

Mix in MagLIF Stagnation

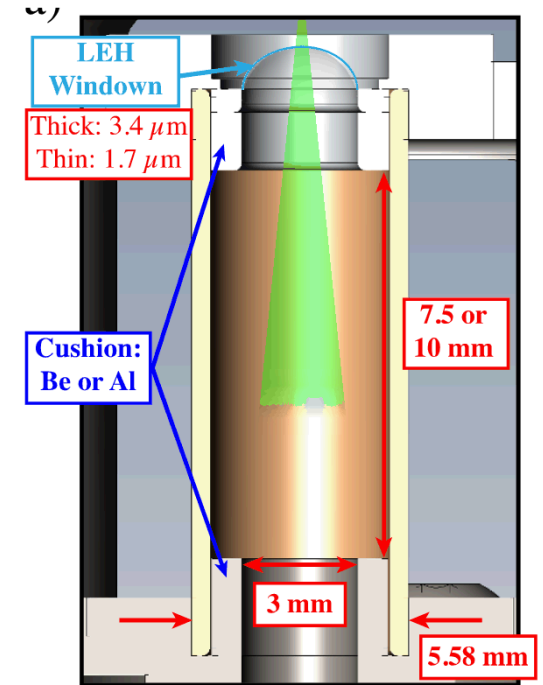
Greg Rochau, Eric Harding,
Stephanie Hansen, Matt Gomez, Pat
Knapp, Kyle Peterson

Questions to Address

- What do we know today and how do we know it (what measurements, model evidence)?
- How does our measurement knowledge differ from our models?
- What are the big questions we need to address?
- What experiments and/or simulations can we conduct to address these questions?
- What new measurements, diagnostics, and/or models are needed?

What do we know today and how do we know it (what measurements, model evidence)?

- Mix from the endcaps affects the yield
 - Be endcaps result in a $\sim 10\times$ higher yield than Al endcaps
- Be from the liner and/or endcaps end up in the hot stagnation plasma.
 - Fe He-like emission is seen in time-integrated spectra
- The intensity ratio of the Fe He-like lines to the continuum is consistent with a few % Be mix.
 - Assumes a lot...
- Early-time mix can significantly affect stag performance
 - 0.1% Ar in the DD duds the yield.
- Mix from the laser-heating phase is probably important
 - Trend of decreasing yield with increasing laser energy
- The LEH window may be a source of mix.
 - Shot with window on bottom of cushion had $\sim 30\times$ less yield
 - Shot using phase-plate had $\sim 20\times$ less yield



