

7/21/16



# SCALING OF $K\alpha$ LINE EMISSION IN Z PINCHES FROM 2 – 60 keV

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*ICOPS, Banff June 2016*



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This work is partially supported by Sandia's Laboratory Directed Research and Development projects 173104 and 165733.

# Overview: Non-thermal $K\alpha$ is a promising route to higher photon energy sources on Z

- Optimization of non-thermal  $K\alpha$  emission is more complicated than creating breaks in a plasma column
- We have studied dynamics responsible for  $K\alpha$  emission
  - New diagnostics have aided us in building this understanding
  - We've started to build a detailed picture of  $K\alpha$  emission
- We have investigated the scaling of  $K\alpha$  emission to 59 keV
  - New diagnostics were needed to allow this assessment
  - Scaling looks similar to inferred at lower photon energies and would not be feasible with thermal emission (e.g.  $He\alpha$ )



# Underlying physics is not as trivial as might be expected

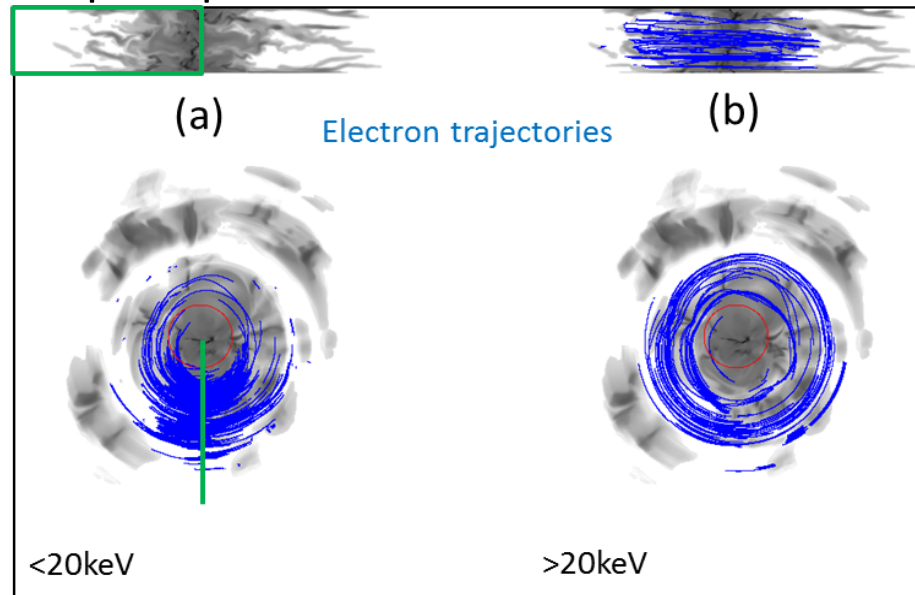
- Large fields are required for electron runaway

- Dreicer field

$$E > E_D = 5.8 \times 10^{-18} \ln \Lambda \frac{n_e}{T_e}$$

- Electrons are highly magnetized,

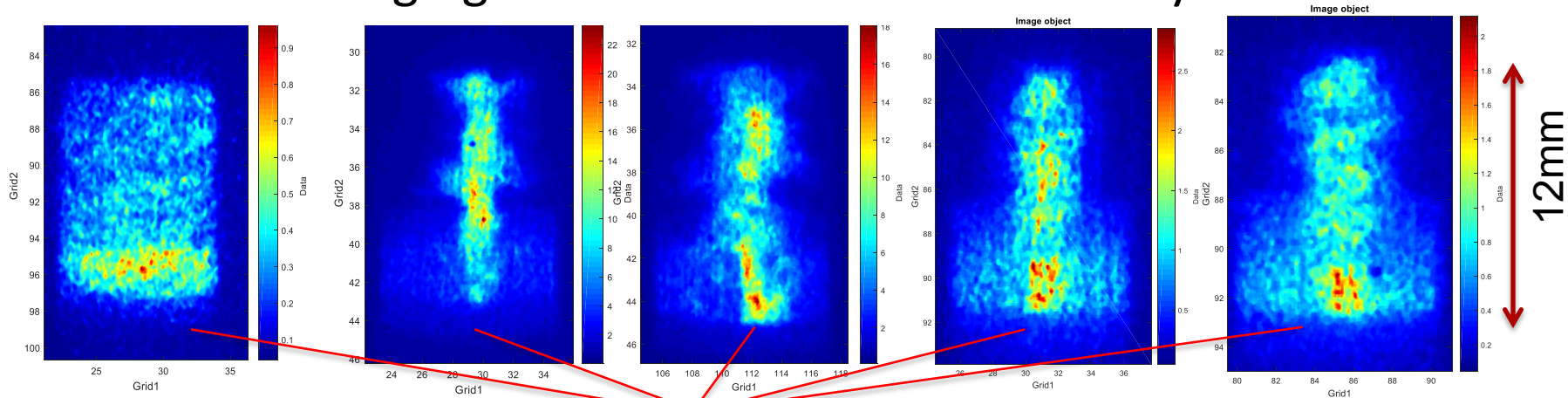
- Aren't accelerated axially across gaps or radially into dense plasma
  - Trajectories post-processed from MHD simulations



