

New Concepts for Improving Dynamic Range in (M)PDV Systems

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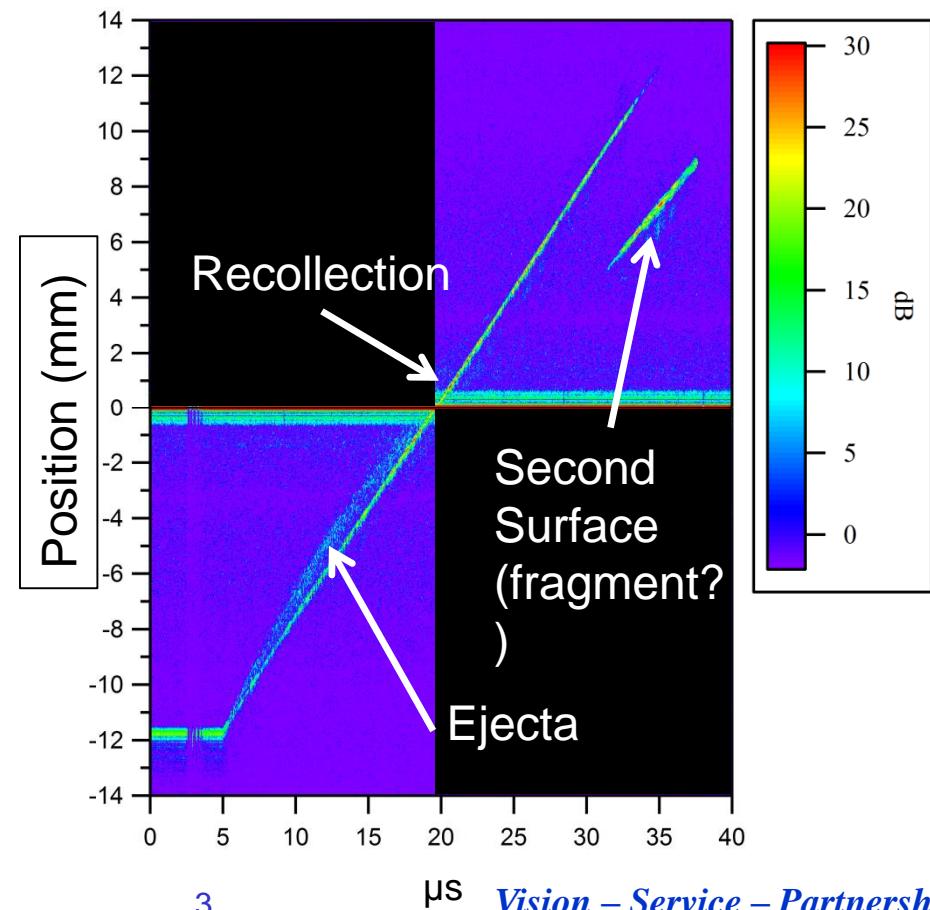
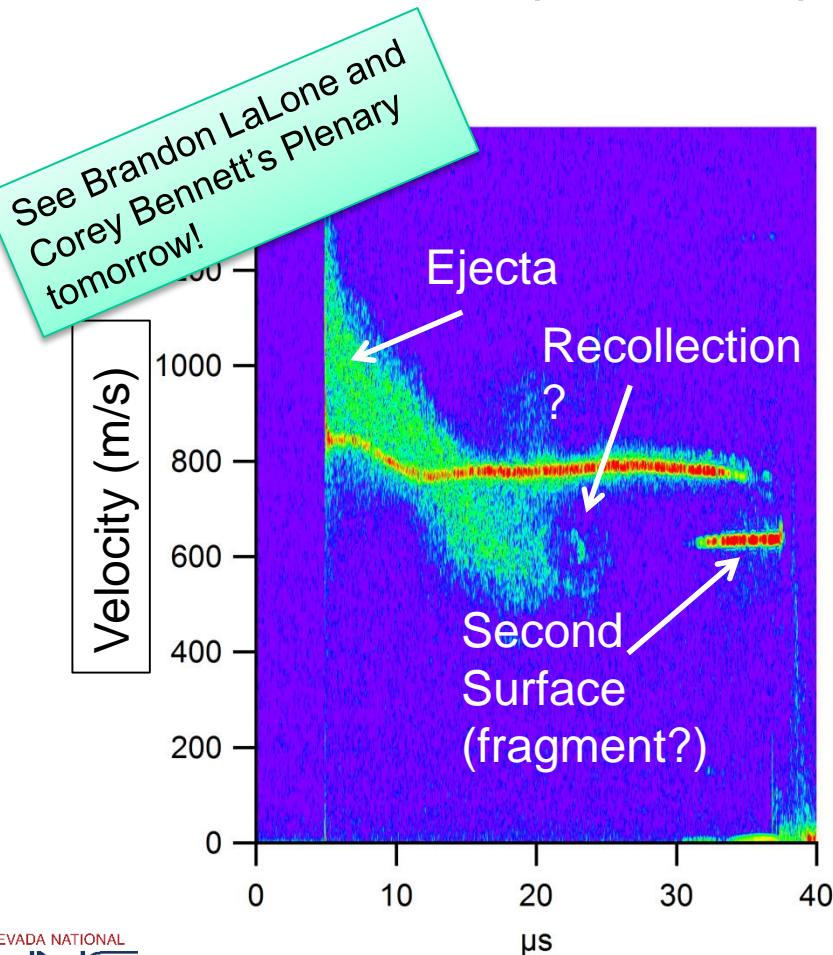
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Overview

- Where are we now
 - Small signal return: Nearly shot-noise limited
 - Large signal return: effective-bits limitation
 - Role of EDFA's
- Deep-time Multiplexed PDV
- Modulated launch-light
- Dynamically refocusing probe optics
 - Resonantly-driven GRIN lens
 - Optically-actuated lens
- Conclusions and next-steps

Why “Dynamic Range?”

- Some experiments have low and/or widely-varying signal returns
- Some experiments have “clouds” of material obscuring the surface
- Limited launch power and probe efficiency



Some back-end schemes have been tried

- Modulate the LO to increase RF amplitude
- Optical auto-gain control on signal return, using SOA
- RF gain control on back-end
- ... none of these provide convincing performance improvements when peak light-returns are <-20 dBm

Low signal returns & the shot-noise limit

- At shot-noise limited signal-to-noise of 1:1:
 - B is bandwidth (Hz), η is quantum efficiency

$$SNR \equiv \frac{\langle i_s^2 \rangle}{\langle i_n^2 \rangle} = 1 = \frac{\eta P_s}{h\nu * B}$$

- For 1550 nm light, we get:

$$P_{shot-noise-limit} = 10 \log\left(\frac{B}{1 \text{ MHz}} * \frac{1}{\eta}\right) - 99 \text{ dBm}$$
- For $\eta=0.7$ (0.9 A/W), the shot-noise limit is -81 dBm in a 50 MHz BW
 - e.g. 2000-point FFT on a 50 GS/s record
- A modern, 6-bit (effective) scope will have ~ 68 dB from its noise floor to full-scale
- So, you should be able to see from the shot-noise limit up to -13 dBm!

EDFA effect on Small Signals

- Noise from LO-ASE:

$$\langle i_{ASE}^2 \rangle = 2\eta^2 e^2 \frac{P_{LO}(P_{ASE}/\text{Hz})}{(h\nu)^2} * 0.5 * B$$

- Add that to the LO Shot-noise to get total noise:

