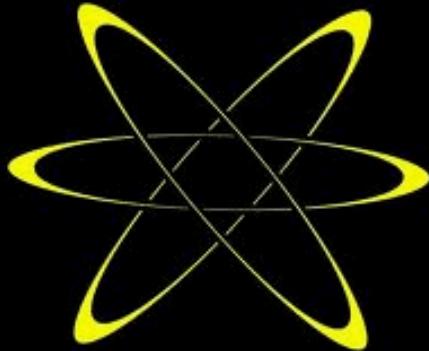


Gamma Spectroscopy

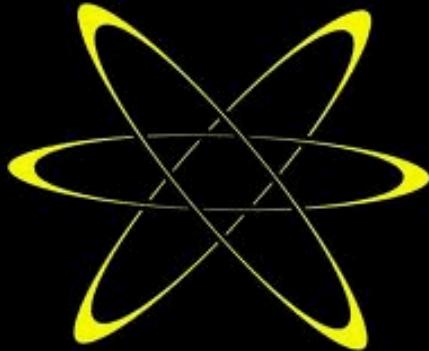
John McCulloch

Radiation Protection and Sample
Diagnostics (4121)



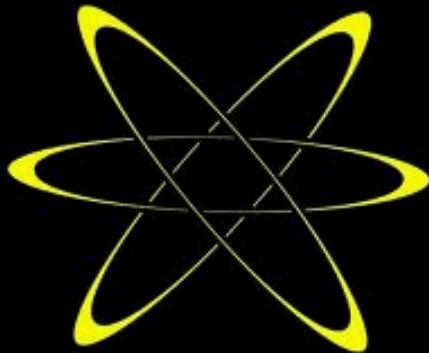
What Do These Have in Common?





First Portable Field Instrument (1945)

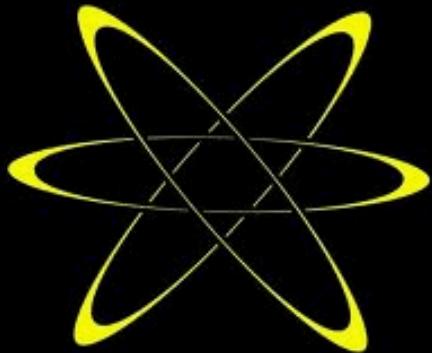




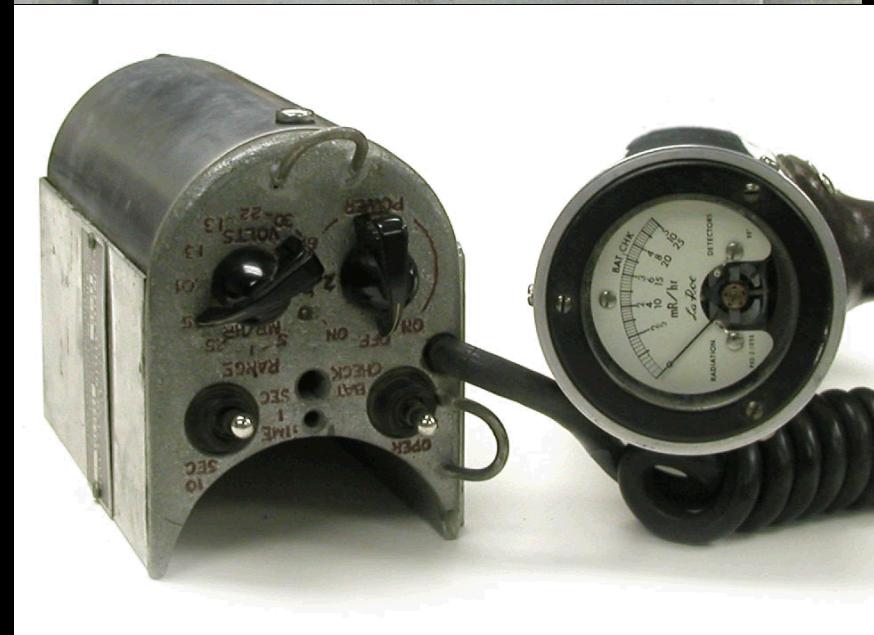
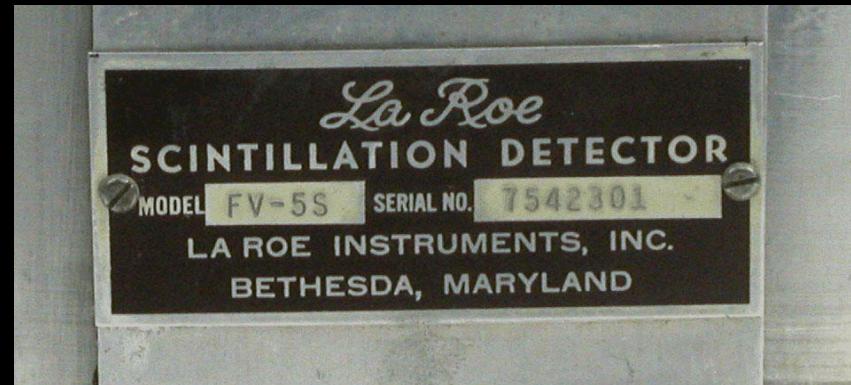
Dawn of Gamma Spec

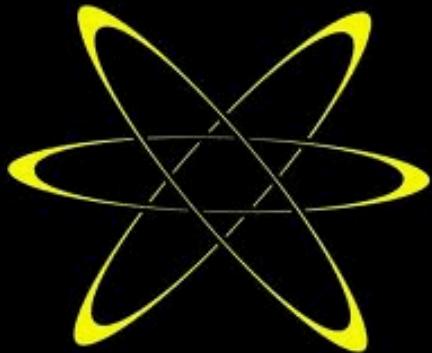
- Gammas “invisible” to all detectors
- 1948 Robert Hofstadter adds a “pinch” of Thallium Iodide to melt of NaI.
- Gamma Spec now possible
- Finds spectra for nuclides
- Identifies escape, annihilation and backscatter peaks.





Early Gamma Analysis

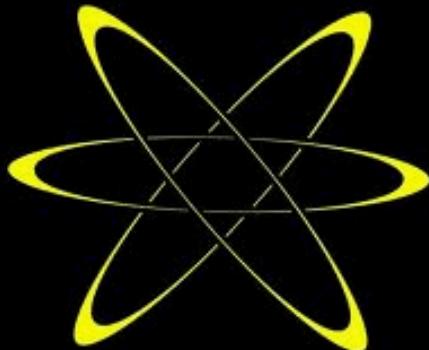




High Purity Germanium

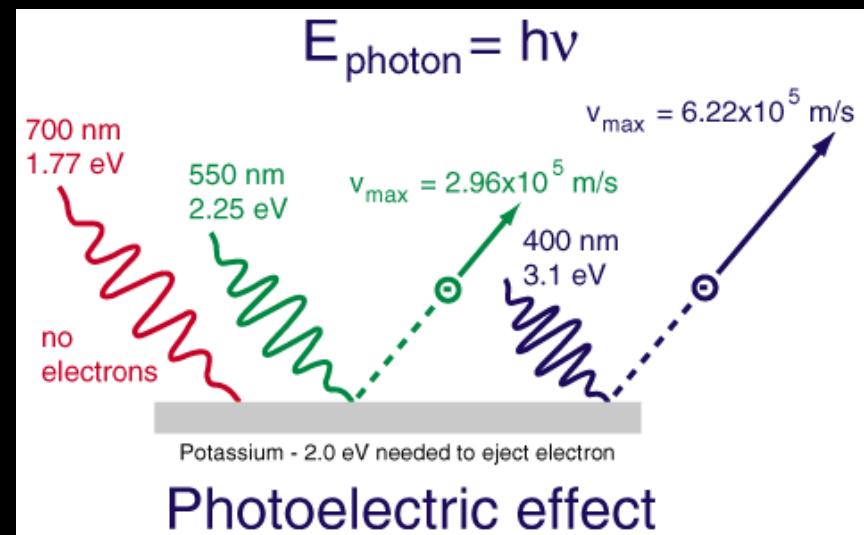
- Allister Tavendale (1963)
- Game Changer
 - Excellent Resolution
- Li Drifted due to impurities
- MUST Keep COLD!

