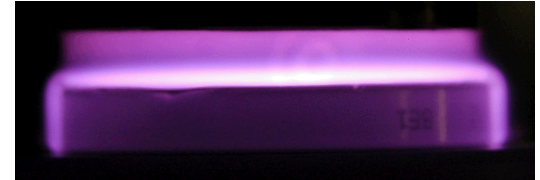


Developing a kinetic approach to radiation transport and its interaction in He/N₂ ionization waves

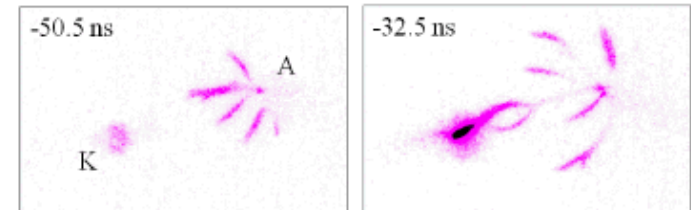
Andrew Fierro, Ed Barnat, Chris Moore, Matthew Hopkins
Sandia National Laboratories, Albuquerque NM

Motivation

- Interested in transient response times of plasma formation as well as steady quantities which may be sensitive to photonic processes.
- Light emission is one of the defining characteristics of plasma discharges. Another tool to make comparisons to experiments for validation purposes.
- Would like to begin to quantify the effect of self-produced radiation on plasma development and its secondary effects.
- Develop a method to discretely model photons in a kinetic code that allows for the incorporation of energy dependent photo-processes.



- Pulsed helium discharge operating at 30 torr (A. Fierro, E. Barnat)



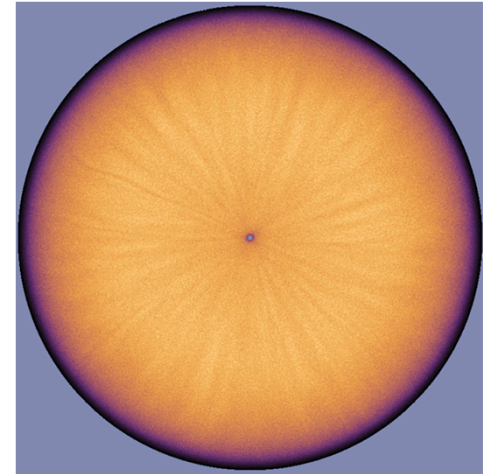
- Fast-gated imaging of streamer propagation.

A. Fierro, *et al.*, J. Phys. D: Appl. Phys., 2012.

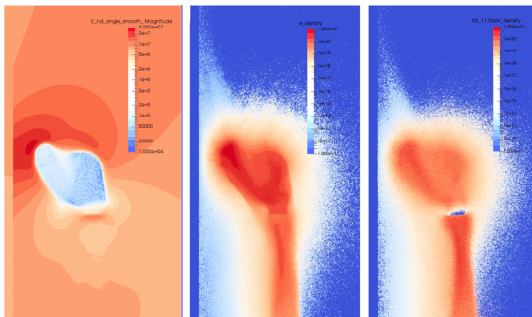
A. Fierro, *et al.*, Plasma Sources Sci. Technol., 2017, to be submitted.

The Kinetic Code - *Aleph*

- Unstructured FEM (compatible with CAD)
- Massively parallel (scales up to 100k proc.)
- Hybrid PIC + DSMC, also PIC-MCC
- Electrostatics with fixed B Field
- Advanced surface (electrode) models
- Collisions, charge exchange, chemistry, excited states, ionization
- Photon transport, photoemission, photoionization
- Advanced particle weighting methods
- Dynamic load balancing restart (with all particles)



- 2D simulation of a Langmuir probe in the electron saturation regime demonstrating unexpected streaming instabilities (B. Yee).



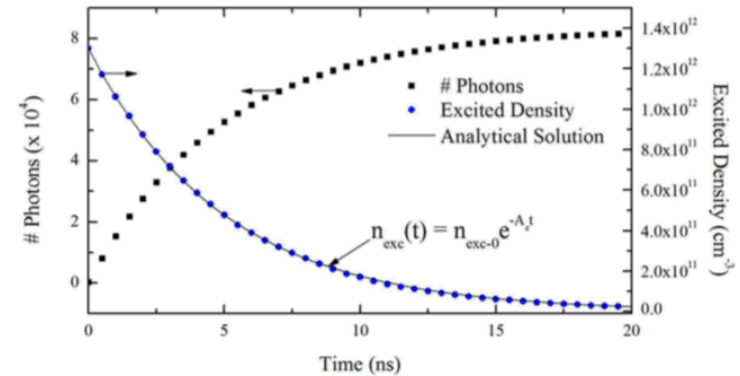
- Simulation of streamer formation along a dielectric surface (C. Moore).