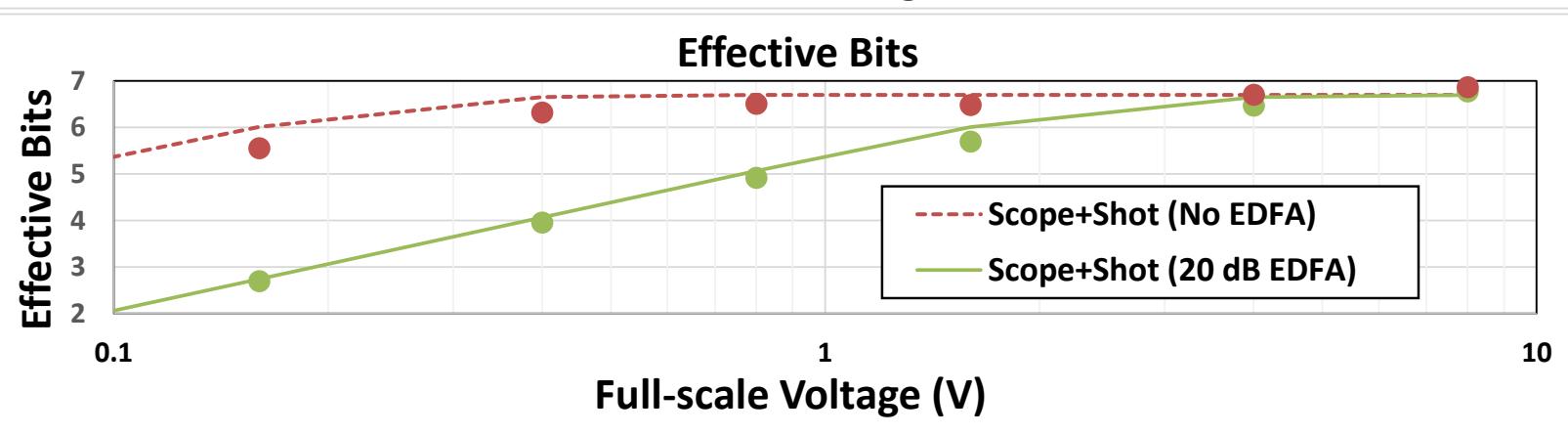
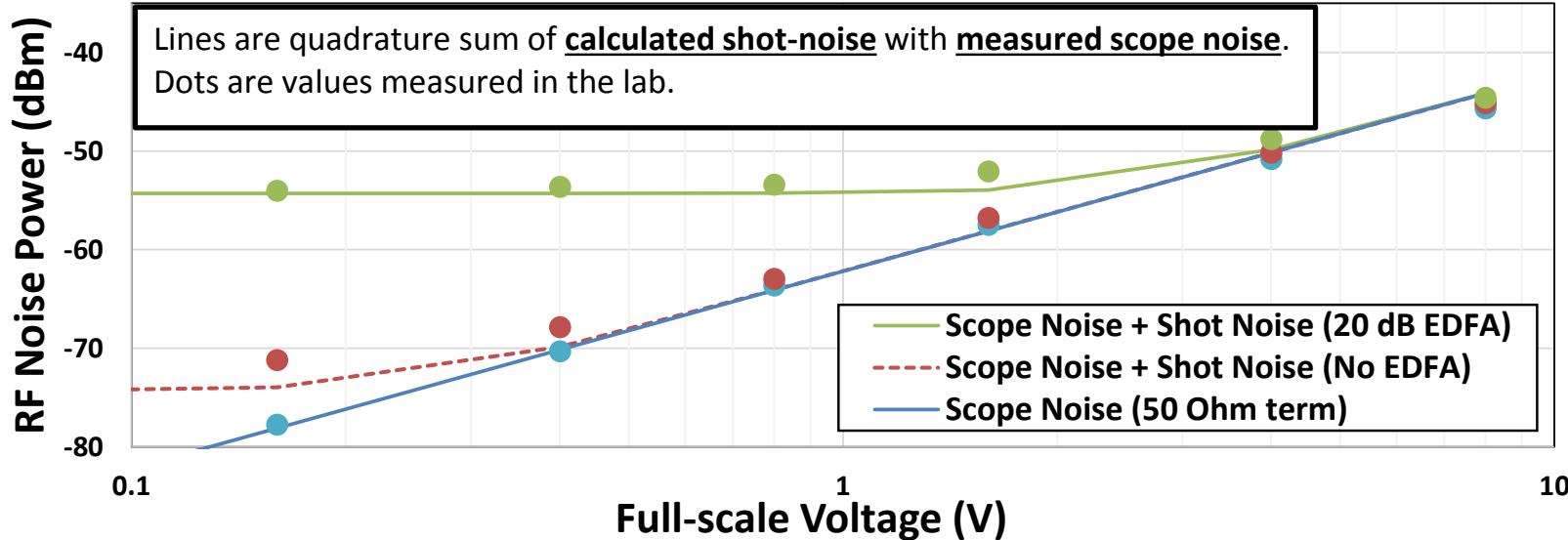
$$\langle i_{n_TOTAL}^2 \rangle = \langle i_{LO_SHOT}^2 \rangle \left(1 + \frac{\eta * (P_{ASE}/\text{Hz})}{2 * h\nu} \right)$$

High-gain, fully
inverted EDFA

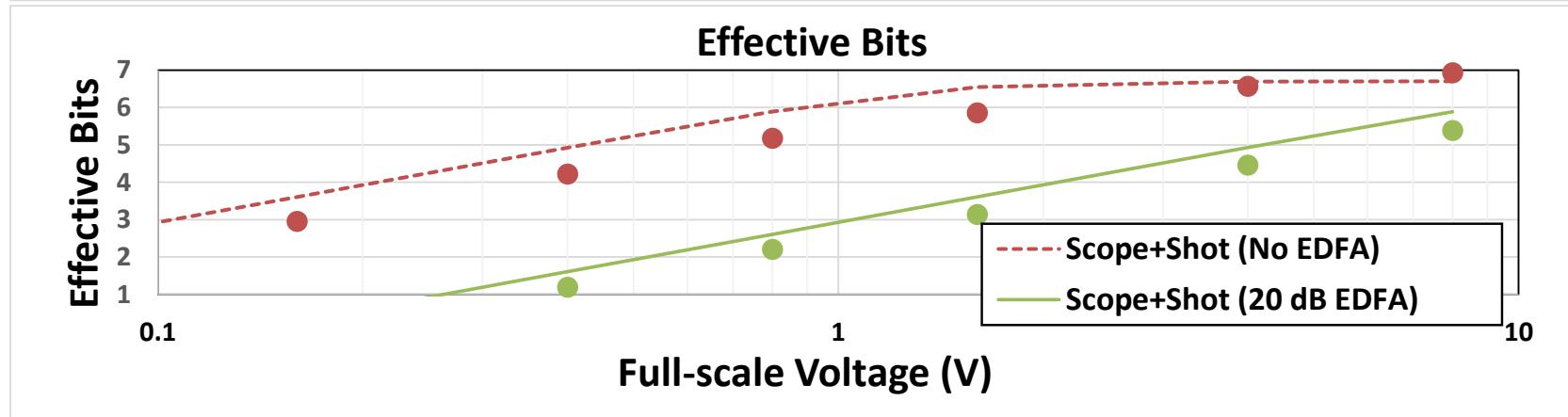
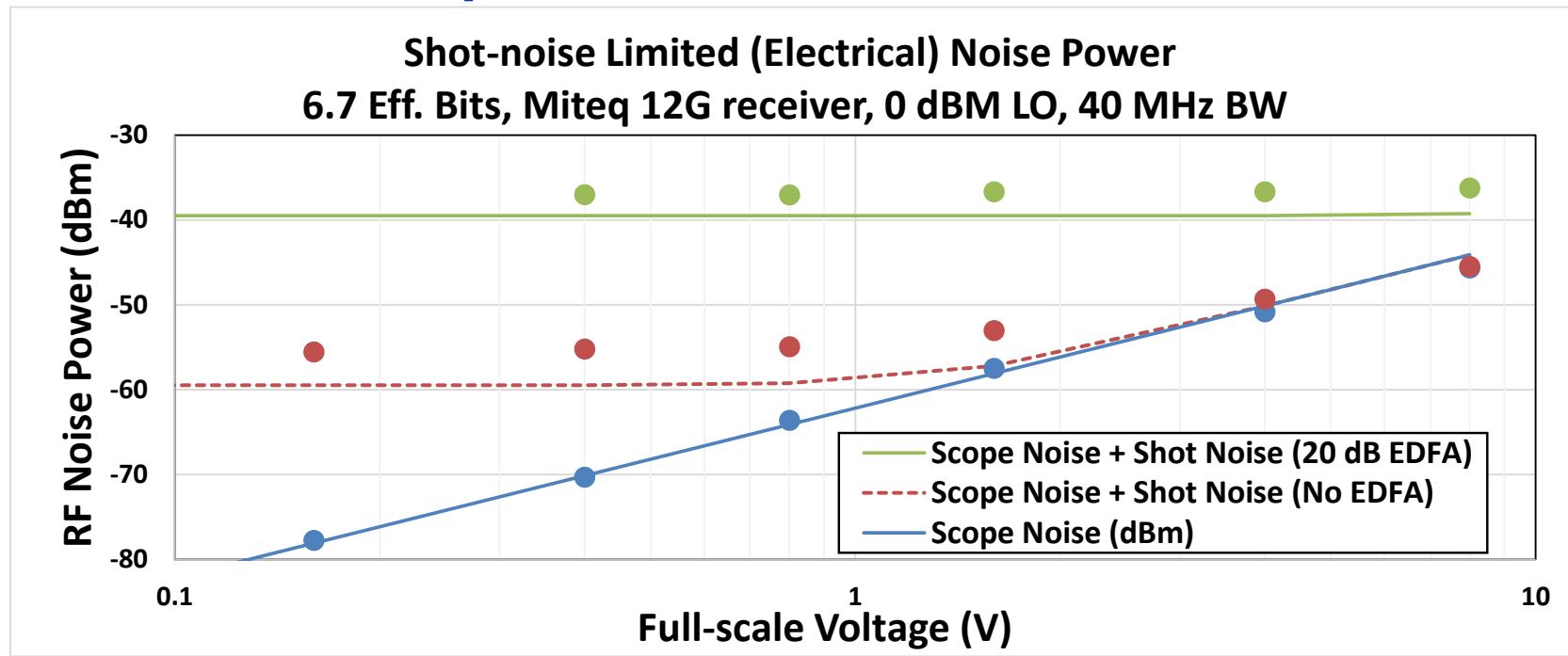
$$\langle i_{n_TOTAL}^2 \rangle \approx \langle i_{LO_SHOT}^2 \rangle * \eta_{PD} * G_{EDFA}$$

