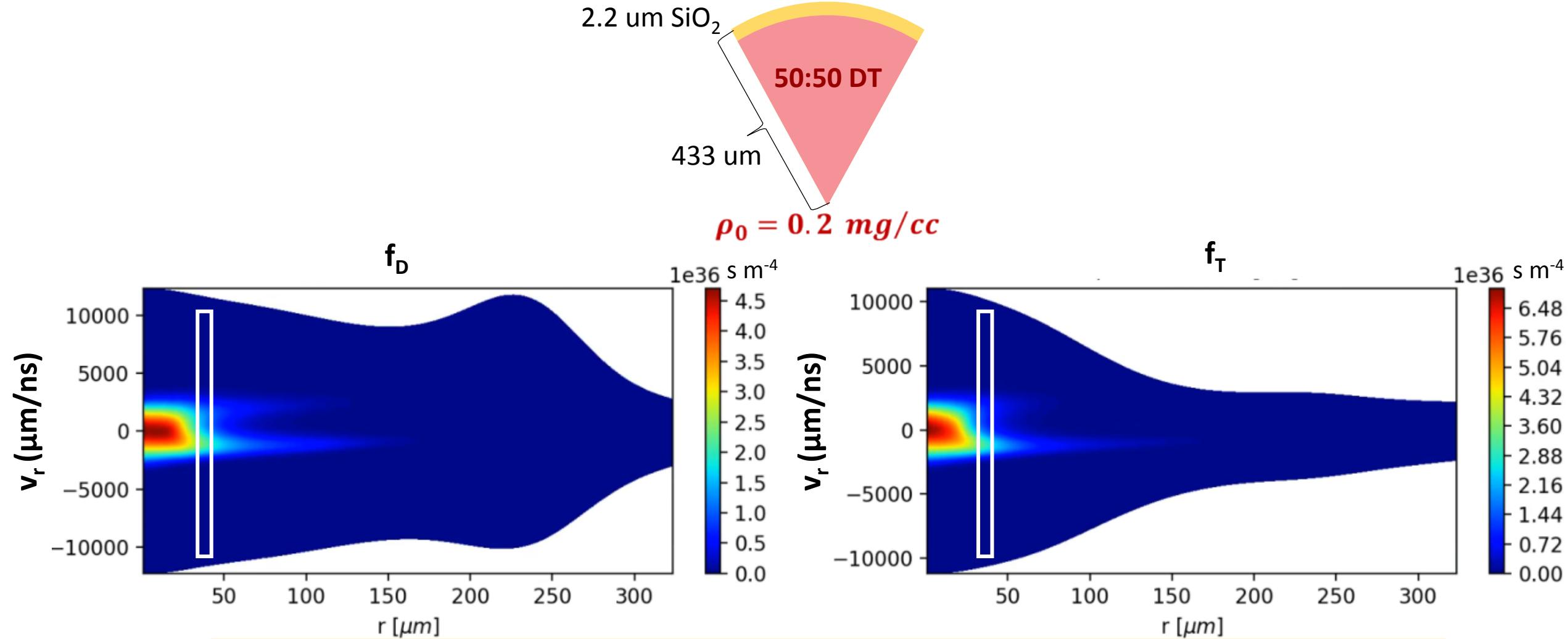
For clarity line-outs at the radius of maximal DTn emission are shown

iFP simulation of the 0.2 mg/cc fill density implosion shows two distinct ion populations



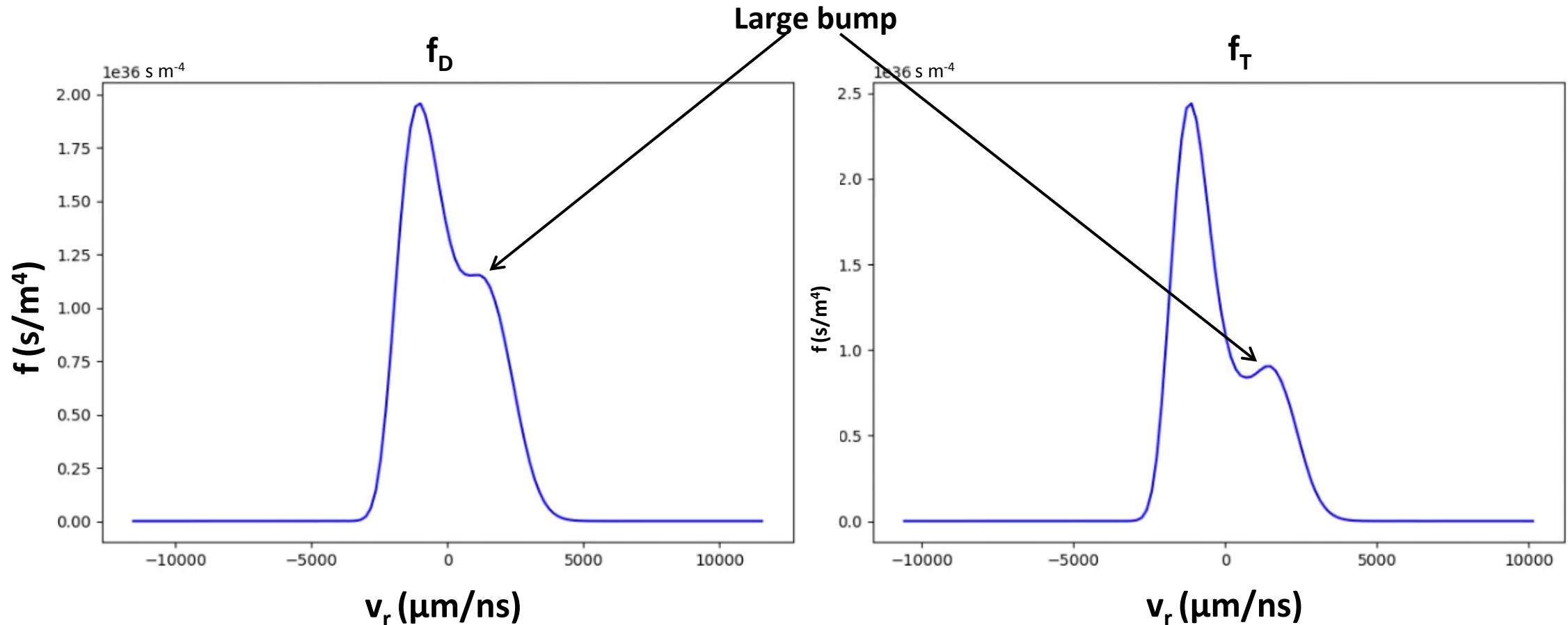
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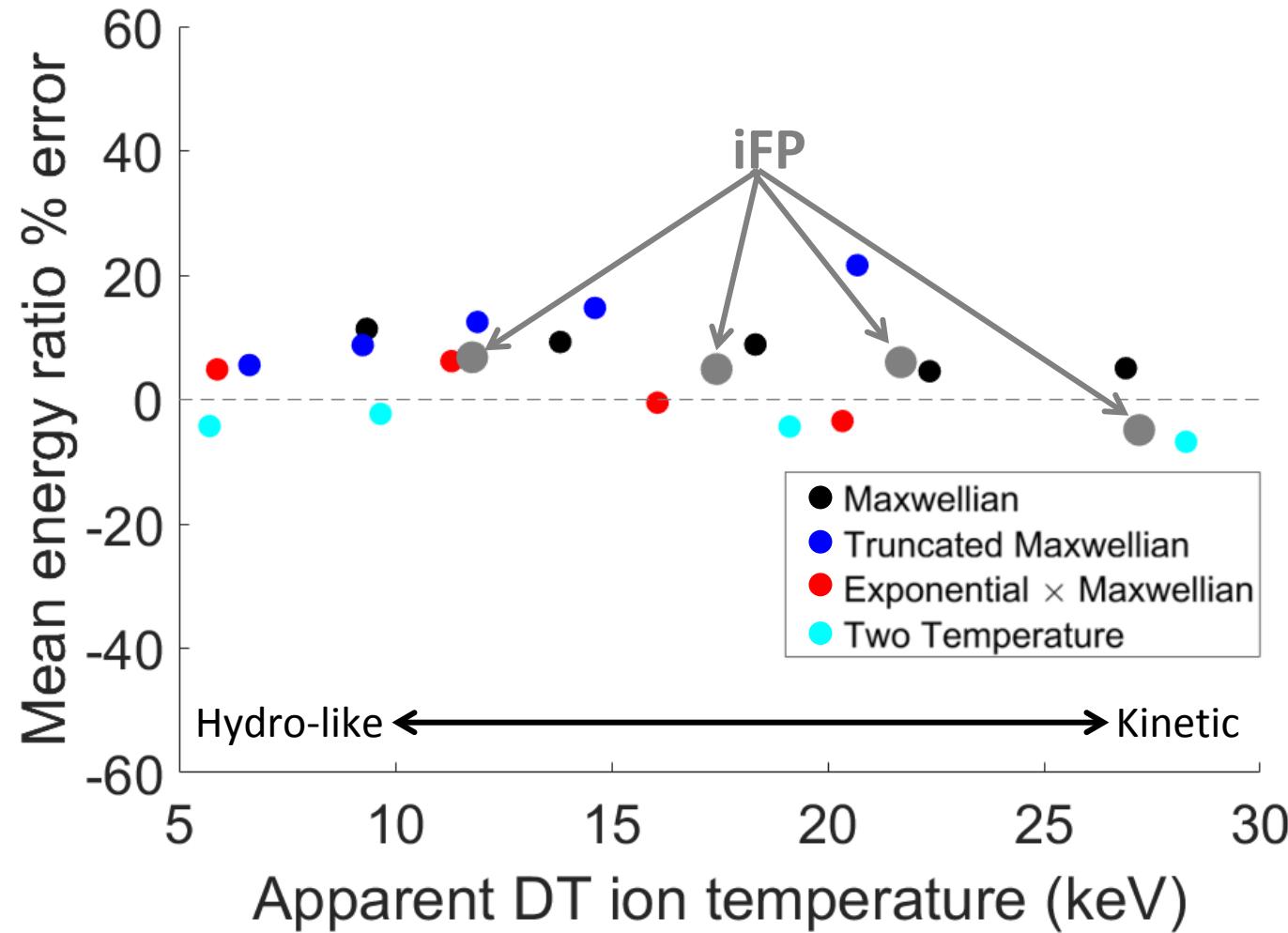
These are the bang time distribution functions integrated over the orthogonal velocity coordinate.

iFP simulation of the 0.2 mg/cc fill density implosion shows two distinct ion populations



For clarity line-outs at the radius of maximal DTn emission are shown

The apparent temperature ratio computed from the iFP 1D2V distribution functions is close to the true mean energy ratio



Future work

- Understand why hydro theory does a good job predicting the equilibration trend for highly kinetic conditions
- Design experiments to measure the impacts of thermal decoupling during shock convergence on the subsequent compression

The observed difference between T_{DTn} and T_{DDn} is explained by shock coupling proportional to mass, and equilibration consistent with hydro theory.

- This is the first direct lab-based experimental measurement of shocks coupling directly proportional to mass in multi-ion plasmas, verifying a long standing prediction.
- For NIF ignition implosions, during the shock convergence phase the D and T ions are out of thermal equilibrium which greatly impacts the ion viscosity.
- iFP simulations indicate that this result is robust, even in the highly kinetic regime.

Questions?