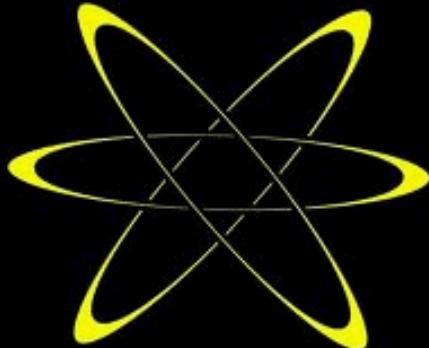




Detector Behavior – Photo Electric Effect

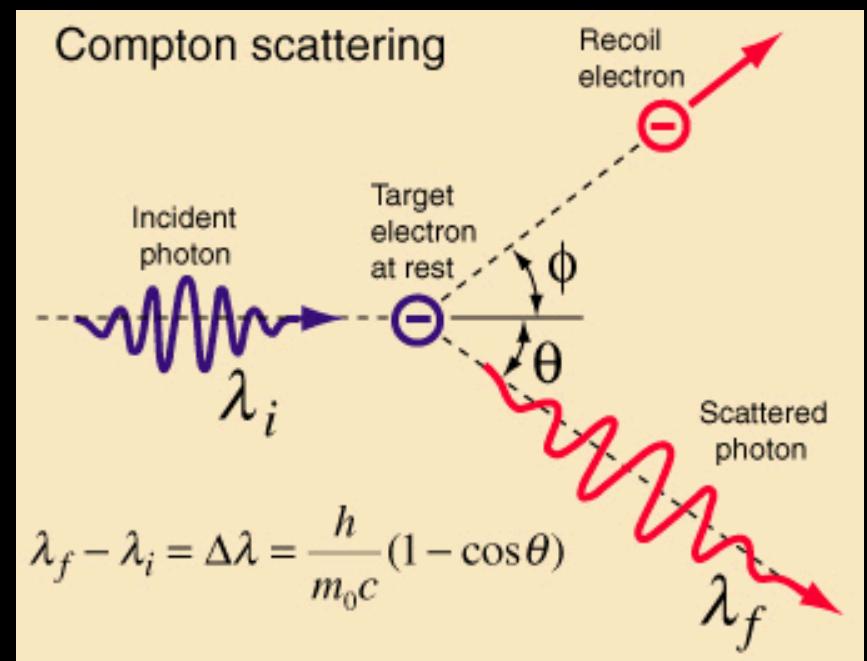
- Heinrich Hertz (1887)
- Einstein (1905)
 - Nobel Prize (1921)
- Most desirable of interactions in HPGe
- Most common for low energy photons

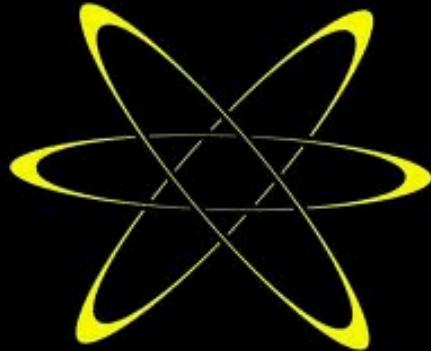




Detector Behavior – Compton Scattering

- Arthur Compton (1923)
 - Nobel Prize (1927)
- Demonstrates Light cannot be strictly a wave
- Occurs with mid-energy level photons
- The most common form of interaction in gamma spec





Compton Scattering - Possibilities

Electron Energy

$$E_e = h\nu - h\nu'$$

Exiting Photon Energy

$$h\nu' = \frac{h\nu}{1 + (h\nu/m_0c^2) * (1 - \cos\theta)}$$

Small Scattering Angle

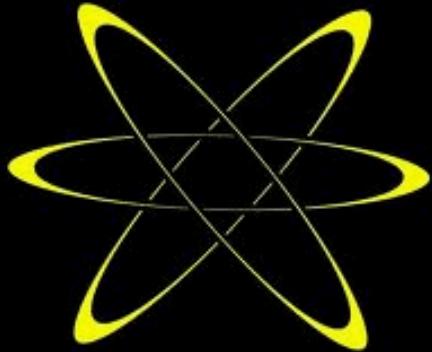
$$\lim_{\theta \rightarrow 0} \frac{h\nu}{1 + (h\nu/m_0c^2) * (1 - \cos\theta)} = h\nu$$

Large Scattering Angle

$$\lim_{\theta \rightarrow \pi} \frac{h\nu}{1 + (h\nu/m_0c^2) * (1 - \cos\theta)} = \frac{h\nu}{1 + (2h\nu/m_0c^2)}$$

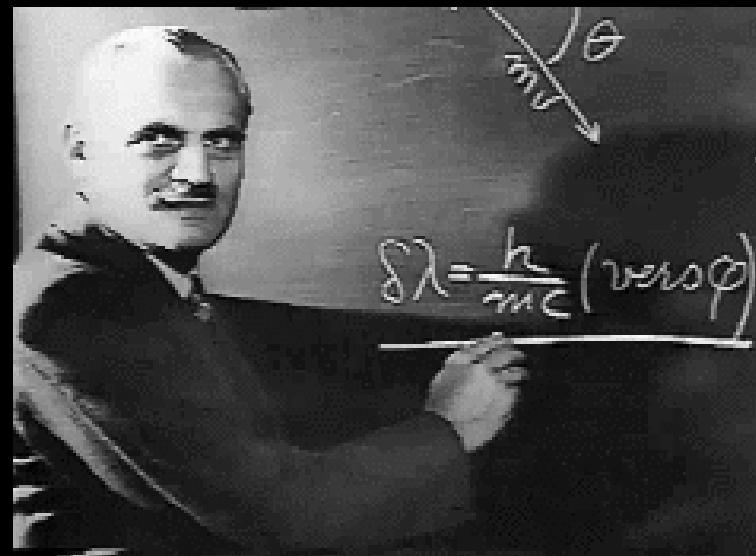
Max Scatter Energy

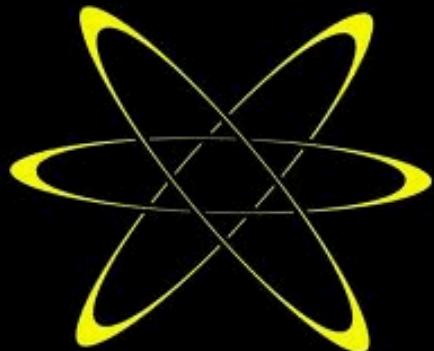
$$h\nu' = \lim_{h\nu \rightarrow \infty} \frac{h\nu}{1 + (2h\nu/m_0c^2)} = \frac{m_0c^2}{2} = h\nu - 256 \text{ keV}$$



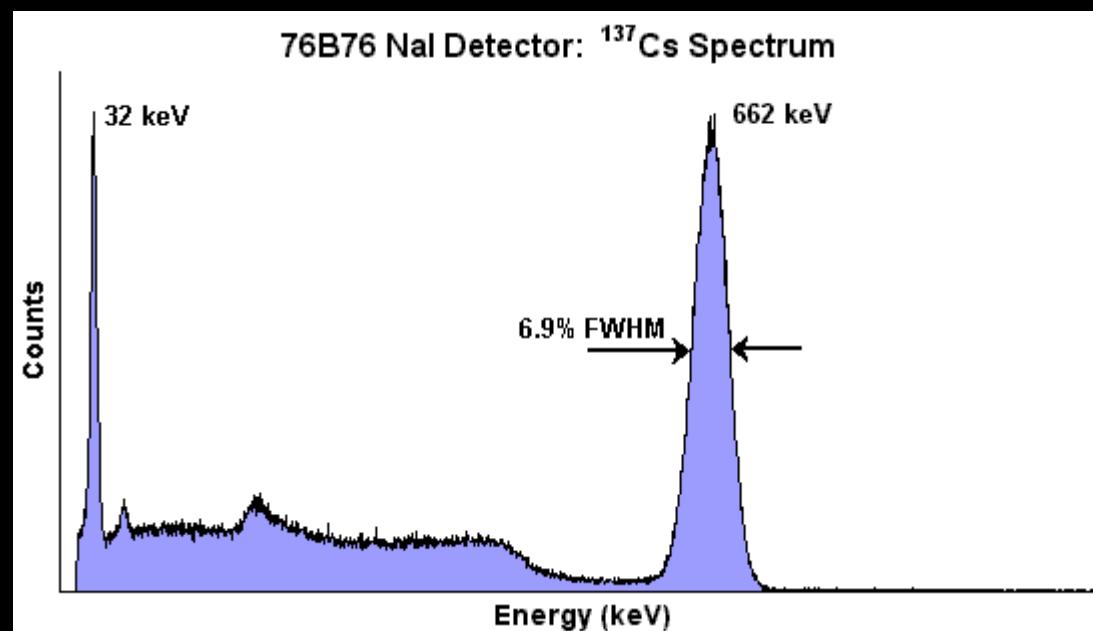
Back Scatter

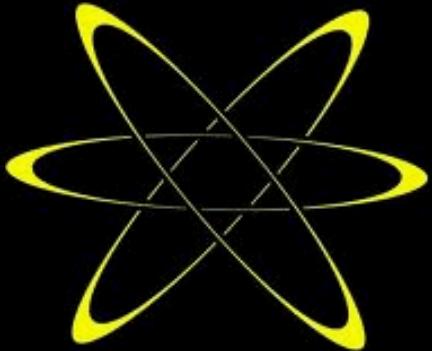
- Special Case of Compton
- Photon outside detector Scatters In
- Always near 256 keV
- Peak leans left





