

# Z-Machine Presentation

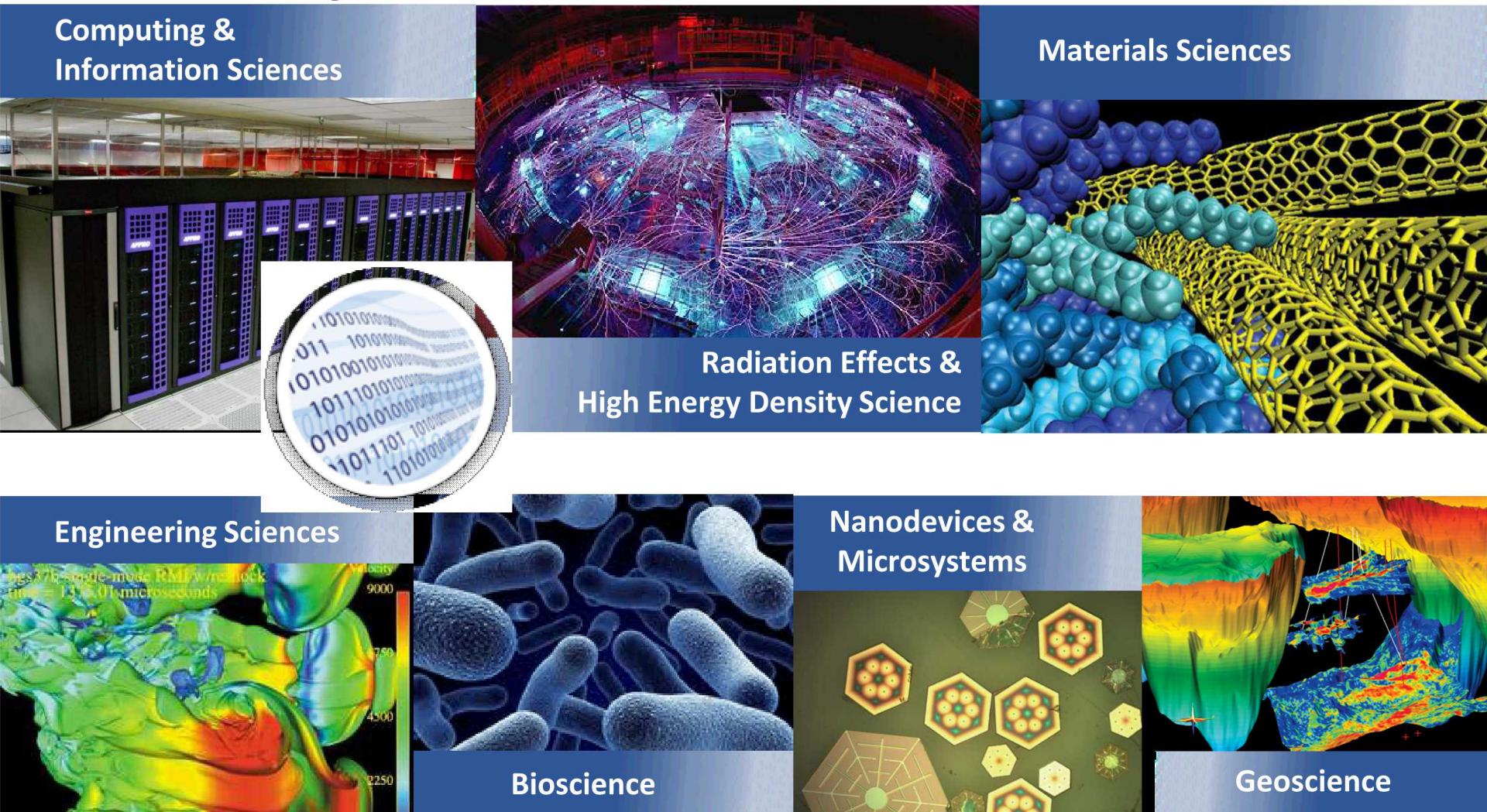
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Campaign 1, Primary Assessment Technology  
CISAC Fellow visit May 12, 2019

SAND2019-

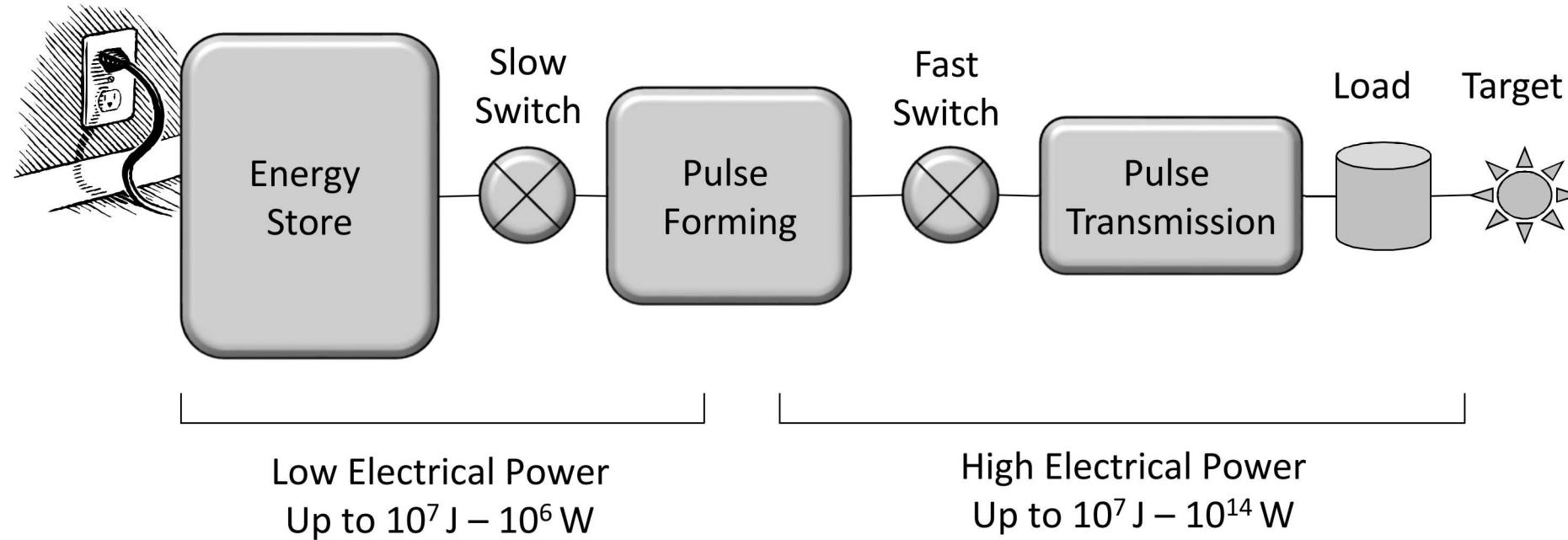


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# Pulsed power plays a key role in Sandia's REHEDS Research Foundation, one of 7 critical disciplines at Sandia supporting our national security missions



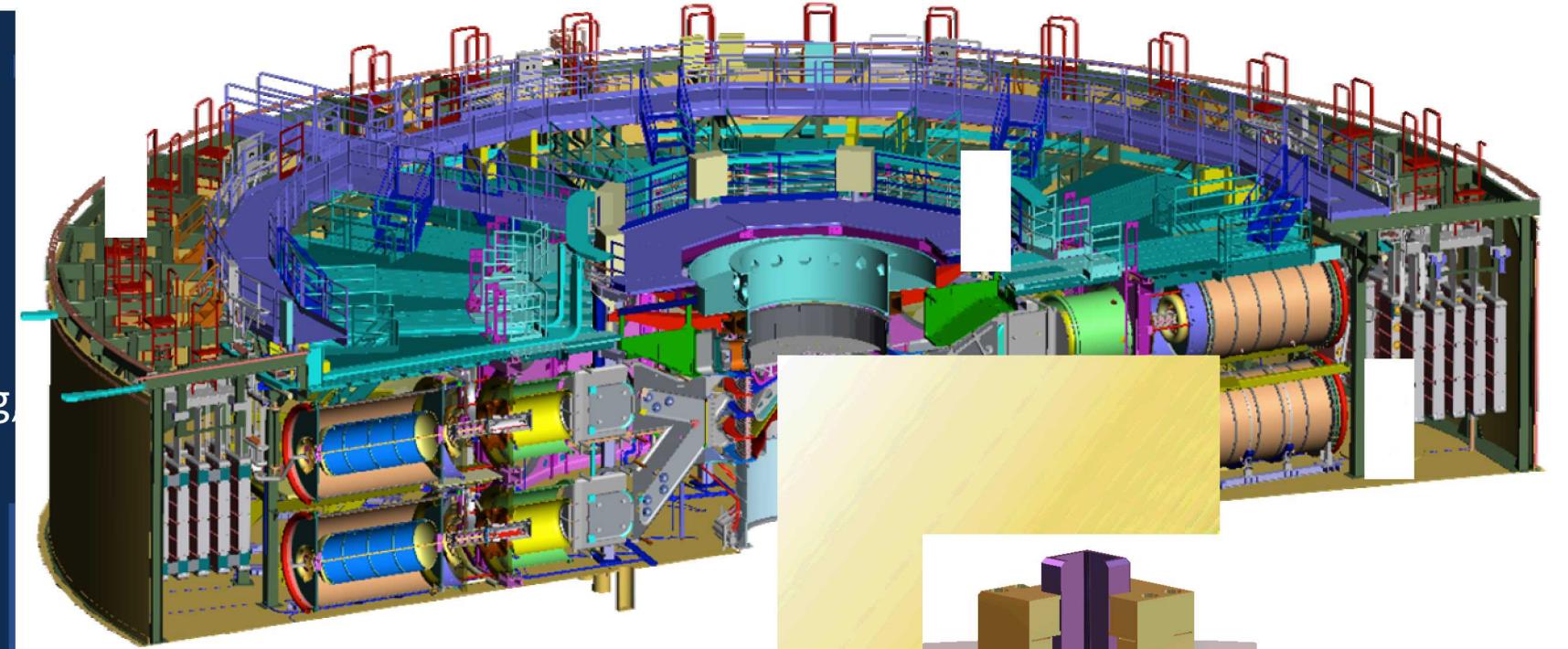
# Pulsed power: The temporal compression of electrical energy to produce short bursts of high power



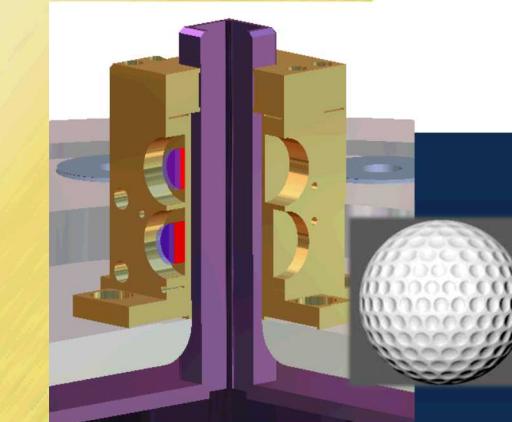
# Sandia's Z Machine is a unique platform for multi-purpose research on high energy density (HED) environments

Acknowledge teams:  
Z operations, cryogenics,  
diagnostics, theory/  
simulations, target design  
and -fabrication, engineering,  
and management

$I \sim 26$  MA,  
 $\tau \sim 100-1000$  ns  
X-ray power  $> 250$  TW  
X-ray energy  $> 2$  MJ