Radiation Transport Method

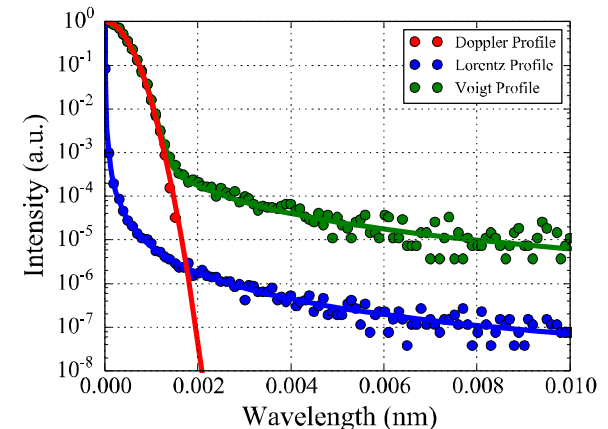
- $e^- + A \rightarrow A^* + e^-$
 - leads to an accumulation of A^*
- for each A^* evaluate:
 - $R < 1 - e^{-t/\tau}$, τ = lifetime of A^*
- If evaluated to be true
 - $A^* \rightarrow A + h\nu$
 - $|v_{ph}| = c$, v_{ph} = isotropic
 - $\lambda_{ph} = hc / (E(A^*) - E(A))$
 - Lorentz shape

$$\lambda_s = \tan[(R - 0.5)\pi] \cdot \Delta\lambda_r + \lambda_0$$
 - Doppler shape

$$\lambda_f = \frac{(c + \hat{v}_{ph} \cdot v_p)\lambda_c}{c}$$
- Each photon is pushed through the computational domain and interactions are handled with traditional DSMC procedures.



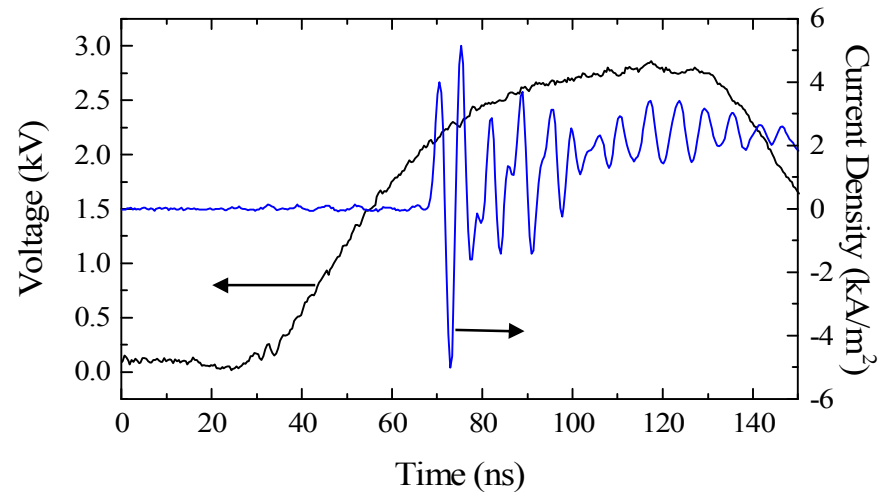
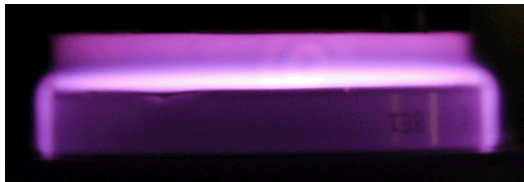
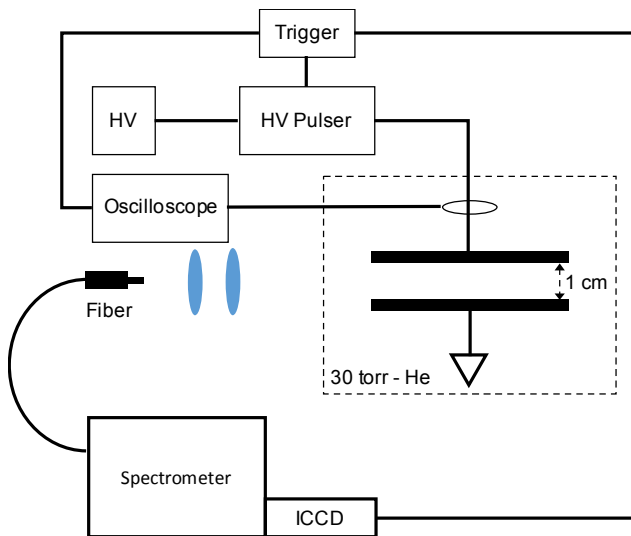
- Verification of the method.



- Simulated line profiles.

