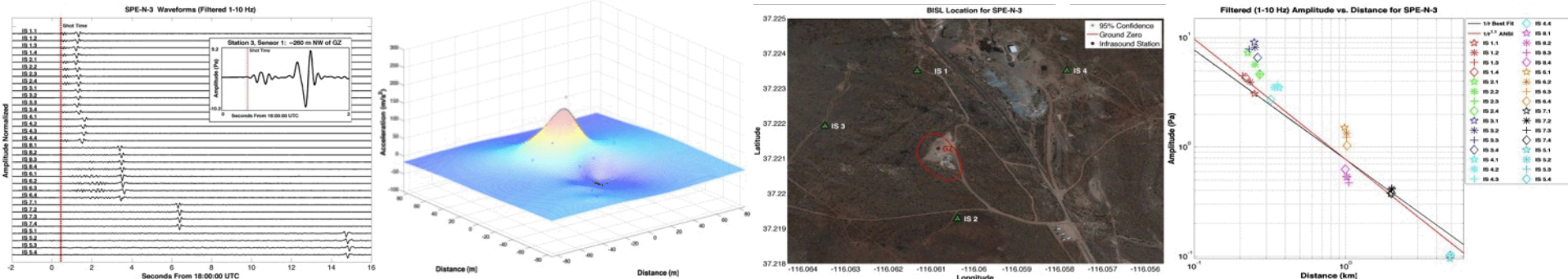


Exceptional service in the national interest



New Sensors, New Sources: The future of infrasound monitoring

Kyle R. Jones – Sandia National Labs

RMR: June 19, 2014



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Roadmap

- Near-Source Infrasound is an important part of understanding signals detected at greater distances
- Several current projects related to source physics
 - The Source Physics Experiment
 - Humming Wombat
- Limitations and Challenges
- What does this mean for our science and what are the next steps

Near-Source Infrasound is an
important part of understanding
signals detected at greater distances

Motivation

- The ability to detect and monitor for various sources both man-made and natural
 - Nuclear and chemical explosions
 - Artillery
 - Aircraft and rockets
 - Bolides
 - Tornadoes
 - Hurricanes
 - Wind Farms
- Not always able to deploy within the “near-source” region
- Must be able to make determinations about the source after propagation

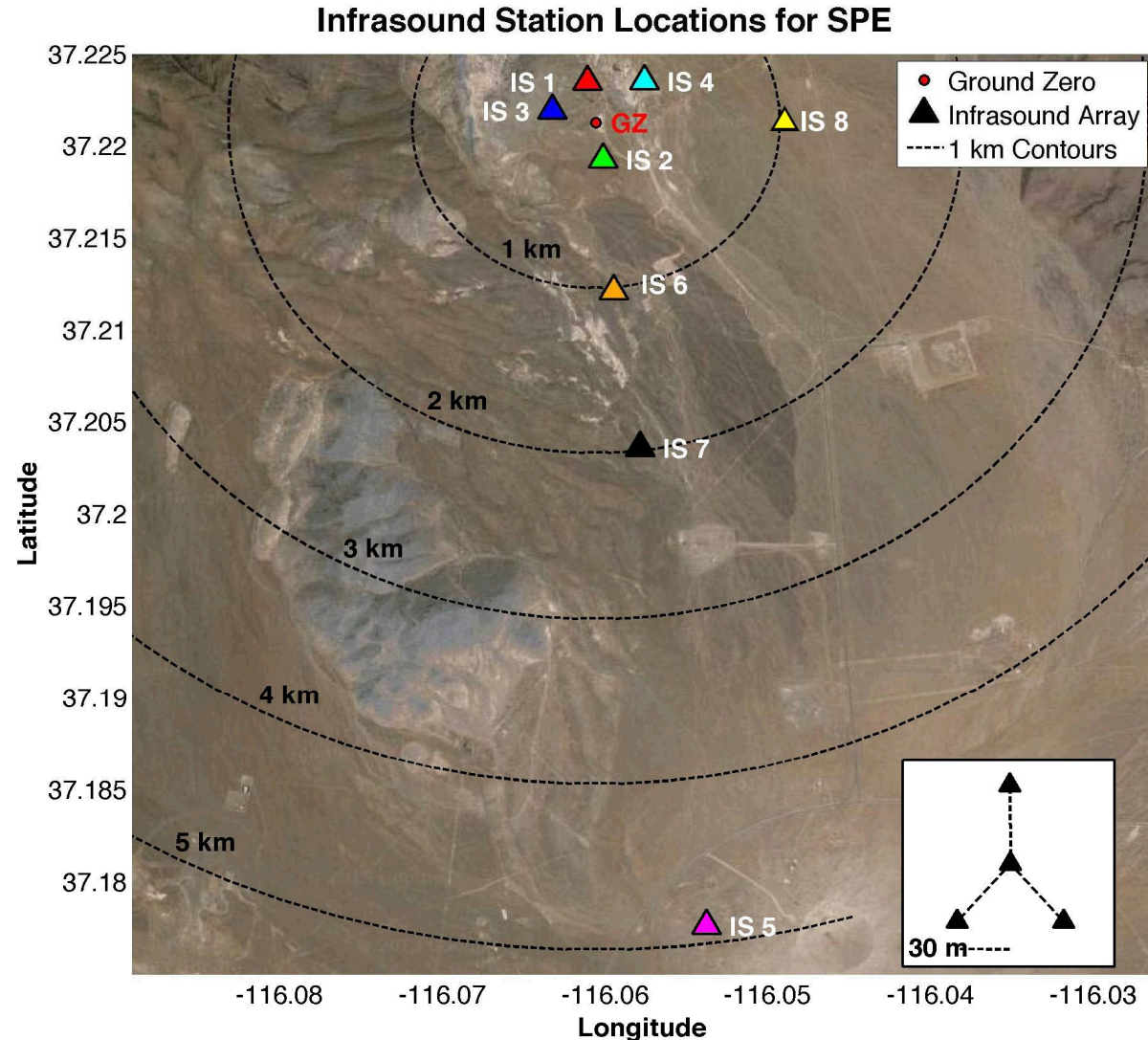
Several current projects related to
source physics:

