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# Magneto-inertial fusion and high energy density science using pulsed power

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For National Academies of Science High Energy Density Science Study  
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# High energy density science has diverse applications and spans enormous ranges of time and length scales

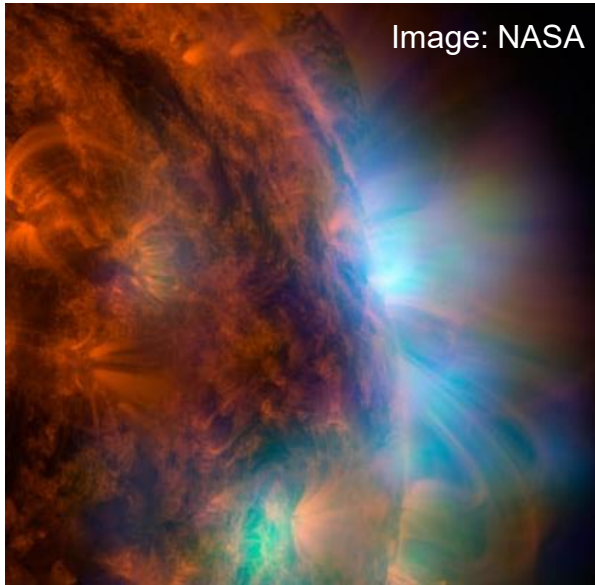


Image: NASA

**Astrophysics**  
 $10^{+9}$  meters  
 $10^{+17}$  seconds

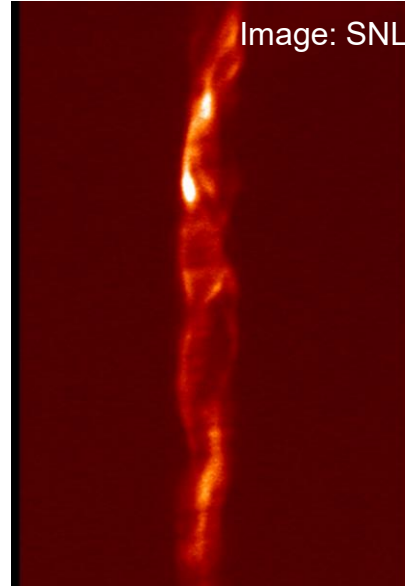
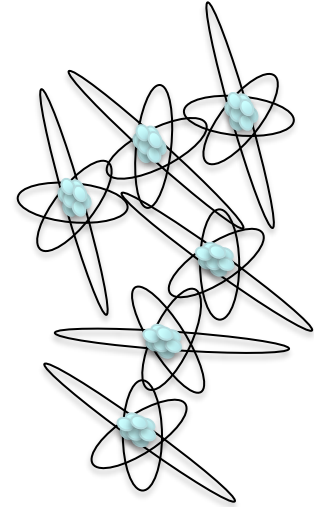


Image: SNL

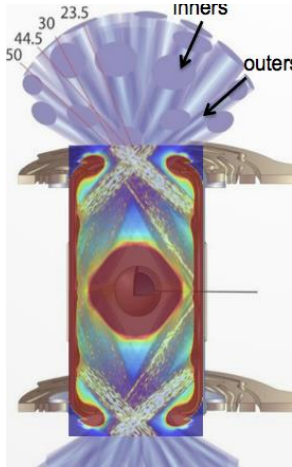
**Fusion science**  
 $10^{-4}$  meters  
 $10^{-9}$  seconds



**Atomic physics**  
 $10^{-10}$  meters  
 $10^{-14}$  seconds

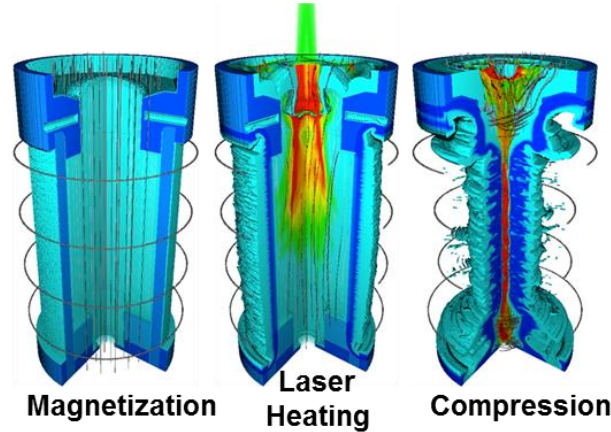
Understanding material and system behavior across scales is a key feature of HED science

