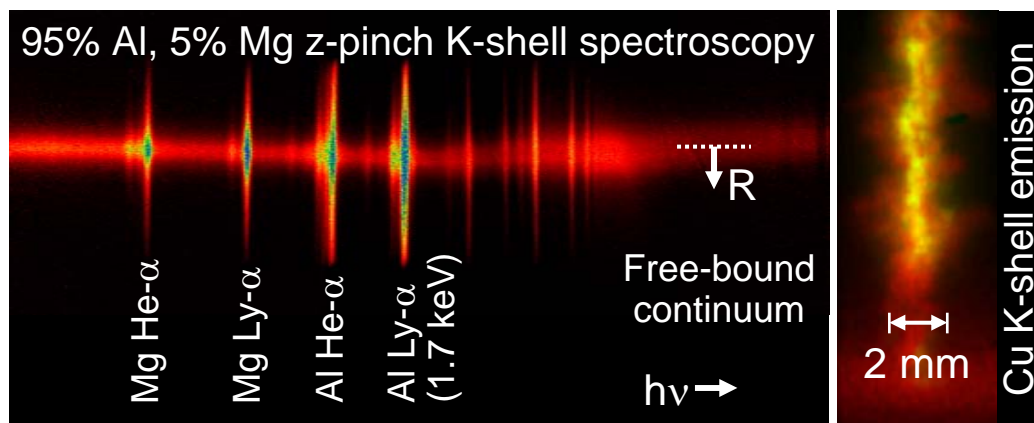


Spectroscopic diagnosis and progress toward understanding thermalization physics in low- to mid-atomic-number z pinches on the Z machine

B. Jones¹, C. A. Coverdale¹, C. Deeney², P. D. LePell³, J. W. Thornhill⁴, A. L. Velikovich⁴, R. W. Clark⁴, J. P. Apruzese⁴, J. Davis⁴, K. G. Whitney⁵, Y. Maron⁶



