

The 2016 Infrasound Wanaka Balloon Flight: Precursor to extraterrestrial acoustic research?



Jonathan M. Lees and Daniel C. Bowman

Department of Geological Sciences, University of North Carolina at Chapel Hill, NC, USA.
Sandia National Laboratory

Why study Acoustic Waves in the Stratosphere?

- Acoustic waves are present throughout the atmosphere
- Carry information about processes at the solid Earth/Atmosphere interface and the free atmosphere
- Sound bends away from the surface (usually)
- Wave field is directional in most cases
- Acoustic waves transmit energy to the upper atmosphereand can cause distortion in the ionosphere

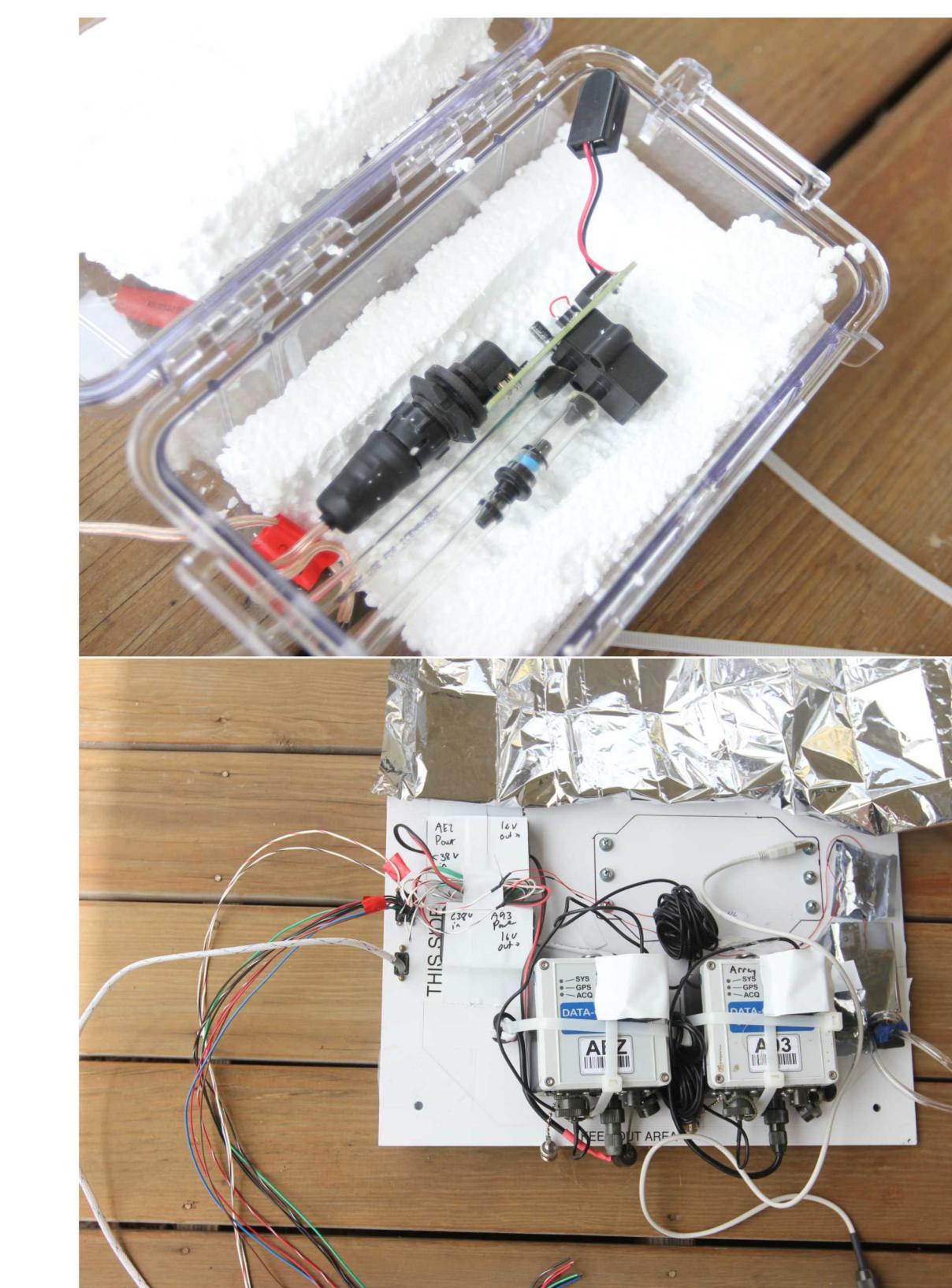
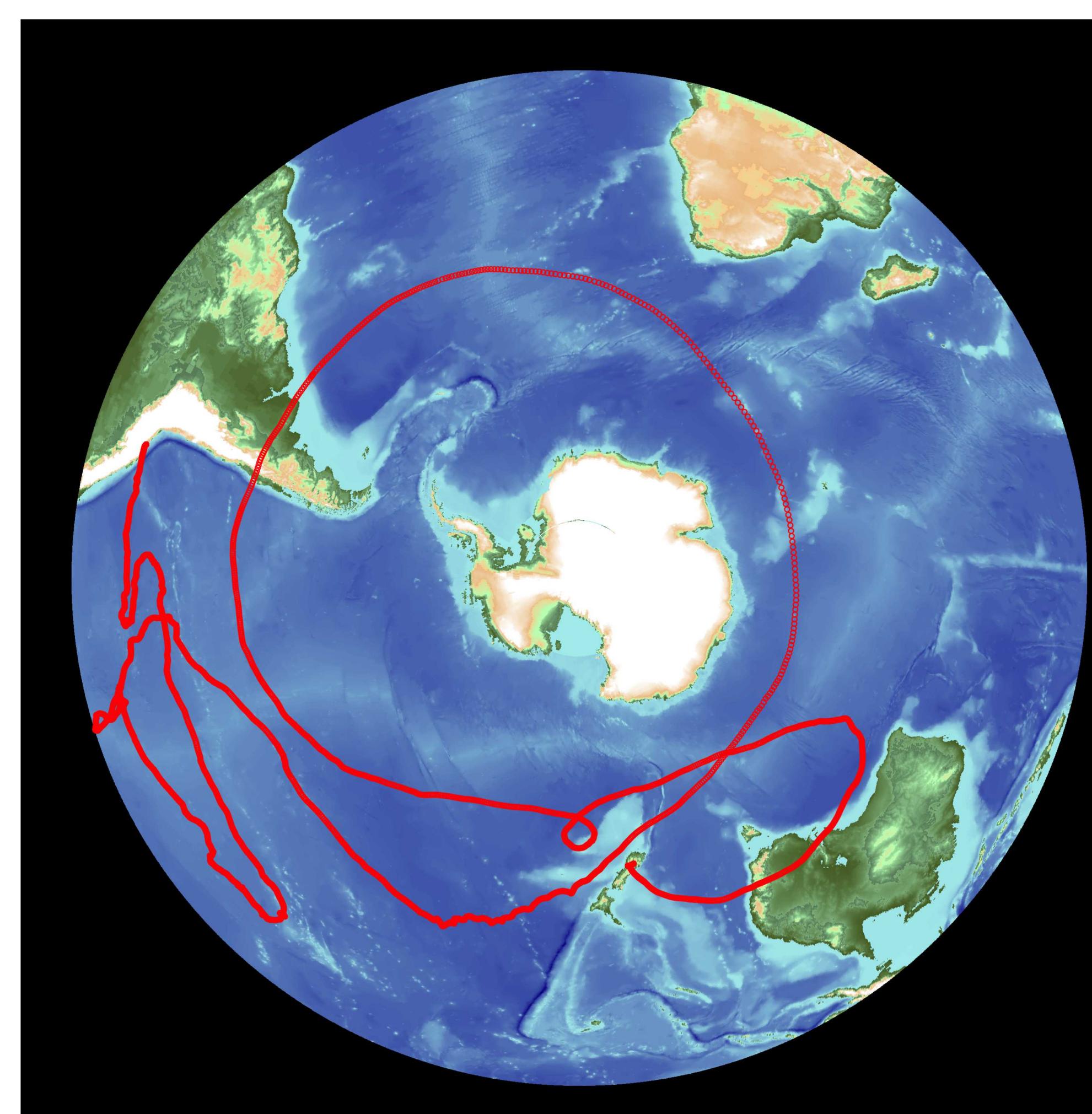
Why Balloons?

