

# SCALING OF $\text{K}\alpha$ LINE EMISSION IN Z PINCHES FROM 2 – 60 keV

David J. Ampleford, Guillaume P. Loisel, Stephanie B. Hansen,  
Patrick F. Knapp, Christine A. Coverdale, Christopher A. Jennings,  
Gregory A. Rochau

*Sandia National Laboratories, Albuquerque, NM 87185*



Sandia  
National  
Laboratories

*Exceptional  
service  
in the  
national  
interest*

*ICOPS, Banff June 2016*



Sandia National Laboratories is a multi-program laboratory managed and operated by Sandia Corporation, a wholly owned subsidiary of Lockheed Martin Corporation, for the U.S. Department of Energy's National Nuclear Security Administration under contract DE-AC04-94AL85000.

This work is partially supported by Sandia's Laboratory Directed Research and Development projects 173104 and 165733.

# Overview: Non-thermal K $\alpha$ is a promising route to higher photon energy sources on Z

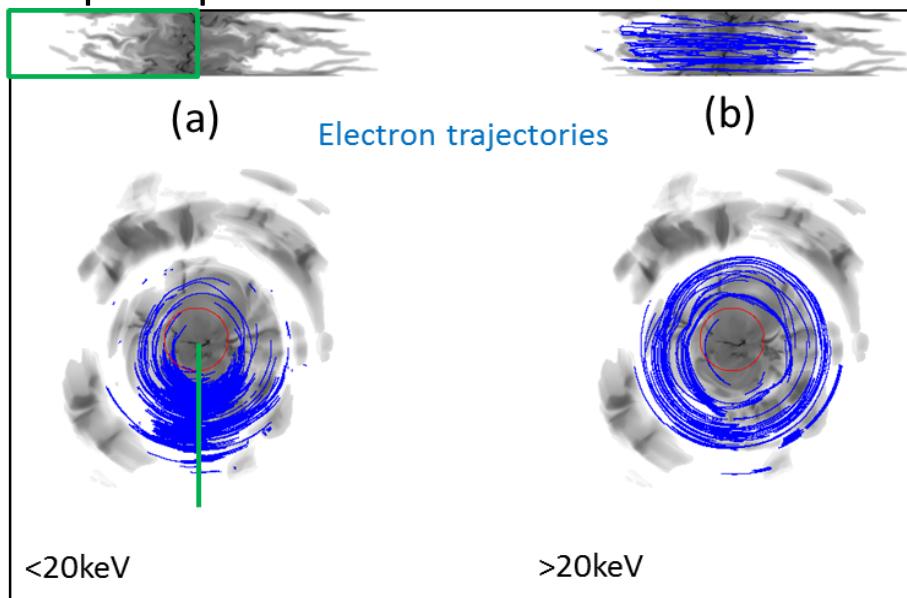
- Optimization of non-thermal K $\alpha$  emission is more complicated than creating breaks in a plasma column
- We have studied dynamics responsible for K $\alpha$  emission
  - New diagnostics have aided us in building this understanding
  - We've started to build a detailed picture of K $\alpha$  emission
- We have investigated the scaling of K  $\alpha$  emission to 59 keV
  - New diagnostics were needed to allow this assessment
  - Scaling looks similar to inferred at lower photon energies and would not be feasible with thermal emission (e.g. He $\alpha$ )

# Underlying physics is not as trivial as might be expected

- Large fields are required for electron runaway
  - Dreicer field

$$E > E_D = 5.8 \times 10^{-18} \ln \Lambda \frac{n_e}{T_e}$$

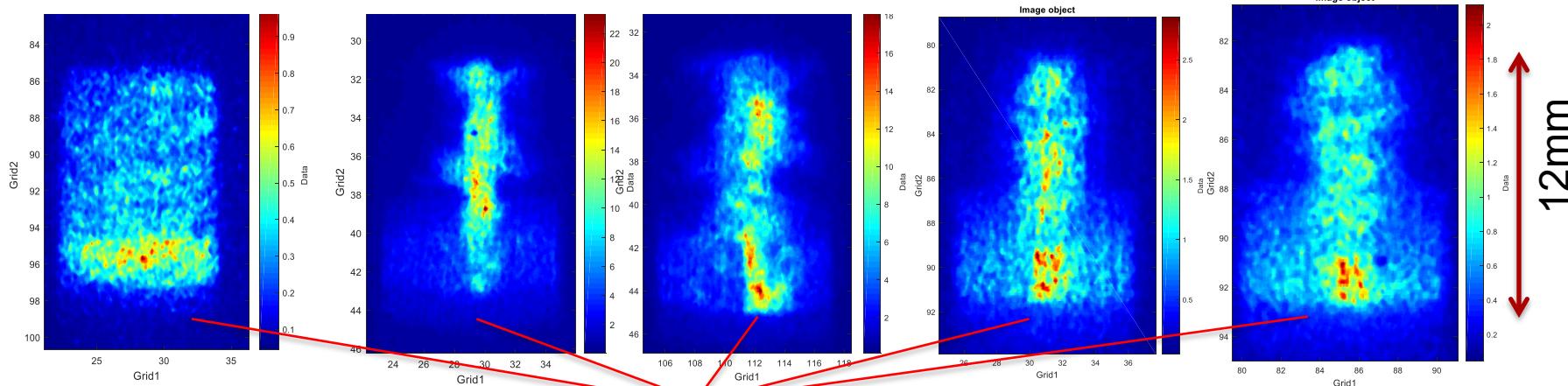
- Electrons are highly magnetized,
  - Aren't accelerated axially across gaps or radially into dense plasma
  - Trajectories post-processed from MHD simulations