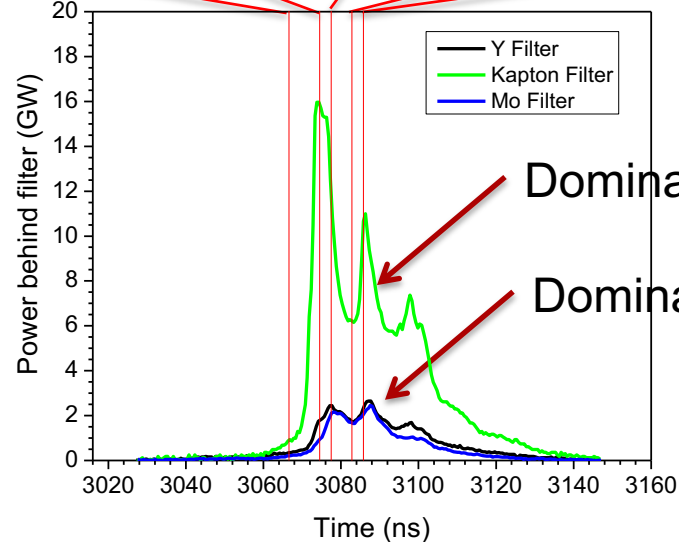
Simulation by  
J.P Chittenden *et al*

# Using 17 keV Mo $K\alpha$ emission we have studied how $K\alpha$ emission connects to array dynamics

- Self emission imaging at 277 eV shows the overall dynamics



- Comparing with PCDs we see how  $K\alpha$  fits in evolution

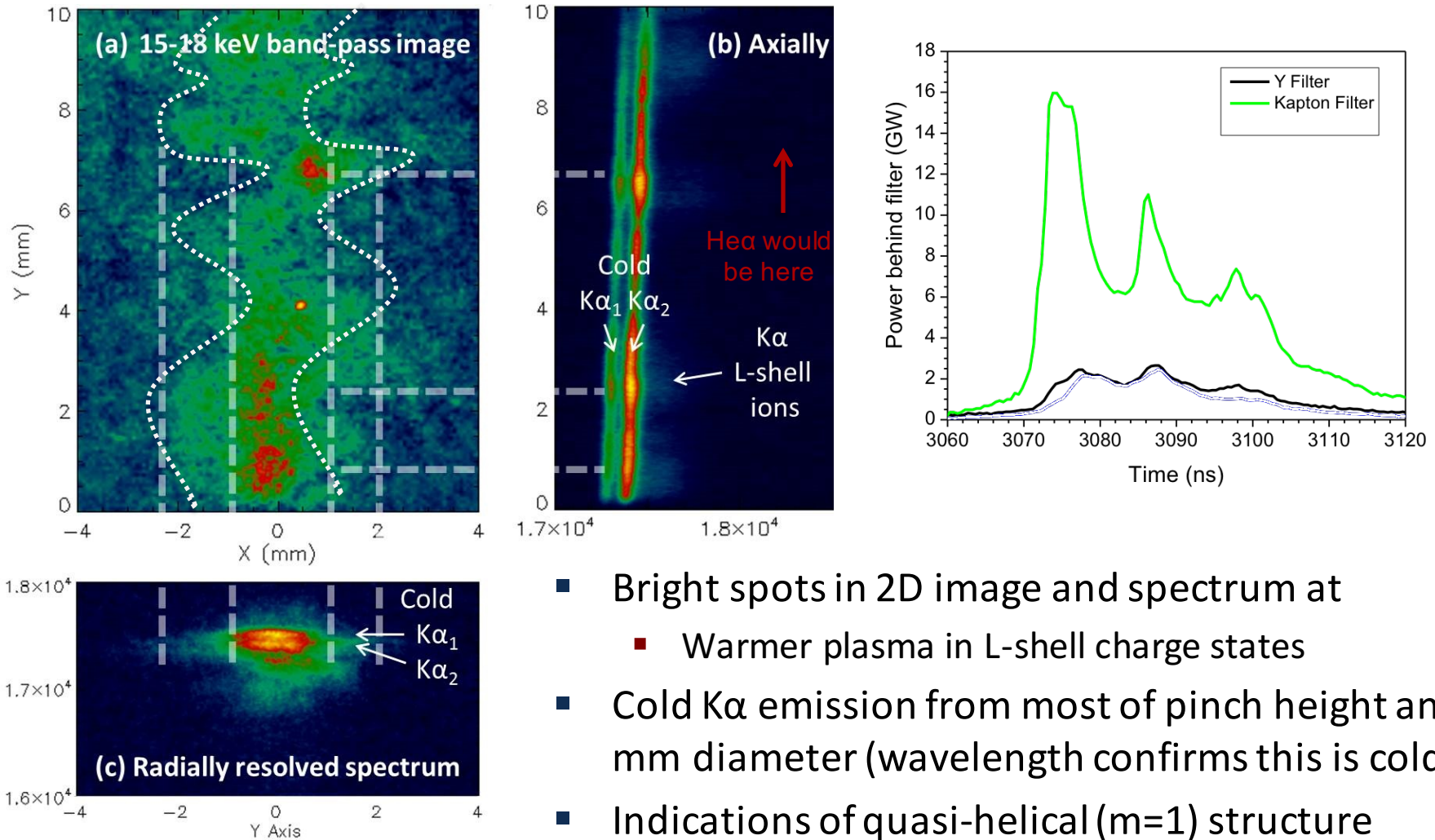