The Brysk form for the temperature mass weighting is only accurate for DT implosions when the temperature ratio is not too large

$$T_{DTn} = \frac{m_T T_T + m_D T_D}{m_T + m_D}$$

$$\begin{array}{c} \boxed{} \\ \boxed{} \end{array} \quad = \quad \begin{array}{c} \boxed{} \\ \boxed{} \end{array}$$

$$T_T = T_{DTn} + \frac{m_D}{m_T} (T_{DTn} - T_{DDn})$$

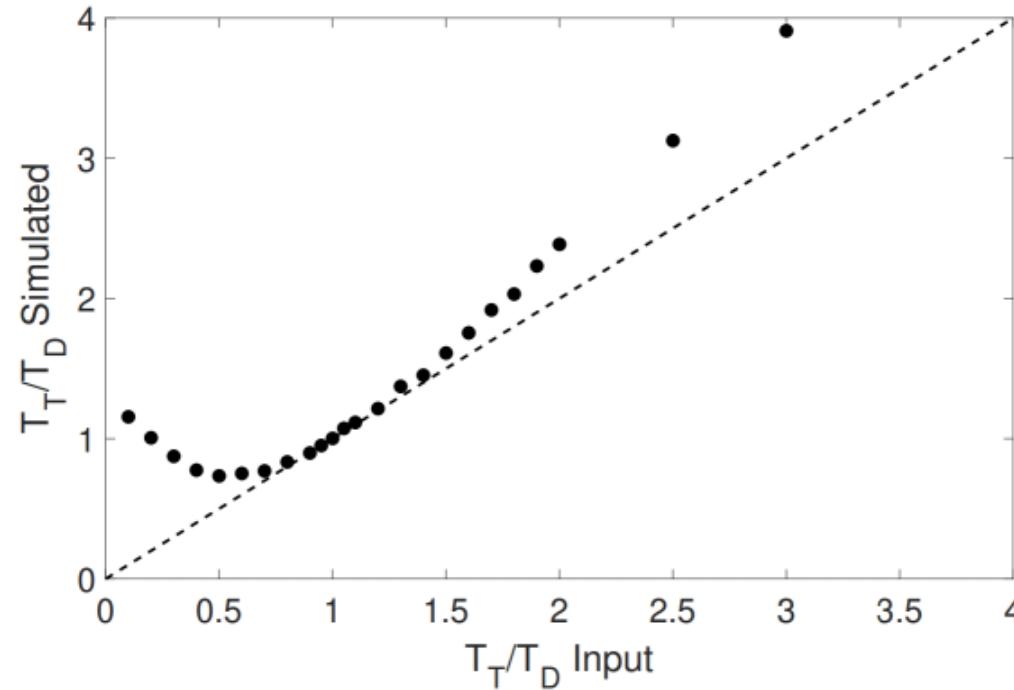
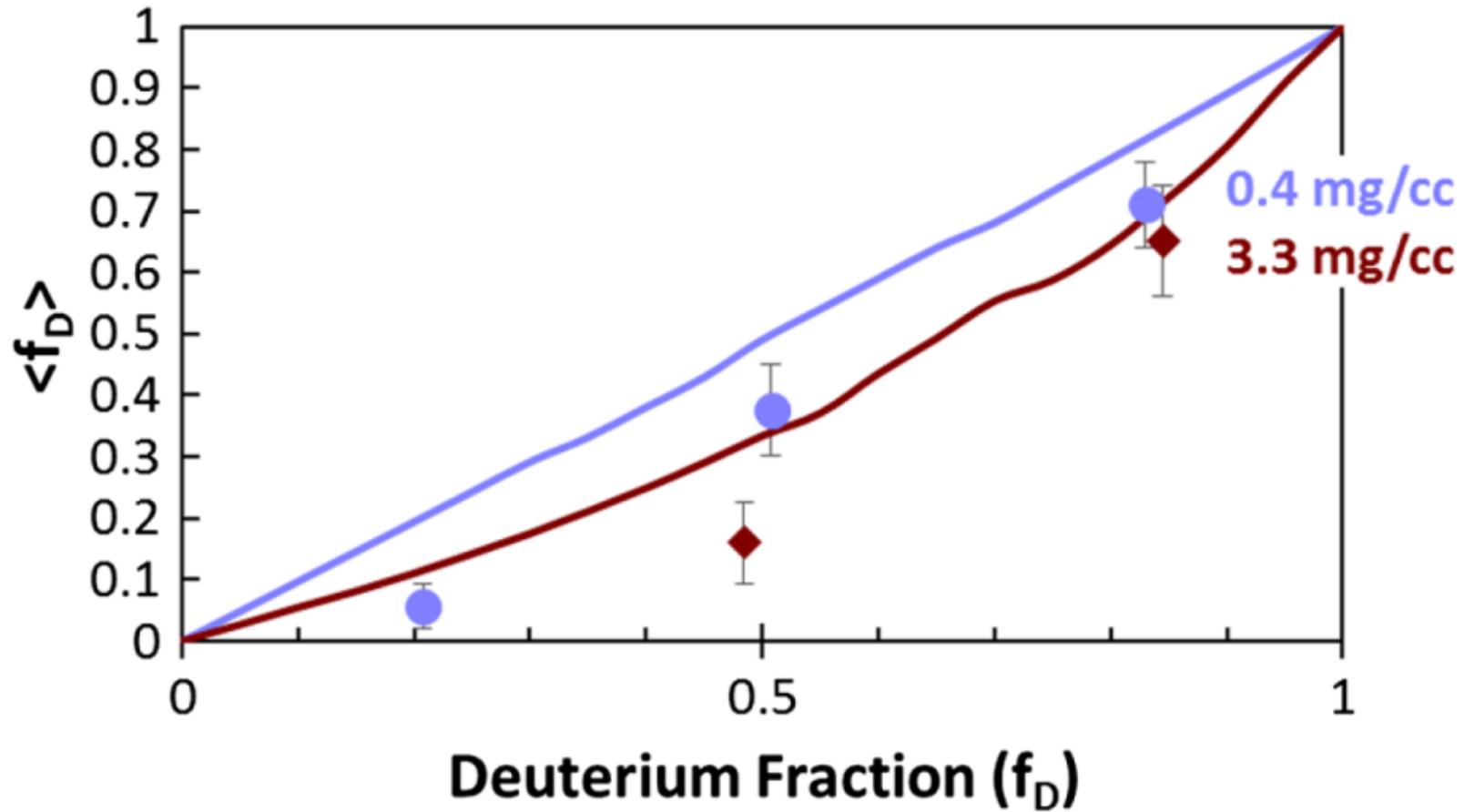


Figure 4. Black points show the temperature ratios inferred from Monte Carlo simulated neutron spectra plotted versus the input temperature ratio. The dashed black line is a reference showing equality between input and simulated values.

Species separation was previously observed in D³He gas-filled shock driven implosions



These experiments used the same OMEGA platform and consistently inferred suppression of the D content in the core during nuclear emission.

The level of species separation can be inferred from the combined yields and ion temperatures taking into account profiles to second order

DT – flat reactivity ratio – Uniform model –

$$\langle f_D \rangle = \frac{\frac{Y_{DDn}}{Y_{DTn}}}{\frac{Y_{DDn}}{Y_{DTn}} + \frac{1}{2} \frac{\sigma v_{DDn}(T_{rDD})}{\sigma v_{DTn}(T_{rDT})}}$$

$$\frac{\langle f_D \rangle}{\langle f_{^3He} \rangle} = 2R_o \frac{Y_{DDn}}{Y_{D^3He}} \left[1 + \frac{1}{4} \frac{\delta^2 R_o}{\delta T^2} \frac{T_{sD^3He} - BT_{sDDn}}{B \frac{\delta R_o}{\delta T}} \right]$$

D³He – steep reactivity ratio – second order model* –

$$B = \frac{m_{^3He} + m_D * \frac{T_{^3He}}{T_D}}{m_D + m_{^3He}}$$

From f_D the mass concentration of the light species (c) can be computed as

$$c = \frac{2 * f_D}{2 * f_D + 3 * (1 - f_D)}$$

One potential source of species separation is multi-ion diffusion within the sharp gradients of the shock front

The diffusion equation can be written as:

$$\rho \frac{Dc}{Dt} = -\nabla \cdot i$$

- c is the light species mass concentration ($\frac{\rho_l}{\rho}$)
- i is the drift velocity of the light species relative to the center of mass

$$i = \rho_l(u_l - u)$$

$$I = -\rho D(\nabla c + k_p \nabla \log P_i + k_{Ti} \nabla \log T_i + k_{Te} \nabla \log T_e + k_E \frac{e}{T_i} \nabla \Phi)$$

noitsevde seg leebi ehit bue $\frac{\Phi \nabla}{e \pi \epsilon_0} \approx \Phi \nabla$ gnisu

$$I = -\rho D(\nabla c + k_p \nabla \log P_i + k_{Ti} \nabla \log T_i + k_{Te} \nabla \log T_e + k_E \nabla \log P_e)$$

Separation driven by gradients in T, and P is compared to homogenization driven by classical diffusion to estimate the concentration gradient

The strength of the separation can be estimated as:

$$\begin{aligned} L_s &= \frac{i}{\rho_l} * \tau_{shock} = (u_l - u) * \tau_{shock} \\ &= -\frac{D}{c} \frac{(k_p + k_{Ti} + k_{Te} + k_E)}{\Delta_{shock}} * \frac{\Delta_{shock}}{v_{shock}} \end{aligned}$$

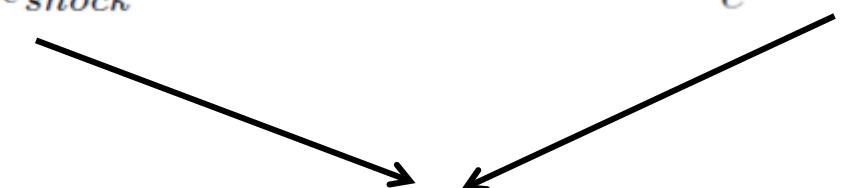
$$L_s = -\frac{D}{c} \frac{(k_p + k_{Ti} + k_{Te} + k_E)}{v_{shock}}$$

Separation driven by gradients in T, and P is compared to homogenization driven by classical diffusion to estimate the concentration gradient

The strength of homogenization can be estimated as:

$$L_r = (u_l - u) * \tau_{imp} = -\frac{D}{c} \frac{R_0}{v_{shock}} \nabla c$$

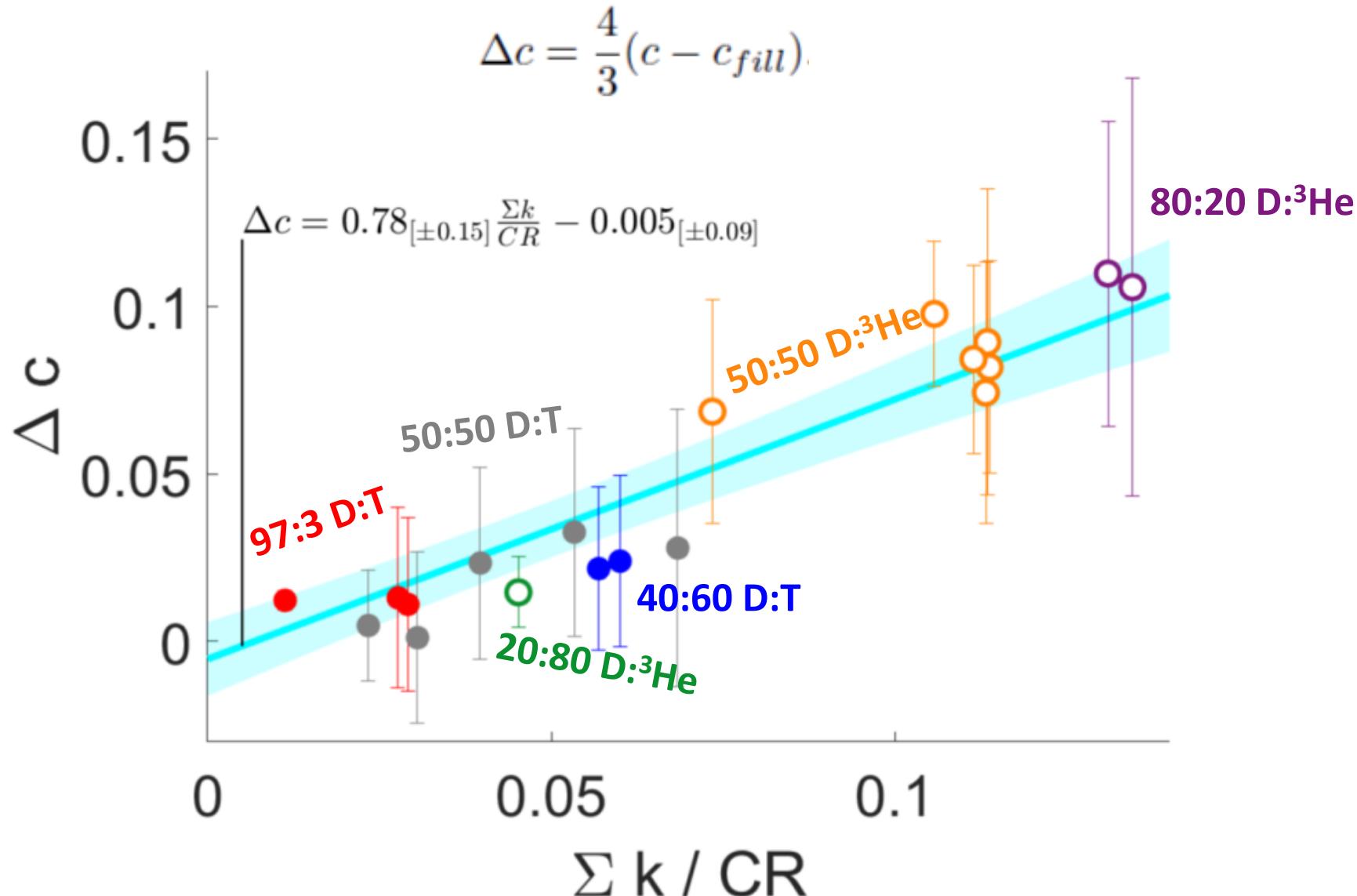
Separation driven by gradients in T, and P is compared to homogenization driven by classical diffusion to estimate the concentration gradient