¹*Sandia National Laboratories*

²*NNSA, US DOE*

³*Ktech Corporation*

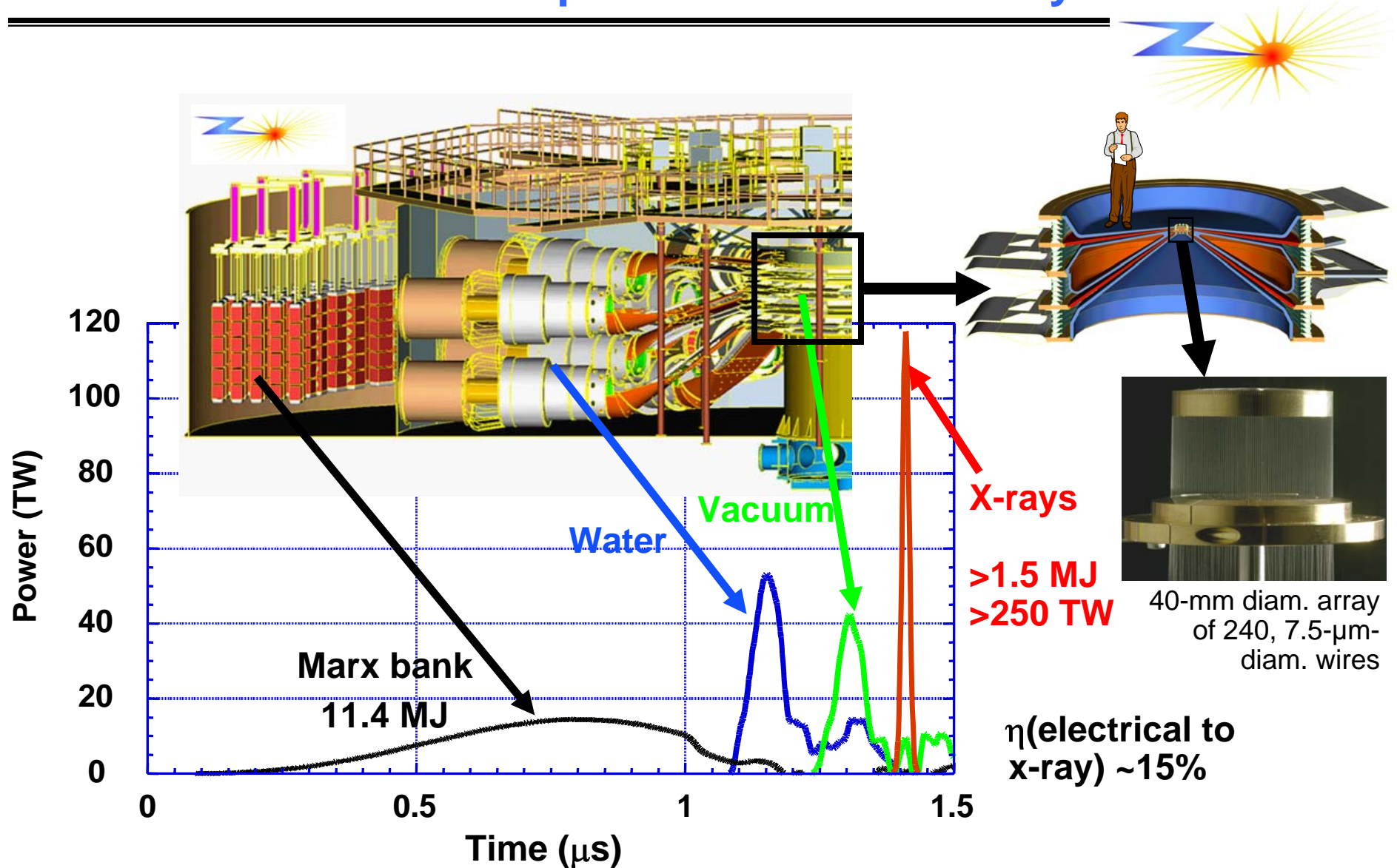
⁴*Naval Research Laboratory*

⁵*Berkeley Scholars*

⁶*Weizmann Institute of Science*

May 14, 2007

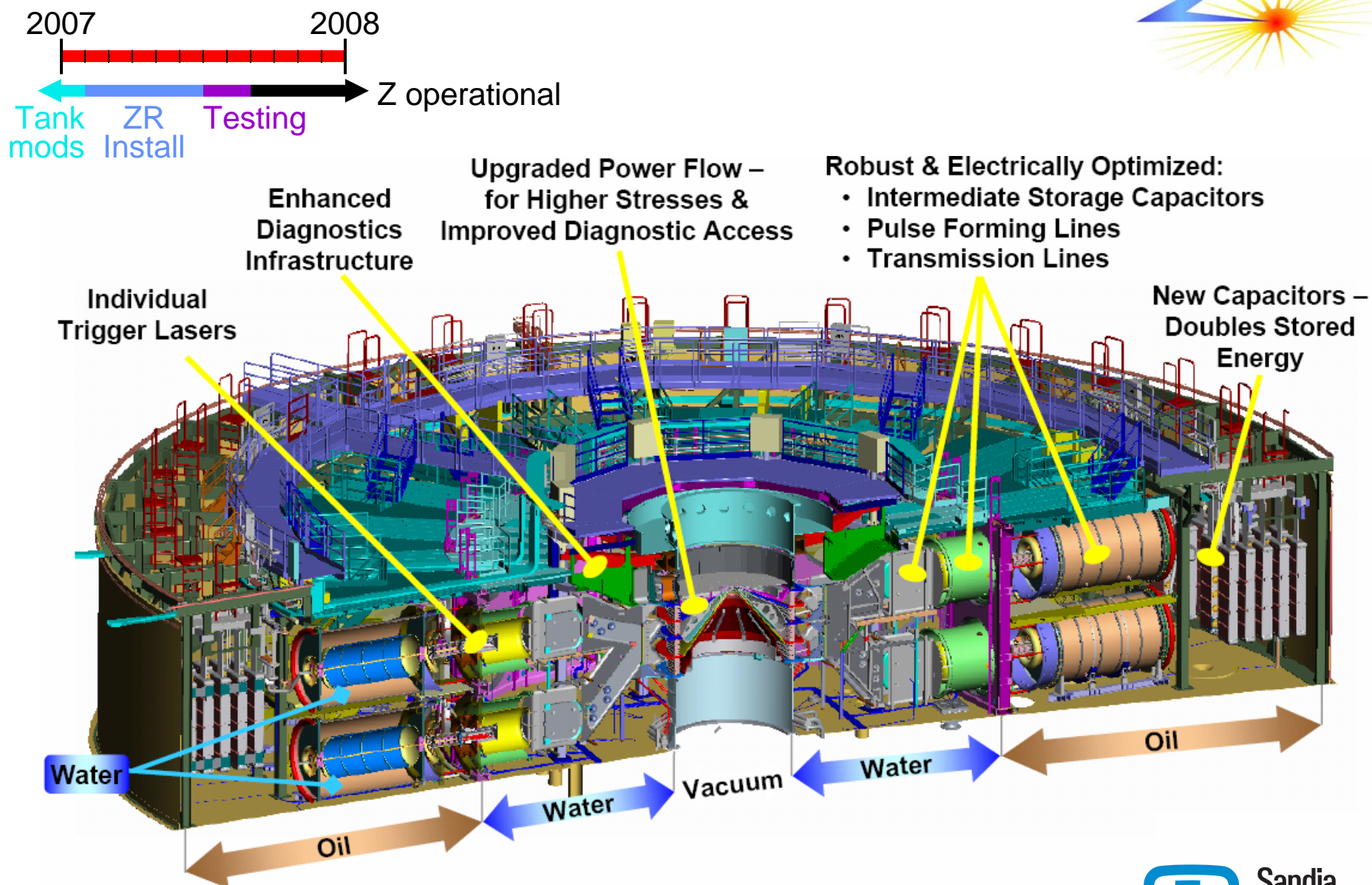
Sandia's Z machine produces intense x-ray radiation



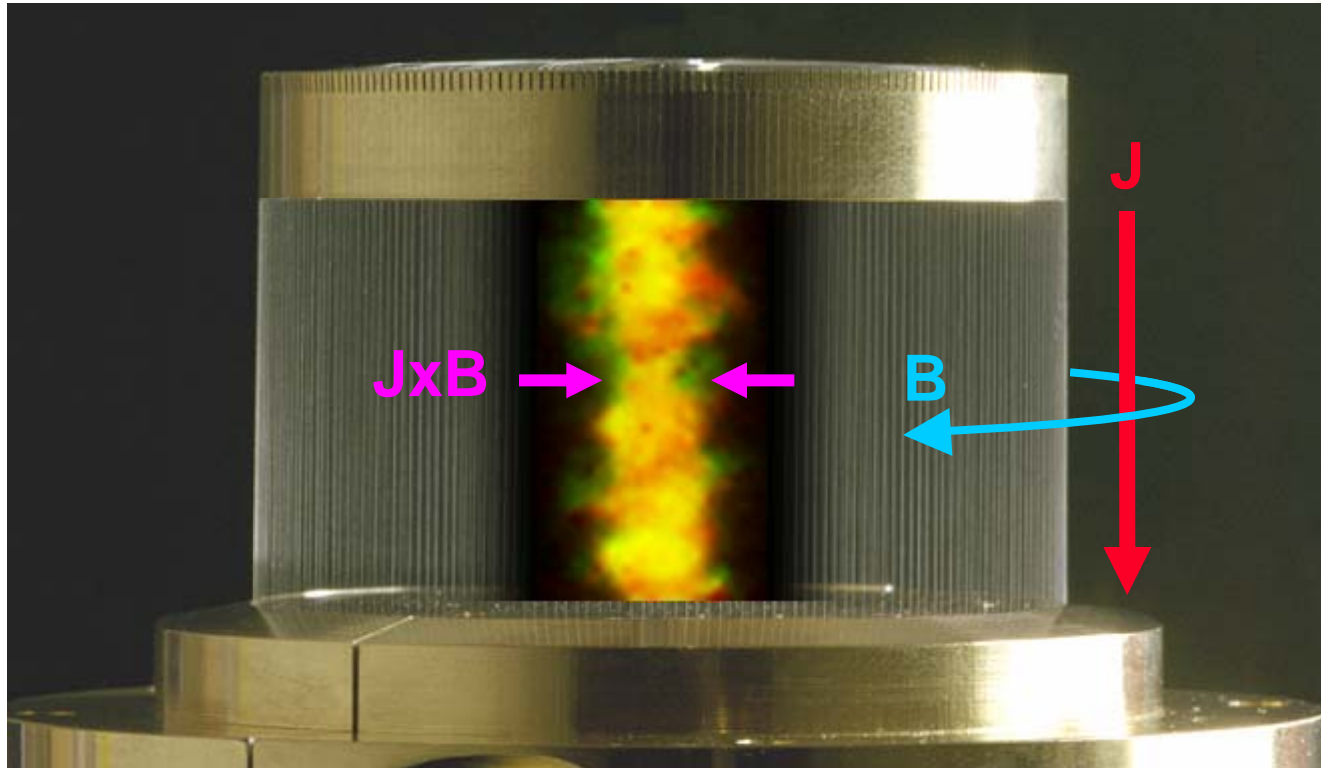
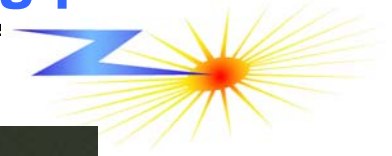
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Z Refurbishment underway—26 MA capability in 2007



J x B pinches wire array into a dense, radiating plasma



Radiation Effects Sciences (C7) z-pinch sources

$n_i \sim 10^{20} \text{ cm}^{-3}$, $T_e \sim 1 \text{ keV}$ (non-LTE)

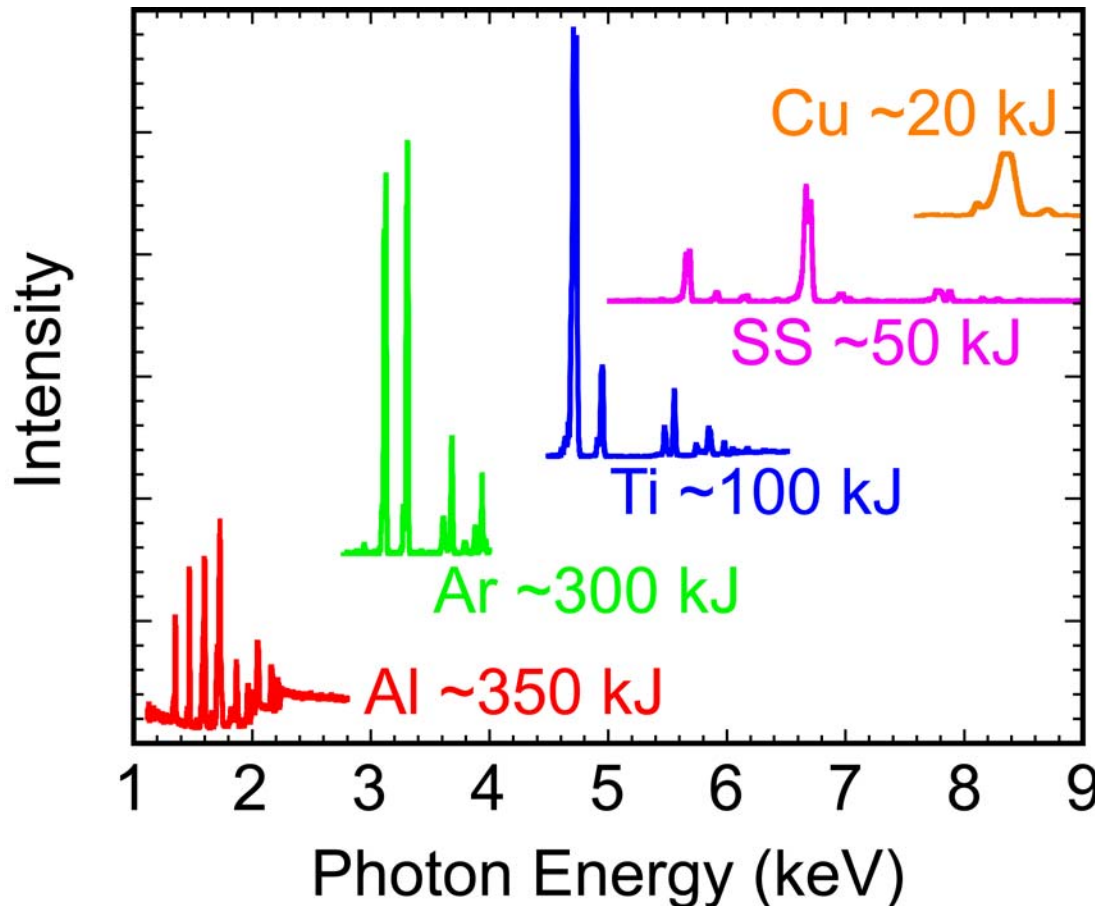
$P_{\text{rad}} \sim 100 \text{ TW}$, $Y_{\text{rad}} \sim 1 \text{ MJ}$, $Y_{\text{rad}, > 1 \text{ keV}} \sim 100 \text{ kJ}$