The Source Physics Experiment

Humming Wombat

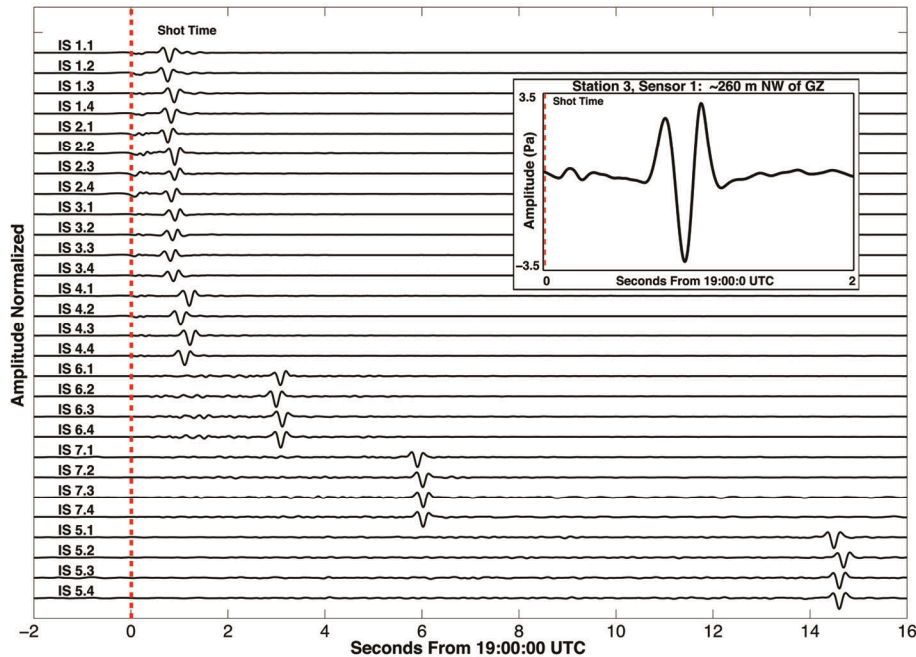
The Source Physics Experiment

- Northeast corner of the Nevada National Security Site (NNSS)
- Various underground explosions in the same source hole
- Network of 8 infrasound arrays consisting of 32 Hyperion sensors
- Additionally, surface and borehole accelerometers, geophones, and broad-band seismometers

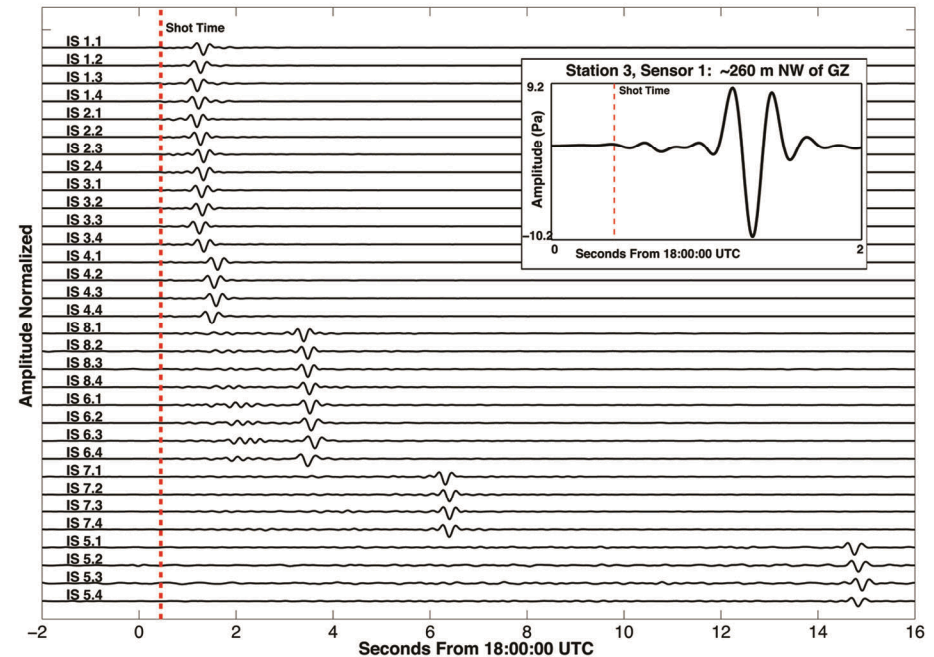


The Source Physics Experiment

SPE-2 Waveforms (Filtered 1–5 Hz)

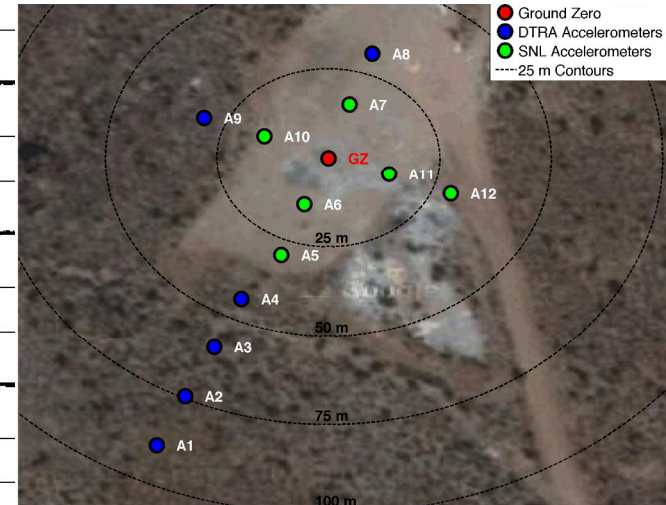
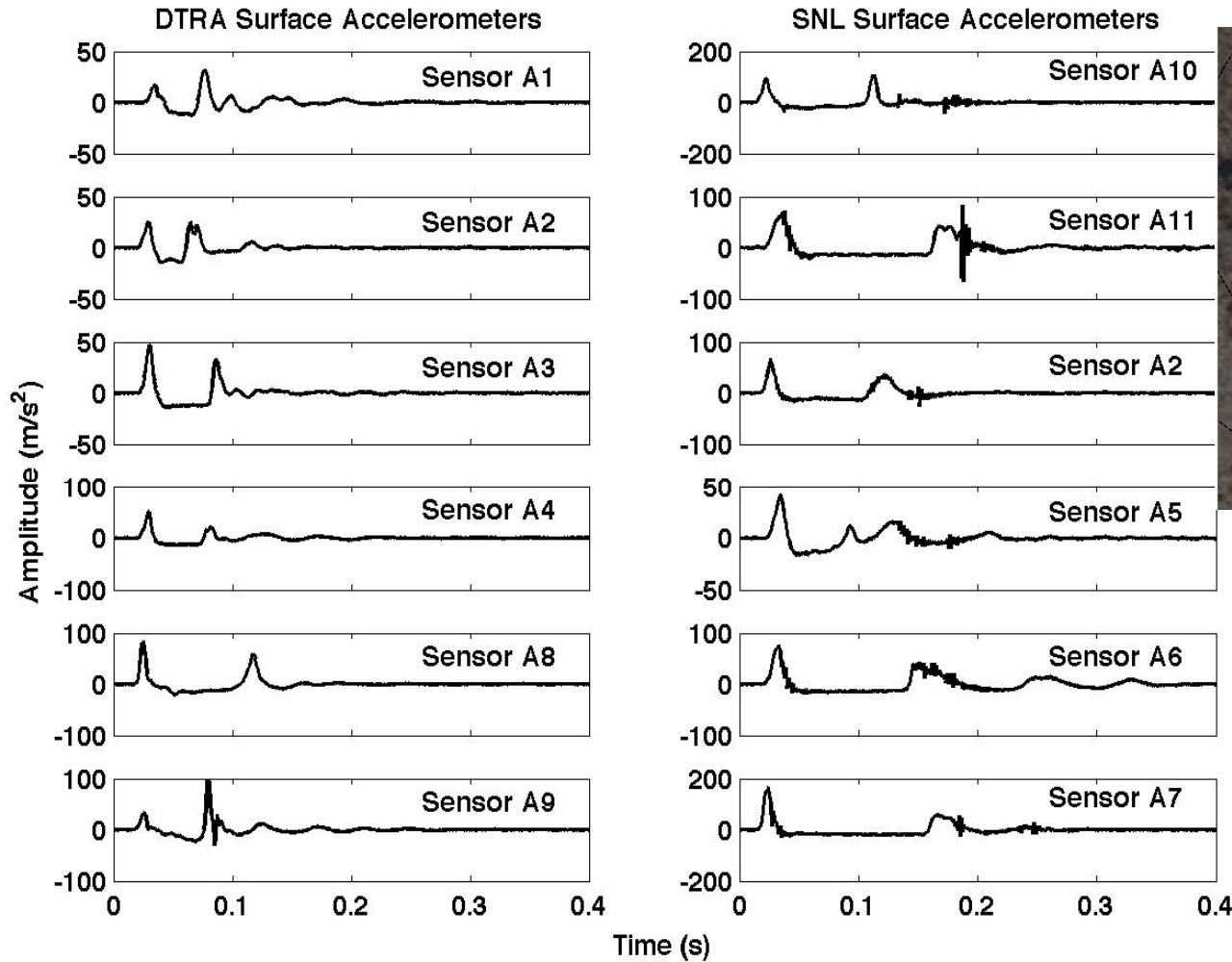