Towards Validation

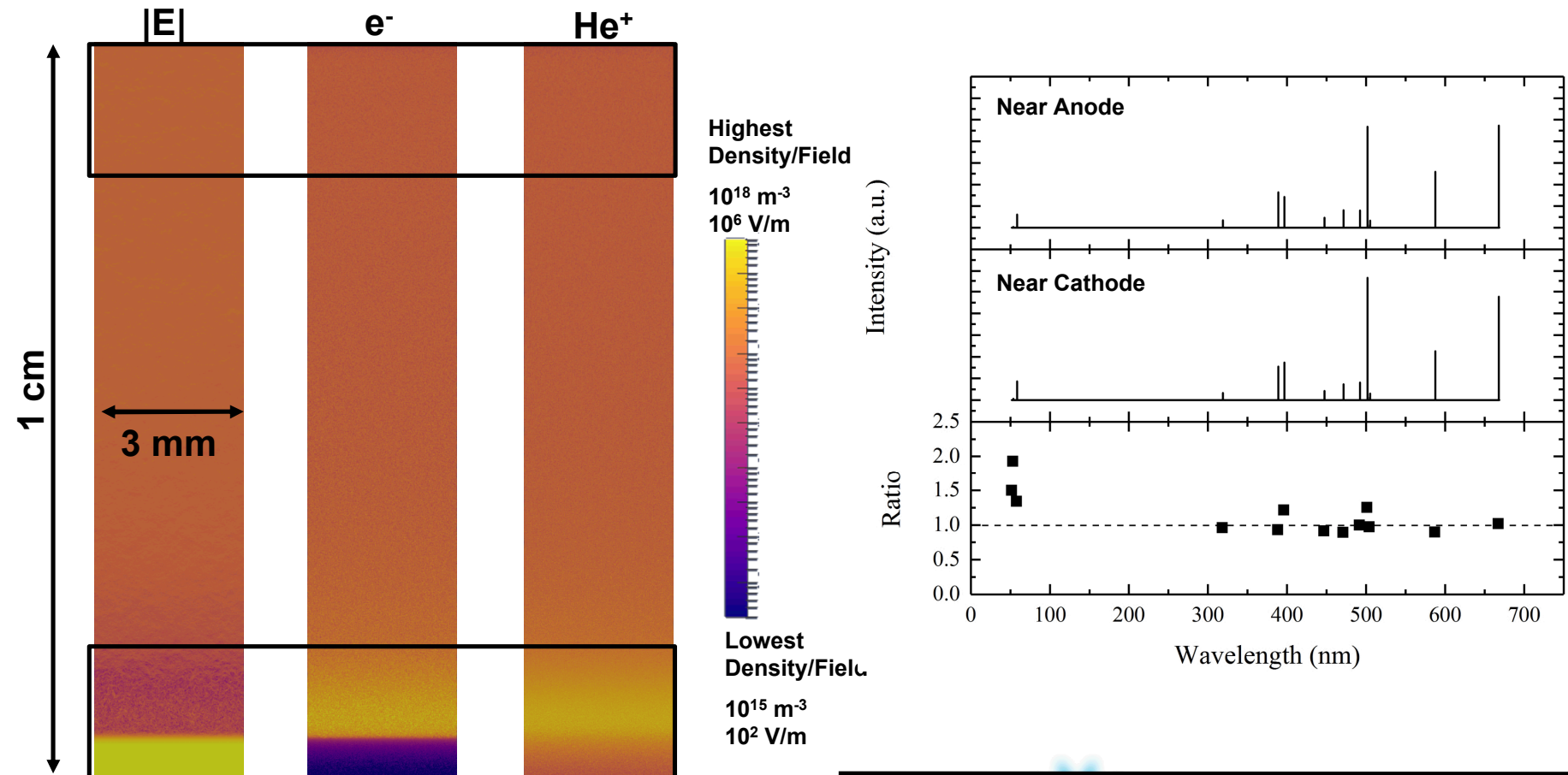
- With this method, we are able to generate spatially-resolved line emission spectra that does not assume LTE.
 - Can we begin to use emission spectra as a viable validation tool?



- Experimental setup focused on capturing time-resolved optical emission spectroscopy in the visible regime.