Dominated by 5-10 keV

Dominated by 10-20 keV

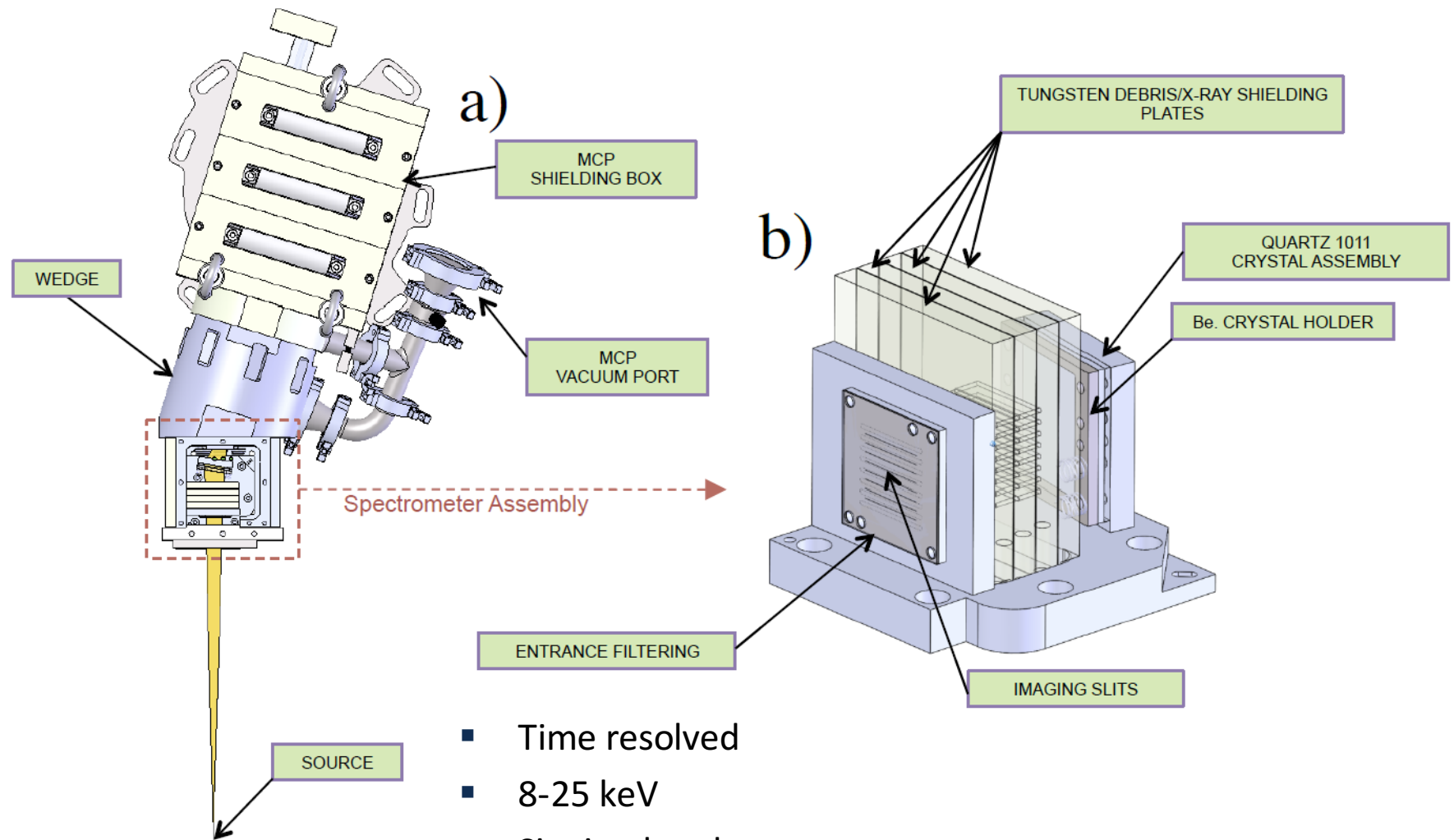


# Spatially resolved spectra combined with high energy pinhole imaging provide insight into structure



- Bright spots in 2D image and spectrum at
  - Warmer plasma in L-shell charge states
- Cold K $\alpha$  emission from most of pinch height