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# Tin-lead Alloying for Efficient and Stable All-inorganic Perovskite Solar Cells

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**ABSTRACT:** Cesium containing all-inorganic perovskites have received considerable interest in photovoltaics research because of their potential for improved stability compared to their organic-inorganic hybrid counterparts. However, the inorganic perovskites studied thus far still suffer from lower power conversion efficiency and long-term instability, due to an unfavorable bandgap and either phase instability or air-sensitivity. Herein, A strategy to mitigate these concerns is investigated by alloying tin and lead on the B site to form tin-lead alloyed low-bandgap (~1.34 eV) inorganic CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> perovskites. Solar cells made using this material in an inverted full-structured architecture with a PEDOT:PSS hole transport materials (HTM) attain power conversion efficiency (PCE) up to 9.41% (stabilized PCE 7.23%). Furthermore, a simple HTM-free device without PEDOT:PSS layer is demonstrated more stable than the full-structured device and exhibits a PCE of 7.60% (stabilized PCE 7.31%) – the highest efficiency to date for an inorganic perovskite with a bandgap below 1.40 eV. This simplified device structure shows good reproducibility and stability. This work provides a possible route for fabricating low-cost, high stability devices with competitive efficiencies.

## INTRODUCTION

In recent years, organic-inorganic hybrid halide perovskites have been the subject of much research due to their unique physical and chemical properties (e.g. large optical absorption, high mobility, long diffusion length),<sup>1-3</sup> as well as cost effectiveness and solution processability. As a result, dramatic improvements have been made to the power conversion efficiency (PCE) of the perovskite solar cells from 3.8% to 25.2%.<sup>4-9,10</sup> Although such high performance is beginning to rival commercial Si-based solar cells, the long-term stability of perovskites is still an active area of research as companies look towards commercialization on large scales.<sup>11, 12</sup> The best performing perovskite solar cells usually contain organic cations, such as methylammonium (MA<sup>+</sup>) and formamidinium (FA<sup>+</sup>),<sup>13-18</sup> which are both volatile under elevated temperature and susceptible to moisture-related degradation. One proposed route to address these problems is by replacing the organic cations with inorganic cations, such as cesium (Cs<sup>+</sup>) or rubidium (Rb<sup>+</sup>), to form all-inorganic perovskites;<sup>19-21</sup> these approaches have already demonstrated enhanced thermal stability, moisture resistance and photostability yet generally with lower overall performance than the hybrid counterparts.<sup>22-25</sup>

The first report of functioning all-inorganic perovskites in solar cells was by Snaith et al,<sup>26</sup> who demonstrated a power conversion efficiency (PCE) of 2.9% with cubic CsPbI<sub>3</sub> perovskites. Unfortunately, CsPbI<sub>3</sub> in the black cubic perovskite phase is unstable in ambient atmosphere and

prone to transform into a yellow, orthorhombic non-perovskite phase,<sup>19, 27-31</sup> which is non-photoactive and deteriorates performance of the devices. Several strategies have been proposed to address this issues including solvent engineering,<sup>23, 32</sup> surface/volume ratio modulation,<sup>33, 34</sup> anion substitution, and others.<sup>22, 27, 29, 35-41</sup> Impressive enhancements have been achieved not only on the stability but also on the performance of the devices; very recently, all-inorganic perovskite PCEs of over 13% with improved stability have been reported by several groups.<sup>23, 35, 36, 39</sup> Such progress is encouraging, but the PCE of the inorganic perovskite solar cells are still inferior to that of the organic-inorganic counterparts. One reason for this limitation is that the bandgaps of the stable Cs-based perovskites lie in a range of 1.7-2.3eV, significantly above that for ideal single-junction solar cells (1.1-1.4eV).<sup>22, 24, 28, 29, 35-38, 40-45</sup>

A strategy that has been used very recently to tailor the bandgap for organic-containing perovskites is replacing Pb in the lattice with Sn, which results in a lower bandgap. Sn-based perovskites have been demonstrated with lower bandgap, high charge mobility, and being more eco-friendly than Pb-based perovskites.<sup>46-48</sup> However, fully Sn-based perovskites suffer from their notorious air-sensitivity because the Sn<sup>2+</sup> tends to be oxidized to Sn<sup>4+</sup>.<sup>49, 50</sup> A recent strategy that provides a route around this issue in organic-containing perovskites is alloying Sn and Pb. In organic-inorganic perovskites, tin-lead alloyed perovskites have been shown to have low bandgap but have oxidation stability more similar to Pb than Sn materials.<sup>19, 51-58</sup> Encouraged by these scenarios, we hypothesized that

incorporation of  $\text{Sn}^{2+}$  into an all inorganic lead halide perovskite would broaden the spectral response of inorganic cesium lead iodide and exhibit enhancement of the PCE and stability.

In this work, we prepare all-inorganic tin-lead perovskites by a facile solution process. Alloying Sn and Pb provides a strategy to mitigate the limitations of all-inorganic perovskites; first, tin-lead alloying gives rise to bandgap bowing, allowing us to reach bandgaps that are ideal for PV operation ( $\sim 1.34$  eV). Second, the incorporation of Sn stabilizes the perovskite structure by increasing the tolerance factor of the lattice, which results in superior phase stability. With this composition we attain power conversion efficiencies of 9.41% in an inverted full-structured device architecture, competitive for all-inorganic perovskites, and in a more stable HTM-free architecture we are still able to achieve 7.60% PCE and stabilized 7.31% PCE. These results are both the most efficient HTM-free all-inorganic perovskite, and, most importantly, the most efficient low-bandgap ( $E_g < 1.40$  eV) inorganic perovskite solar cell to the best of our knowledge. Notably, the un-encapsulated inorganic tin-lead perovskite solar cells exhibit good photostability under continuous illumination, long-term stability and the air stability. Tin-lead alloyed all-inorganic perovskite thus represent a promising avenue to stable and efficient devices, which could be used in both single junction and tandem solar cells.

## EXPERIMENTAL SECTION

### Materials:

All materials were used as received without purification as follows:  $\text{PbI}_2$  (Alfa Aesar, 99.999%),  $\text{SnI}_2$  (Alfa Aesar, 99.999%),  $\text{CsI}$  (Sigma-Aldrich, 99.999%),  $\text{SnF}_2$  (Acros Organics, 99%), PEDOT:PSS (Ossila, A14083),  $\text{PC}_{61}\text{BM}$  (Solenne, 99.5%), Bathocuproine (BCP) (Alfa Aesar, 98%),  $N,N$ -Dimethylformamide (DMF) (Sigma-Aldrich, anhydrous, 99.8%), dimethyl sulfoxide (DMSO) (Sigma-Aldrich, anhydrous, >99.9%), chlorobenzene (Sigma-Aldrich, anhydrous, 99.8%), 2-propanol (IPA) (Sigma-Aldrich, anhydrous, 99.5%), anisole (Sigma-Aldrich, anhydrous, 99.7%) and ITO.

### Materials synthesis:

**Tin-lead precursors.** The tin-lead mixed compounds were prepared by dissolving stoichiometric amount of  $\text{SnI}_2$ ,  $\text{PbI}_2$ ,  $\text{CsI}$  for the desired composition in an aqueous mixed solvent (DMF:DMSO = 7:3 v/v) with a small amount of  $\text{SnF}_2$  ( $\text{SnI}_2:\text{SnF}_2=1:0.1$  in molar ratio) at a constant stirring at  $\sim 60$  °C overnight. In a typical  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$ , for example, 322.7 mg  $\text{PbI}_2$ , 117.8 mg  $\text{SnI}_2$ , 4.7 mg  $\text{SnF}_2$  and 259.8 mg  $\text{CsI}$  were mixed in 1 mL mixed solvent (700  $\mu\text{L}$  DMF and 300  $\mu\text{L}$  DMSO) to obtain 1 M precursor. For  $\text{MASn}_{0.3}\text{Pb}_{0.7}\text{I}_3$ , the  $\text{CsI}$  was replaced with MAI counterparts. All the products were filtered with 450  $\mu\text{m}$  PTFE filter before use.

