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Shock states and thermodynamic properties of liquid silicates

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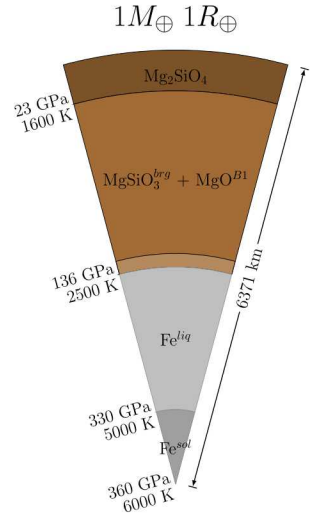
⁴Department of Earth and Planetary Sciences, Harvard University, Cambridge, MA, USA



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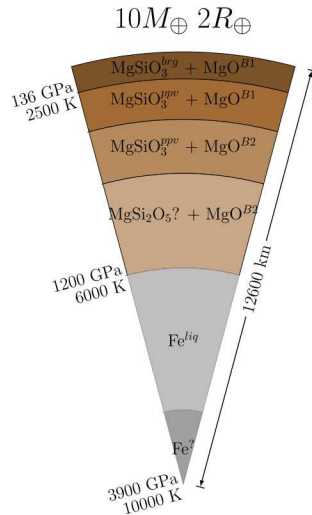
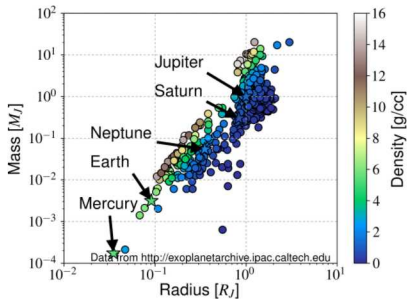
Why dynamic compression?

Giant impacts explore exotic P, T states



Why dynamic compression?

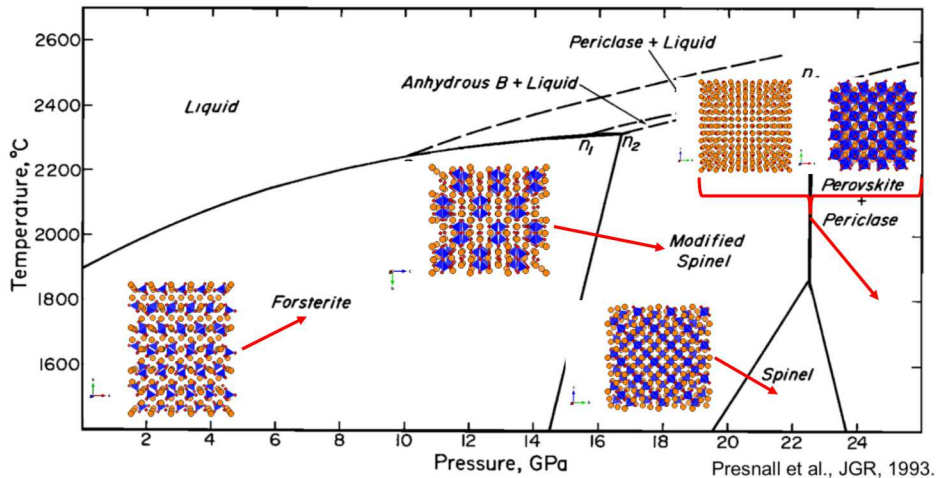
Planetary interior models unconstrained



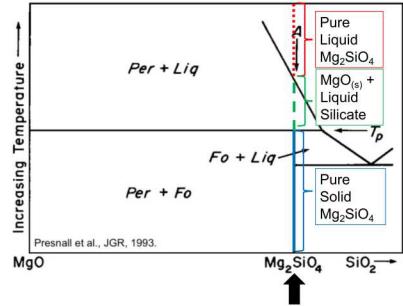
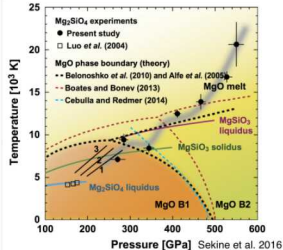
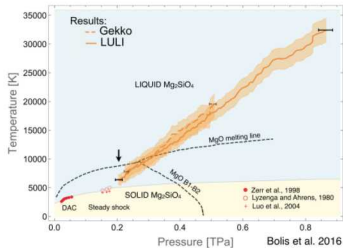
Low pressure phase diagram