In Activation

-DD Yield

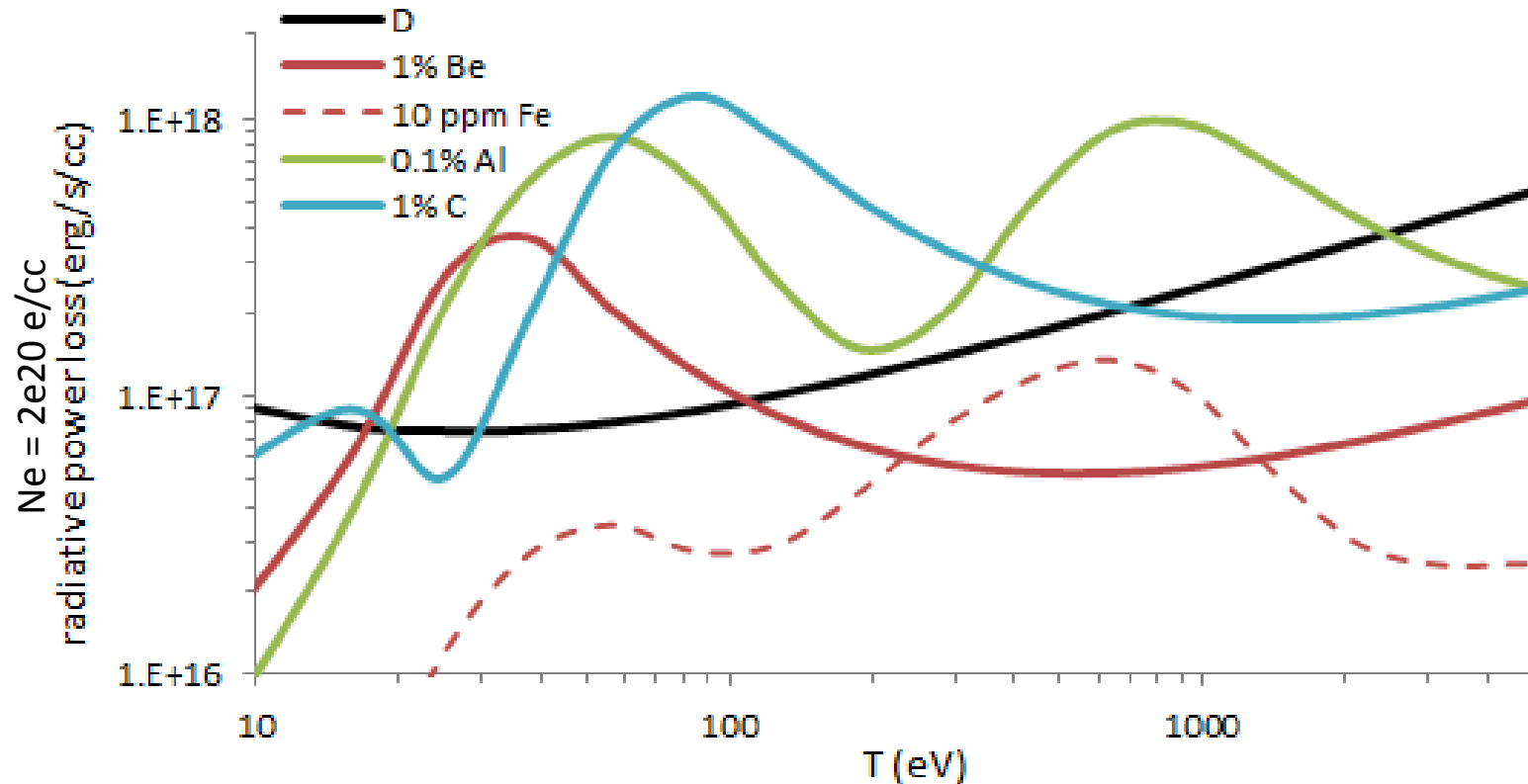
Crystal Spectrometers

-Lines and Continuum

Crystal Imager

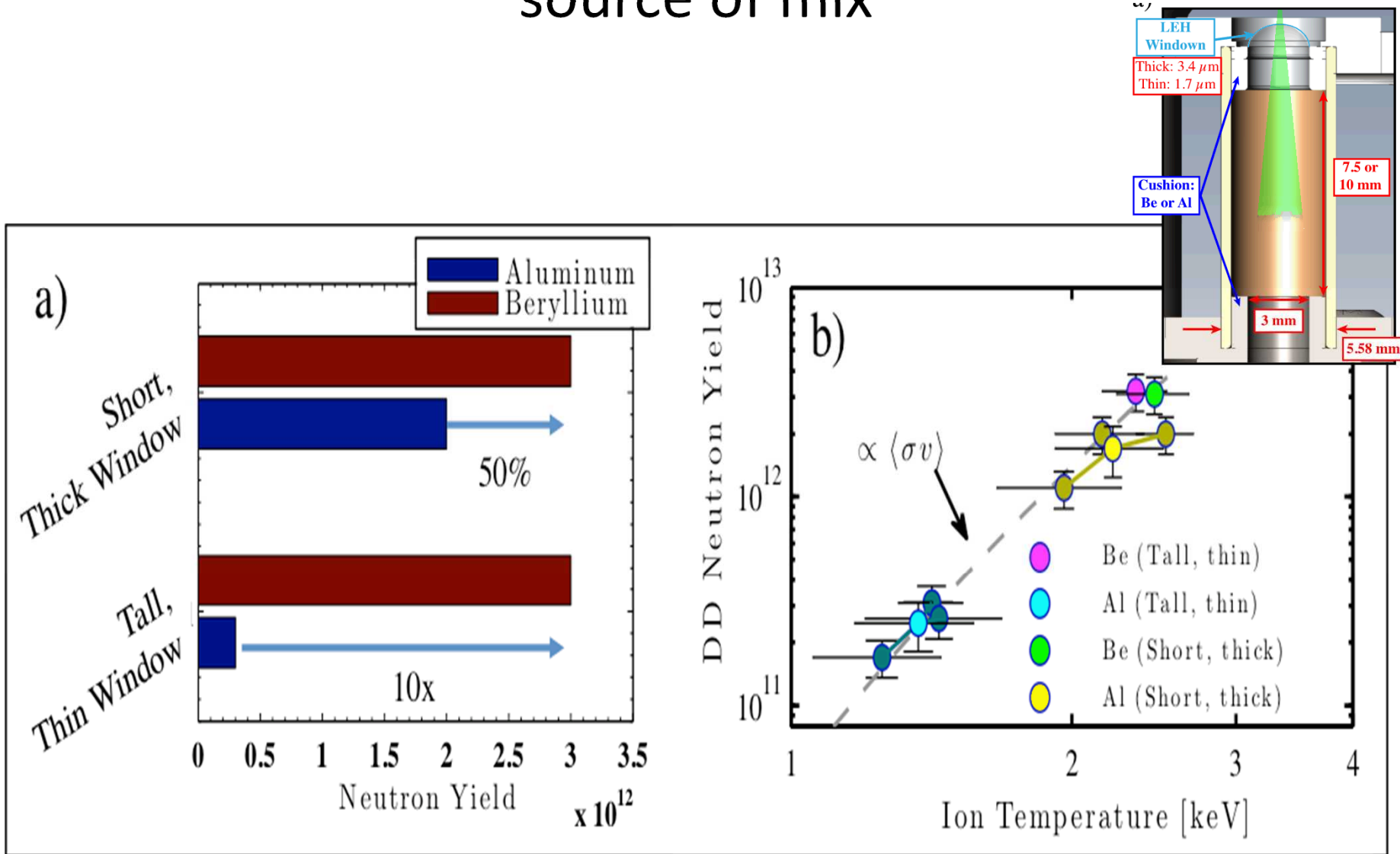
-Hot plasma distribution

Radiative losses vary with material and charge state

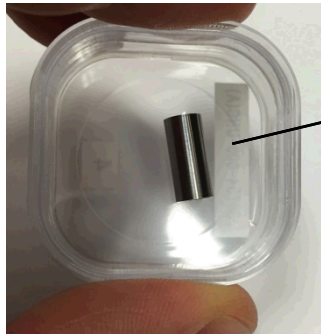


- Mix levels of ~1% Be are tolerable (~2x degradation)
- 0.1% Aluminum is devastating (~10x degradation)
- 1% carbon is devastating (~10x degradation)
Note that uniform mix of half thin (thick) window mass gives ~3% (8%) C
- Mid-Z impurities in Be don't hurt at all, even at 10x detected levels

Yield and Ti are consistent with endcaps being a source of mix

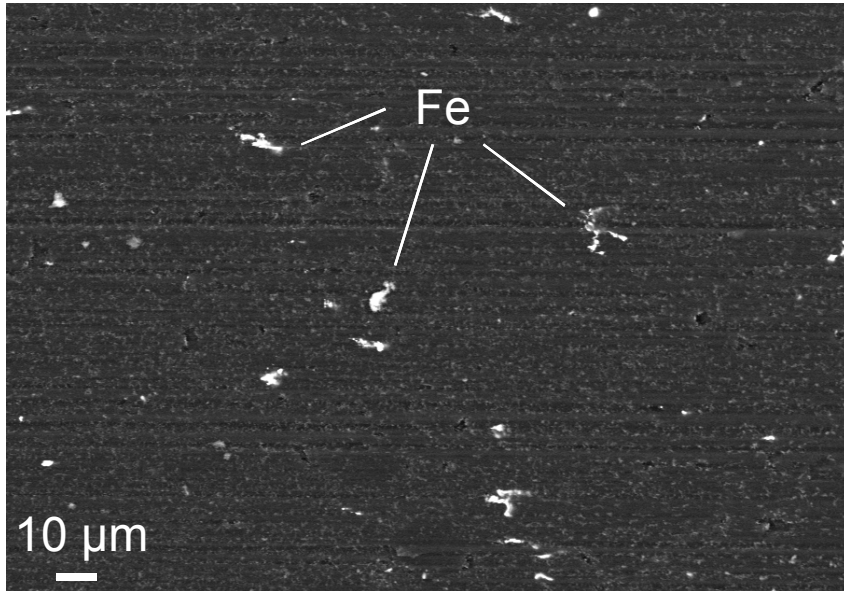


Fe occurs “naturally” in our S-65 grade Be liners and endcaps.