$$\| \mathbf{I} \| = -\frac{D}{c} \frac{R_0}{v_{shock}} \nabla c \quad L_s = -\frac{D}{c} \frac{(k_p + k_{Ti} + k_{Te} + k_E)}{v_{shock}}$$

$$\nabla c = \frac{k_p + k_{Ti} + k_{Te} + k_E}{R_0}$$

$$\Delta c = \nabla c * R_{min} = (k_p + k_{Ti} + k_{Te} + k_E) \frac{R_{min}}{R_0}$$

$$\Delta c = \frac{k_p + k_{Ti} + k_{Te} + k_E}{CR} = \frac{\Sigma k}{CR}$$

The inferred level of species separation is correlated with the estimated value, but consistently lower



Path forward

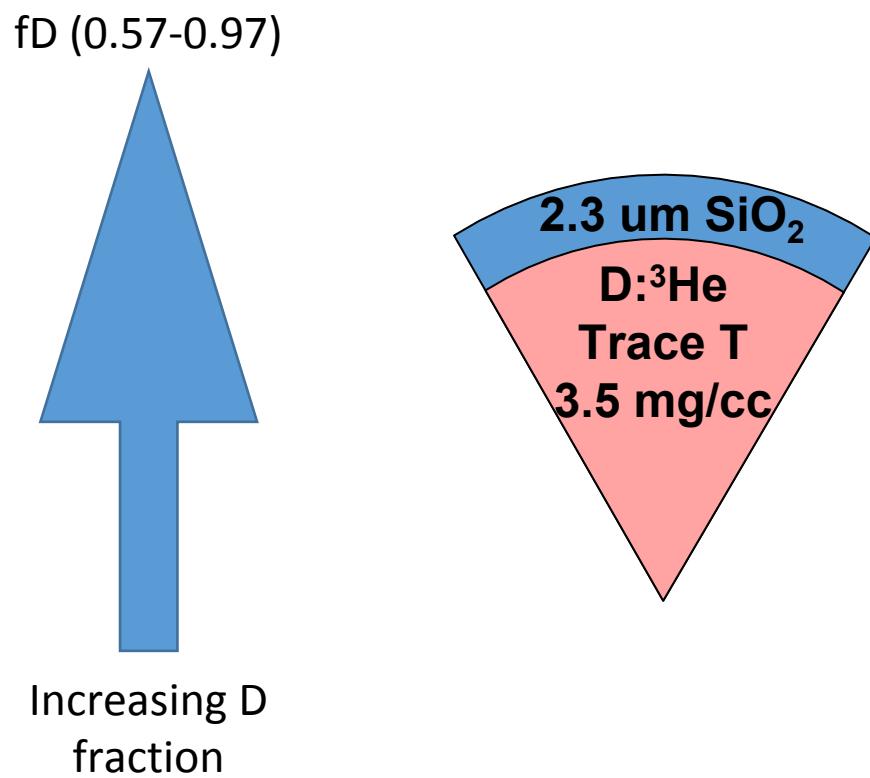
- Multi-ion hydro simulations with PHORCE (VT)
- Experiments investigating the impact of disequilibrium in the shock convergence phase on the subsequent compression phase
 - Experiments have been completed, data analysis is ongoing

Other things

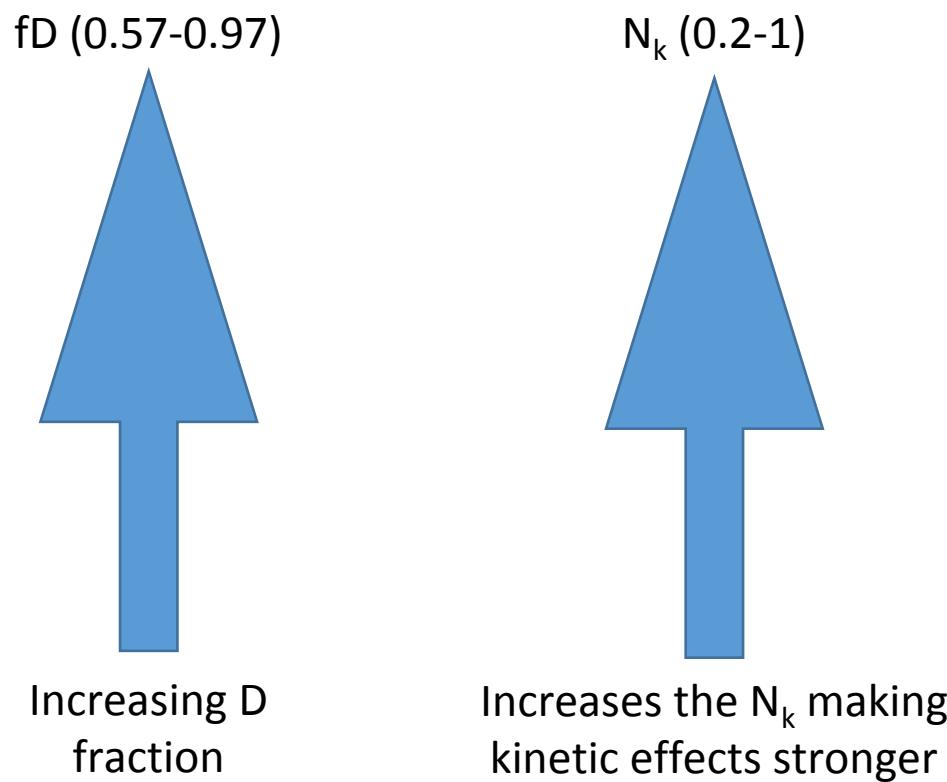
Previous measurements with the Particle X-ray Temporal Diagnostic were used to implicate time-resolved multi-ion effects during shock emission

The PXTD was used to measure the DTn, D³Hep, and DDn bang times simultaneously on experiments with “hydro equivalent” fill ratios.

The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions

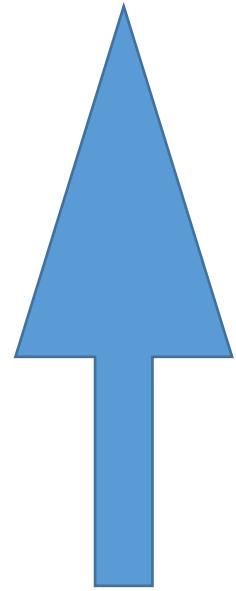


The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions



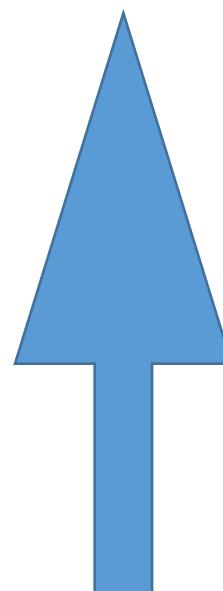
The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions

f_D (0.57-0.97)



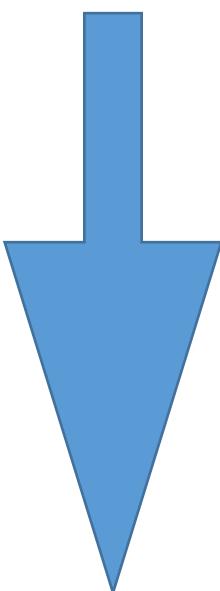
Increasing D fraction

N_k (0.2-1)



Increases the N_k making kinetic effects stronger

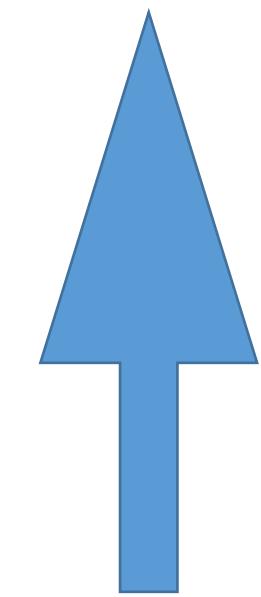
$\Sigma K / CR$ (3-4)



Decreases $\Sigma K / CR$ and species separation

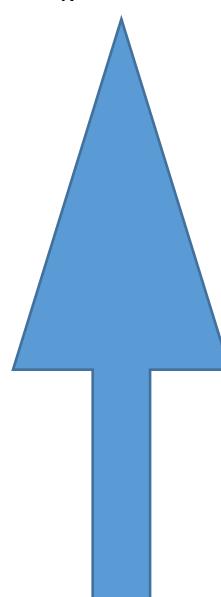
The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions

f_D (0.57-0.97)



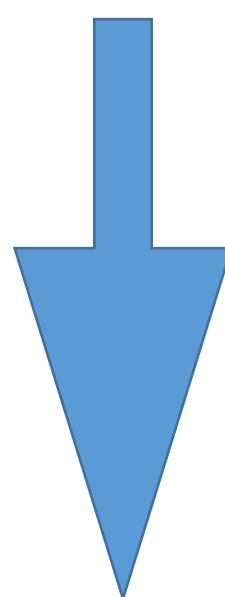
Increasing D fraction

N_k (0.2-1)



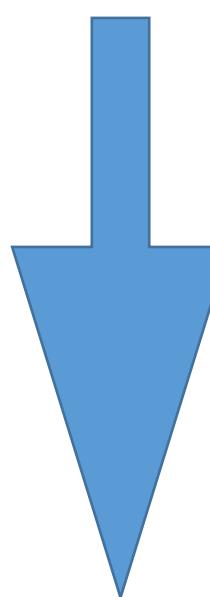
Increases the N_k making kinetic effects stronger

$\sum K / CR$ (3-4)