and few mm diameter (wavelength confirms this is cold K $\alpha$ )
- Indications of quasi-helical ( $m=1$ ) structure
  - Consistent with late time emission of K $\alpha$

# New time-resolved transmission spectrometer allows us to assess evolution of $K\alpha$

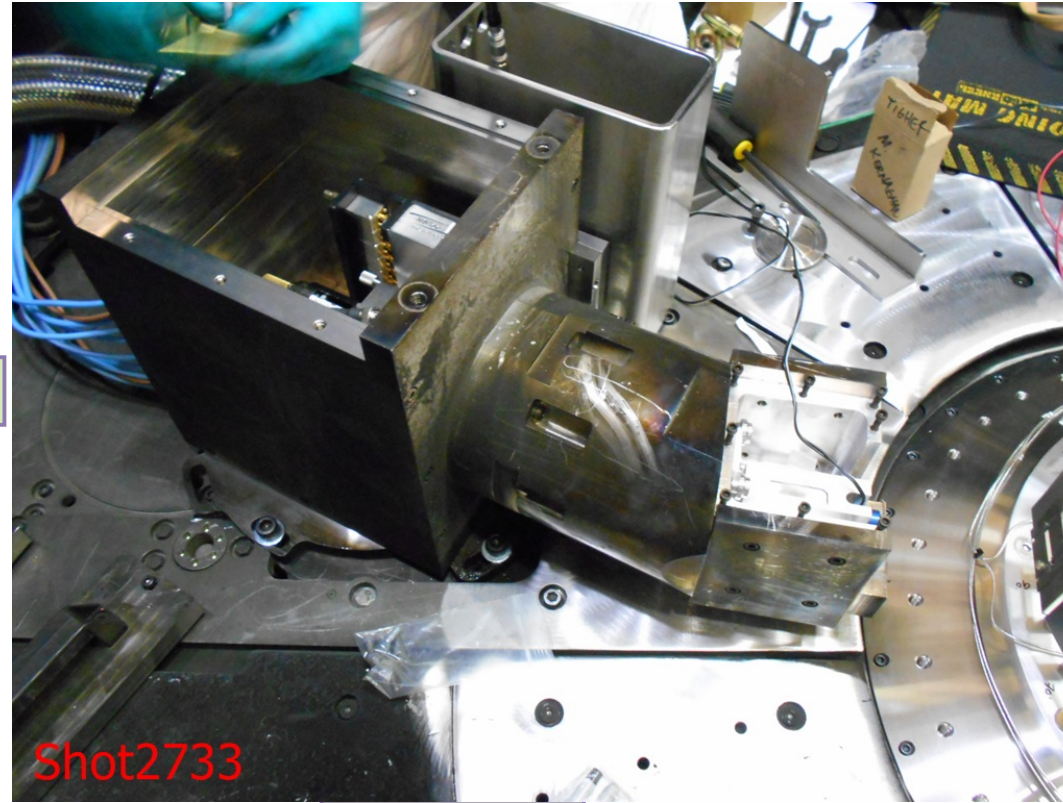
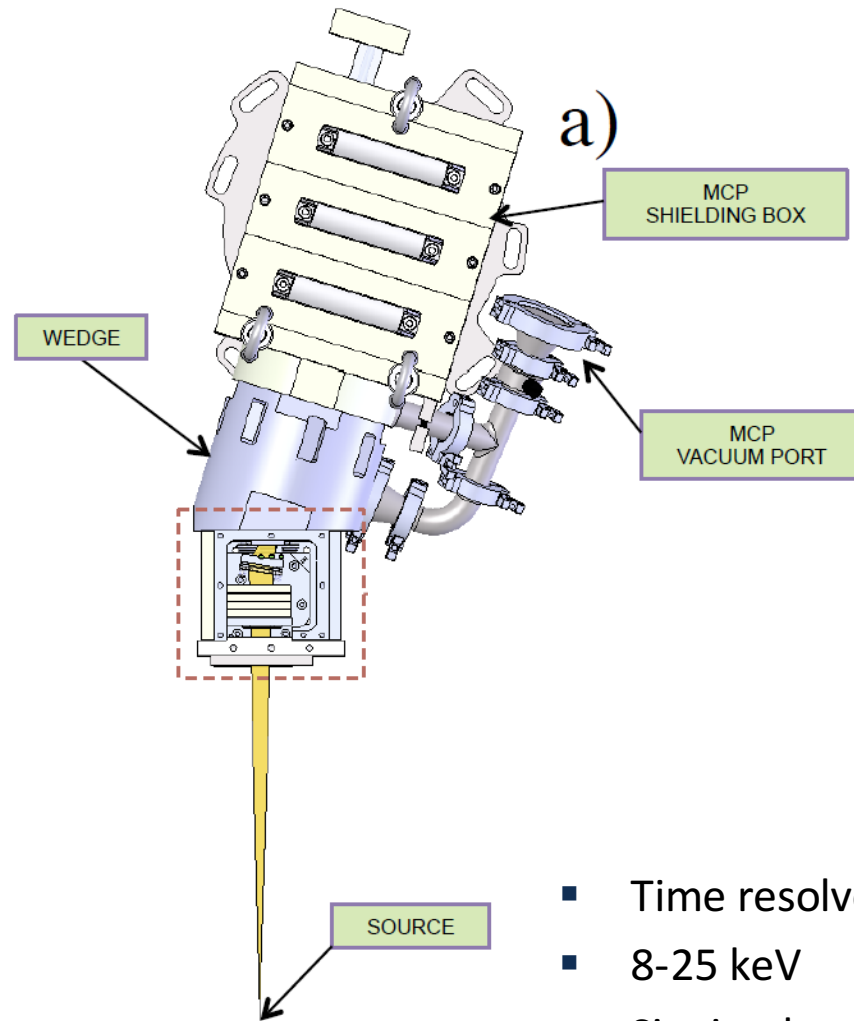


