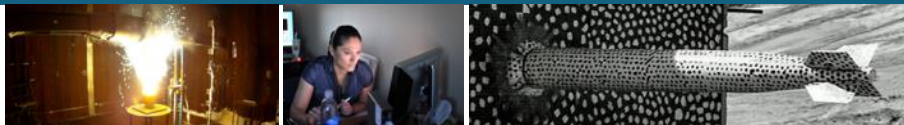




A Five-Moment Multifluid Model for Partially Ionized Plasmas With Arbitrarily Many Species



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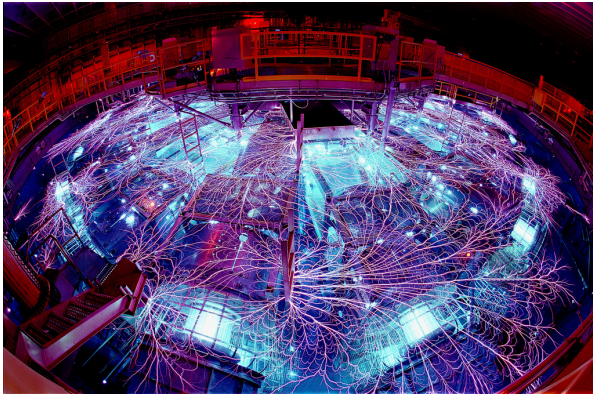


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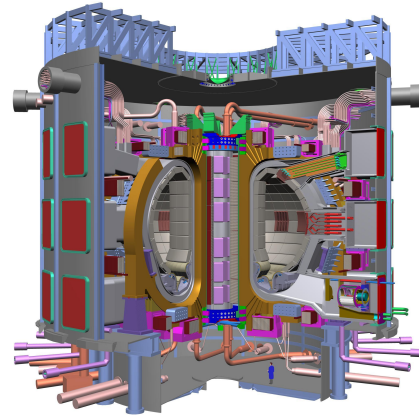
Motivation



- High-energy density plasmas.
- Higher Z elements: He, Ne, Ar.
- Wide range of density/temperature/etc.



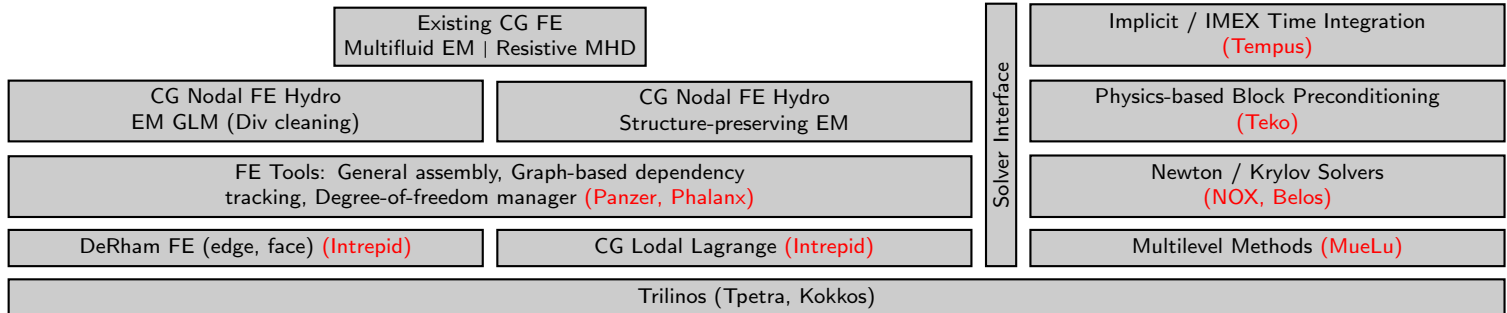
Z-Pinch (Z Machine)
eg., Ar Gas Puff



Magnetic Confinement Fusion (ITER)
eg., MGI for disruption mitigation

Drekar: Resistive MHD / Multifluid with Coupled Multiphysics

- Arbitrarily many equations describing physics (continuity, momentum, energy, electromagnetics).
- ERK, DIRK, IMEX time integration (Tempus).
- 2D & 3D unstructured finite element (Intrepid):
 - Stabilized Q1, Q2 elements (high-order in process).
 - Physics compatible discretizations (node, edge, face).
 - High-resolution positivity-preserving methods.
- Advanced software capabilities:
 - MPI+X (Kokkos).
 - Linear/non-linear solvers (NOX, Belos) with robust, scalable preconditioning (Teko, MueLu).
 - Exact Jacobians through automatic differentiation (Sacado).
 - Asynchronous dependency manages multiphysics complexity (Phalanx).



Overview of Models and Equations



Kinetic equations:

$$\partial_t f_s + \mathbf{v} \cdot \nabla_{\mathbf{x}} f_s + \frac{q_s}{m_s} (\mathbf{E} + \mathbf{v} \times \mathbf{B}) \cdot \nabla_{\mathbf{v}} f_s = \mathcal{C}_s [f_s] + \mathcal{S}_s, \quad (s \in \Lambda),$$

with

$$\mathcal{C}_s = \mathcal{C}_s^{\text{sc}} + \mathcal{C}_s^{\text{ion}} + \mathcal{C}_s^{\text{rec}} + \mathcal{C}_s^{\text{cx}} + \mathcal{C}_s^{\text{rad}},$$

Fluid equations:

$$\partial_t \rho_s + \nabla \cdot (\rho_s \mathbf{u}_s) = \mathcal{C}_s^{[0]} + \mathcal{S}_s^{[0]},$$

$$\partial_t (\rho_s \mathbf{u}_s) + \nabla \cdot (\rho_s \mathbf{u}_s \otimes \mathbf{u}_s + p_s \underline{\mathbf{I}} + \underline{\mathbf{\Pi}}_s) = q_s n_s (\mathbf{E} + \mathbf{u}_s \times \mathbf{B}) + \mathcal{C}_s^{[1]} + \mathcal{S}_s^{[1]},$$

$$\partial_t \mathcal{E}_s + \nabla \cdot [(\mathcal{E}_s + p_s) \mathbf{u}_s + \mathbf{u}_s \cdot \underline{\mathbf{\Pi}}_s + \mathbf{h}_s] = q_s n_s \mathbf{u}_s \cdot \mathbf{E} + \mathcal{C}_s^{[2]} + \mathcal{S}_s^{[2]}.$$

$\mathcal{C}_s^{\text{sc}}$: elastic scattering

$\mathcal{C}_s^{\text{ion}}$: ionization reactions

$\mathcal{C}_s^{\text{rec}}$: recombination reactions

$\mathcal{C}_s^{\text{cx}}$: charge exchange

$\mathcal{C}_s^{\text{rad}}$: radiative loss

Here:

$$s \in \Lambda = \{(\alpha, k) : \alpha = 1, \dots, N_A; k = 0, \dots, z_\alpha\} \cup \{e\}$$