SPE-3 Waveforms (Filtered 1–5 Hz)



The Source Physics Experiment

- Use data from surface accelerometers to get ground motion



DTRA Accelerometers

A1, A2, A3, A4, A8, A9

Endevco 500g

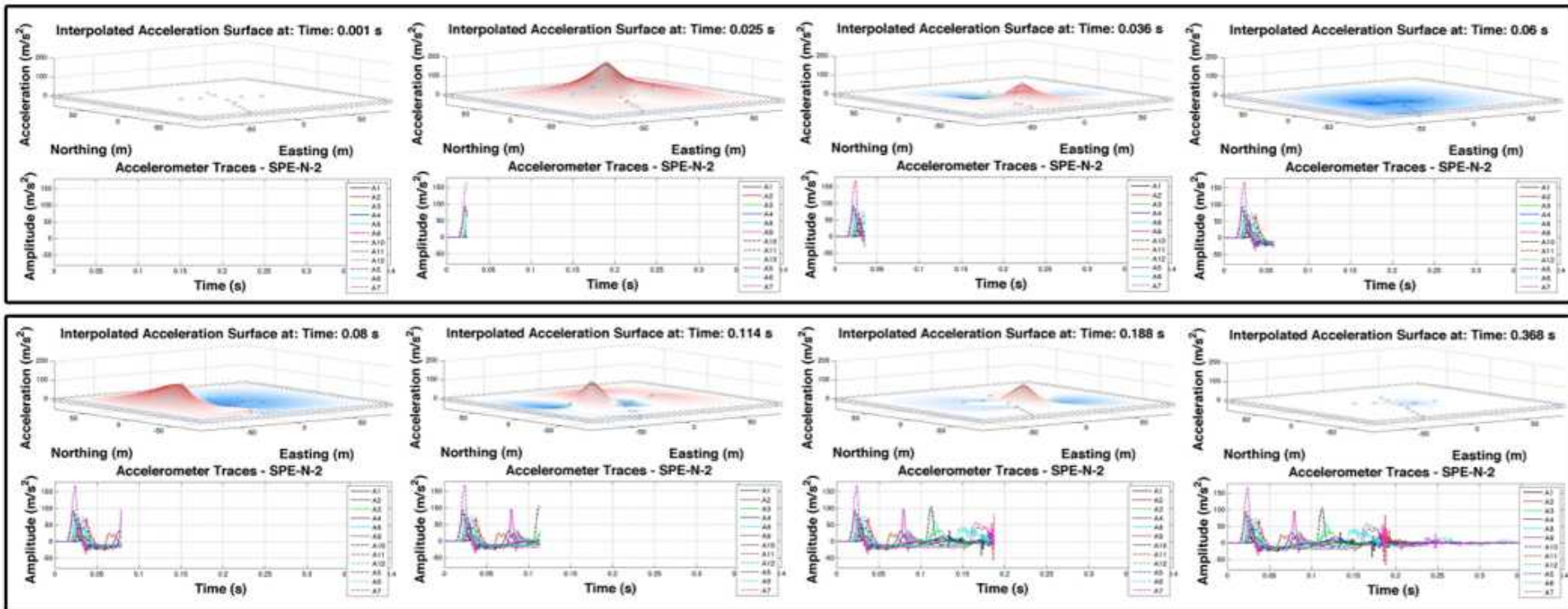
SNL Accelerometers

A5, A6, A7, A10, A11, A12

Endevco 500g

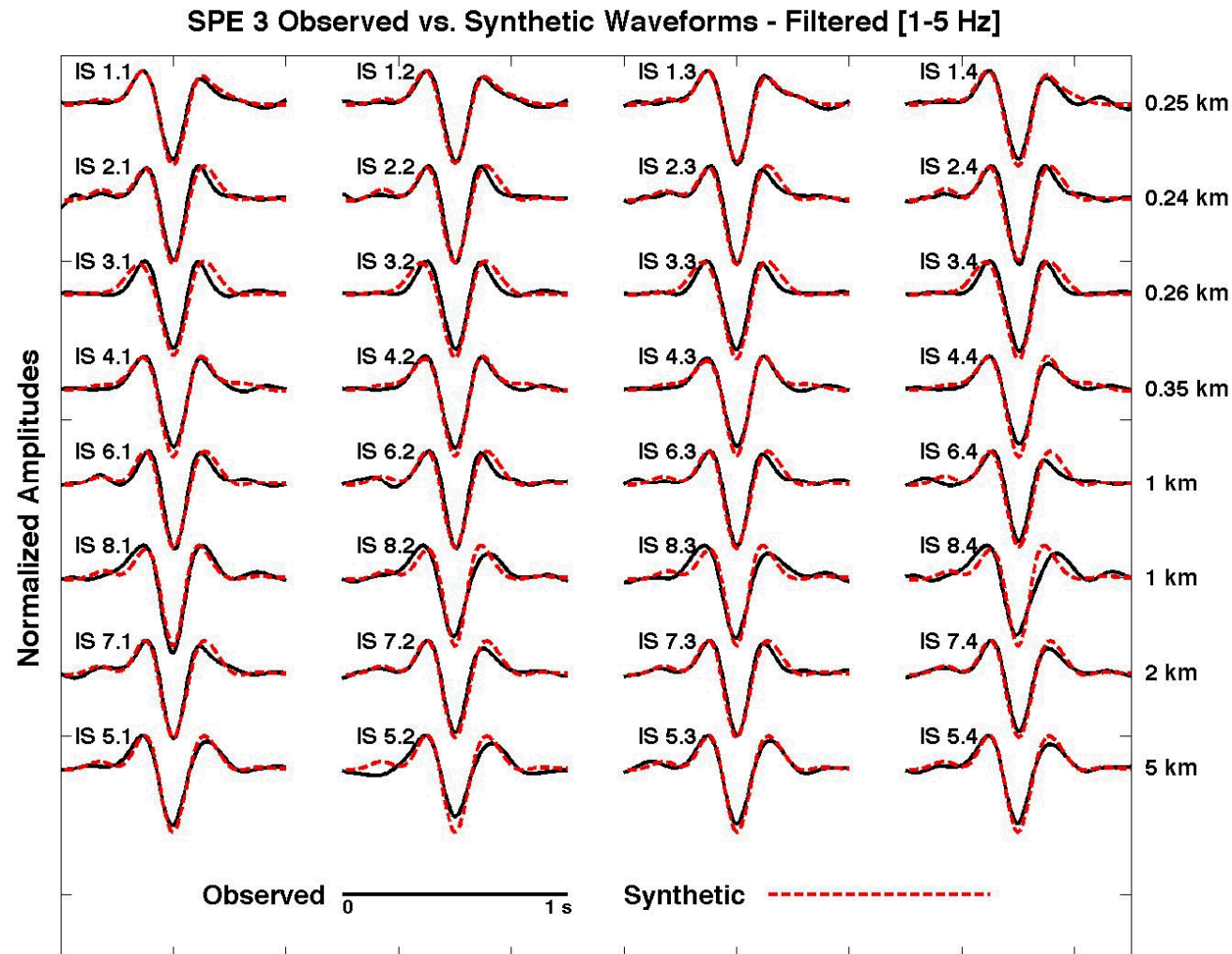
The Source Physics Experiment

- Set up a grid and interpolate/extrapolate the accelerometer data



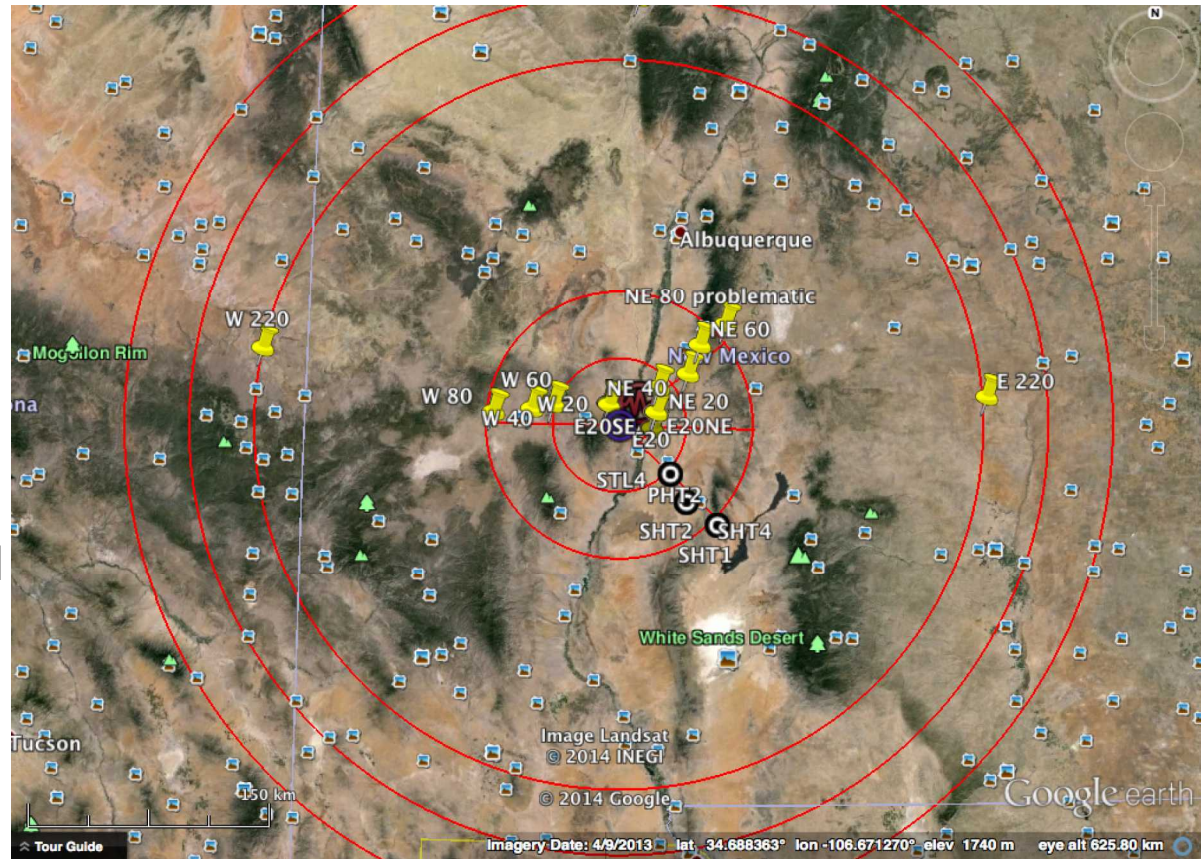
The Source Physics Experiment

- Integrate the pressures at each grid point over both area and time to produce synthetic waveform