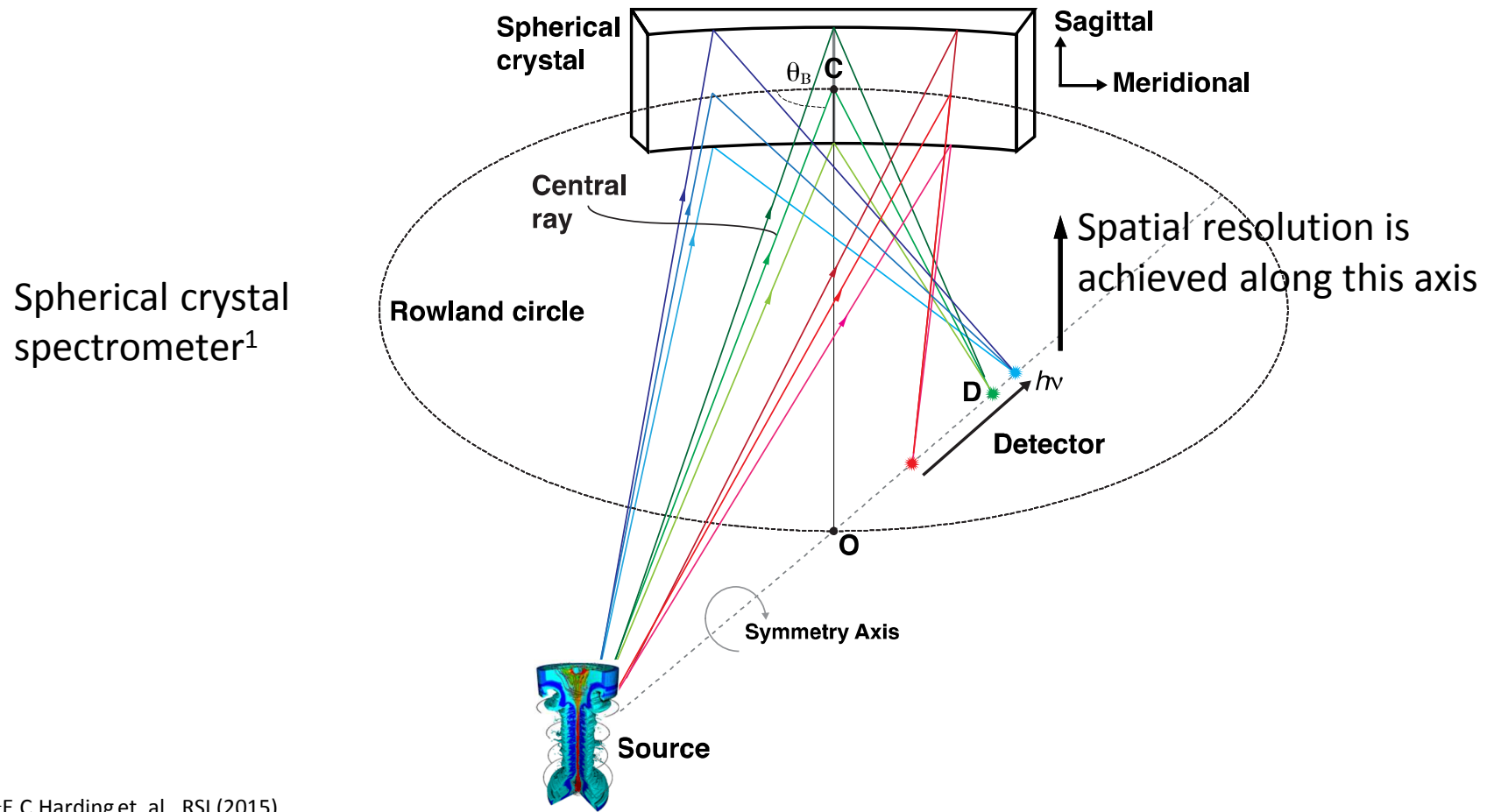
Machined
Be liner

SEM image of the Be liner outer surface



- The Be stock material contains ~100 ppm (0.01% atomic) of Fe that originate from the manufacturing process.
- The Fe appears as micron-sized impurities that are uniformly distributed on the visible surfaces. We are assuming it is uniform in the bulk.

To resolve the Fe emission generated at stagnation we use a spherically-bent crystal spectrometer.



¹E.C.Harding et. al., RSI (2015)

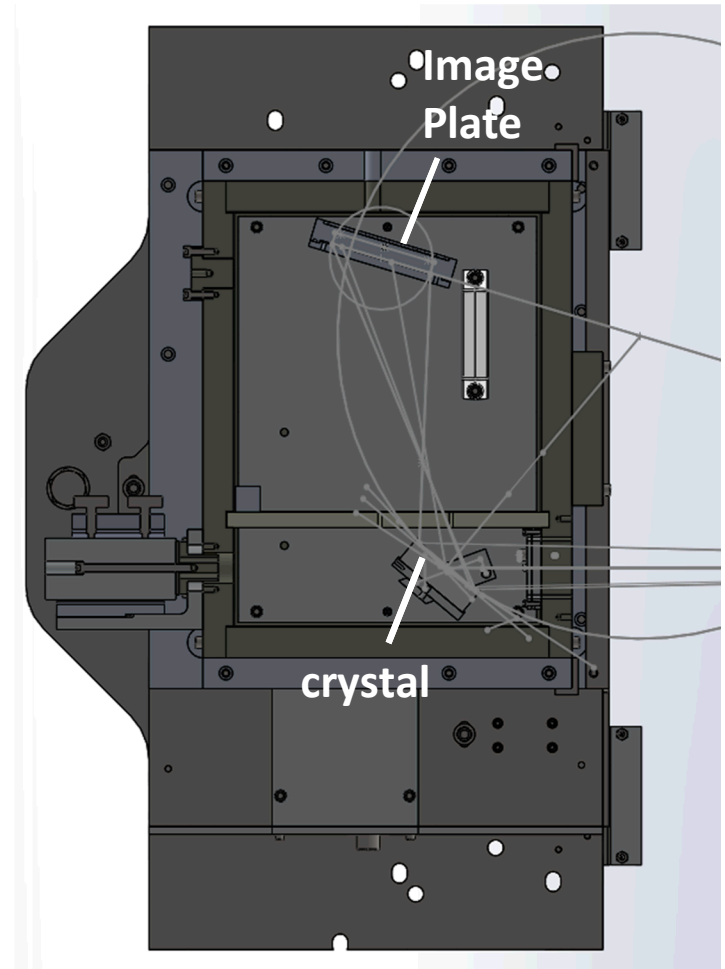
D. Sinars et. al. JSQRT (2006)

FSSR used on dynamic hohlraum capsule implosions

X-Ray Scattering Spherical Spectrometer (XRS3) Spectrometer

Spectrometer setup for He-like Fe emission

Crystal	Q20-23 ($2d = 2.749 \text{ \AA}$)
Source-to-crystal	800 mm
Crystal-to-detector	256.92 mm
Crystal Radius	250 mm
Center Bragg Angle	40°
Crystal size ¹	60 x 36 mm
Spectral Range²	6328 - 7977 eV
Spatial Mag. (M_{sag})	0.30x
Spectral Resolution³	2 eV
Spatial Resolution³	210 μm
Throughput	$1.9\text{e-}7$ steradians

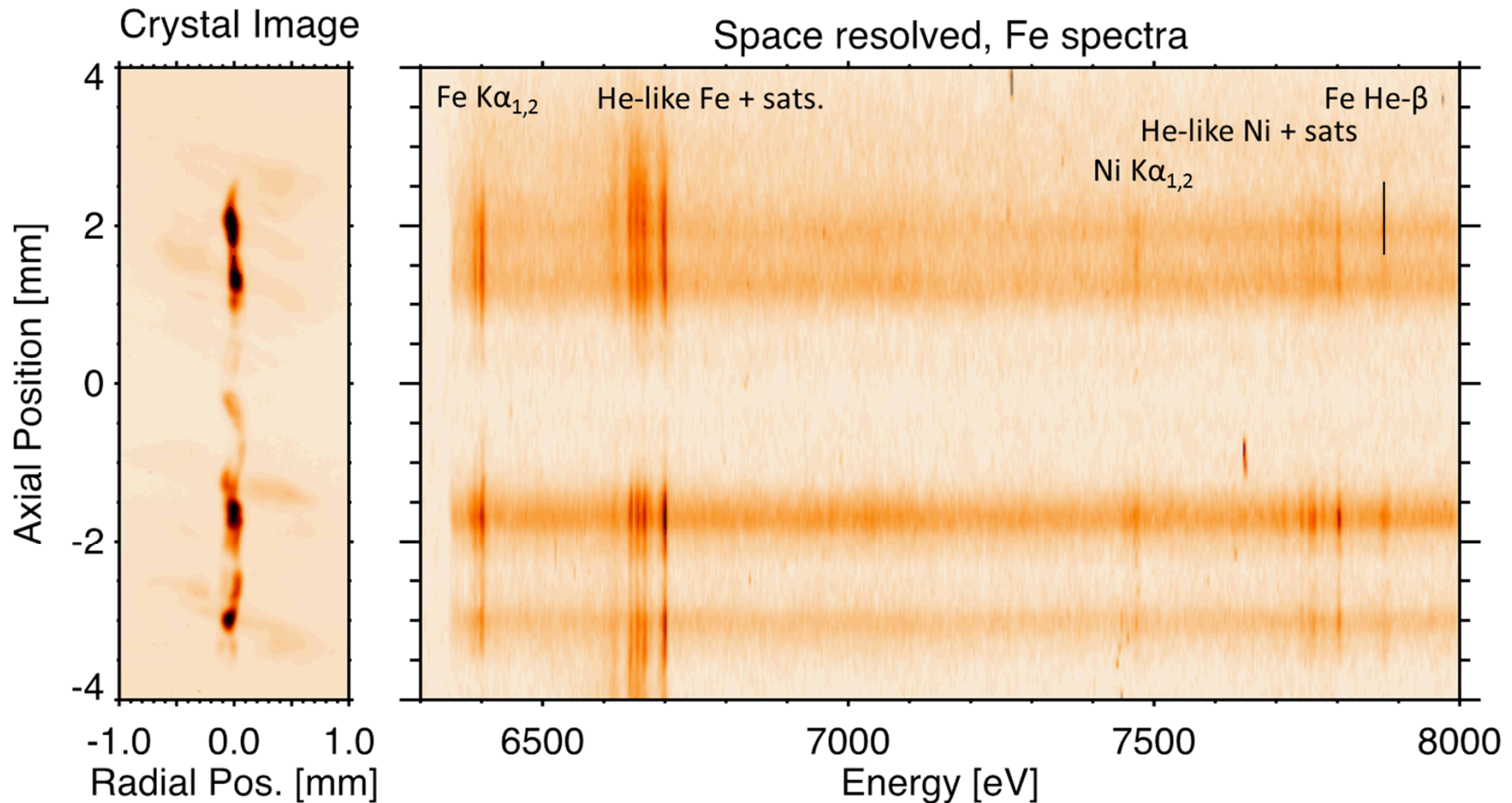


¹This is a tiled crystal consisting of 2 strips, each one is 60 x 18 mm

²Detector length must be 85 mm to capture entire spectral range.