Multifluid EM Plasma Model



Density	$\partial_t \rho_s + \nabla \cdot (\rho_s \mathbf{u}_s) = -\rho_s n_e (I_s + R_s) + m_s n_e (n_{s-1} I_{s-1} + n_{s+1} R_{s+1})$			
Momentum	$\partial_t (\rho_s \mathbf{u}_s) + \nabla \cdot (\rho_s \mathbf{u}_s \otimes \mathbf{u}_s + p_s \mathbf{I} + \underline{\Pi}_s) = q_s n_s (\mathbf{E} + \mathbf{u}_s \times \mathbf{B}) + \sum_{t \neq s} \alpha_{s;t} \rho_s \rho_t (\mathbf{u}_t - \mathbf{u}_s)$ $-\rho_s \mathbf{u}_s n_e (I_s + R_s) + \frac{m_s}{m_{s-1}} n_e \rho_{s-1} \mathbf{u}_{s-1} I_{s-1} + (n_e \rho_{s+1} \mathbf{u}_{s+1} + n_{s+1} \rho_e \mathbf{u}_e) R_{s+1}$			
Energy	$\partial_t \mathcal{E}_s + \nabla \cdot [(\mathcal{E}_s + p_s) \mathbf{u}_s + \mathbf{u}_s \cdot \underline{\Pi}_s + \mathbf{h}_s] = q_s n_s \mathbf{u}_s \cdot \mathbf{E} + \sum_{t \neq s} \frac{\alpha_{s;t} \rho_s \rho_t}{m_s + m_t} [A_{s;t} k_B (T_t - T_s) + m_t (\mathbf{u}_t - \mathbf{u}_s)^2]$ $-\mathcal{E}_s n_e (I_s + R_s) + \frac{m_s}{m_{s-1}} n_e \mathcal{E}_{s-1} I_{s-1} + (n_e \mathcal{E}_{s+1} + n_{s+1} \mathcal{E}_e) R_{s+1}$			
Charge & Current	$q = \sum_s q_s n_s$		$\mathbf{J} = \sum_s q_s n_s \mathbf{u}_s$	
Maxwell's Equations	$\partial_t \mathbf{E} - c^2 \nabla \times \mathbf{B} + \frac{1}{\epsilon_0} \mathbf{J} = \mathbf{0}$	$\partial_t \mathbf{B} + \nabla \times \mathbf{E} = \mathbf{0}$	$\nabla \cdot \mathbf{E} = \frac{q}{\epsilon_0}$	$\nabla \cdot \mathbf{B} = 0$

1. Flexible and extensible multiphysics PDE solver framework (**Drekar**) can accommodate arbitrary numbers of species.
2. Nodal FE or structure-preserving FE discretizations (edge, face, etc.).
3. Stable and accurate AFC local bounds preserving CG FE methods.
4. IMEX capabilities for handling disparate time scales (eg., large plasma/cyclotron freq., EM CFL, etc.).
5. Robust and scalable linear and nonlinear solvers for CG/structure preserving methods
 - ⇒ Newton-Krylov with physics-based block decomposition preconditioning; enabling optimal H(curl) and H(grad) multilevel subsystem solvers.

$$\mathbf{R}_{s;t} = \alpha_{s;t} \rho_s \rho_t (\mathbf{u}_t - \mathbf{u}_s) \Phi_{s;t}$$
$$Q_{s;t} = \frac{\alpha_{s;t} \rho_s \rho_t}{m_s + m_t} \left[A_{s;t} k_B (T_t - T_s) \Psi_{s;t} + m_t (\mathbf{u}_t - \mathbf{u}_s)^2 \Phi_{s;t} \right],$$

- Charge-charge (Coulomb):

$$\alpha_{s;t} = \frac{Z_s^2 Z_t^2 |q_e|^4 \ln \Lambda_{s;t}}{6\pi \sqrt{2\pi} \epsilon_0^2 m_s m_t m_{s;t} (k_B T_s / m_s + k_B T_t / m_t)^{3/2}},$$

- Charge-neutral/Neutral-neutral:

$$\alpha_{s;t} = \frac{1}{m_s + m_t} \frac{4}{3} \left[\frac{8}{\pi} \left(\frac{k_B T_s}{m_s} + \frac{k_B T_t}{m_t} \right) \right]^{1/2} \sigma_{s;t}.$$

⇒ Using constant cross-sections for now (most computed using hard-sphere approximation, some from QM calculations).

Ionization Rates $\Gamma_{(\alpha,k)}^{\text{ion}} = n_e n_{(\alpha,k)} I_{(\alpha,k)}$



Voronov:^a

$$I_{(\alpha,k)} = A_{(\alpha,k)} \frac{1 + P_{(\alpha,k)} \sqrt{U_{(\alpha,k)}}}{X_{(\alpha,k)} + U_{(\alpha,k)}} (U_{(\alpha,k)})^{K_{(\alpha,k)}} \exp(-U_{(\alpha,k)}), \quad U_{(\alpha,k)} = \frac{\phi_{(\alpha,k)}^{\text{ion}}}{T_e}.$$

- Available for H to Ni²⁷⁺.
- Accurate to within 10% for T_e between 1 eV and 20 KeV.

Lotz:^{b,c}

$$I_{(\alpha,k)} = (2.97\text{E-}6) \frac{\xi_{(\alpha,k)}}{\phi_{(\alpha,k)}^{\text{ion}} \sqrt{T_e}} E_1(U_{(\alpha,k)}),$$

- $\xi_{(\alpha,k)}$ is the number of outer electrons in the ionizing atom.
- General analytic model for any atomic species.
- Compatible with ionization potential depression models.

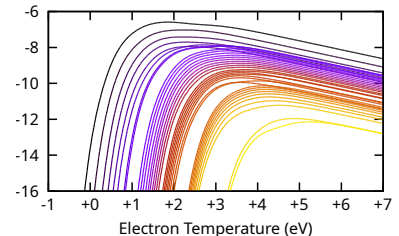
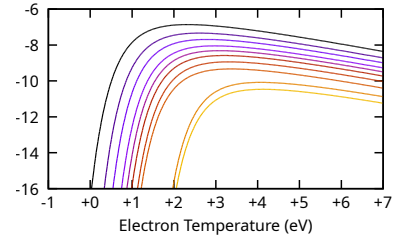
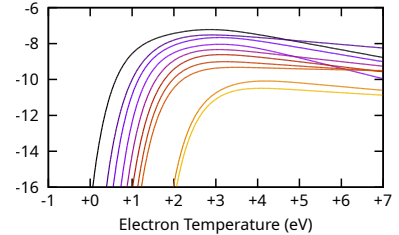
Others: Other sources for specific higher Z elements; eg., Mattioli, et al.^d for Kr.

^aG. VORONOV, *A practical fit formula for ionization rate coefficients of atoms and ions by electron impact: z=1-28*, Atomic Data and Nuclear Data Tables, 65 (1997), pp. 1-35.

^bW. LOTZ, *Electron-impact ionization cross-sections and ionization rate coefficients for atoms and ions from hydrogen to calcium*, Zeitschrift für Physik, 216 (1968), pp. 241-247.

^cW. LOTZ, *Electron-impact ionization cross-sections and ionization rate coefficients for atoms and ions from scandium to zinc*, Zeitschrift für Physik, 220 (1969), pp. 466-472.

^dM. MATTIOLI, G. MAZZITELLI, K. B. FOURNIER, M. FINKENTHAL, AND L. CARRARO, *Updating of atomic data needed for ionization balance evaluations of krypton and molybdenum*, Journal of Physics B: Atomic, Molecular and Optical Physics, 39 (2006), pp. 4457-4489.