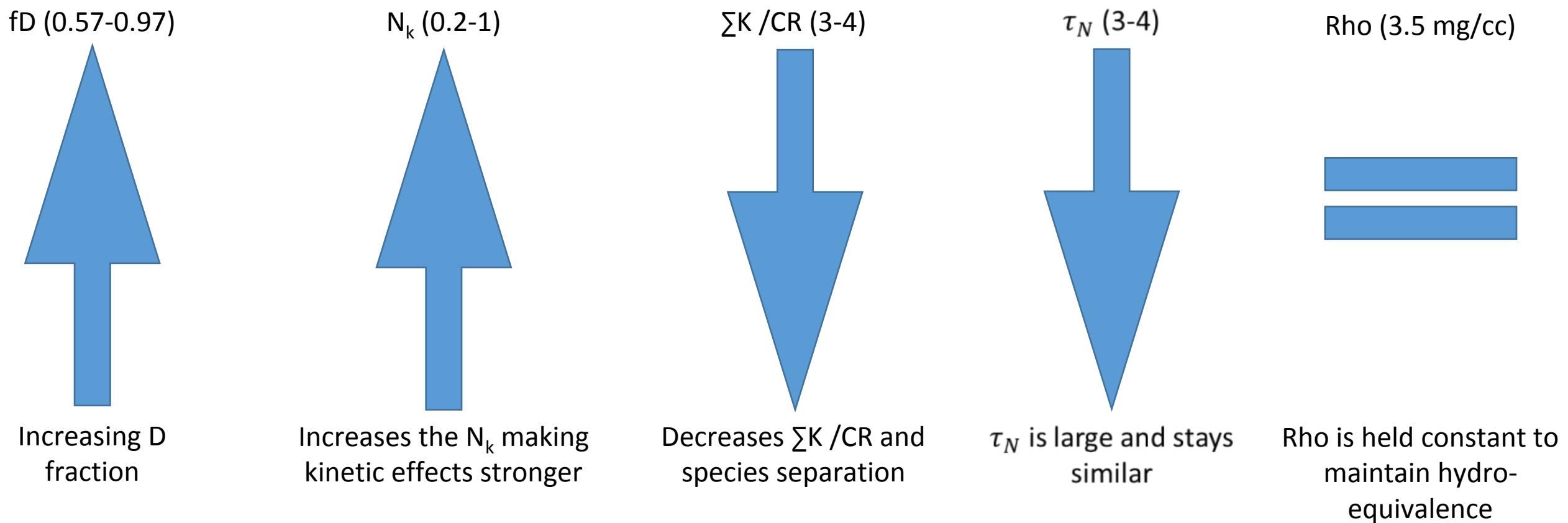
Decreases $\sum K / CR$ and species separation

τ_N (3-4)

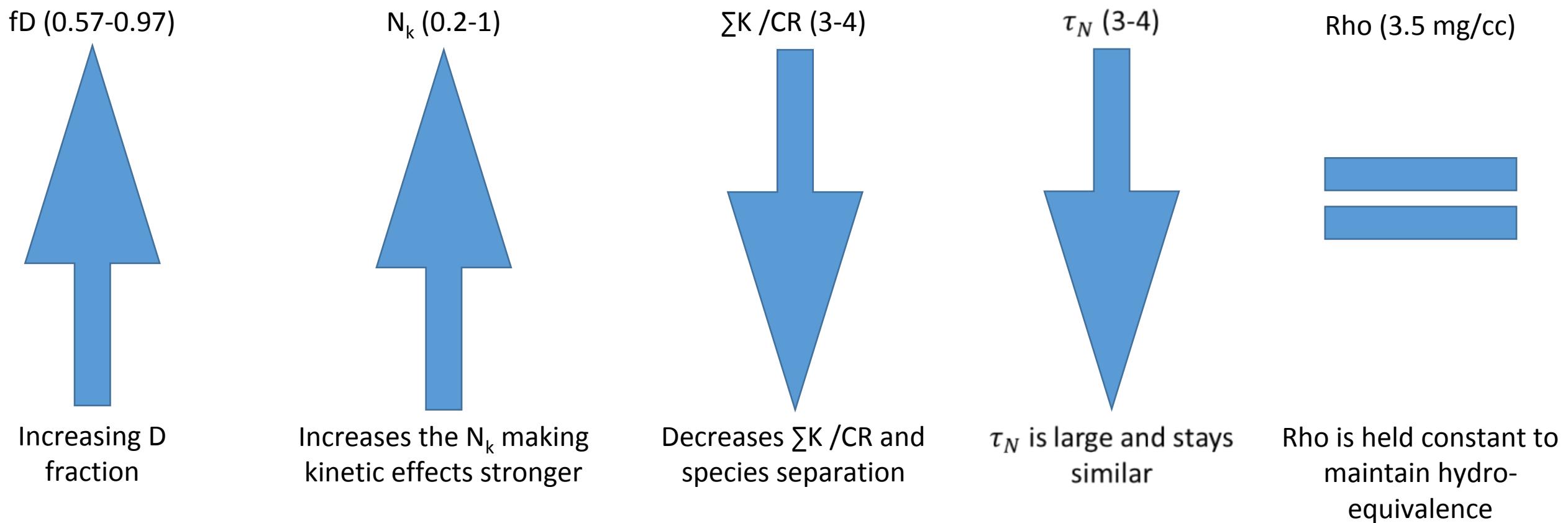


τ_N is large and stays similar

The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions

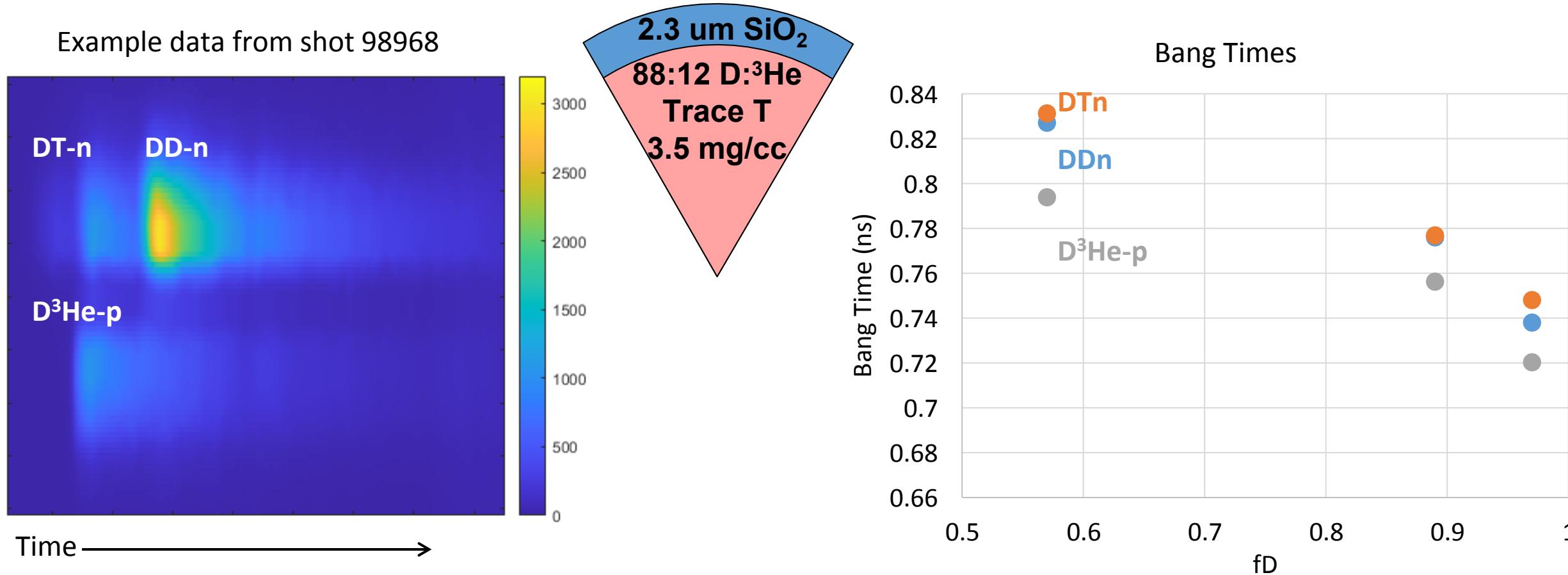


The cause of the emission time difference is determined by measuring the relative emission timing with varying D fraction in hydro equivalent implosions

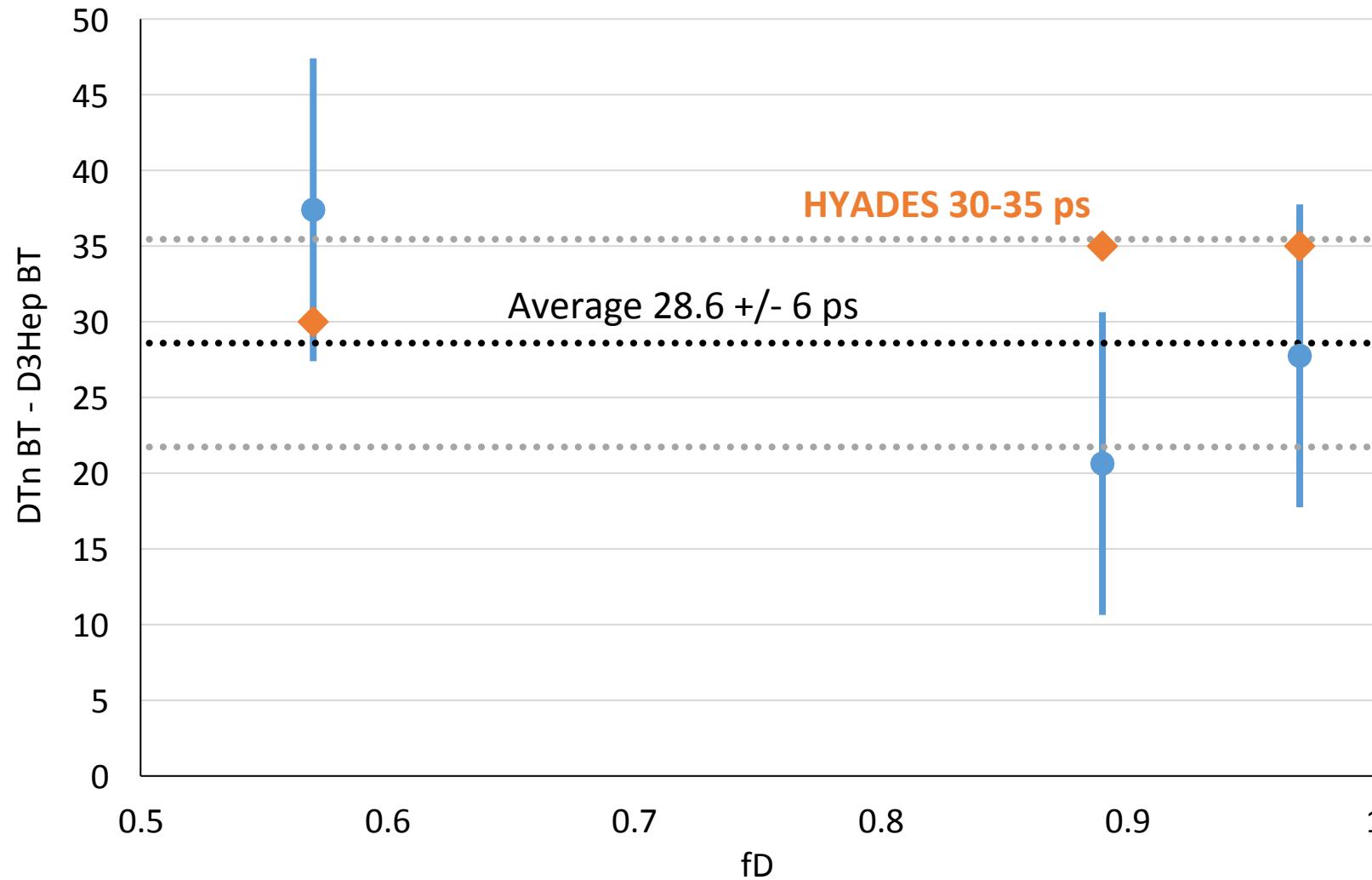


If the delta bang time increases with f_D it is likely kinetic effects, if the bang time decreases with f_D species separation is likely dominant, if it does not change then it could be a hydro effect.