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A variety of K-shell sources have been studied at Z for C7



- Factor of 2-4 increases in yield expected on ZR
- Al: 40 on 20 mm dia. nested wire array
- Ar: L3 1234 nozzle gas puff Z pinch
- Ti: 50 on 25 mm dia. nested wire array
- SS: 55 on 27.5 mm dia. nested wire array
- Cu: 60 on 30 mm dia. nested wire array
- Pulse widths 5 to 25 ns, rise times 2 to 8 ns

C. Deeney *et al.*, Phys. Plasmas **6**, 2081 (1999).

H. Sze *et al.*, Phys. Plasmas **8**, 3135 (2001).

B. Jones *et al.*, J. Quant. Spec. **99**, 341 (2006).

J.W. Thornhill *et al.*, IEEE T. Plasma Sci. **34**, 2377 (2006).



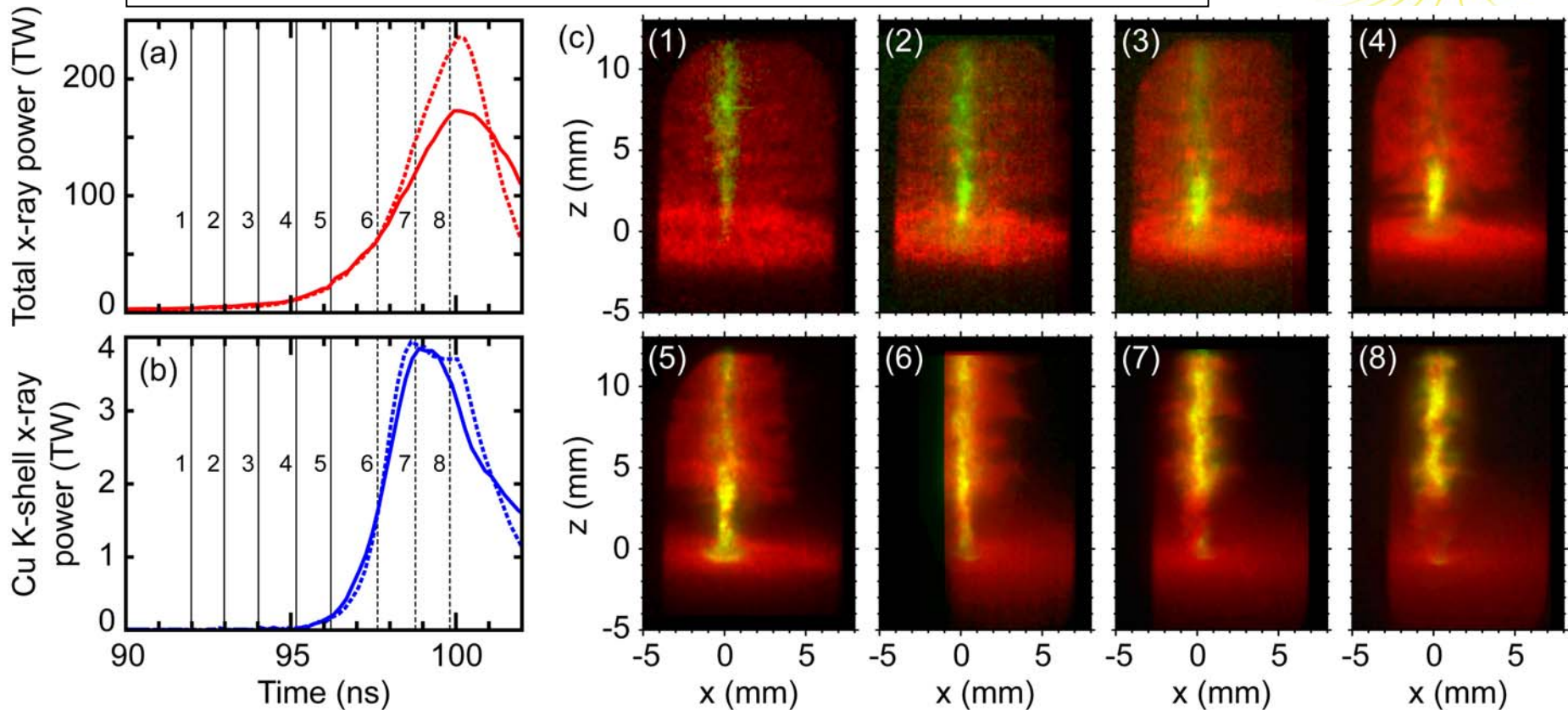
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Plasma heating and radiation excitation on ~1 ns time scale



Cu z pinch on Z (C.A. Coverdale); B. Jones *et al.*, APiP 2007 proc.



- Magnetic RT implosion instability broadens the radial profile
- Plasma accretes on axis, depositing kinetic energy
- e-i-rad thermalization in dense plasma on axis

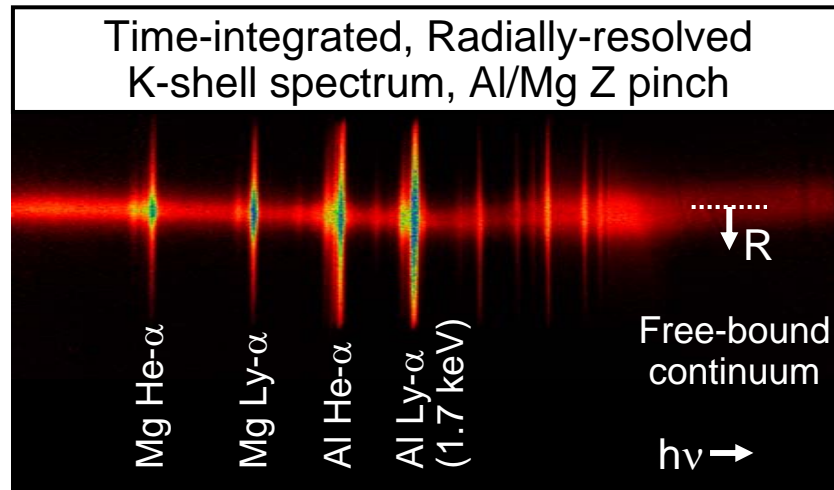
- K-shell emission from column where pinch is assembling
- Details of shock heating are a topic of continuing interest