# High energy density matter is created in experimental facilities that compress energy in space and time



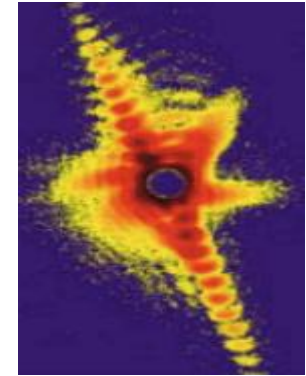
## LLNL's NIF:

422 MJ  $\rightarrow$  1.8 MJ in  $10^{-8}$ s, 1 cm  
~2 MJ/cc, ~200 TW/cc  
~0.4% wall-plug efficiency



## SNL's Z machine:

22 MJ  $\rightarrow$  0.5-1 MJ in  $10^{-7}$ s, 1 cm  
~1 MJ/cc, ~10 TW/cc  
2 – 4% wall-plug efficiency

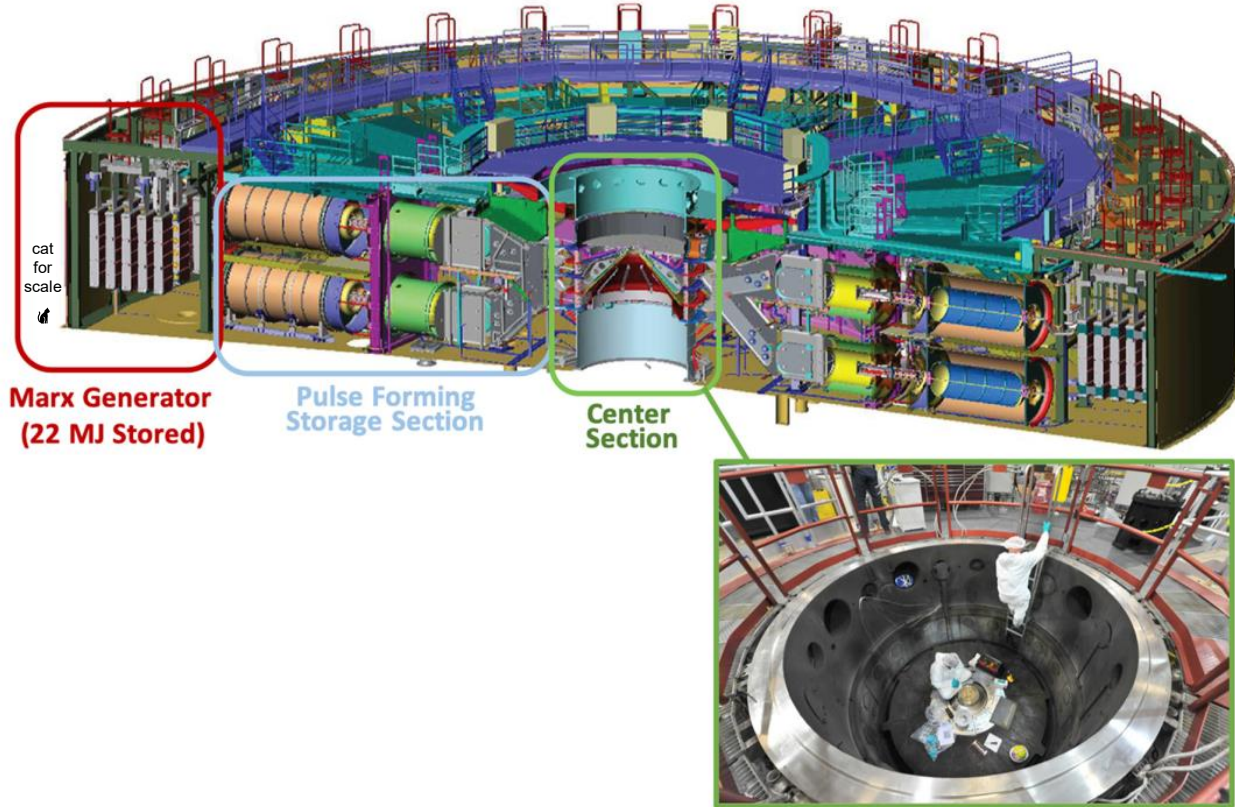


## LCLS/ European XFEL:

2 mJ in  $10^{-13}$ s, 10  $\mu$ m  