- No investigation of acoustic waves above 8 km in over 50 years
- Wind noise pervasive on the surface
- Balloons/Acoustic research → Future of extra-terrestrial planetary exploration

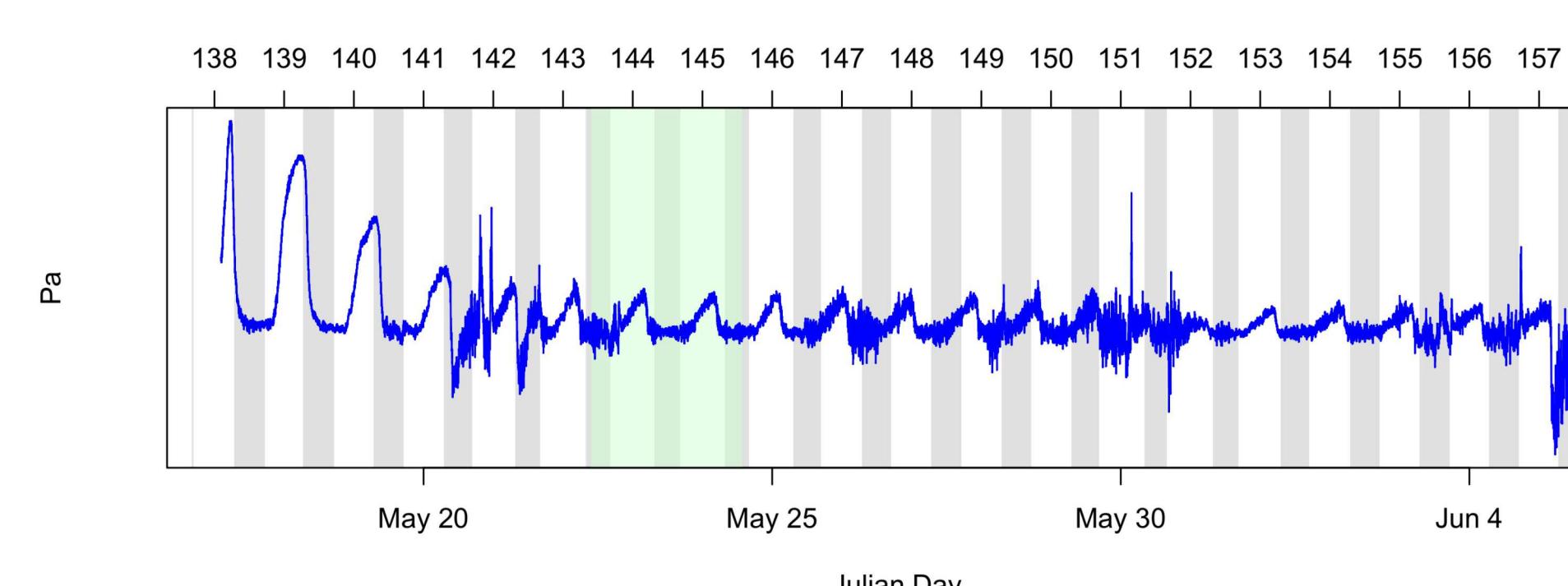
Wanaka Launch, 2016

- Circumnavigation of Antarctic in 2016
- Hover over microbarom source, south of New Zealand
- Southern Ocean suggest a flux of 0.05 mW/m^2 max acoustic energy
- Thermospheric heating: temperature rise of several degrees Kelvin per day
- Lightning strikes, recorded off the coast of New Zealand (Lamb et al. 2018)
- Bolides? Yet to be determined.