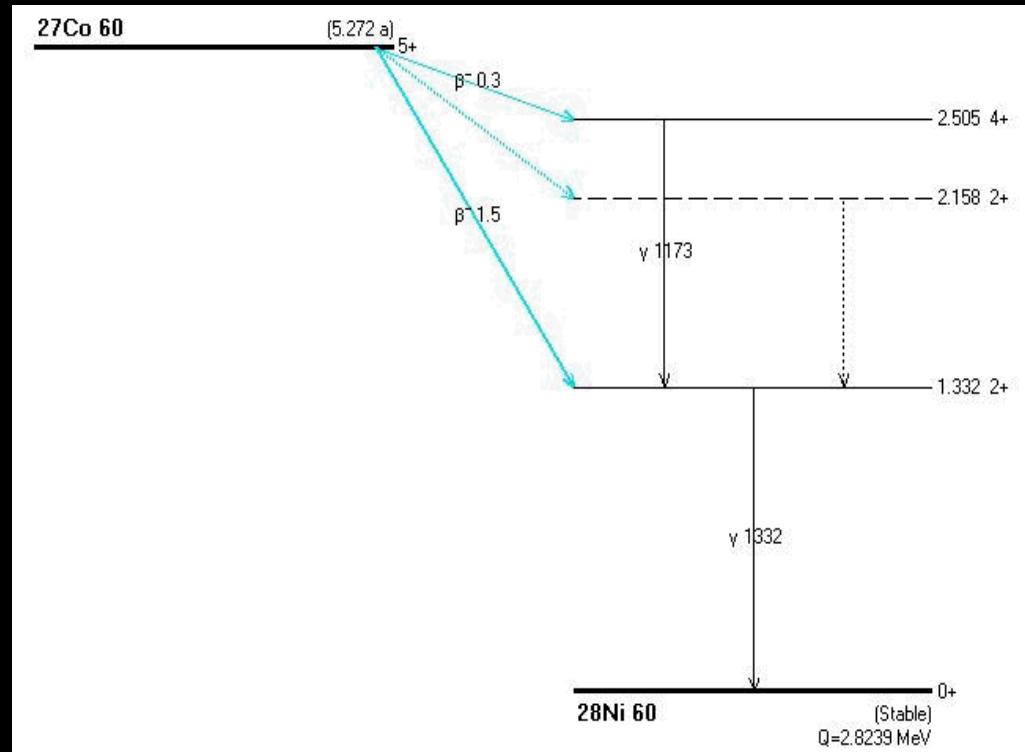
Cs-137





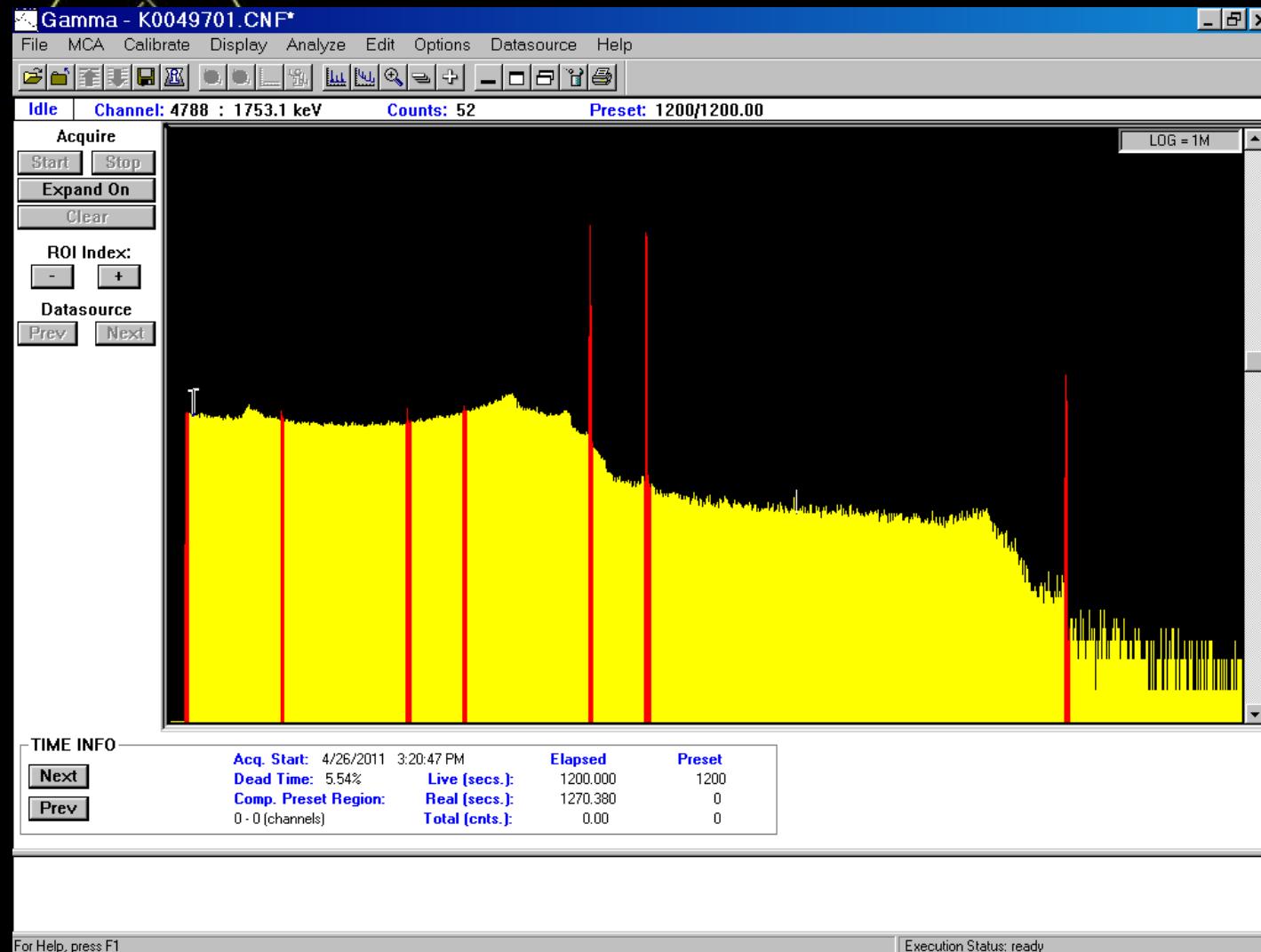
- Cascade Summing
 - Rapid Decay
 - Common
 - In certain isotopes
- Random Summing
 - High Count Rate
 - Fairly Rare
 - 511 peak
- Summing In/Out
- Creates “false” peaks