~2 kJ/cc, ~20 PW/cc

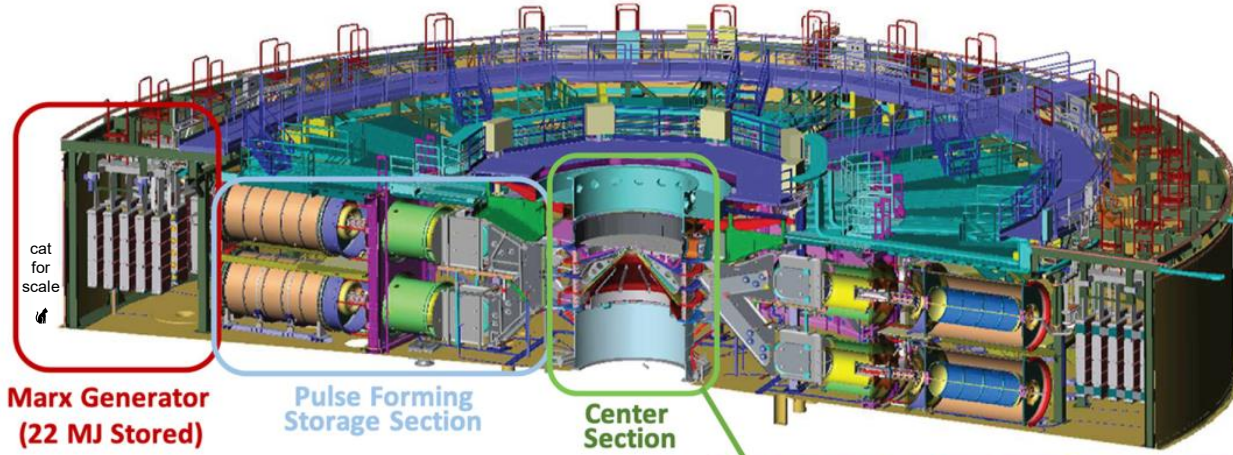
# Pulsed power offers an efficient way to deliver large amounts of energy to relatively large samples

1. capacitors charge up from wall power in ~ 2 minutes (22 MJ)

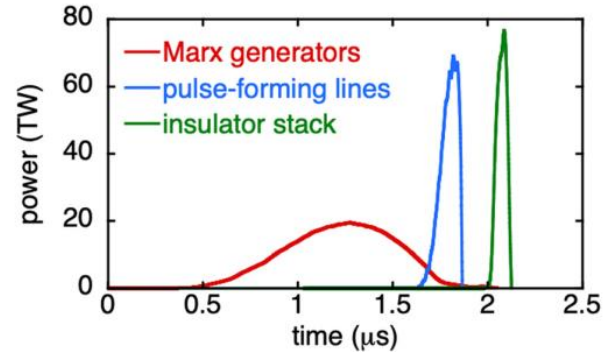




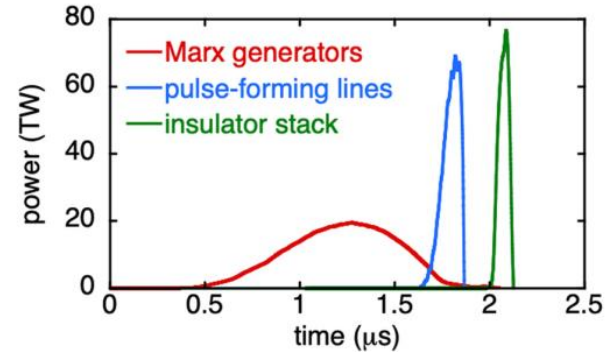
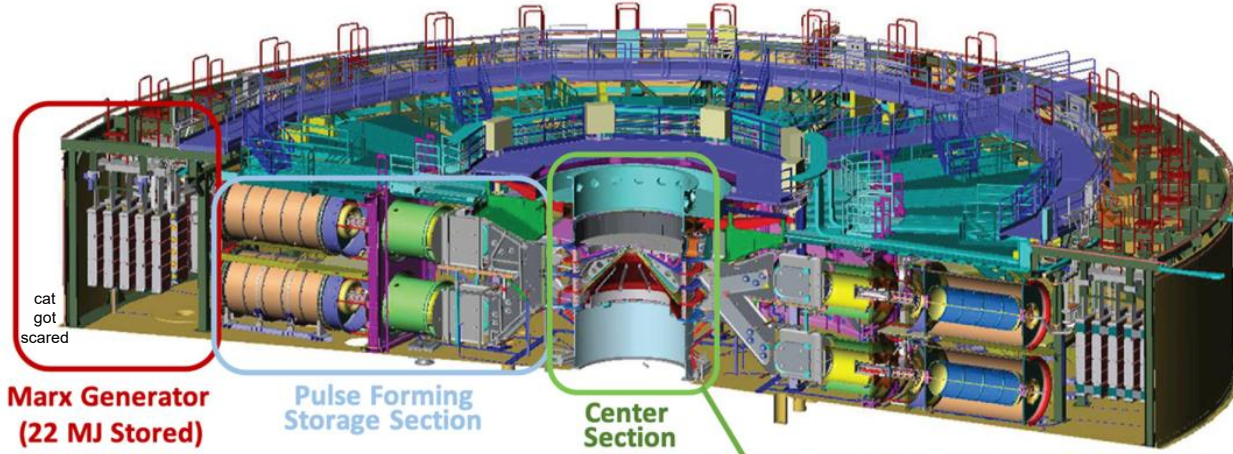
# Pulsed power offers an efficient way to deliver large amounts of energy to relatively large samples