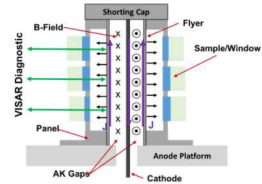
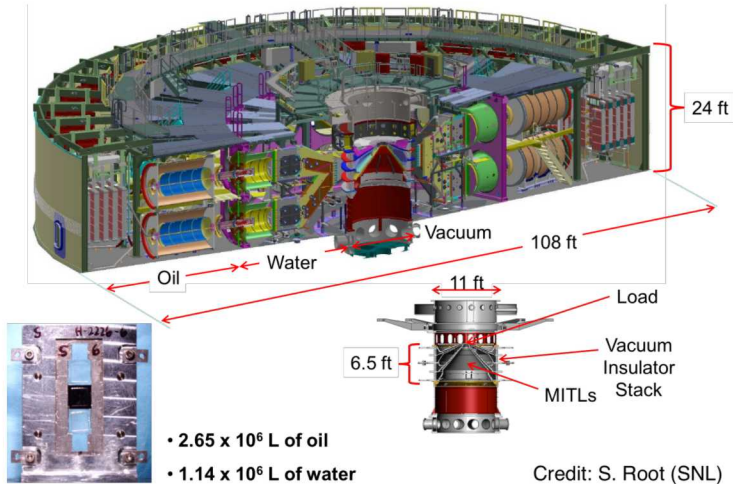


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- Multi-component system
- Incommensurate melting
- Melting ~ 120 -160 GPa
- No XRD, indirect phase ID





$$E = 35 \text{ MJ}$$

$$I = 20 \text{ MA}$$

$$\Delta t = 20 \text{ ns}$$

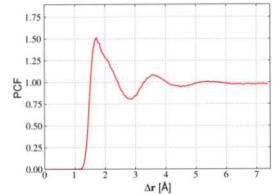
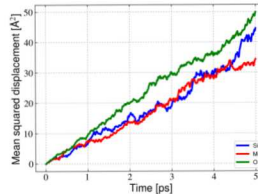
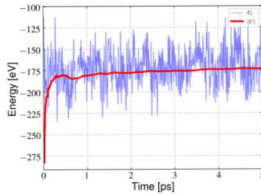
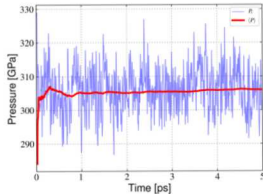
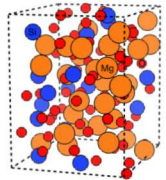
Credit: S. Root (SNL)

DFT-based molecular dynamics

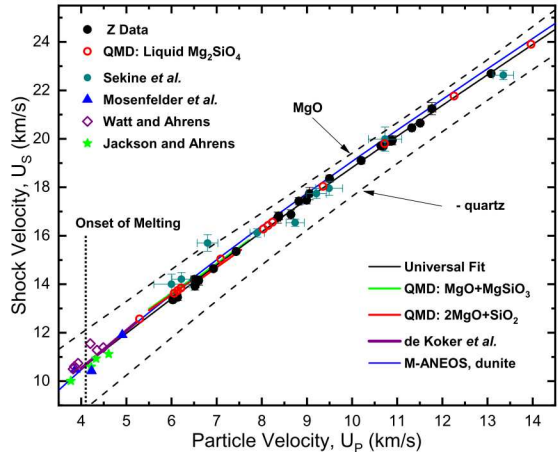
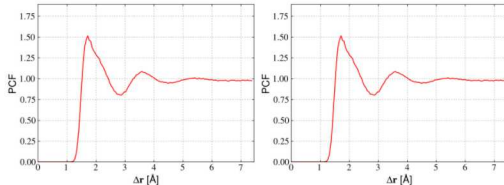
Basic idea: Solve $\mathbf{F} = m\mathbf{a}$ for the ions, with the force due to electrons computed self-consistently from solution of the Kohn-Sham hamiltonian:

$$\left(-\frac{1}{2}\nabla^2 + V_{ext}(\mathbf{r}) + \int d\mathbf{r}' \frac{\rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} + \frac{\delta E_{xc}[\rho]}{\delta \rho(\mathbf{r})} \right) \phi_i(\mathbf{r}) = \epsilon_i \phi_i(\mathbf{r})$$