**CsPbI<sub>2</sub>Br precursor.**  $\text{PbI}_2$ ,  $\text{PbBr}_2$  and  $\text{CsI}$  were dissolved in a mixed solvent (DMF:DMSO=7:3 v/v) in the appropriate molar ratios to result a 0.43 M precursor solution of  $\text{CsPbI}_2\text{Br}$ , which was stirred overnight at 150 °C and filtered before use.

### Device fabrication.

The pre-patterned ITO substrates were ultrasonically cleaned in diluted Micro-90 detergent, deionized water,

acetone and 2-propanol bath sequentially, and dried with nitrogen. Prior to thin film deposition, the substrates were treated for 20 min with oxygen plasma or  $\text{UV}/\text{O}_3$ . For inverted full-structured devices, a thin layer of PEDOT:PSS was deposited and then heated at 150 °C for 15 min on a hotplate before transferring into a nitrogen-filled glovebox. For HTM-free devices, the cleaned ITO substrates were transferred into the glovebox for films deposition immediately after oxygen plasma or  $\text{UV}/\text{O}_3$  treatment. The compositional inorganic perovskite precursors were spin-coated onto the substrates at 10000 rpm for 1s with acceleration rate of 11000rpm and 0 rpm for 20 s with acceleration rate of 500 rpm sequentially in the glovebox. The resulted substrates were annealed at 100 °C for 20 min (it depends on the retention time of black films for that with molar ratio of Sn/Pb less than 0.30) on a hotplate. The perovskite-coated substrates were immersed into an anti-solvent anisole solution for 20 s. Afterwards, PCBM (20 mg  $\text{mL}^{-1}$  in anhydrous chlorobenzene) and BCP (0.5 mg  $\text{mL}^{-1}$  in anhydrous IPA) were sequentially spun on the substrates at 2000 rpm for 20 s and 6000 rpm 15s, respectively. Finally, silver electrode ( $\sim 100$  nm) was deposited by thermal evaporation to complete the fabrication.

### Device characterization.

The photocurrent density-voltage curves and stabilized power output of the devices were measured in the glovebox using a Keithley 2400 source meter under AM 1.5G solar illumination at 100  $\text{mW cm}^{-2}$ . External quantum efficiency (EQE) measurements were performed in ambient environment by loading the samples in a vacuum chamber using a xenon arc lamp and a monochromator, which the light intensity was calibrated with a standard Si photodiode. For devices stability testing, all the samples were transferred out from the glovebox and exposed in air without encapsulation.

**X-ray diffraction (XRD).** XRD characterization was carried out on perovskite thin films on the glass substrates using a Bruker D8 Advance powder diffraction meter with  $\text{Cu K}\alpha$  radiation under operation conditions of 40 kV and 44 mA.

**X-ray/ultraviolet photoelectron spectroscopy (XPS/UPS).** The sample for XPS/UPS was obtained by spin-coating the precursor solution on ITO substrate. The XPS measurements were carried out in an ultrahigh vacuum of  $\sim 6 \times 10^{-11}$  mbar using a Kratos AXIS Ultra DLD. The sample was transferred into the vacuum chamber as soon as possible and excited with X-rays from a monochromatic Al  $\text{K}\alpha$  source. High-resolution XPS spectra were fitted with CasaXPS package. UPS characterization was performed in the same chamber as for XPS using a He 1  $\alpha$  source at 21.22 eV. The electron binding energy scale was calibrated using the Fermi edge of clean gold.

**UV-Vis absorption.** Optical absorption spectra were recorded for perovskite thin films deposited on glass substrates in ambient environment using Agilent 8453 Diode Array UV-Vis spectrophotometer. Sample temperature is controlled from within ChemStation software via an external temperature controller which drives a Peltier heat pump attached to the sample holder. For the films with molar ratio of Sn/Pb less than 0.30, the UV-vis absorption spectra were also obtained using an AvaSpec-ULS2048XL-EVO spectrometer in the glovebox.

**Photoluminescence and time resolved photoluminescence (PL/TRPL).** The steady-state photoluminescence of the samples was performed by placing the sample in a vacuum chamber. Samples were photoexcited using a 450 nm light-emitting diode (LED) and the signals were recorded using a home-built setup based on an Ocean-Optics 2000+ spectrometer. TRPL spectra were acquired using a time-resolved single photon-counting set-up (FluoTime 300, PicoQuant).

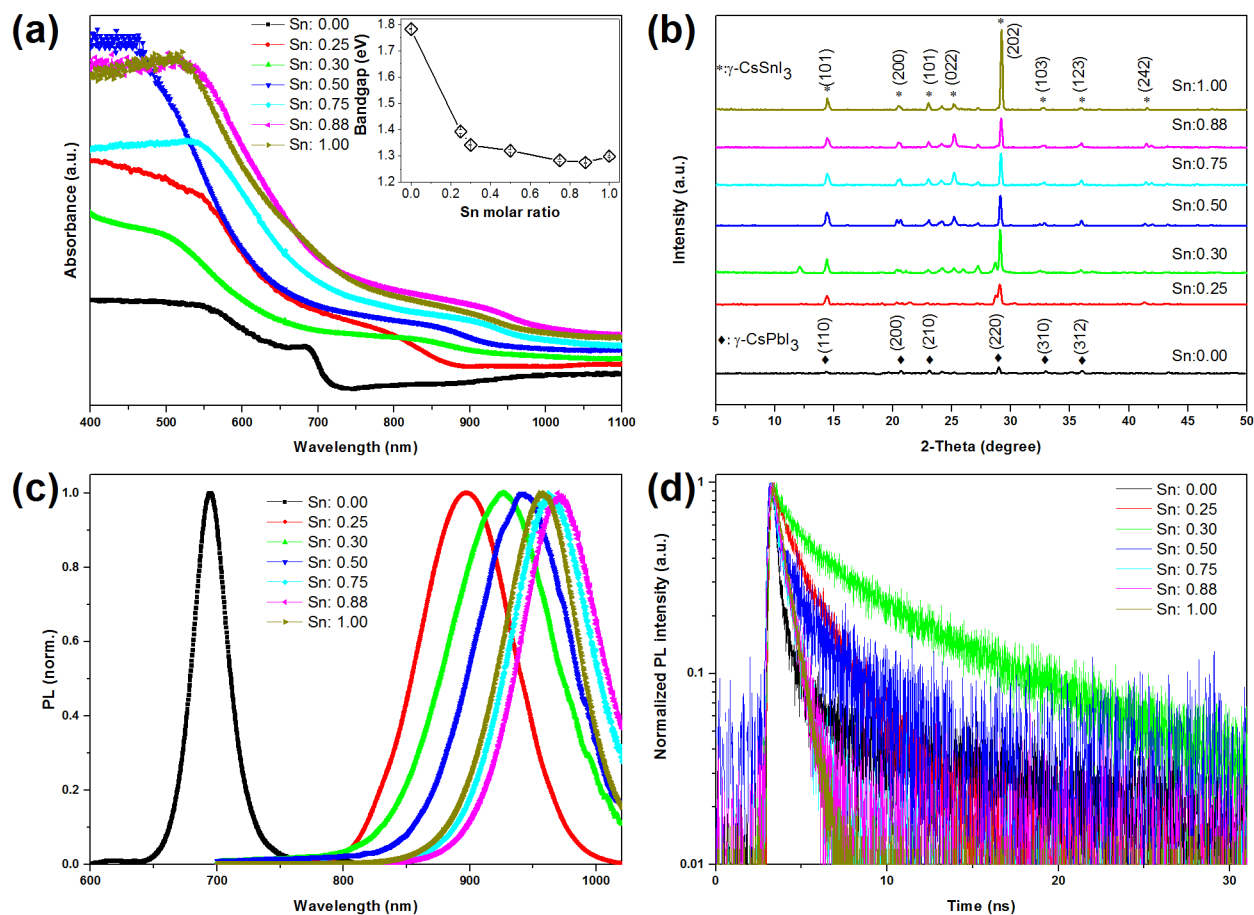
**Scanning electron microscopy (SEM).** SEM was performed on a FEI Sirion XL30 microscopy with a field emission electron source.