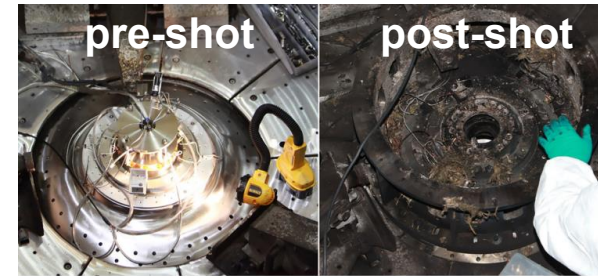
1. capacitors charge up from wall power in  $\sim 2$  minutes (22 MJ)
2. current  $j$  is compressed in space and time to  $>20\text{MA}$  ( $>0.5$  MJ to 1 cm in 100 ns)



# Pulsed power offers an efficient way to deliver large amounts of energy to relatively large samples

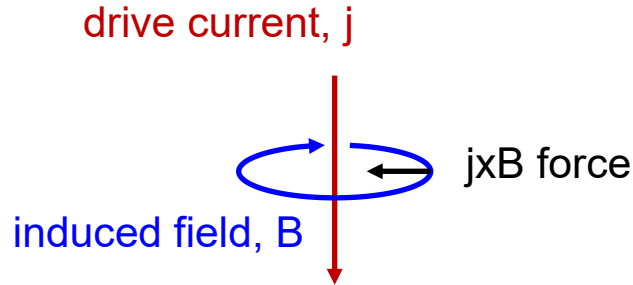


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2. current  $j$  is compressed in space and time to  $>20\text{MA}$  ( $>0.5$  MJ to 1 cm in 100 ns)
3. high energy density science! (the ground shakes)

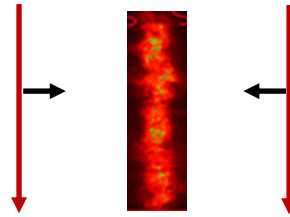


# “Magnetic direct drive” uses the $\mathbf{j} \times \mathbf{B}$ force to deliver energy

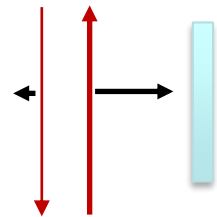
magnetic drive pressure  $\sim 60 \text{ Mbar } (j_{20\text{MA}}/r_{\text{mm}})^2$



26 MA at  $r = 1 \text{ mm} \rightarrow$   
 $\sim 100 \text{ Mbar pressures}$

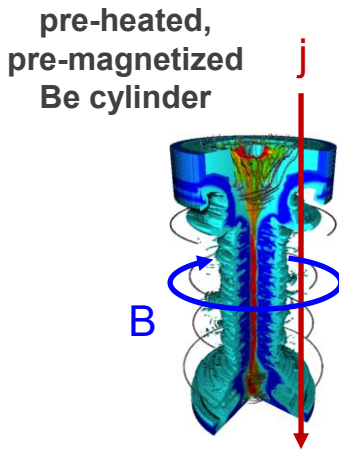


26 MA at  $r \sim 100 \text{ mm} \rightarrow$   
 $v \sim 1000 \text{ km/s}$

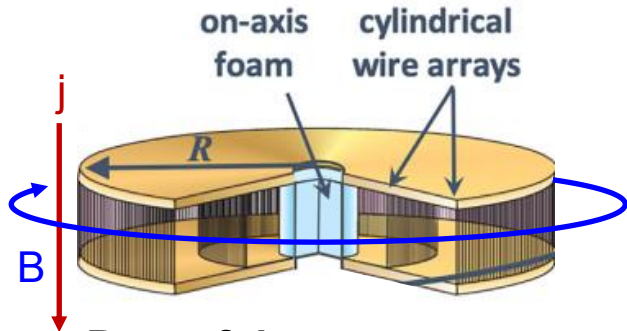


26 MA at  $r \sim 1 \text{ mm} \rightarrow$   
 $> \text{Mbar shocks}$

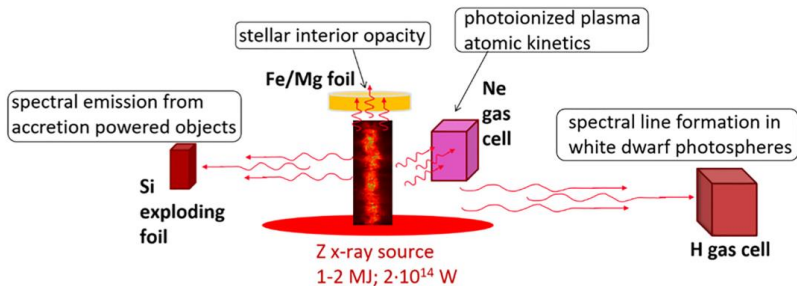
# On Sandia's Z machine, we study a wide range of high energy density science using pulsed power