Lab Data: Discovery 402 Receiver

Shot-noise Limited (Electrical) Noise Power 6.7 Eff. Bits, DSC402 receiver, 0 dBm LO, 40MHz BW



Lab Data: Miteq 12G Receiver



Top-end of the range: Heterodyne Signal Amplitude

PD current, from the textbook: $\langle i_s^2 \rangle = 2 \left(\frac{e\eta}{hv} \right)^2 P_{LO} P_S$

Power into 50 Ohms:

$$P(mW) = \frac{\text{Transimpedance}^2 * \eta^2}{16000 \Omega * V^2} * \text{GAIN}_{OPT+RF} * P_{LO}(mW) * P_S(mW)$$

$$P_{RF}(dBm) = P_{LO}(dBm) + P_S(dBm) + 10 \log_{10} \left(\frac{\text{Transimpedance}^2 * \eta^2}{16000 V^2 \Omega} * \text{GAIN}_{OPT+RF} \right)$$

Power gain of O-E conversion

(~23 dBm for MITEQ 12 GHz receiver)

So, if you have your scope set to 2 V full-scale (+10 dBm), and you are using the Miteq 12G receiver (23 dBm OE gain) with 0 dBm LO, you would expect to fill your scope with -13 dBm of signal light on the receiver.

Back-end configurations that get us NEAR Shot-Noise limit

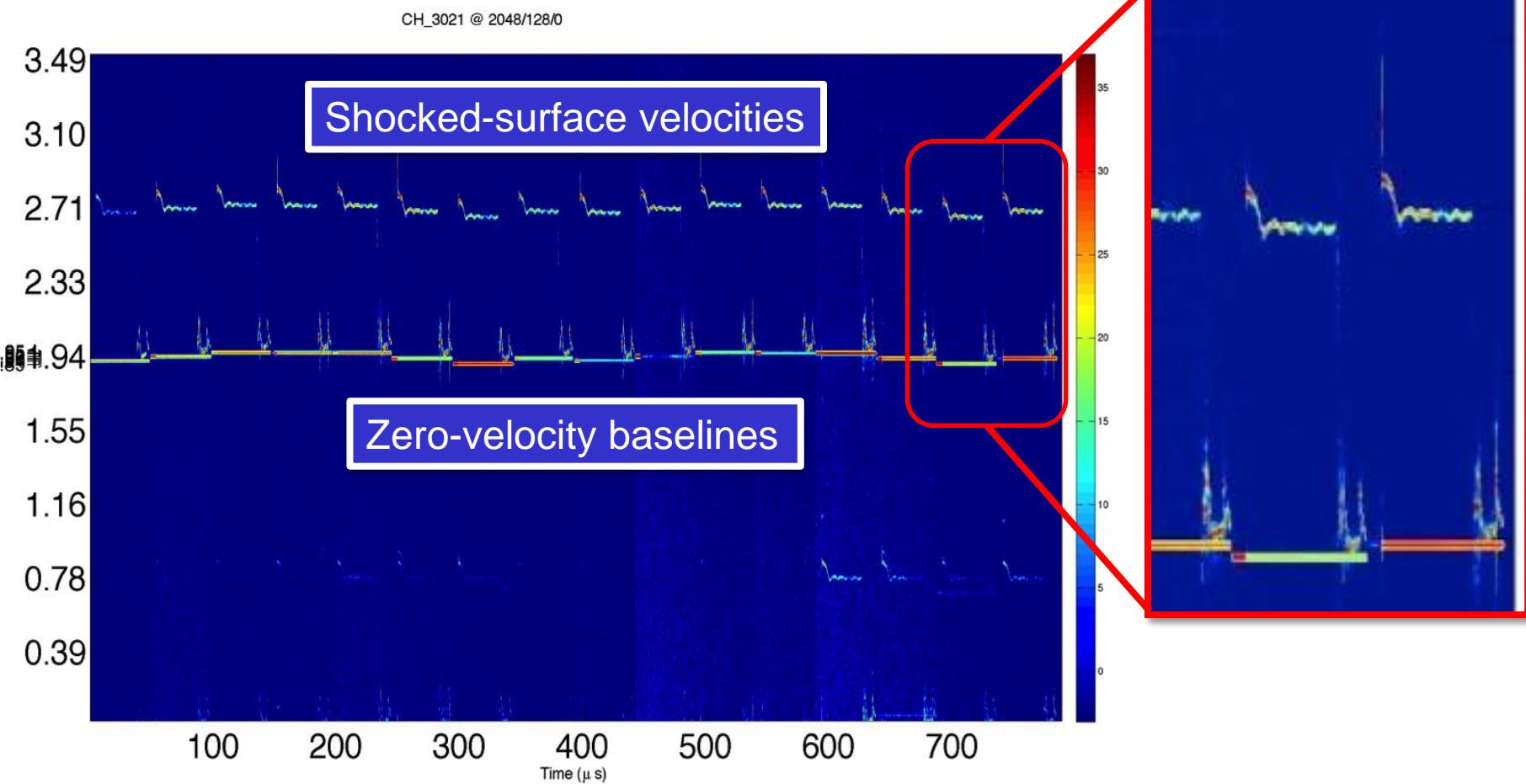
- InGaAs photodiodes at 1550 nm:
 - 75% quantum efficiency, or 0.9 A/W
- Commercial receivers (Miteq, Discovery, NewFocus) with nominal LO power of 0.5 – 2 mW
 - This gets us to regime where **LO shot-noise dominates** over other noise sources
 - Higher LO can bring signal (and noise) up into scope's range
- Bare photodiode(s) with low-noise amplifier and LO up to 30 mW
 - No advantage for (MPDV) over amplified receivers
- Add a commercial, low-noise EDFA preamp to any of the above
 - Raises both signal and noise without changing SNR
 - Good way to compensate downstream (e.g. multiplexing) losses
 - Can help bring signals into scope range
- Modern, high-bandwidth digitizers

Deep-time Multiplexing

- Shot-noise-limited reality: LO shot-noise dominates
- Frequency-multiplexed (and early deep-time multiplexed) MPDV's have multiple LO's on receiver simultaneously
- The NOISE comes from all LO's, but each channel's SIGNAL comes only from its own LO
- We needed to switch the LO light with the signal light
 - Noise dropped by 4x (6 dB)
- Added benefits:
 - Easier to field
 - Data is easier to analyze
 - Less recording bandwidth required

Deep Time MPDV Spectrogram

- Data from deep-time multiplexed experiment
- 16 data channels multiplexed

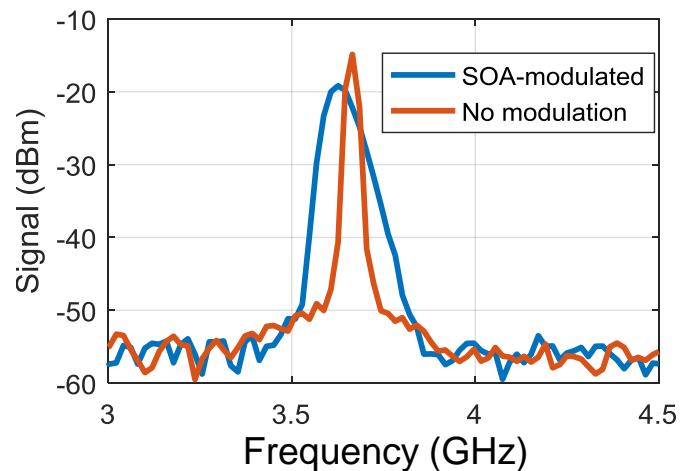


For weak returns, we just need more signal photons!