**Glow Discharge Optical Emission Spectrometer (GDOES).** The composition and the depth profile of inorganic perovskite films were analyzed using GDOES technology. The samples were prepared by spun perovskite precursor solution on a clean glass substrate. The GDOES was performed at room using a Horiba GD-Profilier-2 equipped with a differential interferometry profiling system temperature.

## RESULTS AND DISCUSSION

Tin-lead alloyed perovskites with various Sn:Pb ratios were prepared by spin-coating from stoichiometric solutions in DMF/DMSO mixed solvents, and subsequently annealed as described in the experimental section. To explore the effect of Sn incorporation on the bandgap tunability, UV-vis absorption of the films with a range of compositions (Sn molar ratio in tin-lead compounds) was performed. Figure 1a presents optical absorption spectra of the tin-lead films with Sn molar ratio from 0.00 to 1.00. We observe that alloying of Sn and Pb in tin-lead compounds results in bandgap bowing, which tunes the bandgap between 1.78 eV

and 1.30 eV (inset in Figure 1a) with increase of Sn ratio from 0.00 to 1.00 except for anomalous bandgap behavior shrinking to 1.28 eV and 1.27 eV observed in two of them, namely 0.75 and 0.88, respectively, similar to that reported in tin-lead based organic-inorganic perovskites.<sup>51</sup> This bowing is possibly due to local distortions of the short range ordering of the tin-lead cations. It is worth mentioning that the ideal bandgap for single-junction photovoltaic devices is between 1.1 eV and 1.4 eV according to the theoretical calculation,<sup>44, 59</sup> beyond the range of most reported all-inorganic perovskites recently. Tin alloying strategy, however, could easily tune the bandgap of the all-inorganic CsPbI<sub>3</sub> perovskites from 1.78 eV to the ideal value for photovoltaics. More interestingly, it allows the alloyed perovskite films more stable than the original one, which can be observed simply in the light of transforming time from black phase to yellow phase of the films. Typically, the black CsPbI<sub>3</sub> film that obtained at 100 °C turned to yellow in seconds when it exposed to ambient atmosphere and based on this background, the absorption spectra of CsPbI<sub>3</sub> shown in Figure 1a was measured in a nitrogen-filled glovebox, which is quite different from that of obtained in air (Figure S1). With the incorporation of Sn, the color of the film with 0.25 molar Sn is able to remain unchanged for several minutes in air. Notably, when the Sn molar ration increases to 0.3, the perovskite film could be kept in air for several hours without striking change and its' optical wavelength extend to infrared range. Based on a promising bandgap, air stability (in terms of not discoloring rapidly), and tin % below 50%, which has been shown to be important for oxidation stability,<sup>60</sup> this tin-lead alloyed strategy is therefore promising for efficient and stable photovoltaic devices.



**Figure 1.** (a) UV-Vis absorption spectra, inset shows the corresponding bandgap (b) XRD patterns, (c) steady-state PL spectra and (d) time-resolved PL curves of tin-lead alloyed perovskite thin films with different Sn molar ratio ranging from 0.00 to 1.00.

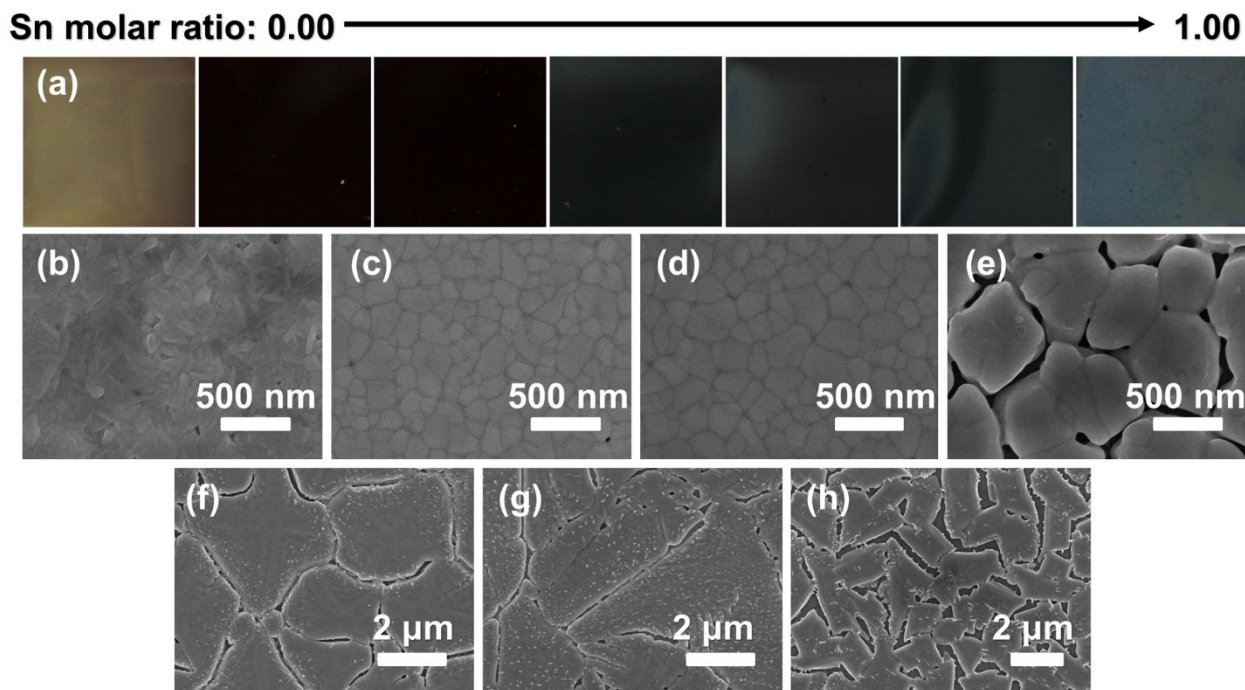
To identify the effect of Sn incorporation on the lattice structure of the perovskites, we performed the x-ray diffraction measurements (XRD) at room temperature in ambient condition. Figure 1b shows x-ray diffraction (XRD) patterns of tin-lead compounds for Sn molar ratio ranging from 0.00 to 1.00. For the film without Sn incorporation and with 0.25 molar ratio of Sn, which are found to discolor in seconds and in minutes in air, respectively, we seal the samples using Kapton tapes<sup>61</sup> in the glovebox to minimize exposure to air and in fact, the XRD patterns of these two samples are quite different from the results without sealing (Figure S2). As shown in Figure 1a, the sealed film without Sn content exhibits the characteristic peaks at 14.37°, 20.69°, 23.01°, 29.02°, 32.87°, and 36.02° and by carefully checking, these peaks can be assigned to the (110), (200), (210), (220), (310) and (312) planes of orthorhombic  $\gamma$ -CsPbI<sub>3</sub> structure<sup>31, 62-65</sup> while the XRD patterns of unsealed CsPbI<sub>3</sub> sample can be indexed to the non-perovskite orthorhombic  $\delta$ -CsPbI<sub>3</sub> structure (Figure S2) due to discoloring during the measurement.<sup>26, 41</sup> Notably, although the CsPbI<sub>3</sub> sample which is obtained at 100 °C pose large bandgap of 1.78 eV, the black appearance and even very close XRD peaks to that obtained at 330 °C, it is actually corresponds to orthorhombic  $\gamma$ -CsPbI<sub>3</sub> rather than cubic  $\alpha$ -CsPbI<sub>3</sub> phase incorrectly as most of the previously reported XRD patterns.<sup>63, 64</sup> For the film with 100% Sn, we index the diffraction peaks at 14.44°, 23.01°, 25.16° and 29.20° to the