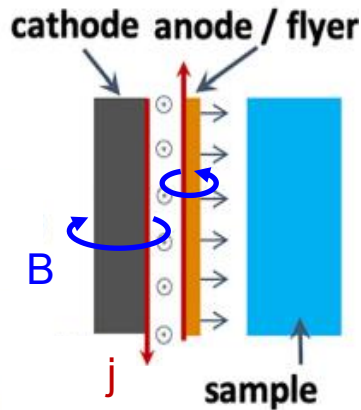
**Magneto-inertial fusion**  
jxB heats and compresses fusion fuel  
(~1 Gbar, > kJ yields)



**Powerful x-ray sources**  
jxB imparts kinetic energy  
(MJ & TW of x-rays)



**X-rays drive fundamental science experiments**

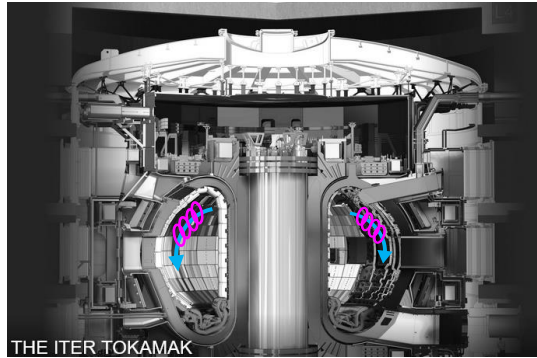


**Materials science**  
jxB provides shock or ramp compression  
(~10 Mbar)



# Magnetized Liner Inertial Fusion (MagLIF) combines aspects of magnetic and inertial confinement fusion

Lawson criterion:  $P\tau > 50 \text{ bar-s}$  (and  $T > 4 \text{ keV}$ )



Magnetic fusion (ITER)

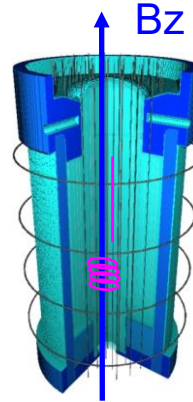
$R \sim 10 \text{ m}$

$P \sim 1 \text{ bar}$

$\tau > 10 \text{ s}$

$P\tau > 10 \text{ bar-s}$

Burn requires plasma  
stability and high  $B$



Magneto-inertial fusion (Z)

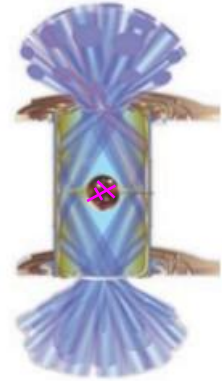
$R \sim 10^{-3} \text{ m}$

$P \sim 10^9 \text{ bar}$

$\tau \sim 10^{-9} \text{ s}$

$P\tau \sim 1 \text{ bar-s}$

Burn requires stability, high  $Bz \cdot R$ ,  
and high  $\rho \cdot Z$



Inertial fusion (NIF)