- Ground-rules:
 - CW light is limited to ~ 20 mW per channel
 - Total power through probe
 - Total power to surface
- Two approaches:
 - Increase the launch power without “breaking the rules”
 - Improve light collection for cases of ejecta obscuring surface

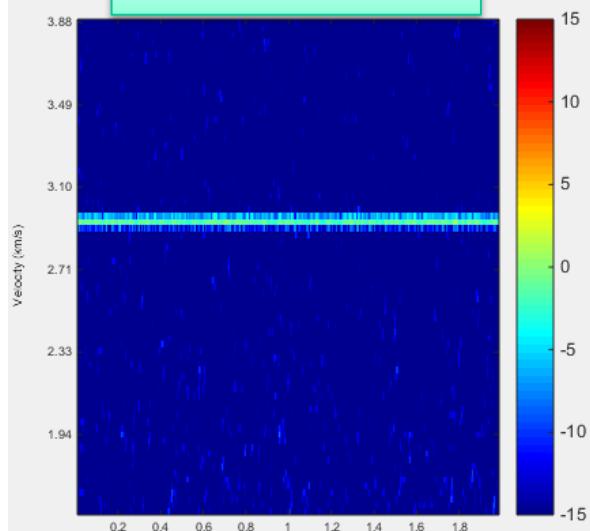
One solution: Modulated launch light

- AO modulator:
 - BW is good enough (50 MHz)
 - don't need Mach-Zehnder
 - High power handling
 - Polarization-insensitive
 - SOA broadens line when modulated
- FPGA-based control
 - 40 MHz master clock
 - 10 MHz ADC & DAC
 - Digital modulation line (up to 50 MHz square-wave)
 - Dual, programmable trigger inputs
- Modulation schemes
 - Free-running, e.g. 50% and 10% duty cycles
 - Triggered waveform
 - Feedback (~ 500 ns)

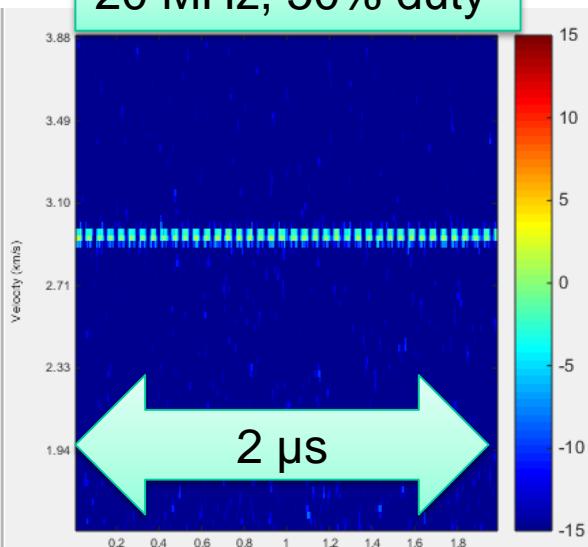


Free-running modulation, -60 dBm time-average signal

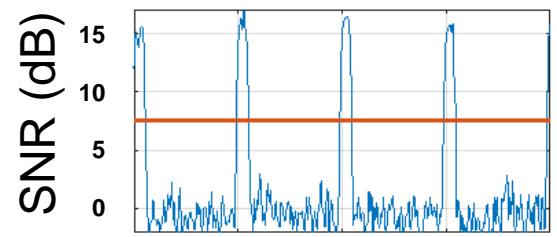
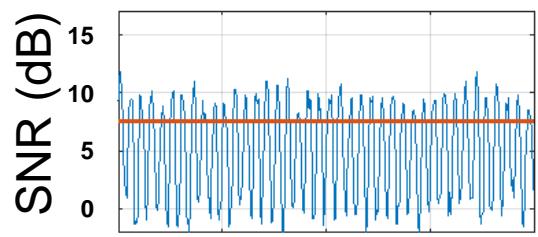
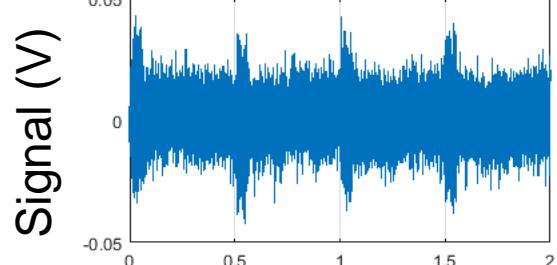
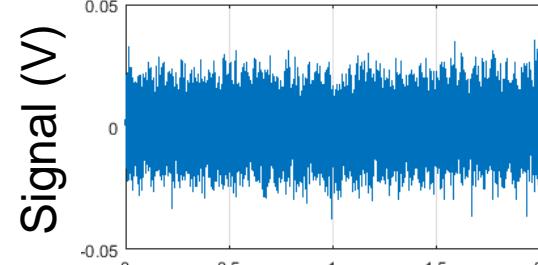
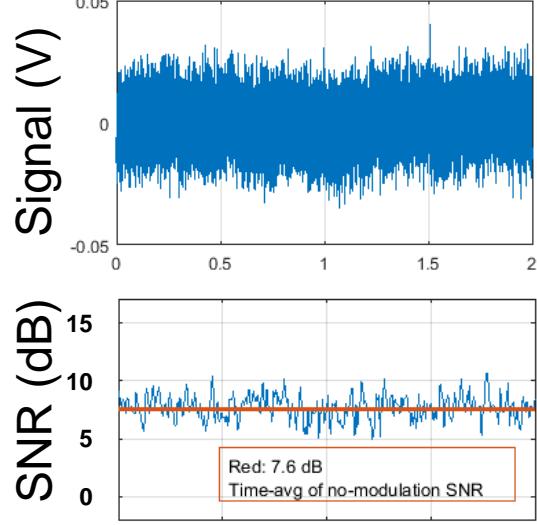
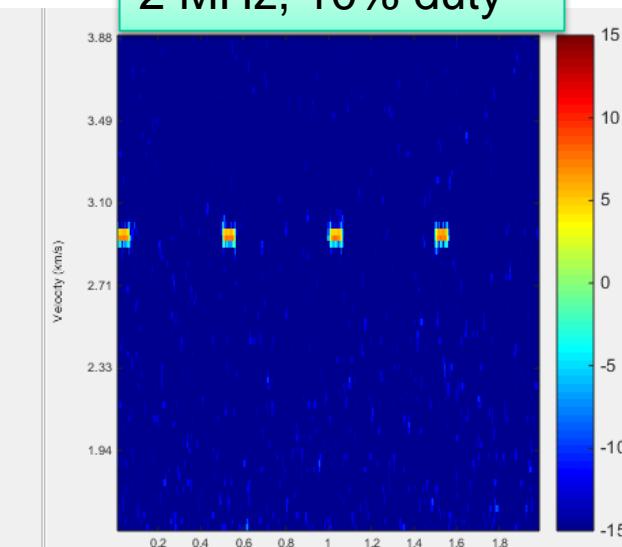
No modulation