Radiative Recombination Rates $\Gamma_{(\alpha,k)}^{\text{rec}} = n_e n_{(\alpha,k)} \left(R_{(\alpha,k)}^{\text{rad}} + R_{(\alpha,k)}^{\text{die}} \right)$



Badnell, et al.:^a

$$R_{(\alpha,k)}^{\text{rad}} = A_{(\alpha,k)} \left[\sqrt{T_e/T_0^{(\alpha,k)}} \left(1 + \sqrt{T_e/T_0^{(\alpha,k)}} \right)^{1-D_{(\alpha,k)}} \left(1 + \sqrt{T_e/T_1^{(\alpha,k)}} \right)^{1+D_{(\alpha,k)}} \right]^{-1},$$

$$D_{(\alpha,k)} = B_{(\alpha,k)} + C_{(\alpha,k)} \exp\left(-T_2^{(\alpha,k)}/T_e\right)$$

- Available for H through Zn, plus Kr, Mo, Xe.
- Fits of calculated data from AUTOSTRUCTURE code.

Kotelnikov, et al.:^b

$$R_{(\alpha,k)}^{\text{rad}} = \frac{8.414k\alpha^4 c a_0^2 [\ln(1+\lambda) + 3.499]}{(1/\lambda)^{1/2} + 0.6517(1/\lambda) + 0.2138(1/\lambda)^{3/2}},$$

$$\lambda = \frac{hR_{\infty}ck^2}{k_B T_e}$$

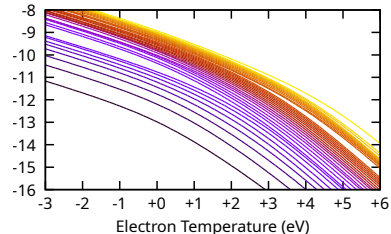
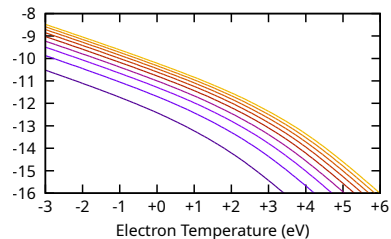
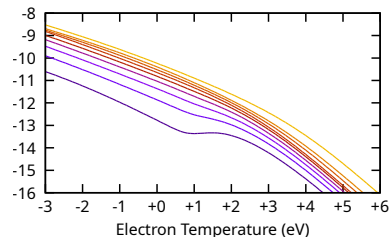
- Generic hydrogenic approximation.
- Valid in both high- and low-temperature limits.

Others: Other sources for specific higher Z elements; eg., Mattioli, et al.^c for Kr.

^aN. R. BADNELL, *Radiative recombination data for modeling dynamic finite-density plasmas*, The Astrophysical Journal Supplement Series, 167 (2006), pp. 334–342.

^bI. A. KOTELNIKOV AND A. I. MILSTEIN, *Electron radiative recombination with a hydrogen-like ion*, Physica Scripta, 94 (2019), p. 055403.

^cM. MATTIOLI, G. MAZZITELLI, K. B. FOURNIER, M. FINKENTHAL, AND L. CARRARO, *Updating of atomic data needed for ionization balance evaluations of krypton and molybdenum*, Journal of Physics B: Atomic, Molecular and Optical Physics, 39 (2006), pp. 4457–4489.



Dielectronic Recombination Rates $\Gamma_{(\alpha,k)}^{\text{rec}} = n_e n_{(\alpha,k)} \left(R_{(\alpha,k)}^{\text{rad}} + R_{(\alpha,k)}^{\text{die}} \right)$



Badnell, et al.:^a

$$R_{(\alpha,k)}^{\text{die}} = T_e^{-3/2} \sum_{i=1}^{N_{(\alpha,k)}} c_i^{(\alpha,k)} \exp\left(-E_i^{(\alpha,k)} / T_e\right)$$

- Available by isoelectronic sequence, through Si sequence.
- Fits of calculated data from AUTOSTRUCTURE code.

Landini, et al.:^b

$$R_{(\alpha,k)}^{\text{die}} = A_{(\alpha,k)} T_e^{-3/2} \exp\left(-T_0^{(\alpha,k)} / T_e\right) \left(1 + B_{(\alpha,k)} \exp\left(-T_1^{(\alpha,k)} / T_e\right)\right)$$

- Less resolved, but data available for a wider range of species (eg., low charge states of Ar).

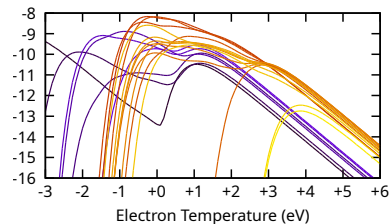
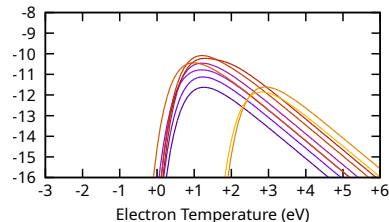
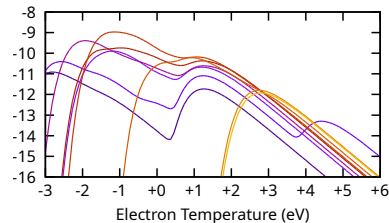
Others: Other sources for specific higher Z elements; eg., Mattioli, et al.^c or Sterling^d for Kr.

^aN. R. BADNELL, M. G. O'MULLANE, H. P. SUMMERS, Z. ALTUN, M. A. BAUTISTA, J. COLGAN, T. W. GORCZYCA, D. M. MITNIK, M. S. PINDZOLA, AND O. ZATSARINNY, *Dielectronic recombination data for dynamic finite-density plasmas: I. Goals and methodology*, *Astronomy & Astrophysics*, 406 (2003), pp. 1151–1165.

^bM. LANDINI AND B. C. MONSIGNORI FOSSI, *The X-UV spectrum of thin plasmas*, *Astronomy & Astrophysics Supplement Series*, 82 (1990), pp. 229–260.

^cM. MATTIOLI, G. MAZZITELLI, K. B. FOURNIER, M. FINKENTHAL, AND L. CARRARO, *Updating of atomic data needed for ionization balance evaluations of krypton and molybdenum*, *Journal of Physics B: Atomic, Molecular and Optical Physics*, 39 (2006), pp. 4457–4489.

^dN. C. STERLING, *Atomic data for neutron-capture elements II. photoionization and recombination properties of low-charge krypton ions*, *Astronomy & Astrophysics*, 533 (2011), p. A62.