- ▶ **Pulsed Power Technology**
- ▶ **Radiation Sources/-Physics**
- ▶ **Inertial Confinement Fusion**
- ▶ **Materials at high pressure/EOS**



# MHD: currents and the corresponding magnetic fields create matter and radiation in extreme conditions

$$\rho \left( \frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right) = \frac{\mathbf{J} \times \mathbf{B}}{c} - \nabla P \approx \frac{1}{4\pi} \mathbf{B} \cdot \nabla \mathbf{B} - \nabla \left( P + \frac{B^2}{8\pi} \right)$$

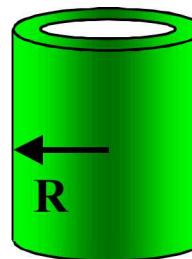
velocity field

drive current  $I$

Pressure

Current x magnetic field

Magnetic field as scalar pressure

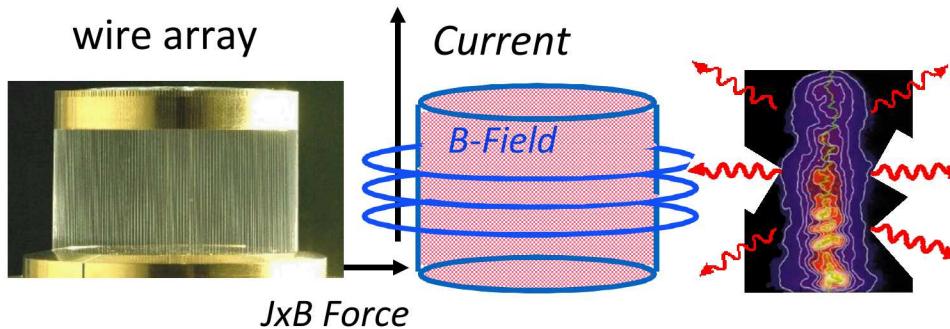


- 25 MA at 1cm radius is 1 Mbar
- 25 MA at 1mm radius is 100 Mbar

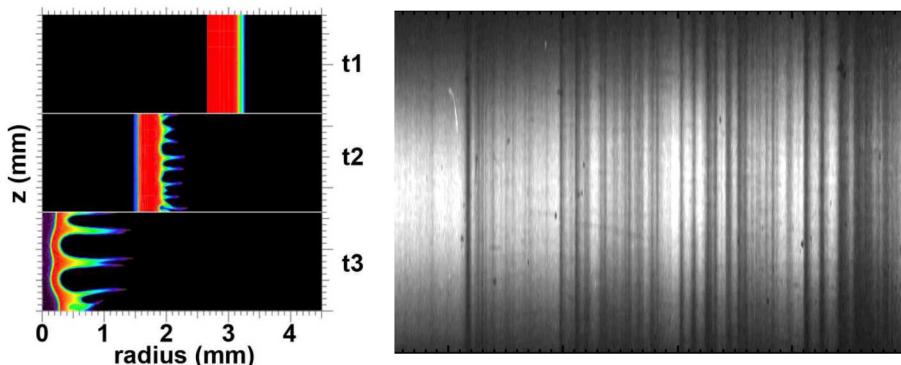
- Using pulsed power (current) as a source has advantages
  - *Can create high pressures without making material hot*
  - Generated over long time scales with control over the time history
  - Large samples and energetic sources (2 MJ to load of 20 MJ stored)
- Integrated projects with theory/simulations/experiment
  - Develop, design, analyze, and optimize experiments

# We use magnetic fields to create HED matter in different ways for different applications

## Radiation physics from Z-Pinch X-ray Sources



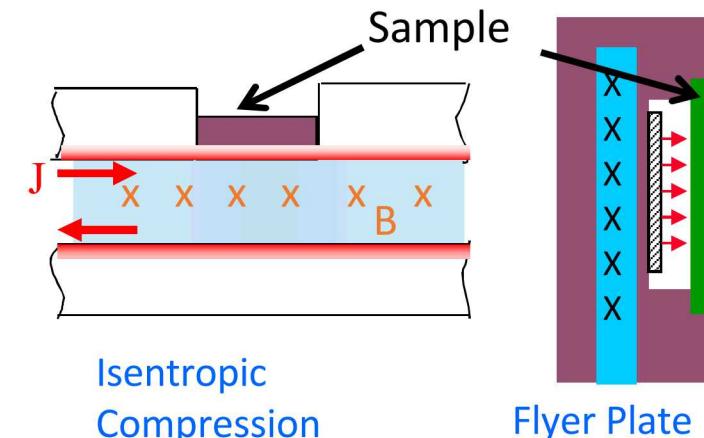
## Atomic- and plasma physics



MRT instabilities

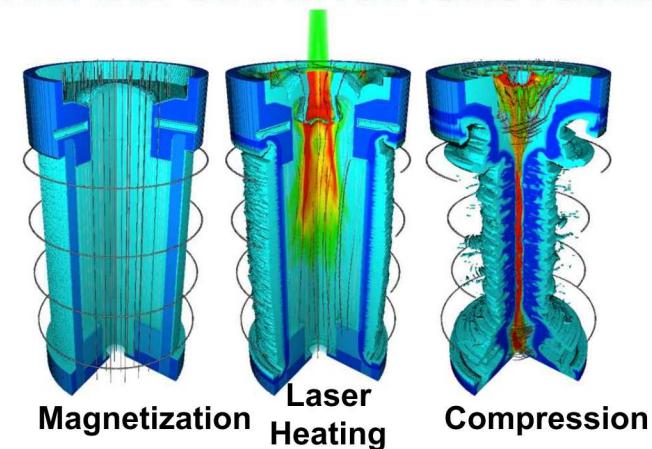
Iron spectrum

## Materials Properties: EOS



Isentropic  
Compression

## Inertial confinement fusion



Magnetization

Laser  
Heating

Compression

# The stockpile stewardship program needs data for materials under HED (High Energy Density) conditions



- A broad range of materials
  - Pu and U for LANL and LLNL
  - Be, DT, Li, CH for National ICF program
  - Stainless steel, W, Ta, water for pulsed power technology
  - Other materials from Al to Zn for a range of applications in HED science

*The Nation's Nuclear Deterrence is not only the hardware – it is also the people of the NNSA Labs*

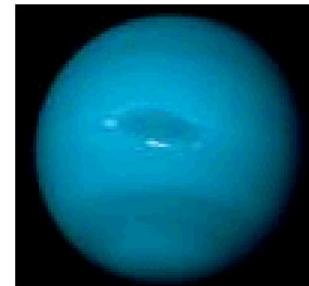
*Scientists, engineers, and others with the education, knowledge, skills, and resources to stay at the very forefront of science.*

*Avoiding Technological Surprise is an important part of what we do.*

# Planetary science also needs data for materials under HED conditions

## ▪ Planetary science – Earths and super-earths

- Silicates, MgO, and iron/iron alloys
- Determining the vaporization threshold for iron – and implications for planetary formation, *Nature Geoscience* 2015.
- Sarah Stewart – recent MacArthur Fellow “Genius award” is a part of the collaboration on planetary science.



## ▪ Planetary science – Jupiter, Saturn, Uranus, Neptune, and exo planets [e.g. hot Neptunes]

- Water, metallization of hydrogen/deuterium: *Science* 2015

*We seek to turn planetary science quantitative by high fidelity modeling and high-precision experiments*

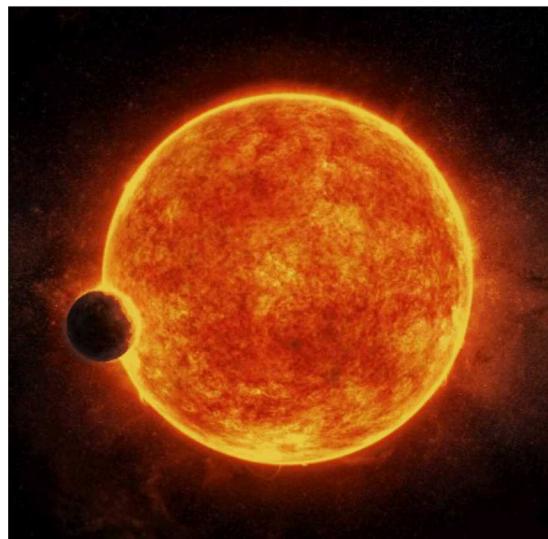
# Vaporization during planet formation and evolution is a key mechanism – large uncertainty in the onset of vaporization



- Giant impacts and the origin of the moon
- Addition and removal of planetary atmospheres
- Chemical evolution of planets
- Exoplanets! Wide range of interiors and atmospheres, mass-radius diagrams

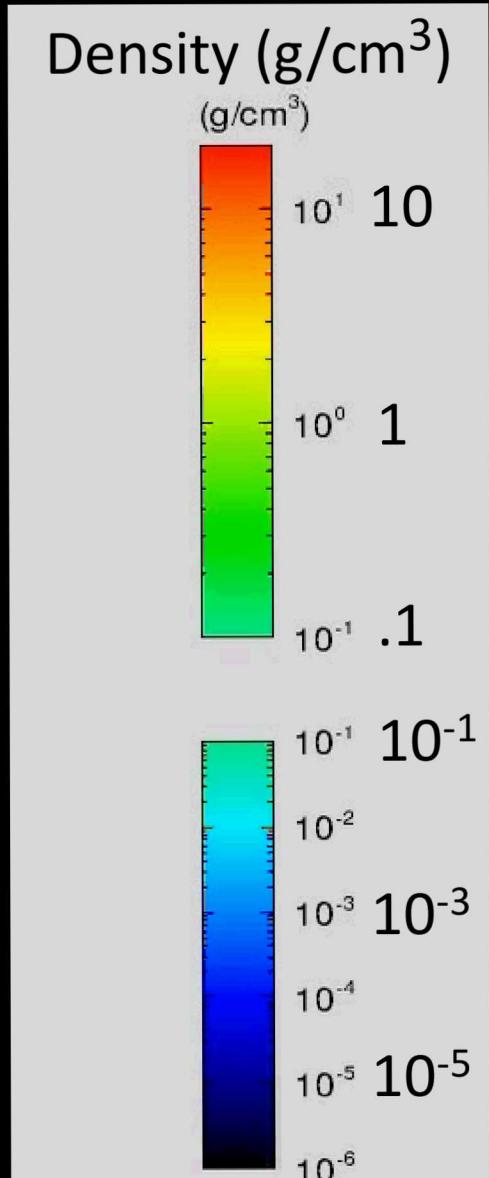
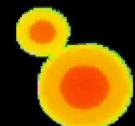
  

- Ted Talk by Professor Sarah T. Stewart “Where did the Moon come from? A new Theory”



# Giant impacts vaporize portions of planets

0.00 hours



D. Crawford, 2011  
CTH calculation

# Z experiments provide material properties in HED conditions to address the moon formation mystery

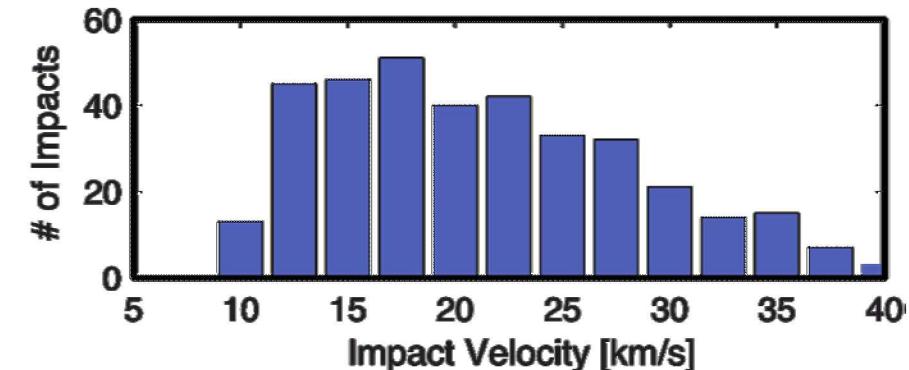


Does an iron meteor:

- plow into a planet as a bullet?
- splatter as a drop of rain?
- vaporize into a cloud of iron to return as iron rain?