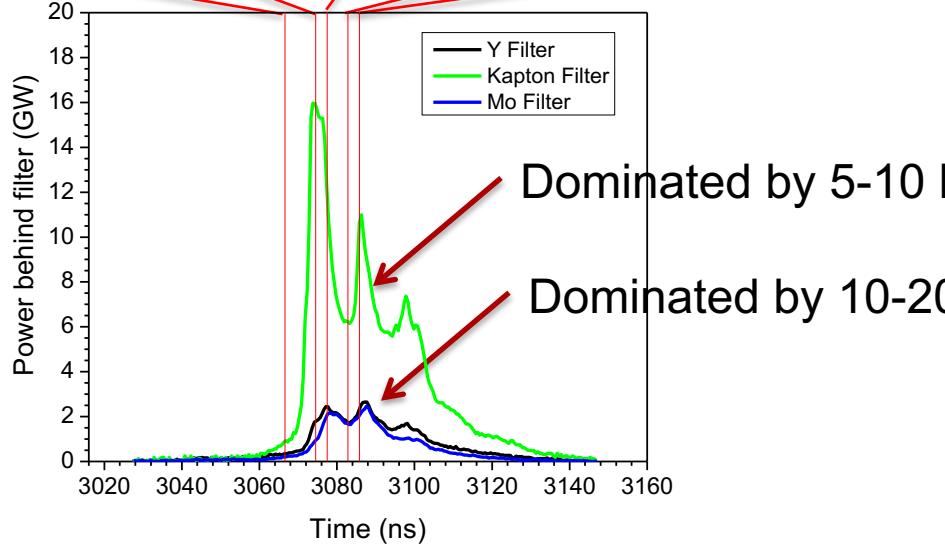
Simulation by  
J.P Chittenden *et al*

# Using 17 keV Mo K $\alpha$ emission we have studied how K $\alpha$ emission connects to array dynamics

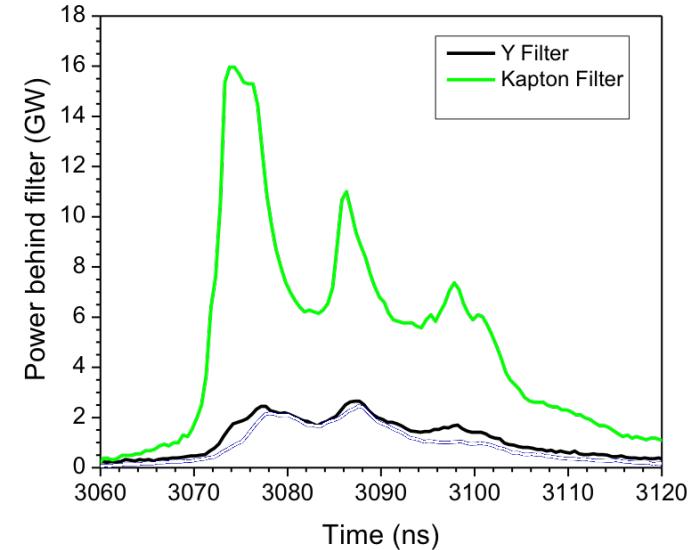
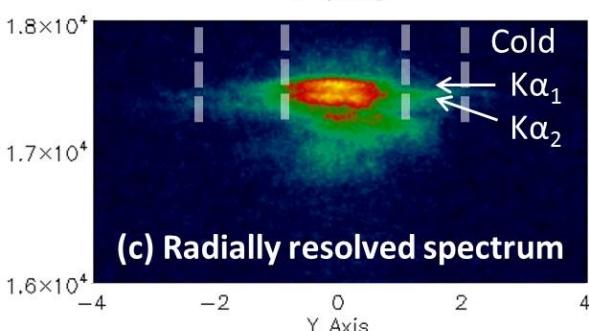
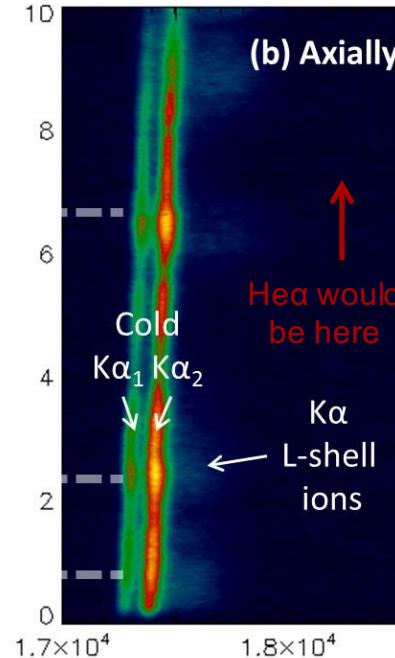
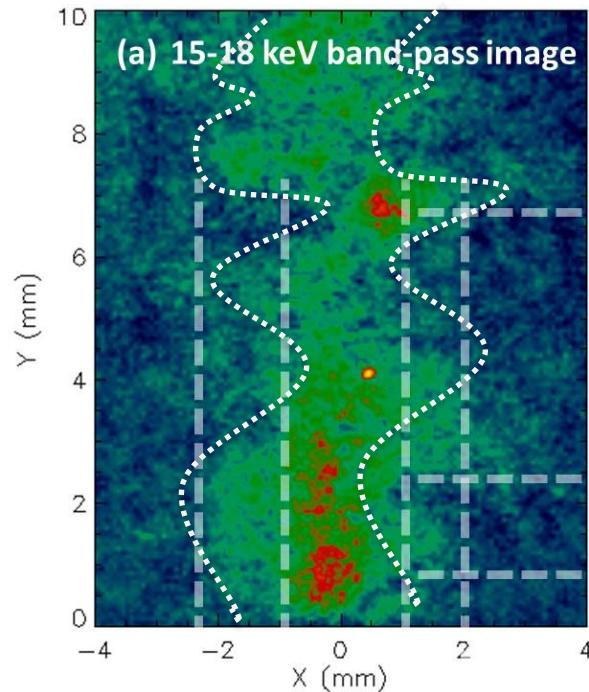
- Self emission imaging at 277 eV shows the overall dynamics