Verification: Two-Species Damped Oscillation (Constant coefficients)



Two-fluid system with no spatial gradients:

$$\begin{aligned}\partial_t (\rho_s \mathbf{u}_s) &= q_s n_s (\mathbf{E} + \mathbf{u}_s \times \mathbf{B}) + \alpha_{s;t} \rho_s \rho_t (\mathbf{u}_t - \mathbf{u}_s) \\ \partial_t \mathcal{E}_s &= q_s n_s \mathbf{u}_s \cdot \mathbf{E} + \frac{\alpha_{s;t} \rho_s \rho_t}{m_s + m_t} [A_{s;t} k_B (T_t - T_s) + m_t (\mathbf{u}_t - \mathbf{u}_s)^2] \\ \partial_t \mathbf{E} &= -\frac{1}{\epsilon_0} \mathbf{J}\end{aligned}$$

Analytic solution can be derived if $\alpha_{s;t}$ is constant:

$$\begin{aligned}\mathbf{u}_e(t) - \mathbf{u}_i(t) &= (\mathbf{u}_e(t_0) - \mathbf{u}_i(t_0)) \exp\left(-\frac{\nu_{e;i} t}{2}\right) \frac{\cos(\eta_{e;i} t - \phi_{e;i})}{\cos \phi_{e;i}} \\ \mathbf{E}(t) &= -(\mathbf{u}_e(t_0) - \mathbf{u}_i(t_0)) \frac{q_e n_i}{\epsilon_0} \frac{\exp(-\nu_{e;i} t/2)}{\omega_{e;i}^2 \cos \phi_{e;i}} \left[\eta_{e;i} \sin(\eta_{e;i} t - \phi_{e;i}) - \frac{\nu_{e;i}}{2} \cos(\eta_{e;i} t - \phi_{e;i}) + \exp(\nu_{e;i} t/2) \left(\frac{\nu_{e;i}}{2} \cos \phi_{e;i} + \eta_{e;i} \sin \phi_{e;i} \right) \right] \\ T_e(t) - T_i(t) &= \exp(-A_{e;i} t) C_{e;i} + \frac{D_{e;i} \exp(-\nu_{e;i} t)}{\cos^2 \phi_{e;i}} \left[4\eta_{e;i}^2 + 2G_{e;i}^2 \cos^2(\eta_{e;i} t - \phi_{e;i}) + 4\eta_{e;i} G_{e;i} \sin(\eta_{e;i} t - \phi_{e;i}) \cos(\eta_{e;i} t - \phi_{e;i}) \right]\end{aligned}$$

Equilibrium velocity and temperature can be computed by conservation of momentum and energy (ideal gas):

$$\mathbf{u}_* = \frac{\sum_s \rho_s \mathbf{u}_s}{\sum_s \rho_s}, \quad T_* = \frac{\sum_s \mathcal{E}_s - \frac{1}{2} \mathbf{u}_*^2 \sum_s \rho_s}{\frac{1}{\gamma-1} k_B \sum_s n_s}.$$

⇒ Compare Drekar and ODE solver (Python odeint) with analytic solution and equilibrium values.

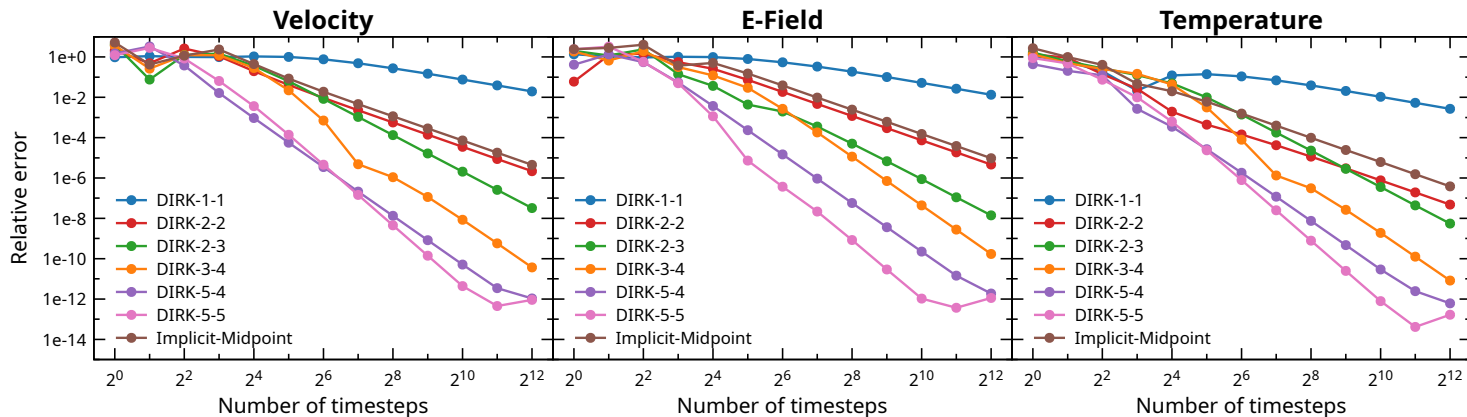
Verification: Two-Species Damped Oscillation (Constant coefficients)



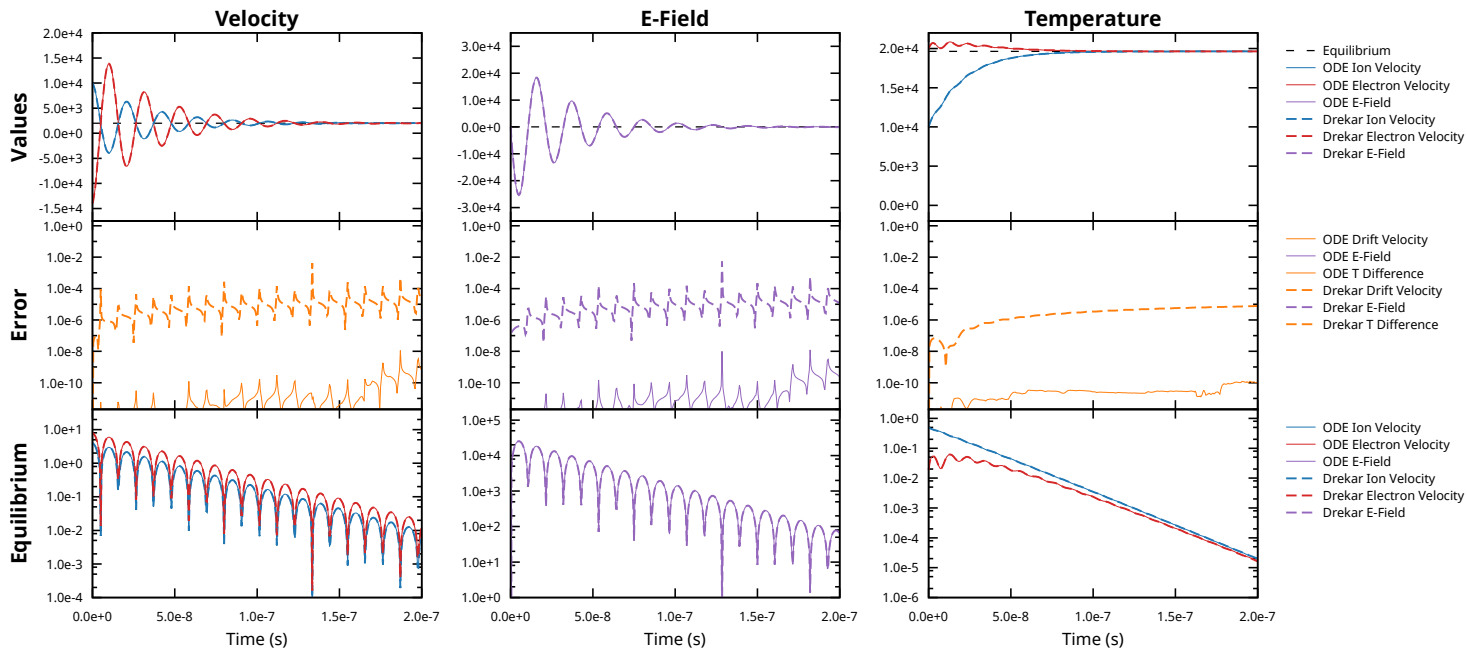
$$\alpha_{e;i} \equiv 1.0\text{E}+18$$

	Ion (i)	Electron (e)
m_s	2.0E-27	1.0E-27
T_s	1.0E+4	2.0E+4
\mathbf{u}_s	$\begin{bmatrix} 1.0\text{E}+4 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} -1.4\text{E}+4 \\ 0.0 \\ 0.0 \end{bmatrix}$
n_s	2.0E+16	2.0E+16

