

# Developing a kinetic approach to radiation transport and its interaction in He/N<sub>2</sub> ionization waves

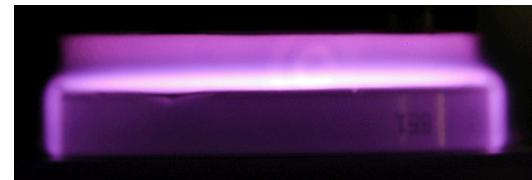
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Sandia National Laboratories, Albuquerque NM



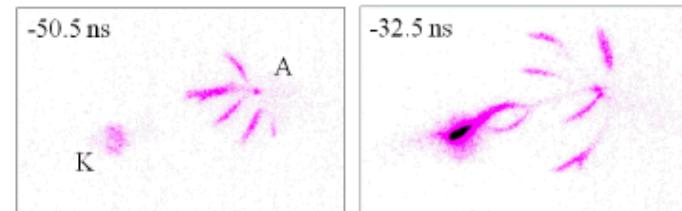
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# Motivation

- Interested in transient response times of plasma formation as well as steady quantities which may be sensitive to photonic processes.
- Light emission is one of the defining characteristics of plasma discharges. Another tool to make comparisons to experiments for validation purposes.
- Would like to begin to quantify the effect of self-produced radiation on plasma development and its secondary effects.
- Develop a method to discretely model photons in a kinetic code that allows for the incorporation of energy dependent photo-processes.



- Pulsed helium discharge operating at 30 torr (A. Fierro, E. Barnat)



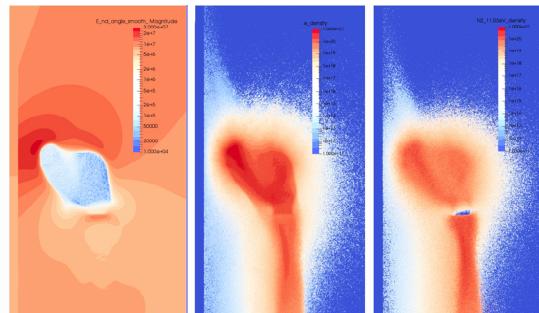
- Fast-gated imaging of streamer propagation.

A. Fierro, *et al.*, J. Phys. D: Appl. Phys., 2012.

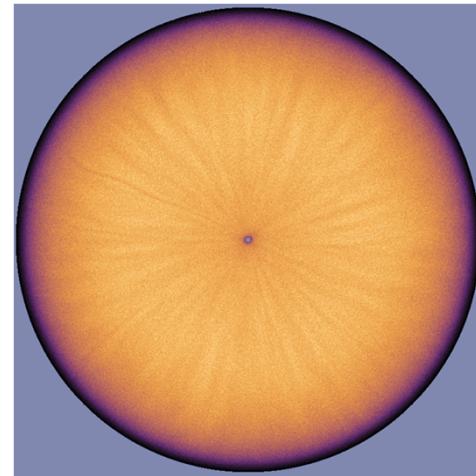
A. Fierro, *et al.*, Plasma Sources Sci. Technol., 2017, to be submitted.

# The Kinetic Code - *Aleph*

- Unstructured FEM (compatible with CAD)
- Massively parallel (scales up to 100k proc.)
- Hybrid PIC + DSMC, also PIC-MCC
- Electrostatics with fixed B Field
- Advanced surface (electrode) models
- Collisions, charge exchange, chemistry, excited states, ionization
- Photon transport, photoemission, photoionization
- Advanced particle weighting methods
- Dynamic load balancing restart (with all particles)



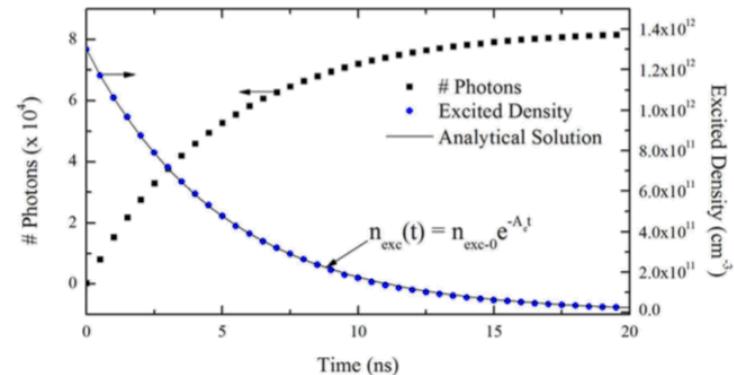
- Simulation of streamer formation along a dielectric surface (C. Moore).