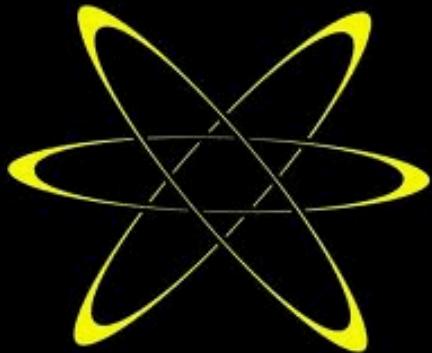
Summing





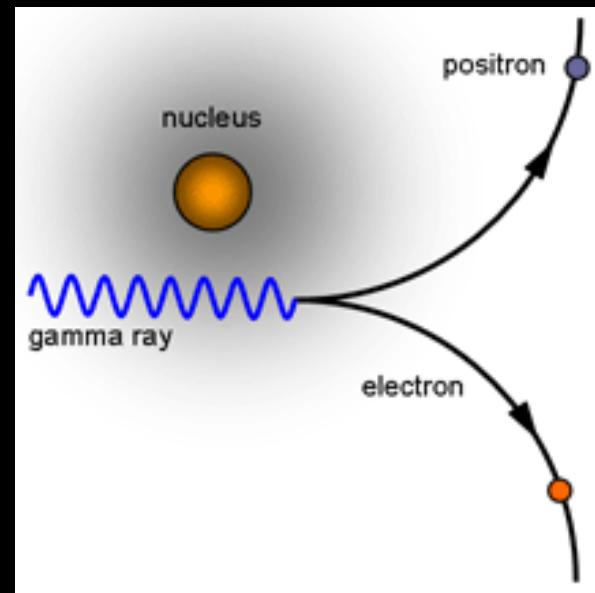
Co-60

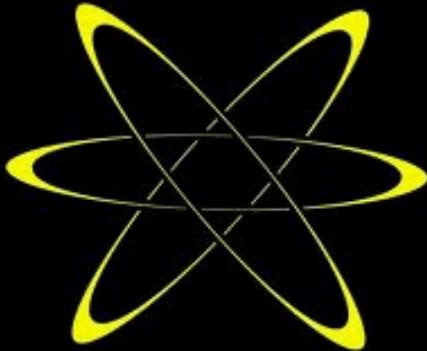




Pair Production

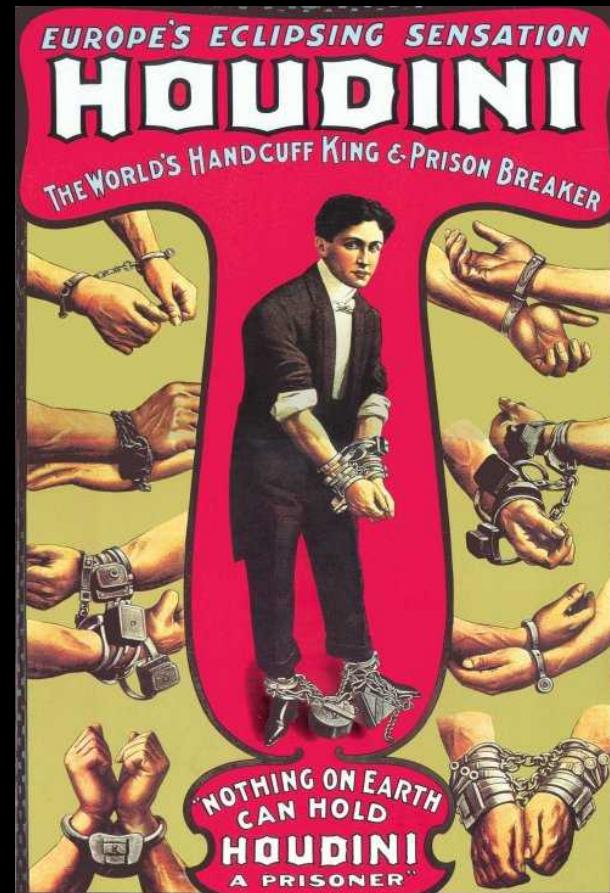
- Coolest form of interaction
- Discovered Patrick Blackett 1933
 - Student of Rutherford
 - Nobel Prize 1947
- Minimum Energy of 1022 keV
- Full energy deposited in 511 keV peak

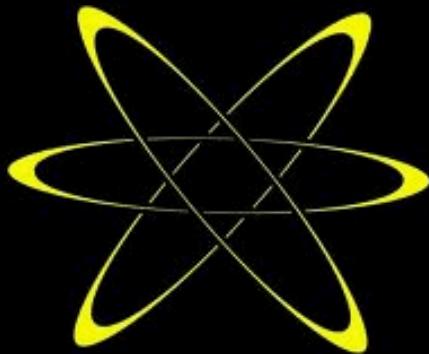




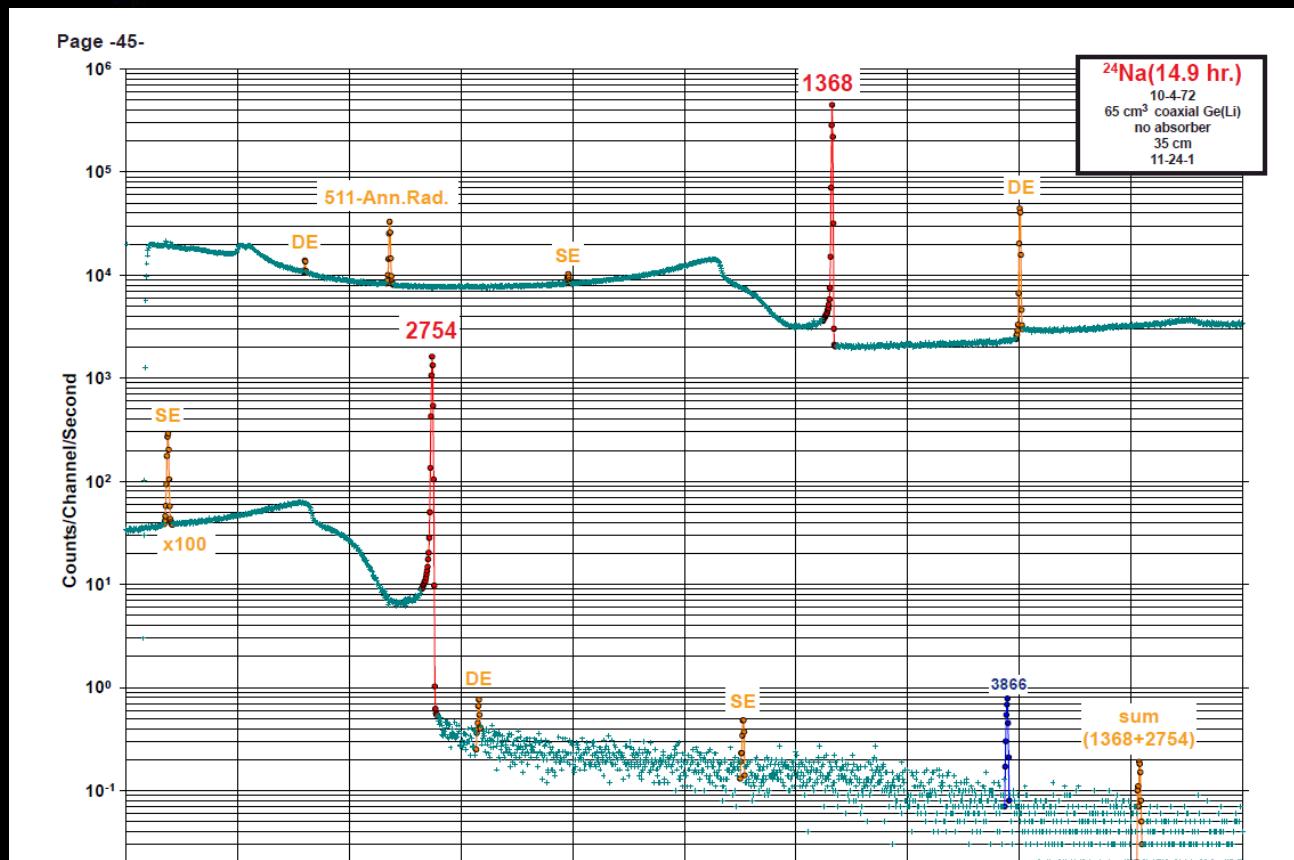
Escape Peaks

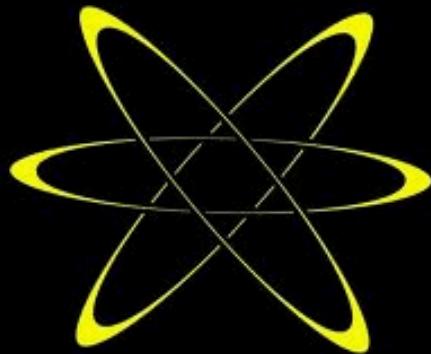
- Only for high energy photons
- Electron/Positron annihilate
- One or both 511 keV photons escape
- Peak shifted down by 511 or 1022 keV.