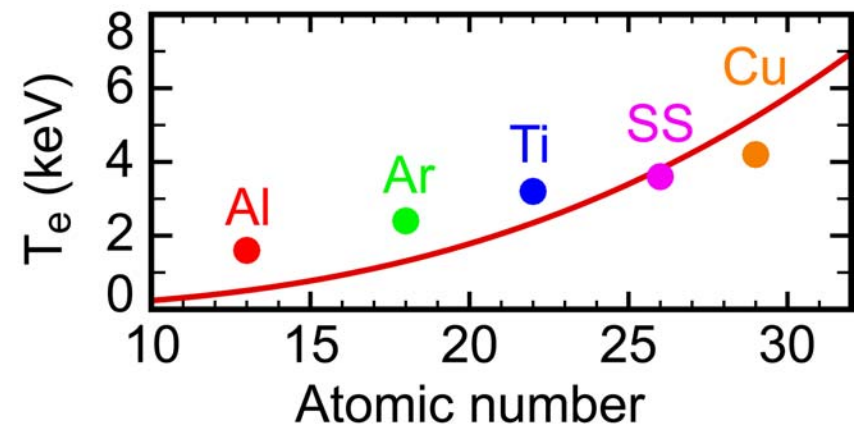
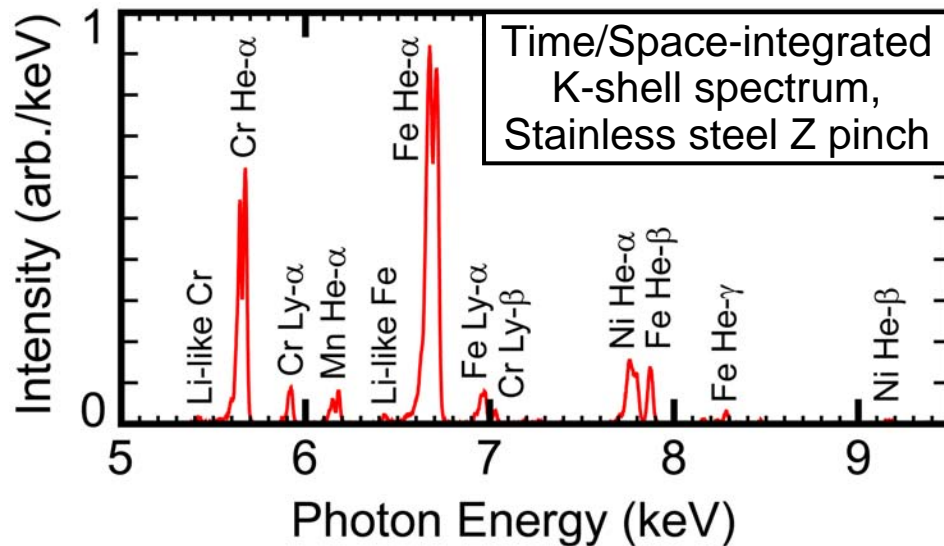
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Load design provides $T_e > 1$ keV for K-shell excitation



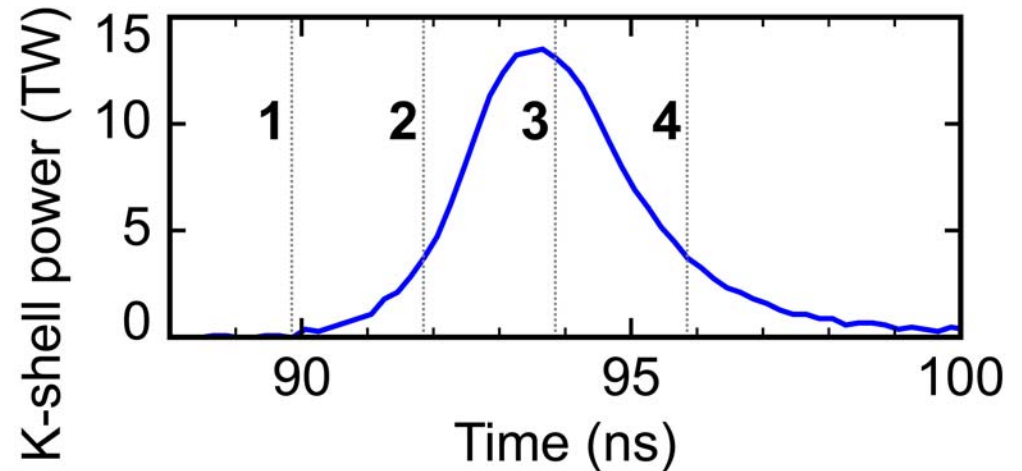
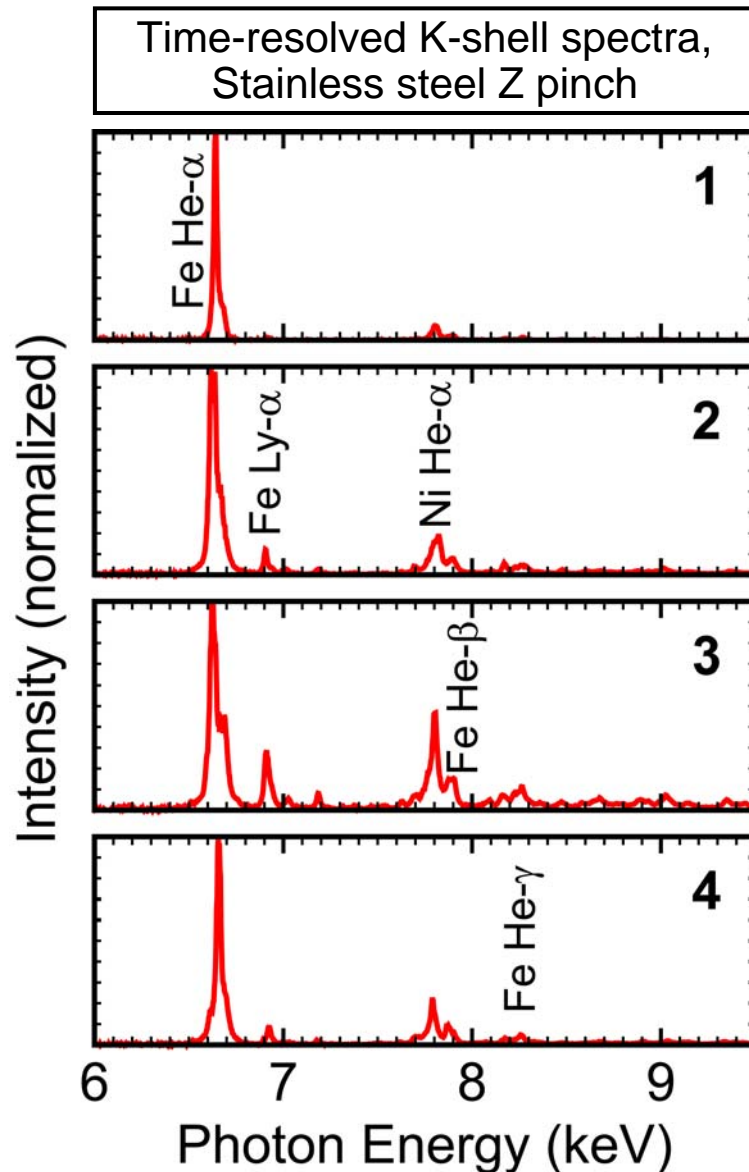
- Time-integrated crystal spectrometers fielded on all Z shots for K-shell survey
- Plasma conditions inferred through non-LTE collisional radiative modeling (NRL)
- Higher photon energy sources require higher T_e to ionize to the K-shell



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Time-resolved x-ray spectra support thermalization studies



- Ni and high order Fe lines light up as n_e , T_e increase near peak K-shell x-ray power
- Time-resolved elliptical crystal spectrometers provide <1 ns resolution with low background