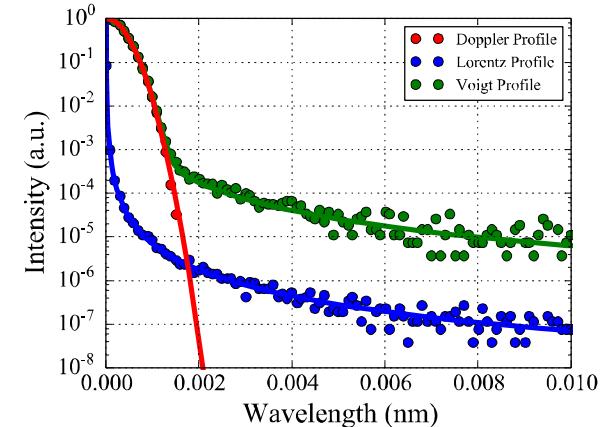
- 2D simulation of a Langmuir probe in the electron saturation regime demonstrating unexpected streaming instabilities (B. Yee).

# Radiation Transport Method

- $e^- + A \rightarrow A^* + e^-$ 
  - leads to an accumulation of  $A^*$
- for each  $A^*$  evaluate:
  - $R < 1 - e^{-t/\tau}$ ,  $\tau$  = lifetime of  $A^*$
- If evaluated to be true
  - $A^* \rightarrow A + h\nu$
  - $|v_{ph}| = c$ ,  $v_{ph}$  = isotropic
  - $\lambda_{ph} = hc / (E(A^*) - E(A))$ 
    - Lorentz shape
$$\lambda_s = \tan[(R - 0.5)\pi] \cdot \Delta\lambda_r + \lambda_0$$
  - Doppler shape
- Each photon is pushed through the computational domain and interactions are handled with traditional DSMC procedures.



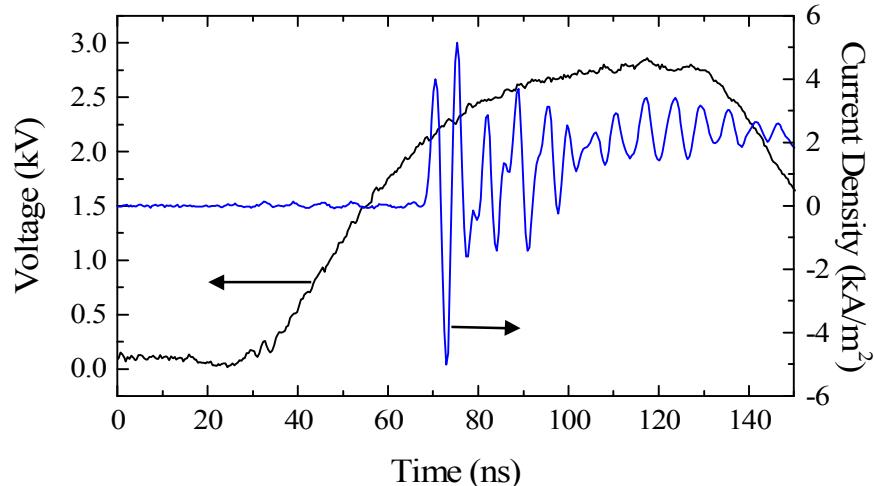
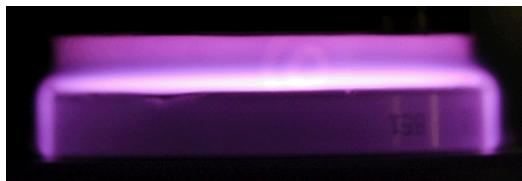
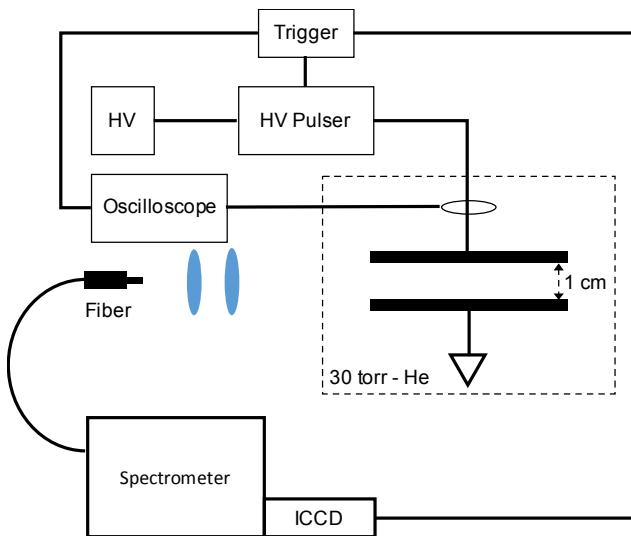
- Verification of the method.



- Simulated line profiles.