- Comparing with PCDs we see how K $\alpha$  fits in evolution

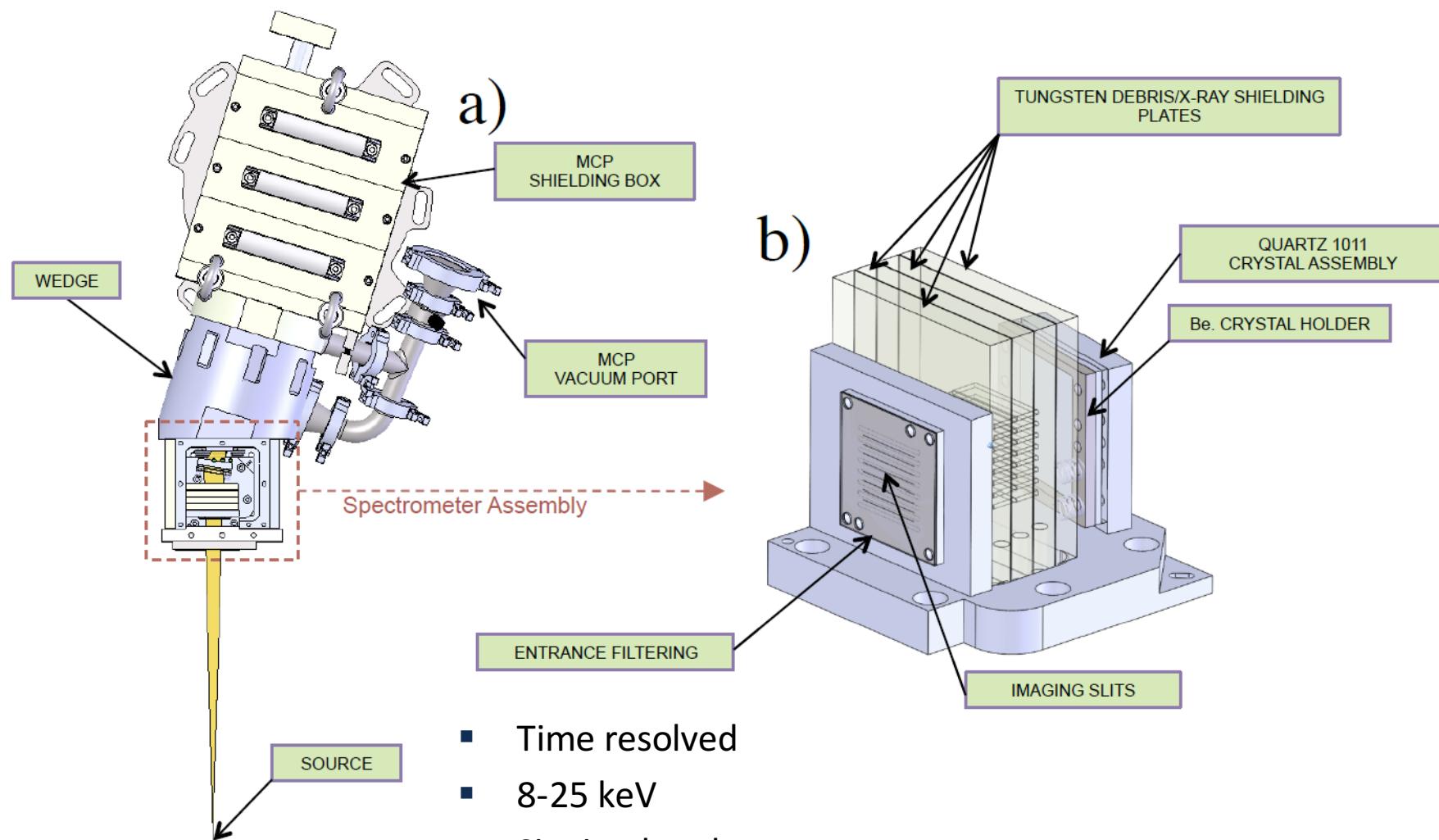


# Spatially resolved spectra combined with high energy pinhole imaging provide insight into structure



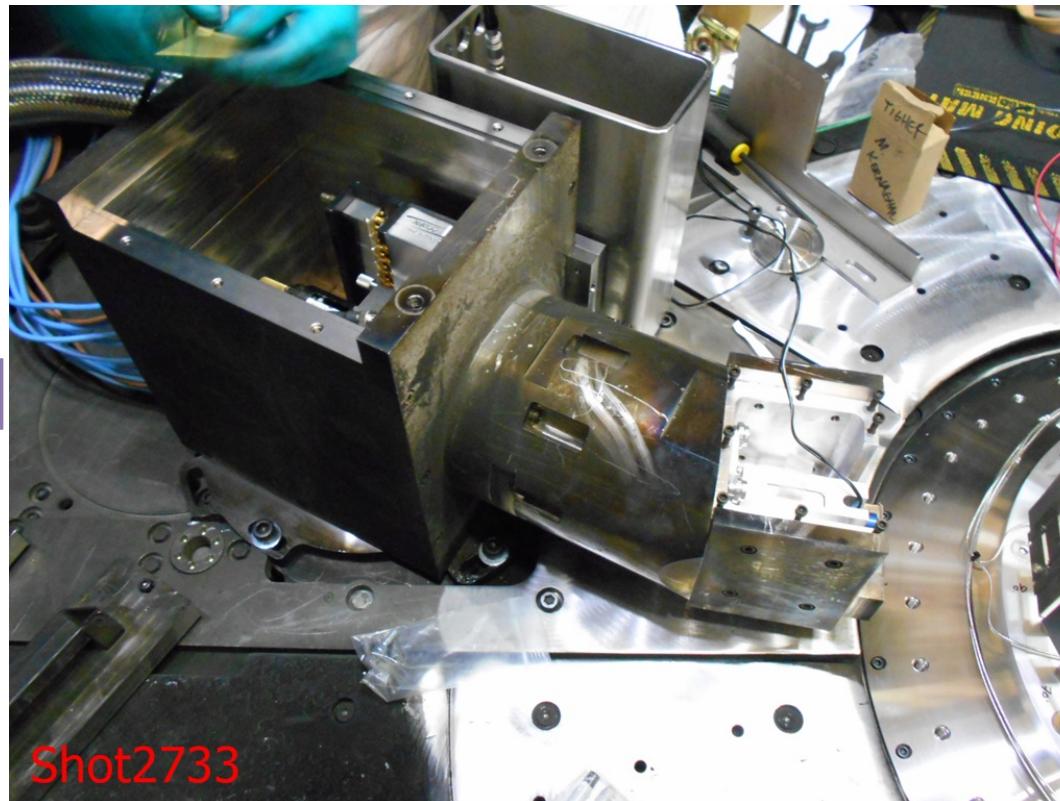
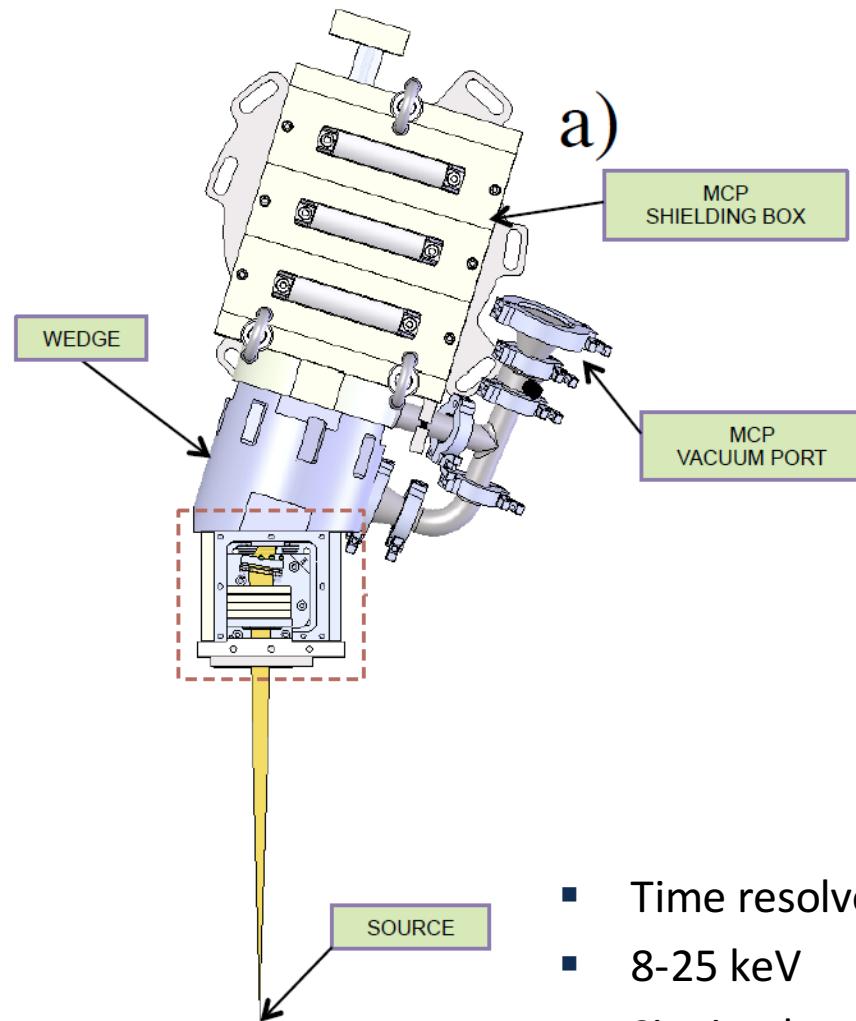
- Bright spots in 2D image and spectrum at
  - Warmer plasma in L-shell charge states
- Cold K $\alpha$  emission from most of pinch height and few mm diameter (wavelength confirms this is cold K $\alpha$ )
- Indications of quasi-helical ( $m=1$ ) structure
  - Consistent with late time emission of K $\alpha$