Towards Validation

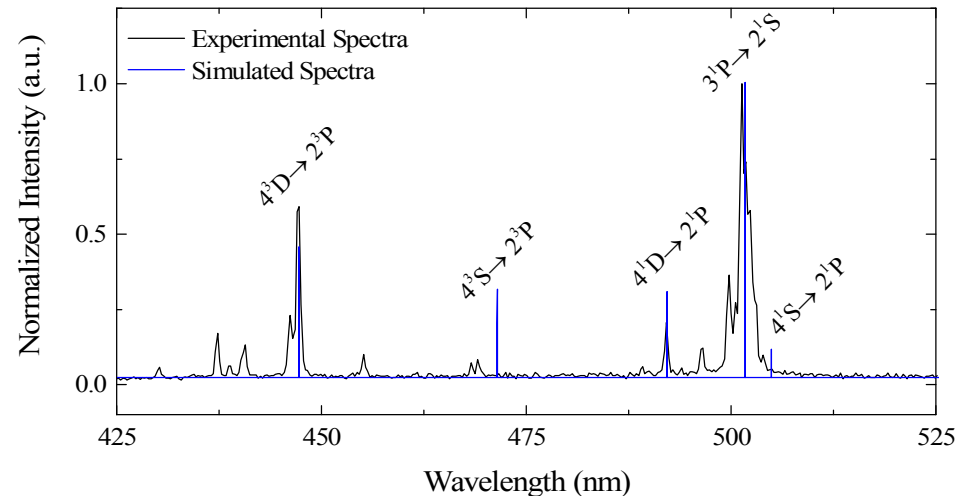
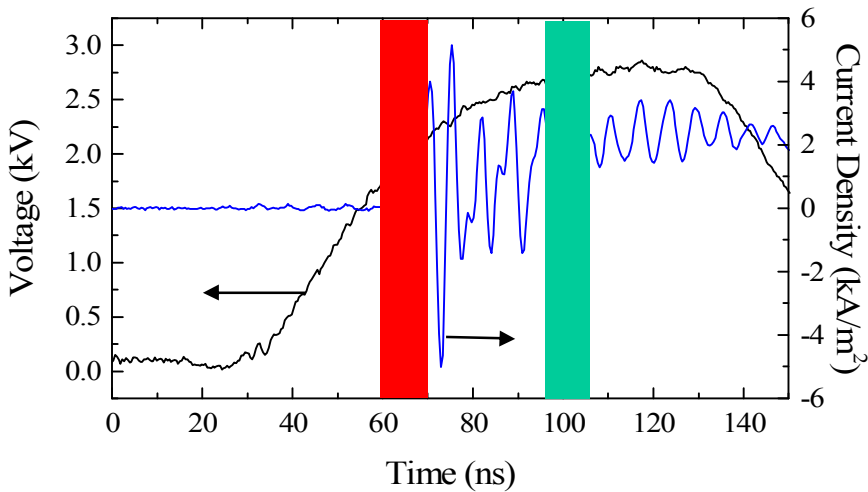
- With this method, we are able to generate spatially-resolved line emission spectra that does not assume LTE.



- Log contour plots from a simulation of pulsed plate-plate discharge, $t = 60\text{ ns}$

Towards Validation

- Initial comparisons to experimental data
 - 10 ns camera gate, timing indicated by green window.
 - Simulation spectra taken at $t = 60$ ns
 - Working on time correlation between experiment and simulation



- The simulation produces realistic line ratios for the He lines shown.
- Assume pure He gas, no impurities are included.

Application to Large Scale Simulations

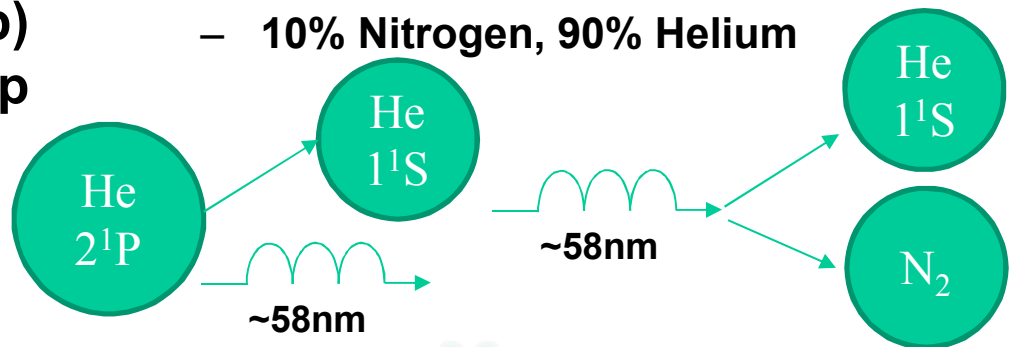
- **How do photonic mechanisms impact near atmospheric pressure discharges? The goal is to simulate a three-dimensional system of a pulsed, near atmospheric pressure discharge.**
 - How big of a system is capable of being simulated in *reasonable* time.
 - What assumptions can be made to alleviate computational requirements?
 - What are the numerical challenges (timestep, space step)?
 - What are the computational resource challenges (memory, processors)?
- **Simulations at higher pressures are extremely challenging due to the anticipated large electron densities. This results in a smaller mesh to resolve Debye lengths and avoid numerical heating.**