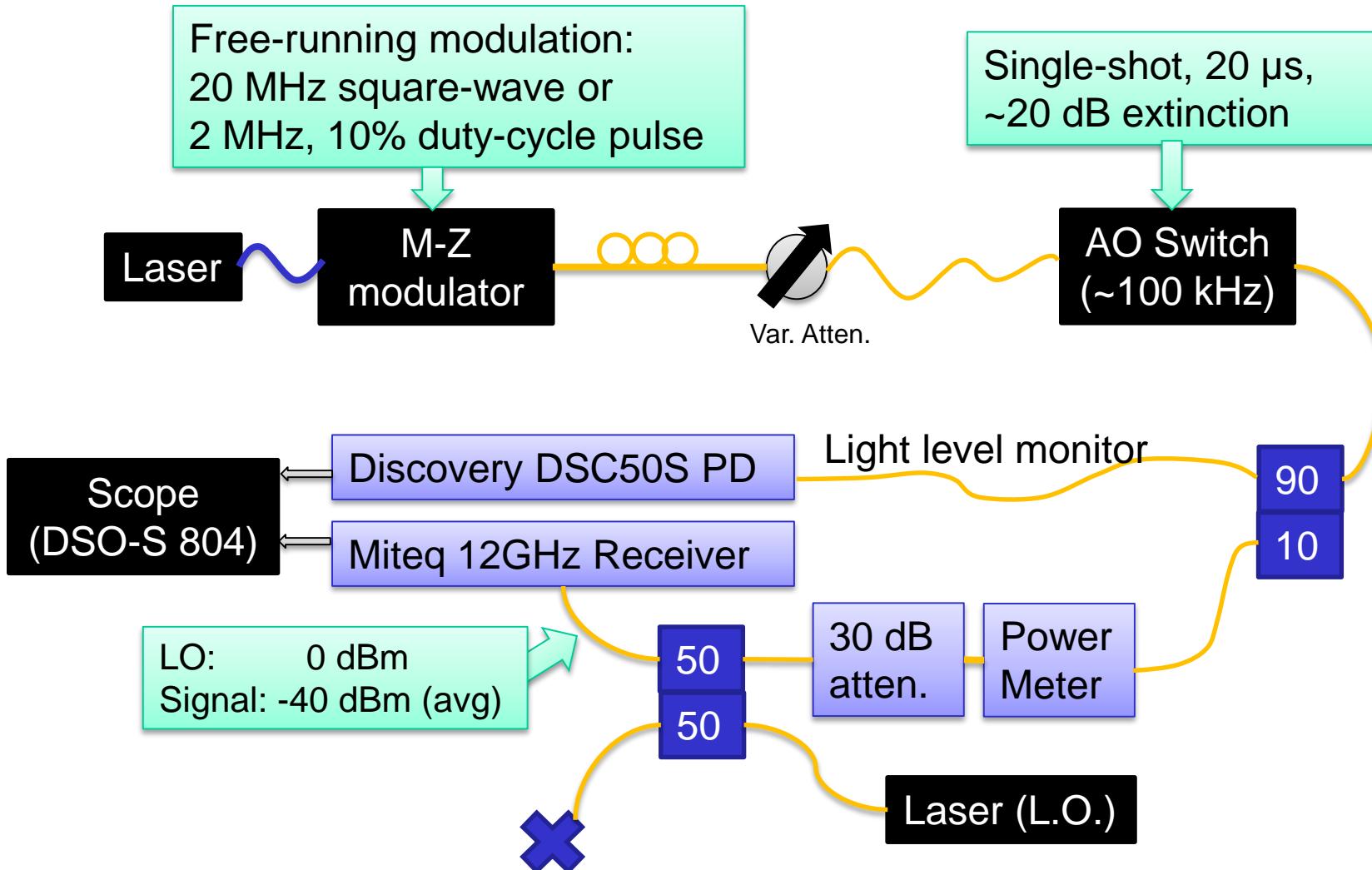
20 MHz, 50% duty



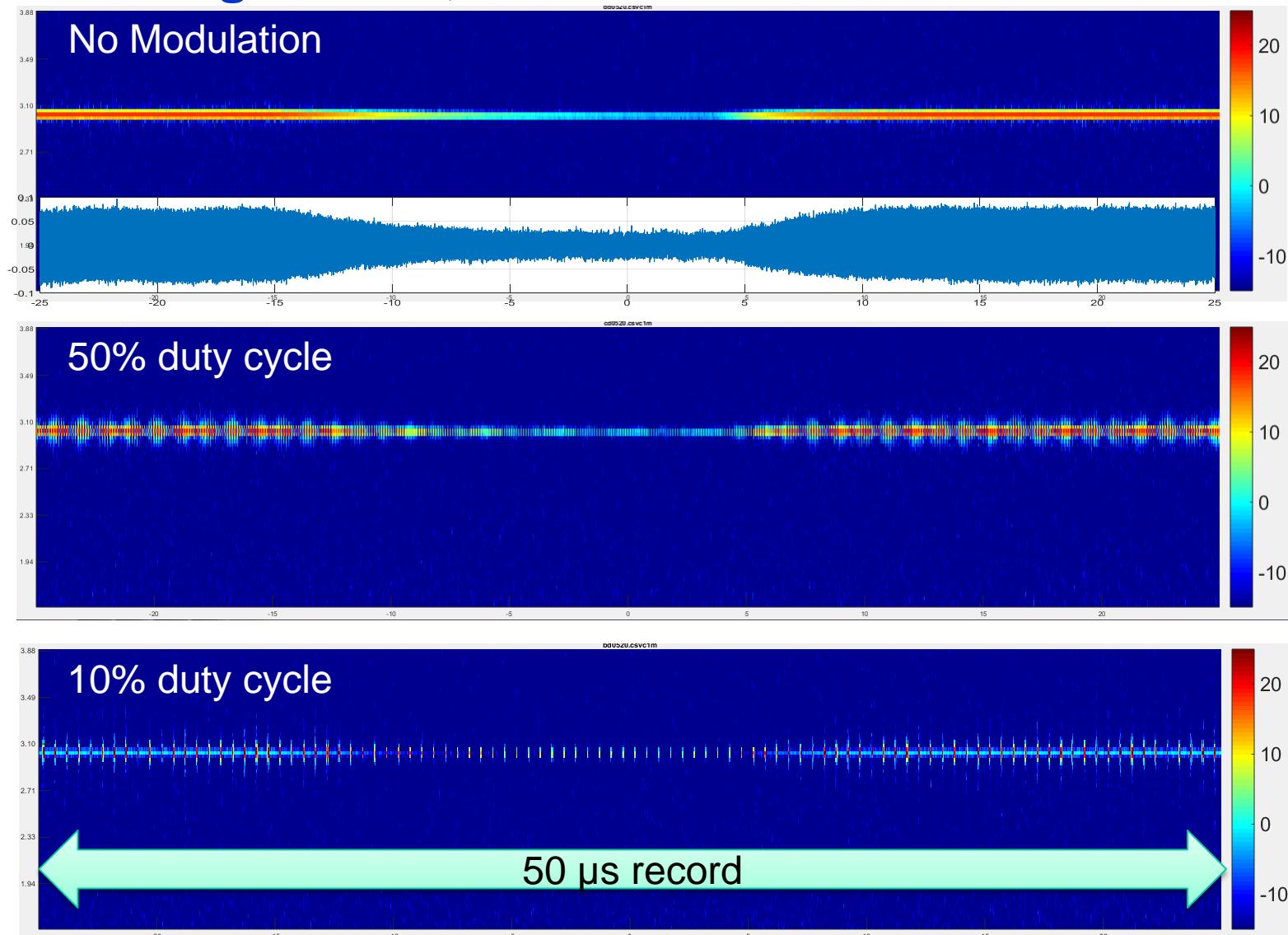
2 MHz, 10% duty



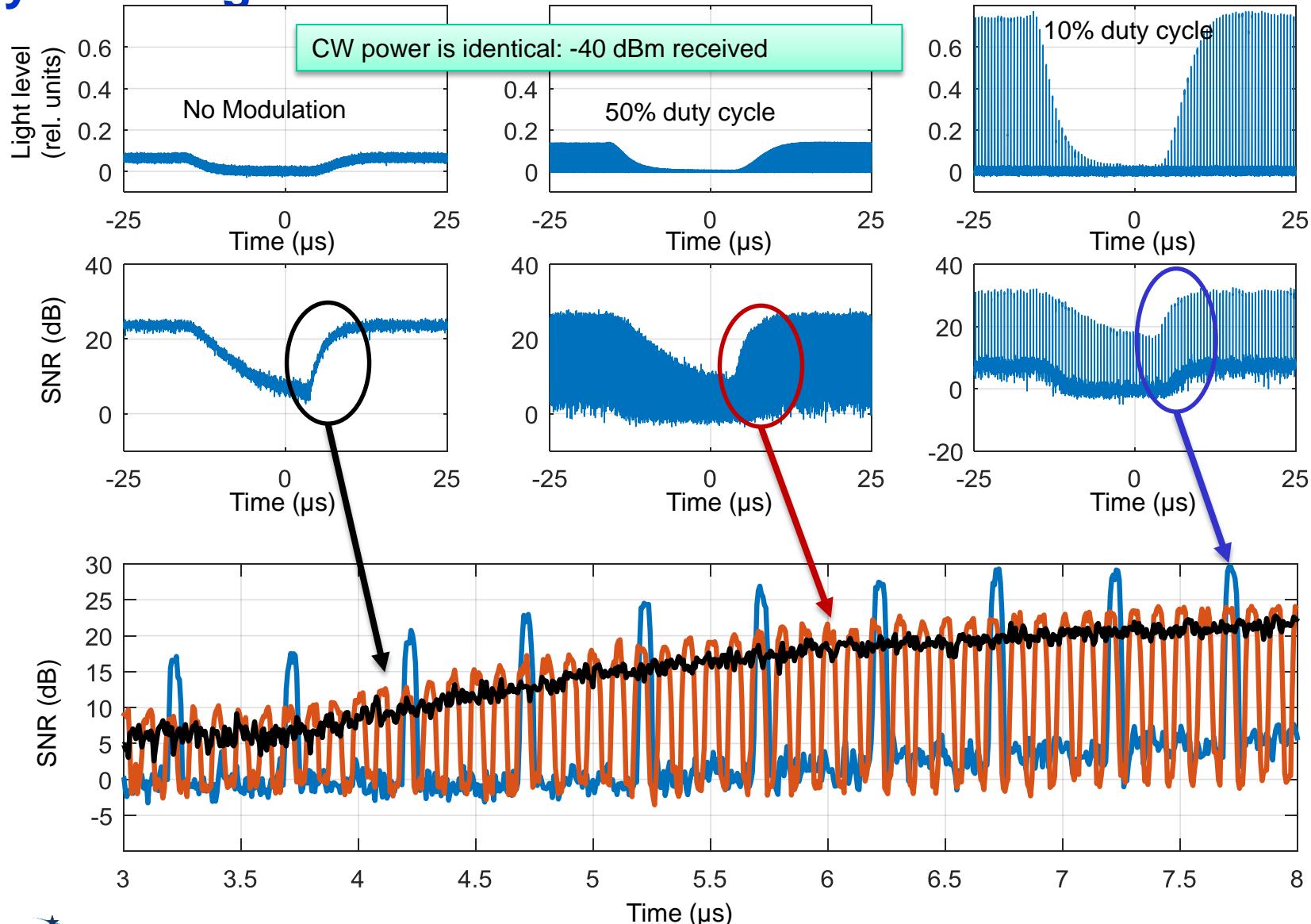
Lab simulation of dynamic signal loss



Dynamic signal loss, -40 dBm to -60 dBm



Dynamic light levels and SNR



Launch-modulation: next steps

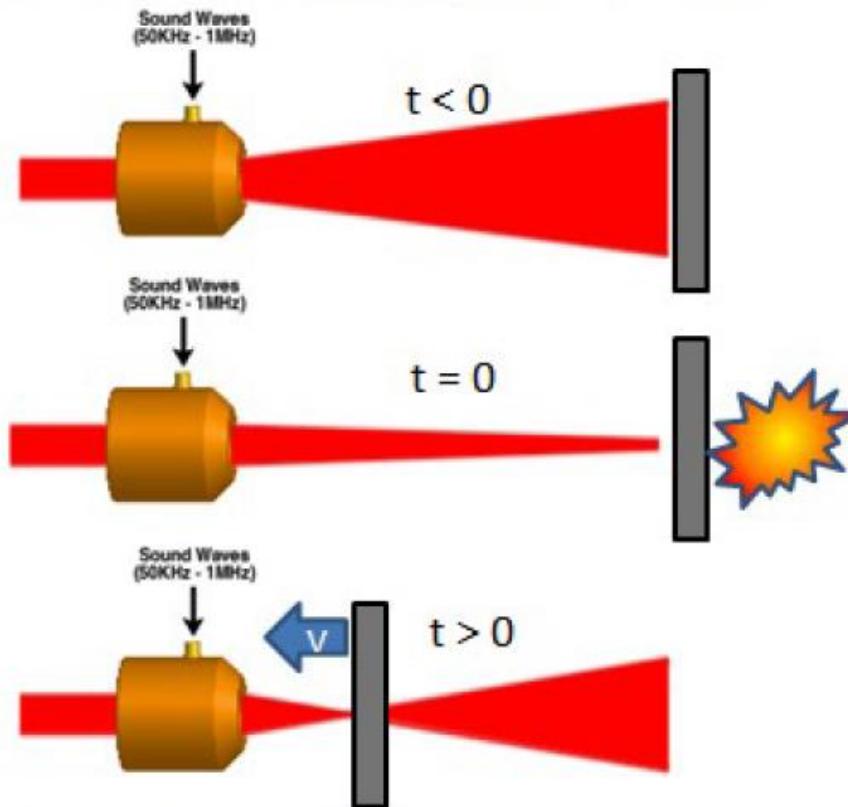
- Build up real-time modulation system:
 - FPGA
 - 50 MHz AO Switch
- Implement modulation schemes:
 - Free-running, variable duty cycle
 - Triggered, programmable output
 - Feedback on return light level (< 1 μ s)
- Field this system on small-scale shots at NSTec / STL
 - Summer 2016

Electrically-actuated lens: TAG Optics