(110), (120), (202) and (220) planes of the black orthorhombic (B- $\gamma$ ) CsSnI<sub>3</sub> perovskite.<sup>46</sup> The XRD patterns in the as-prepared thin films with Sn content between these two end members indicate either single-phase tin-lead alloys or the mixed phase of Sn- and Pb-based perovskites. Regardless the concrete phase, it has been found that the XRD patterns of the unsealed Sn-Pb alloyed perovskites that Sn content are larger than 0.30 retain to be identical between repeated measurements, suggesting that incorporation of Sn in Pb-based perovskite substantially improves the stability of photoactive black phase at room temperature as that of reported FA<sup>+</sup> or metal cation incorporation.<sup>36, 66, 67</sup> Note that the film with 0.30 Sn in molar ratio exhibits very similar XRD patterns to that in  $\gamma$ -CsPbI<sub>3</sub> and B- $\gamma$  CsSnI<sub>3</sub> perovskites, accompanied by some new reflection peaks at 12.12°, 21.16° and 28.68° which differ from any phases of CsSnI<sub>3</sub> and CsPbI<sub>3</sub> and cannot be indexed to any diffraction peaks of the possible undissolved source materials, namely PbI<sub>2</sub>, CsI, SnF<sub>2</sub> or SnI<sub>2</sub>/SnI<sub>4</sub>.<sup>34</sup> The result suggests that the crystal structure might be in the following possible forms: black orthorhombic (B- $\gamma$ ) CsSnI<sub>3</sub>; a mixture of black orthorhombic (B- $\gamma$ ) CsSnI<sub>3</sub> and  $\gamma$ -CsPbI<sub>3</sub>; or a new tin-lead alloyed structure or its mixture (as discussed later).

To gain more insight about the Sn incorporation effect on the optoelectronic properties of inorganic halide perovskite films, we carried out steady-state and time-resolved photoluminescence (TRPL) measurements of tin-lead

alloyed perovskite onto glass substrates, which are shown in Figure 1c and Figure 1d, respectively. As shown in Figure 1c, for every individual composition of the tin-lead films, sharp PL peaks are observed and with the increase of Sn molar ratio in tin-lead alloys, corresponding PL peaks exhibited a redshift step by step except that the emission peaks with Sn ratio of 0.75 and 0.88 are beyond the end member, namely  $\text{CsSnI}_3$ , toward to redshift, in good agreement with absorption onset (Figure 1a). The single PL peak change, continuous absorption onset and corresponding XRD peaks (Figure 1b) strongly suggest that we can rule out the possibility of single  $\text{B-}\gamma$   $\text{CsSnI}_3$  or its mixture with  $\gamma$ - $\text{CsPbI}_3$  perovskites in the films with Sn ratio between two end members and confirm that these products are in new tin-lead crystal structure, i.e. tin-lead alloyed perovskites with single-phase. It is worth mentioning that the PL spectrum of  $\text{CsPbI}_3$  film, which was measured in a pumping container against discoloring, exhibits band edge emission at  $\sim 696$  nm, corresponding to a bandgap of 1.78 eV. This result is consistent with the absorption shown in Figure 1 and in good agreement with that of the most recently reported  $\gamma$ - $\text{CsPbI}_3$  by Hu et al.<sup>63</sup> Figure 1d shows

the TRPL spectra of these tin-lead perovskite films with different Sn molar ratio prepared on glass substrates, and corresponding parameters fitted with biexponential function are given in Table S1, in which two dynamics, a fast decay component  $\tau_1$  and slow decay part  $\tau_2$  are mainly attributed to bimolecular recombination of photogenerated free carriers and trap-assisted recombination,<sup>68-70</sup> respectively. The pristine film without Sn introduce,  $\text{CsPbI}_3$ , and the films with Sn proportion larger than 50% present negligible slow components and the dynamics of them can be traced to exponential function, indicating low trap density in these films. Notably, the calculated  $\tau_2$  of  $\text{CsSn}_{0.25}\text{Pb}_{0.75}\text{I}_3$  and  $\text{CsSn}_{0.30}\text{Pb}_{0.70}\text{I}_3$  films are 2.88 and 7.65 ns, and the average lifetime of these films are evaluated to be 1.28 and 3.90 ns, respectively. The results might be associated with the trap-assisted recombination, in which an electron falls into a "trap", an energy level within the bandgap caused by the presence of the foreign Sn and either occurs in a second-step falling into an empty in the valence band (Sn molar ratio <50%) to complete the recombination process or directly annihilates in over-formed Sn vacancies (Sn molar ratio >50%).



**Figure 2.** (a) Optical images and top-view of SEM images of tin-lead alloyed thin films with Sn molar ratio of (b)0.00, (c)0.25, (d) 0.30, (e) 0.50, (f) 0.75, (g) 0.88 and (h) 1.00 .

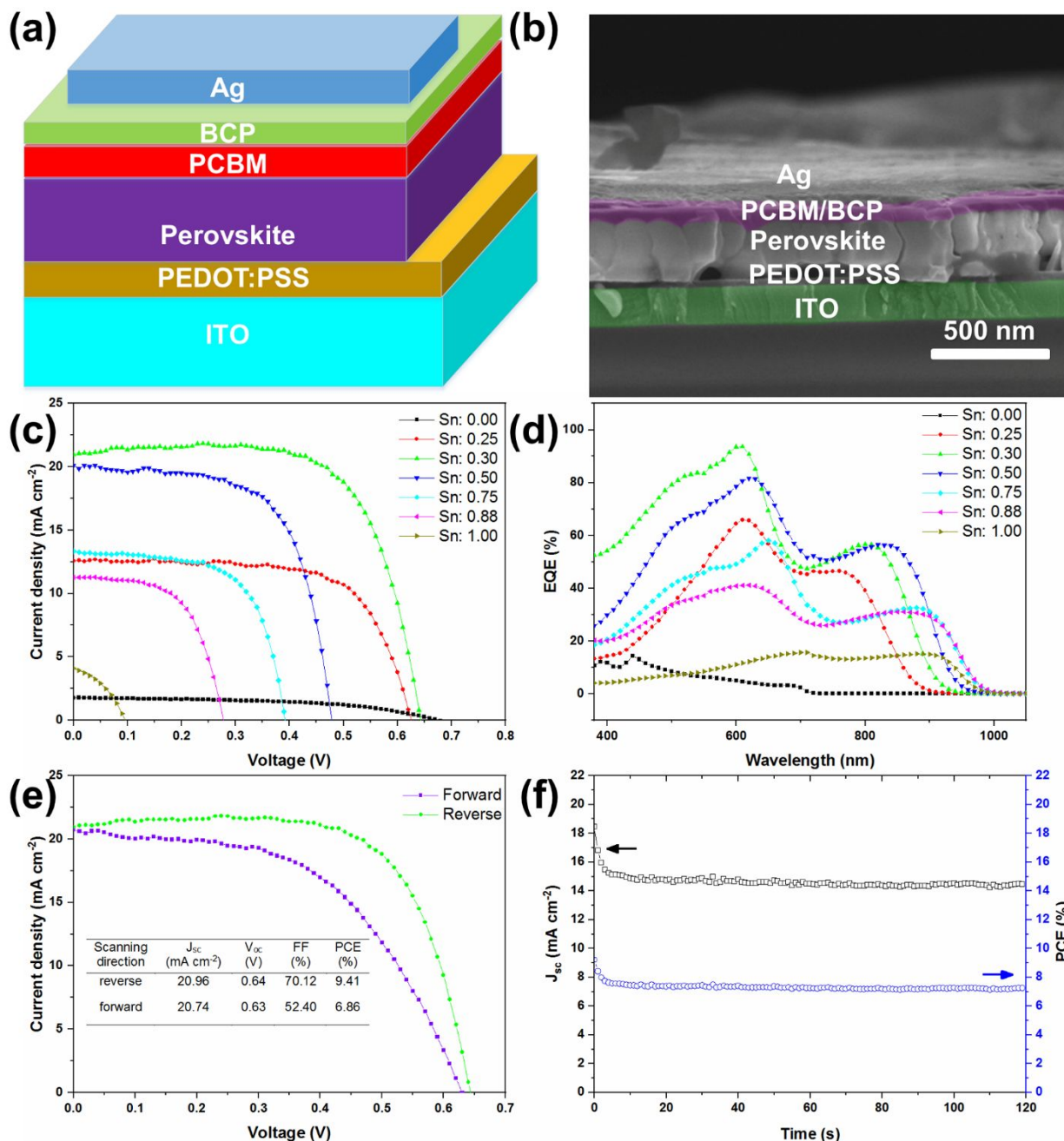
To investigate the effect of the Sn on the change of morphologies, which are prepared using a precursor-phase anti-solvent immersion (PAI) route,<sup>49,51</sup> optical and top-view of scanning electron microscopy (SEM) images of post-annealed tin-lead alloyed films were performed (Figure 2a and Figure 2b~h). From SEM images, we found that the Sn fraction in the precursor source has a crucial influence on the apparent grain size. The pristine  $\text{CsPbI}_3$  films without Sn incorporation exhibit small crystal domain size about  $\sim 200$  nm (Figure 2b). As Sn fraction increase from 0.25 to 1.00, the average domain size increase from  $\sim 300$  nm to several micrometers (Figure 2c~h). However, large