# Towards Validation

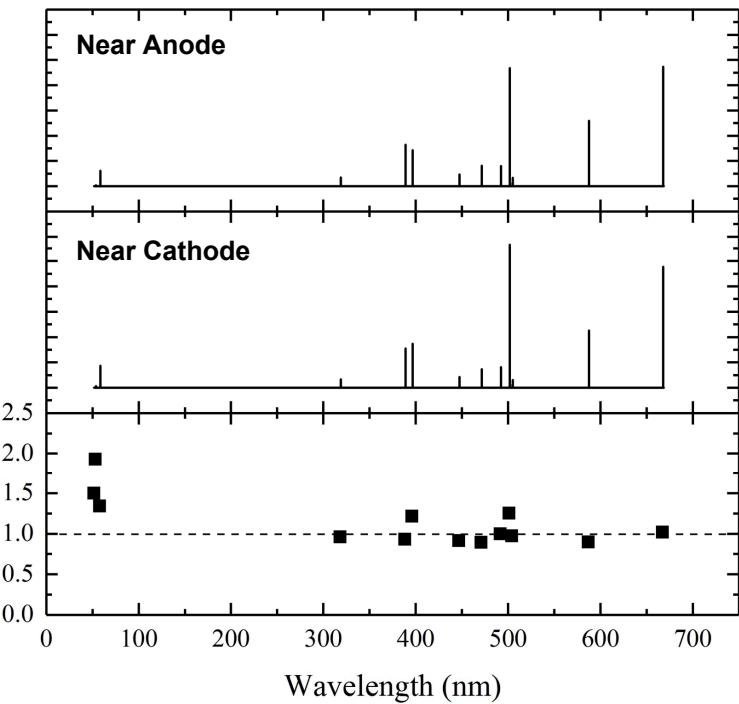
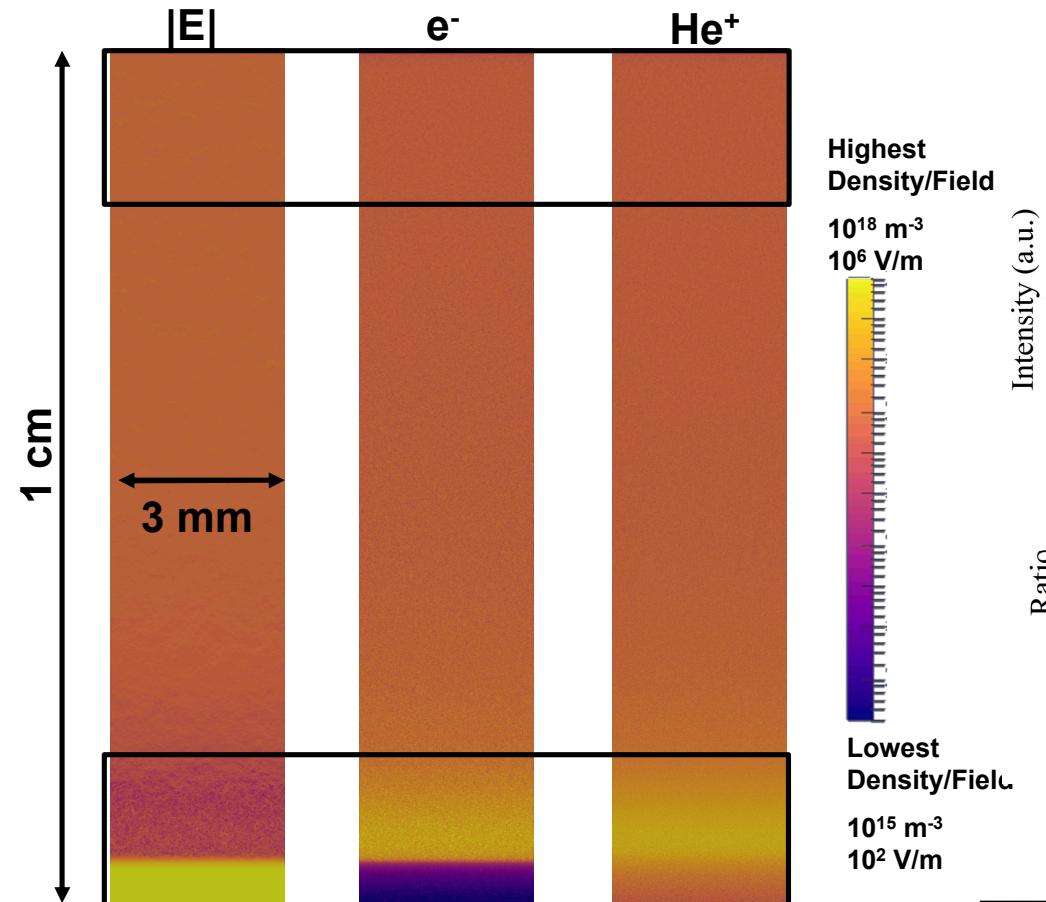
- With this method, we are able to generate spatially-resolved line emission spectra that does not assume LTE.
  - Can we begin to use emission spectra as a viable validation tool?



- Experimental setup focused on capturing time-resolved optical emission spectroscopy in the visible regime.