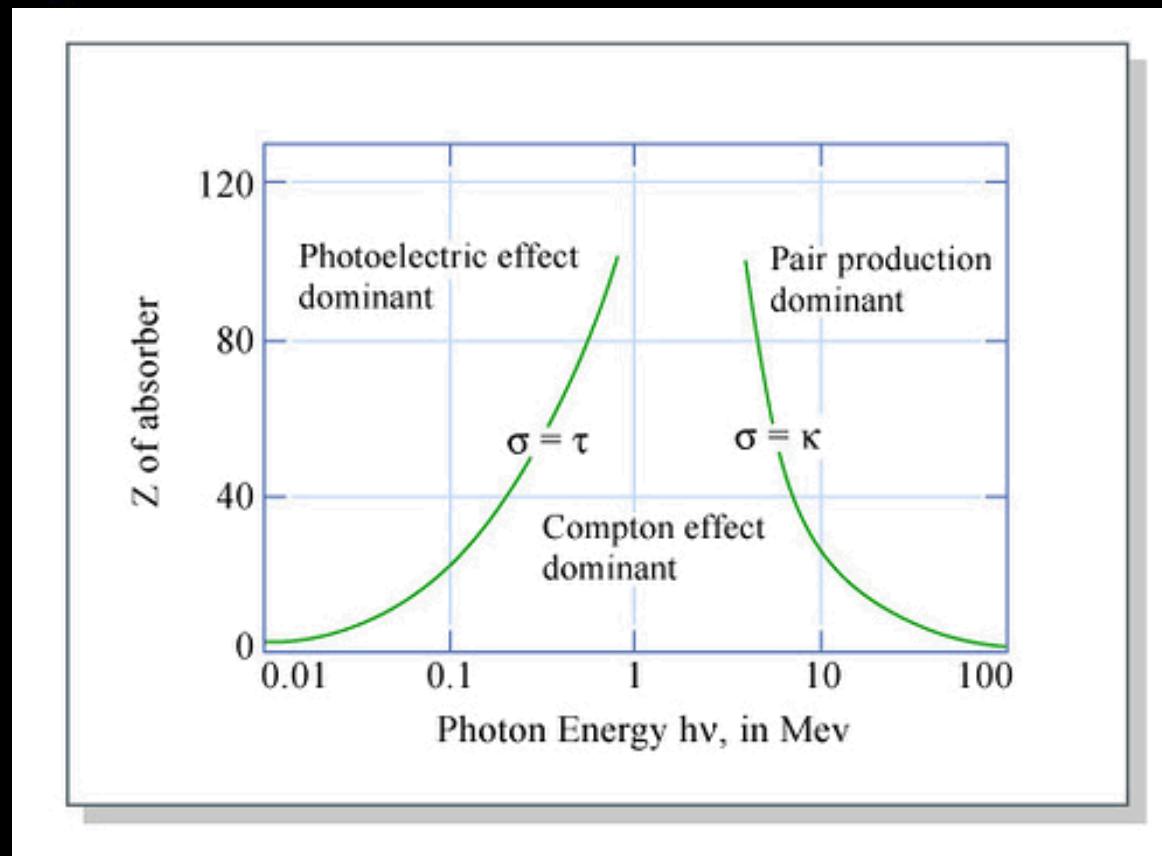


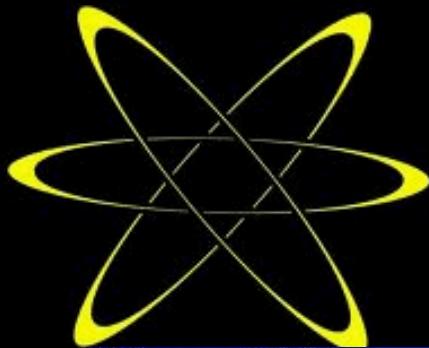
Na-24



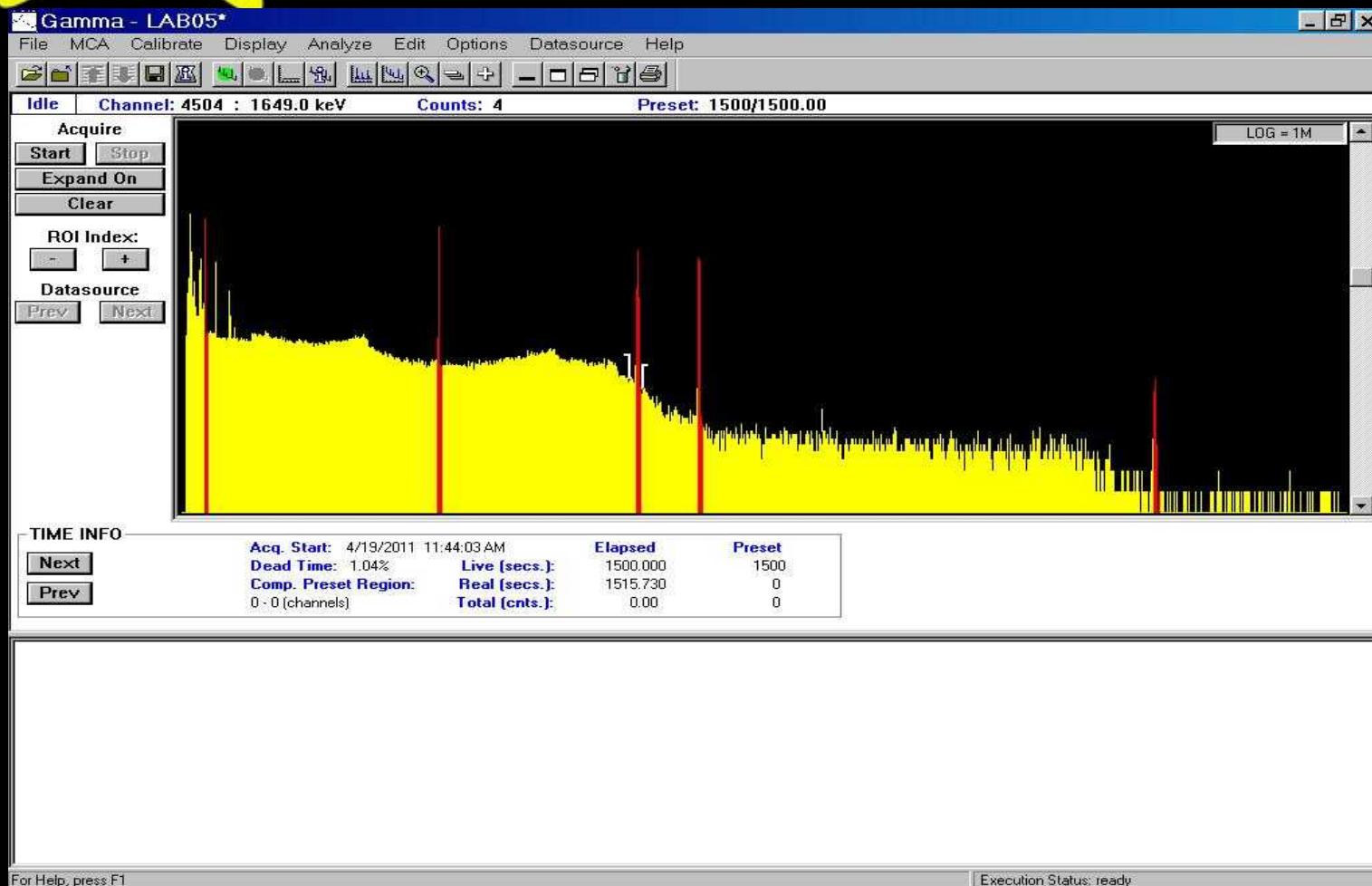


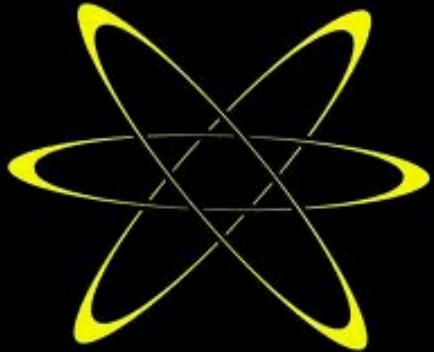
What Dominates?