 - The final summation step takes into account the constructive and destructive interference inherent in such a large source area.
 - Can be thought of a baffled piston in an infinite plane



Humming Wombat

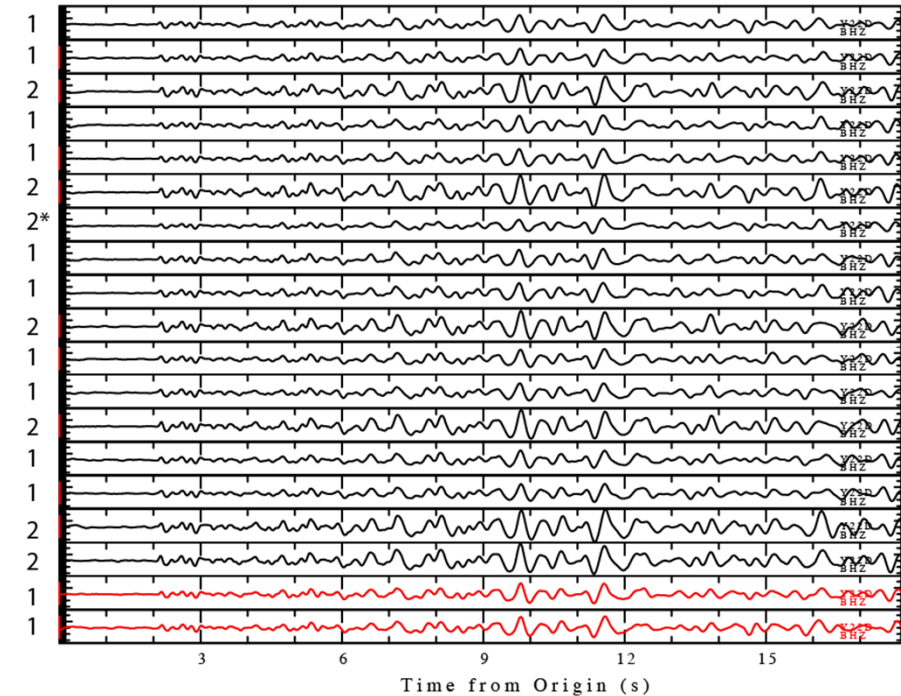
- Central New Mexico at the Energetic Materials Research and Testing Center at New Mexico Tech
- Similar surface explosions over several seasons
- Large network of infrasound blast sensors, near-field and far-field arrays
- Additionally, broadband seismometers and other equipment