- Resonantly-driven, cylindrical liquid cell
- Density modulation creates gradation of index:
 - Compression: converging focus
 - Rarefaction: diverging focus
- For IR operation, standard resonant frequencies are 140 – 340 kHz
- Higher resonant frequencies have smaller effective apertures
- Need to figure out relationship between focal length (or effective focal length), S2", and collection efficiency

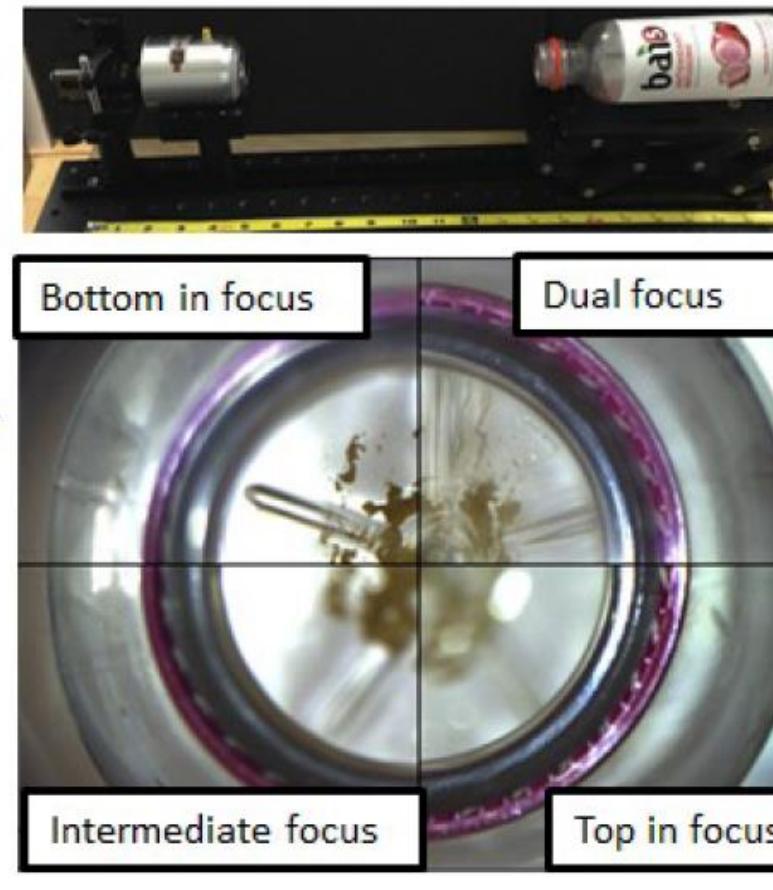
Original concept: Track the Surface

GRIN Lens refocusing with period of 1 – 20 μ s



Source: tag-optics.com

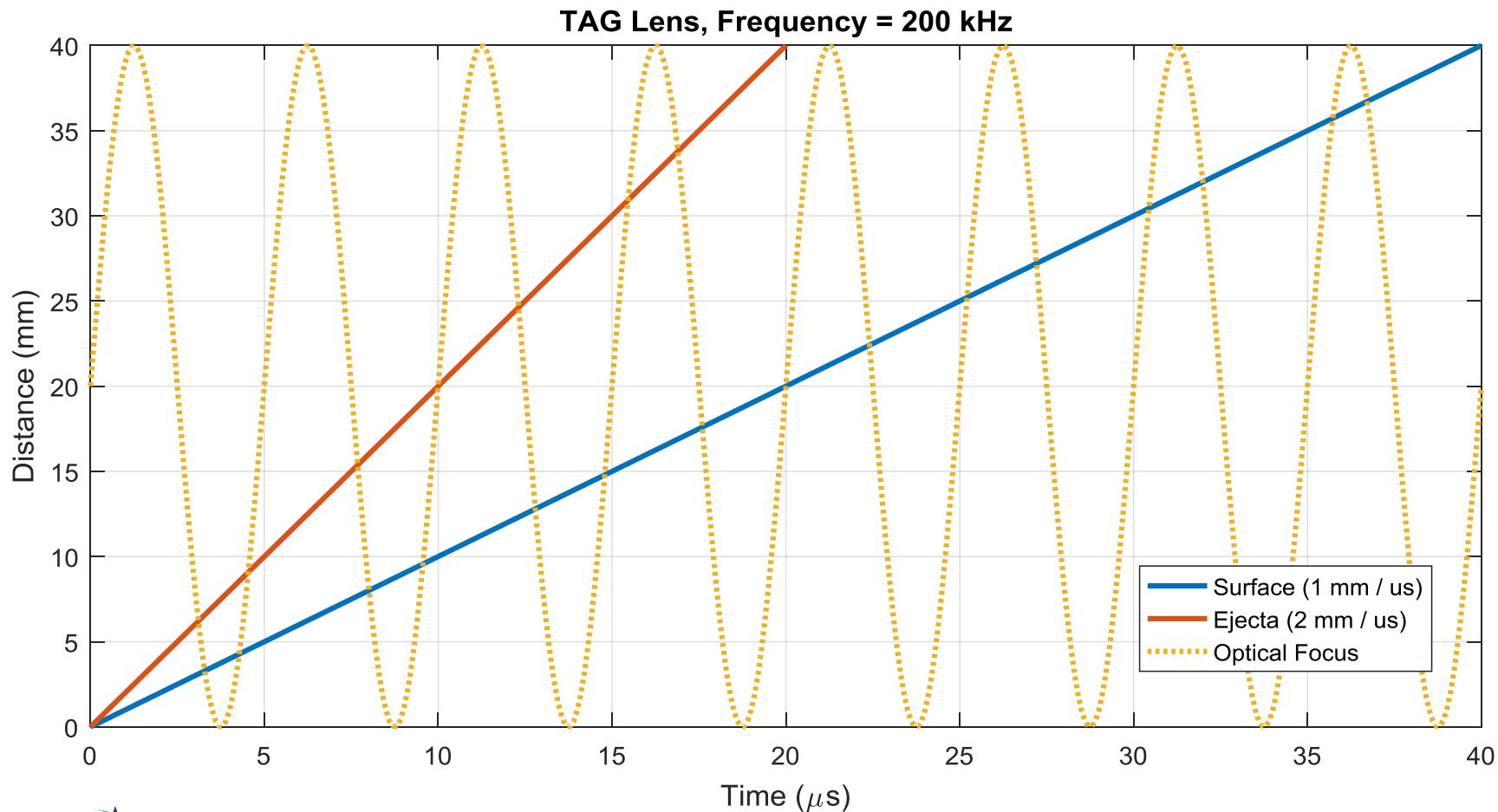
Lab demo with juice bottle



But... resonant frequencies are too high for 10 μ s experiments!