# New time-resolved transmission spectrometer allows us to assess evolution of K $\alpha$



P.F. Knapp et al., In preparation for RSI

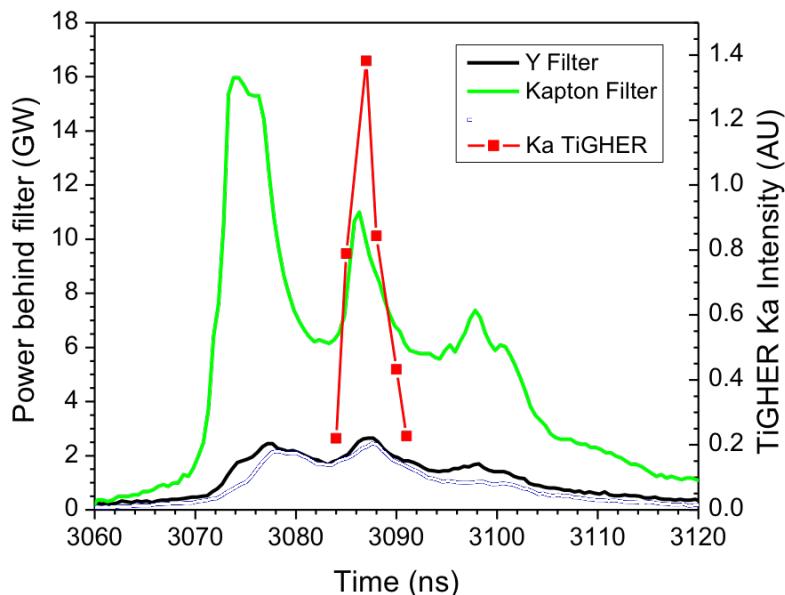
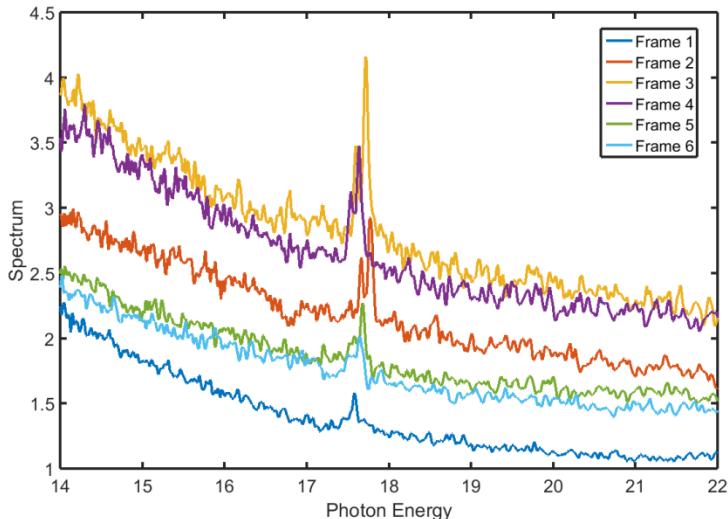
# New time-resolved transmission spectrometer allows us to assess evolution of K $\alpha$