nanoscale domains in the films with higher Sn fraction are accompanied by a significant amount of pinhole formation (Figure 2e~h), which might result in a detrimental effect on the performance of devices. In contrast, the films with Sn ratio of 0.25 and 0.30, especially 0.30, displays smoother surface and good coverage with few pinholes (Figure 2c, d). The corresponding color (Figure 2a) of the tin-lead films gradually changed from orange to deep black as the Sn fraction increase from 0.00 to 0.30, whereas further increase Sn fraction, it gradually turned from dark to slight greyish white. It should be noted that the decent black color, uniform and densely coverage surface has been obtained in

1 the film with Sn fraction of 0.30 (Figure 2c), which would  
2 play an important role in the photoelectrical performance  
3 of perovskite devices.

4 To evaluate the feasibility of the tin-lead alloyed absorber  
5 in photovoltaic devices, we fabricated inverted planar  
6 perovskite solar cells with a structure ITO/Poly(3,4-  
7 ethylenedioxythiophene) poly(styrenesulfonate)  
8 (PEDOT:PSS)/perovskite/Phenyl-C61-butyric acid methyl  
9 ester (PCBM)/bathocuproine (BCP)/Ag as shown in Figure  
10 3a. Figure 3b shows cross-sectional SEM image of a typical  
11 devices based on tin-lead thin films with Sn ratio of 0.3, in  
12 which all layers are obviously observed corresponding to  
13 schematic structure layer by layer and the perovskite  
14 absorber shows good crystallinity with ~300 nm in  
15 thickness. The photocurrent density-voltage (J-V) curves as  
16 a function of the molar ratio of Sn in the tin-lead alloyed  
17 perovskite solar cells are shown in Figure 3c, and  
18 corresponding photovoltaic parameters are given in Table  
19 1. As shown in Figure 3c, the pure Pb based perovskite solar

cell exhibits the highest open circuit voltage ( $V_{oc}$ ) of ~0.67  
V but shows a very low short circuit current density ( $J_{sc}$ ) of  
1.8 mA cm<sup>-2</sup>, thus leading to a low PCE of 0.61%. In contrast,  
the lowest  $V_{oc}$  of 0.10 V, a  $J_{sc}$  of 4.07 mA cm<sup>-2</sup> corresponding  
to a PCE of ~0.15% in the pure Sn based perovskite solar  
cell. Substituting Pb with Sn partially, the  $V_{oc}$  varies between  
these two endpoints which can be attributed to the  
variation of the decreasing bandgap of the tin-lead  
absorbers with increase of Sn proportion. In stark contrast  
to  $V_{oc}$ , the tendency of  $J_{sc}$ , Fill factor (FF) and PCE are ascend  
at first and then descend at last as Sn percentage increases.  
We achieved the highest PCE up to 9.41% with a  $J_{sc}$  of 20.96  
mA cm<sup>-2</sup>, FF of 70.12% and  $V_{oc}$  of 0.64 V, using the  
CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> perovskite film as an active layer in solar cell.  
This impressive PCE mainly determined by the enhanced  $J_{sc}$   
and decent FF, which possibly come from the nature of low  
bandgap (~1.34 eV), long lifetime and the densely coverage  
of the CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> perovskite film.



**Figure 3.** (a) Schematic structure, (b) typical cross-sectional SEM images, (c) photocurrent density-voltage curves (J-V curves) and (d) EQE spectra of tin-lead based perovskite solar cells with Sn molar ratio ranging from 0.00 to 1.00; (e) the J-V curves with forward scanning and reverse scanning and (f) stabilized current density and PCE of the best performing full-structured perovskite solar cells.

Figure 3d shows the external quantum efficiency (EQE) spectra of tin-lead alloyed perovskite solar cells. Obviously, EQE onset wavelength of the tin-lead perovskite solar cells shows a gradual redshift tendency with the increase of the Sn content in tin-lead alloys with an exception of CsSnI<sub>3</sub>, in good agreement with the UV-vis absorption and PL spectra (Figure 1a, c). For the best performing devices based on CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> perovskite film, the EQE in the peak larger than 90% has been obtained, superior to that of most reported all-inorganic perovskite solar cells. To check the reliability of the champion solar cell, the J-V scans with

forward and reverse direction and stabilized power output (SPO) at maximum power point were measured under continuous illumination of AM1.5G (Figure 3e, f). We note that CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> based full-structured perovskite solar cell presents a PCE of 9.41% with reverse scanning while has only a PCE of 6.86% (table inset in Figure 3e). As shown in Figure 3f, the steady-state current density dropped quickly at first 15 s and gradually stabilized at 14.6 mA cm<sup>-2</sup>, which induced a similar tendency in steady-state PCE decreasing from 9.22% to 7.23%. It should be noted that CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> based perovskite solar cells in this device configuration

exhibit pronounced hysteresis in the J-V scanning measurements with a hysteresis index<sup>71</sup> about 21.58%. It is now understood that hysteresis in perovskite solar cells is likely related to both the presence of mobile ionic defects and their interaction with the p- and n- type interfaces.<sup>72</sup> Snaith et al. demonstrated the specific device architecture affects the severity of this behavior in perovskite solar cells.<sup>73</sup> Others<sup>74, 75</sup> have also pointed out the PEDOT:PSS in the inverted planar structure solar cells may induce stability issues due to the poor chemical stability of