	Ion (i)	Electron (e)
ρ_s	4.0E-11	2.0E-11
$\rho_s \mathbf{u}_s$	$\begin{bmatrix} 4.0\text{E}-7 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} -2.8\text{E}-7 \\ 0.0 \\ 0.0 \end{bmatrix}$
$\rho_s e_s$	4.1430E-3	8.2860E-3
\mathcal{E}_s	6.1430E-3	1.0246E-2



Verification: Two-Species Damped Oscillation (Constant coefficients)

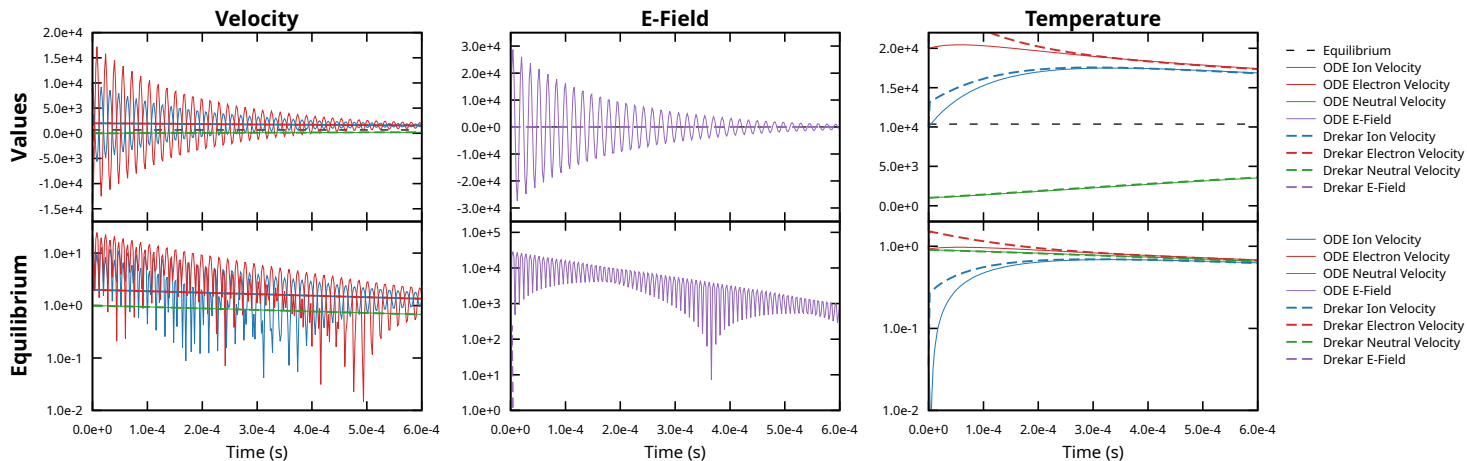


Verification: Three-Species Damped Oscillation (Physical coefficients)



	Neutral (n)	Ion (i)	Electron (e)
m_s	3.0E-27	2.0E-27	1.0E-27
T_s	1.0E+3	1.0E+4	2.0E+4
\mathbf{u}_s	$\begin{bmatrix} 0.0 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} 1.0E+4 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} -1.4E+4 \\ 0.0 \\ 0.0 \end{bmatrix}$
n_s	4.0E+16	2.0E+16	2.0E+16

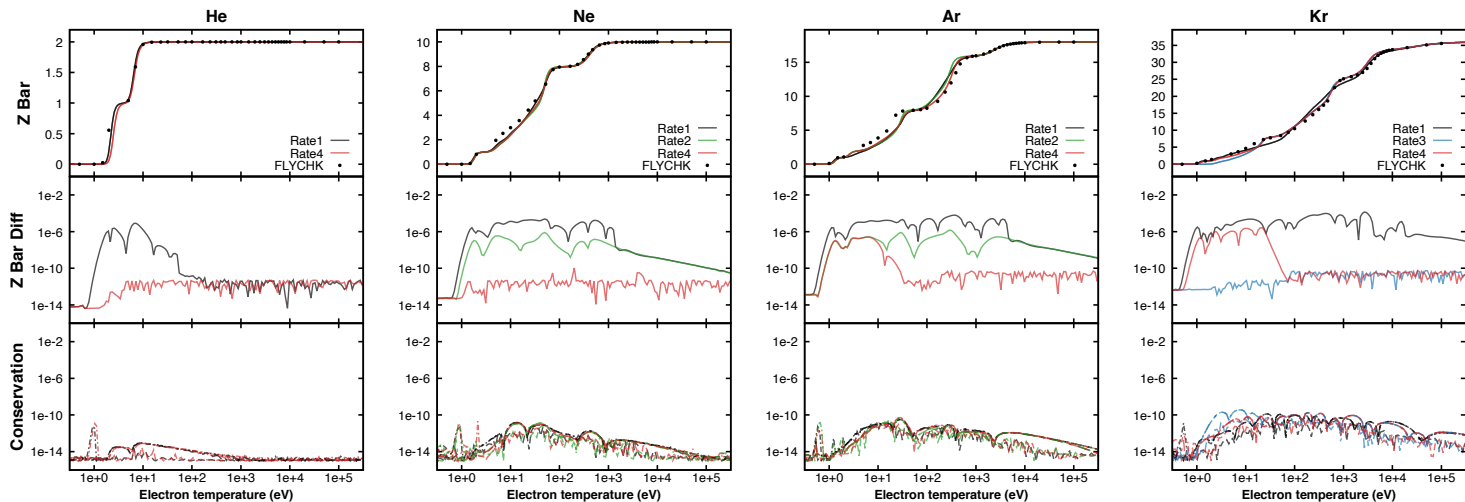
	Neutral (n)	Ion (i)	Electron (e)
ρ_s	1.2E-10	4.0E-11	2.0E-11
$\rho_s \mathbf{u}_s$	$\begin{bmatrix} 0.0 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} 4.0E-7 \\ 0.0 \\ 0.0 \end{bmatrix}$	$\begin{bmatrix} -2.8E-7 \\ 0.0 \\ 0.0 \end{bmatrix}$
$\rho_s e_s$	8.2860E-4	4.1430E-3	8.2860E-3
\mathcal{E}_s	8.2860E-4	6.1430E-3	1.0246E-2



Verification: Ionization/Recombination



- Equilibrium ion fractions computed by linear solve.
- Compare long-time integration (with non-equilibrium initial condition) to equilibrium \bar{z} .
- Long-time solution recovers equilibrium values.
- Conservation of total mass and momentum is also good.
- Rate coefficients give equilibrium values comparable to FLYCHK library.