³Limited by the Image Plate resolution of 63 microns.

We believe we are observing He-like Fe emission from stagnation.



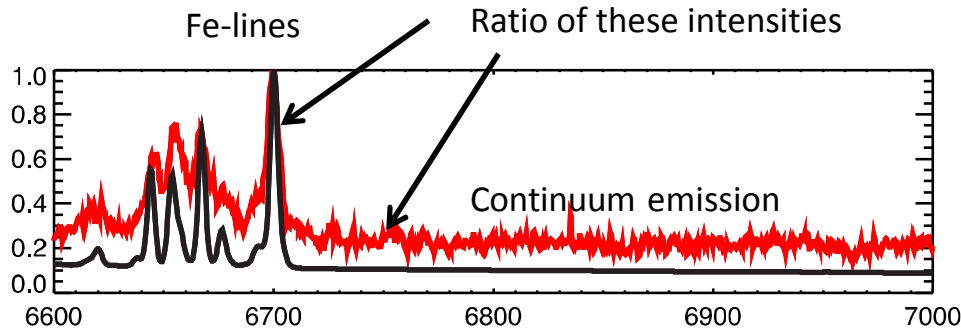
Z2850

Target: 7.5 mm tall Be AR6 liner with Be cushions and a 3.5 μm LEH window

Yield: 3×10^{12} DD (highest performing MagLIF shot to date)

Ratio of Fe line emission and nearby continuum emission is sensitive to mix fraction.

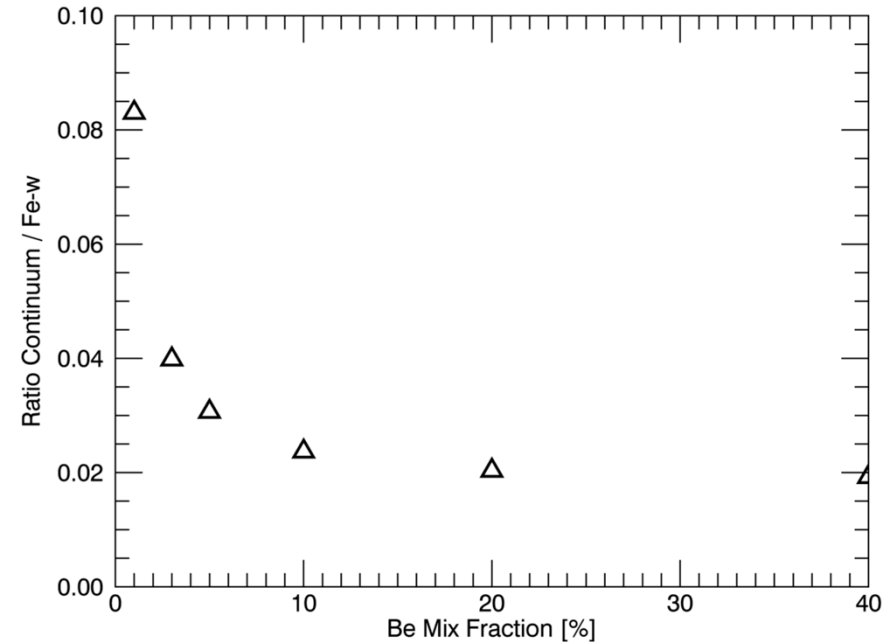
XRS3 data and PrismSPECT



Key Assumptions:

- Lines and continuum are from the same plasma at the same time
 - If continuum has other contributions, then this is a lower bound on mix
- Fe/Be ratio is fixed throughout the volume
- Others?

Intensities Ratios vs. Be mix



PrismSPECT runs use
D + Be + Fe

Relative amounts of each are varied
w/ fixed total mass density and T_e .

Example: (Z2839) Be mix fraction appears to be 1-3%

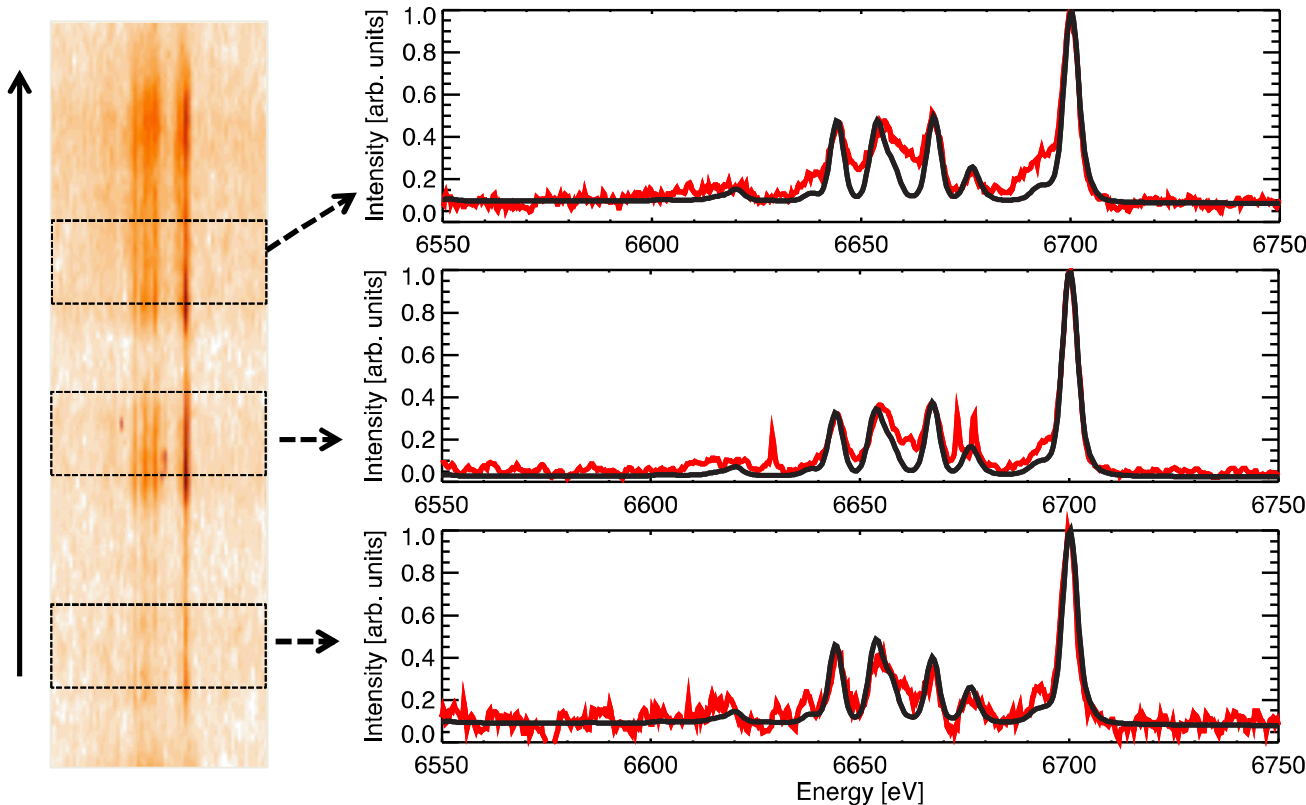
Experimental spectra fitted with PrismSPECT
simulations using $E/\Delta E = 3000$.