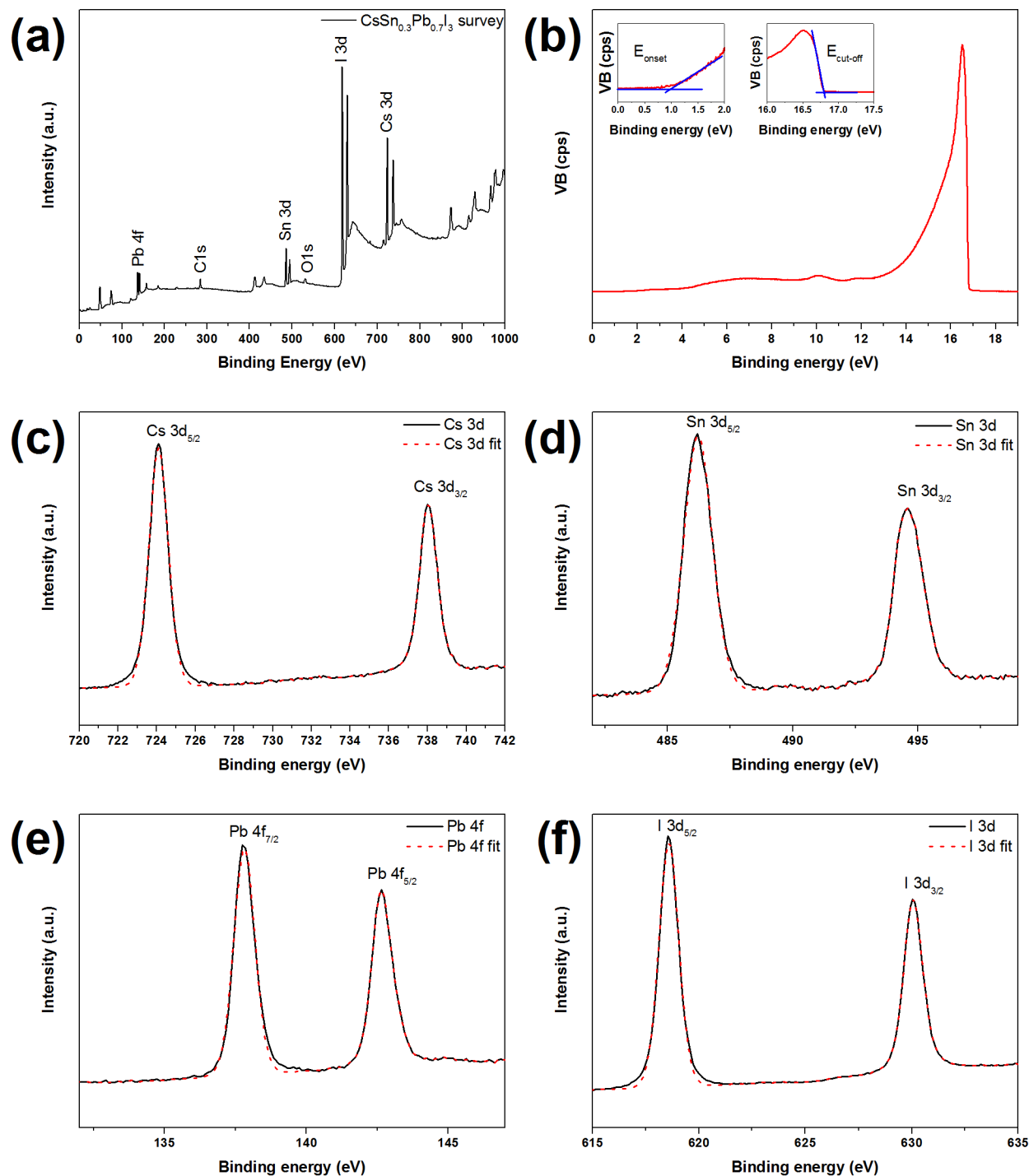
PEDOT:PSS: the reaction at the interface between the active layer and PEDOT:PSS has been demonstrated to be a major cause of device failure in inverted devices.<sup>76, 77</sup> We thus hypothesize that PEDOT:PSS might be a non-ideal contact here and we found that, remarkably, we could in fact attain effective and less hysteresis devices by removing the hole transport layer entirely, following the demonstration of effective hole-transport-material-free (HTM-free) Sn-based perovskite solar cells.<sup>46</sup>

**Table 1.** Key photovoltaic parameters derived from the J-V curves of fully-structured perovskite solar cells in Figure 3c with different Sn molar ratio in tin-lead compounds.

Sn molar ratio	$J_{sc}$ (mA cm <sup>-2</sup> )	$V_{oc}$ (V)	FF (%)	PCE (%)
0.00	1.78	0.67	50.88	0.61
0.25	12.58	0.63	67.39	5.34
0.30	20.96	0.64	70.12	9.41
0.50	20.10	0.48	64.44	6.22
0.75	13.34	0.39	64.05	3.33
0.88	11.28	0.28	58.71	1.85
1.00	4.07	0.10	36.53	0.15

Prior to fabricate the HTM-free devices based on the thin films with different Sn molar ratio, which are designated with the stoichiometry ratio in precursor solution, we conduct the X-ray photoelectron spectra (XPS) and ultra-violet photoelectron spectra to identify the atomic composition and electronic structure of the samples. A typical survey spectrum of the perovskite film with Sn ratio of 0.3 (Figure 4a) identify the presence of both Sn and Pb, compared to the only metal Sn in CsSnI<sub>3</sub> and Pb in CsPbI<sub>3</sub> (Figure S3). High-resolution XPS spectra of the Cs 3d, Sn 3d, Pb 4f and I 3d were performed, which can distinguish the oxidation state of the corresponding elements (Figure 4c~f). The characteristic peaks of Sn 3d observed at 486.2 eV and 494.6 eV with a peak splitting of 8.4 eV which is assigned to Sn<sup>2+</sup> 3d<sub>5/2</sub> and 3d<sub>3/2</sub>, respectively, and distinct from that of Sn<sup>4+</sup> and Sn<sup>0</sup> suggesting no oxidation or disproportionation has taken place here. The narrow doublet peaks of Pb 4f appear at 137.8 eV and 142.6 eV with

a separation of 4.8 eV, indicating the Pb is also in the Pb<sup>2+</sup> state as expected. The chemical composition from XPS spectra shows that the atomic ratios of the elements of Sn and Pb in the present film is about 2.95 and 7.14, respectively, which are very close to the 3:7 in the precursor solutions. The elemental distribution was also confirmed by glow discharge optical emission spectroscopy (GDOES) analysis, which probes the chemical composition as a function of sputtering time (film depth) (Figure S4). As shown in Figure S4, the atomic ratio of Sn and Pb from surface to depth etched is consistent with the XPS measurement suggesting Sn and Pb coexist and the distribution of them is fairly homogeneous through the cross-section layer of the film. Figure 4b shows the typical UPS of the film with Sn ratio of 0.30, in which the valence band maximum (VBM) is determined to be 5.44 eV and corresponding conduction band minimum (CBM) was estimated to be 4.10 eV based on the optical bandgap.