Ionization Wave Simulation

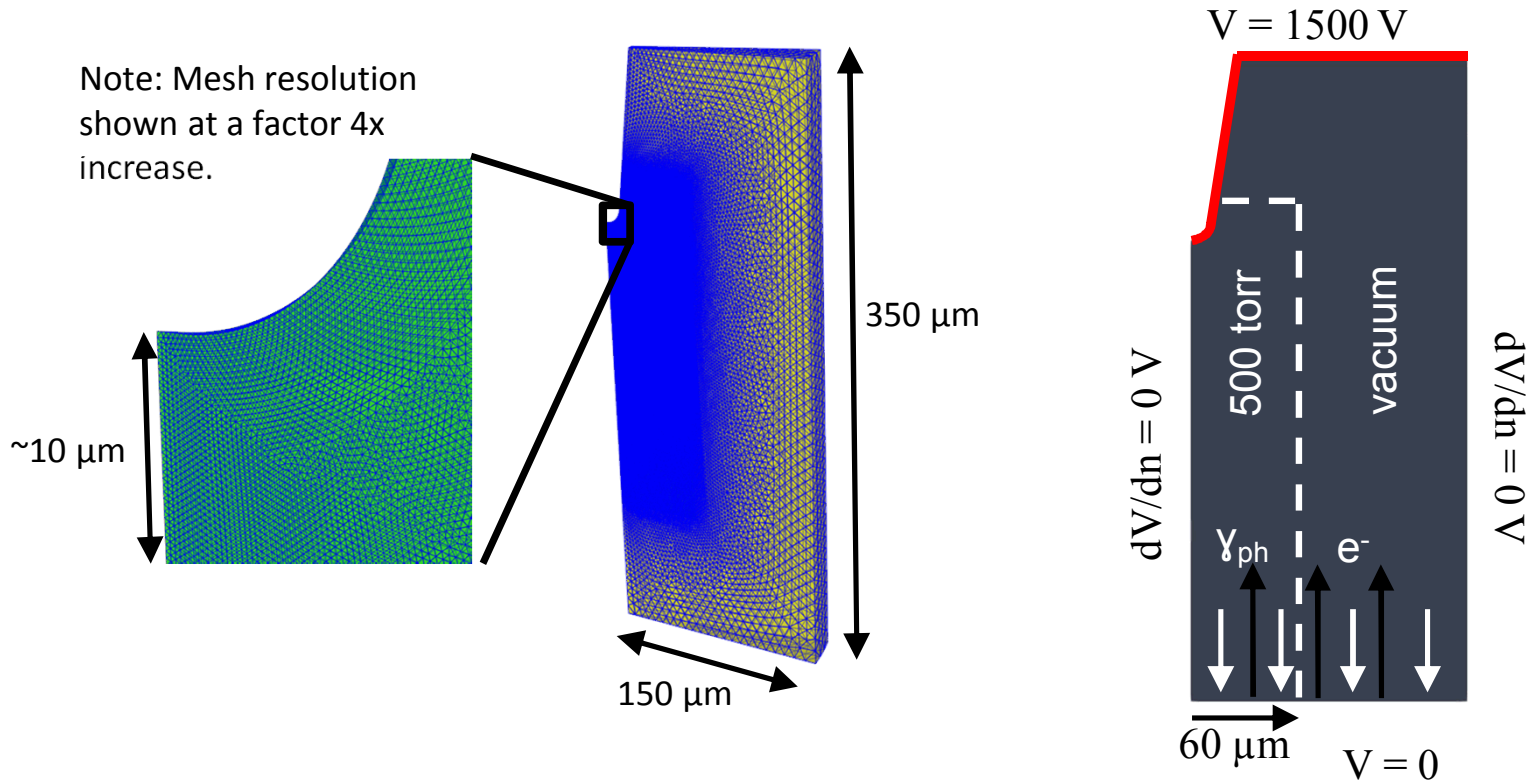
- Numerical parameters are guided by experiments of similar type discharges
 - 500 torr background pressure
 - 10^{15} cm^{-3} electron density
 - 0.5 eV electron temperature
- Debye length and Photon CFL conditions (photon does not cross more than 1 element in a time step) set the minimum time step and spatial step.
 - $dx = 50 \text{ nm}$
 - $dt = 2 \times 10^{-15} \text{ s}$

| Parameter | Spatial or Time Scale |
|---|------------------------------------|
| Debye Length - λ_D | ~150 nm |
| Electron Mean Free Path - λ_{mfp} | ~200 nm |
| Photon Mean Free Path - γ_{mfp} | ~25 μm |
| Inverse plasma frequency - $(\omega_{pe})^{-1}$ | $\sim 1 \times 10^{-13} \text{ s}$ |
| Inverse collision frequency - $(\nu_c)^{-1}$ | $\sim 5 \times 10^{-13} \text{ s}$ |
| Electron CFL @ $5 \times 10^6 \text{ m/s}$ | $\sim 1 \times 10^{-13} \text{ s}$ |
| Photon CFL | ~2 x 10 ⁻¹⁵ s |
| Charged particles per element | 50 |

- Compare 2 different cases:
 - 90% Nitrogen, 10% Helium
 - 10% Nitrogen, 90% Helium