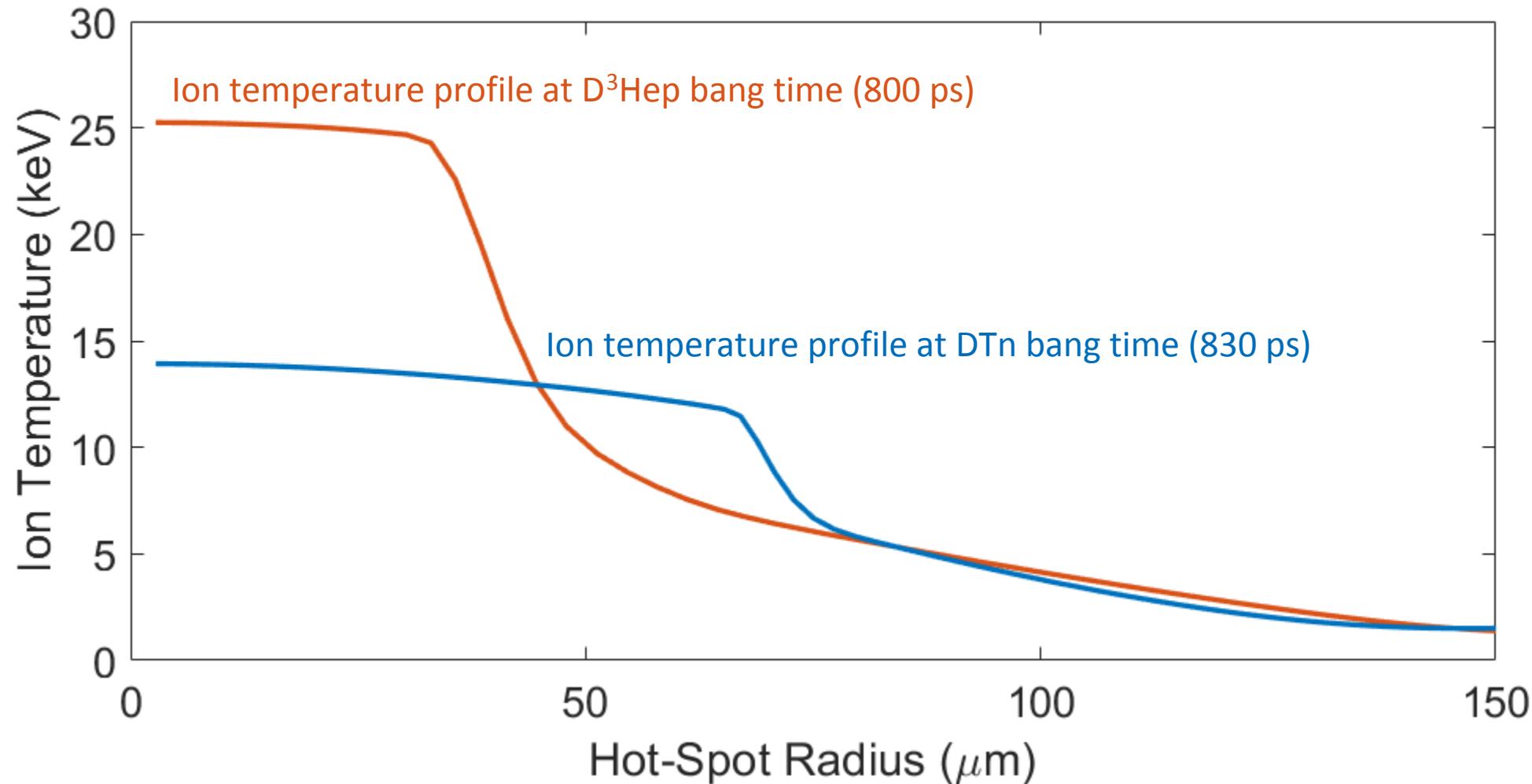
On these experiments the DDn, DTn, and D³Hep emission histories were simultaneously measured on PXTD



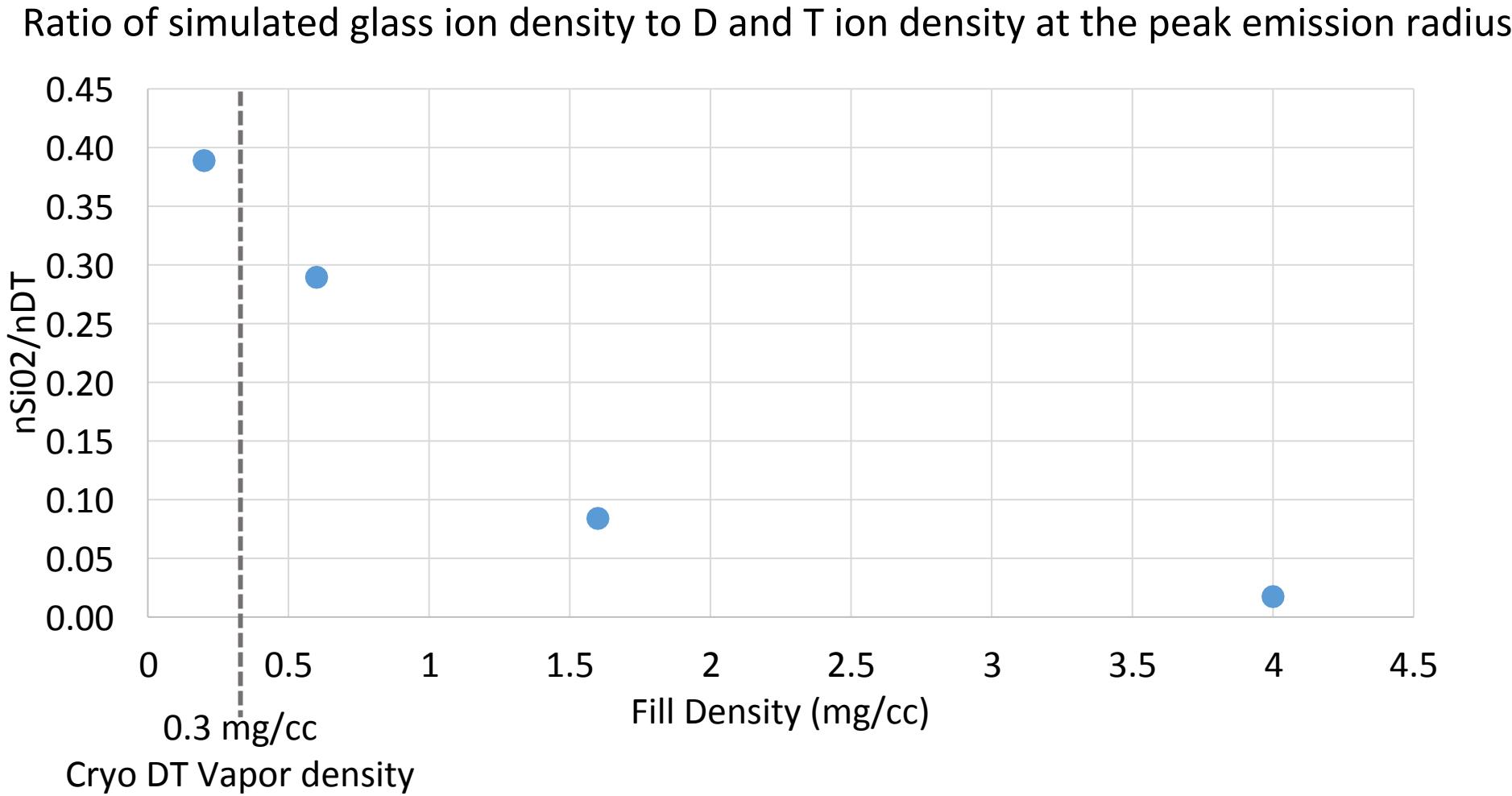
The measured bang time differentials are consistent with a constant value, indicating that hydrodynamics is most likely responsible



The D^3He emission is weighted closer to shock rebound when the central temperatures are higher, but the heated volume is smaller.

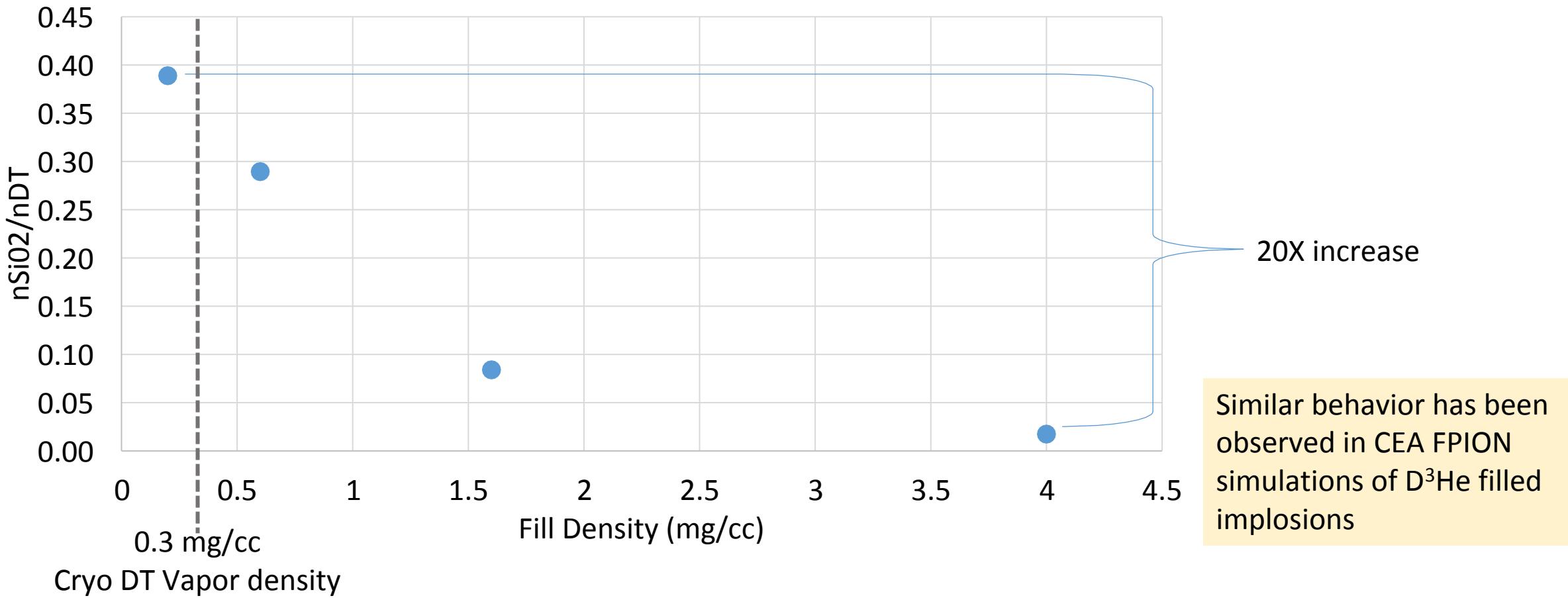


These iFP simulations show dramatically increasing SiO₂ mix with decreasing DT fill density