Hyperbolic system for each fluid:

$$\partial_t \mathbf{U}_s + \nabla \cdot \mathbf{F}_s(\mathbf{U}_s) = \mathbf{S}_s(\mathbf{U}, \mathbf{E}, \mathbf{B}), \quad \mathbf{U}_s = (\rho_s, \rho_s \mathbf{u}_s, \mathcal{E}_s)^T.$$

Semi-discrete scheme:

$$\mathbf{M}_C \partial_t \hat{\mathbf{U}}_s + \mathbf{K}_s(\hat{\mathbf{U}}_s) + \mathbf{B}_s(\hat{\mathbf{U}}_s) + \mathbf{S}_s(\hat{\mathbf{U}}, \hat{\mathbf{E}}, \hat{\mathbf{B}}) = \mathbf{0},$$

where

$$\mathbf{M}_C = \{M_{k,\ell} = m_{k,\ell} I_{N \times N}\}_{k,\ell}$$

$$\mathbf{K}_s(\hat{\mathbf{U}}_s) = \{\mathbf{k}_{s,k}\}_k$$

$$\mathbf{B}_s(\hat{\mathbf{U}}_s) = \{\mathbf{b}_{s,k}\}_k$$

$$\mathbf{S}_s(\hat{\mathbf{U}}, \hat{\mathbf{E}}, \hat{\mathbf{B}}) = \{\mathbf{S}_{s,k}\}_k$$

$$m_{k,\ell} = \int_{\Omega} \phi_k \phi_{\ell} d\mathbf{x}$$

$$\mathbf{k}_{s,k} = - \int_{\Omega} \nabla \phi_k \cdot \mathbf{F}_s(\mathbf{U}_s) d\mathbf{x}$$

$$\mathbf{b}_{s,k} = \int_{\Gamma} \phi_k \mathbf{F}_s(\mathbf{U}_s) \cdot \mathbf{n} d\Gamma$$

$$\mathbf{S}_{s,k} = - \int_{\Omega} \phi_k \mathbf{S}_s(\mathbf{U}, \mathbf{E}, \mathbf{B}) d\mathbf{x}$$

AFC Stabilization



Stabilized system:

$$\mathbf{M}_L \partial_t \hat{\mathbf{U}}_s + \mathbf{K}_s(\hat{\mathbf{U}}_s) + \mathbf{B}_s(\hat{\mathbf{U}}_s) + \mathbf{S}_s(\hat{\mathbf{U}}, \hat{\mathbf{E}}, \hat{\mathbf{B}}) + \mathbf{D}_s(\hat{\mathbf{U}}_s) \hat{\mathbf{U}}_s + \alpha_s \mathcal{F}_s = \mathbf{0},$$

where

$$\mathcal{F}_s = -\mathbf{D}_s(\hat{\mathbf{U}}_s) \hat{\mathbf{U}}_s + (\mathbf{M}_C - \mathbf{M}_L) \partial_t \hat{\mathbf{U}}_s \quad \text{and} \quad \alpha_s \in [0, 1].$$

Question: How to choose \mathbf{D}_s , α_s , \mathcal{F}_s for each species?

First try: Stabilize each fluid separately.

- \mathbf{D}_s is constructed based on the sound speed of species s .
- α_s is chosen based on the behavior of species s .

Alternative: Synchronize stabilization between fluids.

- \mathbf{D}_s is constructed based on the maximum sound speed of *all* species (all \mathbf{D}_s equal).
- α_s is chosen based on the behavior of *all* species (all α_s equal).