# Towards Validation

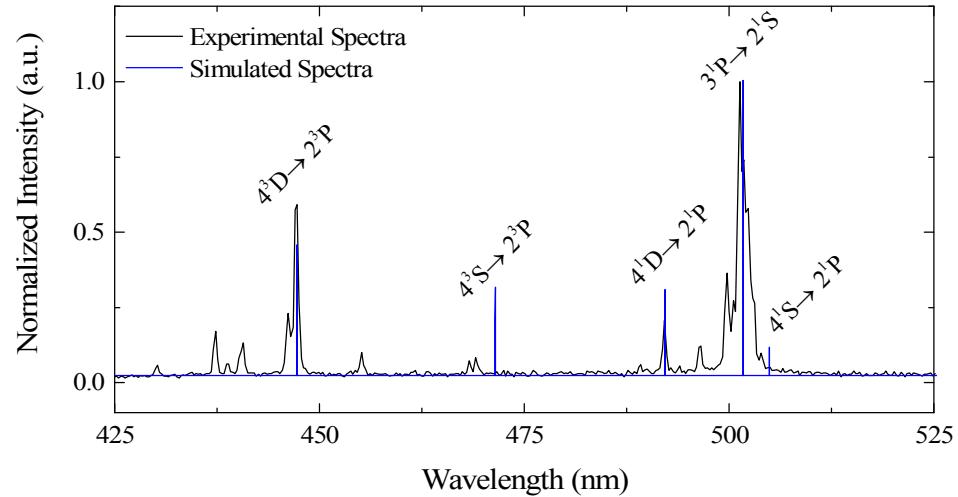
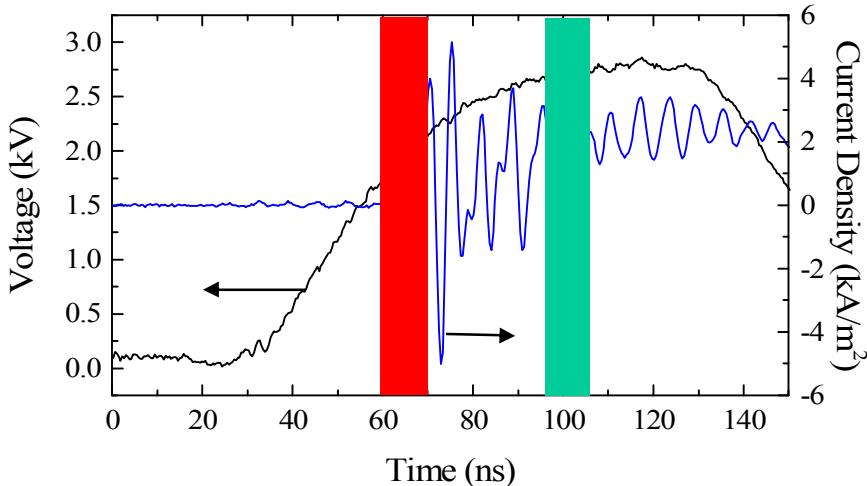
- With this method, we are able to generate spatially-resolved line emission spectra that does not assume LTE.



- Log contour plots from a simulation of pulsed plate-plate discharge,  $t = 60\text{ns}$

# Towards Validation

- Initial comparisons to experimental data
  - 10 ns camera gate, timing indicated by green window.
  - Simulation spectra taken at  $t = 60$  ns
  - Working on time correlation between experiment and simulation



- The simulation produces realistic line ratios for the He lines shown.
- Assume pure He gas, no impurities are included.

# Application to Large Scale Simulations

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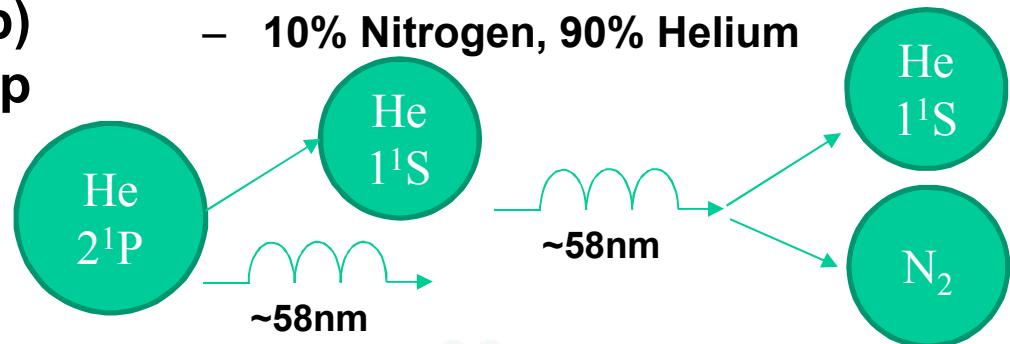
- How do photonic mechanisms impact near atmospheric pressure discharges? The goal is to simulate a three-dimensional system of a pulsed, near atmospheric pressure discharge.
  - How big of a system is capable of being simulated in *reasonable* time.
  - What assumptions can be made to alleviate computational requirements?
  - What are the numerical challenges (timestep, space step)?
  - What are the computational resource challenges (memory, processors)?
- Simulations at higher pressures are extremely challenging due to the anticipated large electron densities. This results in a smaller mesh to resolve Debye lengths and avoid numerical heating.