- Time resolved
- 8-25 keV
- Sits in-chamber
- $\lambda/\Delta\lambda \sim 500$

P.F. Knapp et al., In preparation for RSI



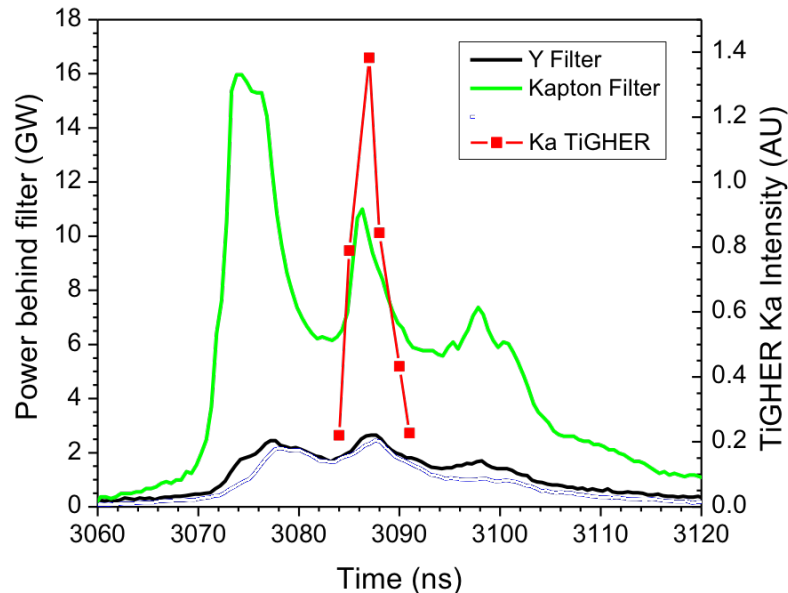
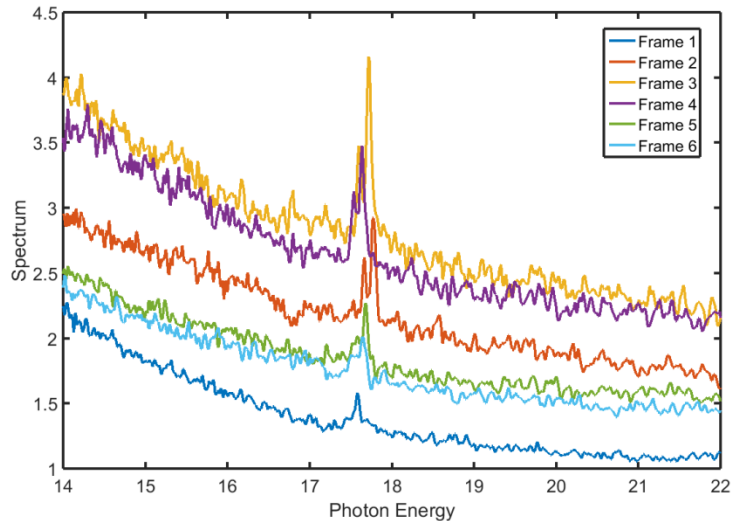
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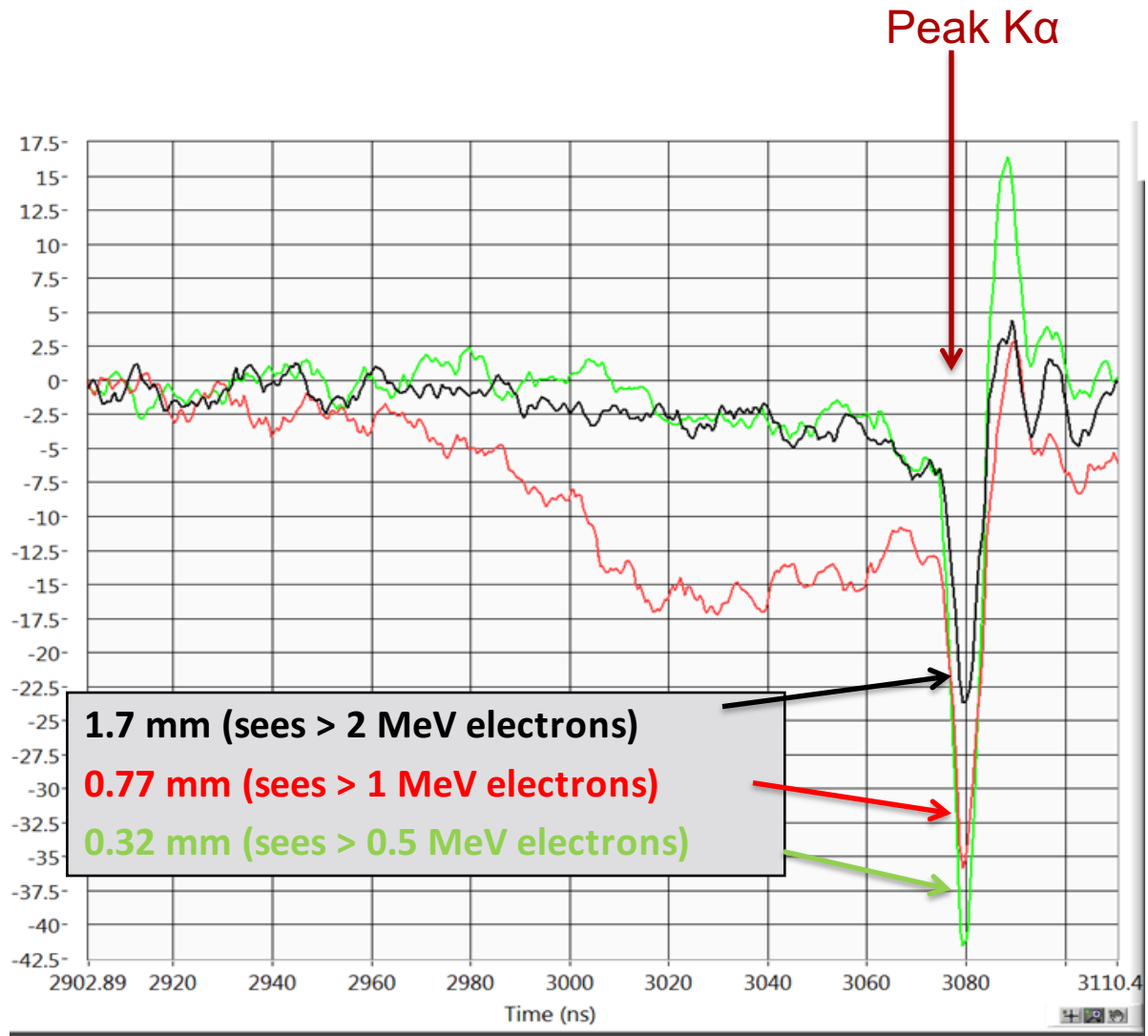
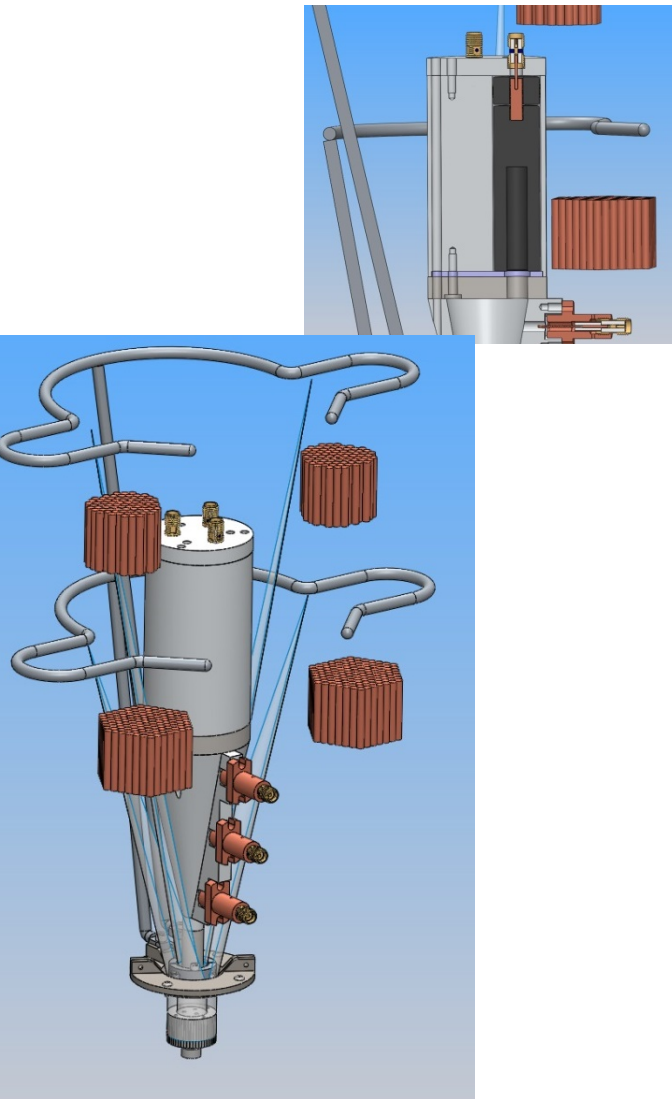
# Time resolved spectra show fast rise and fall of $K\alpha$