Humming Wombat

Motors

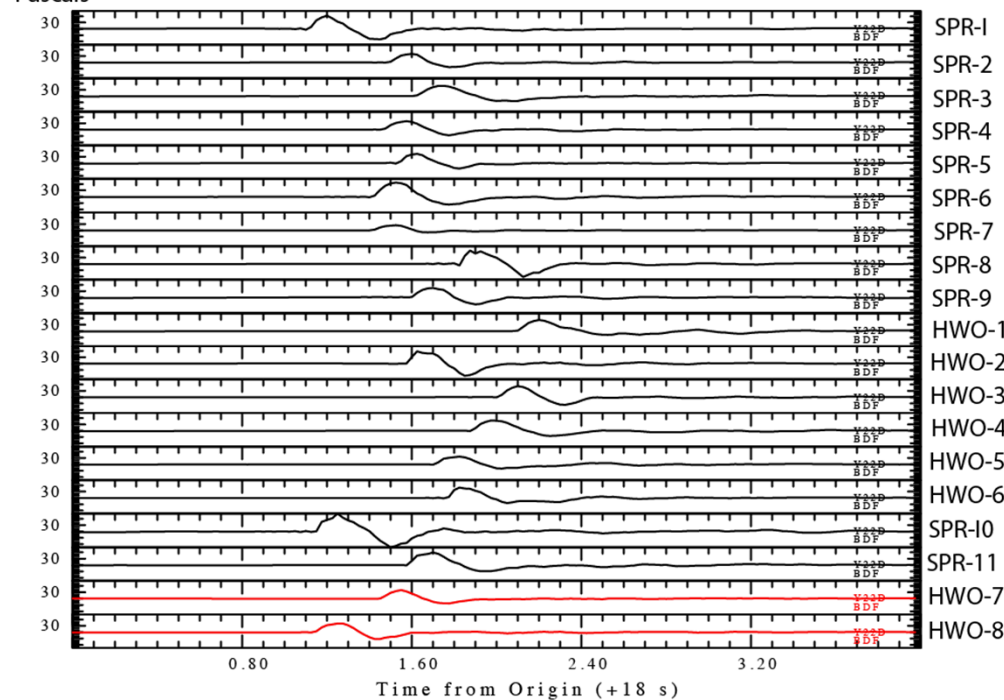
SEISMIC



Station Y22D; 6.68 km from GZ

Pascals

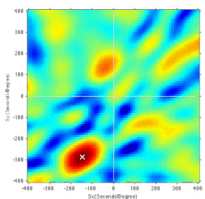
ACOUSTIC



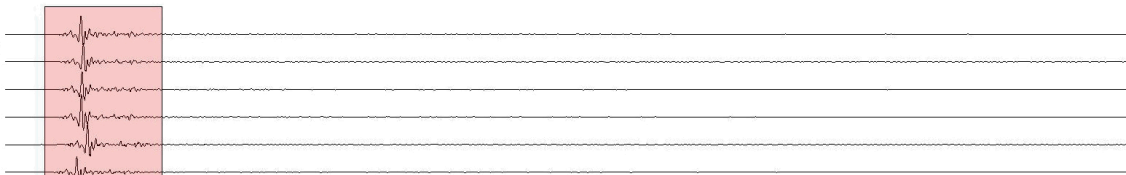
- Sensor network includes:
 - Infrasound
 - Seismic
 - Blast Sensors
- Close in measurements out to hundreds of km
- Surface Shots

Humming Wombat

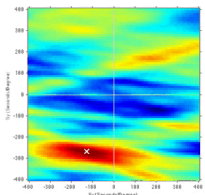
Humming Wombat Explosion Signals Detected at Sandia FACT Array 110 Km at 206 Degrees Back Azimuth



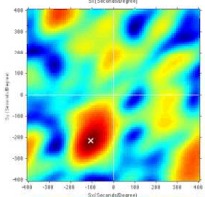
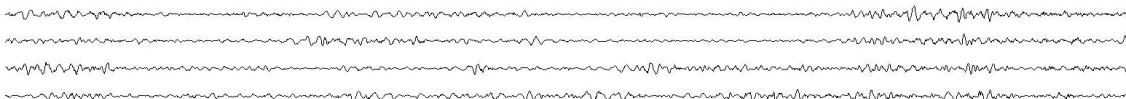
HWO 1-1 Double Motor
12/11/2013 - 22:26:10 UTC
Back Azimuth: ~207°
Trace Velocity: ~343 m/s
Peak to Peak: 19.6 Pa
Traces Magnified 0x



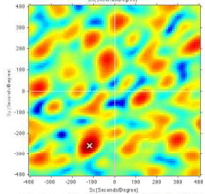
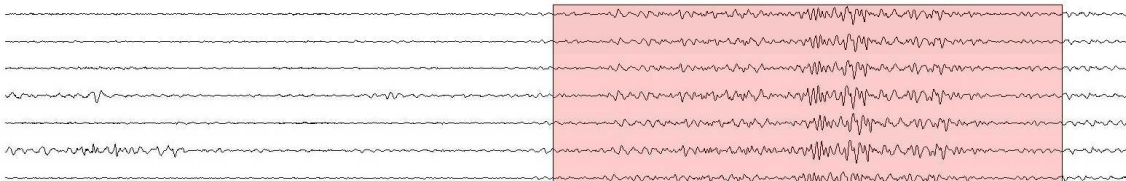
HWO 1-2 Single Motor
1/10/2014 - 20:33:04 UTC
No Event(s) Detected
Traces Magnified 10x



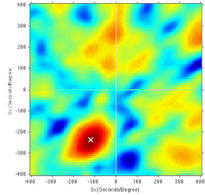
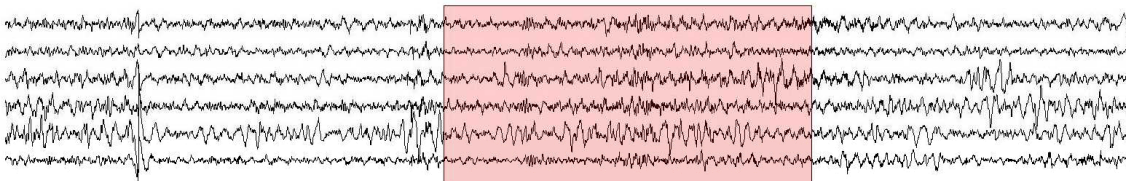
HWO 1-3 Single Motor
1/16/2014 - 20:56:41 UTC
Back Azimuth: ~203°
Trace Velocity: ~342 m/s
Peak to Peak: 0.16 Pa
Traces Magnified 100x



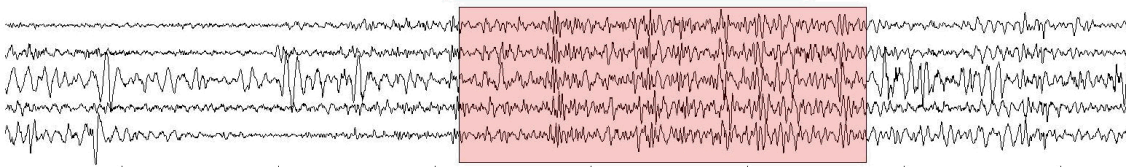
HWO 2-1 Double Motor
3/5/2014 - 22:07:46 UTC
Back Azimuth: ~206°
Trace Velocity: ~374 m/s
Peak to Peak: 0.55 Pa
Traces Magnified 25x



HWO 2-2 Single Motor
3/6/2014 - 20:54:19 UTC
Back Azimuth: ~204°
Trace Velocity: ~392 m/s
Peak to Peak: 0.20 Pa
Traces Magnified 200x



HWO 2-3 Single Motor
3/10/2014 - 20:09:34 UTC
Back Azimuth: ~207°
Trace Velocity: ~369 m/s
Peak to Peak: 0.24 Pa
Traces Magnified 100x