Flight Trajectory

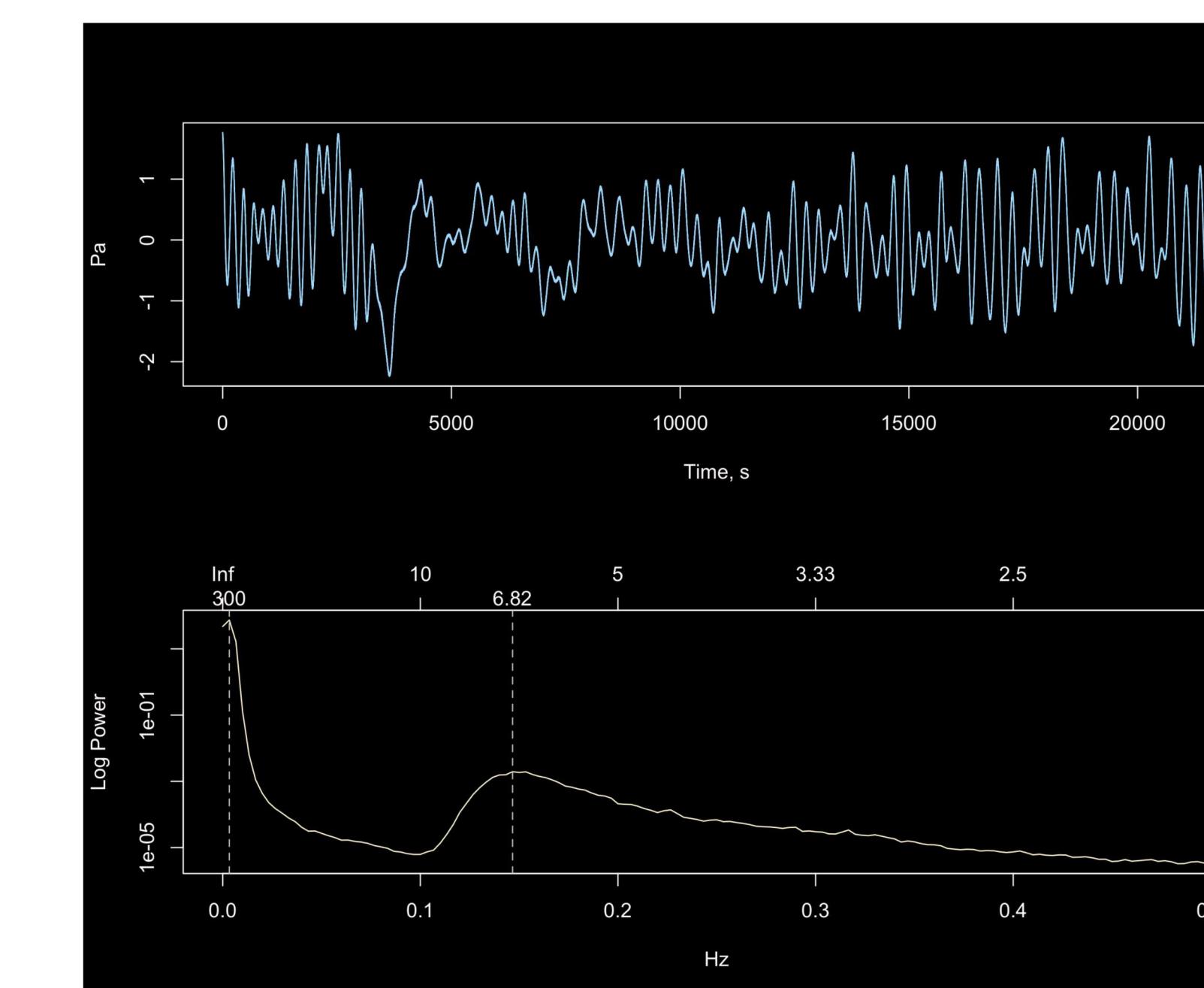


Acoustic Signals, Unfiltered, Full-Flight