Next Approach: Free-running

Each object passes through focus twice every 5 μ s



Optical relay for TAG Lens

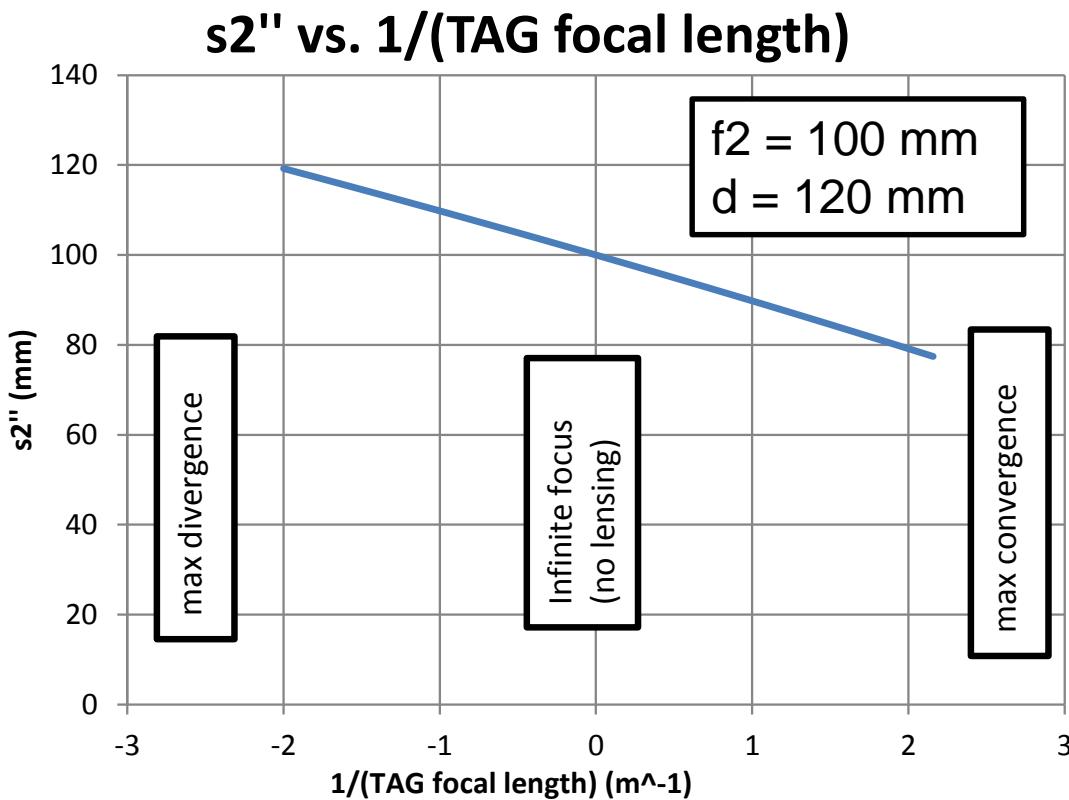
From Fiber
collimator

TAG Lens
(-500 mm → ∞ → +500 mm)

Relay Lens: f_2

s_2''

Focal Plane of
Combined lens system



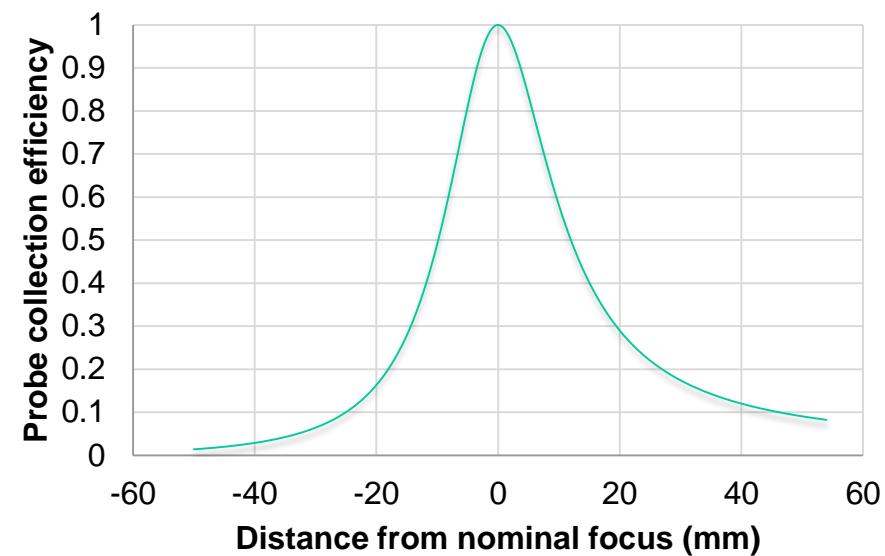
For this configuration:

- Focal point slews between 75 mm and 120 mm from relay lens.
- Demonstration shots will use right-angle pellicle to protect fixed optical system

Expected optical performance

- Predictions are for 5-10 dB rejection of objects 20 mm from nominal focus
- Should improve tracking of surface behind ejecta