320

340

360

Seconds From Shot Time

380

400

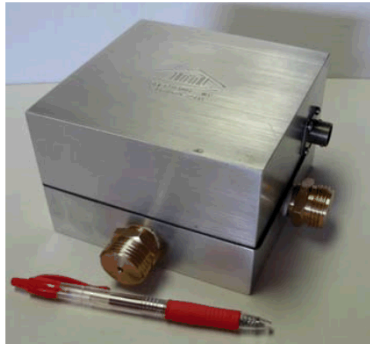
420

440

460

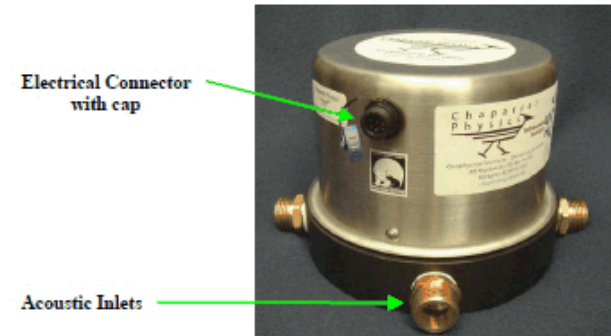
Sensor Tech

Inter-Mountain Labs (IML-ST)



Inter-mountain Labs. 555 Abraska, Sheridan, WY 82801
(307) 674-7506 <http://www.intermountainlabs.com/>

Chaparral (CP25)



http://bric.postech.ac.kr/upload/geditor/201010/0.15810600_1286701047.jpg

CEA - MB2000



Seismically decoupled
versions available or
soon to be

Very near-source
overpressure sensors

http://www-dase.cea.fr/public/dossiers_thematiques/microbarometres/images/mubaro.gif

Hyperion (IFS3000)



Hyperion Technology Group, Inc., 3248 West Jackson Street, Tupelo, Mississippi 38801
662.823.0600 <http://www.hyperiontechgroup.com>

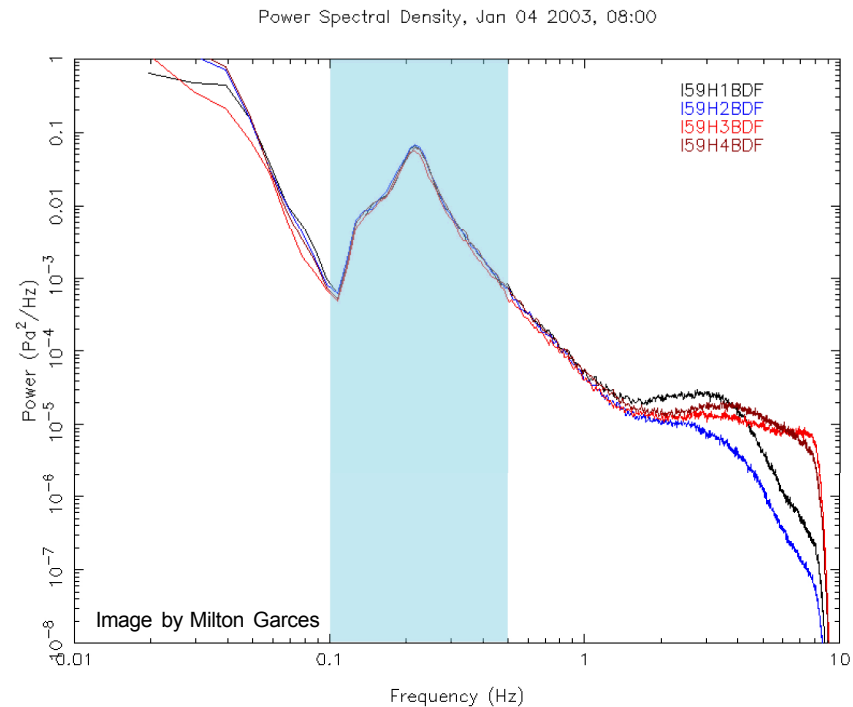
Detectability Issues

- Detectability depends on many factors

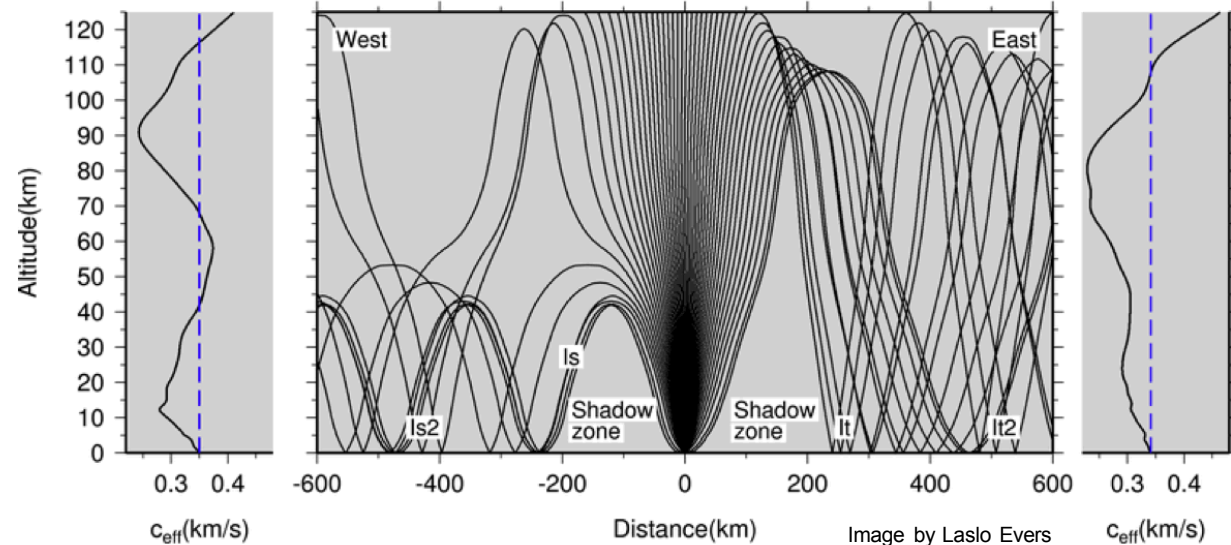
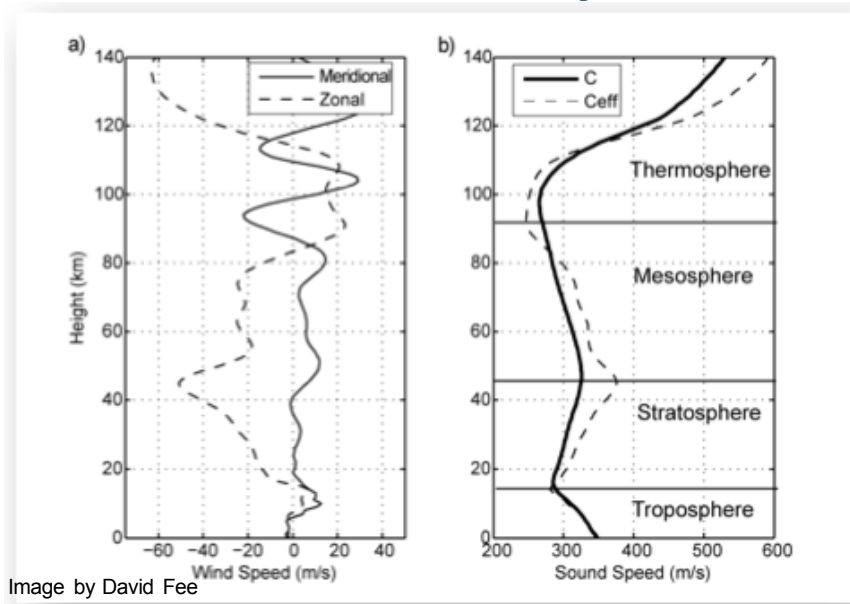
- Yield/Size of the event
- Above vs. below ground
- Frequency range
- Microbaroms (0.1 – 0.5 Hz)

- WIND!!!!