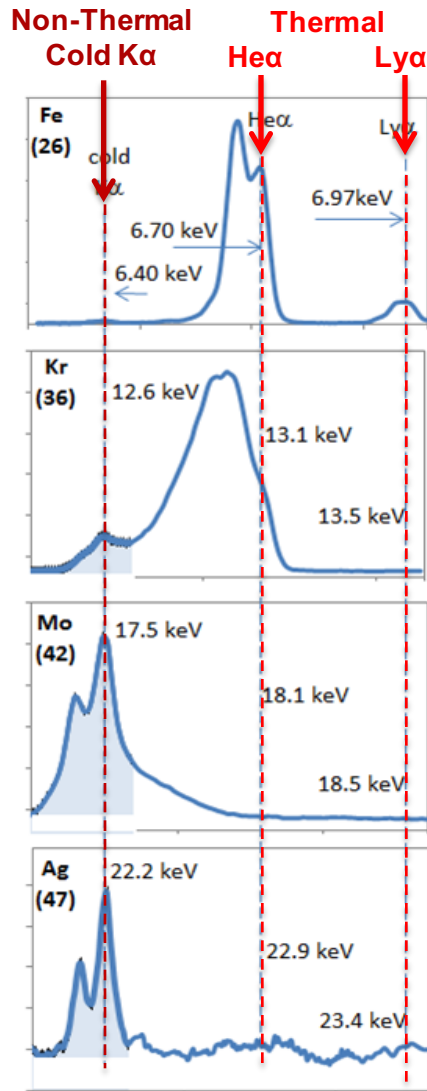
- Mo  $K\alpha$  seen on time resolved spectrometer
- Rises and falls rapidly
  - $\sim 3$  ns FWHM
  - Faster varying than continuum
- May be other short bursts
  - TiGHER captured 10 ns window
- $K\alpha$  pulse resembles shape of one pulse on the PCD trace
  - More work needed on checking cross timing of  $K\alpha$  pulse



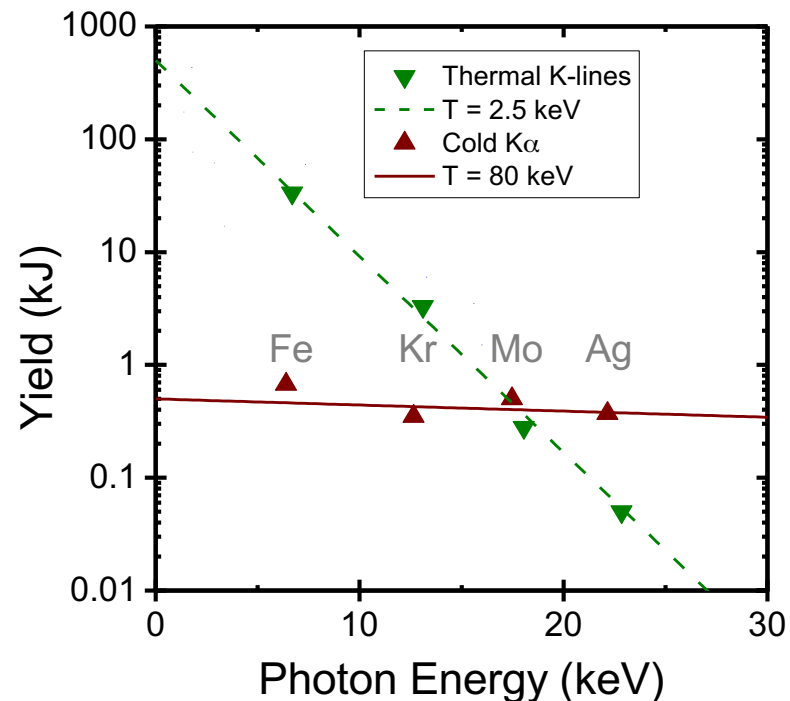
# We have developed a Faraday Cup looking at energetic electrons that are present in Z pinches on Z



# Spectra from Z experiments using different materials shows transition from thermal to non-thermal emission



- As atomic number increases see transition
  - from thermal K-shell (Fe almost entirely He $\alpha$ )
  - to non-thermal K-shell (Ag entirely cold K $\alpha$ )
- Scaling to higher energies is much better with cold K $\alpha$

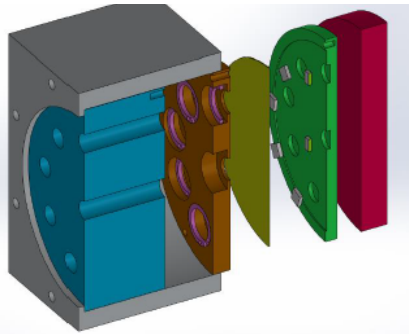


# To assess higher photon energy $K\alpha$ & Bremsstrahlung we have developed a number of diagnostics

## Filtered TLD detectors

Absolute fluence measurement 10keV-20MeV

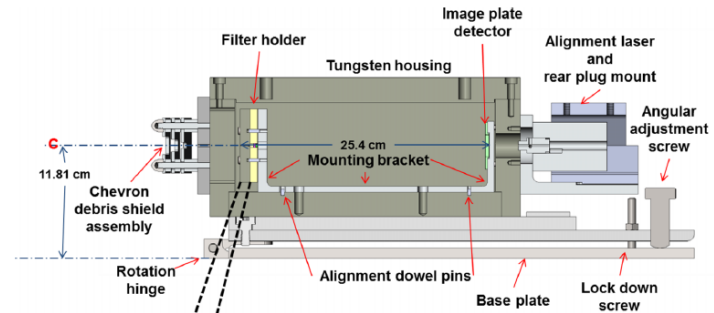
V. Harper-Slaboszewicz *et al.* *In prep.*



## Time Integrated PHC

Spatial structure 15 keV-100 keV

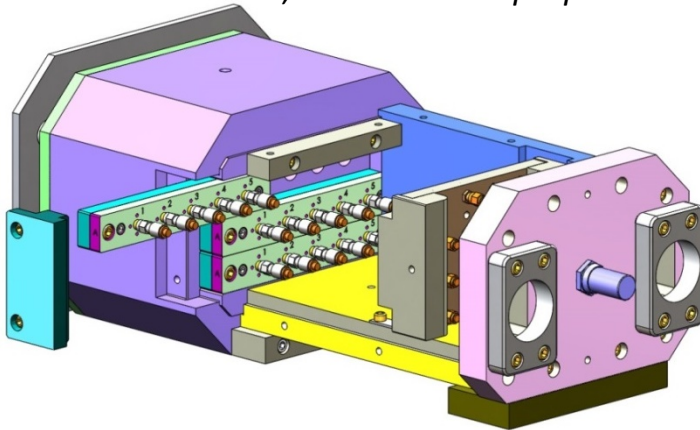
L.A. McPherson *et al.*, RSI **87**, 063502 (2016)