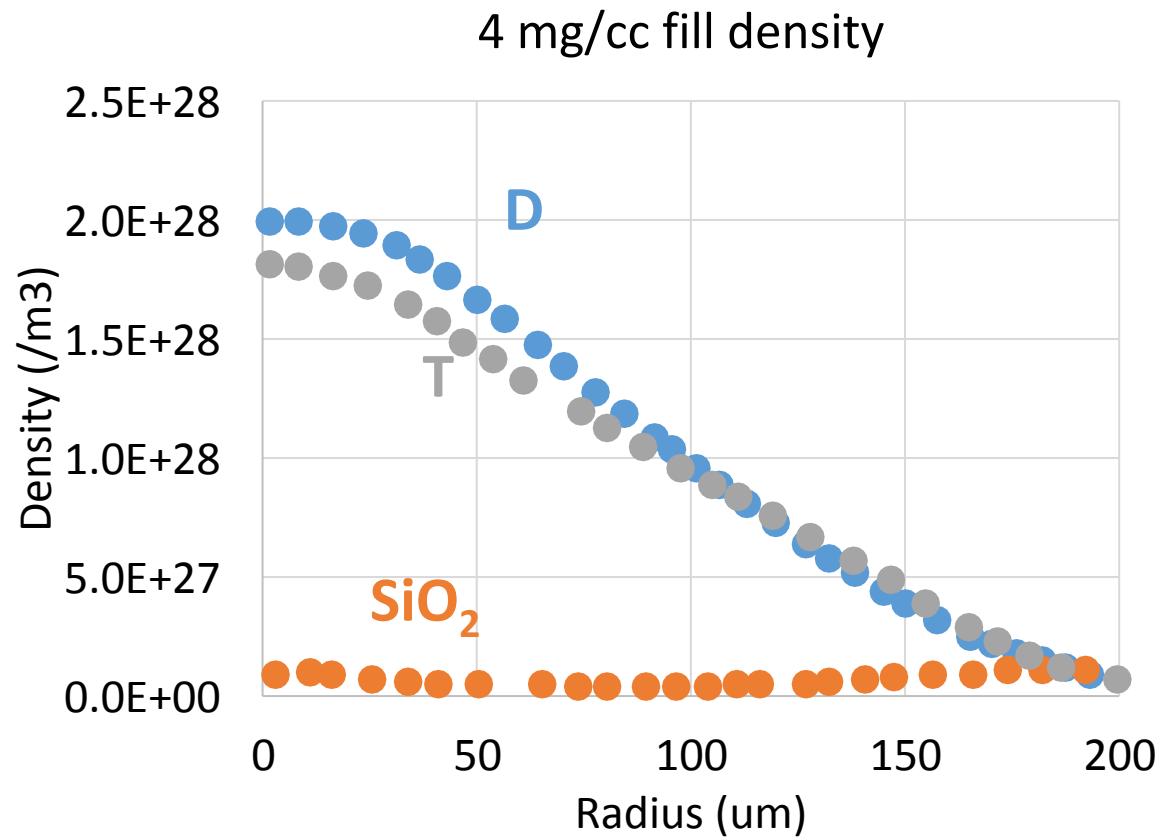


These iFP simulations show dramatically increasing SiO_2 mix with decreasing DT fill density

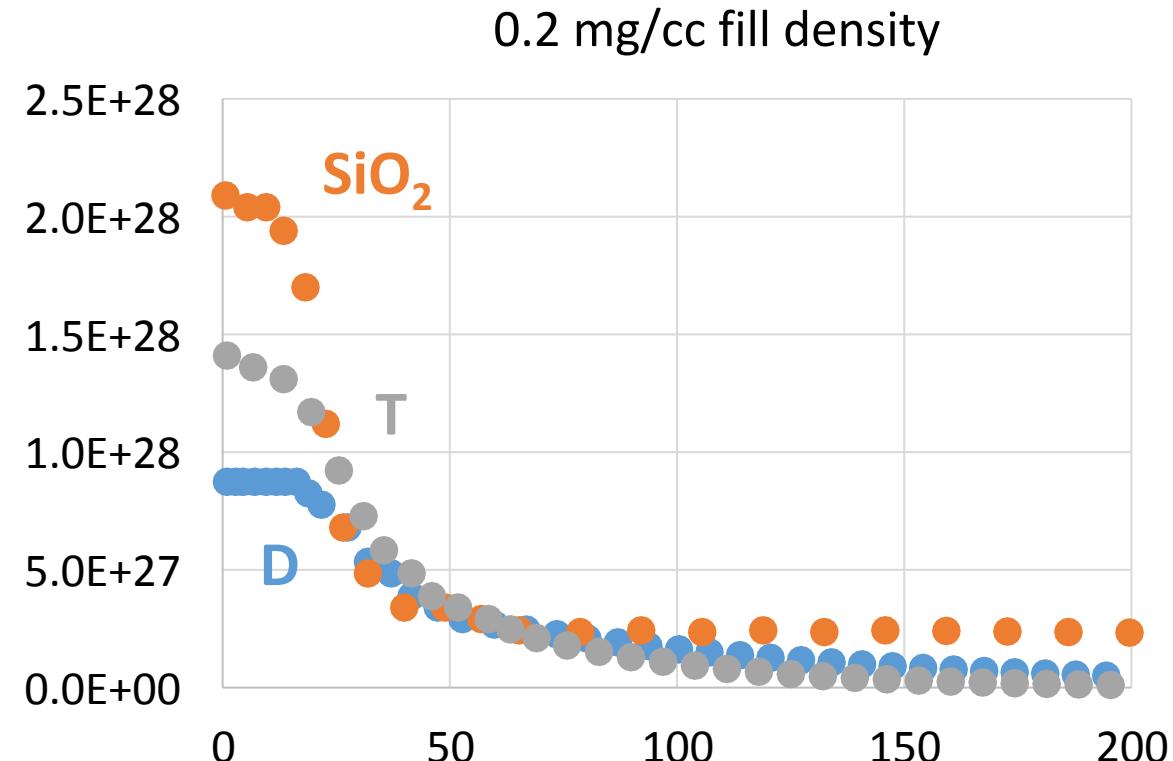
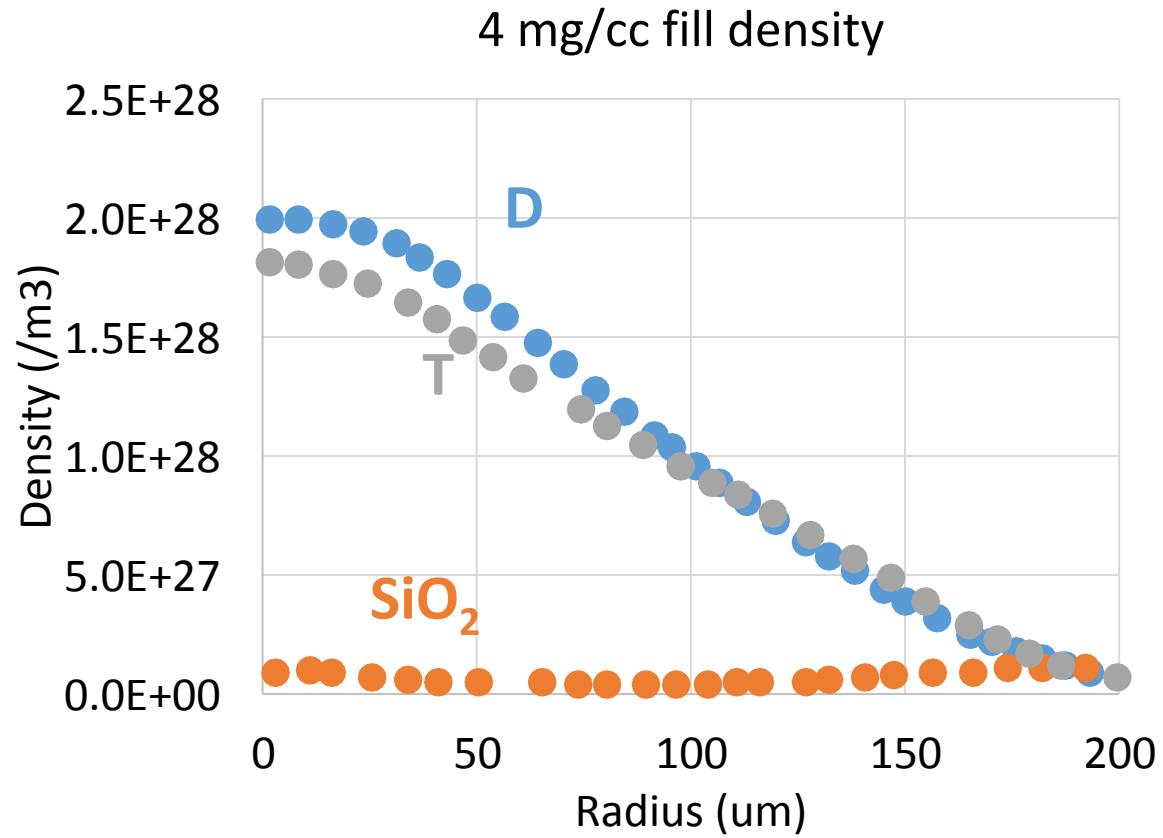
Ratio of simulated glass ion density to D and T ion density at the peak emission radius



The iFP simulated ion density profiles at bang time show that the SiO_2 ions dominate the material in the core at low density



The iFP simulated ion density profiles at bang time show that the SiO_2 ions dominate the material in the core at low density



Hans Rinderknecht saw almost identical D³He yield with and without D in the initial 0.5 mg/cc gas fill of CD shell shock driven implosions