$R \sim 10^{-5} \text{ m}$

$P \sim 10^{11} \text{ bar}$

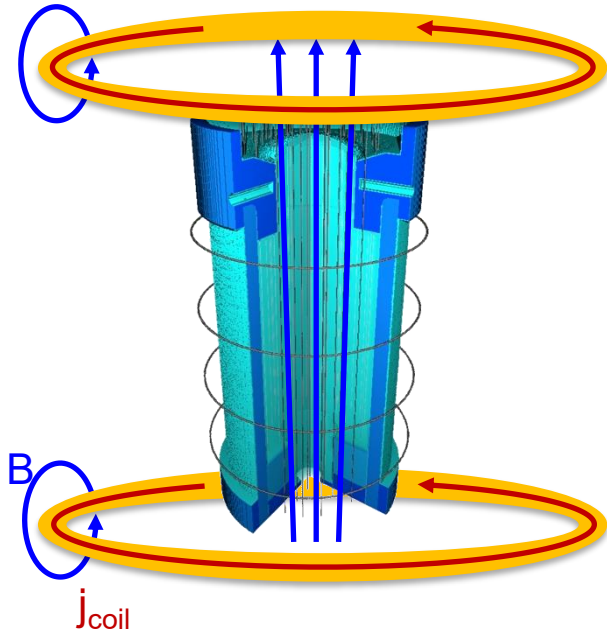
$\tau \sim 10^{-10} \text{ s}$

$P\tau \sim 10 \text{ bar-s}$

Burn requires implosion  
stability and high  $\rho \cdot R$

# MagLIF's 3-stage design uses an imposed axial field and laser preheat to enable a relatively slow, stable implosion

$$B_z = 10 - 20 \text{ T}$$



## Stage 1: pre-magnetize

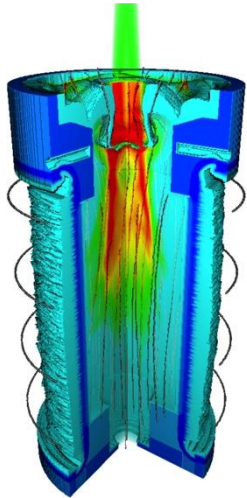
A ~1 cm beryllium liner is filled with fusion fuel and pre-magnetized using magnetic field coils

The imposed axial magnetic field prevents conduction losses that can cool the plasma below fusion temperatures

At stagnation, the imposed B field is flux-compressed to ~10 kT, effectively trapping charged fusion products

# MagLIF's 3-stage design uses an imposed axial field and laser preheat to enable a relatively slow, stable implosion

$$E_{\text{laser}} = 1 - 2 \text{ kJ}$$



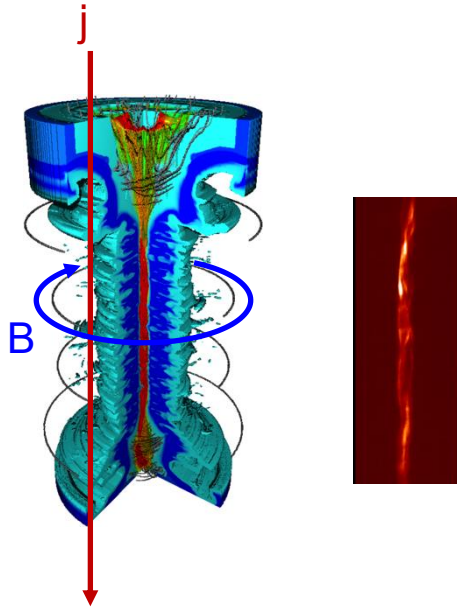
Stage 2: pre-heat the fuel

The 4kJ Z-Beamlet laser penetrates a thin window and preheats the fuel core to  $T \sim 200 \text{ eV}$ , reducing the convergence required to reach fusion temperatures from  $p dV$  compressional work

While the axial B-field effectively prevents conduction losses, even small amounts of window mix can lead to radiative losses that cool the preheated plasma

# MagLIF's 3-stage design uses an imposed axial field and laser preheat to enable a relatively slow, stable implosion

$$j = 15 - 21 \text{ MA}$$



## Stage 3: compression

The Z machine delivers a compressional force that implodes the liner, axial field, and fuel at  $v \sim 100 \text{ km/s}$

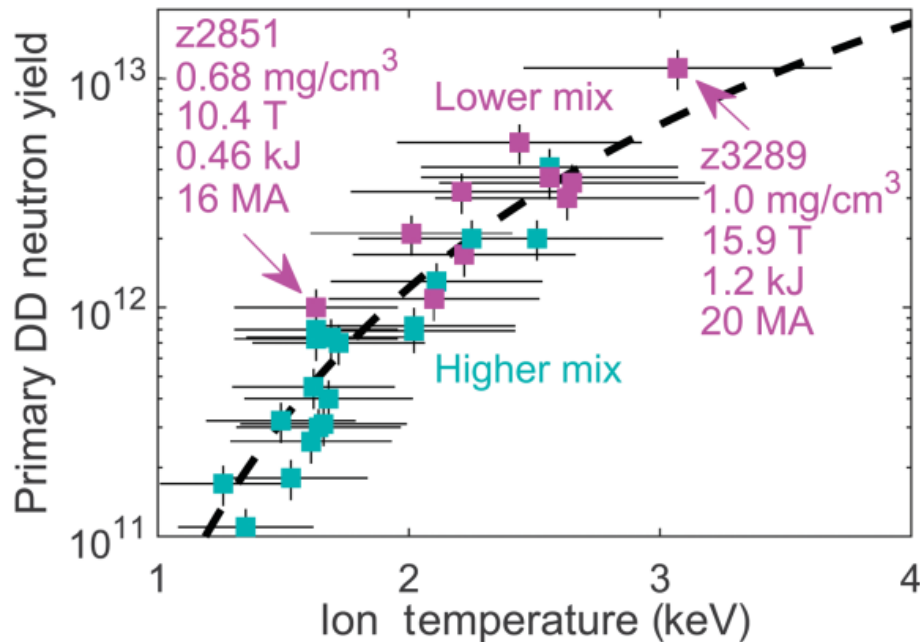
At stagnation, relatively stable plasma columns with helical structure yield  $\sim 10^{13}$  D-D neutrons ( $10^{15}$  D-T)