*We determined that vaporization is significantly easier than previously thought, changing the way we understand giant impacts.*

*See Sarah's TED talk for an example*

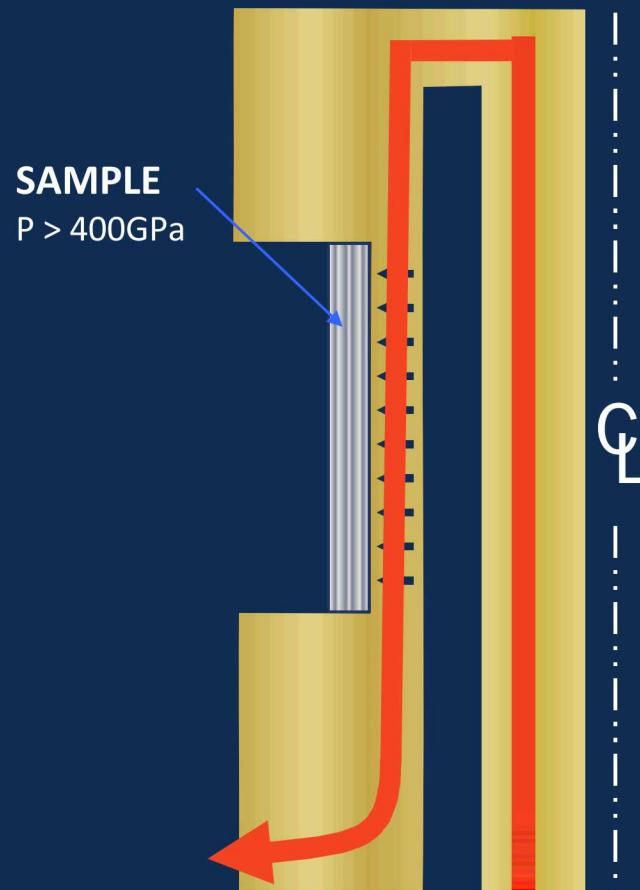


Simulations of planetary dynamics suggest high impact velocities.

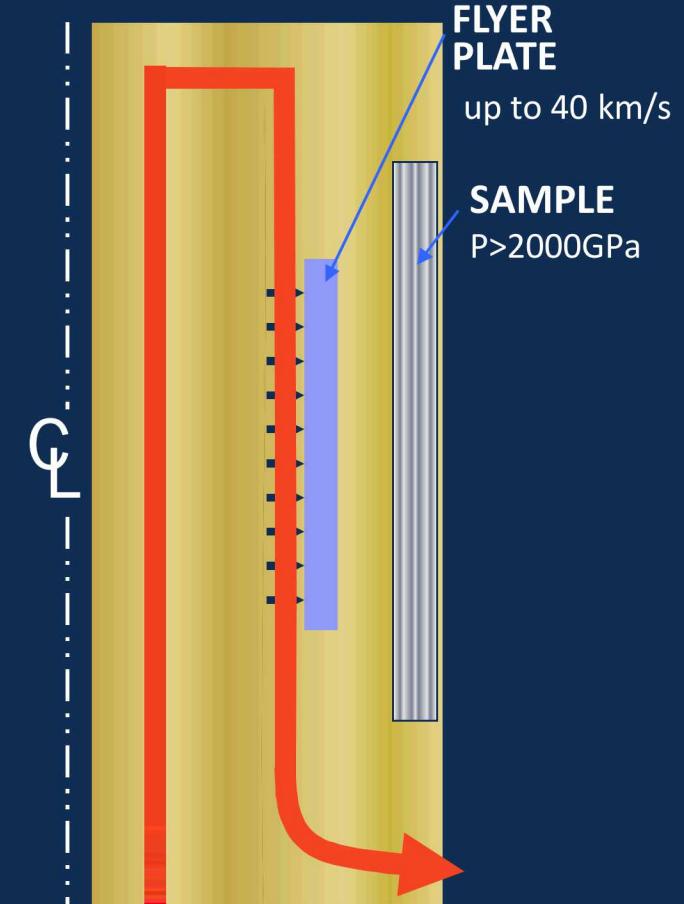
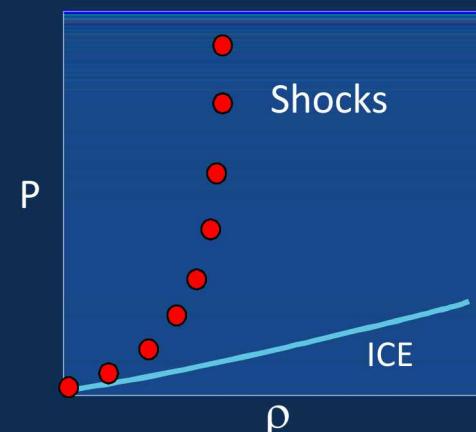
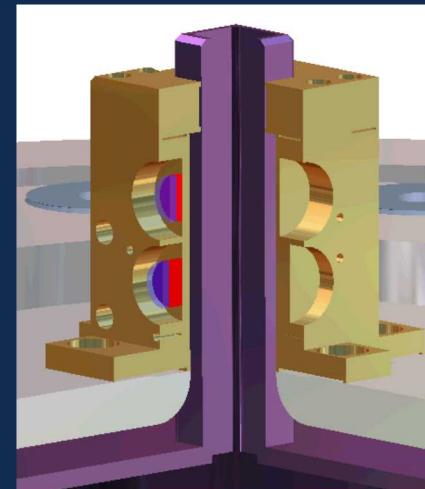
*These velocities are directly accessible on the Z-machine!*

*Impact vaporization of planetesimal cores in the late stages of planet formation*, R.G. Kraus, S. Root, R.W. Lemke, S.T. Stewart, S.B. Jacobsen, and T.R. Mattsson, *Nature Geoscience* 2015 DOI: 10.1038/NGEO2369

# Isentropic compression and shock wave experiments map different regions of phase space



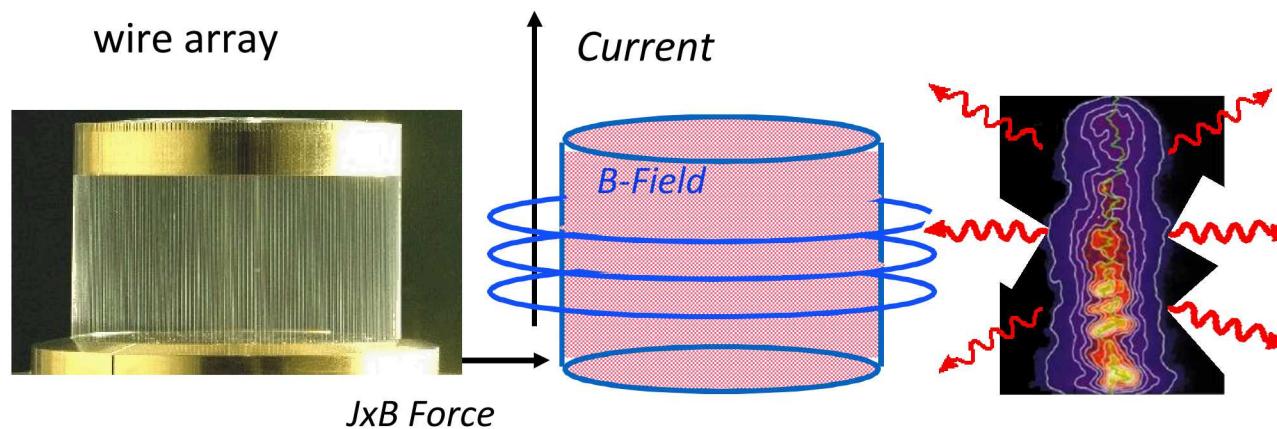
**Isentropic Compression Experiments:**  
Gradual pressure rise in sample



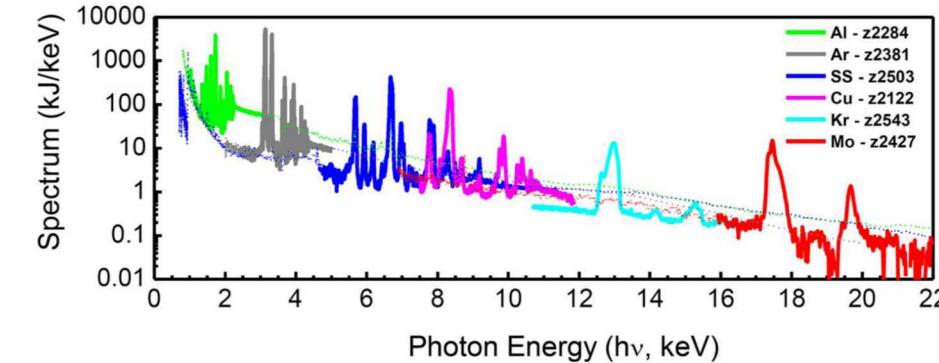
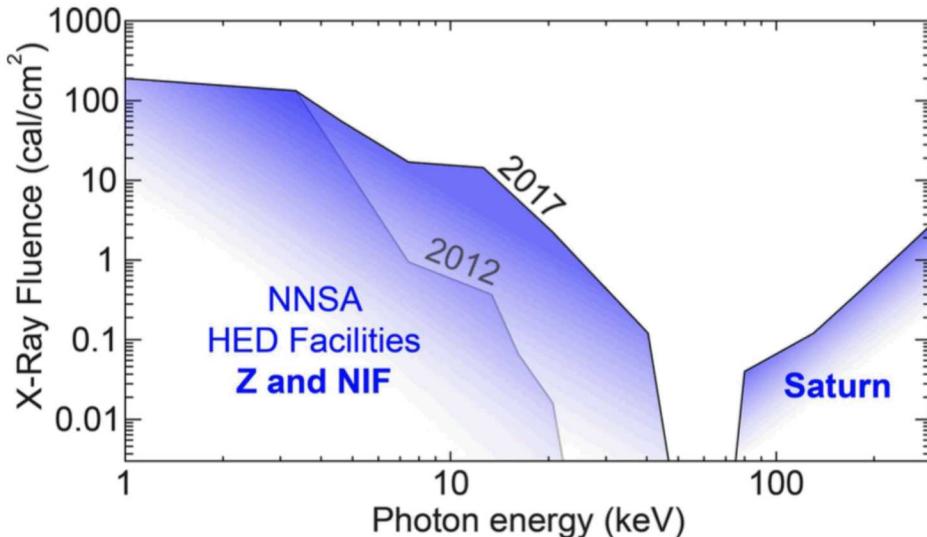
**Shock Hugoniot Experiments:**  
Shock wave in sample on impact