Inferred values

$T_e = 1.5 \text{ keV}$
 $n_e = 1.2e23 \text{ cm}^{-3}$
Be mix $\sim 1\%$

$T_e = 1.6 \text{ keV}$
 $n_e = 1.7e23 \text{ cm}^{-3}$
Be mix $\sim 3\%$

$T_e = 1.4 \text{ keV}$
 $n_e = 2.0e23 \text{ cm}^{-3}$
Be mix $\sim 1\%$



Z2839

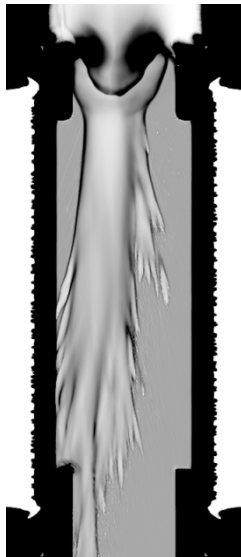
Target: 10 mm tall Be AR6 liner with Be cushions and a 1.5 μm LEH window

Yield: $3E12$ DD (highest performing MagLIF shot to date)

How does our measurement knowledge differ from our models?

- Models indicate mix can occur from multiple origins:
 - Blast wave from laser preheat causes blowoff from liner wall and endcaps
 - Laser can pass through the gas and cause blowoff from the bottom end cap
 - Laser can deflect through LEH plasma and hit the liner/endcap causing blowoff
 - The exploded LEH window can mix into the gas
 - The liner is RT unstable

**Laser
deflected**



**Laser on
bottom cap**



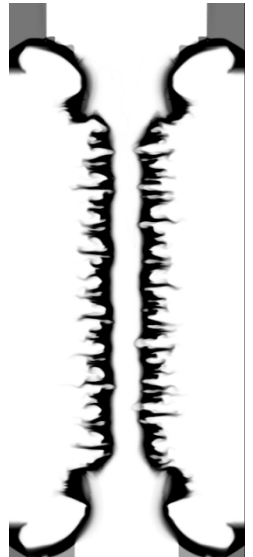
**Blast Wave
Reflected**



**LEH window
injected**



**Unstable
liner**

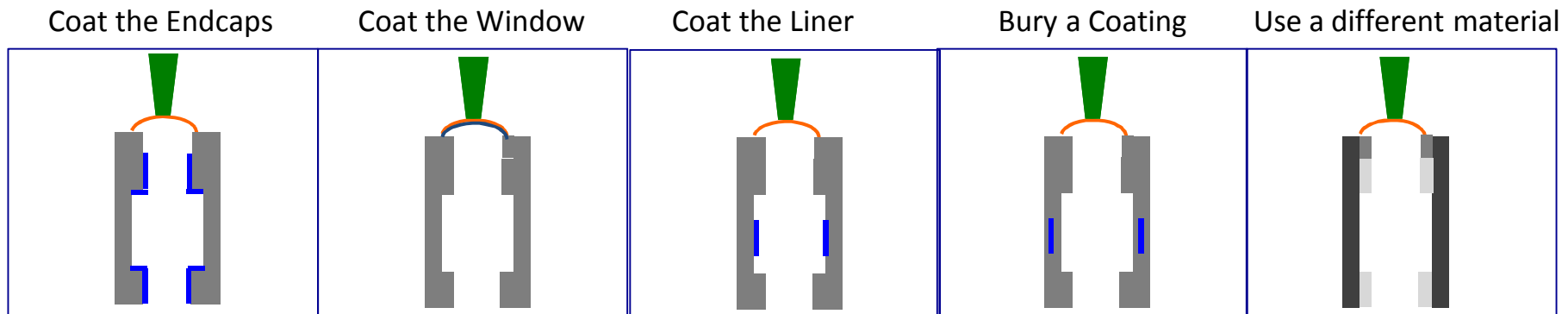


What are the big questions we need to address?

- What are the true origins of the Be mix and at what level of relative importance?
 - Present data is ambiguous on the source and simulations indicate it may come from multiple places.
- How is the Be mix distributed in space and time?
 - Present interpretation is that Fe emission all comes during stagnation
 - Spectrometers integrate over at least 2 spatial dimensions
- Is the LEH window a significant contributor to mix?
- How can we reduce the amount of mix from preheat where it has the largest impact?

What experiments and/or simulations can we conduct to address these questions?

- Use localized coatings/tracers to assess the origins of the mix.
 - Liner (Co), endcap (Co), window (Co, Cl)
 - This is our baseline approach; ~1 week per year of dedicated development on Z
 - Should use OMEGA-EP and Pecos (ZBL) to develop techniques for window coatings
- Use different liner/endcap materials and assess impact on yield and/or intensity of line emission.
 - Li (target development in progress), high purity Be
- Assess mix with different laser preheating.
- Assess yield with various LEH window positions.
 - A coupled problem

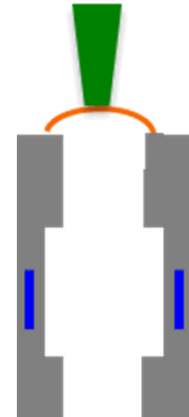


Future experiments and diagnostics

- Time-integrated x-ray emission lines from dopant, impurities, and tracers can unambiguously track material fractions at stagnation
 - Co coating on endcaps (and windows) will emit in range of XRS3
 - Ultra-pure and high-impurity Be can help isolate mix source
- In general, x-ray diagnostics must dance along multiple fine lines:
 - AR-6 liner is opaque to $h\nu < 3$ keV initially (and $h\nu < 6$ keV at stagnation) $\rightarrow Z \geq 18$ (26)
 - Detectable emission $\rightarrow T \geq \sim h\nu/3$ and ~ 100 ppm at preheat (\sim few ppm at stagnation)
 - Concentrations > 100 ppm of $Z = 18$ (or ~ 10 ppm of $Z = 36$) degrade neutron yields by 2-3x and decrease (or eliminate) x-ray signals \rightarrow
 - Fuel dopants can be precisely fixed but are restricted to gases (Ar/Kr)
 - There are many good options for endcap/liner tracers (Ti/Co/Zn) but since we lack a predictive simulation capability for dynamic mix, we must rely on empirical data
 - Window tracers can be volumetric (Cl) or coated, but must not interfere with laser energy deposition
- With 500 J of preheat energy and AR12 liners, time-dependent K-shell spectroscopy of 100 ppm Ar in fuel and 1% Cl window could provide temporal evolution of fuel temperature and mix

What new measurements, diagnostics, targets, and/or models are needed?

- Time-resolved spectroscopy of tracers
 - Requires single line-of-sight gating...
- 2-D spatial imaging of tracers...
- Be liners with buried tracer layers
 - Assess depth of the liner that penetrates stag column
- High purity Be Liners...
- Li liners (in development)
 - Reduce impact of liner mix
 - Oxygen contaminants need to be controlled
- Cryo DD layers?