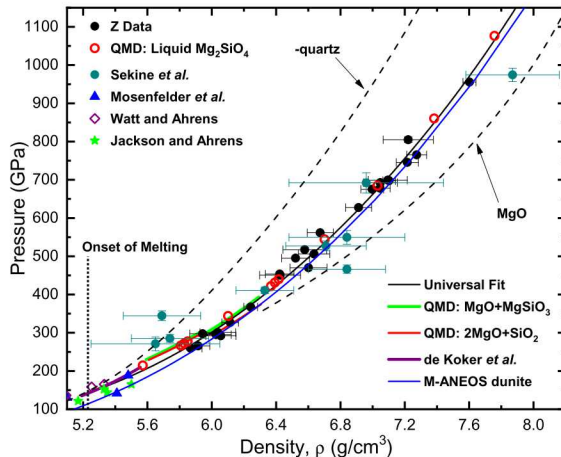
Thermodynamic quantities such as E , P , T , V , directly accessible without making any assumptions about the response or properties of the material *a priori*.



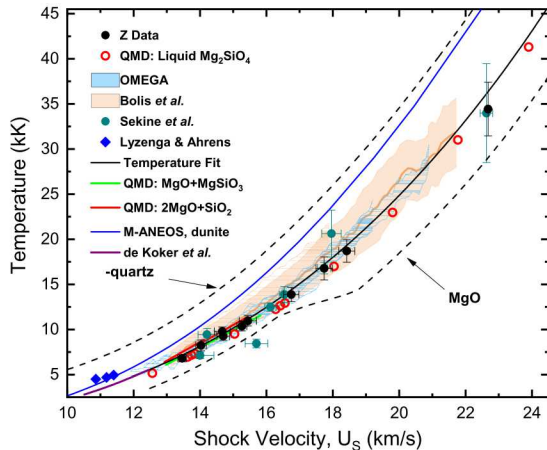
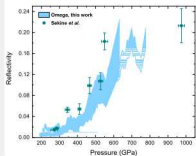
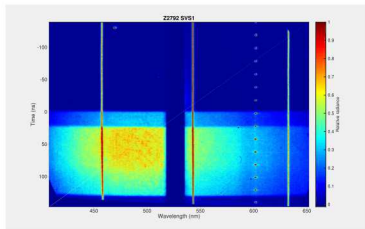
- Data span $13 < U_S < 24$ km/s
- Resolve curvature in Hugoniot
- No discontinuities
- Agree with extrapolated MANEOS



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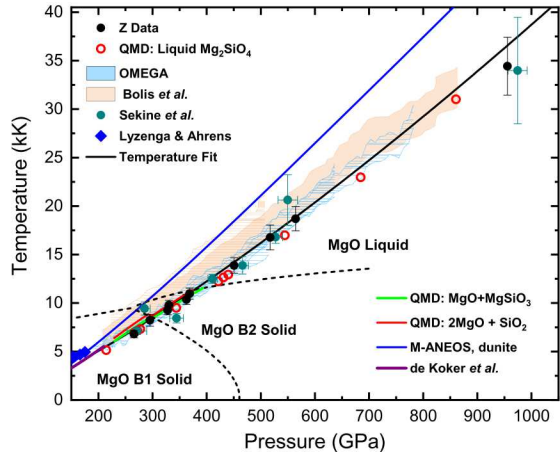


- MANEOS thermal prop's off
- Z, OMEGA, and DFT-MD in agreement



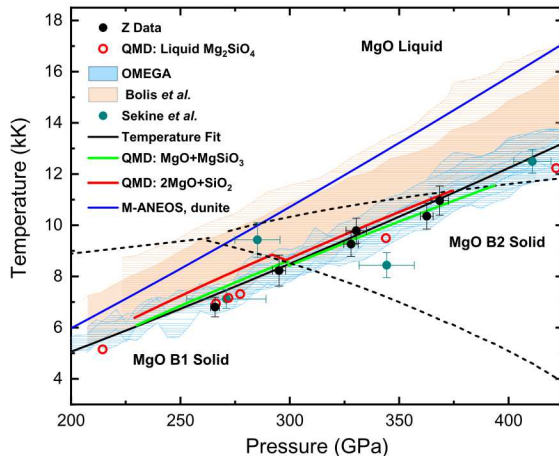
- Data span $13 < U_S < 24$ km/s
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Well... Is it de-mixed, or not?