- Time resolved
- 8-25 keV
- Sits in-chamber
- $\lambda/\Delta\lambda \sim 500$

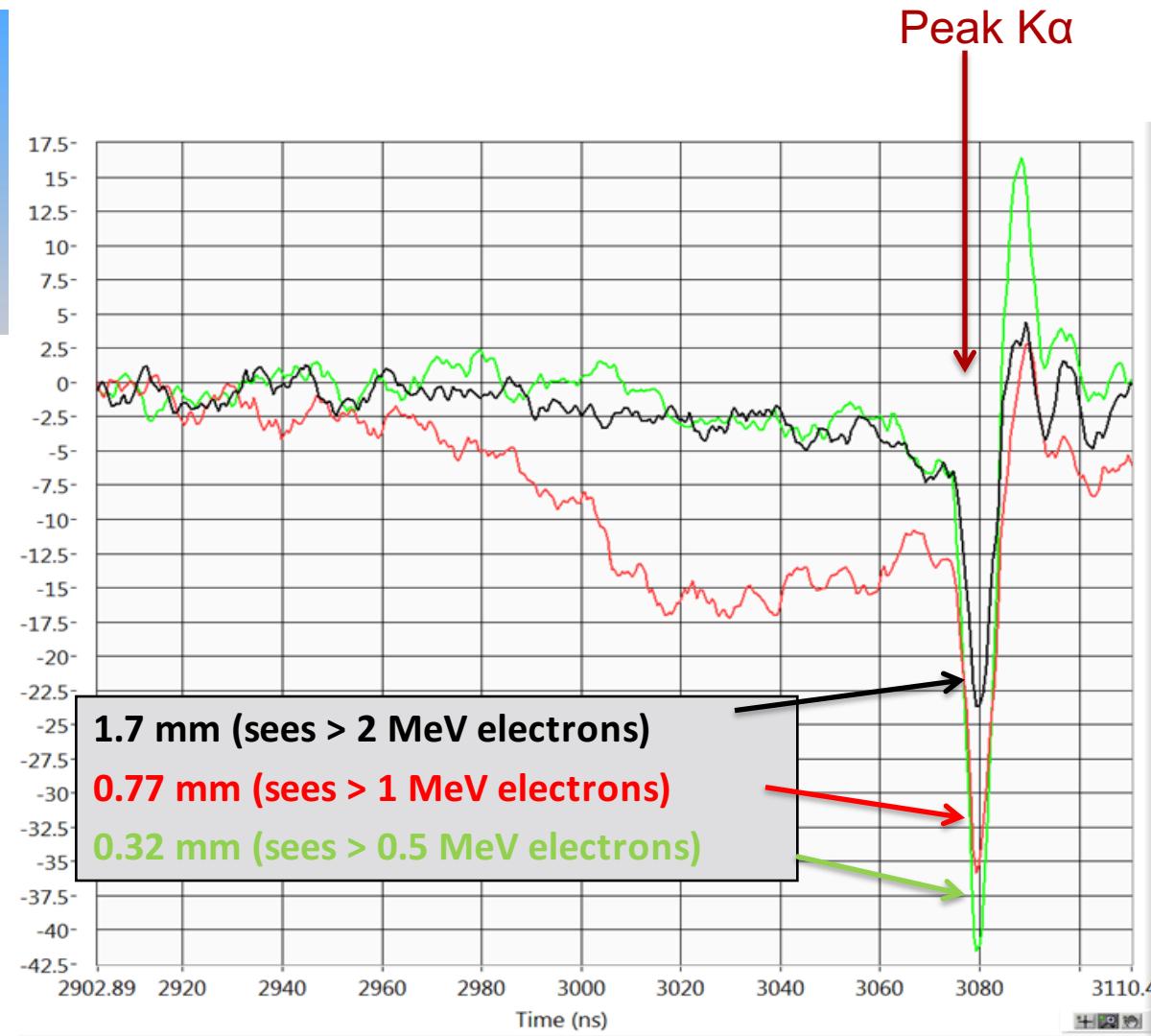
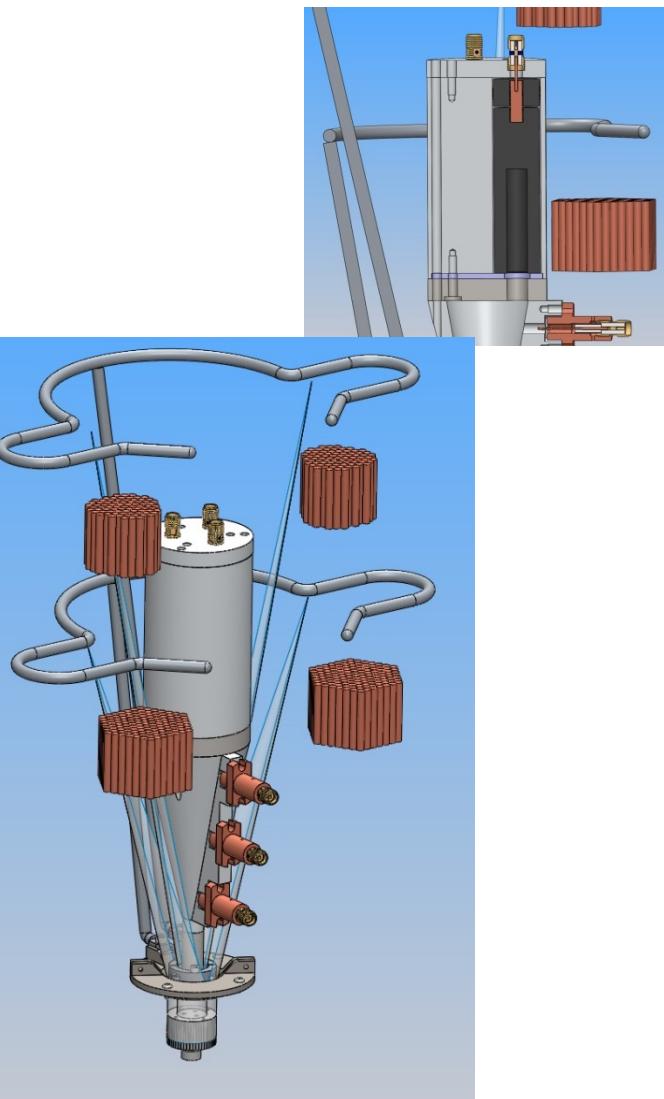
P.F. Knapp et al., In preparation for RSI