## Differentially Filtered Diodes

Absolute fluence measurement 10keV-20MeV

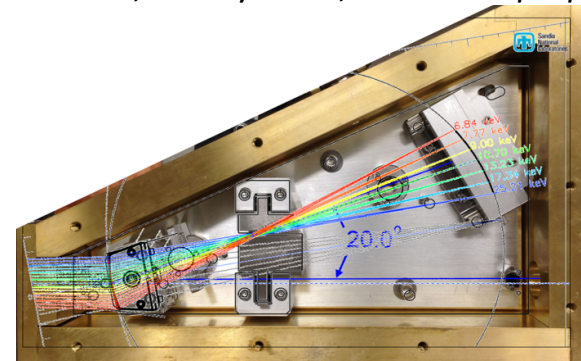
K.S. Bell, Talk 5F-7 & *In prep.*



## Space resolved spectra

Spectra 7 keV-100 keV

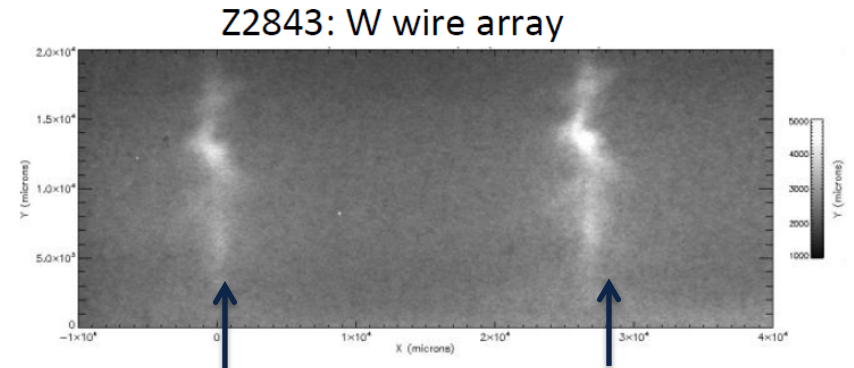
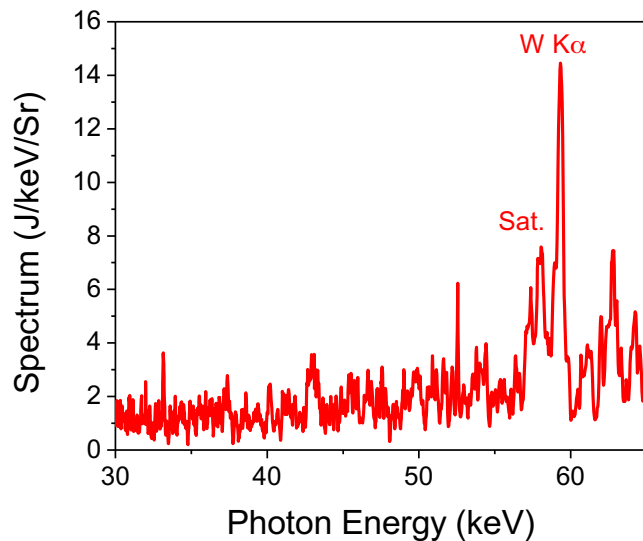
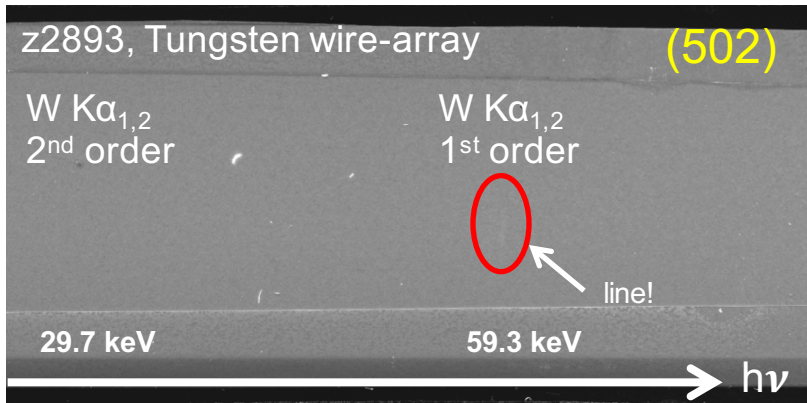
G. Loisel, J Seely *et al.*, HTPD & *In prep.*