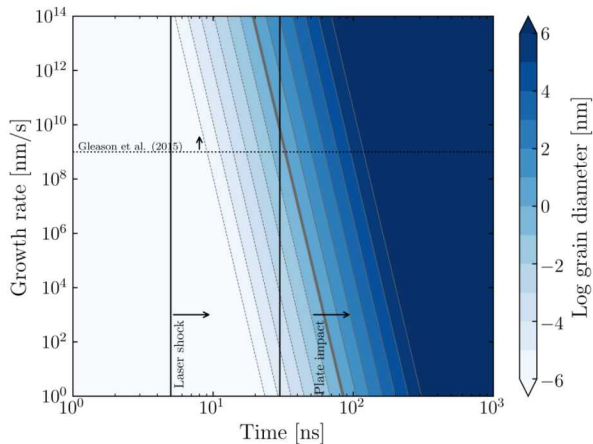
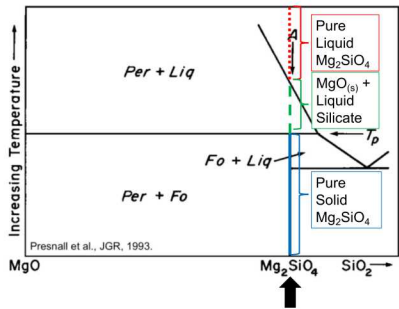
- Data span $13 < U_S < 24$ km/s
- Resolve curvature in Hugoniot
- No discontinuities
- MANEOS thermal prop's off
- Smaller than expected signal due to de-mixing

Well... Is it de-mixed, or not?

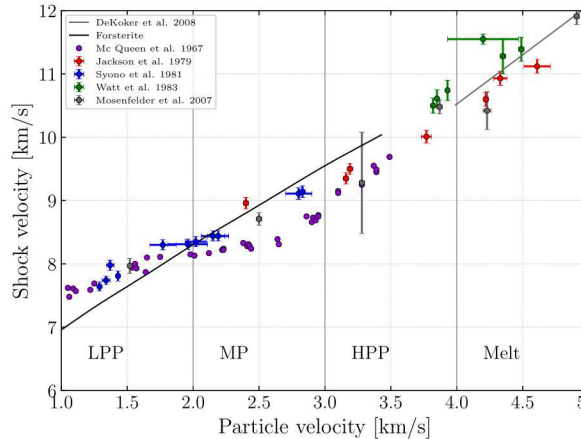


Two things to consider:

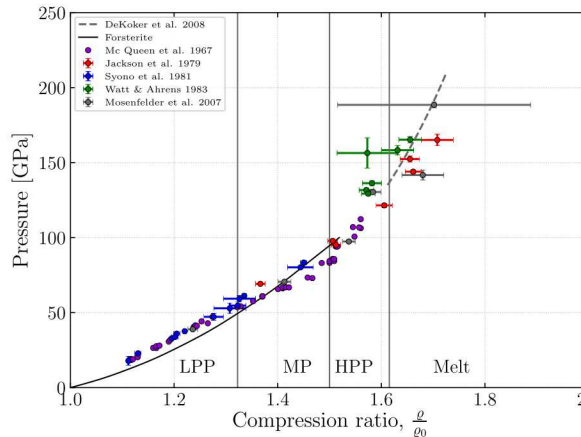
- Incommensurate melting at high pressure?
- Experiments probing equil. states?



- Persistence of Fo to 10's GPa?
- MP and HPP difficult to reproduce
- New XRD experiments suggest metastable Ol to 75 GPa(!)



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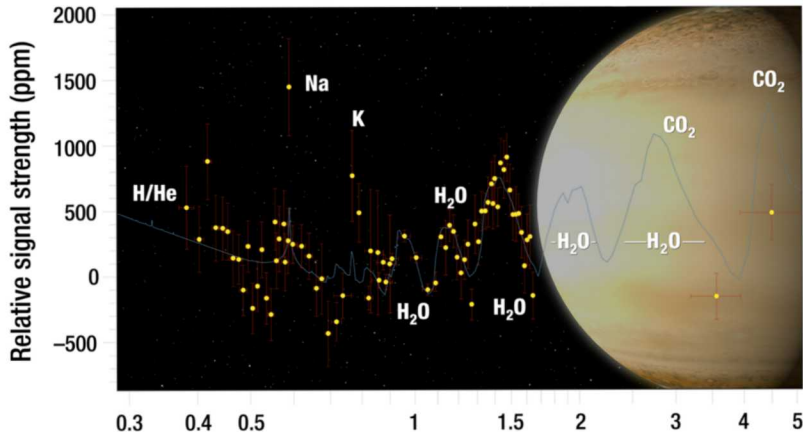


- Hugoniot states in liquid well constrained
- Possible kinetic effects persist at least up to melting
- Entropy, melting, and release states in progress

Questions?