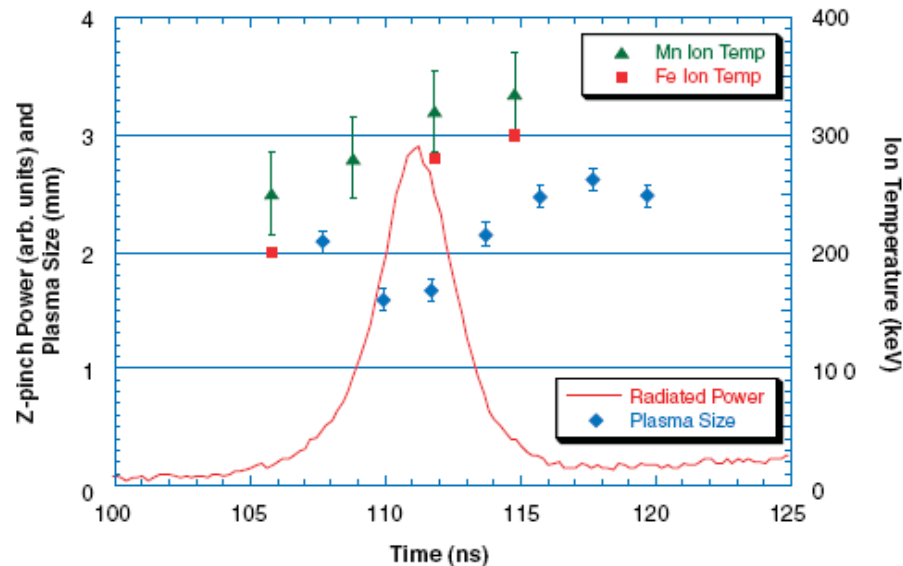
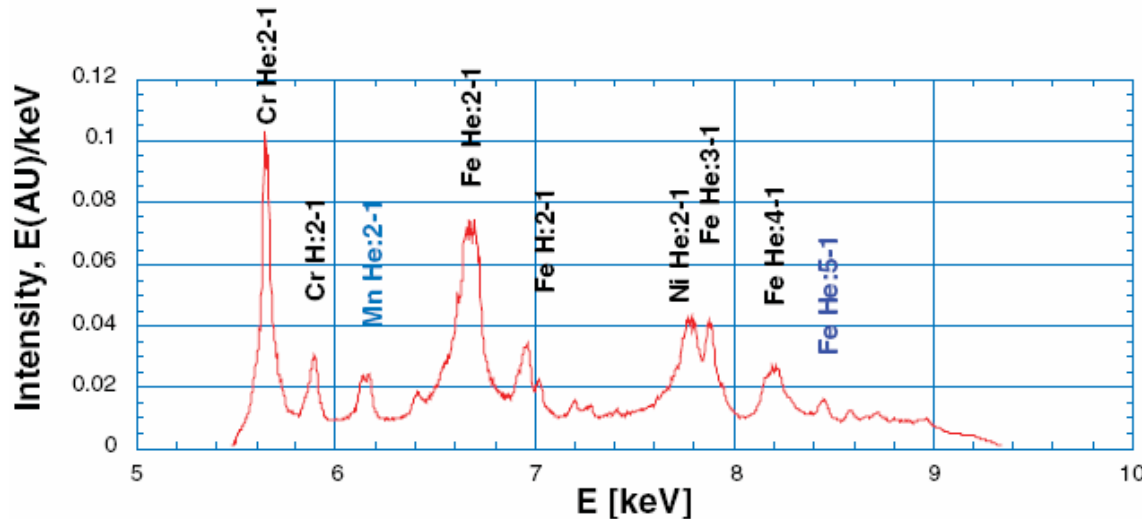
— J. E. Bailey *et al.*, PRL
92, 085002 (2004).



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High ion temperatures are expected in z-pinch plasmas



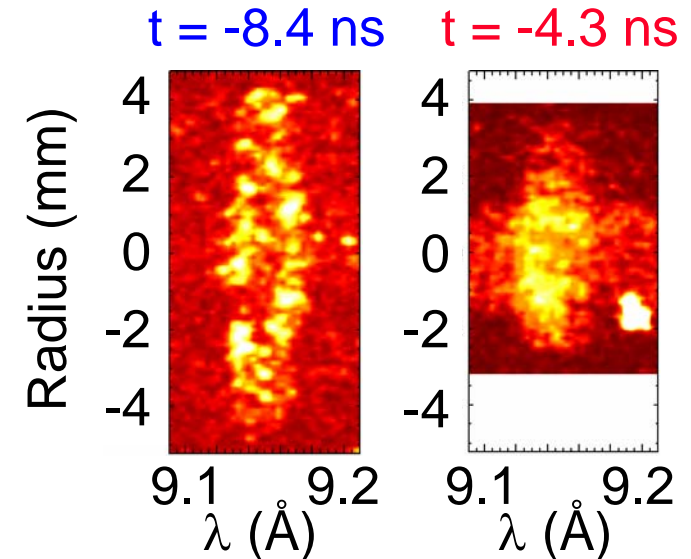
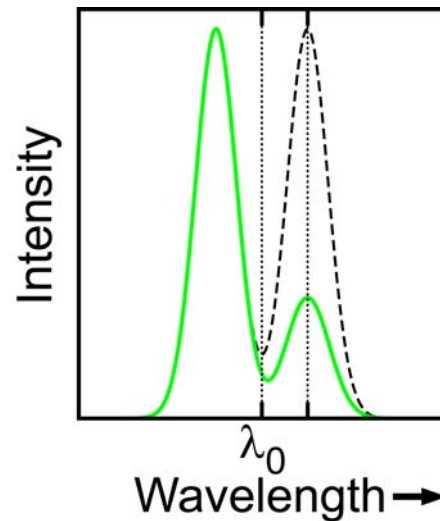
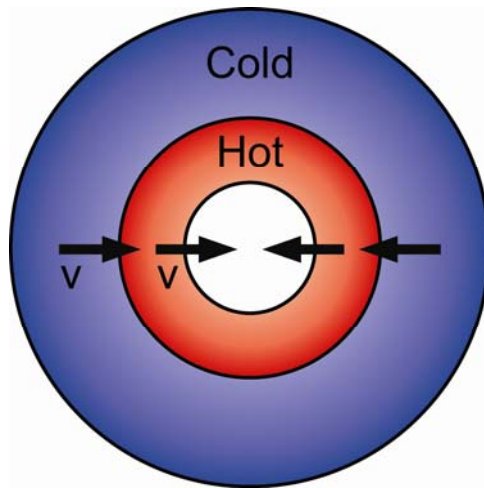
- Ions carry KE; other heating proposed
- We have started to address ion heating in K-shell radiators (e.g. stainless steel)

- M.G. Haines *et al.*, PRL **96**, 075003 (2006): $T_i > 100$ keV inferred from Doppler-broadened line widths on Z
- Results are controversial—further study is desired
- E. Kroupp *et al.*, PRL **98**, 115001 (2007) also discusses > 1 keV ions in a small gas puff z pinch

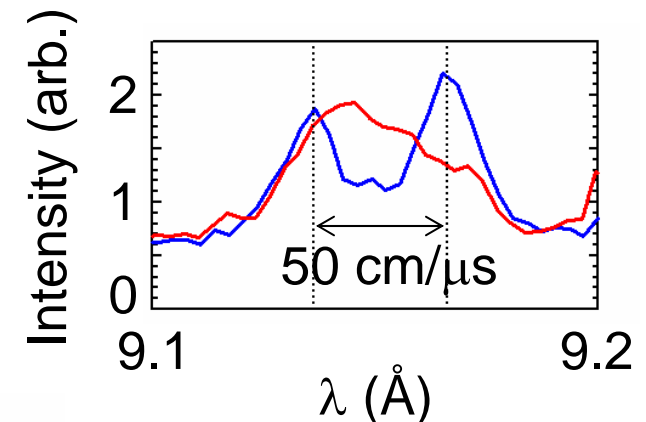
Optically thin Doppler splitting seen in low-mass Al wire array



Z1520 (Coverdale), Mg He- α (Dunham)



- Oval shape—Doppler split seen on axis, but not in tangential view of shell
- At early time, red/blue-shifted lines are similar magnitude \Rightarrow optically thin
- Speckle could be azimuthal structure
- At later times, red-shifted line is attenuated by shell/trailing mass opacity
- Splitting not so obvious in Al lines—brighter precursor emission on axis?