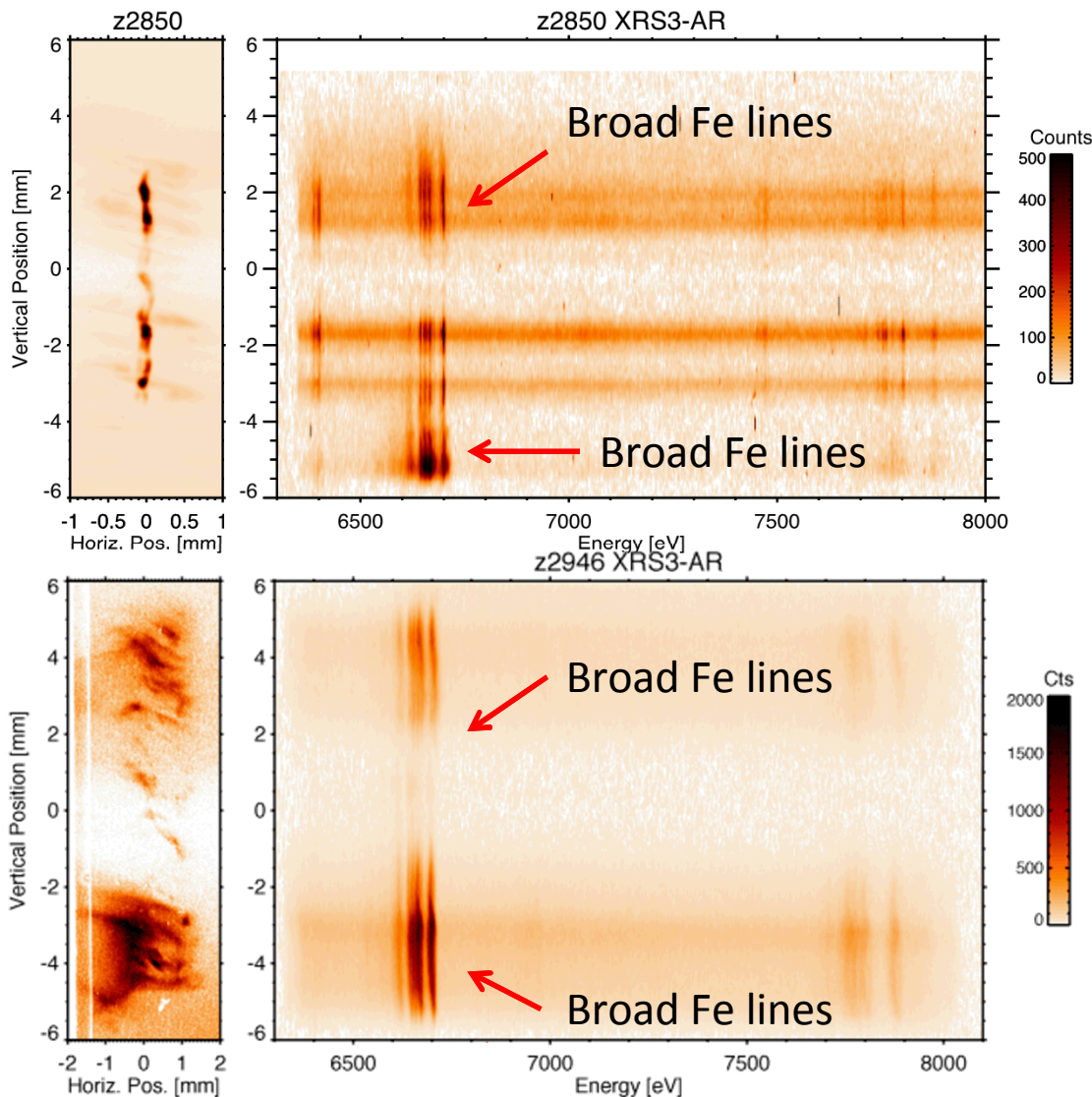


Materion's UHP9999 material is nearly a single crystalline piece of Be. This is the highest purity Be we can get.



Initial discussions with Don Hashiguchi and Jerry Holman (Materion Elmore, OH):
UHP 9999 is fragile. Machining will likely require a careful EDM process and possibly diamond turning. EDM only (i.e., no SPDT) may yield $R_a \sim 400\text{nm}$. 1 liner is ~\$8k.

Fe lines can come from somewhere other than the stagnation column



Full MagLIF shot

(z2850), $Y_{DD} = 3e12$:

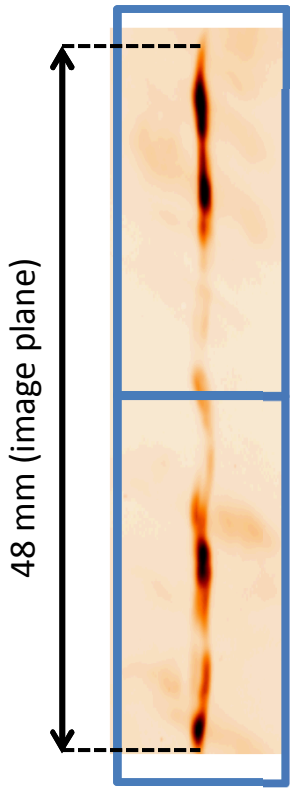
Fe spectral lines near the top and bottom of the target appear broadened.

Implosion only shot (z2946), no laser heating:

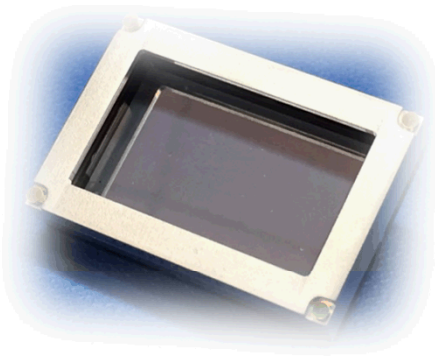
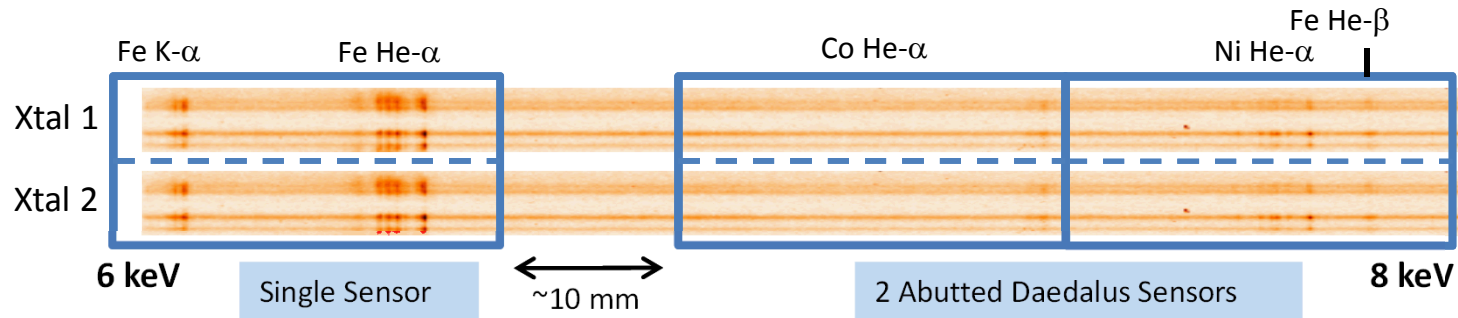
No stagnation column but there are strong Fe signals appearing near the ends. This spatially broad source of Fe emission maybe contaminating other shots like z2850.