EXTRAS

Comprehensive spectrum of WASP-39b



$$P = \rho_0 U_S u_P, \quad \frac{\rho}{\rho_0} = \frac{U_S}{U_S - u_P}$$

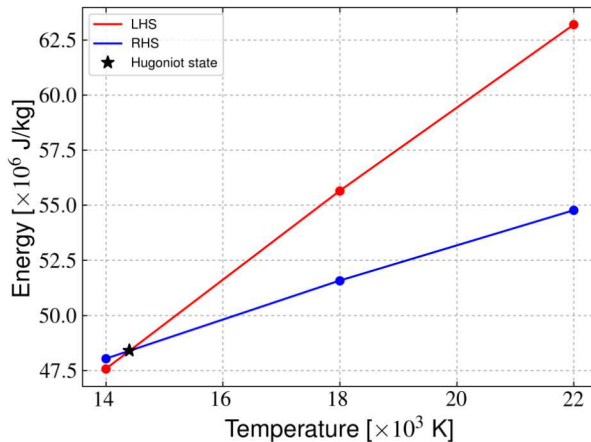
$\Downarrow$

$$E - E_0 = \frac{1}{2}(P + P_0) \left(\frac{1}{\rho_0} - \frac{1}{\rho} \right)$$

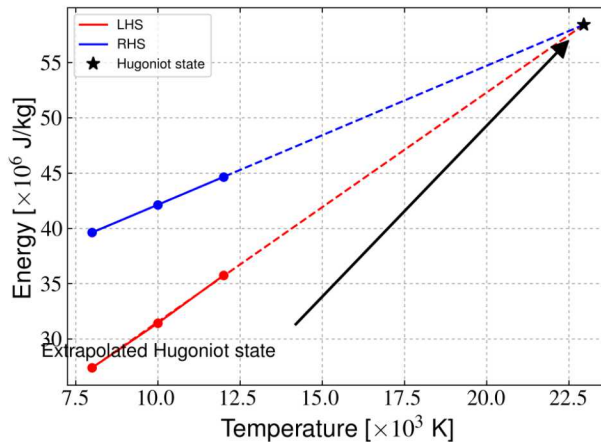
- 1d curve, set by choice of reference

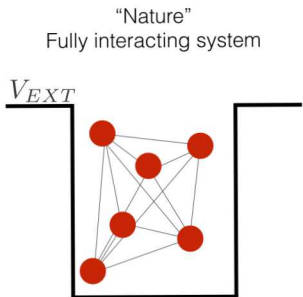


How to estimate the Hugoniot



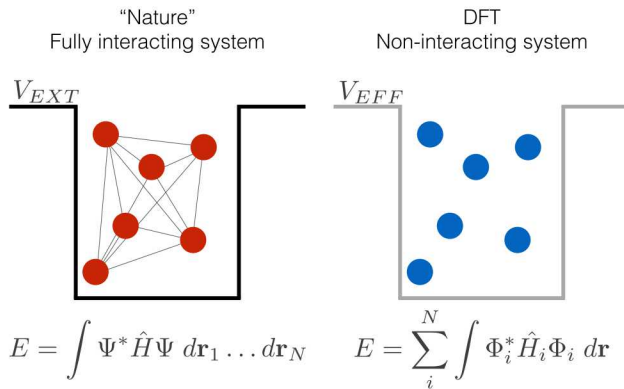
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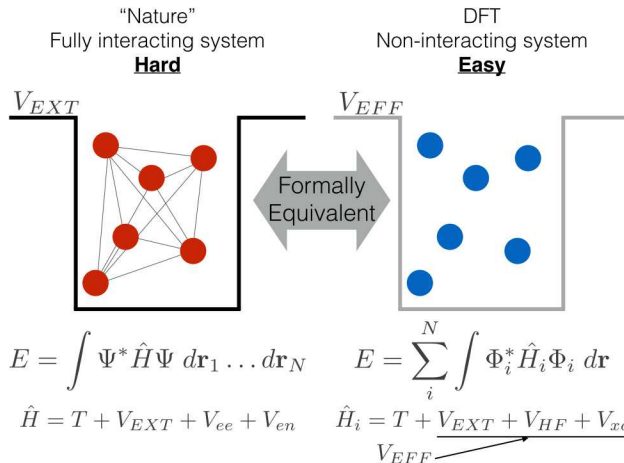


$$E = \int \Psi^* \hat{H} \Psi \, d\mathbf{r}_1 \dots d\mathbf{r}_N$$

Adapted from Mattsson et al. 2005



Adapted from Mattsson et al. 2005



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Planetary science on the Z Machine