# Radiation driven experiments to study fundamental atomic- and plasma physics is a major part of the ZFSP

## Radiation physics from Z-Pinch X-ray Sources



# Sandia and Lawrence Livermore National Laboratories are collaborating to produce record levels of >10 keV x rays

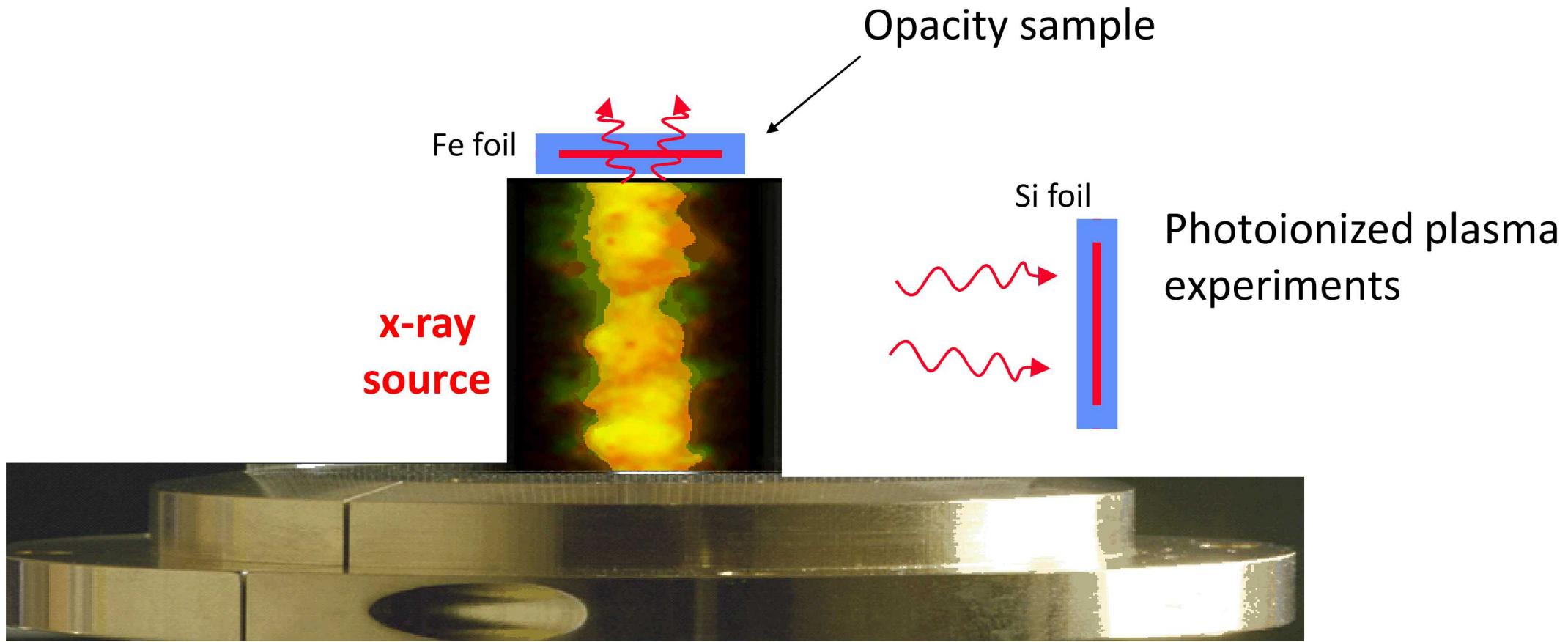


*These x-ray sources are being used to study physics models for matter exposed to rapid, intense doses of x rays*

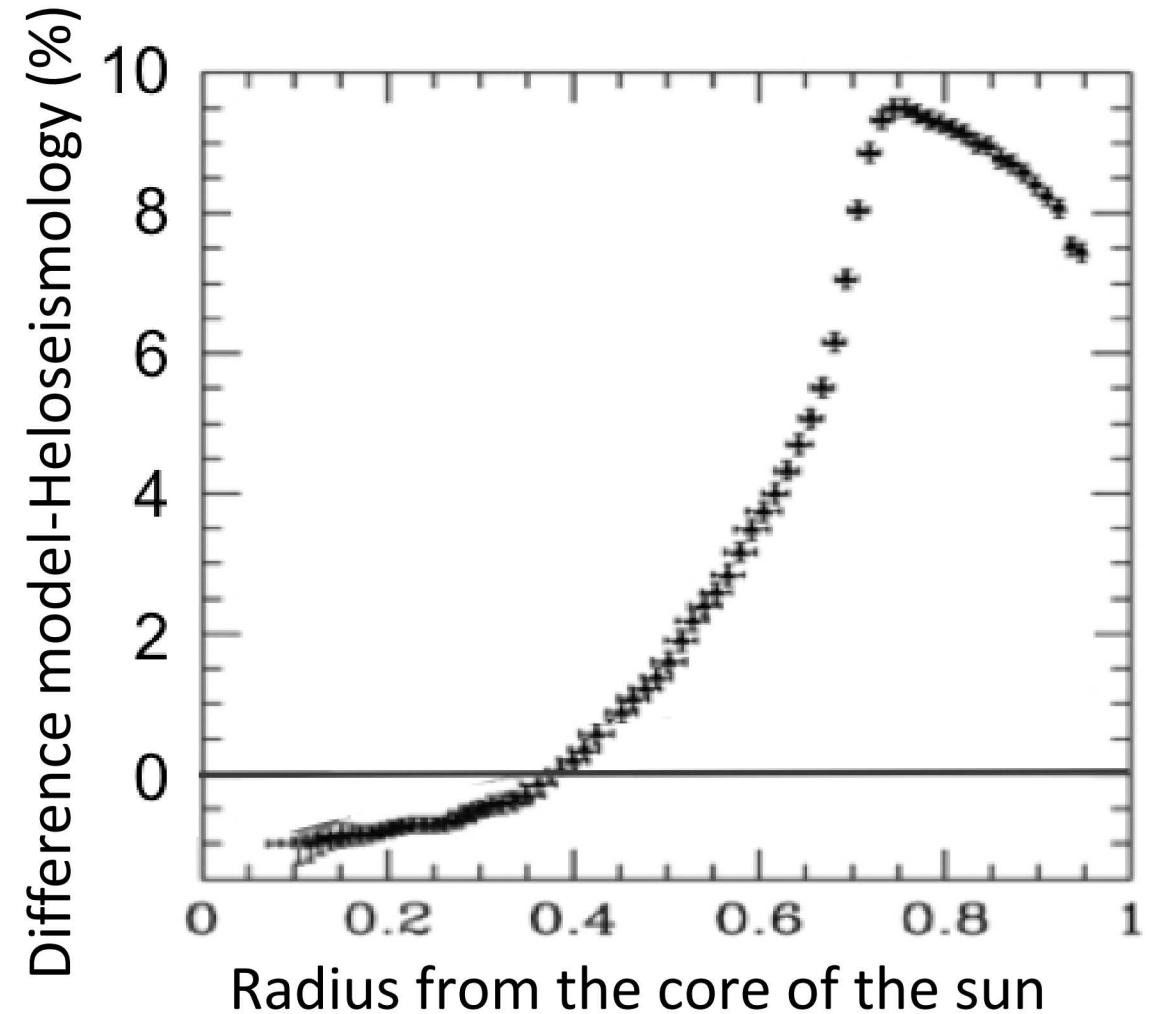
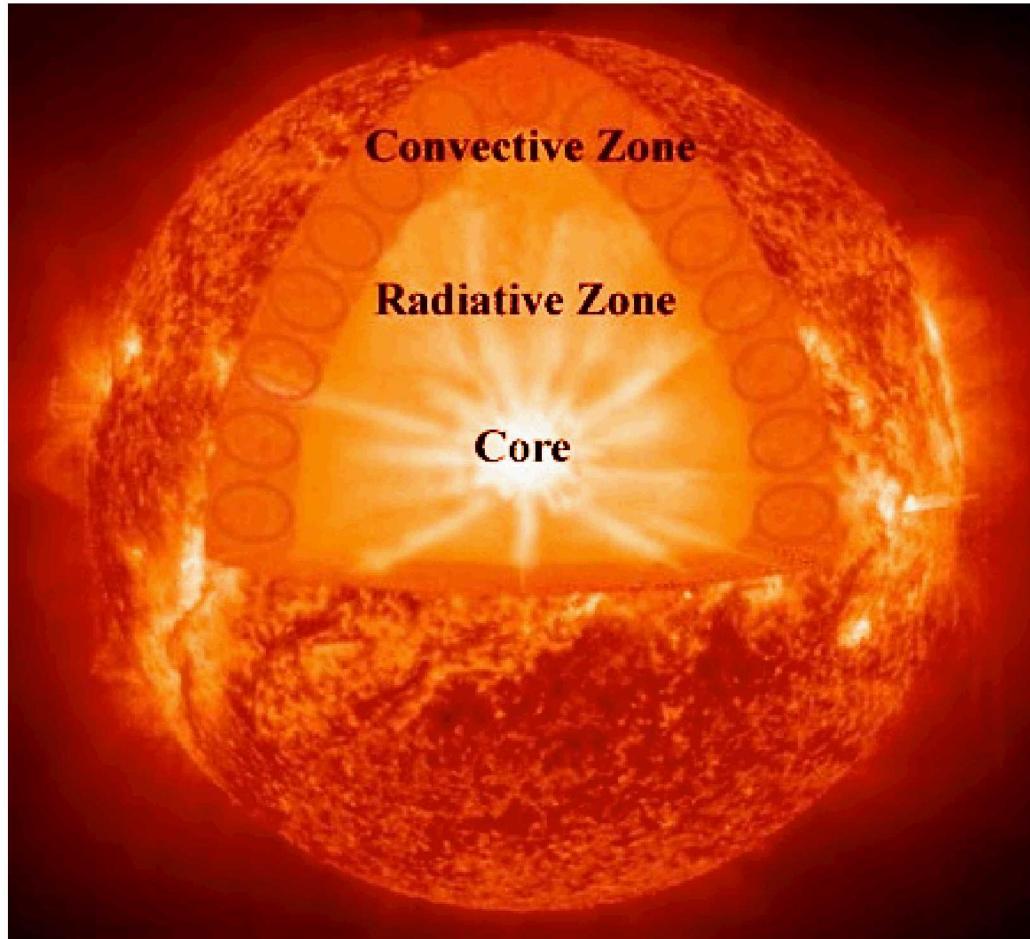
e.g., Studies of high-rate thermal degradation of polyethylene, where ~3 keV x-rays can heat ~100 microns of material at ~ $10^{12}$  K/s.  
Lane & Moore, Phys. Chem. A 122 (2018).

D.J. Ampleford *et al.*, Phys. Plasmas 21, 056708 (2014).