Genie 2K Spectrum



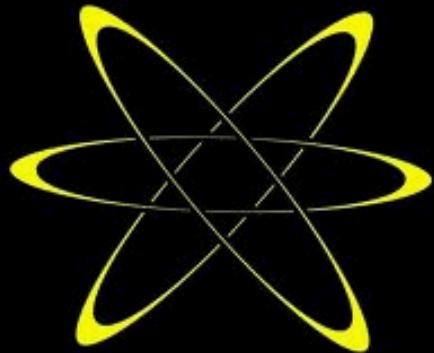


Application Time



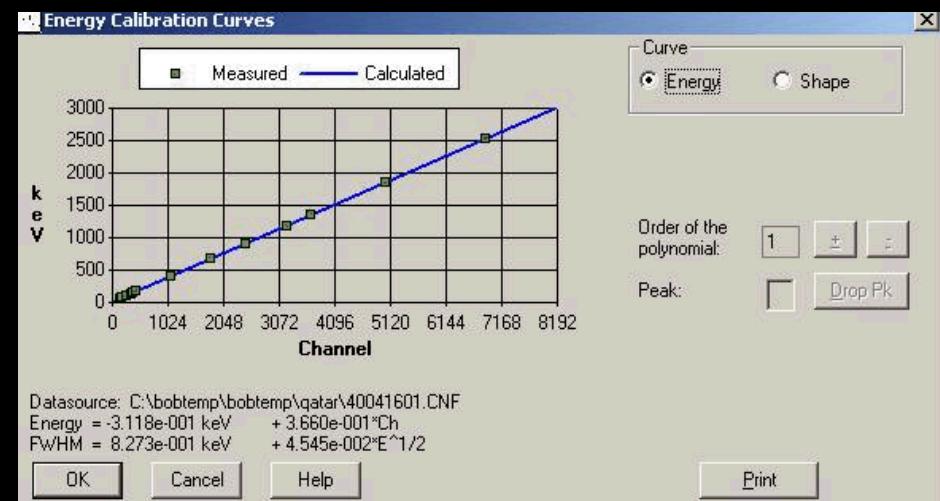
* = Primary line						
NID Confidence Index Threshold : 0.40						
Energy tolerance : 1.000 FWHM						
Nuclide	Id	Energy	Eff.	Yield	Activity	Activity
Name	Confid.	(keV)	(Absolute)	(%)	(pCi/Each)	Error
I-131	1.000	80.18	9.23E-002	2.62	2.87E+005	4.04E+004
		284.30	9.16E-002	6.05	3.16E+005	4.17E+004
		364.48*	7.25E-002	81.20	3.37E+005	4.40E+004
		636.97	4.31E-002	7.26	3.49E+005	4.56E+004
		722.89	3.85E-002	1.80	3.66E+005	4.84E+004
BI-212	0.517	727.17*	3.83E-002	6.65	2.79E+003	4.13E+002
		785.46	3.58E-002	1.11		
		1078.62	2.72E-002	0.54		
		1620.62	1.89E-002	1.51		

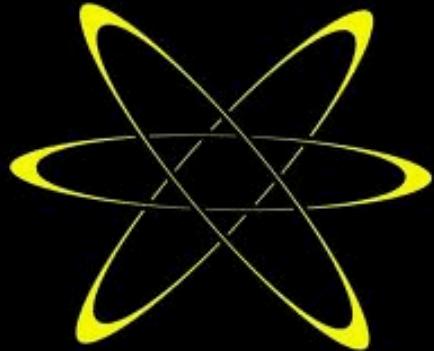
Activity Report / Sample ID = D0072301 Page 5



Energy Calibration

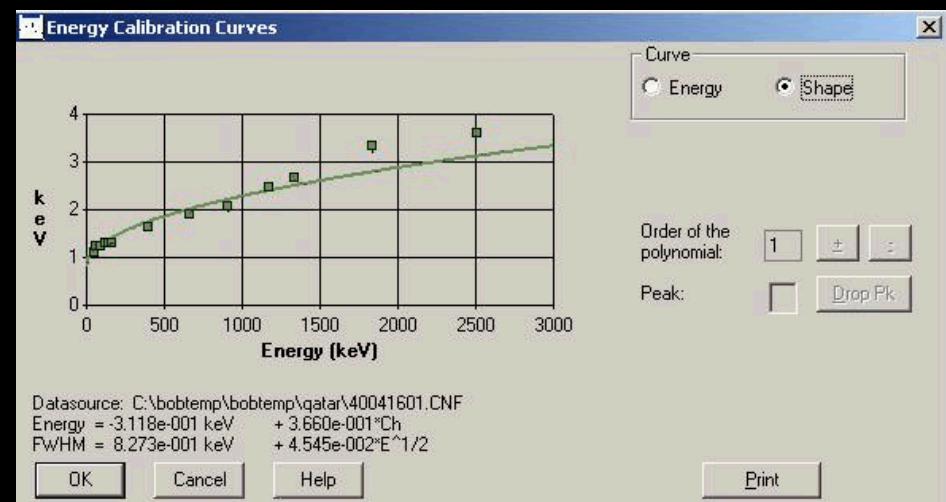
- ENERGY CALIBRATION
 - $E = b + m \cdot ch$
 - Energy response is very linear

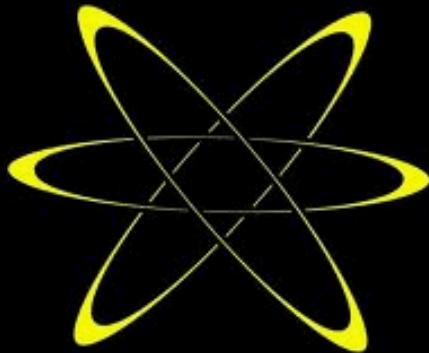




Shape Calibration

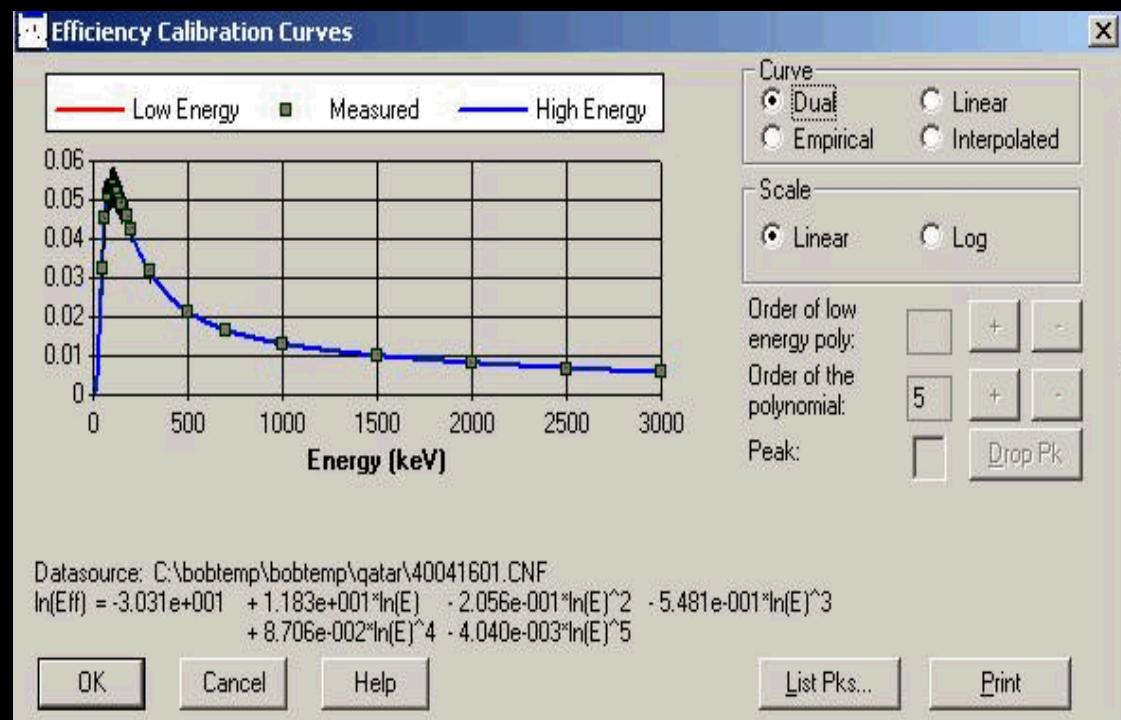
- SHAPE CALIBRATION
 - FWHM
 - Peaks broaden with root of energy