Fast gated CMOS sensors will be used to separate stagnation from other events

Spherical Crystal 2-D Imager (Mag 6)



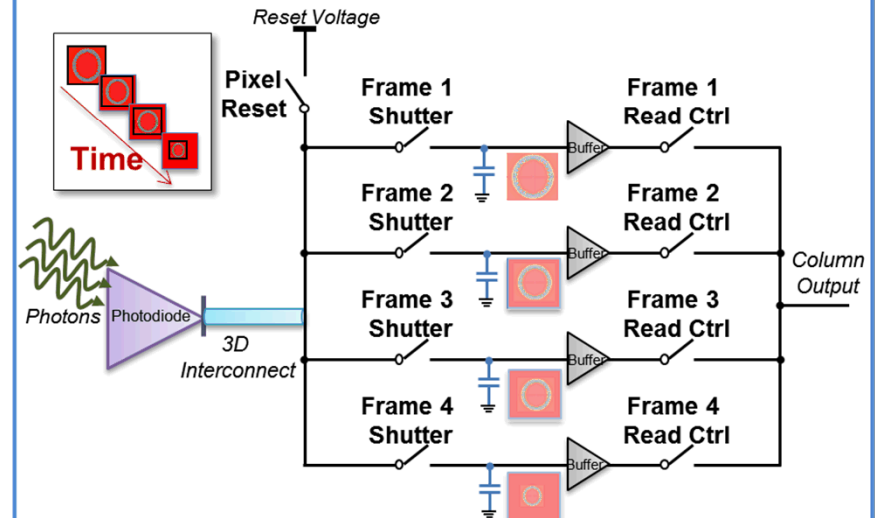
Dual Crystal Spherical Spectrometer w/ 1-D Resolution



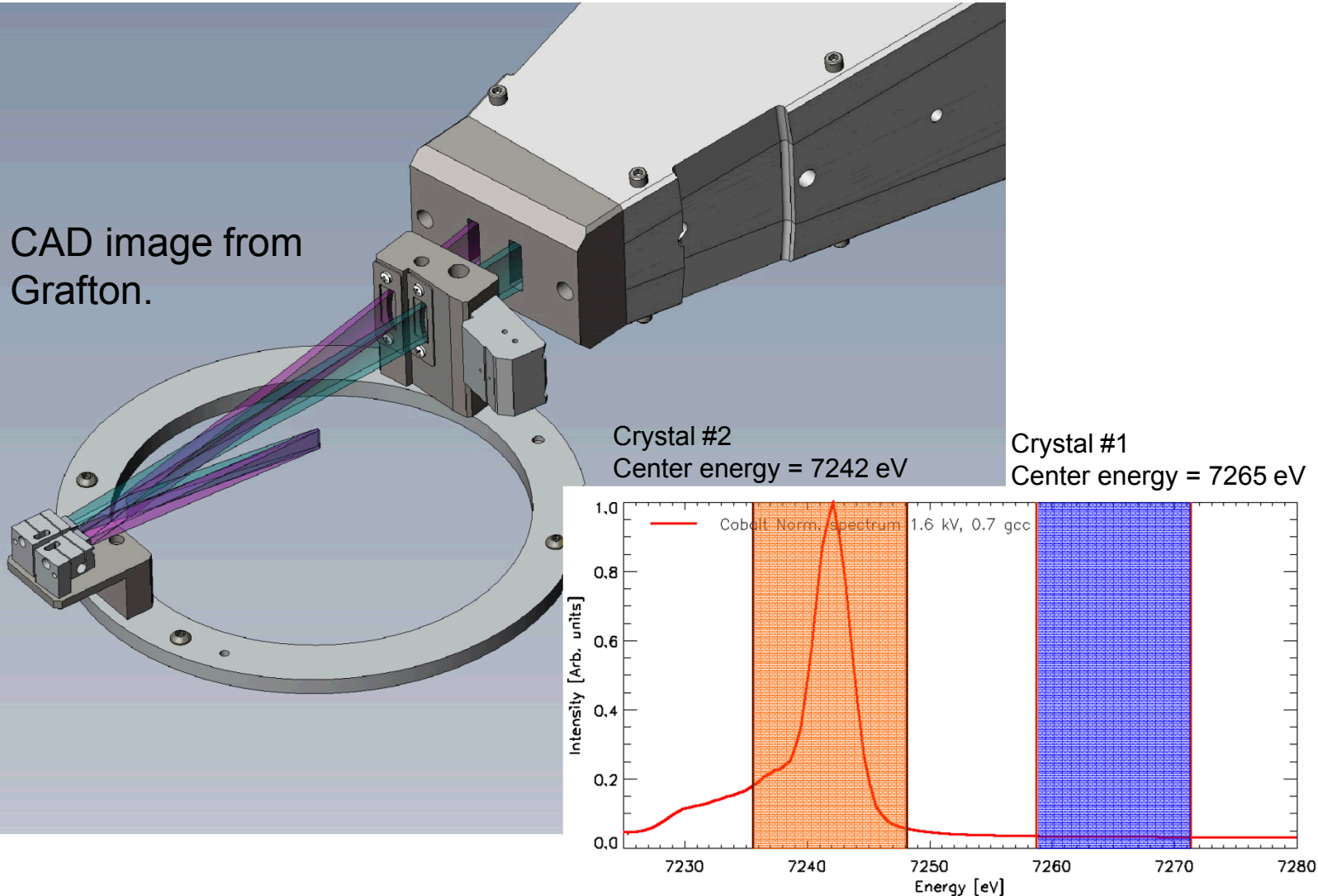
Next Generation Sensor (FY18)

- 3 – 9 frames per sensor depending on resolution requirements
- ~1 ns gate times
- Good dynamic range up to ~10 keV

Each Of The 448 x 1024 Pixels Has This Four Sample, Hold & Read-Out Circuit

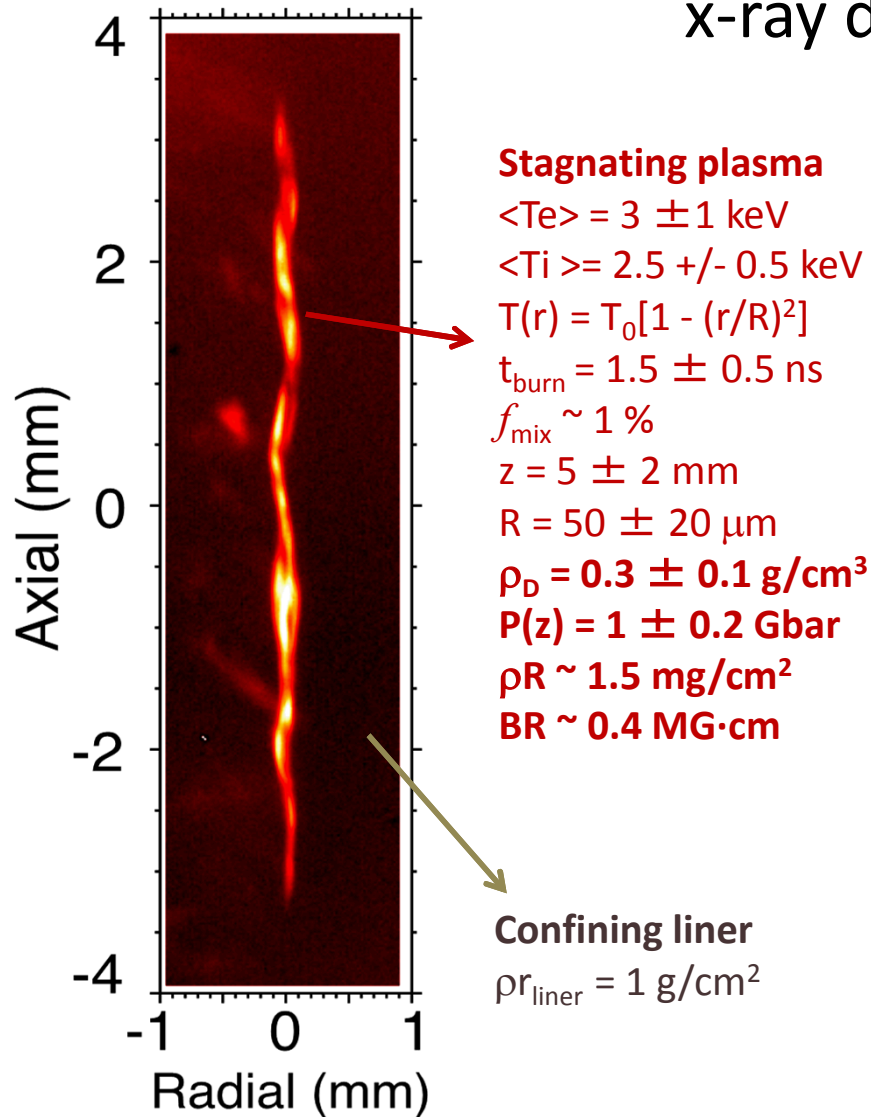


The new Dual crystal imager will capture two images of stagnation by using two side-by-side Ge 335 crystals.