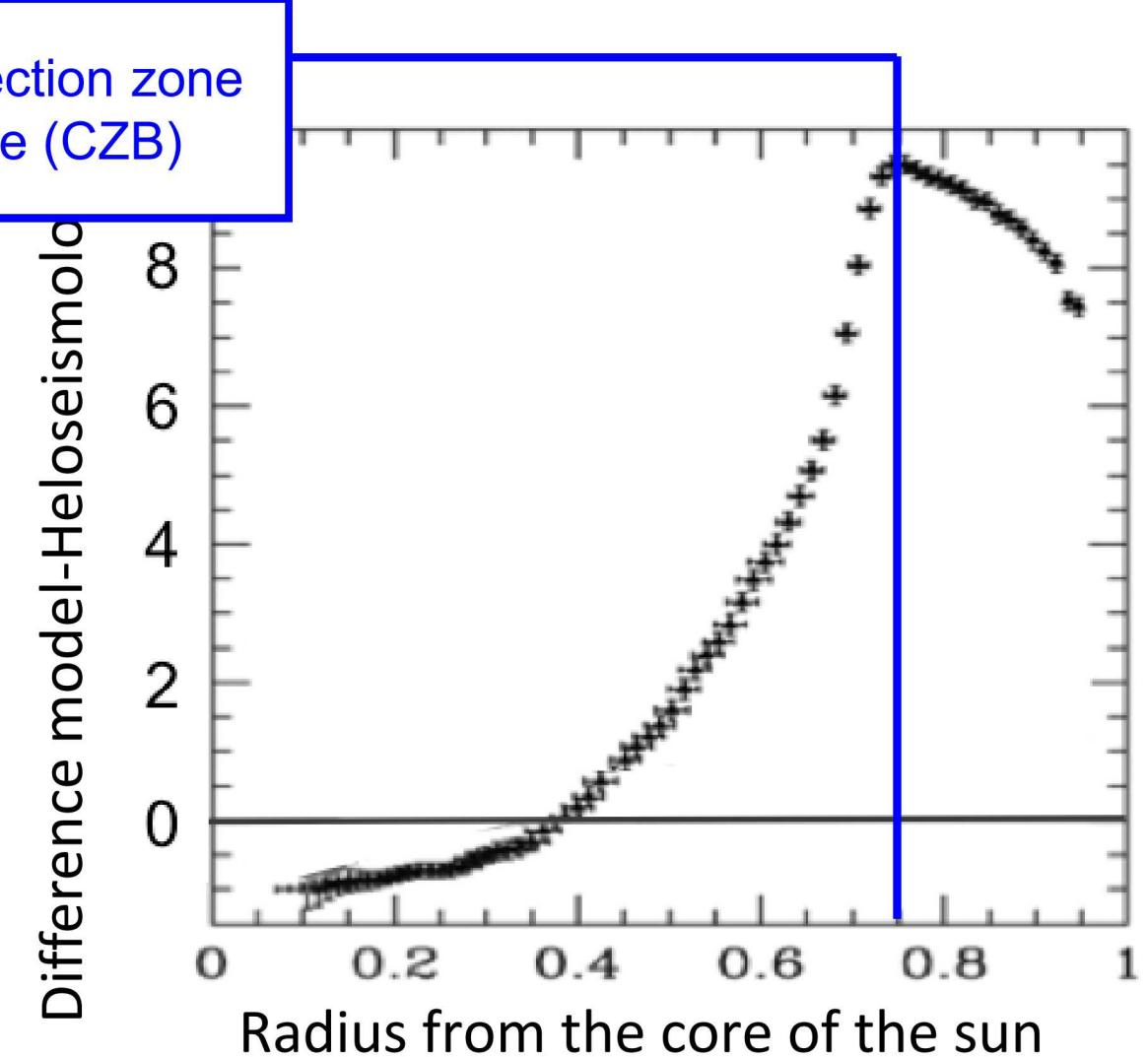
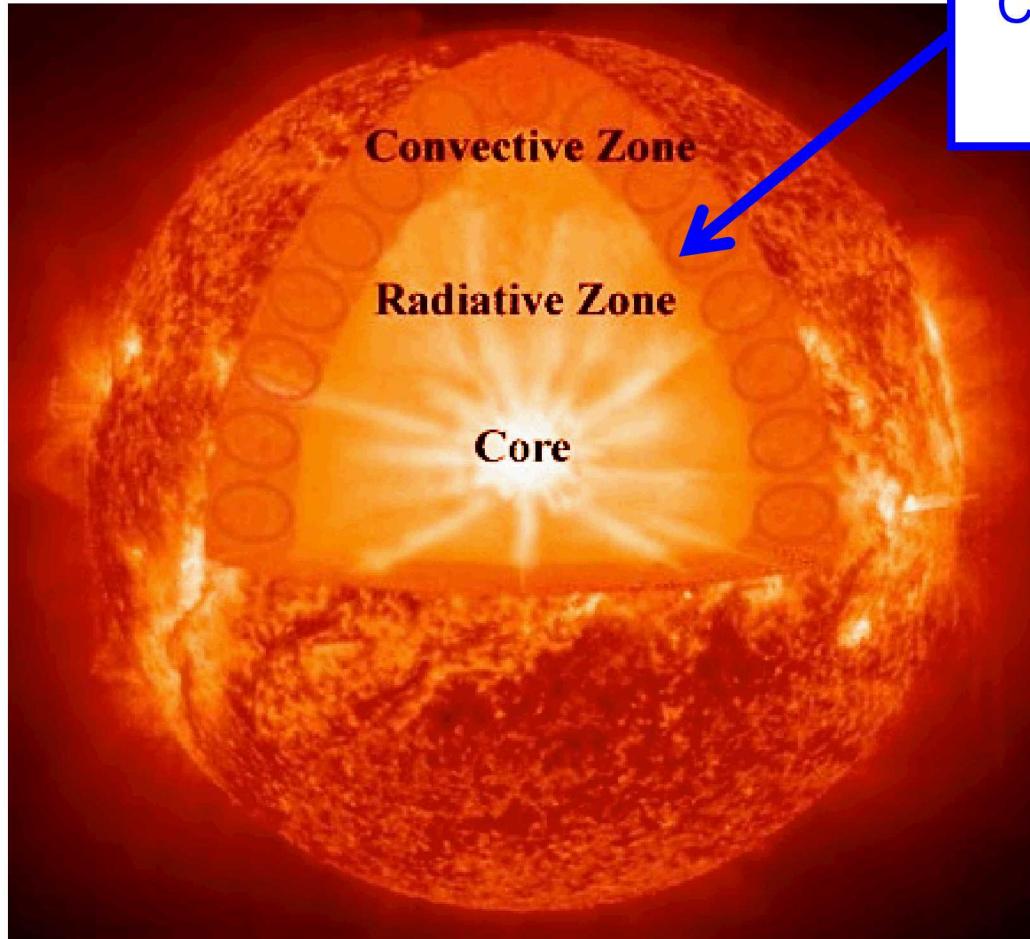
# The Z machine uses 27 million Amperes to create x-rays, and perform multiple benchmark experiments simultaneously



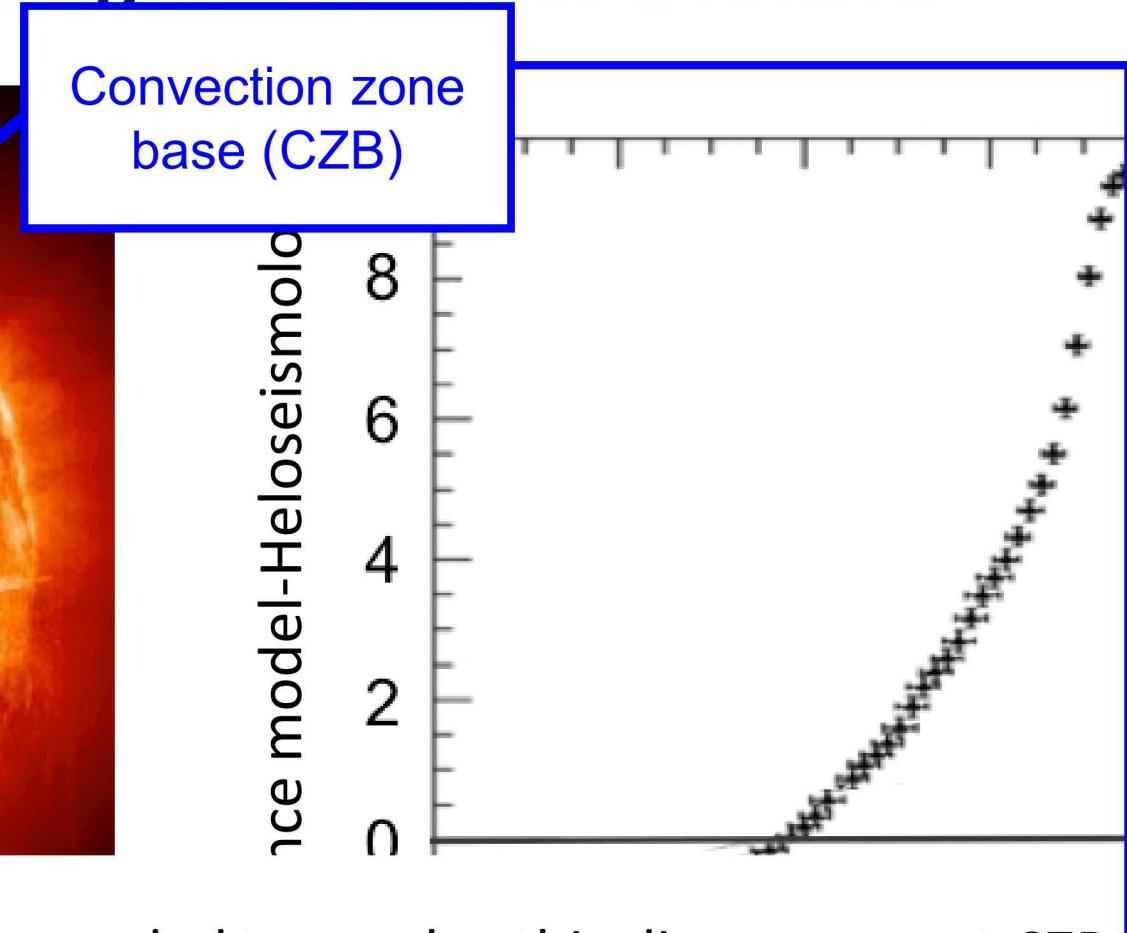
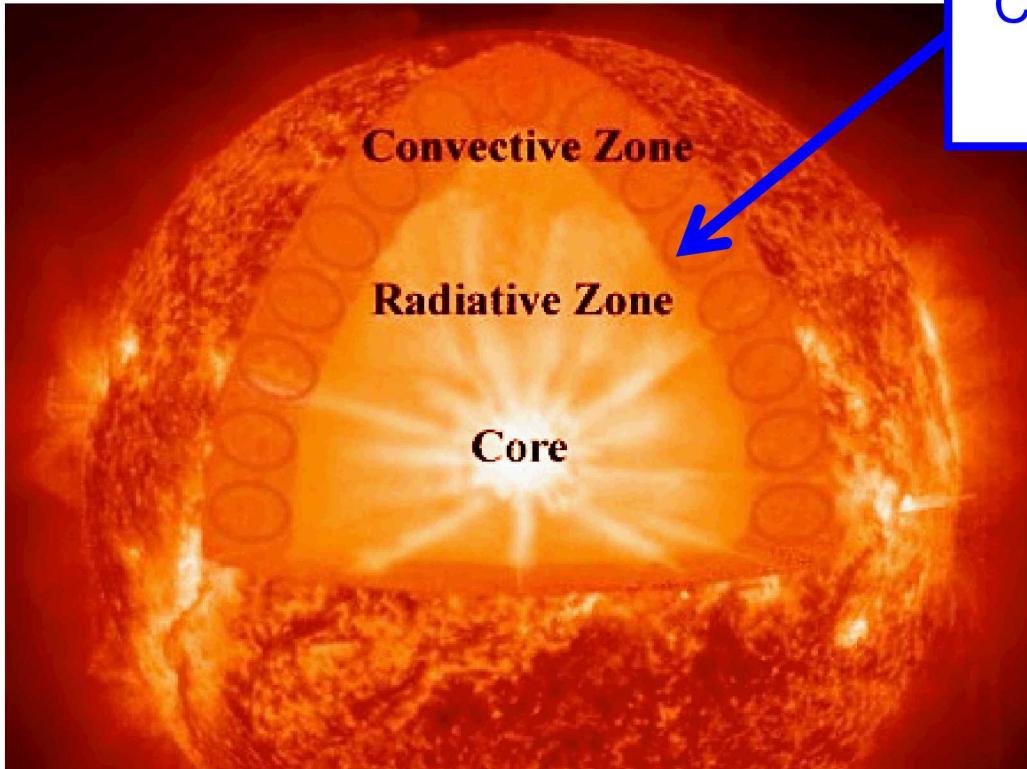
# Modeled solar structure disagrees with observations