# Ionization Wave Simulation

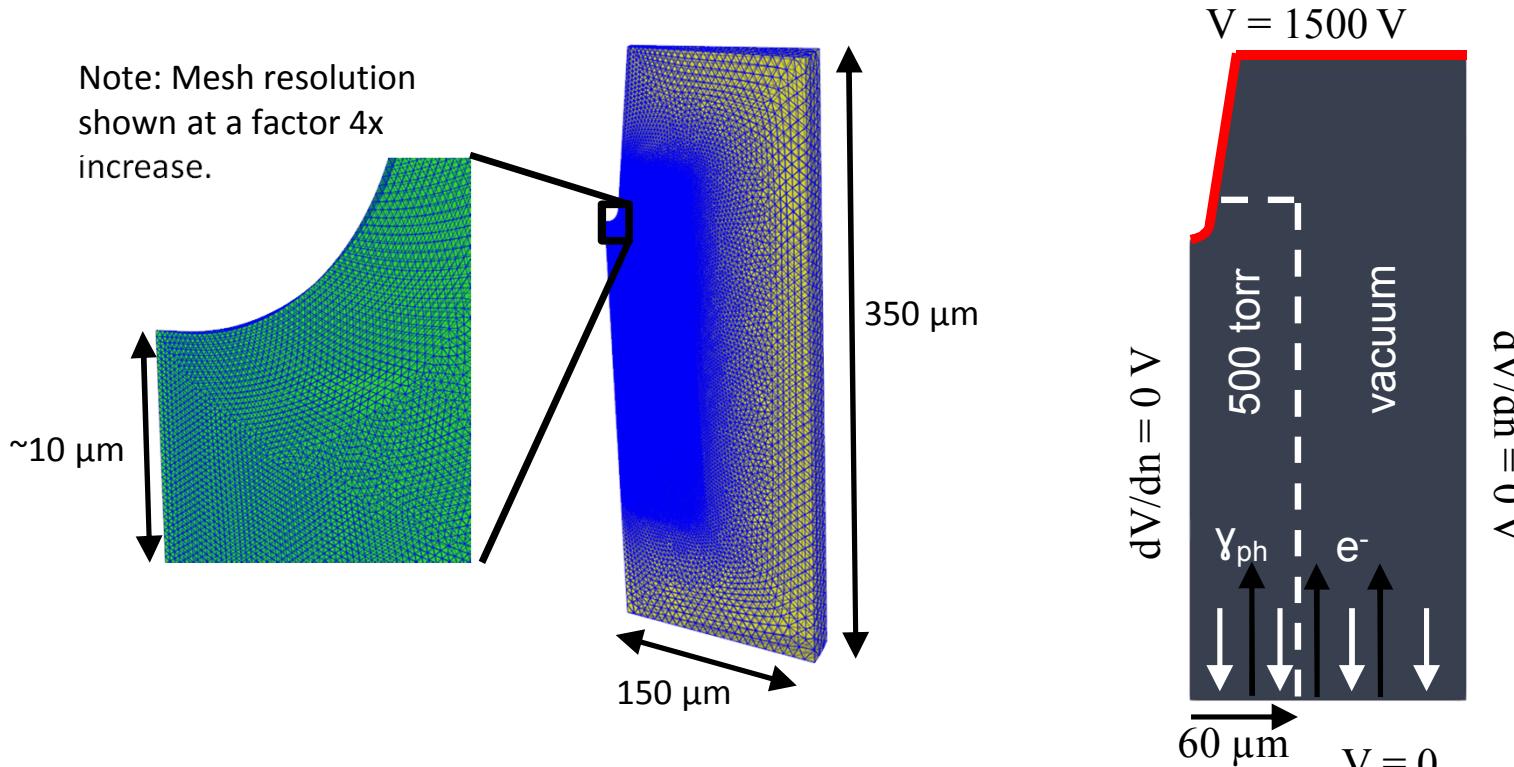
- Numerical parameters are guided by experiments of similar type discharges
  - 500 torr background pressure
  - $10^{15} \text{ cm}^{-3}$  electron density
  - 0.5 eV electron temperature
- Debye length and Photon CFL conditions (photon does not cross more than 1 element in a time step) set the minimum time step and spatial step.
  - $\text{dx} = 50 \text{ nm}$
  - $\text{dt} = 2 \times 10^{-15} \text{ s}$

Parameter	Spatial or Time Scale
Debye Length - $\lambda_D$	$\sim 150 \text{ nm}$
Electron Mean Free Path - $\lambda_{\text{mfp}}$	$\sim 200 \text{ nm}$
Photon Mean Free Path - $\gamma_{\text{mfp}}$	$\sim 25 \mu\text{m}$
Inverse plasma frequency - $(\omega_{\text{pe}})^{-1}$	$\sim 1 \times 10^{-13} \text{ s}$
Inverse collision frequency - $(\nu_c)^{-1}$	$\sim 5 \times 10^{-13} \text{ s}$
Electron CFL @ $5 \times 10^6 \text{ m/s}$	$\sim 1 \times 10^{-13} \text{ s}$
Photon CFL	$\sim 2 \times 10^{-15} \text{ s}$
Charged particles per element	50

- Compare 2 different cases:
  - 90% Nitrogen, 10% Helium
  - 10% Nitrogen, 90% Helium