Neutron diagnostic indicate high magnetization sufficient to trap  $\sim 50\%$  of charged fusion products

X-ray diagnostics indicate  $T \sim 3 \text{ keV}$ ,  $\tau \sim 1 \text{ ns}$ , and  $P \sim 1 \text{ Gbar}$



# We can vary experimental parameters to investigate mix, stability, and yield scaling

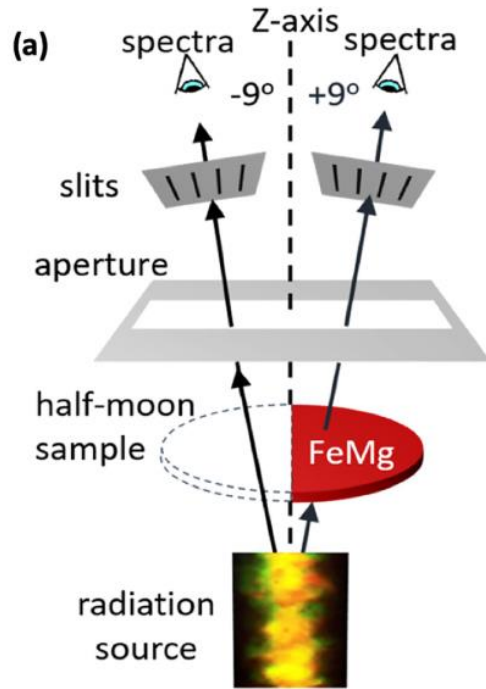


$j(\text{MA})$	$E_{\text{laser}}(\text{kJ})$	$B_z(\text{T})$	$Y(\text{kJ})$	Z
15	0.5	10	0.2	
20	1.2	20	2	
22	6	30	80	
60	30*	30	>1000	

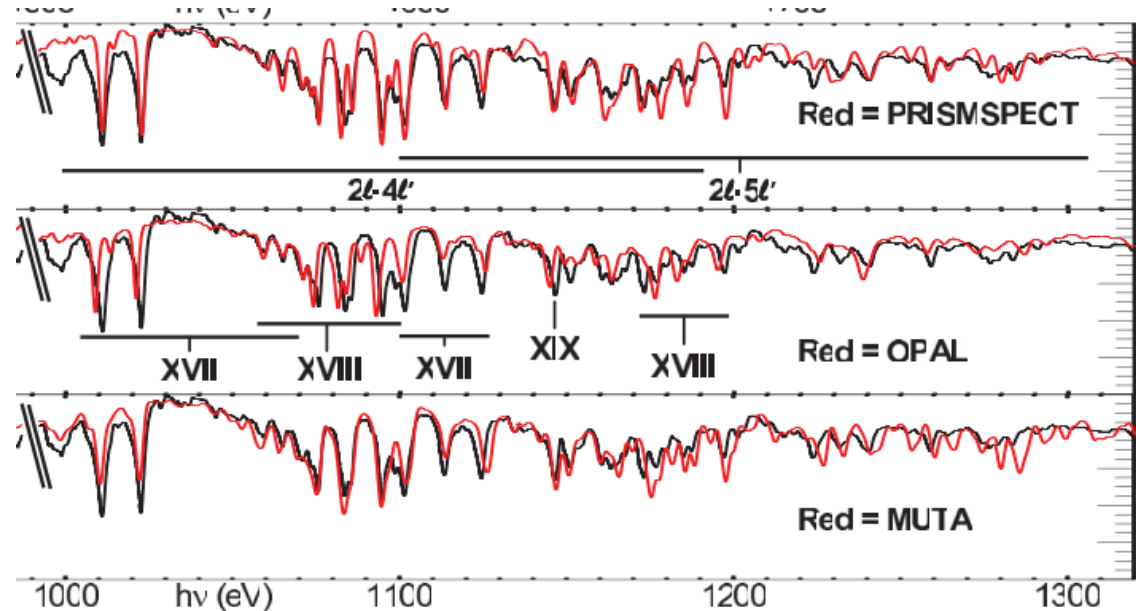
The efficiency of pulsed power opens paths to controlled high-yield fusion