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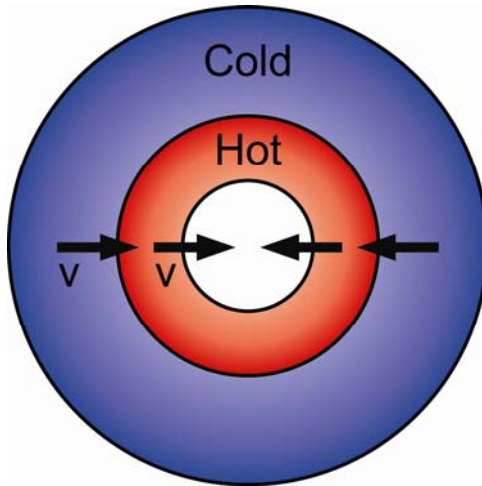
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Doppler-shifted absorption seen in high-mass Al wire array

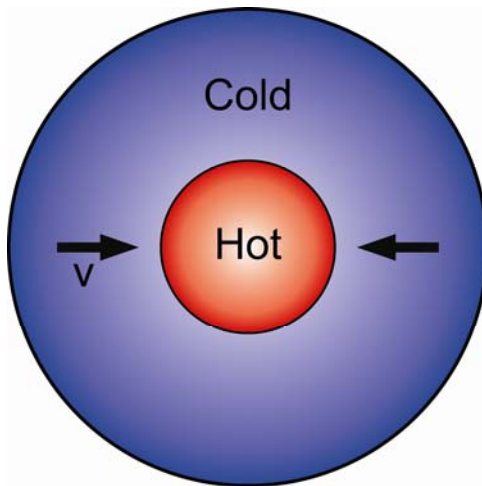


Z1518, Al Ly- α
 $t = -9.0$ ns

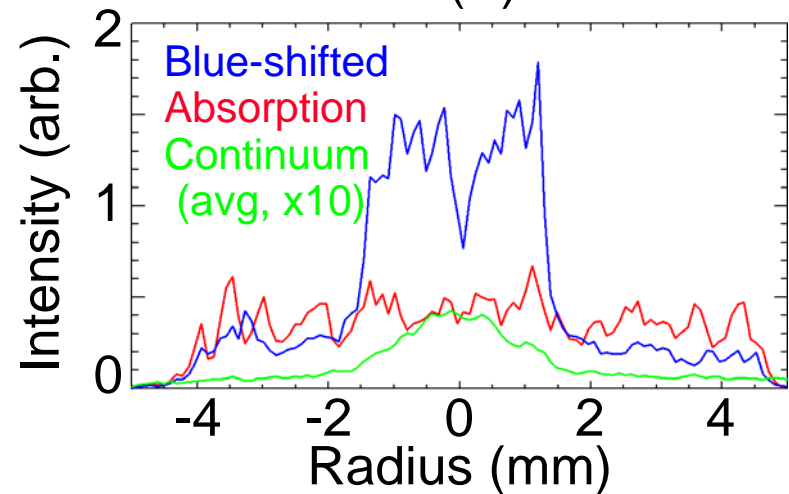
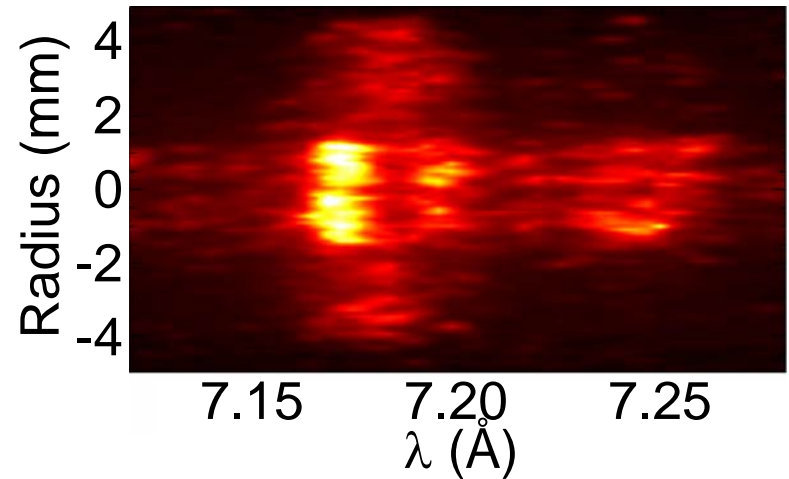
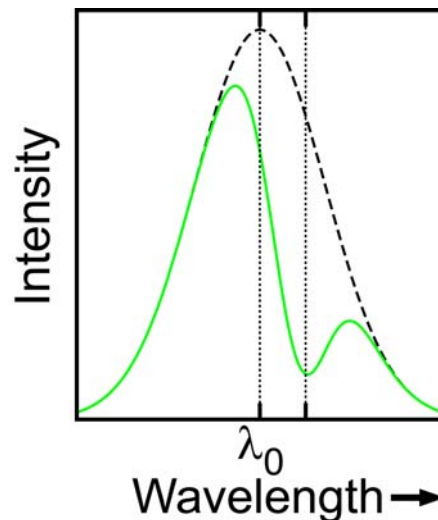
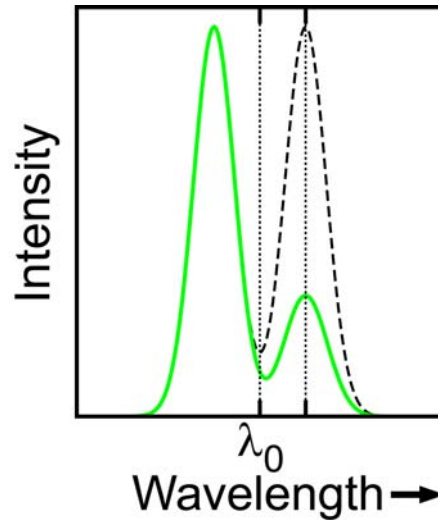
Can interpret as...



or as...