# Time resolved spectra show fast rise and fall of K $\alpha$

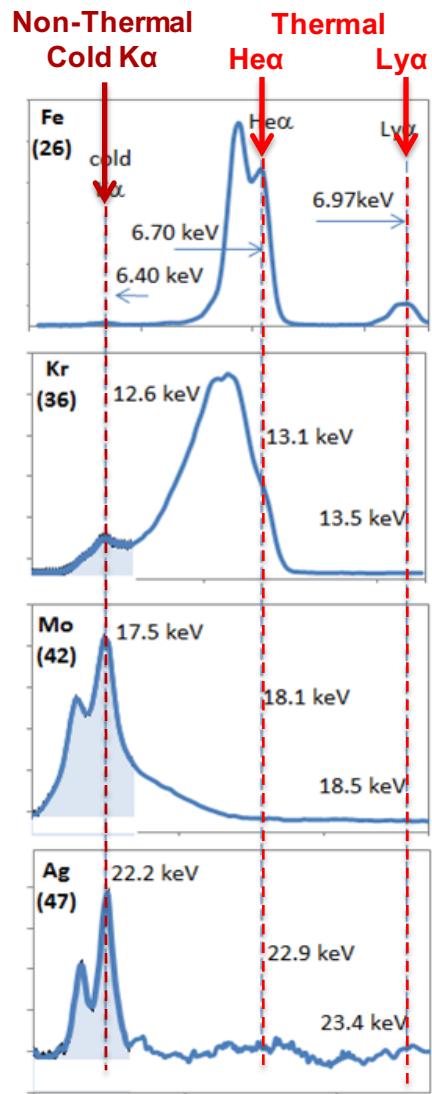


- Mo K $\alpha$  seen on time resolved spectrometer
- Rises and falls rapidly
  - ~3ns FWHM
  - Faster varying than continuum
- May be other short bursts
  - TiGHER captured 10 ns window
- K $\alpha$  pulse resembles shape of one pulse on the PCD trace
  - More work needed on checking cross timing of K $\alpha$  pulse

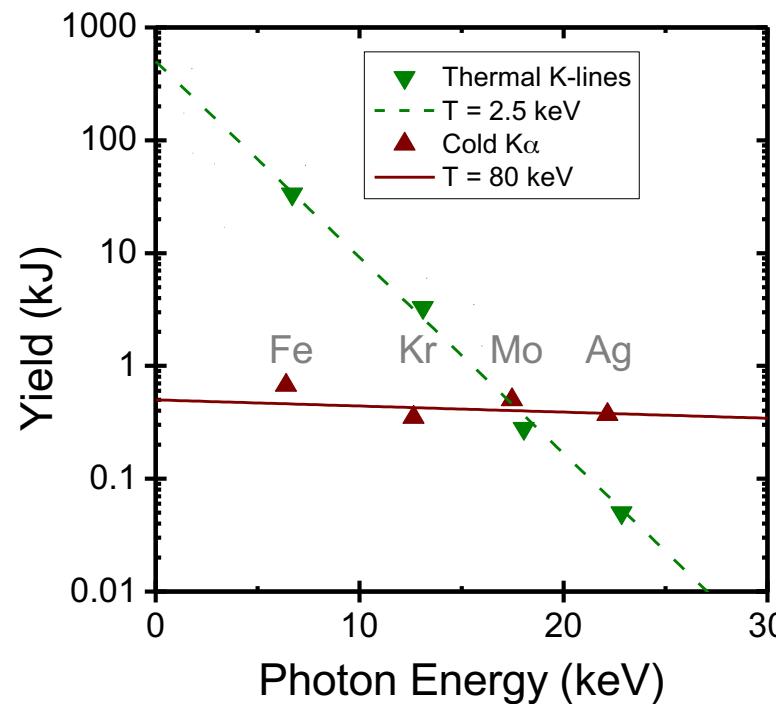
We have developed a Faraday Cup looking at energetic electrons that are present in Z pinches on Z



# Spectra from Z experiments using different materials shows transition from thermal to non-thermal emission



- As atomic number increases see transition
  - from thermal K-shell (Fe almost entirely He $\alpha$ )
  - to non-thermal K-shell (Ag entirely cold K $\alpha$ )
- Scaling to higher energies is much better with cold K $\alpha$

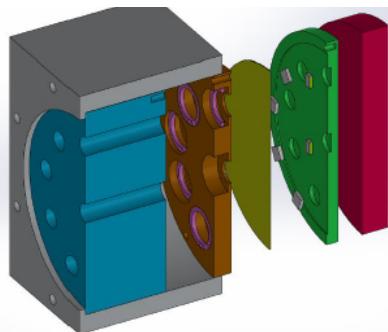


# To assess higher photon energy K $\alpha$ & Bremsstrahlung we have developed a number of diagnostics

## Filtered TLD detectors

Absolute fluence measurement 10keV-20MeV

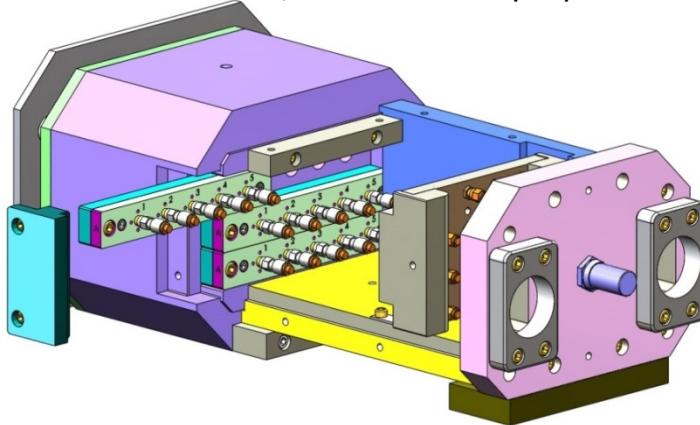
V. Harper-Slaboszewicz *et al.* *In prep.*



## Differentially Filtered Diodes

Absolute fluence measurement 10keV-20MeV

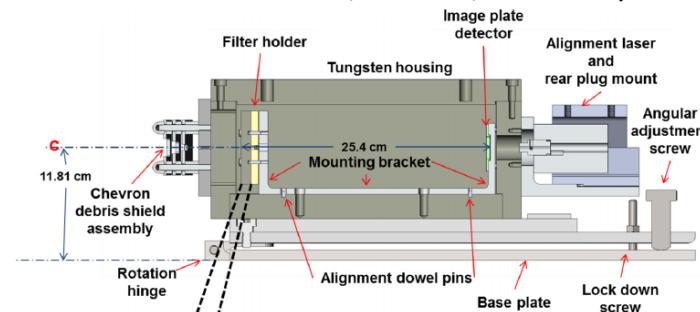
K.S. Bell, Talk 5F-7 & *In prep.*



## Time Integrated PHC

Spatial structure 15 keV-100 keV

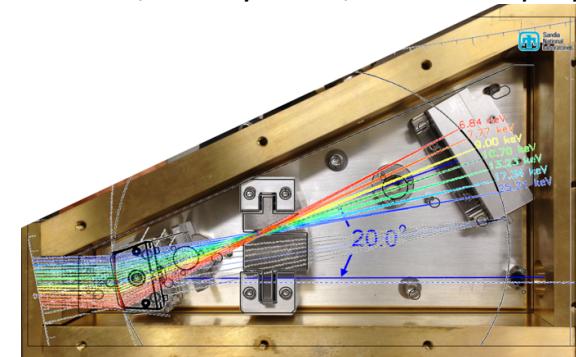
L.A. McPherson *et al.*, RSI **87**, 063502 (2016)