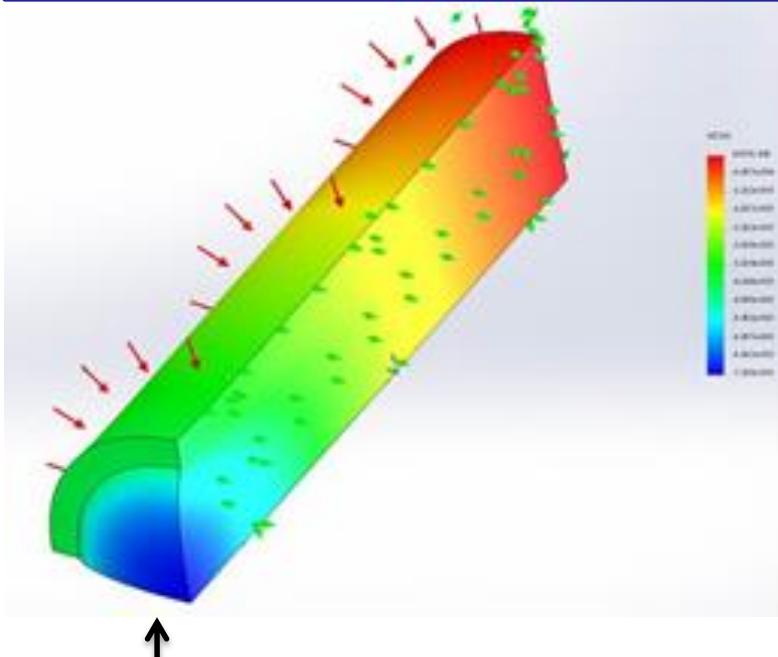
Predicted collection efficiency



Optically Actuated Lens

- Objective: actuate dynamic lensing *through the fiber*

100 μm polymer fiber in steel tube

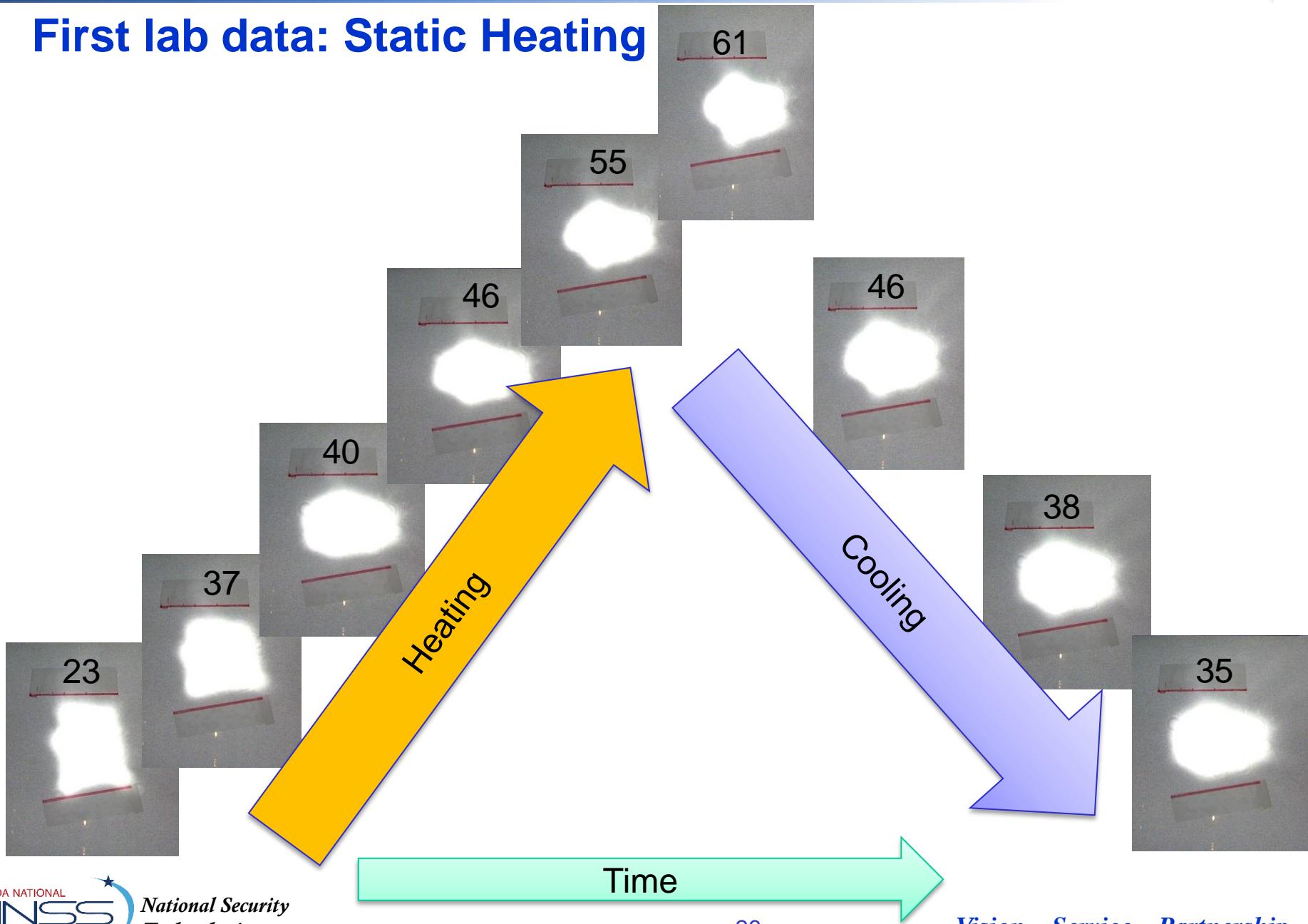


Polymer expansion ($+30^\circ\text{C}$) causes deformation and end. Focal length ~ 40 mm

Next steps:

- Test simulation predictions:
 - Measure optical focusing
 - Measure physical deflection
- Custom fiber-draw with doped polymer
- Assemble dynamically actuated probe
- Test on dynamic shots

First lab data: Static Heating



Physical feasibility

- Volume $\sim 1 \text{ mm}^3$
- Mass $\sim 1 \mu\text{g}$
- Energy to heat by $10^\circ \text{ C} \sim 10 \mu\text{J}$
- Power in $10 \mu\text{s} \sim 1 \text{ W}$

- Use current-pulsed, high-power laser diodes
- Use cladding-pumped fiber to deliver pump + signal

Dynamic-refocus: Next steps

- TAG lens
 - Assemble optical test bench in lab to quantify dynamic-range improvement during resonant operation
 - Begin designing dynamic experiments
- Optically actuated lens scheme
 - Mechanical modeling of larger polymer fibers
 - Verification of mechanical modeling: optical and physical
 - Begin considering pump and absorber system that could create the lensing desired

Summary: Places to gain dynamic range

- We are within a few dB of the shot-noise limit with back-end hardware
- Where can we squeeze more dynamic range?
 - Modulate launch power
 - More launch photons → more DR
 - May still run up against backscatter and probe power-handling limitations
 - Dynamically refocusing probe optics could improve collection over a wider range of probe-surface distances
 - Potentially useful in discrete-probe configurations
 - Balanced detection
 - Need better selection of lab-friendly receivers
 - 3 dB more signal (using 50/50 combiner)
 - More efficient use of LO power
 - Rejection of common-mode power swings

Backup slides

Shot-noise limit (Optical power)

$$SNR \equiv \frac{\langle i_s^2 \rangle}{\langle i_n^2 \rangle} = 1 = \frac{\eta P_s}{h\nu * B}$$

$$P_{shot-noise-limit} = \frac{B * h\nu}{\eta} = \frac{B * 1.28 \times 10^{-19} J}{\eta} = \frac{B}{1 \text{ MHz}} \frac{1.28 \times 10^{-10} mW}{\eta}$$

$$P_{shot-noise-limit} = 10 \log \left(\frac{B}{1 \text{ MHz}} * \frac{1}{\eta} \right) - 99 \text{ dBm}$$

SNR relationship to Effective-Number-of-Bits (ENOB)

Normalize for fraction of full-scale used

SNR increase by using frequency-domain analysis

$$SNR_{f,dB} = (6.02 \times ENOB) + 1.76 + 20 \log \left(\frac{2A}{V_{FS}} \right) + 10 \log \left(\frac{N_{FFT}}{2} \right)$$

ENOB = effective bits for digitizer, V_{FS} = full scale voltage, A = RMS amplitude of applied signal
 (See Wiley Encyclopedia of Electrical and Electronics Engineering, Vol. 18, J. Blair)

“Frequency-domain Number of Bits”

Number of bits as function of SNR, fraction of full-scale, FFT Points

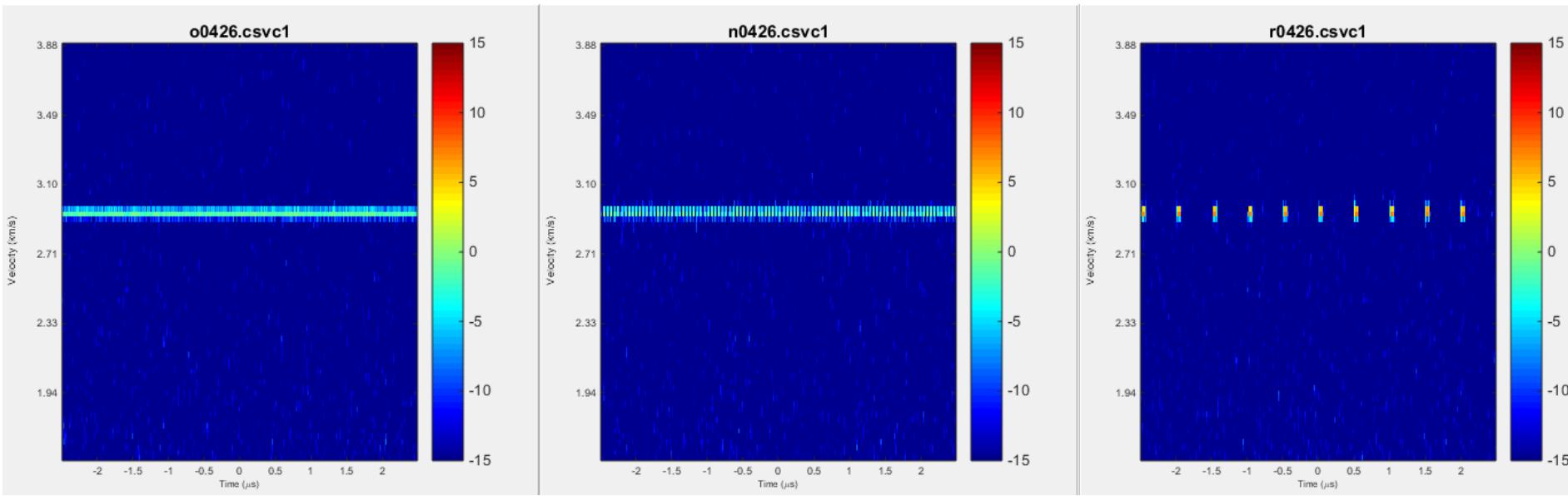
$$FNOB = \frac{1}{6.02} \left[SNR_{f,dB} - 1.76 - 20 \log \left(\frac{2A}{V_{FS}} \right) - 10 \log \left(\frac{N_{FFT}}{2} \right) \right]$$

“signal” cancels out...

$$FNOB = \left(\frac{1}{6.02} \right) * \left[10 \log \left(\frac{V_{FS}^2}{50\Omega} * 1000 \right) - noise_{dbm} - 7.78 - 10 \log \left(\frac{N_{FFT}}{2} \right) \right]$$

... Just a noise measurement! Measure with receiver on, LO power at nominal.

Full-time (5 μ s) spectrograms of free-running modulation



Advantages to balanced receiver

- For deep-time MPDV, no problems when LO switches
- Don't throw away LO or signal photons
 - 3 dB signal gain for 50/50 combiner
 - Not a problem if you are already using 90/10 combiners
- ASE-ASE is common-mode
 - Not commonly a problem
 - In balanced receiver, suppressed by 20-30 dB