Opacity in cold trailing mass

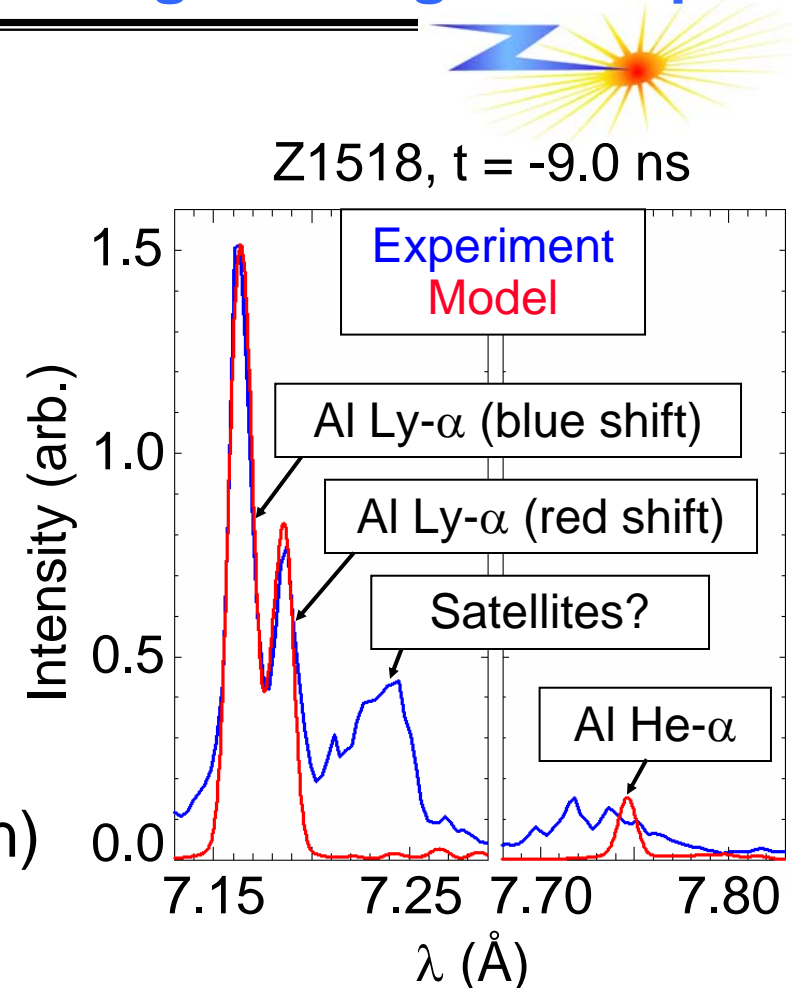
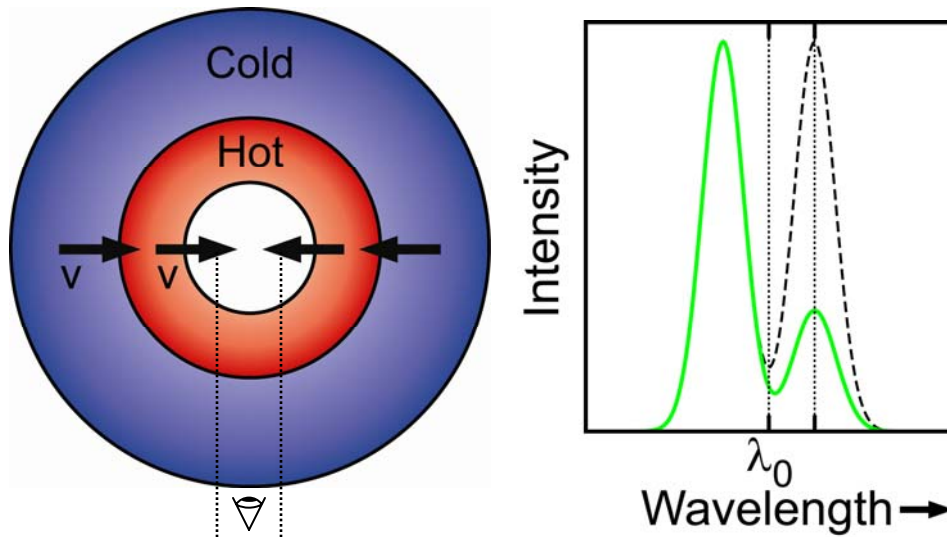


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CR model with Doppler shifts/broadening is being developed

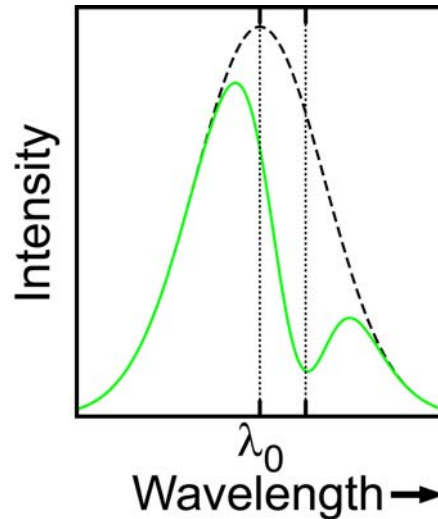
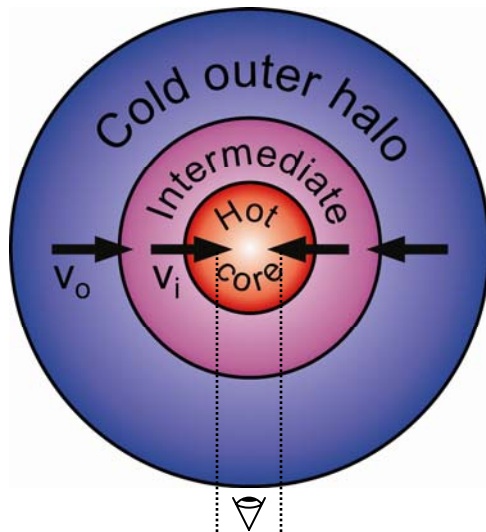
Opacity in cold trailing mass attenuates red-shifted Al Ly- α and He- α



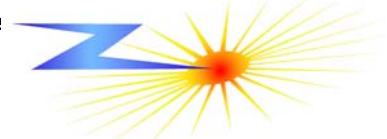
- Collisional-radiative model, radiation transport in discrete zones (Y. Maron)
 - Hot: $1.5 \text{ mm} < R < 2 \text{ mm}$
 - Cold: $2 \text{ mm} < R < 9 \text{ mm}$
- Line shape calculations
 - Stark broadening (not dominant)
 - Doppler broadening/splitting (implemented for first time, $\delta v/v = 10\%$)
- Preliminary results (need to consider satellites):
 - Hot: $n_i = 5 \times 10^{19} \text{ cm}^{-3}$, $T_e = 700 \text{ eV}$, $v = 40 \text{ cm}/\mu\text{s}$
 - Cold: $n_i = 5 \times 10^{19} \text{ cm}^{-3}$, $T_e = 150 \text{ eV}$, $v = 30 \text{ cm}/\mu\text{s}$

Matching spectral features constrains plasma parameters

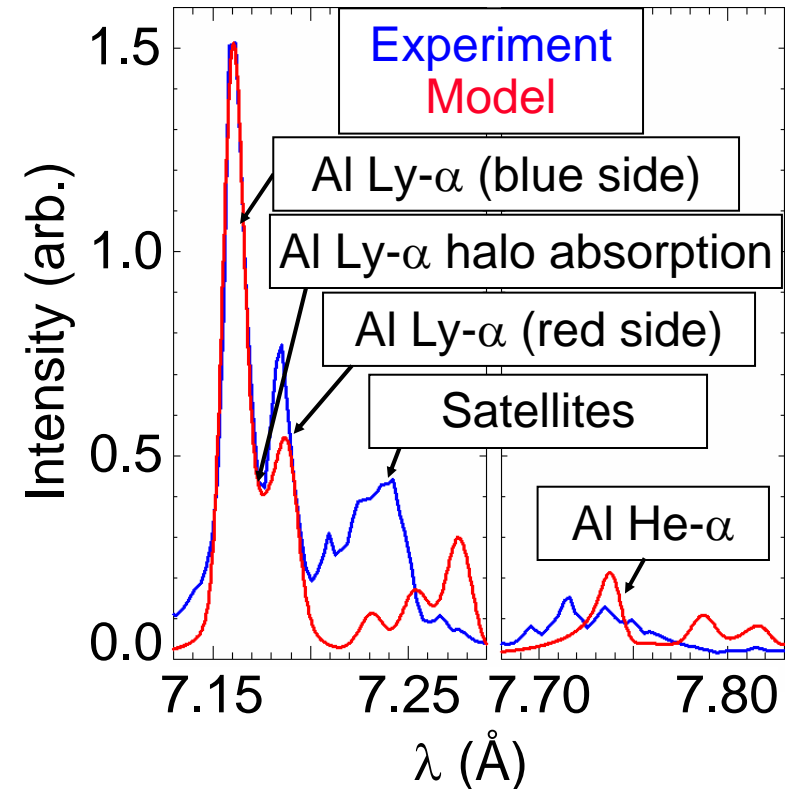
Opacity in cold trailing mass forms
absorption dip in broad Al Ly- α line