- Wind reduction systems

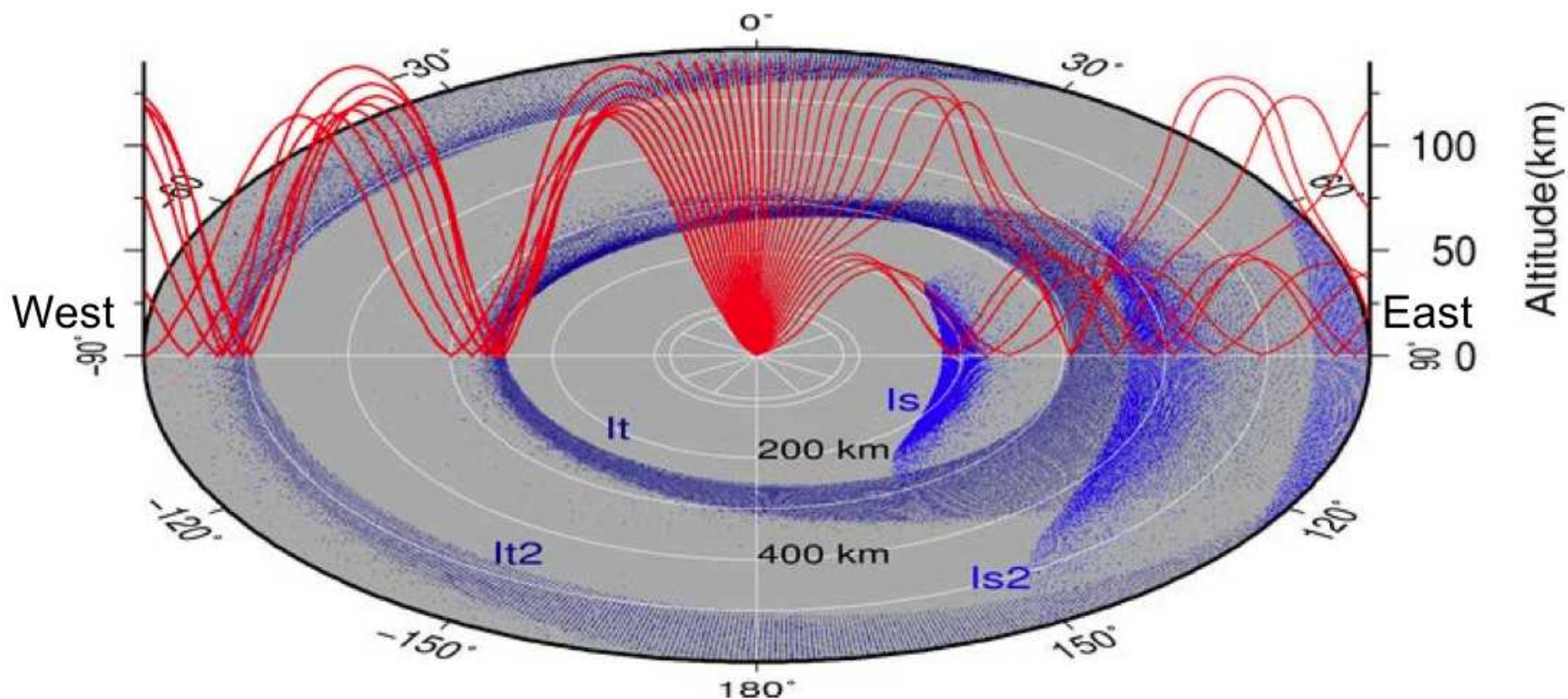


Detectability Issues



Detectability Issues

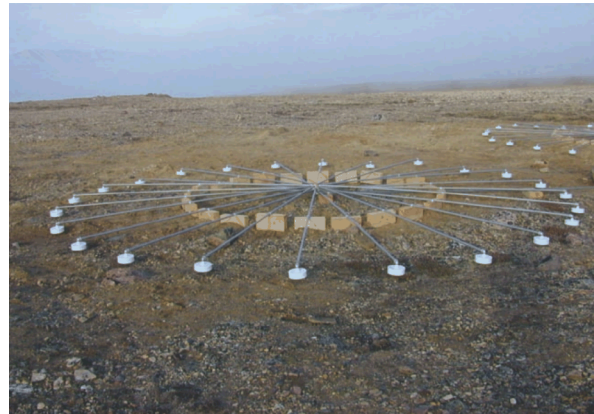
Raytracing Northern Hemisphere winter



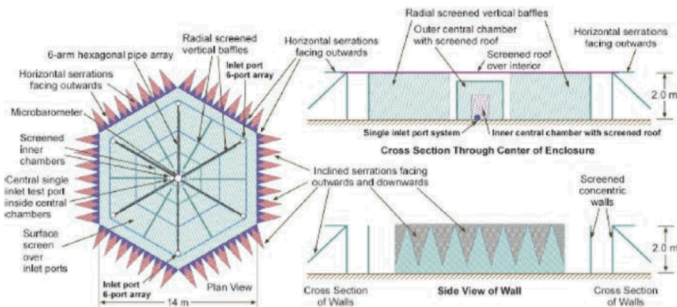
Wind Reduction

Current State of Windscreens

- IMS Pipe Rosette
- Porous Hoses
- IRIS “Puck” in Gravel
- Wind Fences



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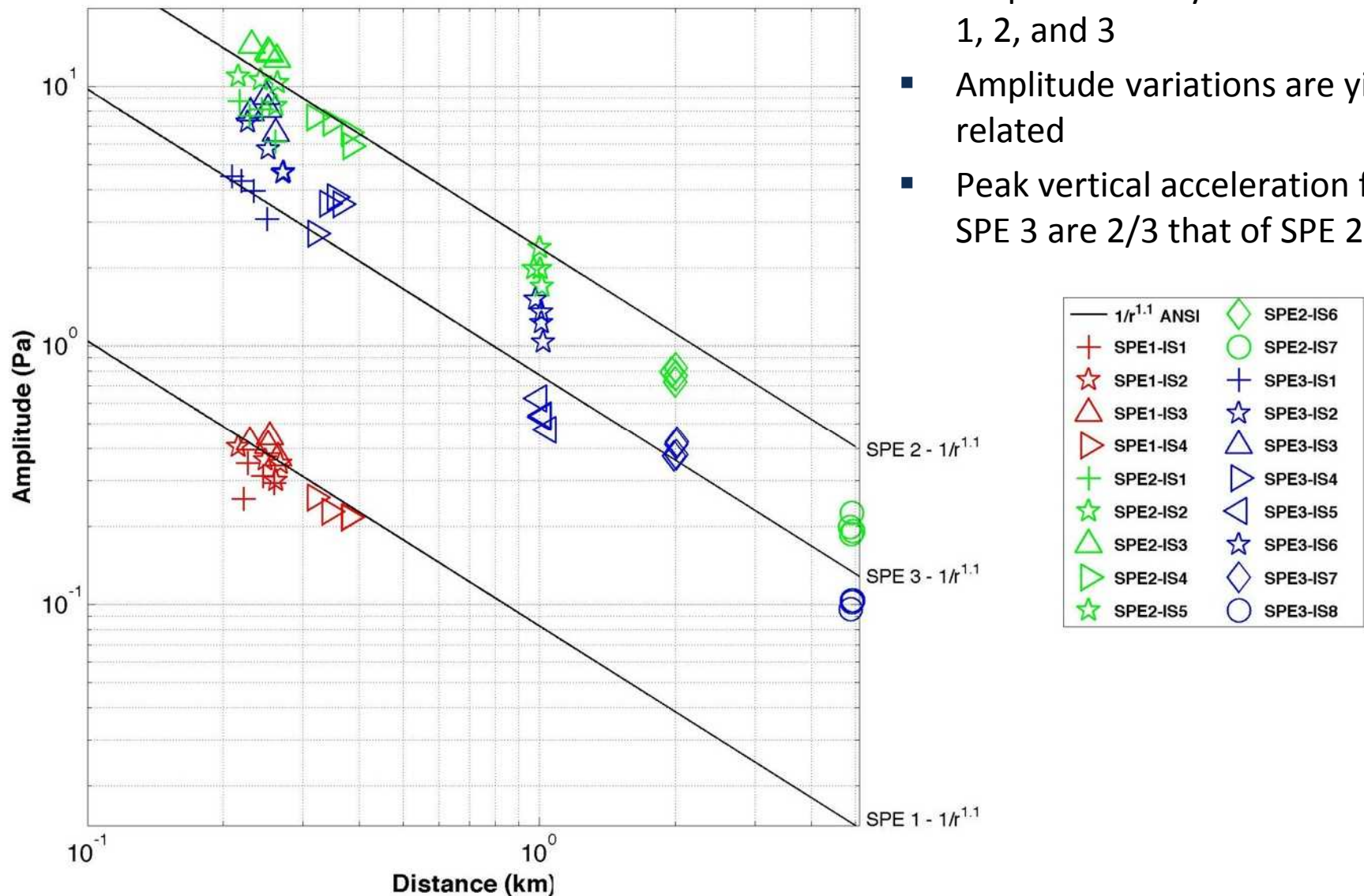
FACT SITE



FACT SITE

Amplitude vs. Distance

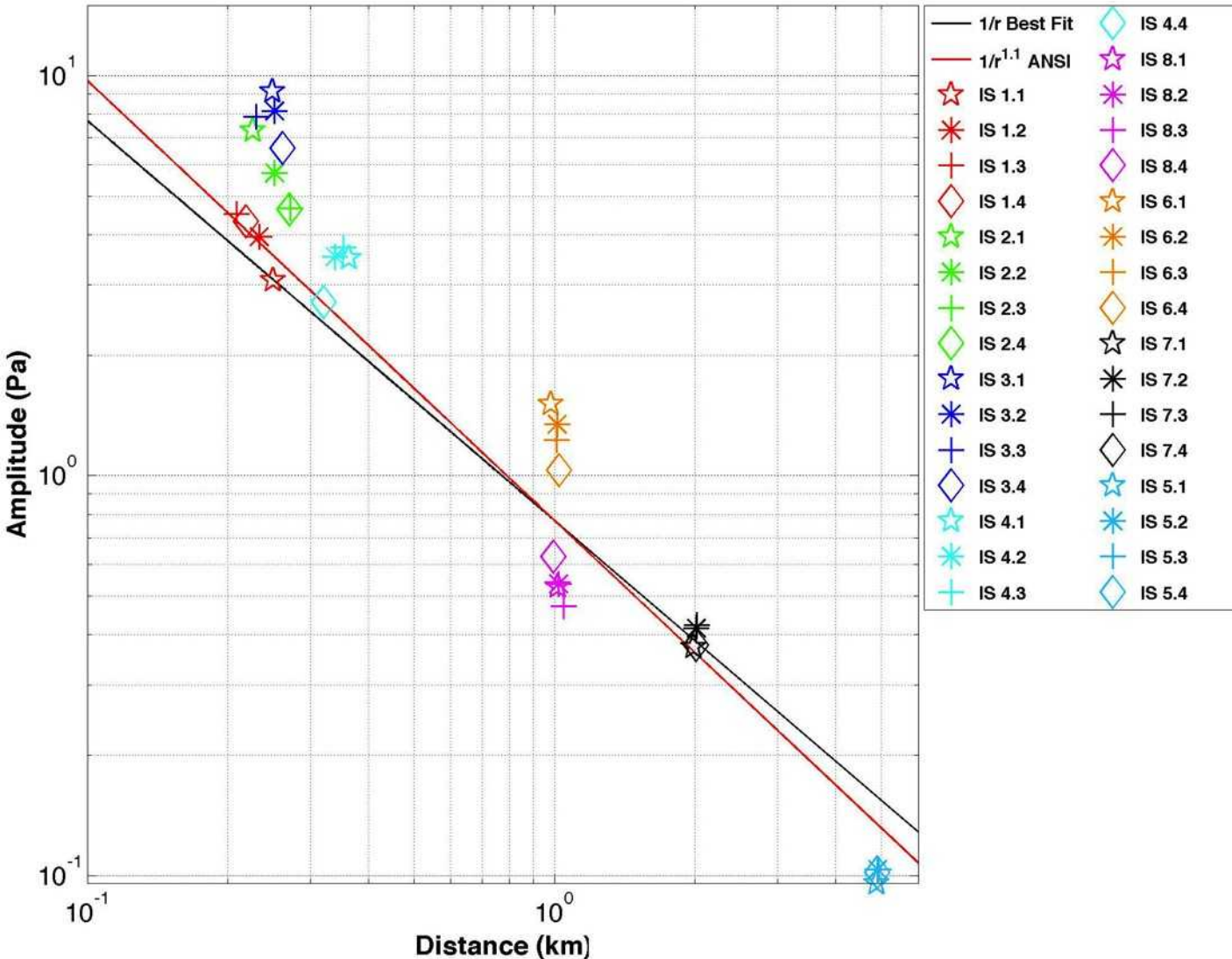
Filtered (1-10 Hz) Amplitude vs. Distance for SPE-N-1, 2, & 3



- Amplitudes vary between SPE 1, 2, and 3
- Amplitude variations are yield related
- Peak vertical acceleration for SPE 3 are 2/3 that of SPE 2

Amplitude vs. Distance

Filtered (1-10 Hz) Amplitude vs. Distance for SPE-N-3



- Amplitude decays linearly
- Reasons for amplitude differences:
 - Off-axis vs. On-axis pressures observations
 - Line of sight
 - Wind reduction issues

Path Forward

- Investigate whether or not porous hoses are necessary at these distances
 - Porous hoses attenuate amplitude and add a level of complication to analysis
 - Deteriorate in the sun, leading to porosity changes

- Account for wind in the short propagation to determine if it improves solution
 - This should be investigated but may prove not to matter

- Continue to model the source using infrasound to better match the shape of the observed wave

Questions?