## Space resolved spectra

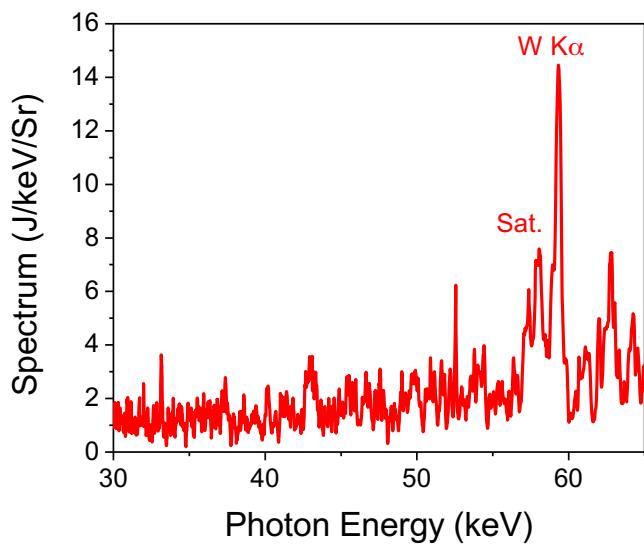
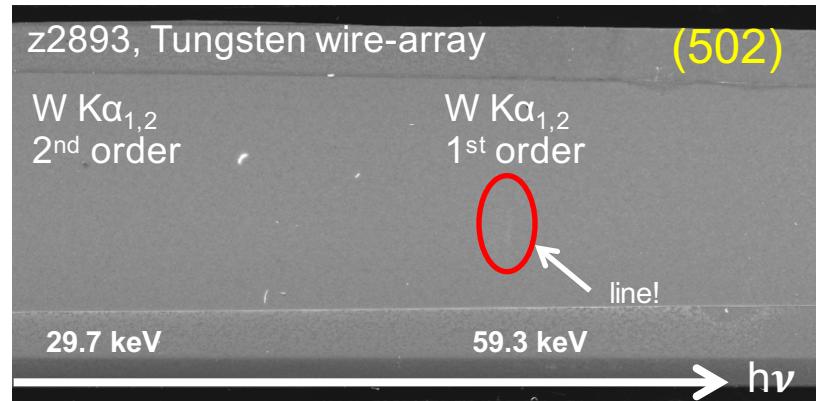
Spectra 7 keV-100 keV