	Core	Intermediate	Outer
T_e (eV)	400	250	100
R_{out} (mm)	1.5	3.5	10
n_i (cm $^{-3}$)	4e19	4e19	4e19
v_r (cm/ μ s)	0	20	10
δv (cm/ μ s)	30	0	0



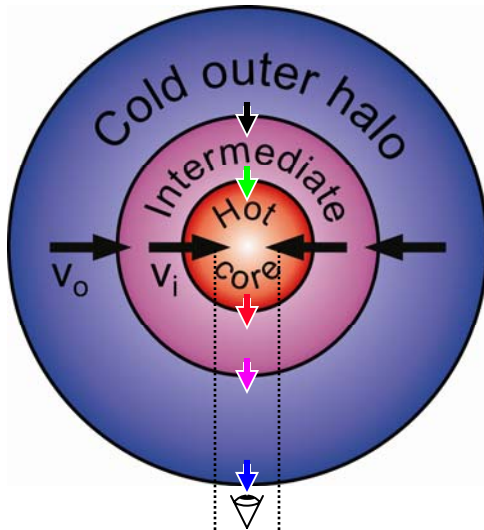
Z1518, $t = -9.0$ ns



- Al Ly- α satellites can be better explained with dense core
- The δv in core may indicate T_i , turbulence, or ∇v
- Study required to determine how well plasma parameters can be constrained

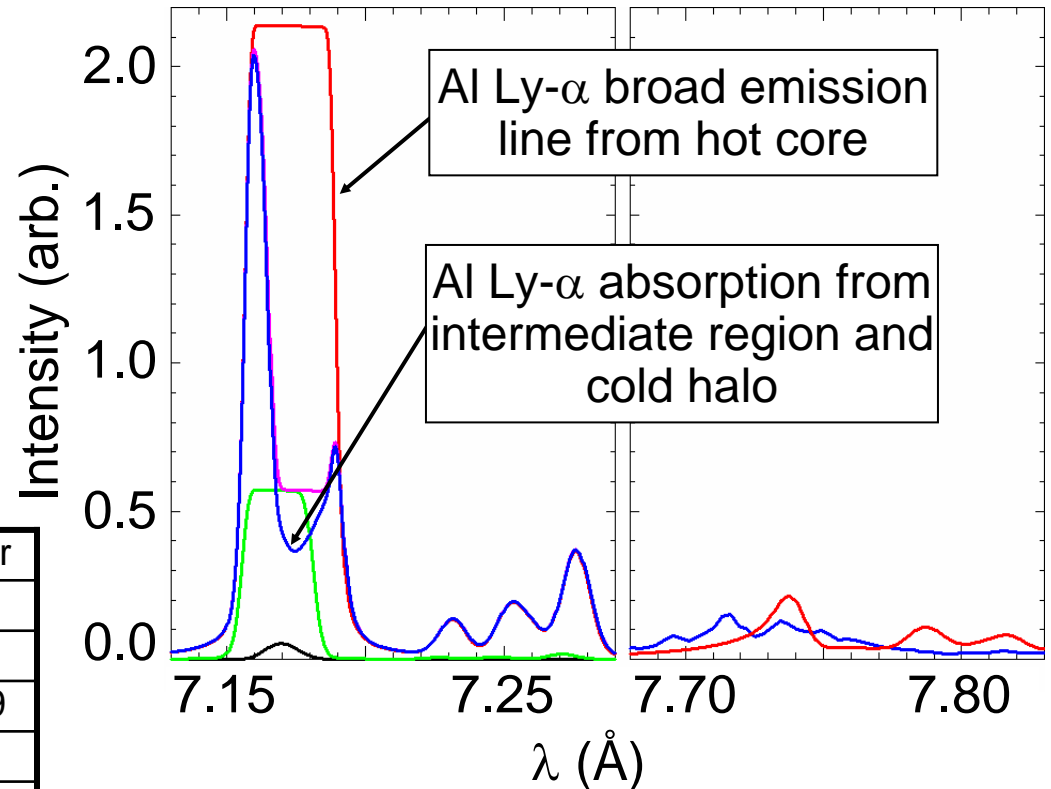
Radiation transport is tracked through each plasma region

Opacity in cold trailing mass forms
absorption dip in broad Al Ly- α line



	Core	Intermediate	Outer
T_e (eV)	400	250	100
R_{out} (mm)	1.5	3.5	10
n_i (cm ⁻³)	4e19	4e19	4e19
v_r (cm/ μ s)	0	20	10
δv (cm/ μ s)	30	0	0

Z1518, $t = -9.0$ ns

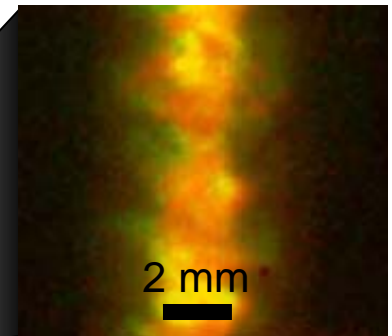
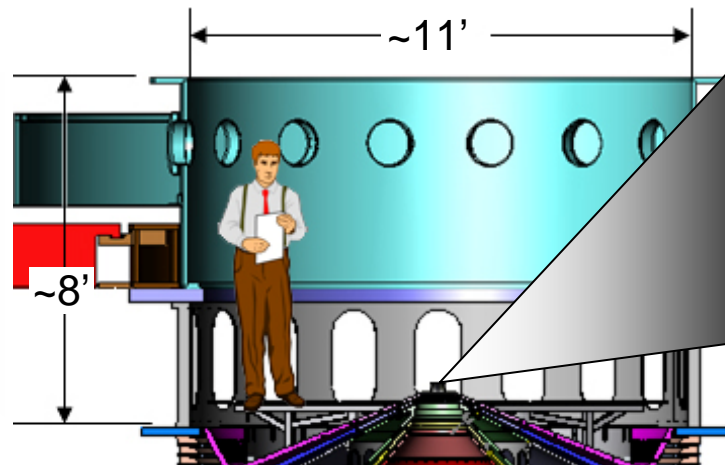
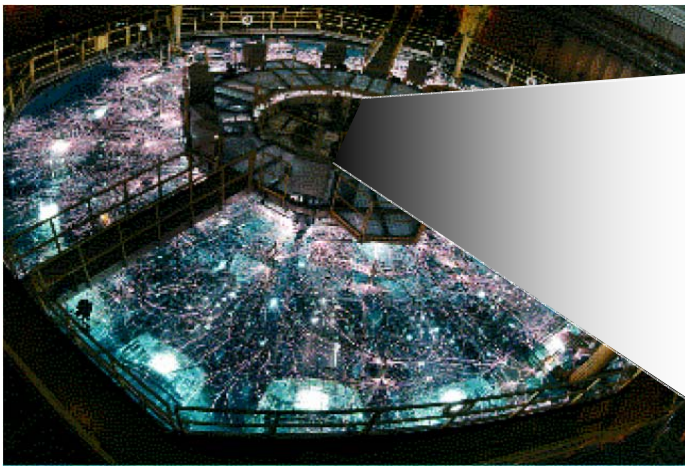


- Broad Al Ly- α line from hot core backlights the cold halo
- May be able to infer T_i in the core through this analysis
- Al Ly- α satellites originate in the hot core and will help constrain core density

Summary



- The Radiation Effects Sciences program at Sandia is developing 1-10 keV K-shell radiation sources (wire array and gas puff z pinches) on the Z machine
- Physics associated with conversion of ion kinetic energy to electron thermal energy, then radiation is important for z pinches of all classes
- Z-pinch physics, plasma heating and energetics will continue to be investigated via K-shell spectroscopy and other x-ray measurements on Z along with numerical and analytical theory



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