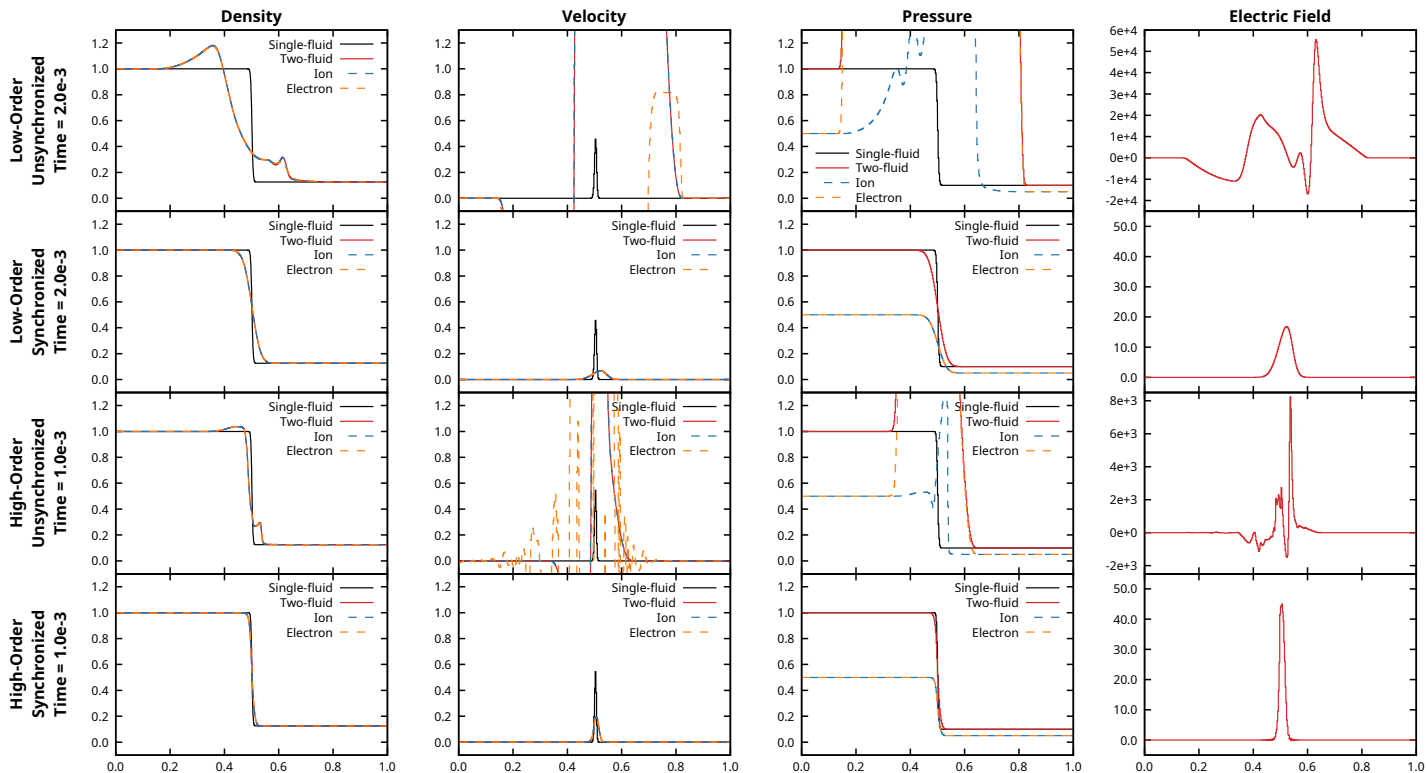
Two-Fluid Sod Problem



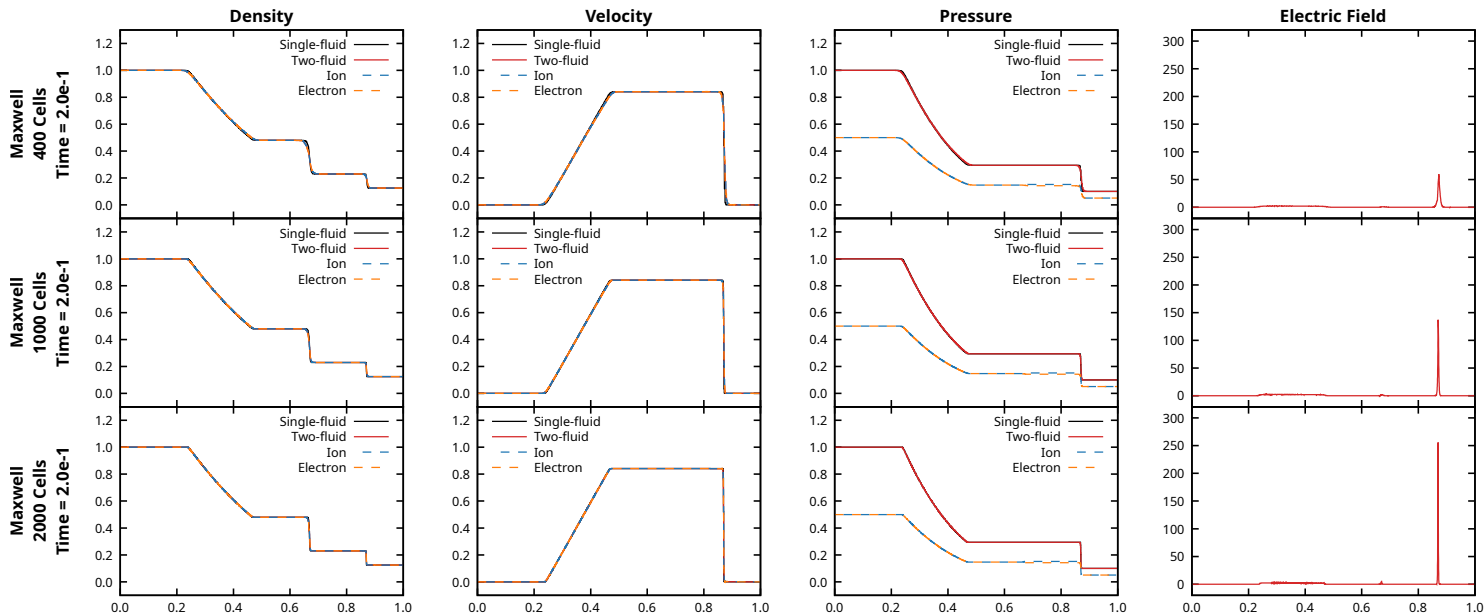
- Position $x \in [0, 1]$
- Time $t \in [0, 0.2]$
- Density $\rho = \begin{cases} 1.0, & x < 0.5 \\ 0.125, & x > 0.5 \end{cases}$
- Pressure $p = \begin{cases} 1.0, & x < 0.5 \\ 0.1, & x > 0.5 \end{cases}$
- $\varepsilon =$ Particle mass ratio.
- $\varepsilon_0 = \varepsilon^2$
- Two-fluid solution recovers single-fluid behavior as $\varepsilon \rightarrow 0$.
- Use $\varepsilon = 1\text{E-}4$

Quantity	Electron	Ion
Mass (m_s)	$\frac{\varepsilon}{1+\varepsilon}$	$\frac{1}{1+\varepsilon}$
Density (ρ_s)	$\frac{\varepsilon}{1+\varepsilon}\rho$	$\frac{1}{1+\varepsilon}\rho$
Pressure (p_s)	$p/2$	$p/2$

Two-Fluid Sod Problem: Stabilization (Electrostatic)



Two-Fluid Sod Problem: Electrostatic Mesh Refinement



Divergence Cleaning for Maxwell's Equations



Divergence constraints *must* be adequately satisfied!

$$\nabla \cdot \mathbf{B} = 0,$$

$$\epsilon_0 \nabla \cdot \mathbf{E} = q.$$

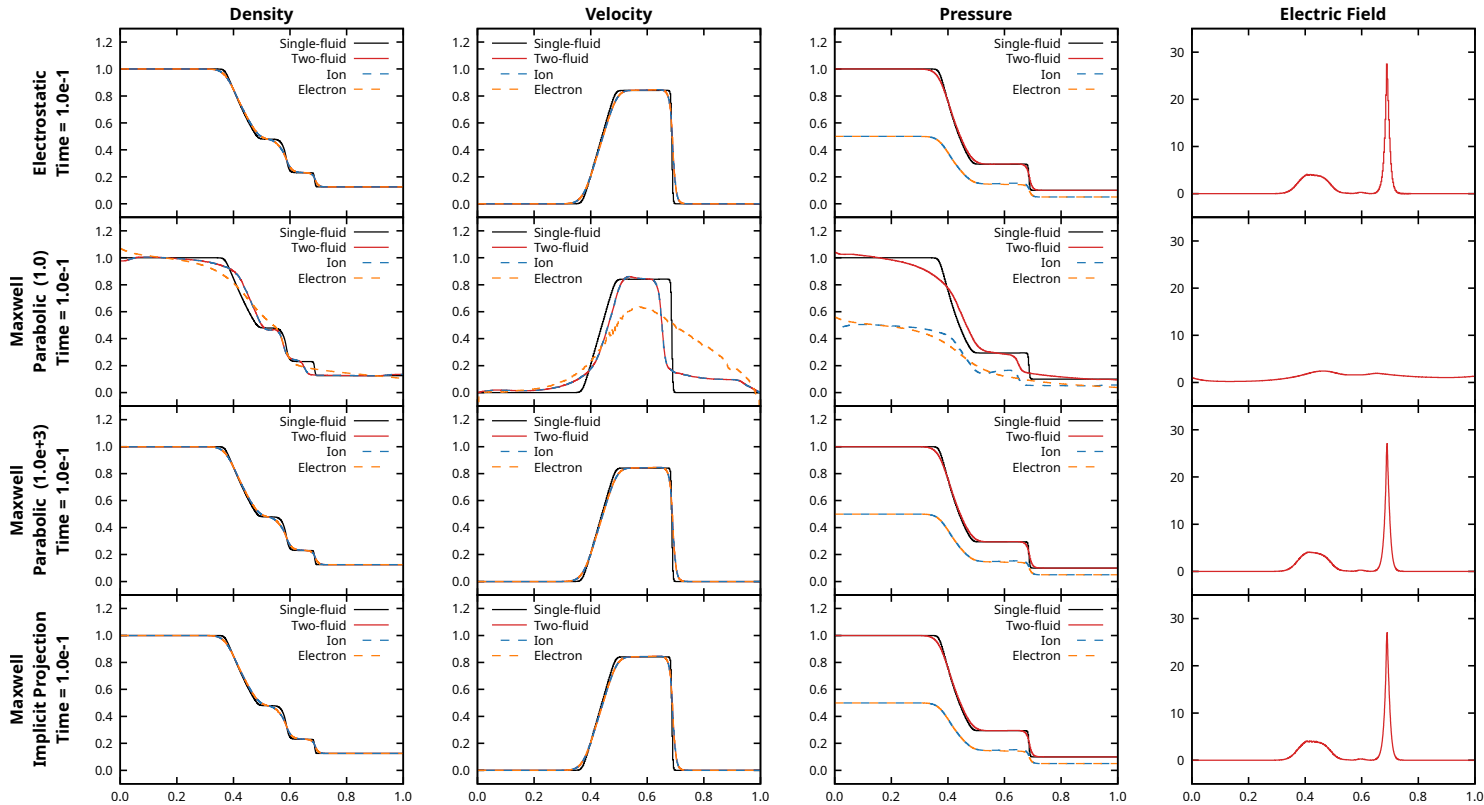
Eliminated Parabolic Cleaning: Add penalty term to Ampere's Law: $\nabla \cdot [c_p^2 (\epsilon_0 \nabla \cdot \mathbf{E} - q) \mathbf{I}]$

- Requires c_p parameter to be specified.
- c_p must be “large enough”.