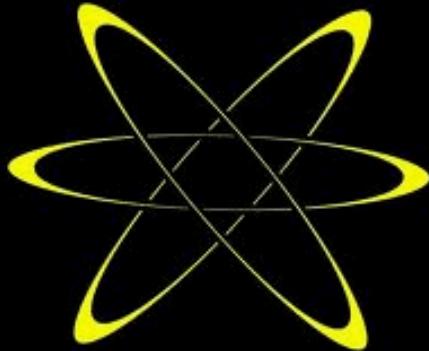




Efficiency Calibration

- Factors That Affect Efficiency
- Geometry
 - Size
 - Shape
 - Distance
- Density
 - Material
 - Atomic Number
 - Packaging



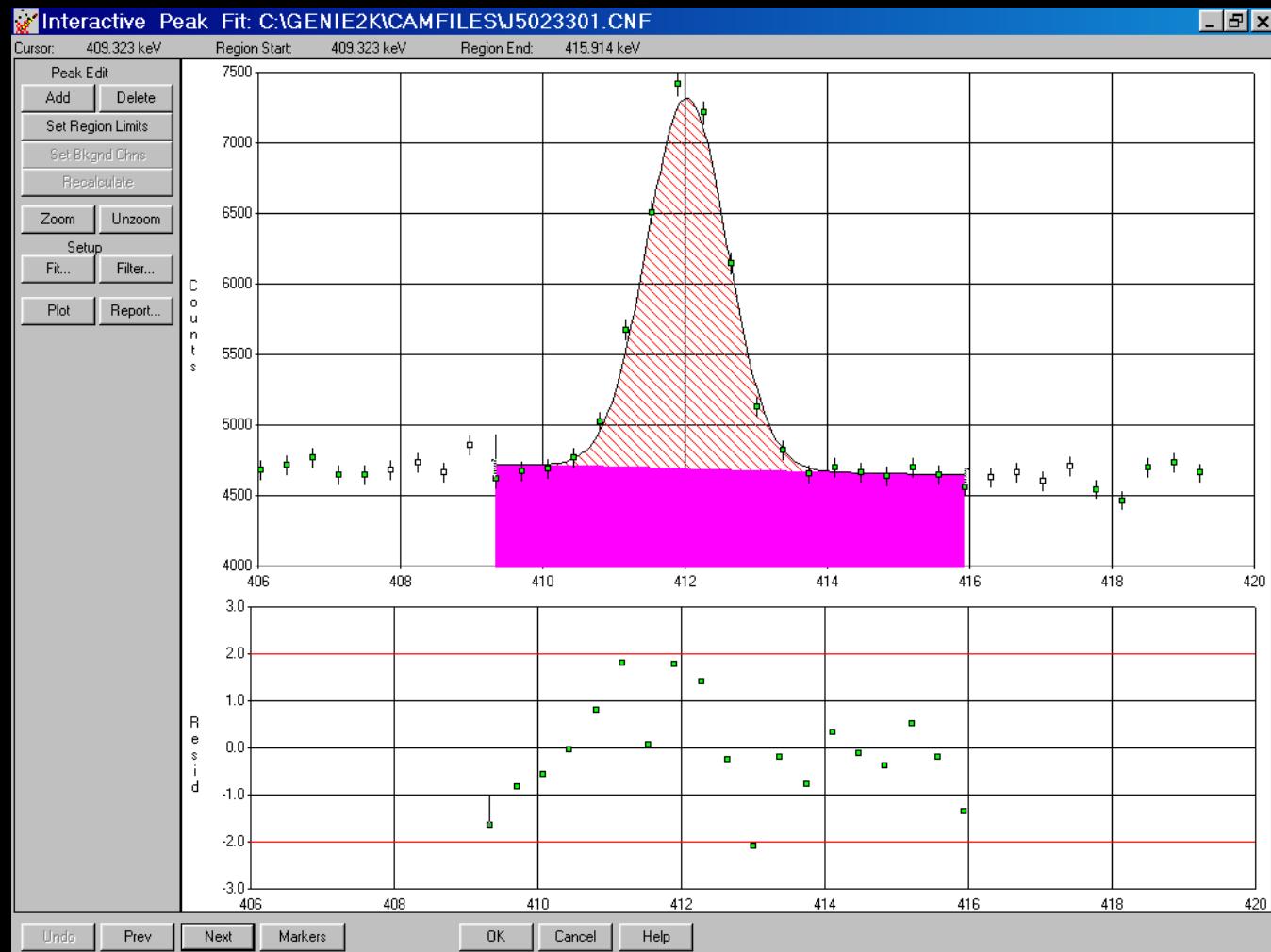


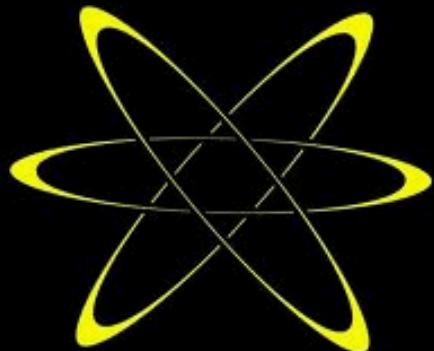
RPSD Calibrated Geometries

- Point Source
- Point Source 5 cm off detector face
- 1 inch Air Filter
- 2 inch Air Filter
- 2 inch Air Filter in Marinelli
- Carbon Cartridge Air Filter
- 4 x 5 Inch Air Filter
- Soil 250 mL Jar
- Liquid 250 mL Jar
- Low Density 250 mL Jar
- Soil 500 mL Marinelli
- Liquid 500 mL Marinelli
- Low Density 500 mL Marinelli
- LSC Vial
- Four Varieties of Drums
- ISOCS Capability (As Needed)