# Modeled solar structure disagrees with observations



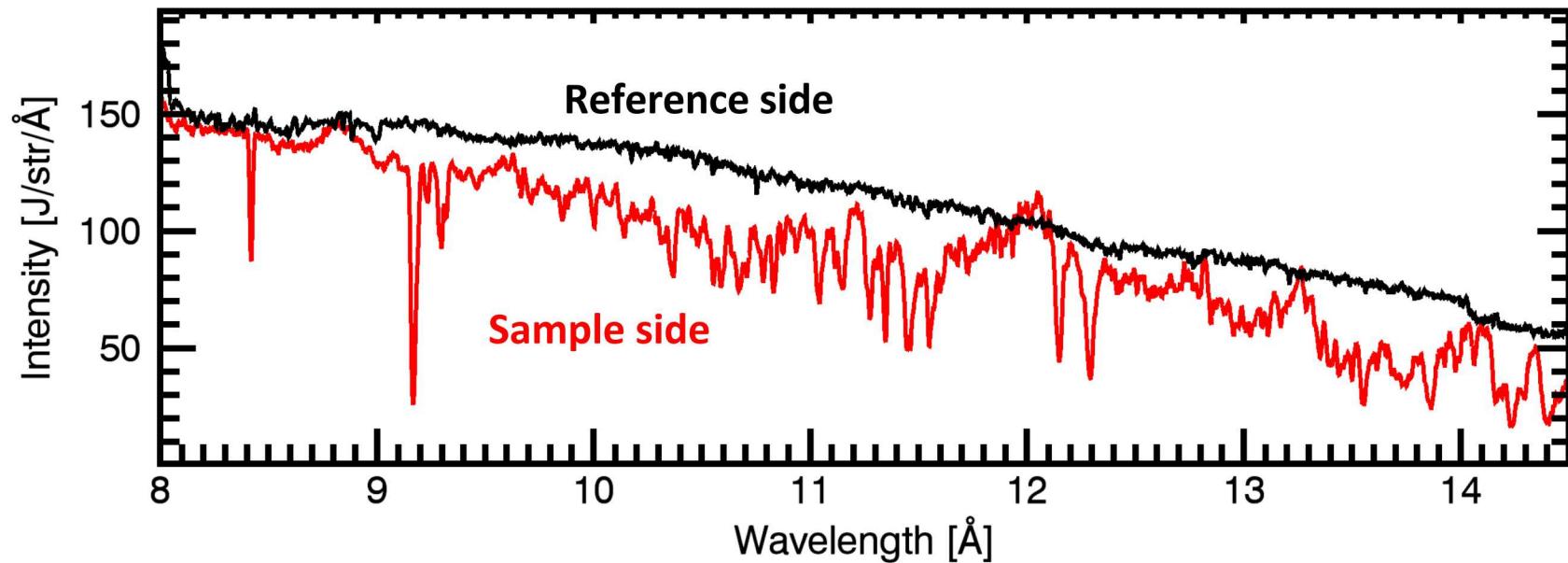
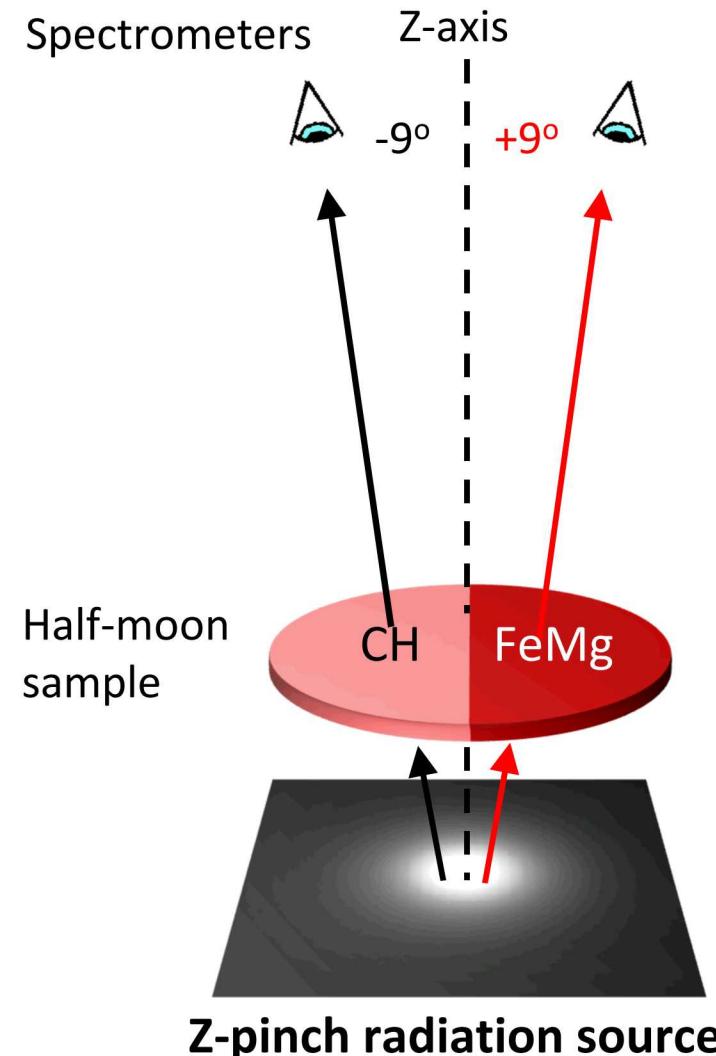
# Modeled solar structure disagrees with observations



- 17% mean-opacity increase is needed to resolve this discrepancy at CZB
- Calculated opacity has never been tested at solar interior conditions

**Objective: Measure Fe opacity at conditions approaching the CZB**

# High-temperature Fe opacities are measured using the Z-Pinch opacity science platform

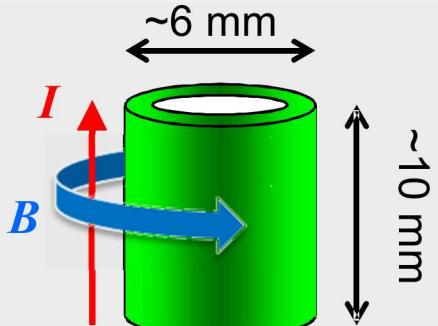


Modeled opacity shows severe disagreement as conditions approach ones in the solar interior

**Bailey et al., *Nature* 517, 56 (2015).**

# Magnetic direct drive provides an alternative way to do ICF using an axial B-field to reduce confinement requirements

## Magnetic Direct Drive (MDD)

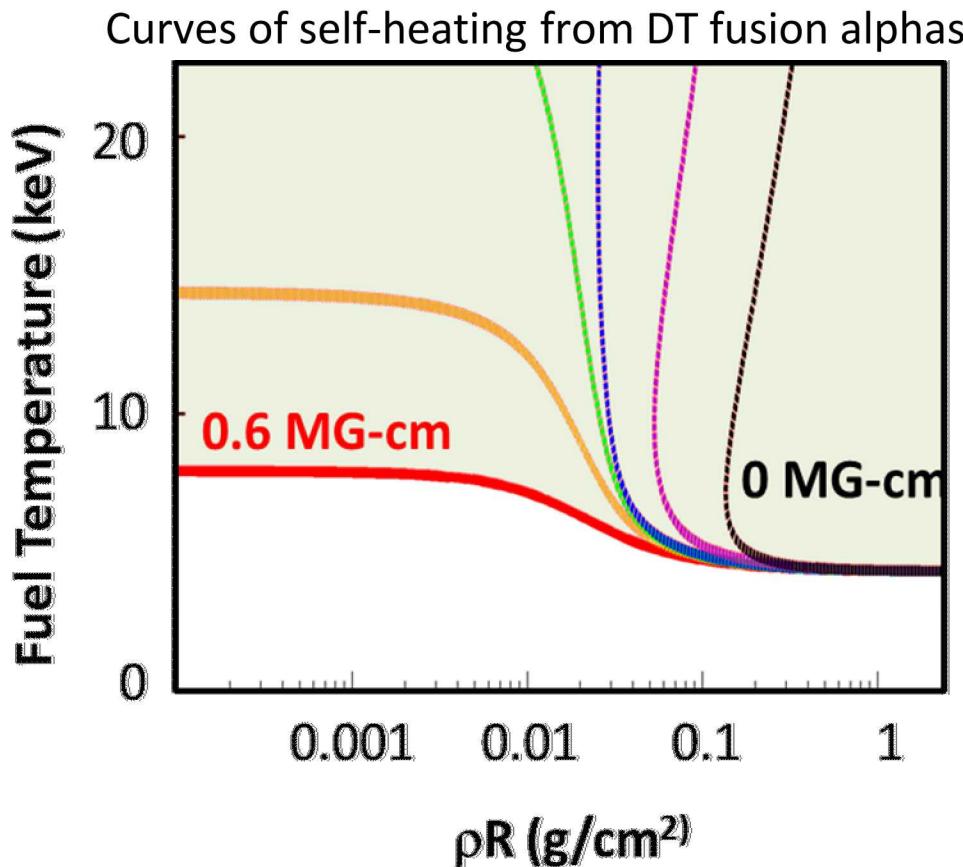


### Drive Pressure

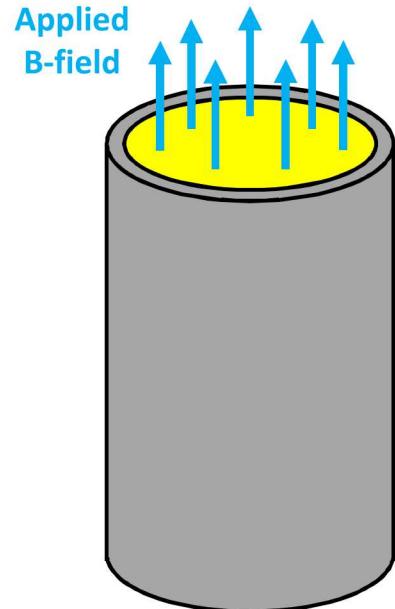
$$P = \frac{B^2}{8\pi} = 105 \left( \frac{I_{MA}/26}{R_{mm}} \right)^2 \text{ Mbar}$$

- Cylindrical convergence
  - Harder to achieve high  $\rho r$
- Thick liners (~500  $\mu\text{m}$ )
  - Harder to achieve high velocity

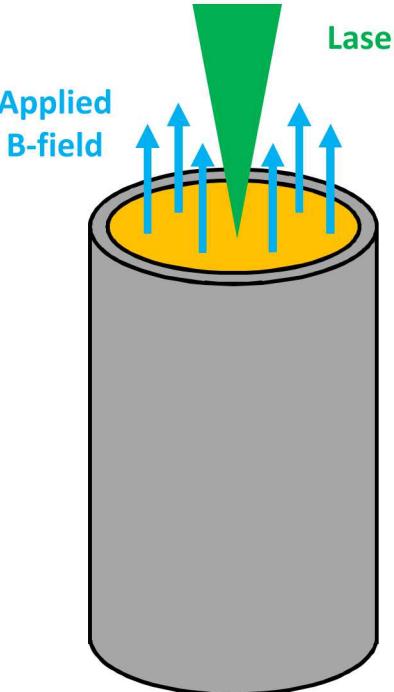
Imposing an axial B-field relaxes  $\rho r$  requirements