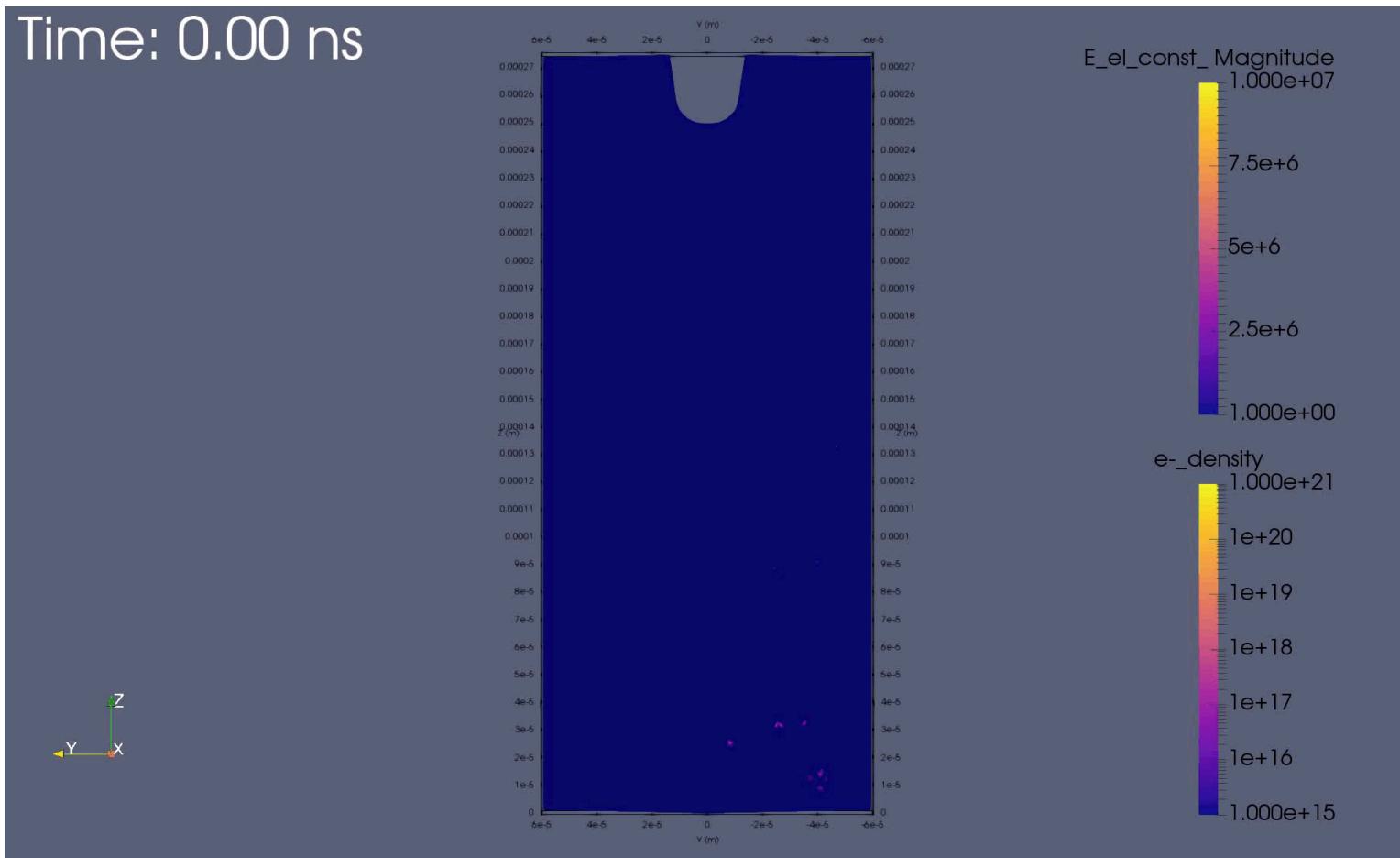
# Simulation Setup



- **~170 million elements for only a 10 degree wedge.**
- **To ease particle requirements, for  $r > 60 \mu\text{m}$ , we assume vacuum. Executed on 5120 cores on the Skybridge super computer.**
- **In summary, this is a very challenging simulation for even modern computers and numerical techniques.**