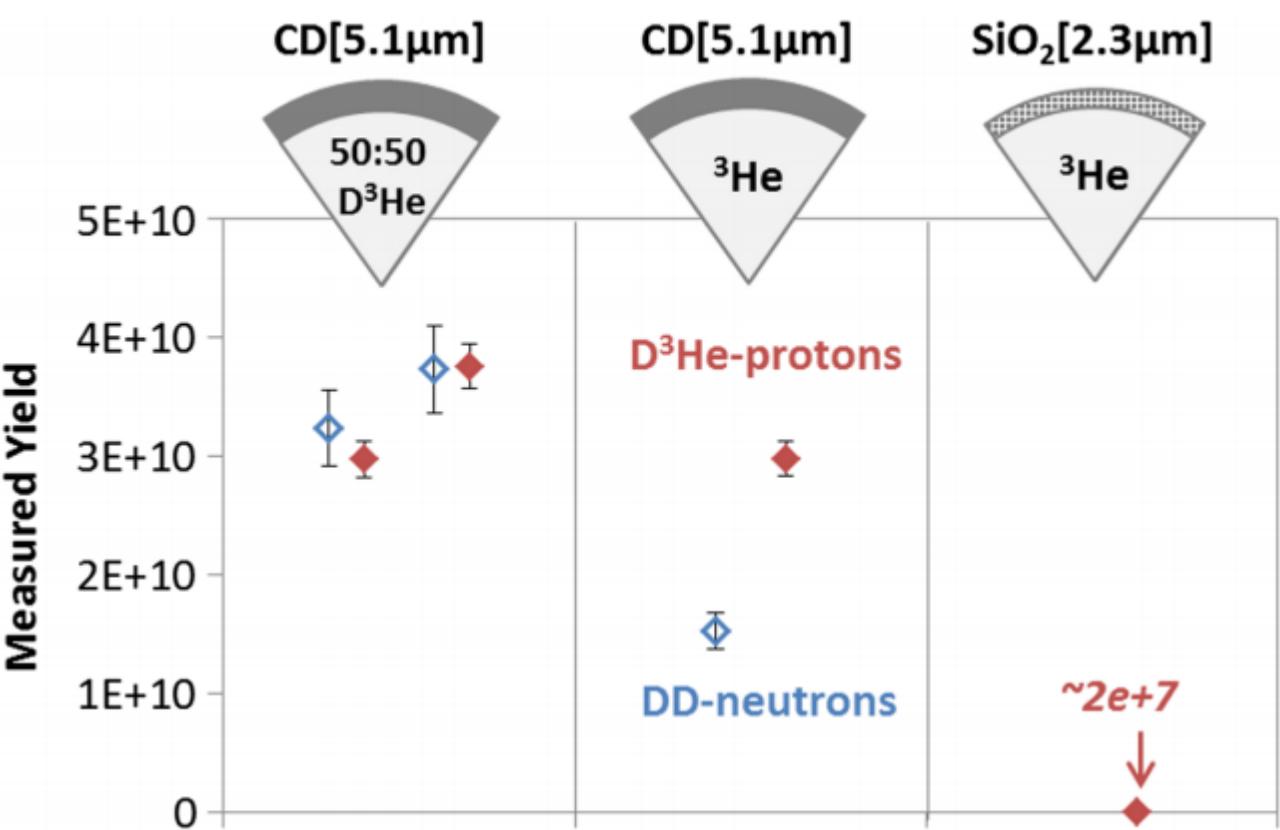
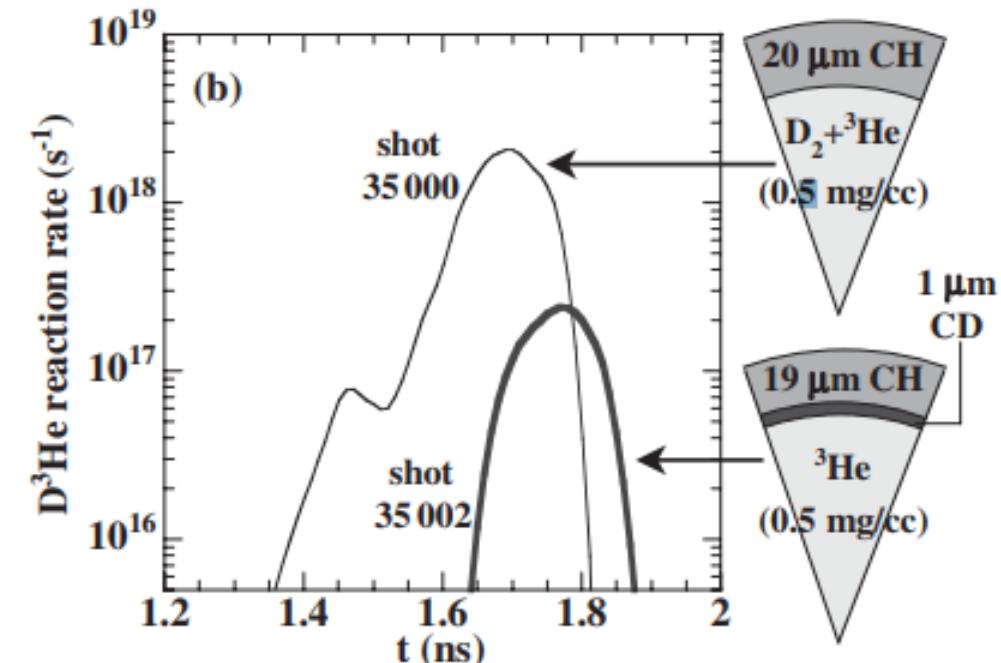
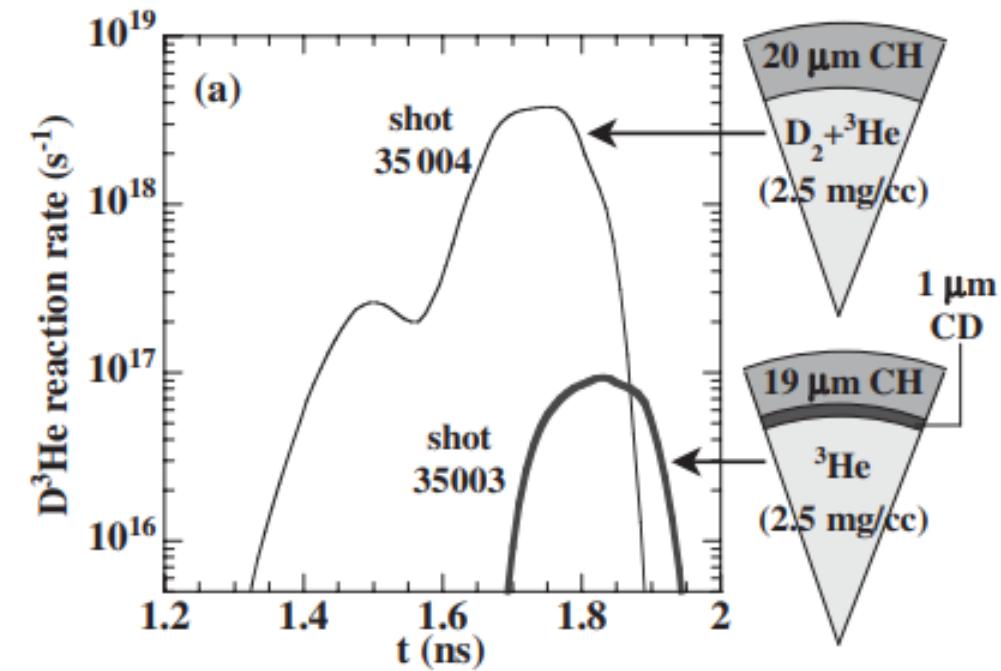


FIG. 1 (color online). Yields of D³He protons (solid red) recorded from 30 kJ implosions of deuterated plastic shells filled with a 50:50 D³He mixture are comparable in magnitude to implosions of the same shells filled with pure ³He. Implosions of glass shells filled with ³He produce yields 3 orders of magnitude lower, ruling out D₂-gas contamination of the targets as an explanation for this result. The observed D³He-*p* yields require shell-deuterium mix into the ³He fuel to be of the order of 10% the initial ³He gas density.

Ryan Rygg saw that in compressive implosions the compression yield dominated in the separated reactant implosions.



Chikang Li saw that recessing a CD layer in ablatively driven implosions reduces the yield by an order of magnitude.

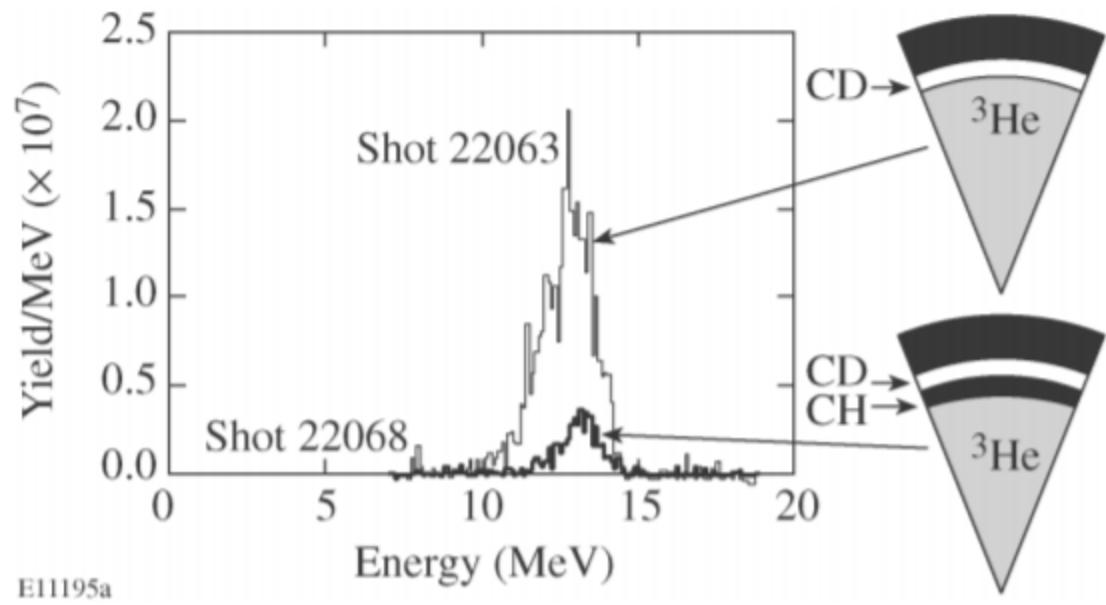


FIG. 1. The structures of two different capsules filled with 4 atm of pure ^3He gas, and measured spectra of primary D^3He protons from implosions at OMEGA. The ratio of number densities (D to C) is 1.56 in the CD layer.

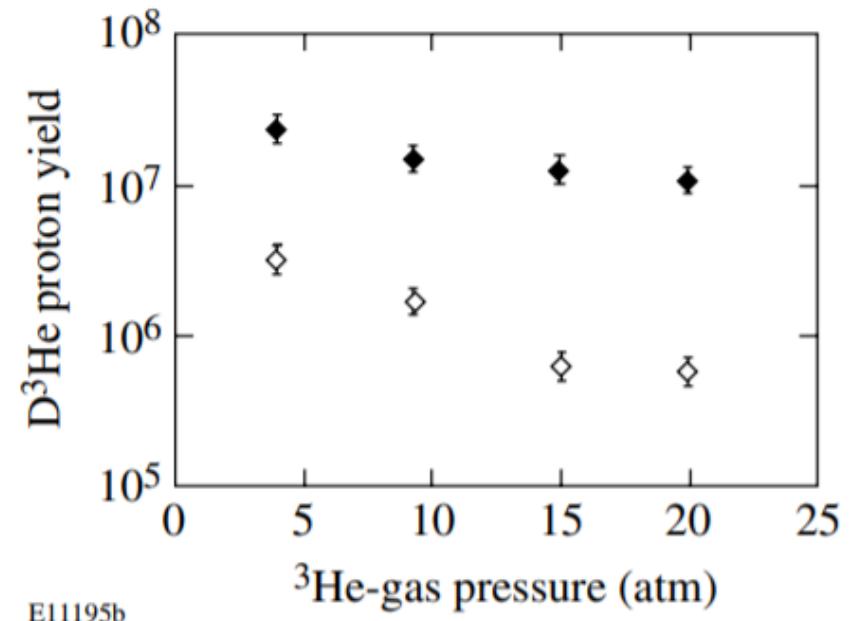


FIG. 2. D^3He proton yields plotted as a function of gas-fill pressures for implosions of capsules with ^3He fuel and zero- (closed dark diamond) and 1 μm -offset (open diamond) CD layer shells.

Similar results are seen in FPION simulations of D3He-filled implosions

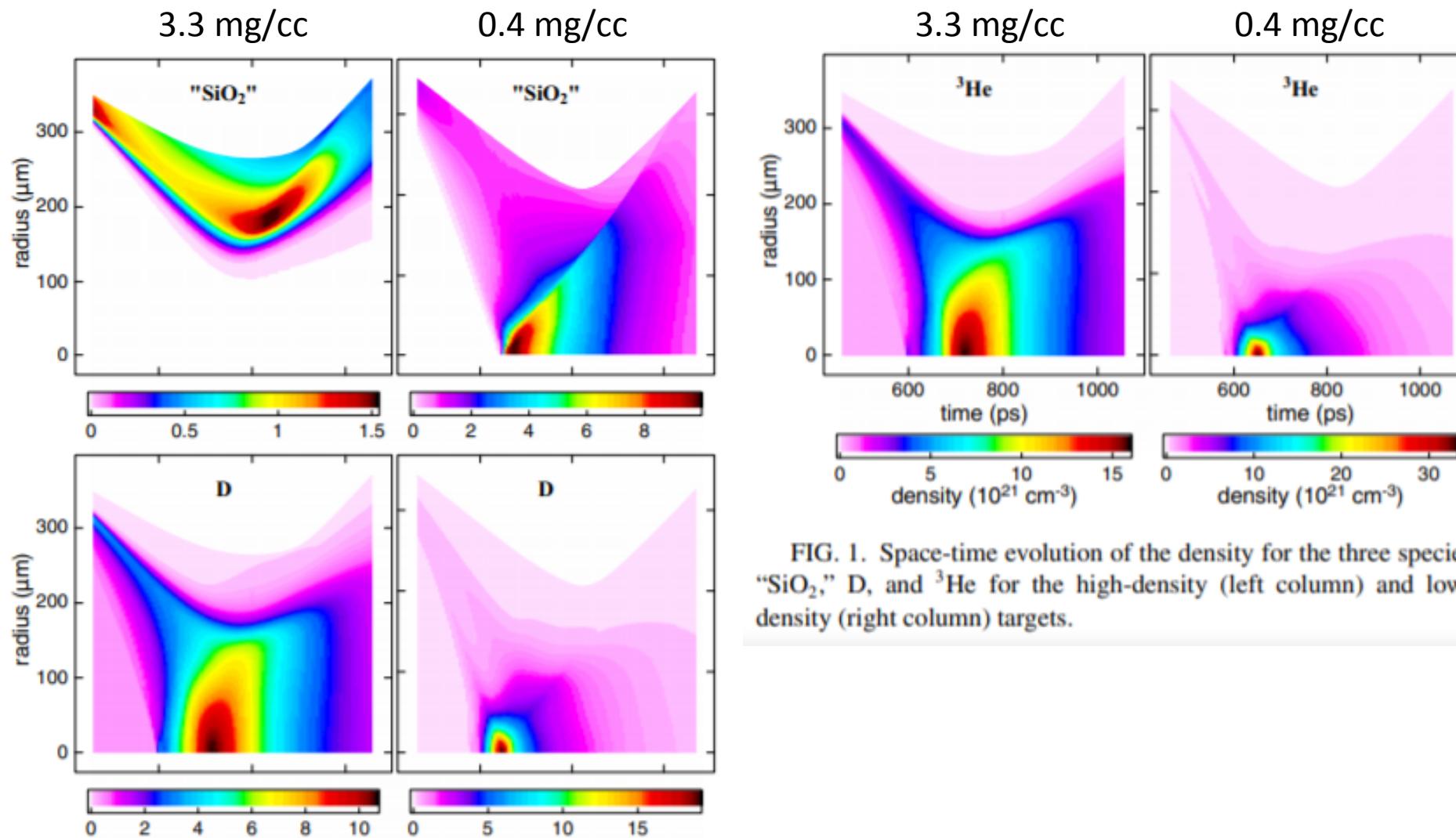


FIG. 1. Space-time evolution of the density for the three species $^{33}\text{SiO}_2$, D, and ^3He for the high-density (left column) and low-density (right column) targets.