# Recent data has shown first evidence of W $K\alpha$ emission on Z

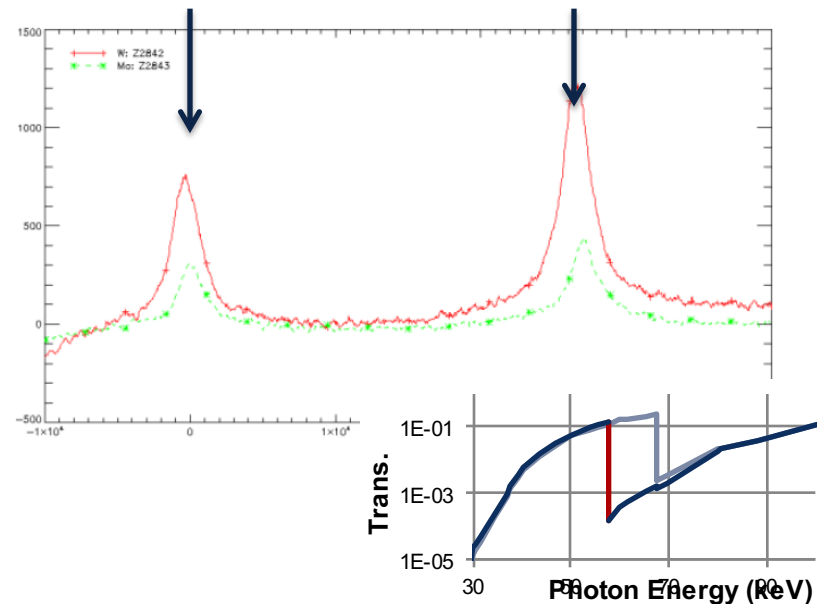
Spectra show evidence of line

Ross paired for  $\sim 60$  keV gives image



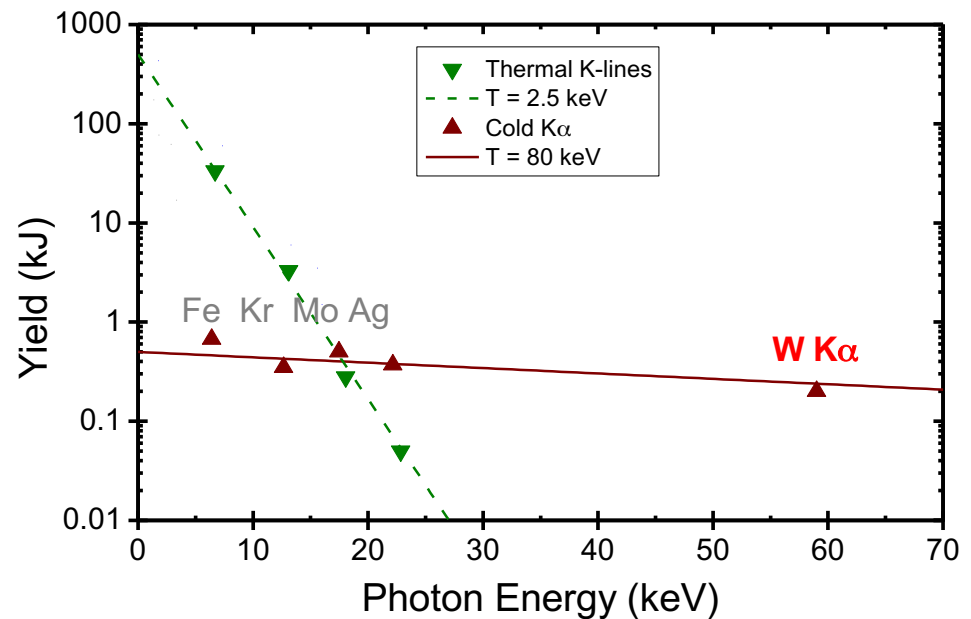
Insensitive to 59 keV

Sensitive to 59 keV



# Scaling to W $K\alpha$ is consistent with lower energy data

- We have previously assumed an energetic electron population at energies  $\sim 80$  keV to provide scaling through  $K\alpha$  data
  - Not tightly constrained, especially with small range in  $h\nu$  and small dependence on Te
- W data point fits well and provides first data point at  $>30$  keV for wire arrays



Calibration of Quartz 502 CRITR-X setup by Loisel, Seely *et al.*  
W yield unfold by G. Loisel

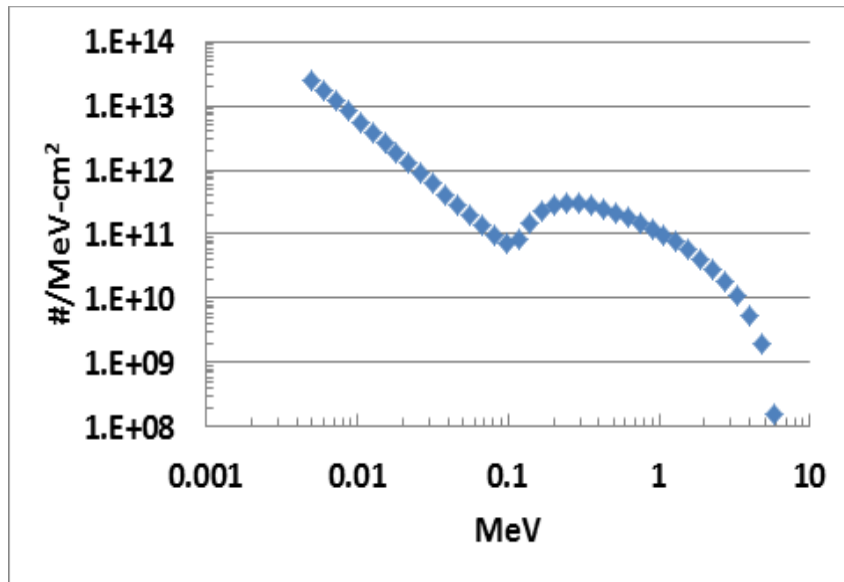


# We are also taking steps to better measure Bremsstrahlung from wire arrays on Z

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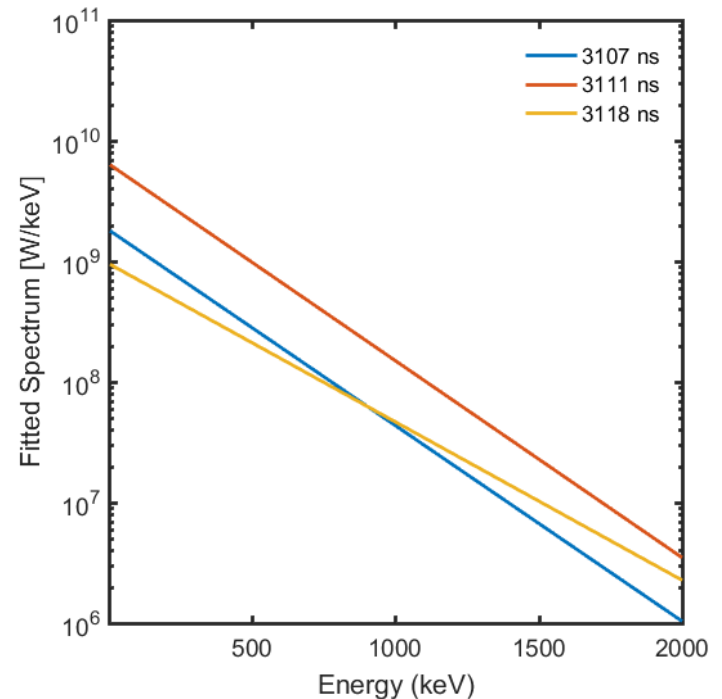
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