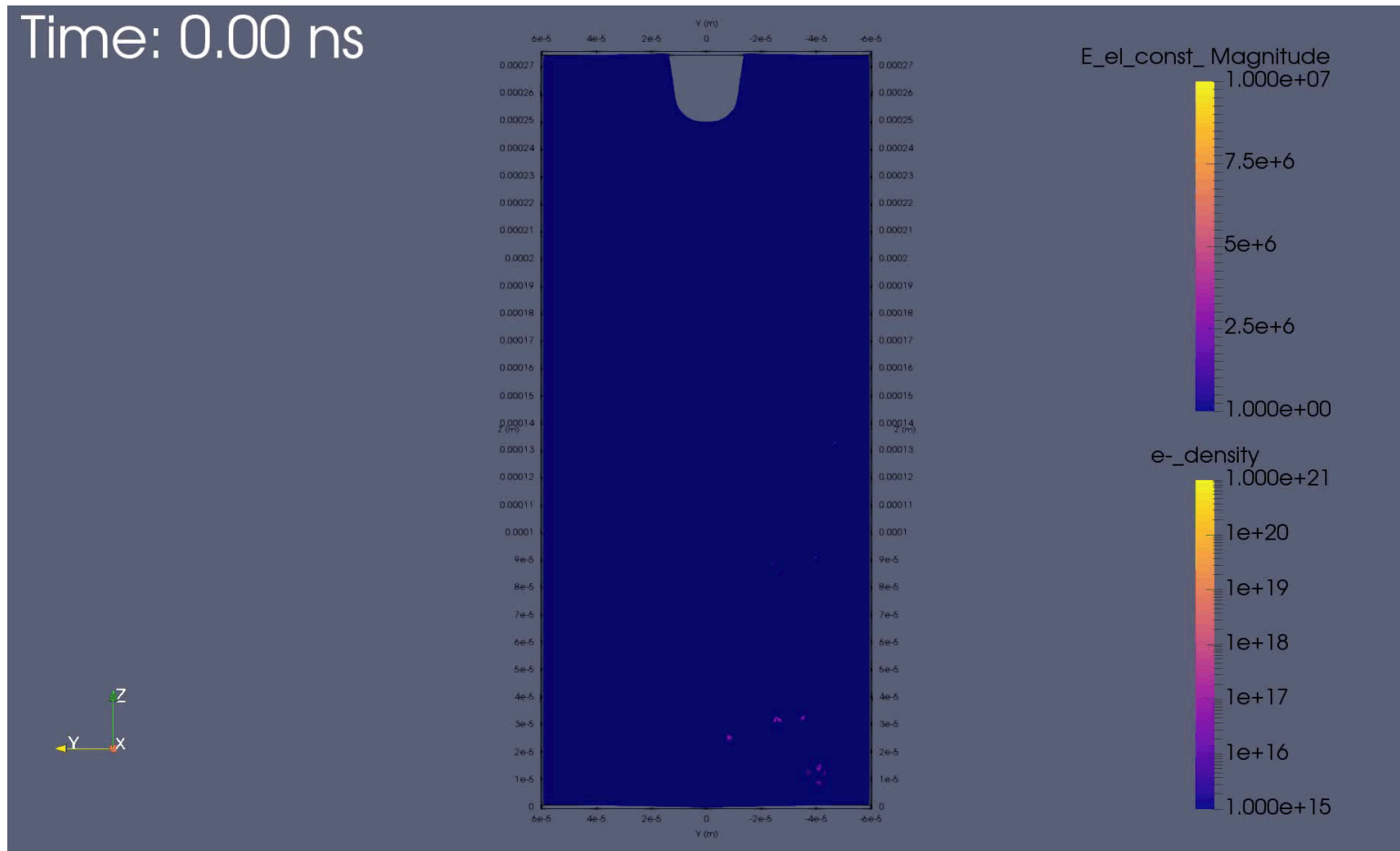
Simulation Setup



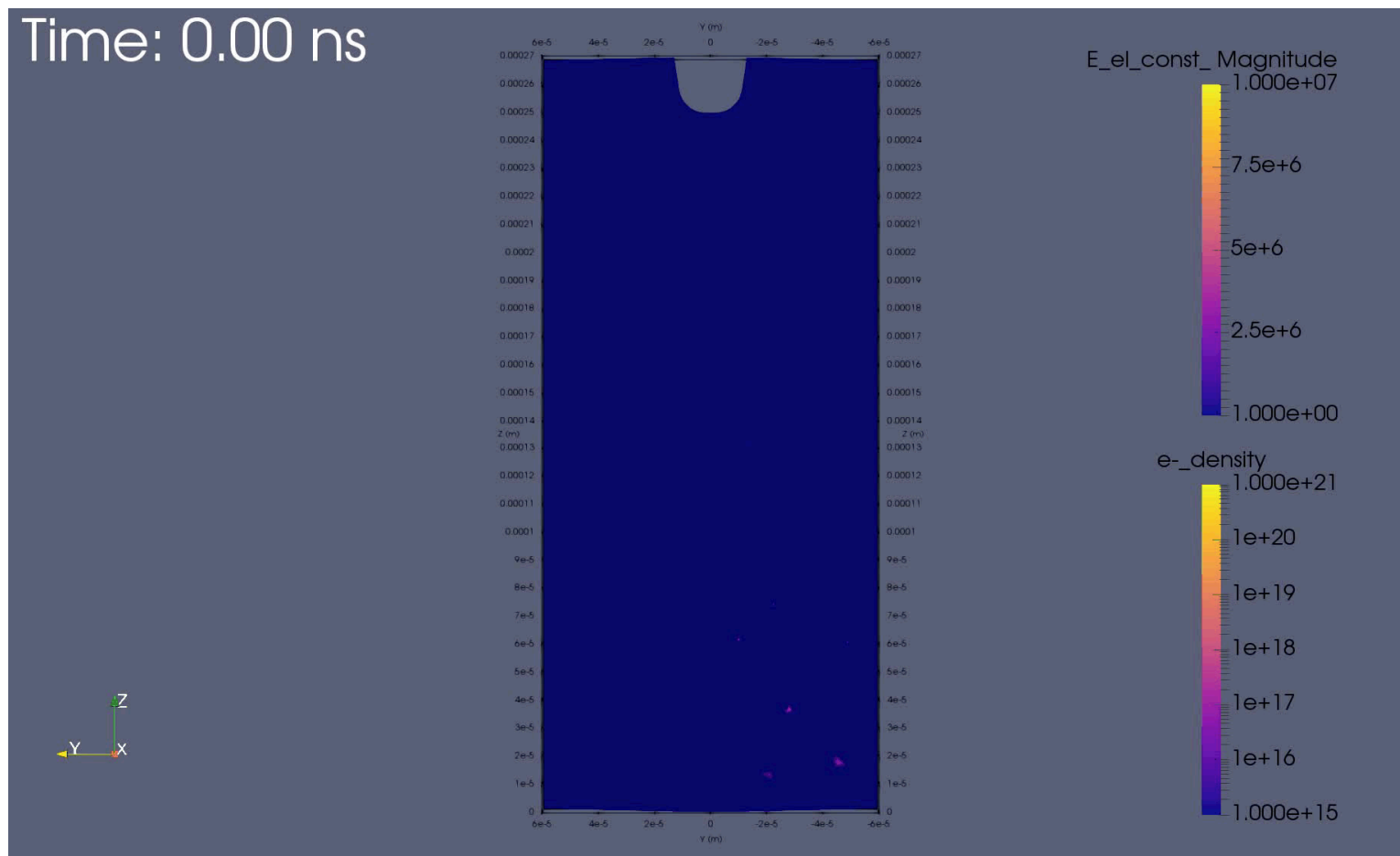
- ~170 million elements for only a 10 degree wedge.
- To ease particle requirements, for $r > 60 \mu\text{m}$, we assume vacuum. Executed on 5120 cores on the Skybridge super computer.
- In summary, this is a very challenging simulation for even modern computers and numerical techniques.