Mike Rosenberg also saw decreasing YOC with decreasing fill density (and therefore increasing Knudsen number).

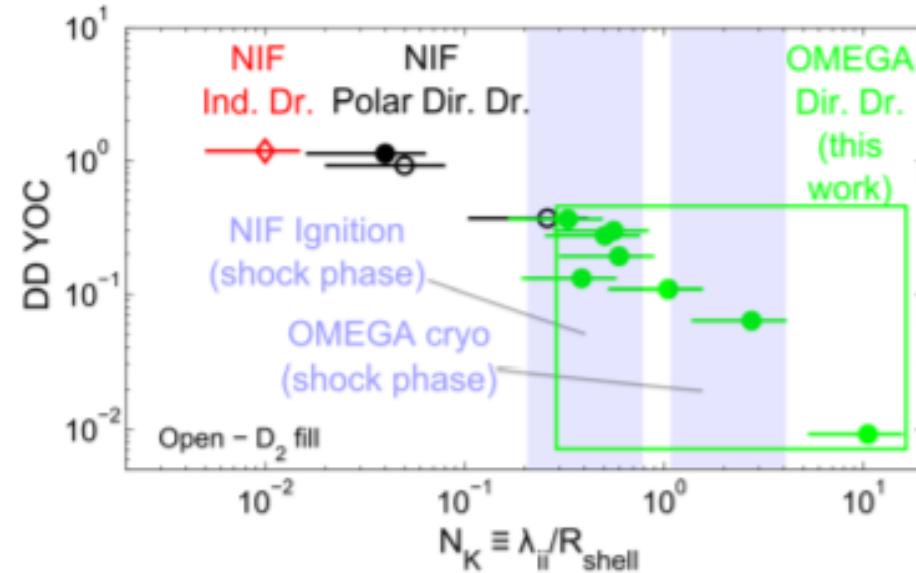
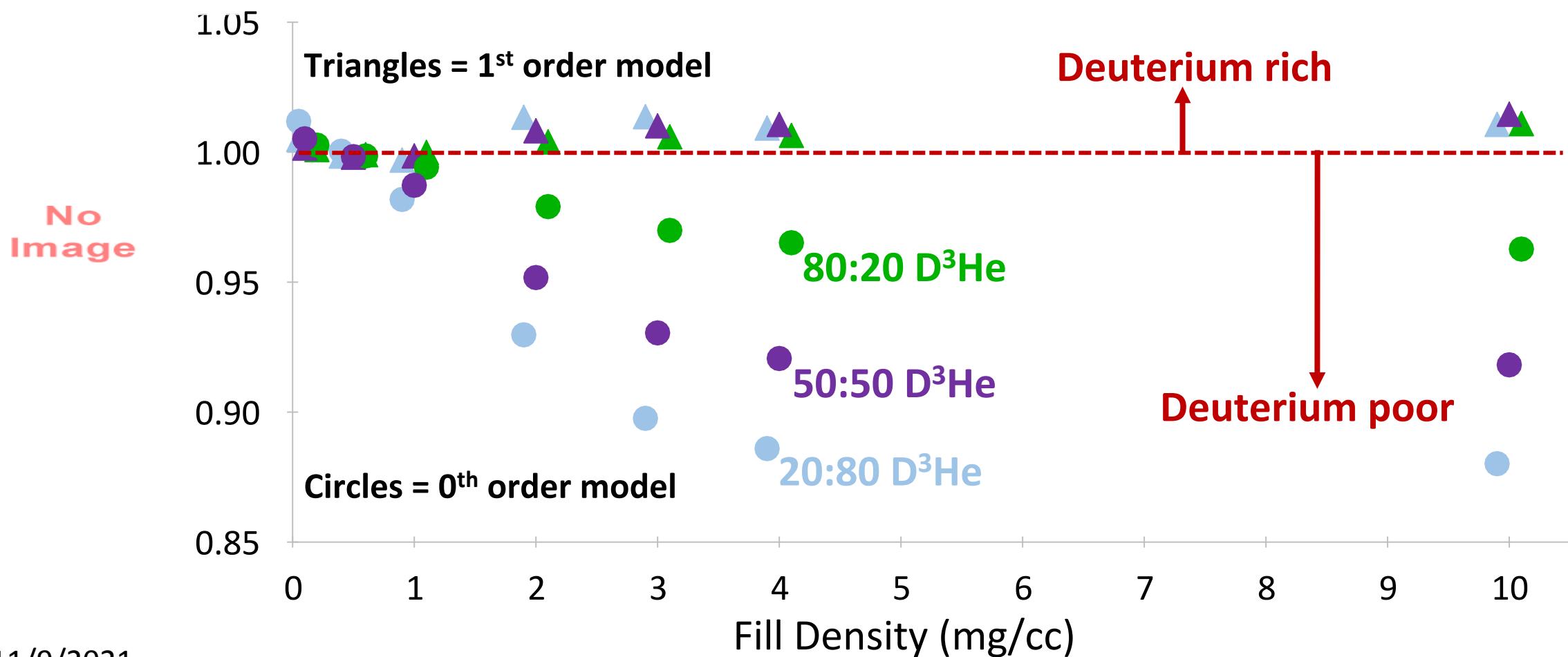


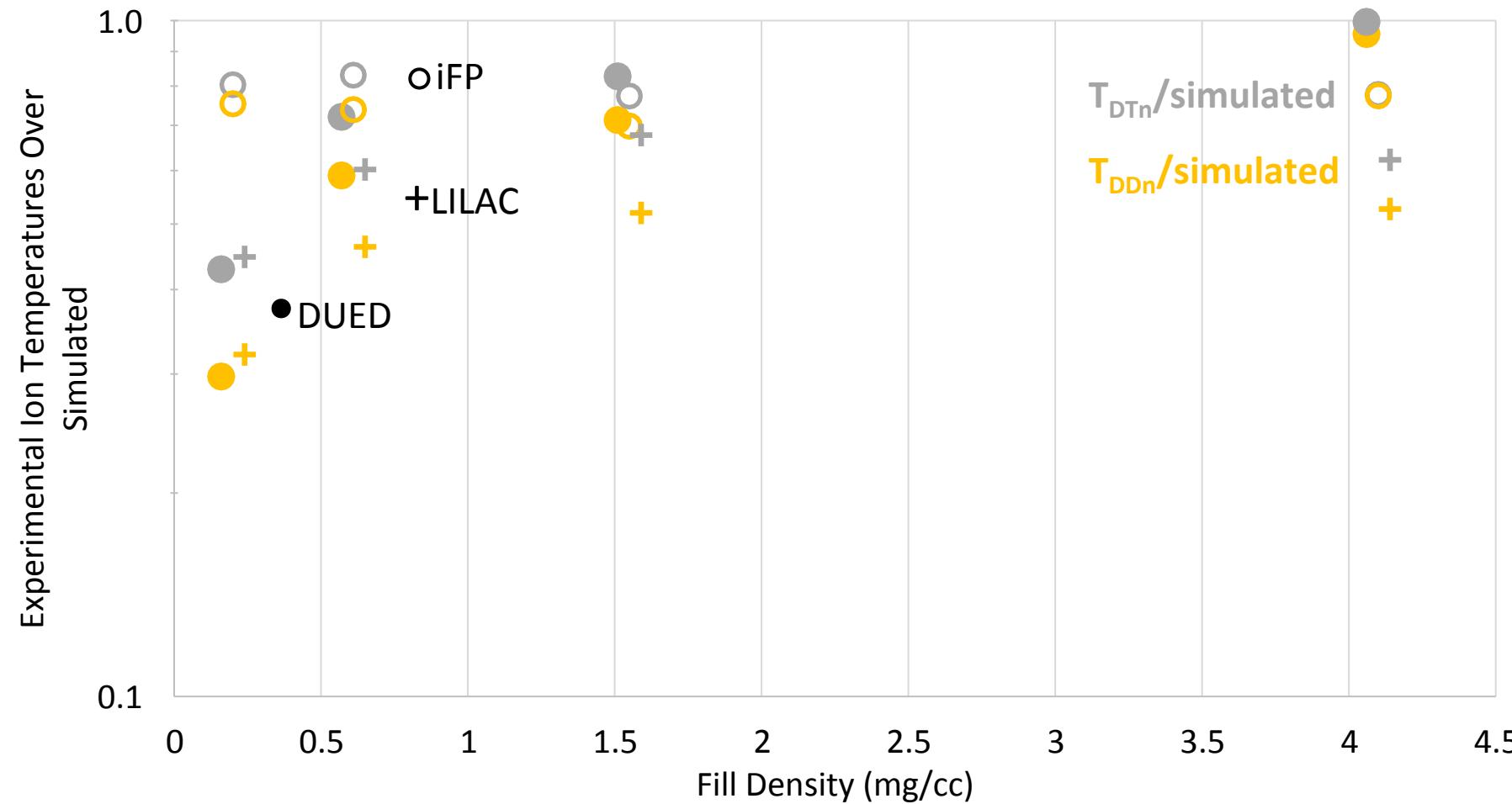
FIG. 1. Ratio of measured DD fusion yields to hydrodynamics-simulated yields (yield over clean, YOC) as a function of the Knudsen number (N_K) for an indirect-drive exploding pusher on NIF (red diamond),²³ three polar-direct-drive (PDD)²⁴ exploding pushers on NIF,²² and direct-drive exploding pushers on OMEGA (green circles).⁶ Fusion burn profile measurements of the OMEGA experiments are described herein. Filled markers represent D^3He -filled implosions, while open markers denote D_2 -filled implosions. Though the drive conditions are quite different, these experiments show a unified picture of the increasing impact of ion kinetic effects as a function of increasing Knudsen number for $N_K \geq 0.1$. A band centered around $N_K = 0.5$ shows the approximate Knudsen number at the center of a NIF ignition-relevant indirect-drive implosion²⁵ or a NIF polar-direct-drive implosion²⁴ immediately after shock convergence, while a band centered around $N_K = 2$ shows the approximate Knudsen number after shock convergence at the center of a cryogenic layered implosion on OMEGA.²⁶ This data were originally presented in Ref. 22. Reproduced with permission from Phys. Plasmas **21**, 122712 (2014).

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iFP consistently overpredicts the DT_n and DD_n ion temperatures indicating that some details of the ion distributions may not be properly captured



iFP temperatures are calculated by simulating the neutron spectra and using them to infer apparent temperatures.

The temperature ratio from iFP simulations is not consistent with the experimentally observed trend