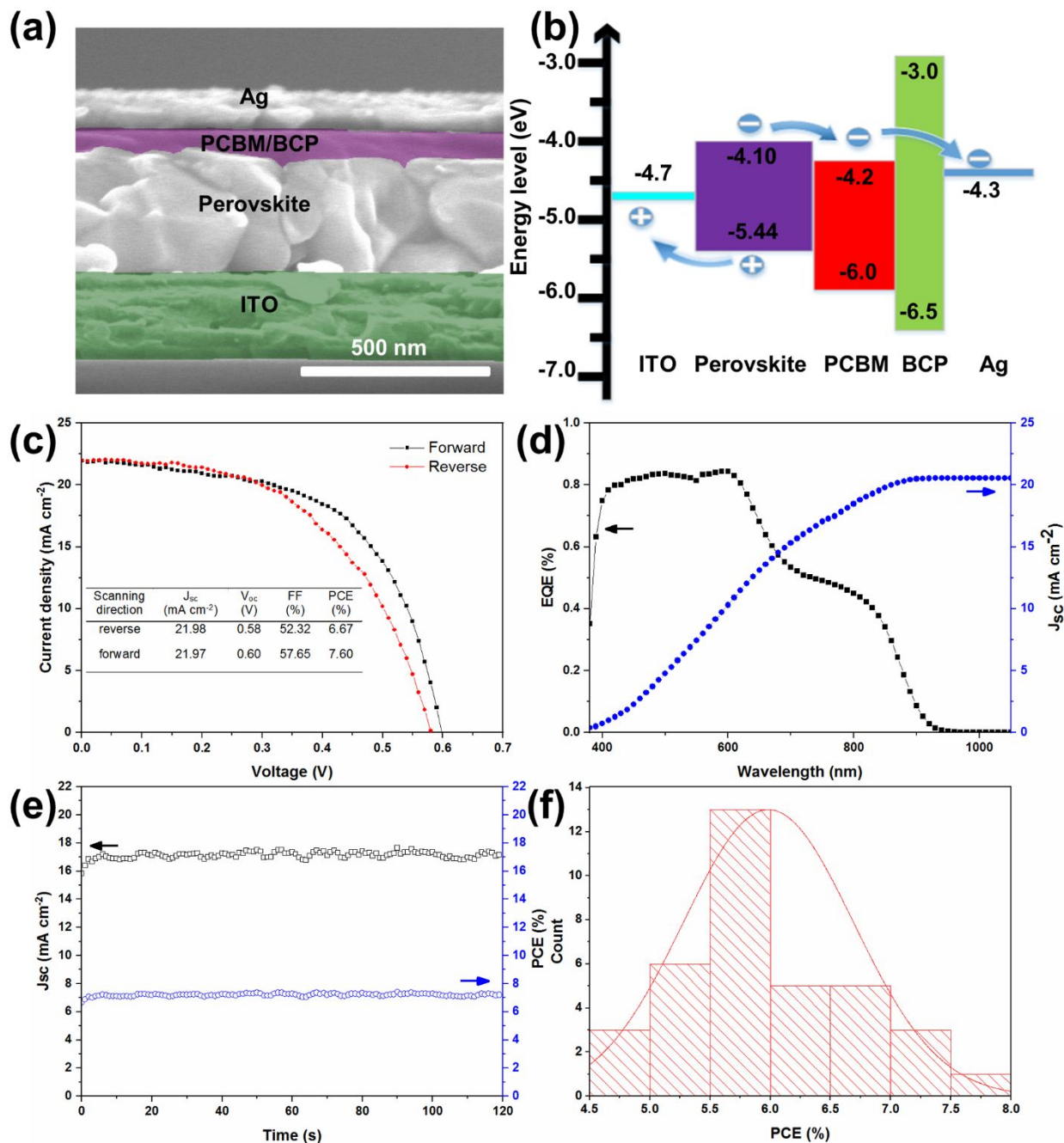
**Figure 4.** (a) XPS survey spectrum (b) UPS spectrum, and high resolution XPS core level for (c) Cs 3d, (d) Sn 3d, (e) Pb 4f, and (f) I 3d of the tin-lead perovskite film with Sn molar ratio of 0.30 (dashed red lines in (c)~(f) correspond to curve fitting with CasaXPS).

The HTM-free devices are fabricated with the structure of ITO/perovskite/PCBM/BCP/Ag (Figure 5a) by removing the PEDOT:PSS hole transport material in the inverted full structure (Figure 3a). The typical energy-level band alignment of the HTM free perovskite solar cells based on the film with Sn of 0.30 are shown in Figure 5b. This energy level alignment between  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  and PCBM should facilitate electron extraction and block holes.

Corresponding cross-sectional SEM image are shown in Figure 5c, in which all functional films are well layered and the thickness of the  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  perovskites is estimated to be  $\sim 300$  nm, in good agreement with that of full-inverted structure devices. We tested different two end member films, namely  $\text{CsPbI}_3$  and  $\text{CsSnI}_3$ , and other Sn-Pb ratios in this structure (Figure S5 and Figure S6). As shown in Figure S5 and Table S2, HTM-free devices with  $\text{CsPbI}_3$  and  $\text{CsSnI}_3$

are almost under the short-circuit condition and in fact, with Sn fraction above 40%, most devices are in the same situation of short-circuit (Figure S6) – this result is consistent with literature suggesting that tin-lead alloyed materials become exponentially more susceptible to heavy doping via oxidation as Sn content increases; this heavy doping likely creates shunting paths through the perovskite, which would be more of an issue when there no hole-selective contact present,<sup>60</sup> whereas tin-lead films with Sn ratio ranging from 0.10 to 0.35 are valid in the HTM-free structure (Figure S6, key parameters are summarized in Table S3). The PCE of these devices increase gradually with the increase of the Sn molar ratio till the optimal value has been achieved in the films with Sn ratio of 0.30, and then drop down gradually with the increasing Sn. We associate the result of the validity of the devices based on the lower Sn incorporation (typical Sn ratio <50%) with the intermediate energy level formed between the VBM and CBM coming from the foreign atom, which will facilitate the carrier transport as described in the TRPL when the HTM is absent in the devices, but the mechanism is not fully understand yet and it would be further investigated in our future work. The J-V characteristics of the HTM-free solar cells based on CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> was carried out under the same condition as that of full-structured solar cells (Figure 5c). Compared with the full-structured devices, an even higher J<sub>sc</sub> of 21.97 mA cm<sup>-2</sup> is achieved, and an impressive PCE of 7.60% has been obtained in spite of slightly decreasing V<sub>oc</sub> and FF of 0.60 V and 57.65%. Moreover, the J-V curves of devices measured under forward and reverse scanning with different scan rates show less hysteresis with a small

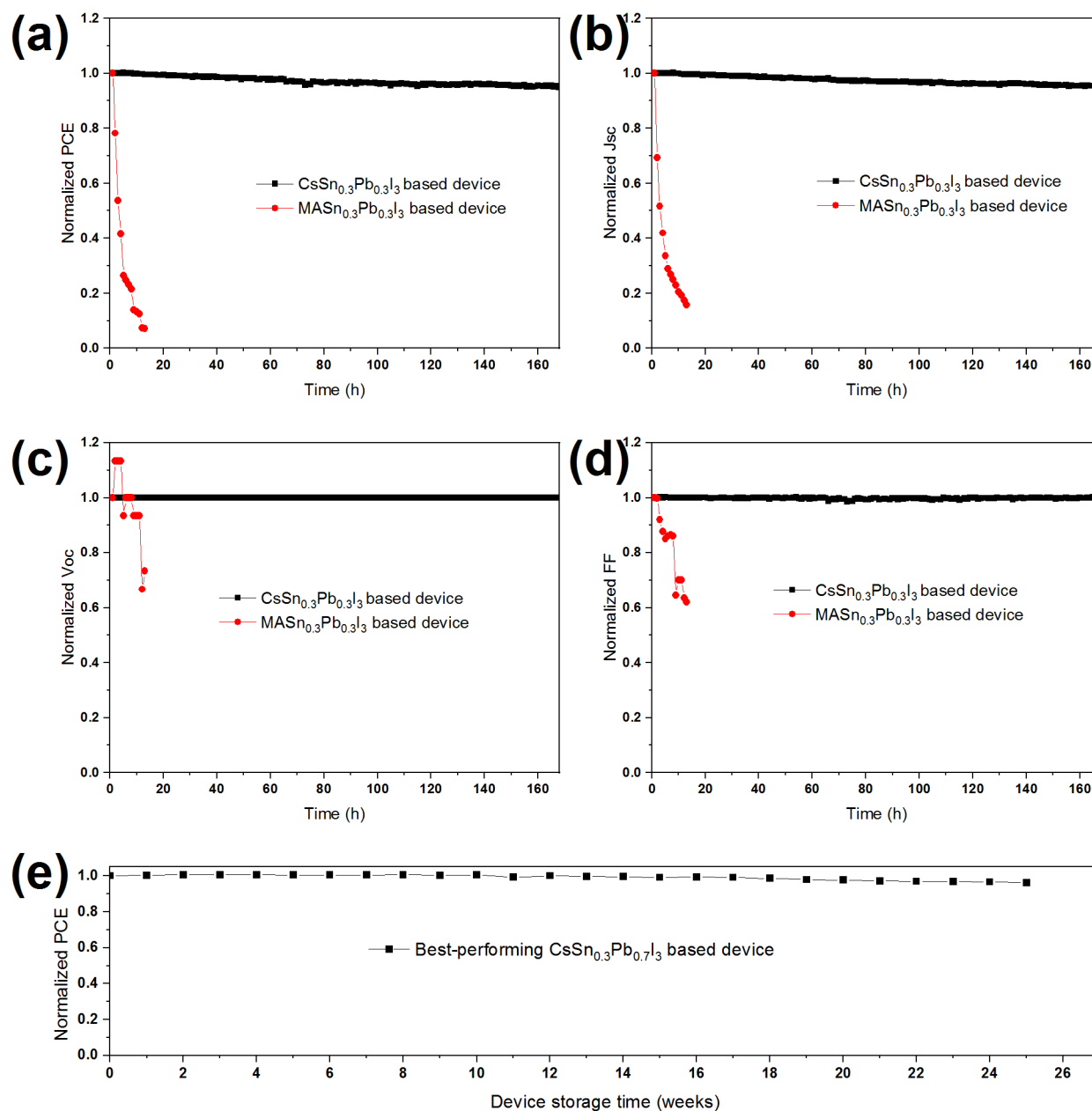
hysteresis index only 6.72% than the full-structured devices, which is further confirmed with J-V bi-directional scanning with different sweep rate as shown in Figure S7 (corresponding parameters summarized in Table S4). The EQE spectrum with an integrated current density and SPO of the HTM-free device based on CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> are presented in Figure 5d and Figure 5e, respectively. As shown in Figure 5d, almost all of the EQE values with little fluctuation are larger than 82% in the wavelength between 420 nm to 600 nm, where it shows a linear-like change in the full-structured devices (Figure 3d). This can be ascribed to less parasitic absorption in the absent PEDOT:PSS layer and is consistent with higher current density in the HTM-free devices. The EQE also shows a broad response with wavelength onset up to 925 nm, indicating a bandgap of ~1.34 eV for this film, which is consistent with the UV-Vis absorption (Figure 1a) and that in the full-structured device (Figure 3d). The photocurrent density integrated from EQE spectrum with the AM1.5G solar photo flux are found to be consistent with the J<sub>sc</sub> extracted from J-V curves as shown in Figure 5c. Importantly, instead of the dramatic decrease of PCE at steady state seen with the full-structured cells, the power output stabilizes at ~7.31% with the HTM-free devices during maximum power point tracking as shown in Figure 5e. This stabilized PCE is very close to that extracted from the J-V plot and notably, even without HTM layer, it is in fact superior to that in the full-structured devices. More importantly, to the best of our knowledge, this is the highest efficiency to date for an inorganic perovskite solar cell with the bandgap below 1.40 eV, especially for the HTM-free inorganic perovskite devices.