G. Loisel, J Seely *et al.*, HTPD & *In prep.*

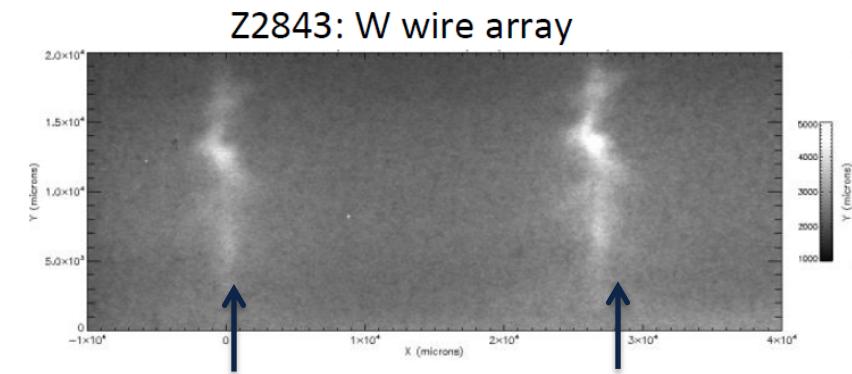


# Recent data has shown first evidence of W K $\alpha$ emission on Z

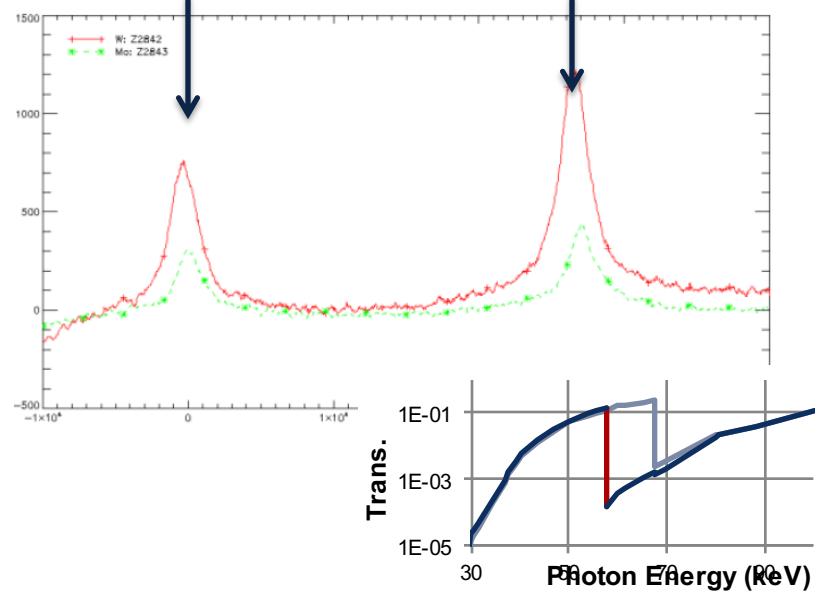
Spectra show evidence of line



Ross paired for  $\sim$ 60keV gives image

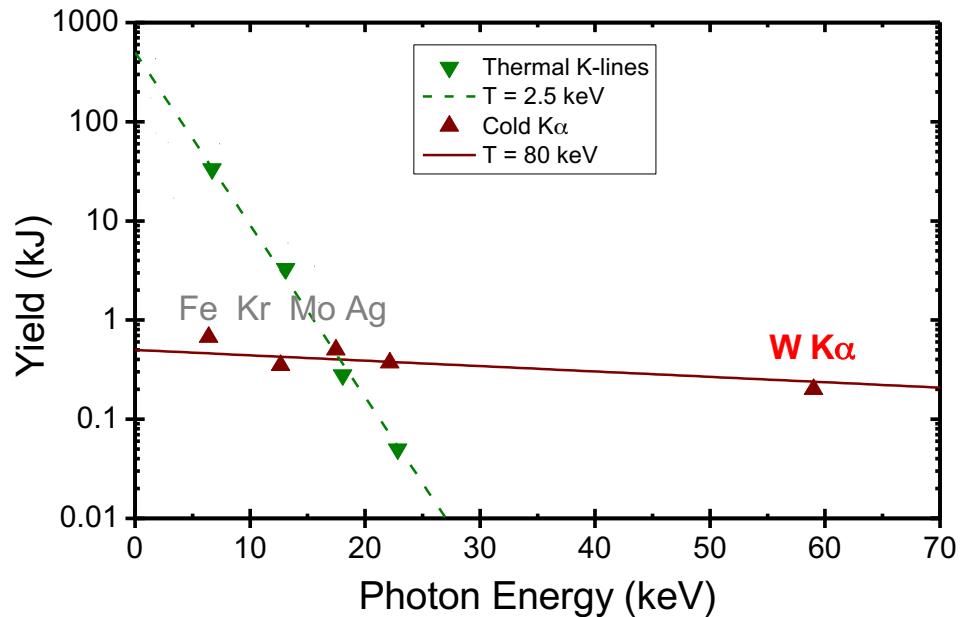


Insensitive to 59 keV      Sensitive to 59 keV



# Scaling to W K $\alpha$ is consistent with lower energy data

- We have previously assumed an energetic electron population at energies  $\sim$ 80 keV to provide scaling through K $\alpha$  data
  - Not tightly constrained, especially with small range in  $h\nu$  and small dependence on  $T_e$
- W data point fits well and provides first data point at  $>30$  keV for wire arrays



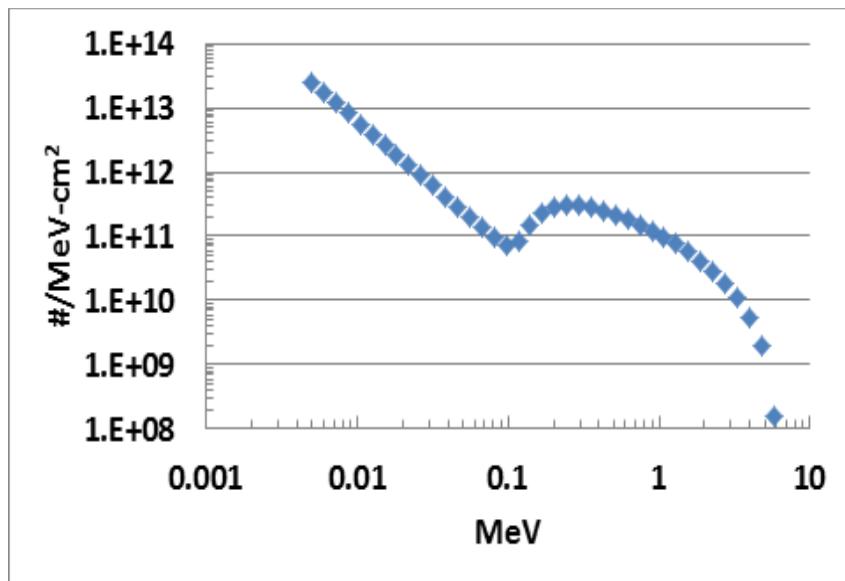
Calibration of Quartz 502 CRITR-X setup by Loisel, Seely *et al.*  
W yield unfold by G. Loisel

# We are also taking steps to better measure Bremsstrahlung from wire arrays on Z

## Filtered TLD detectors

Absolute fluence measurement 10keV-20MeV

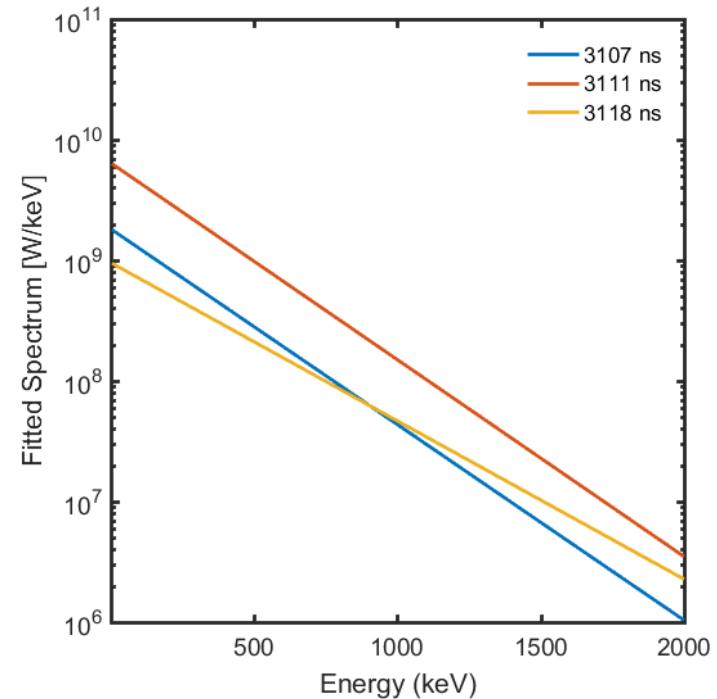
V. Harper-Slaboszewicz *et al.* *In prep.*



## Differentially Filtered Diodes

Absolute fluence measurement 10keV-20MeV

K.S. Bell, Talk 5F-7 & *In prep.*



# Overview: Non-thermal K $\alpha$ is a promising route to higher photon energy sources on Z

- Optimization of non-thermal K $\alpha$  emission is more complicated than creating breaks in a plasma column
- We have studied dynamics responsible for K $\alpha$  emission
  - We've started to build a detailed picture of K $\alpha$  emission
- We have investigated the scaling of K $\alpha$  emission to 59 keV
  - Scaling looks similar to inferred at lower photon energies and would not be feasible with thermal emission (e.g. He $\alpha$ )