90% N₂, 10% He

Time: 0.00 ns



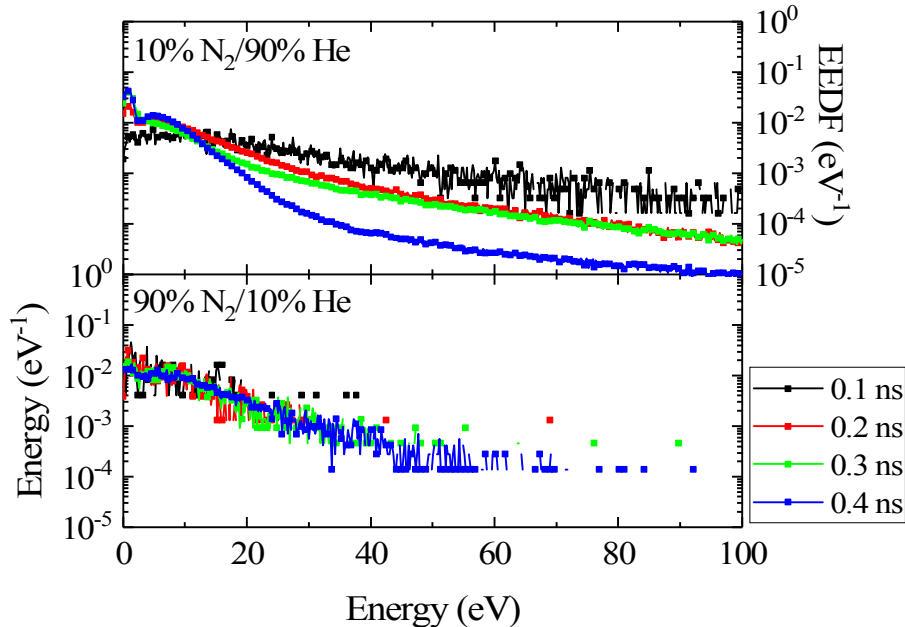
10% N₂, 90% He



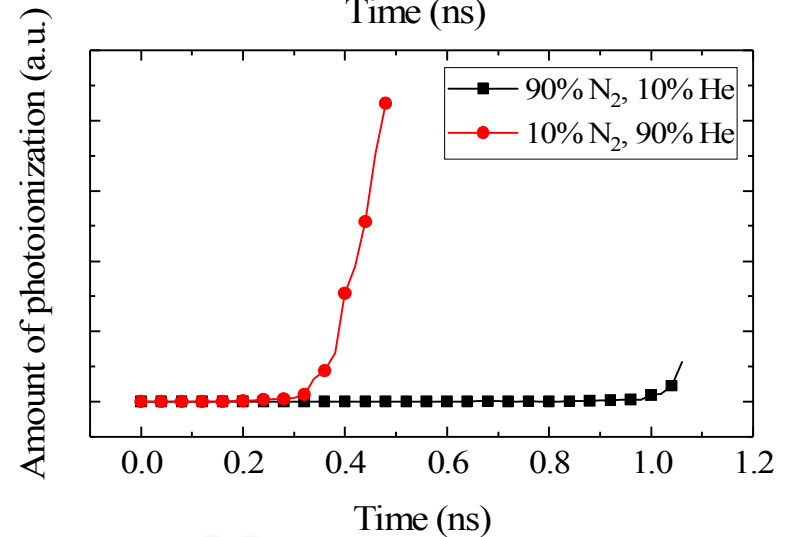
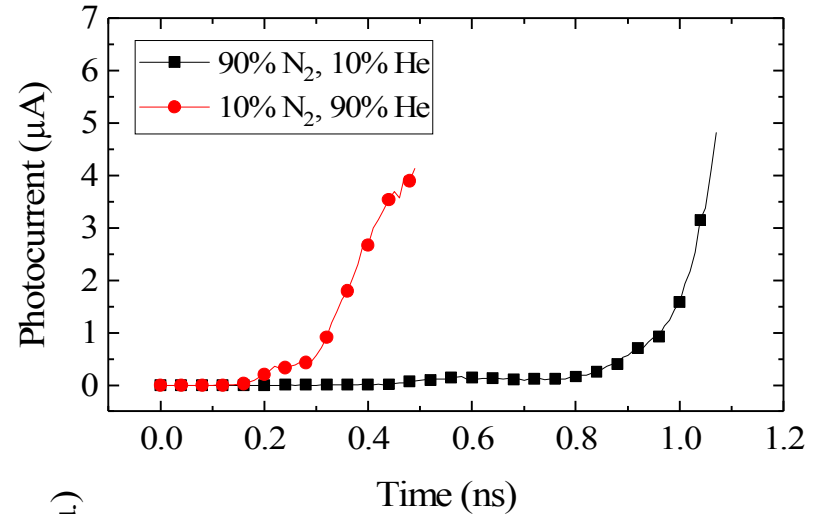
Results

- Comparing photo-effects from each case

- Both photo-emission and photo-ionization occur earlier for 90% He

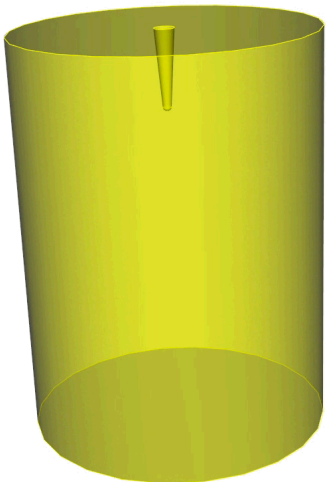
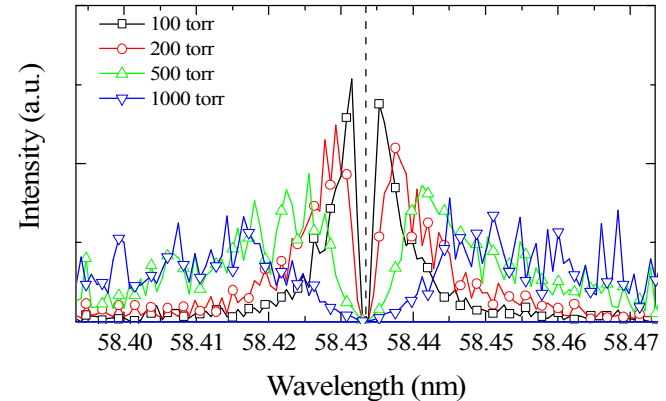


- EEDF's are shifted towards higher energies for the 90% He case due to less energy loss to vibrational or rotational energy modes.



Conclusion and Outlook

- Have incorporated a method to discretely track photons and include energy-dependent photo-processes.
- Can generate non-LTE emission spectra that is both spatially and temporally resolved. On-going work is comparing simulation data versus experiment.
- Gain knowledge in simulating large-scale, near atmospheric pressure plasmas.



- Even on large super computing systems, modeling larger plasma devices (> 1 mm) at near atmospheric pressures with a kinetic code is still very much a challenging problem, and likely still years away with current algorithms and hardware.