**Figure 5.** (a) Typical cross-sectional SEM image, (b) energy level diagram, (c) J-V curves, (d) EQE and integrated current density, (e) stabilized power output at maximum power tracking point of the reported HTM-free perovskite solar cells based on  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  film, (f) Statistical histogram distribution of PCEs extracted from 36 individual  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  based HTM-free perovskite solar cells

To test the reproducibility of the HTM-free devices, we measured J-V curves of 36 individual  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  based PEDOT-free perovskite solar cells. Figure 5f shows the histograms of the J-V scan PCEs of these cells. We note that

the distribution of the PCEs lie in a range from 4.87% to 7.60% with an average PCE of  $5.98\% \pm 0.70\%$ , implying reasonable reproducibility of the tin-lead alloyed solar cells.



**Figure 6.** Normalized photovoltaic performance of (a) PCE, (b)  $J_{sc}$ , (c)  $V_{oc}$  and (d) FF as a function aging time of the perovskite solar cells based on  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  and  $\text{MASn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  without encapsulation under one-sun continuous illumination in glovebox and (e) Photostability of the  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  based perovskite solar cells without encapsulation as a function of storage time in glovebox measured for every week.

Finally, we probed the photostability of HTM-free perovskite solar cells based on inorganic  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  films under continuous one-sun illumination in a nitrogen-filled glovebox. For comparison, we fabricated organic-inorganic counterparts-  $\text{MASn}_{0.3}\text{Pb}_{0.7}\text{I}_3$  based devices as the controlled samples. Figure 6 shows the normalized photovoltaic parameters as a function of the illuminated time of the devices for more than 168 h, in which measurements were automatically performed every hour by Keithley 2400 via a computer-controlled program. During the measurements,  $\text{MASn}_{0.3}\text{Pb}_{0.7}\text{I}_3$ -based device showed an improvement in  $V_{oc}$  at the beginning of the measurement, the device performance dropped quickly with rapid decline in  $J_{sc}$  and FF and after 8 h light soaking, the PCE of the device retains

under 8% of original efficiency. In contrast, the FF of  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$ -based device fluctuated between 98.45% and 100.17% of the initial value, and  $J_{sc}$  exhibited slight increase for the first 8 h and then followed slow drop at the rest of the time, whereas  $V_{oc}$  remained unchanged with a constant value of 0.66 V. Eventually, the overall PCE of the device exhibited no significant change and remained 94.86% of that from the start point after 168 h continuous illumination, suggesting a negligible light soaking effect which might ascribe to the large ion migration barrier of the inorganic perovskite.<sup>23, 78</sup> These results indicate that  $\text{CsSn}_{0.3}\text{Pb}_{0.7}\text{I}_3$ -based device poses good photostability which is a very important for practical application of the solar cells. We further investigated long term stability of the all-

inorganic CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub>-based device by measuring best-performing device after every 7 days (Figure 6e) which was stored in an opaque container in the glovebox. The corresponding PCE of the device remains more than 97% of the fresh one even after almost 6 months. Besides the stability in the glovebox, we also examined the air stability of CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub>-based device compared to that of most recently reported inorganic CsPbI<sub>2</sub>Br based perovskite solar cells, which are considered to be the state-of-the-art all-inorganic photovoltaic (PV) devices with good stability.<sup>28, 29, 39, 79, 80</sup> The stability tests of are carried out in the ambient atmosphere by measuring the PCE (extracted from J-V curves) for every 2 hours (Figure S8). We observed that CsPbI<sub>2</sub>Br based device display a nearly linear decrease in PCE and quenched after stored for 22 hours in air. Notably, the PCE of CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> based devices show no obvious loss at the first 4 hours exposing in air and then slight loss finishing at 87.0% of the initial value at the end of the period, indicting excellent moisture resistance of our simplified tin-lead alloyed device, which might come from the enhanced phase stability by incorporation of Sn.

## CONCLUSIONS

In summary, we have demonstrated that partially substituting Pb with Sn in all-inorganic perovskites is a promising strategy to stabilize the photoactive black phase at low temperature and attain efficient and stable solar cells with ideal bandgap. With alloyed CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> perovskite, we are able to fabricate functional solar cells both with and without a PEDOT:PSS hole transport material. These devices showed promising PCEs of 9.41% and 7.60% (stabilized PCE of 7.23% and 7.31%) respectively, making these devices the most efficient low bandgap all-inorganic perovskite solar cells reported to date. Moreover, we found that the simplified HTM-free CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> based devices exhibit good photostability with negligible light soaking and long-term stability in the glovebox. The good moisture resistance of unencapsulated devices in air is also presented. There is still much room for improvement in the performance of CsSn<sub>0.3</sub>Pb<sub>0.7</sub>I<sub>3</sub> based solar cells and the incorporation of Sn into inorganic perovskites opens up new possibilities for attaining extremely stable and highly efficient perovskite solar cells in the future.

## ASSOCIATED CONTENT

### Supporting Information

The Supporting Information is available free of charge at:

Additional data for UV-Vis absorption, XRD, TRPL, XPS, GDOES, J-V and air stability characterization.

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## Notes

The authors declare no competing financial interest.

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A tin-lead alloyed low bandgap ( $\sim 1.34$  eV) inorganic perovskite has been reported to attain efficiencies up to 9.41% (stabilized-PCE 7.23%) and 7.60% (stabilized-PCE 7.31%) in full-structured and HTM-free perovskite solar cells, respectively, which are the highest efficiencies to date for all-inorganic perovskites with a bandgap below 1.40 eV.