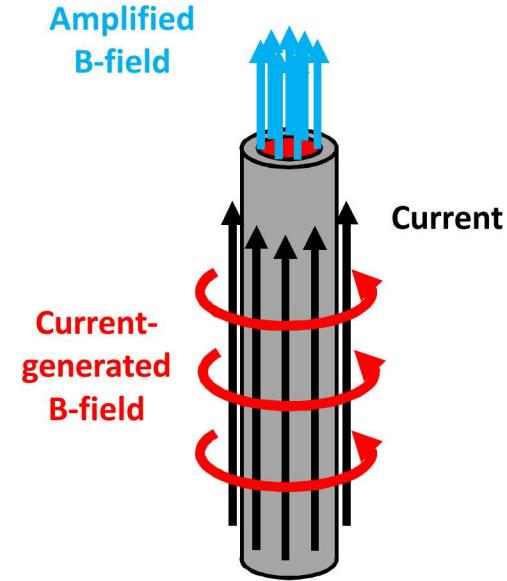
# Magnetized Liner Inertial Fusion (MagLIF) relies on three stages to produce fusion relevant conditions



Apply axial magnetic field

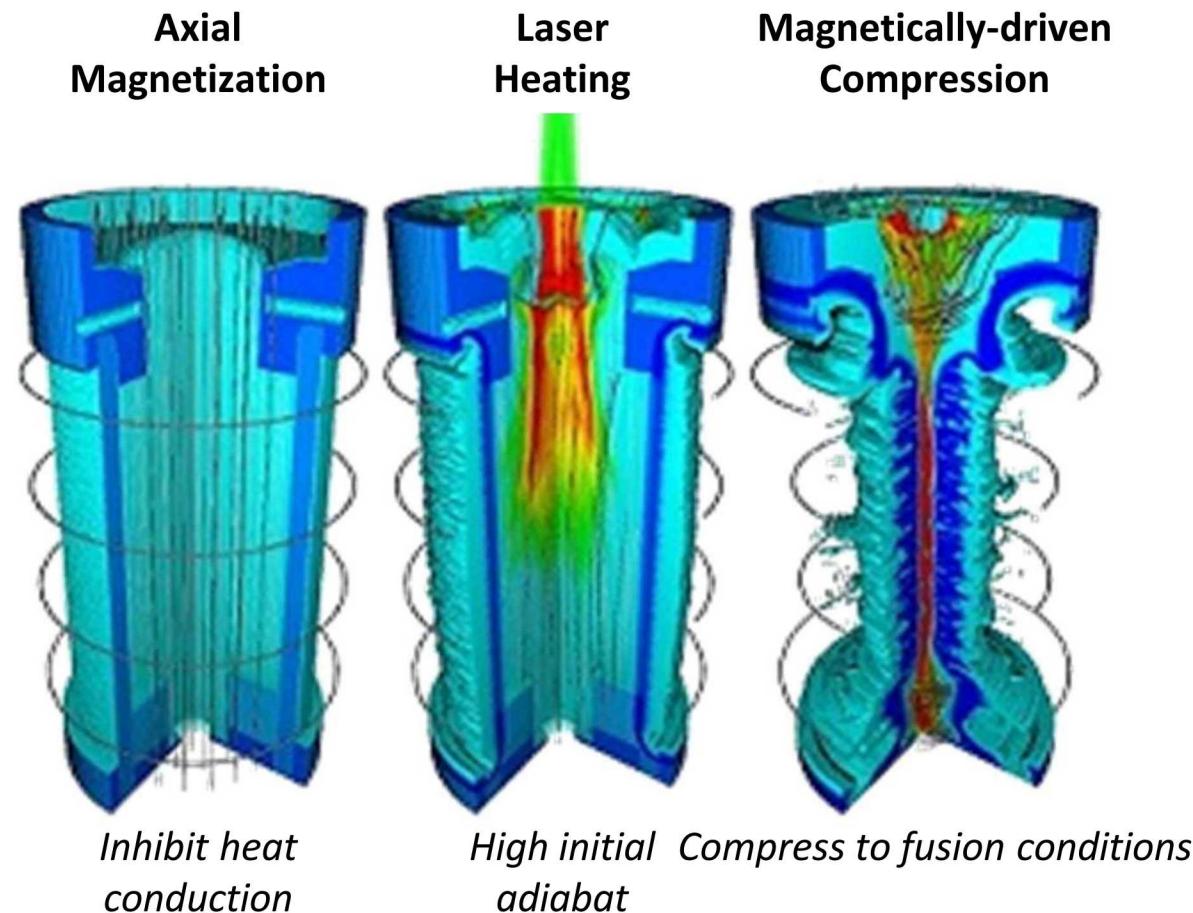


Heat the magnetized fuel  
(e.g., with a laser)



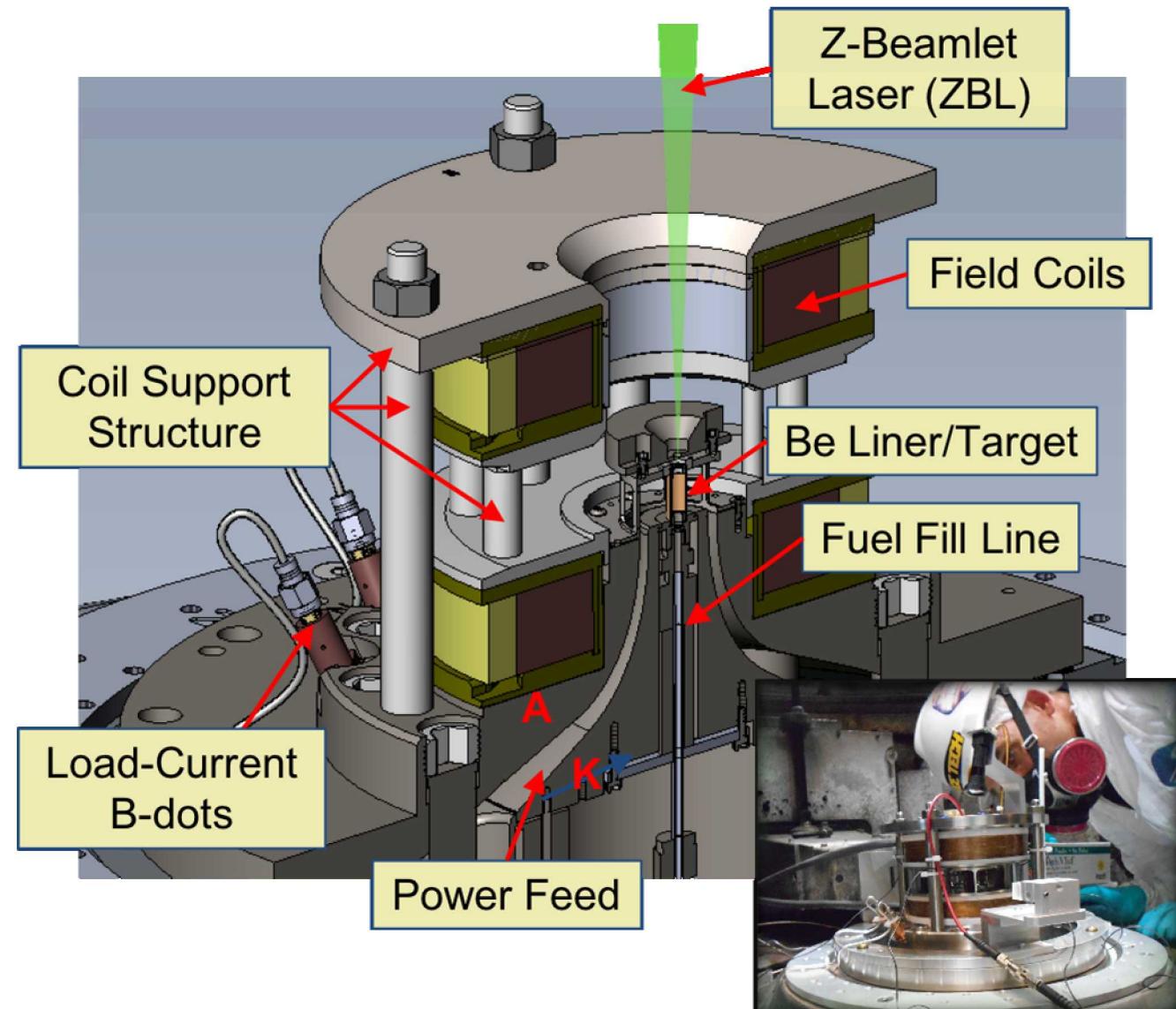
Compress the heated  
and magnetized fuel

# Magnetized Liner Inertial Fusion (MagLIF) relies on three stages to produce fusion relevant conditions



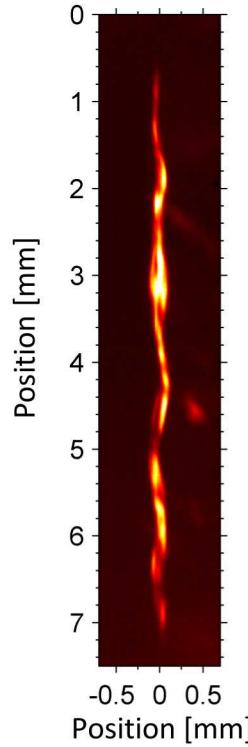
# Configuration of initial MagLIF Experiments (ca. ~2014)

- **Field Coils:** Helmholtz-like coil 10-30 T axial field  $\sim 3$  ms rise time
- **ZBL:** 1-4 kJ green laser, 1-4 ns square pulse w/ adjustable prepulse (prepulse used to help disassemble laser entrance window)