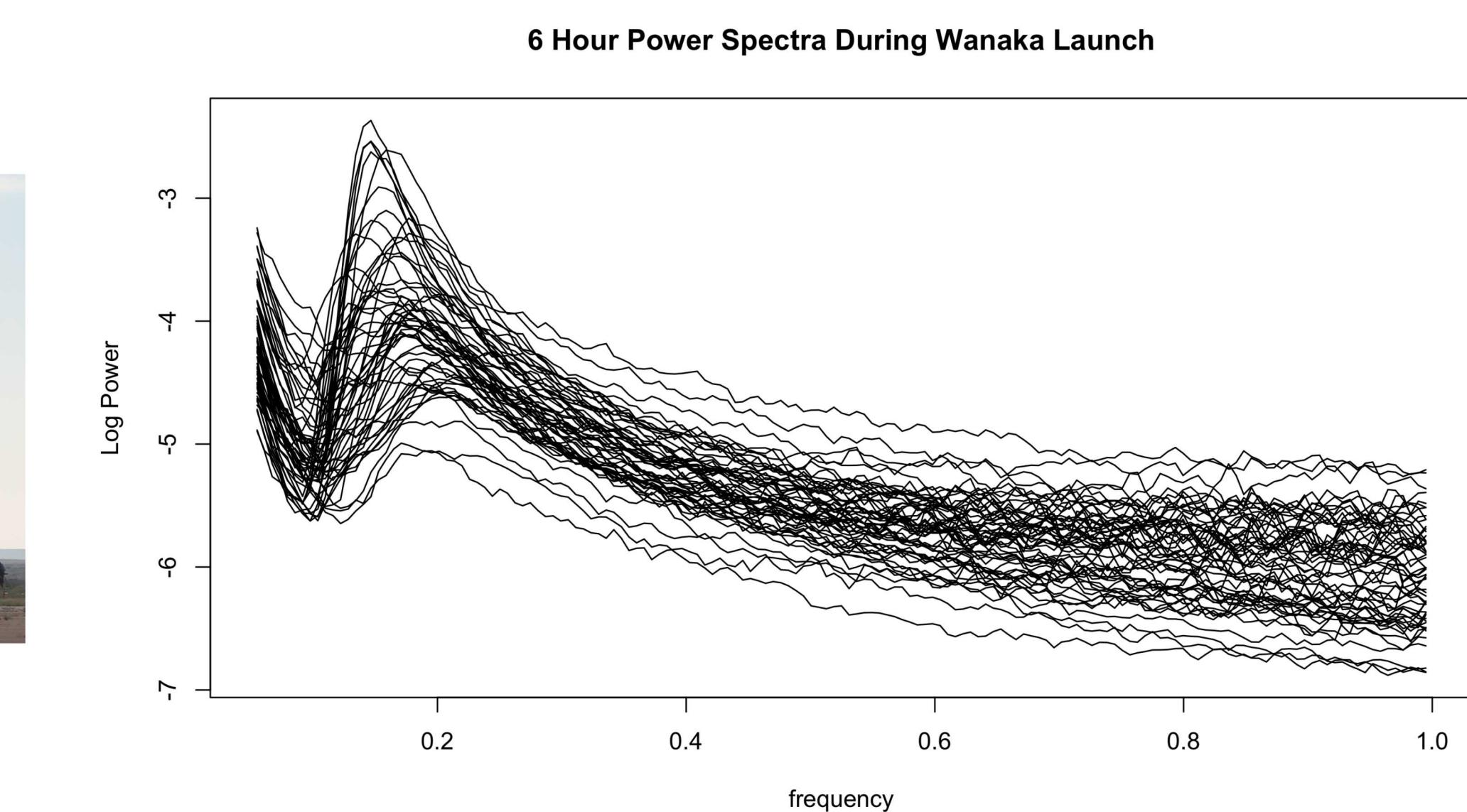
Note Day/Night oscillatory behaviour.