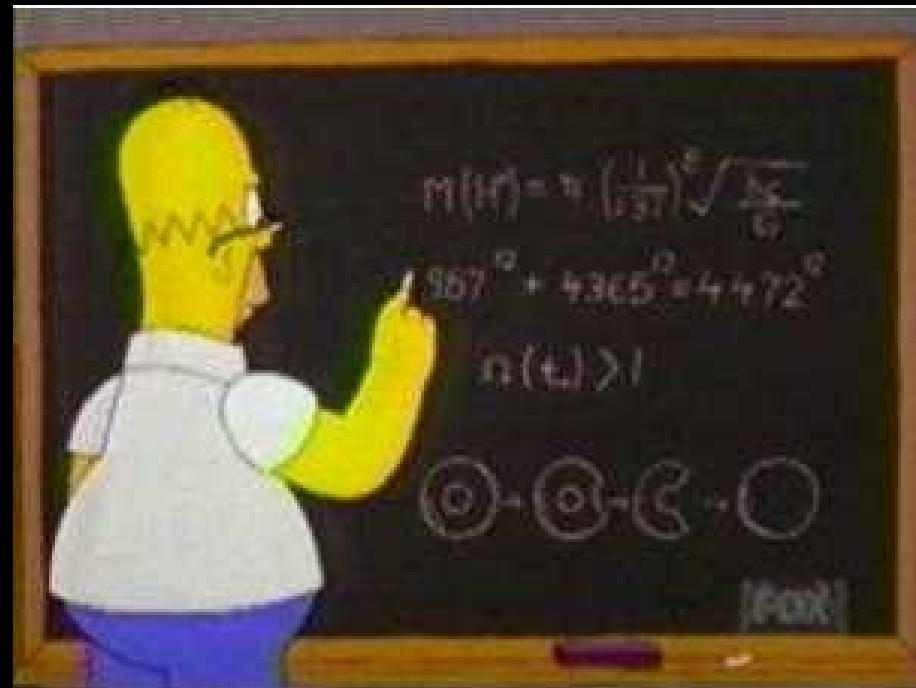
Identifying and Quantifying Peaks

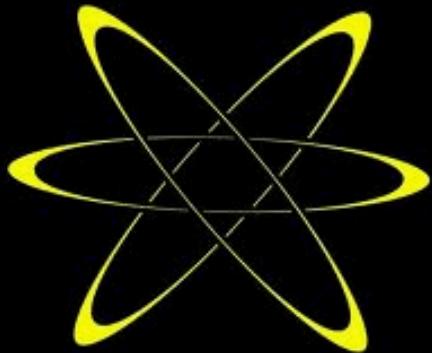




Background Subtraction

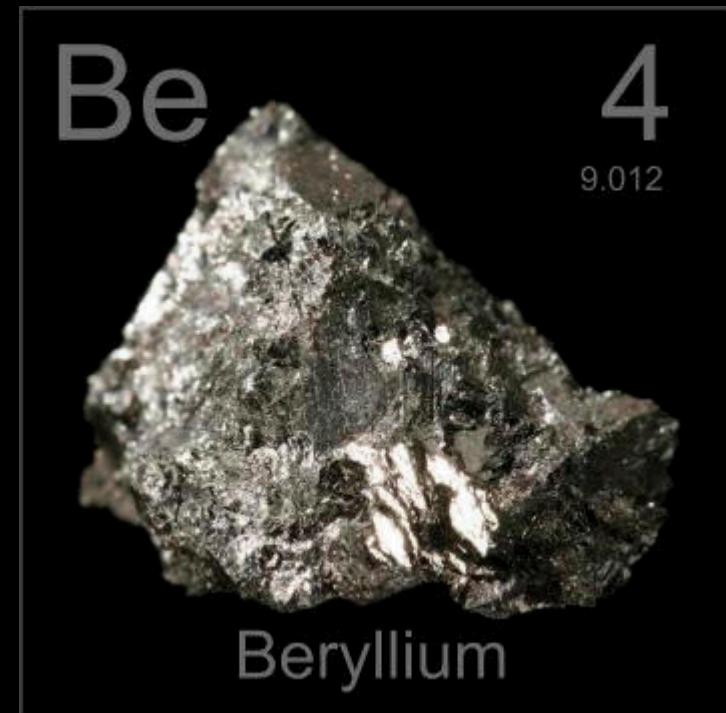
- Monthly Background
- Long Count Time
- Count Rate Subtracted from each applicable peak

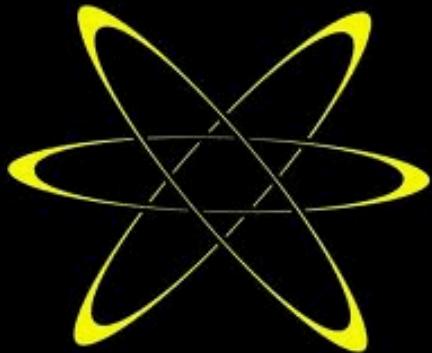




Fun Background Fact

- Beryllium
- Be-7 Occurs in BG
forms from spallation
in atmosphere
- Many Unique
Properties/Attributes

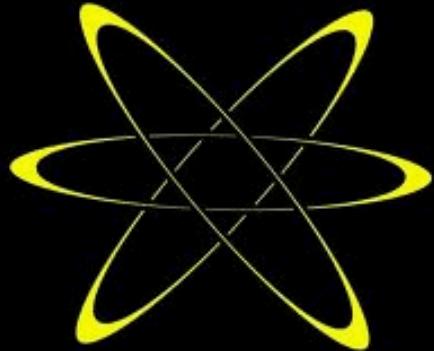




Isotope Identification

- Peak Centroid Compared to User Defined Library
- Abundances cross checked
- Confidence Level Assigned
- Not Perfect!
- Needs a Human to Review!

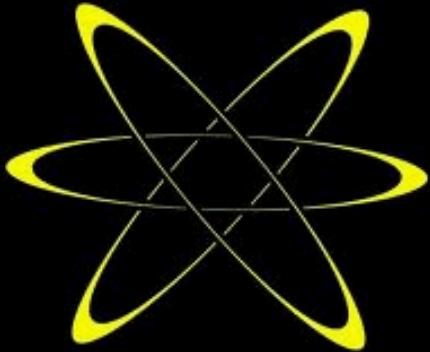




Unidentified Peaks

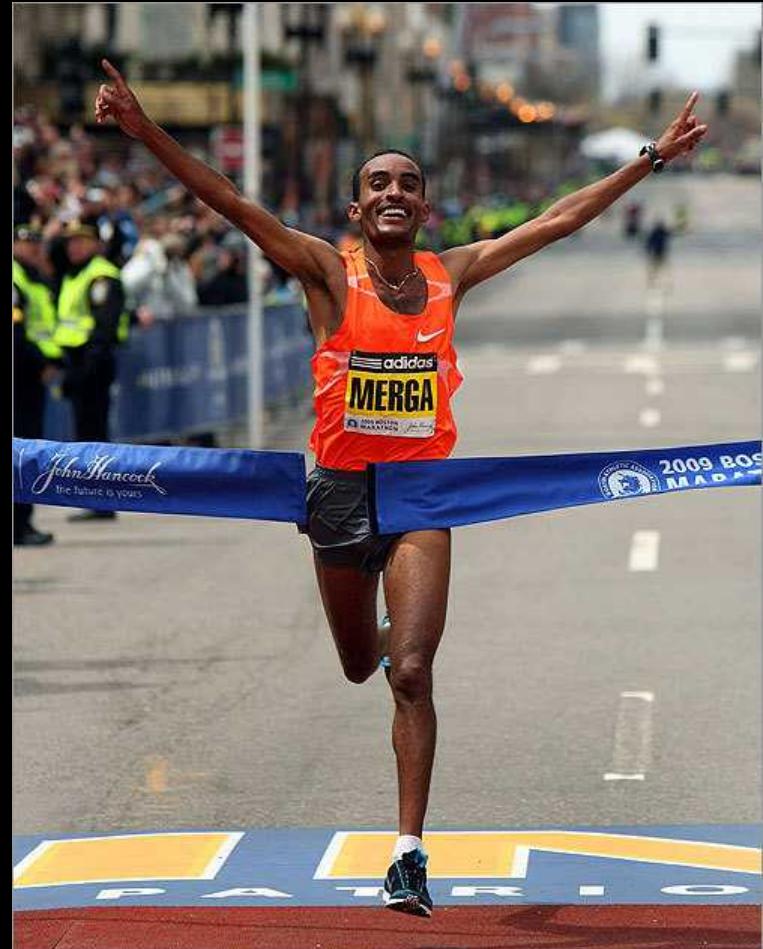
- Very Valuable Information!
- Peaks that didn't fit algorithm
- Helps user identify potentially missed nuclides

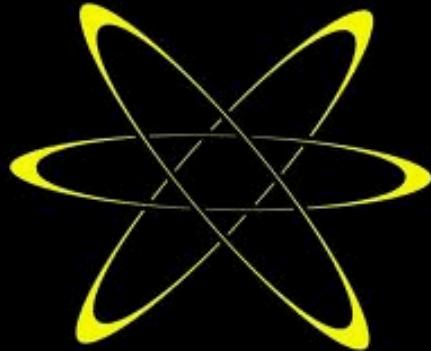




- 8 Nobel Prizes (at least)
- 100 min count time
- 3 Sec computer time
- 10 -100 min analyst time
- 10-15 pages of paper
- Results FINALLY!

Results

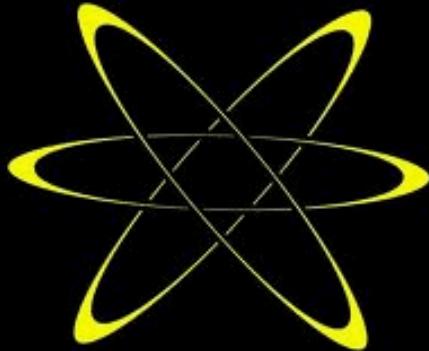




Some Thoughts on Uranium

- Naturally Occurring
- If U-235 is present in spectrum possibly enriched
 - Above 15% enriched probably won't see U-238
- U-235 frequently mis-identified with Ra-226
 - They share a line at 186 keV
- Most of activity belongs to U-234
- ASK QUESTIONS of RPSD!

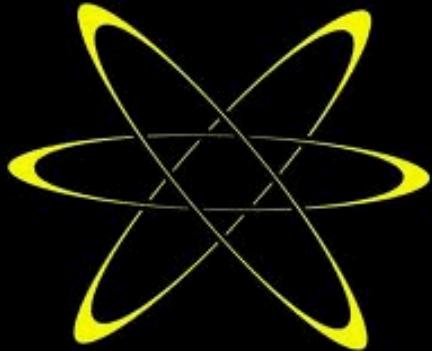




What is Gamma Spec Good For?

- Identifying unknown nuclides
- Quantifying individual isotopes
- Little or no sample prep time
- “Looking” into containers
- ISOCS





Limitations of Gamma Spec

- Long Count Times
- Low Efficiency
- Expensive
- Can't identify pure beta/alpha emitters
- Limitations on detection of many heavy isotopes (Pu, Cm, Np, Th)