\*Experiments at NIF are ongoing, along with “mini-MagLIF” at LLE

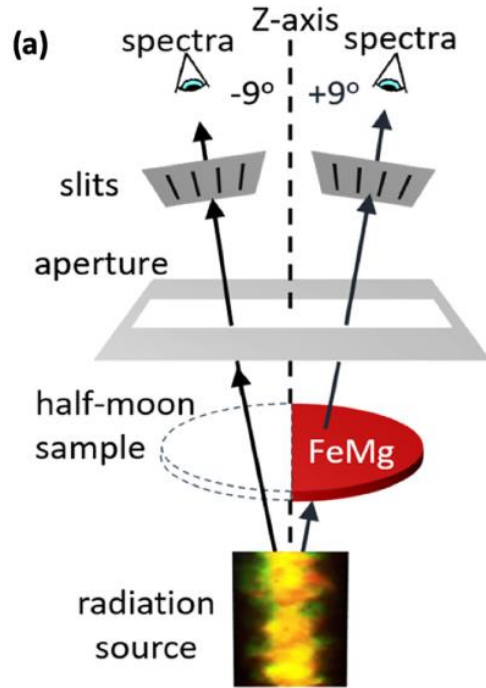
# Benchmark measurements of stellar interior opacities inform models of our sun (helioseismology, elemental abundances)



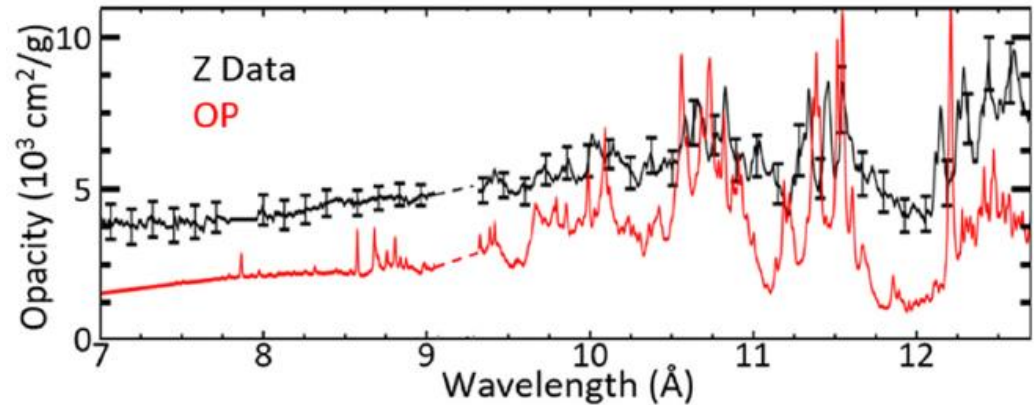
In 2007, Bailey et al found good agreement between models and experiments for iron at temperatures and densities slightly lower than that at the solar radiation/convection zone boundary:



# Benchmark measurements of stellar interior opacities inform models of our sun (helioseismology, elemental abundances)



After a refurbishment of the Z machine enabled experiments at higher densities and temperatures, Bailey et al found surprising disagreement between models and experiments



This is one of only a handful of benchmark experiments for high energy density plasmas: we will be surprised again.

# Key opportunities and existing gaps in pulsed power ICF/HEDP Sandia National Laboratories

- Z has been an engine of discovery: designed in 1996 to drive wire arrays, today 90% of Z shots are on ~5 platforms developed in the last 10-20 years. This innovation and diversity offers risk mitigation for ICF scaling and profound opportunities for fundamental science, but challenges personnel, target fab, diagnostics, & machine operation & design
- Community (esp. science of pulsed power) is small, with people/funding/historical effort 10-100x less than laser community – and mid-scale university facilities that drive innovation & support pipeline have uncertain futures
- Pulsed power drivers are inherently coupled to targets/loads through the load inductance; matching machines to loads increases efficiency (but requires people & design tools) (Even “low efficiency” PP is ~1% → high yield ICF)
- While current loss mechanisms tend to be linear and self-limiting, we do not yet have predictive whole-machine models that predict loss & attendant plasma flow to load
- Measurements of current flow at the target surface are difficult; existing simulation codes have significant & fundamental uncertainties in low-density/near-vacuum regions that are compounded by uncertainty in loss models
- Target design and diagnostic interpretation rely on constitutive data across a very wide range of temperature-density regimes in which only a handful (literally ~ 5) of true benchmarks are available



# High energy density science with pulsed power relies on strong ties with larger communities