6 Hours, Unfiltered



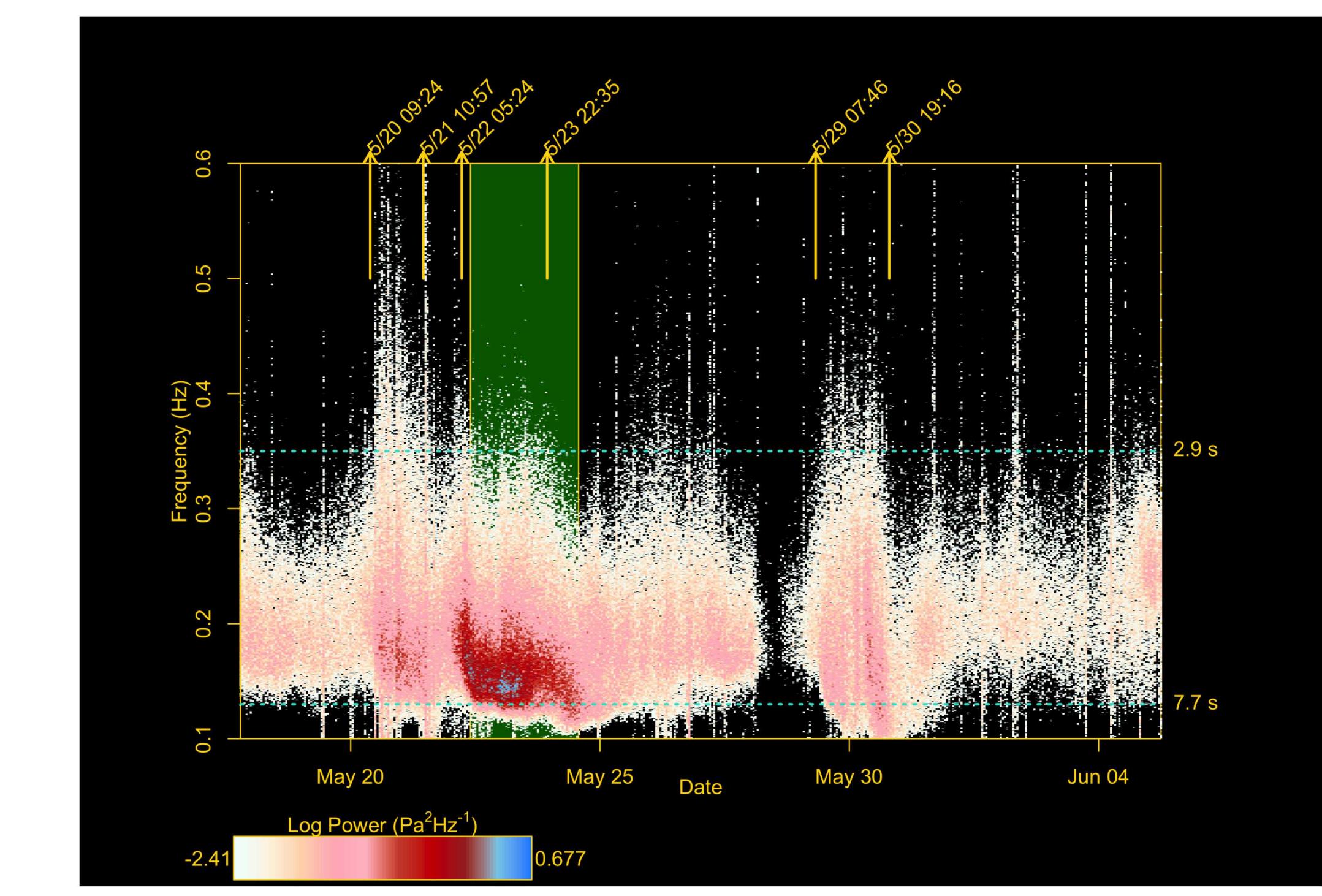
6 Hours extracted during Loop-De-Loop, when the microbarom peaked.

Balloon Low Frequency Acoustic Spectra



What causes variability in microbarom spectra during flight?

Balloon Acoustic Spectrogram



Flux Estimated at Ocean Surface versus Stratosphere

- Acoustic Energy Flux on balloon estimated from frequency spectrum of wave forms:

$$E = \frac{2}{N} \sum_{\omega_a}^{\omega_b} \frac{P \omega_k \omega_k^2}{\rho c}$$

- Flux_{Balloon} = 0.047 mW/m^2
- Estimate acoustic power of ocean microbarom (source):
 - 100 km radius circle
 - sum the Hasselmann integral
- Flux_{Ocean} = 0.052 mW/m^2

Energy Transfer

Acoustic wave dissipation heats the upper atmosphere:

Earlier Work

- 30 K/day heating from microbarom (Rind, 1977)
- 0.03 K from explosions (Drobzheva and Krasnov, 2006)
- $\approx 13 \text{ K/day}$ from thunderstorms (Krasnov et al., 2007)

Bowman and Lees (GRL, 2018)

- Up-going energy flux 0.047 mW/m^2
- 1 to 10 K thermospheric heating rate per day

Planetary Exploration: Balloon Borne Missions

Seismology on gaseous planets: Acoustic Waves on floating platforms



Quote: "This hotspot, by process of elimination, must be heated from below, and this detection is therefore strong evidence for coupling between Jupiter's lower and upper atmospheres, probably the result of upwardly propagating acoustic or gravity waves." O'Donoghue, J. et al., Nature, 2016 <http://dx.doi.org/10.1038/nature18940>

Acknowledgements

Prof. Jeff Johnson (actually, UW; now Boise State U.)
Sharon Kedar, JPL
Nicholas Graham, Marine Weather Consultant
COSI gamma ray telescope
High Altitude Student Platform (HASP)
CEDAR (NSF)
National Science Foundation
Columbia Scientific Ballooning Facility (CSBF)
National Aeronautics and Space Administration (NASA)

Sandia National Laboratories is a multi-mission laboratory managed and operated by National Technology and Engineering Solutions of Sandia, LLC., a wholly owned subsidiary of Honeywell International, Inc., for the U.S. Department of Energy's National Nuclear Security Administration under contract DE-NA0003525. The views expressed here do not necessarily reflect the views of the United States Government, the United States Department of Energy, or Sandia National Laboratories.