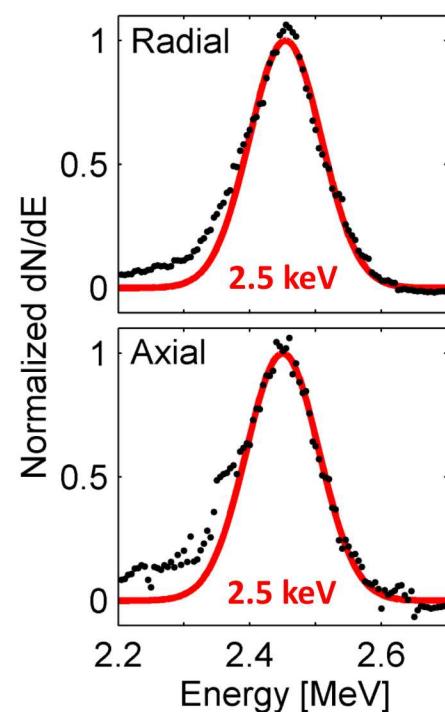


# We have demonstrated key aspects of magneto-inertial fusion on Sandia's Z facility

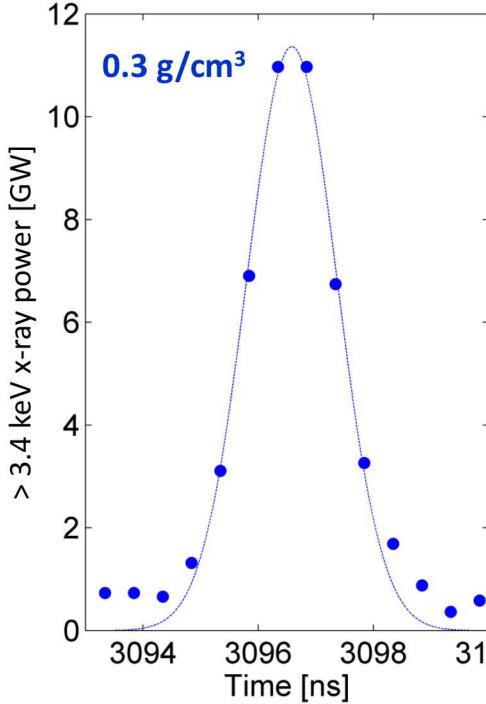
## Well-behaved stagnation volume



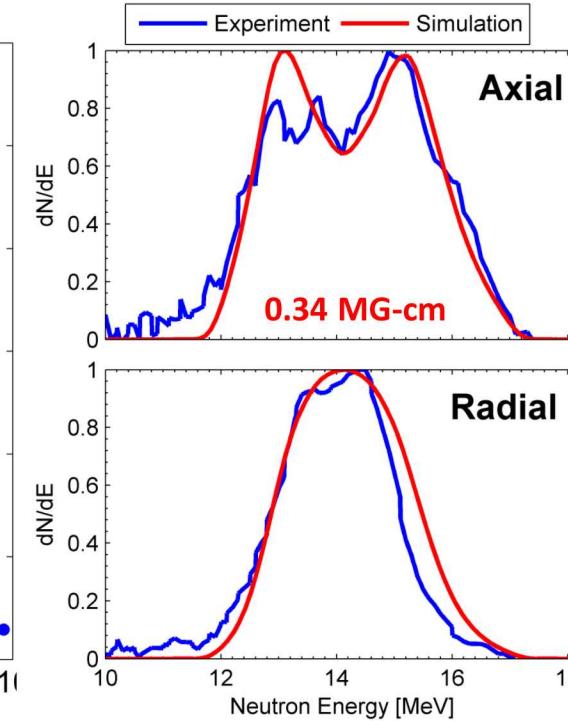
## Relevant temperatures



## Relevant densities



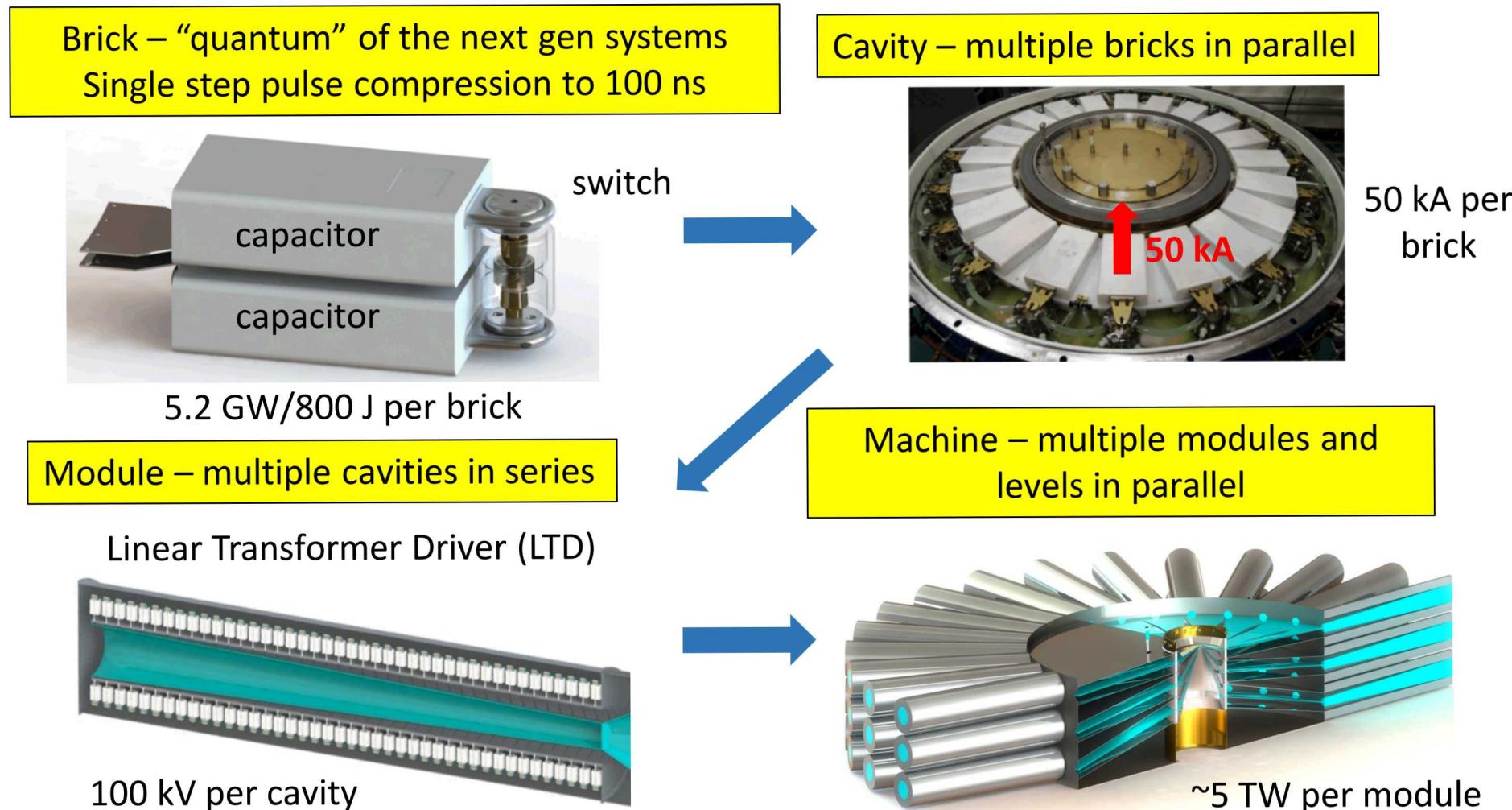
## Relevant fuel magnetization



M.R. Gomez *et al.*, Phys. Rev. Lett. 113, 155003 (2014);  
K.D. Hahn *et al.*, RSI 85 (2014);  
S.B. Hansen *et al.*, Phys. Plasmas 22, 056313 (2015);  
P.F. Schmit *et al.*, Phys. Rev. Lett. 113, 155004 (2014).

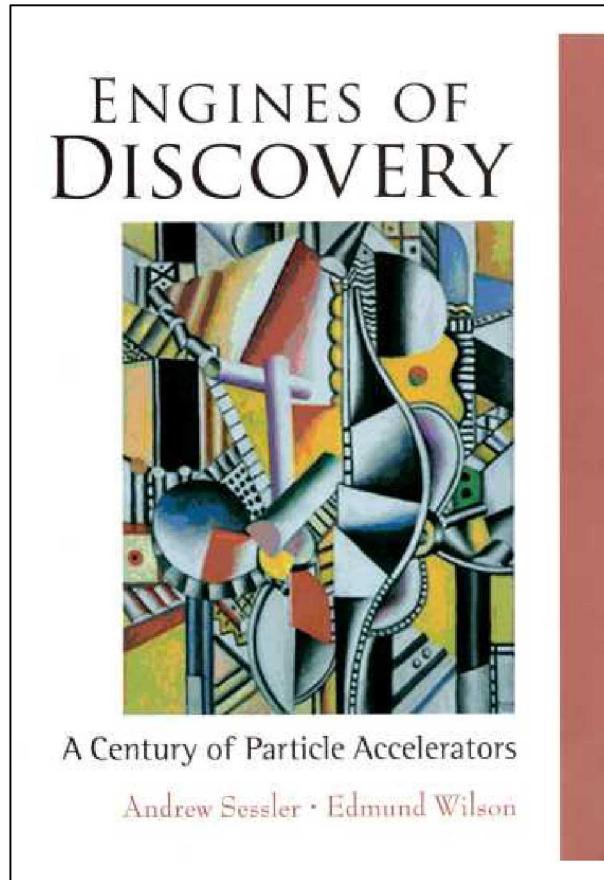
Differences in nTOF shape of secondary DT peak due to magnetization of tritons

# We are exploring a modular architecture that can scale to 300-1000 TW and is twice as electrically efficient as Z



Next-gen machines: 20,000-200,000 bricks, 33-60 cavities/module, and 65-800 modules!

# Pulsed power machines can be “engines of discovery” for HED science, just as particle accelerators have been – the future is exciting!



- Classical particle accelerators have numerous applications to fundamental understanding of the universe and structure of matter, to probing of matter, and to industrial and medical applications.
  - Induced radioactivity, Isotope enrichment, antiproton, nuclear structure,  $J/\psi$  (quarks), tau lepton,  $W, Z$  particles, top quark, Higgs boson
  - Industrial (hardening, sterilization)
  - Medical use (imaging, therapy)
  - Research (x-ray and particle probes)
  - 14 Nobel prizes for accelerators or using accelerators
- Pulsed power accelerators are engines of discovery for HED science
  - Intrinsic material properties (EOS, conductivity, strength, structure of materials)
  - Radiation transport, atomic physics, opacity
  - Magnetized plasma physics
  - Fusion ignition (“chief unsolved problem in plasma physics”)

# The Z Fundamental Science Program engages a broad community and has advanced HED science



- **10 teams won shots on the 18-19 allocation**
  - Carnegie Institution of Washington
  - Lawrence Livermore National Laboratory
  - Northwestern University
  - Sandia National Laboratories
  - UC Davis/ Harvard
  - University of Rostock, Germany
  - UN Reno
  - UT Austin x 2
  - Washington State University
- **12+ students are currently involved**
  - Former students have done very well
- **Resources over 8 years**
  - 100 dedicated ZFSP shots (5-7% of all Z shots)
  - Ride-along experiments on Z program shots, guns, DICE, and THOR
- **Science with far-reaching impact**
  - Nature, Nature Geoscience, SCIENCE
  - 6 Phys. Rev. Lett, 3 Physics of Plasmas, 5 Physical Review (A,B,E)
  - About 40 total peer reviewed publications and 10 conference proceedings
  - 70+ invited presentations
- **Popular outreach**
  - National Public Radio, "All things considered", 2014
  - Discover Magazine
    - Reportage 9/16/2012
    - *Iron rain #62 in top 100 Science stories in 2015*
  - Albuquerque Journal Front Page 9/2017
  - Twice local TV coverage on planetary science

# The ZFSP greatly benefits Sandia's and NNSA mission on both short- and long term



- **Direct recruiting**
  - Rick Kraus (Ph.D. Harvard) – Lawrence Fellow, now staff at LLNL
  - Ross Falcon (Ph.D. UT Austin) – postdoc at Sandia
  - Taisuke Nagayama (Ph.D. UN Reno) – staff at Sandia
  - Mark Schoeble (Ph.D. UT Austin) – staff at Sandia
- **Growth in the HED science community**
  - HEDLP funding to Harvard and UT Austin
  - Active participation in the academic community of HED science – attracting new academic partners
- **Direct methods development**
  - The platform for shock experiments developed jointly with Harvard is now our standard setup for science campaign experiments
  - The work on Fe opacity has served an important role for platform development and provides international peer review benefitting research in science campaigns
- **Development of technical staff**
  - An opportunity for Sandia staff to do leading research and participate fully in the international research community