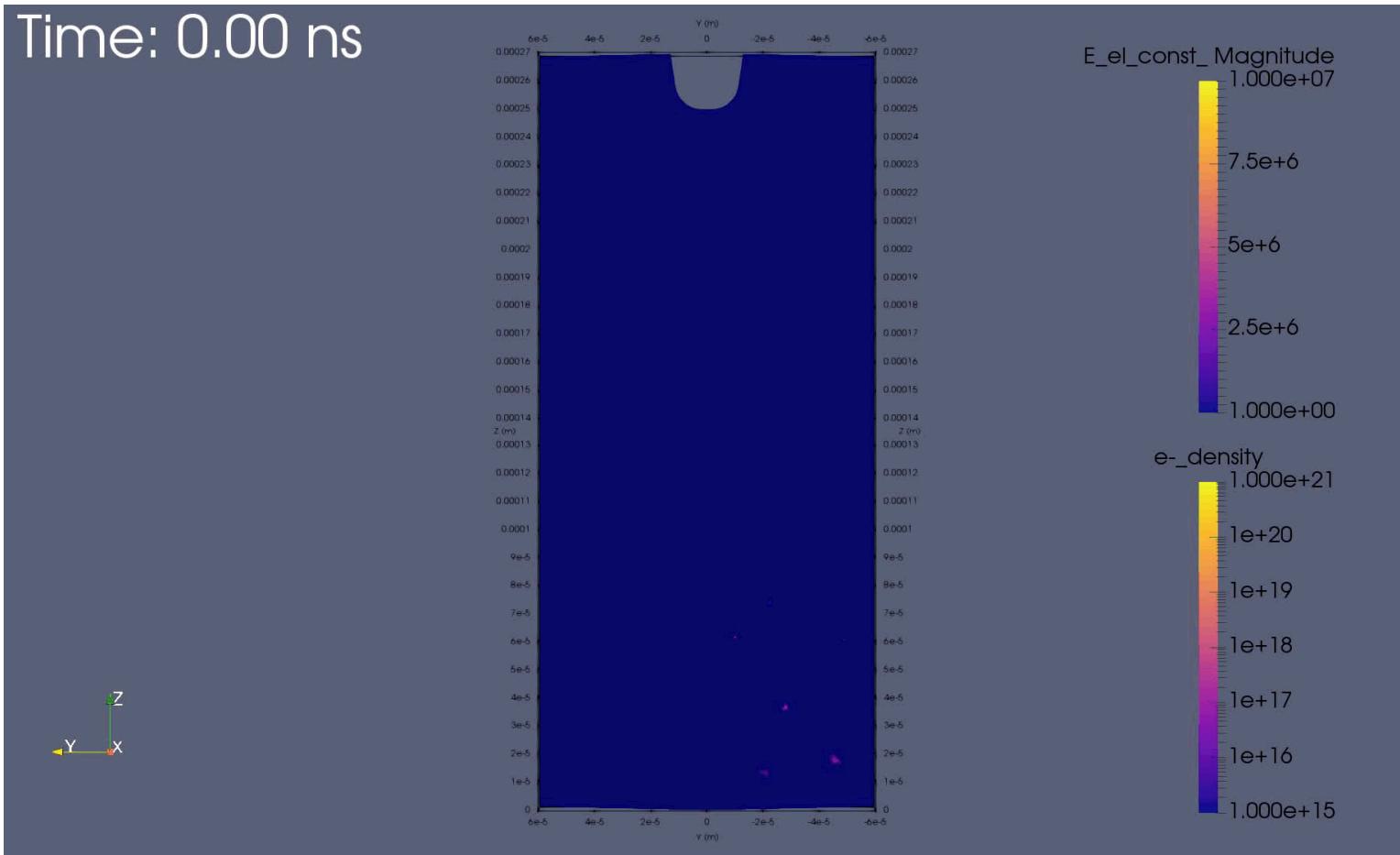
# 90% N<sub>2</sub>, 10% He

Time: 0.00 ns



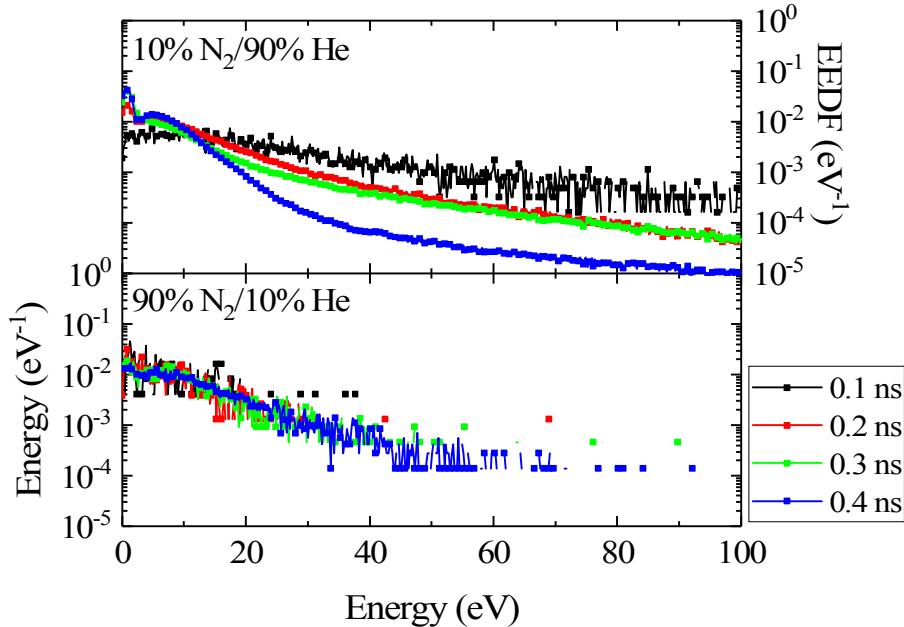
# 10% N<sub>2</sub>, 90% He

Time: 0.00 ns

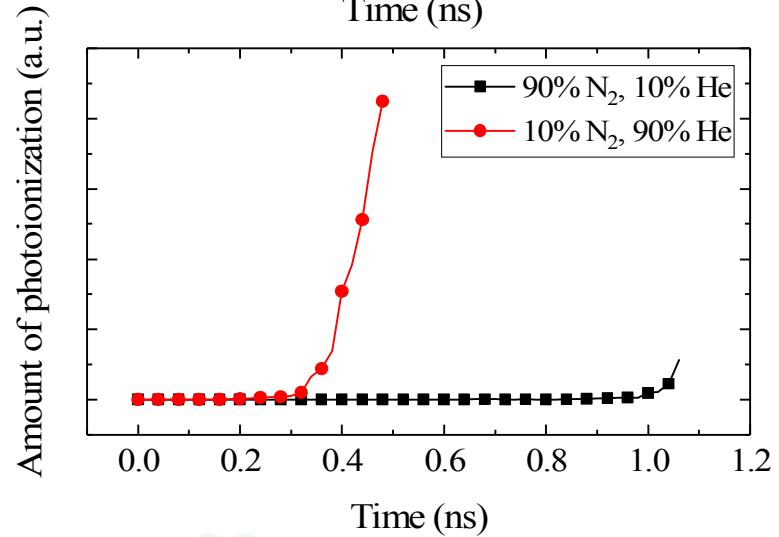
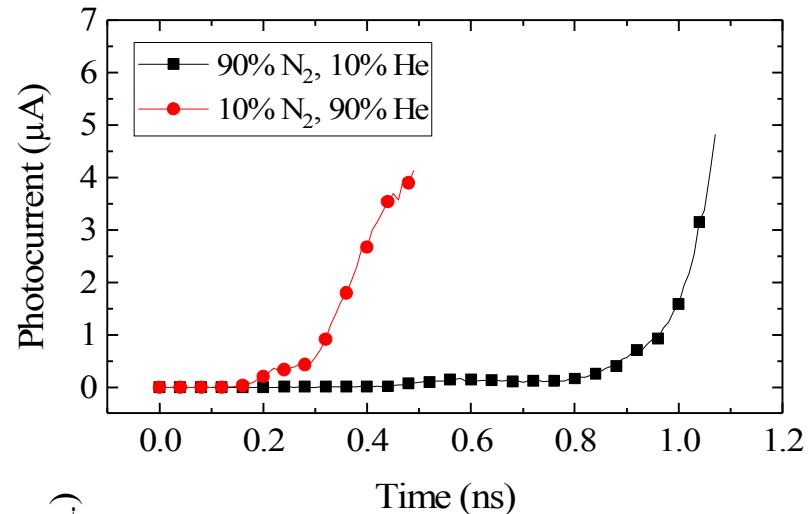


# Results

- Comparing photo-effects from each case
  - Both photo-emission and photo-ionization occur earlier for 90% He



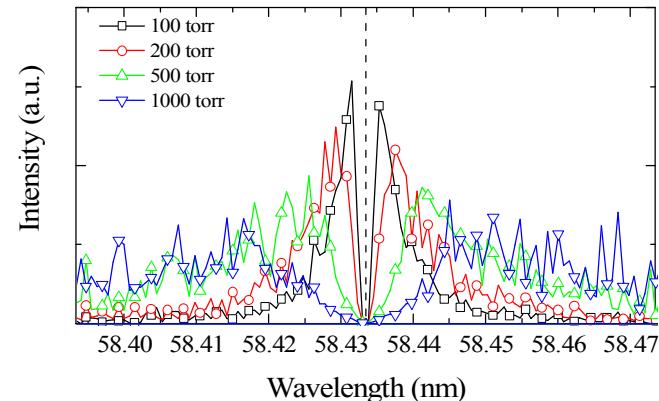
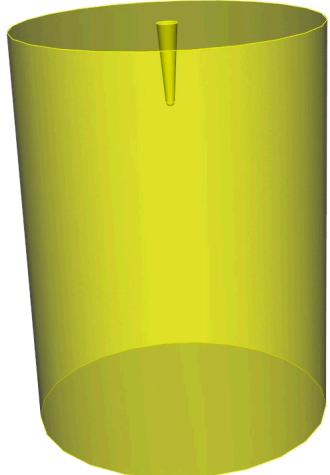
- EEDF's are shifted towards higher energies for the 90% He case due to less energy loss to vibrational or rotational energy modes.



# Conclusion and Outlook

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- Have incorporated a method to discretely track photons and include energy-dependent photo-processes.
- Can generate non-LTE emission spectra that is both spatially and temporally resolved. On-going work is comparing simulation data versus experiment.
- Gain knowledge in simulating large-scale, near atmospheric pressure plasmas.



- Even on large super computing systems, modeling larger plasma devices ( $> 1$  mm) at near atmospheric pressures with a kinetic code is still very much a challenging problem, and likely still years away with current algorithms and hardware.