Implicit Projection: Set $\hat{\mathbf{E}} = \mathbf{E} + \nabla\phi$, where $\epsilon_0 \Delta\phi + \epsilon_0 \nabla \cdot \mathbf{E} - q = 0$.

- Requires elliptic solve.
- Controls divergence error to machine/solver precision.
- Boundary conditions can be non-trivial.

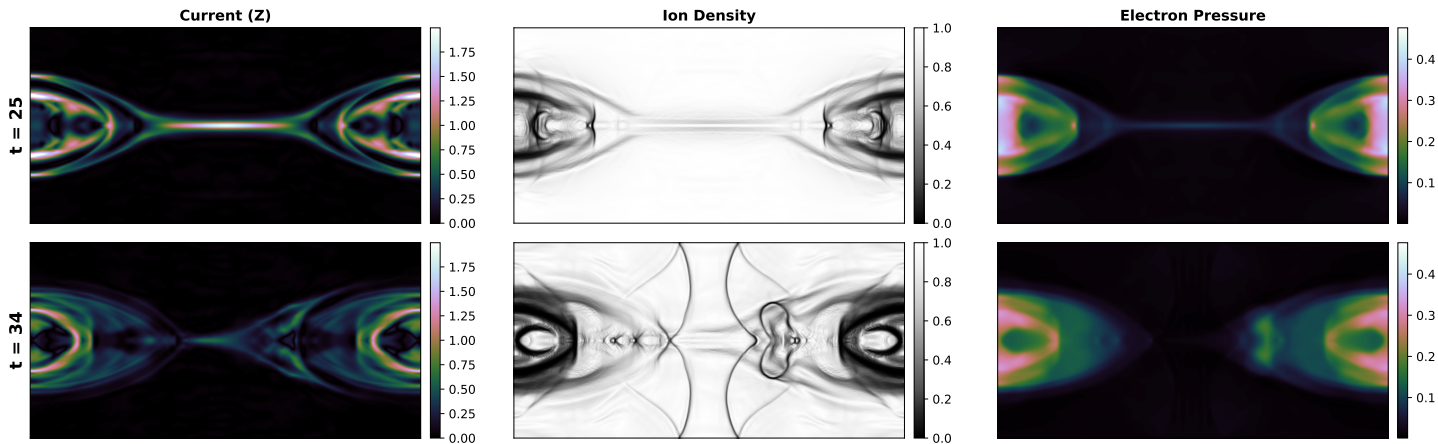
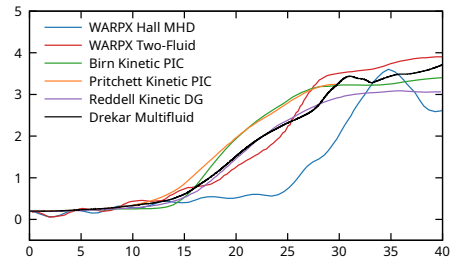
Two-Fluid Sod Problem: Maxwell Divergence Cleaning



GEM Challenge Problem



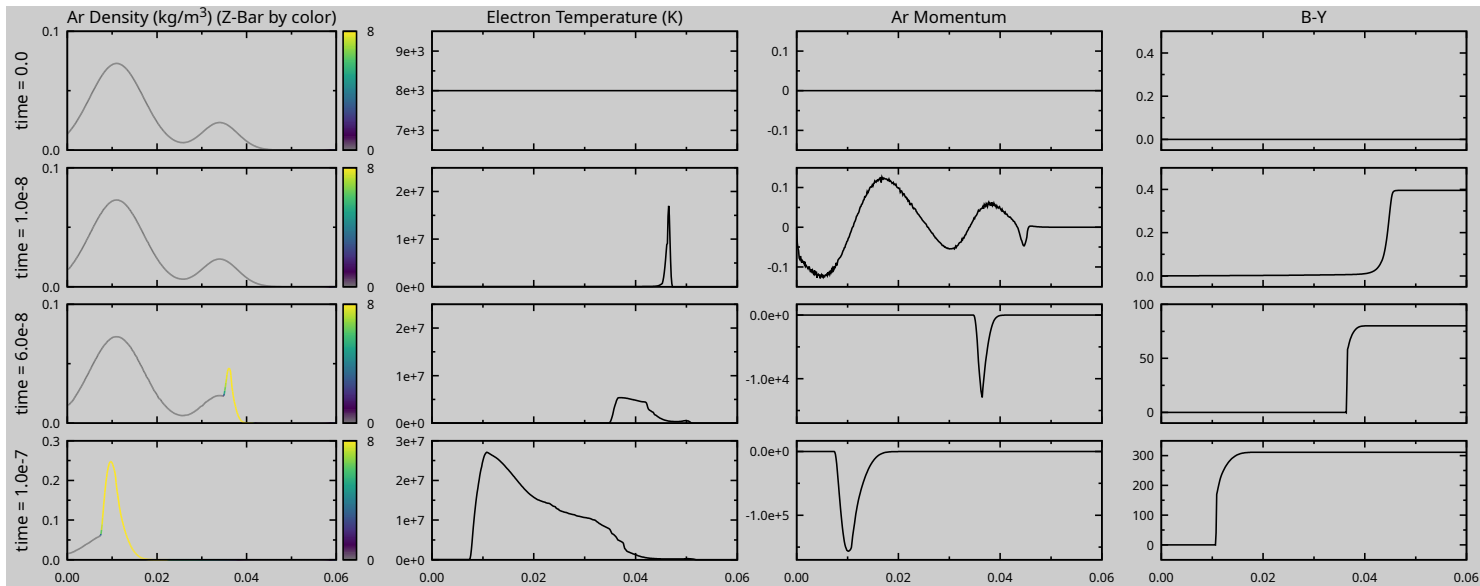
- Test magnetic reconnection of a perturbed current sheet equilibrium.
- Electron/Hall effects are crucial (resistive MHD behaves differently).
- Standard comparison is based on integral of magnetic field.
- Our results compare favorably with kinetic and other two-fluid codes.



1D Argon Gas Puff (Proof of Concept)



- Argon gas ($z = 0$ to 8^+) plus electrons.
- Potential form of Maxwell's equations.
- Driven by EM field applied at the boundary.
- Ionization + collisions yields resistive heating.
- Implicit time integration follows ion fluid CFL.
- Overall behavior is qualitatively reasonable.





Thank you