Backup

MagLIF stagnation is diagnosed with extensive neutron & x-ray diagnostics



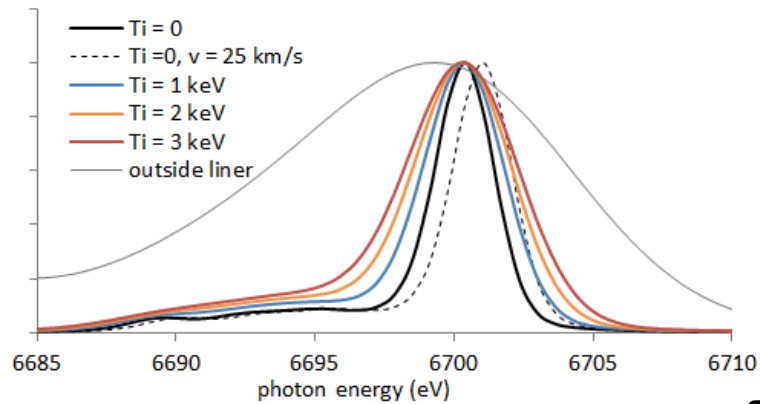
X-ray image of MagLIF plasma

Magnetized Liner Inertial Fusion (MagLIF) has the potential to produce high fusion yields by exploiting:

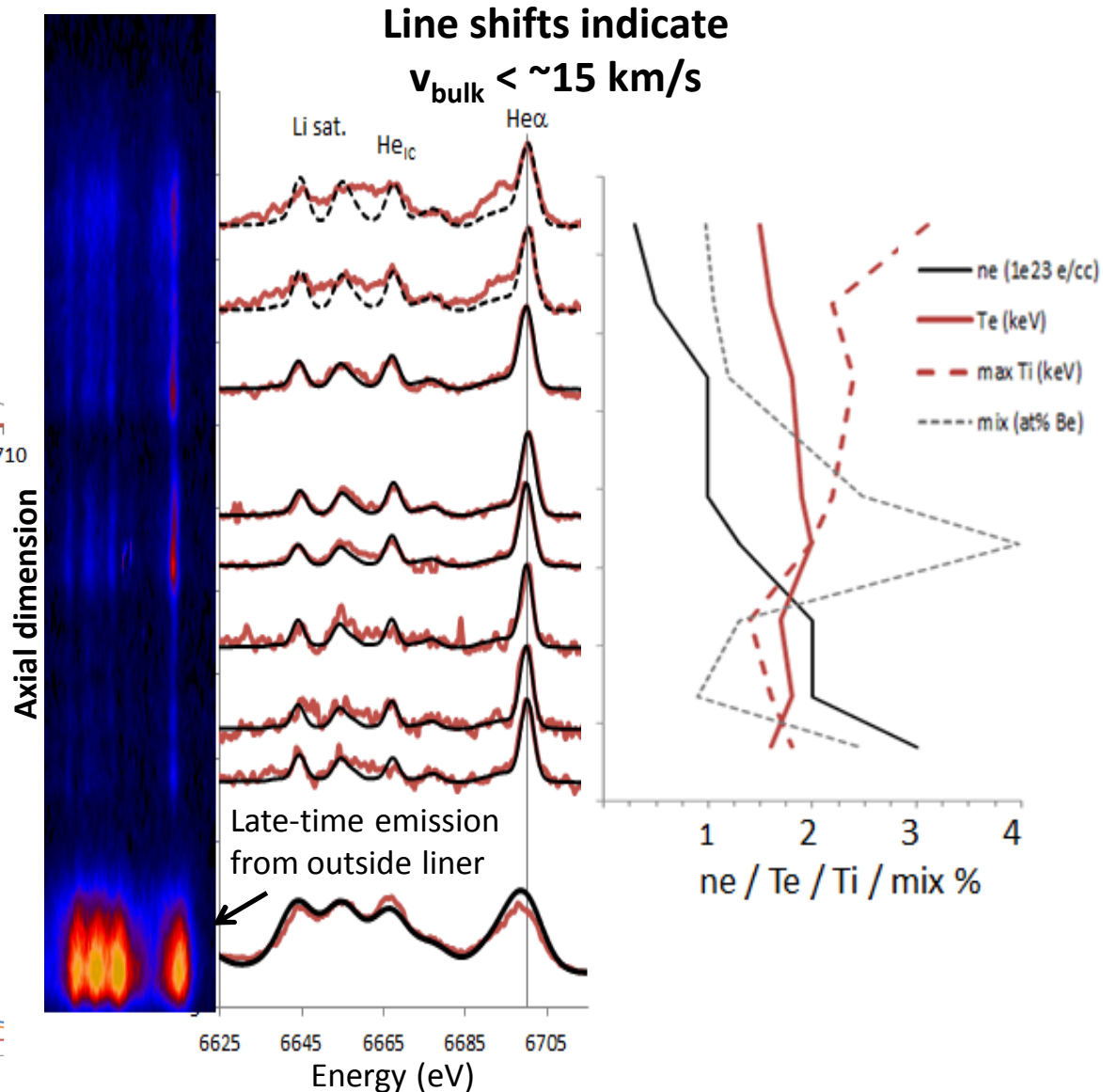
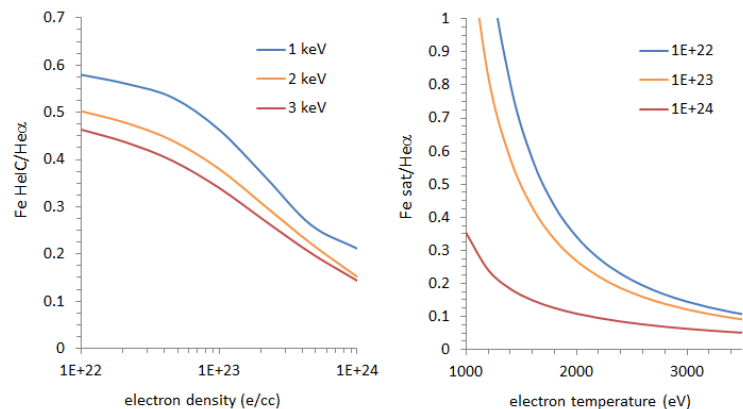
- 1) magnetic confinement that relaxes required pressures 100x (present experiments trap ~40% of fast fusion products)
- 2) a highly efficient driver delivering ~1% of its stored energy *to the fuel*
- 3) Symmetric drive and slow, low-convergence implosions that are robust against instabilities
- 4) Preheat and stagnation stages are both highly sensitive to radiative losses from impurities (mix)

High-resolution, axially resolving spectrometer provides a wealth of information

Line shapes $\rightarrow T_{\text{ion}}, v_{\text{bulk}}, \& r_{\text{source}}$

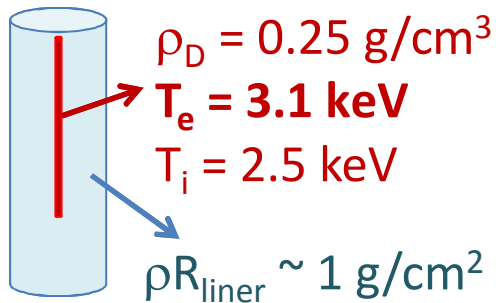


Intensity ratios $\rightarrow T_e, n_e, \& \text{mix}$



Data constrains some gradients as well as burn averages

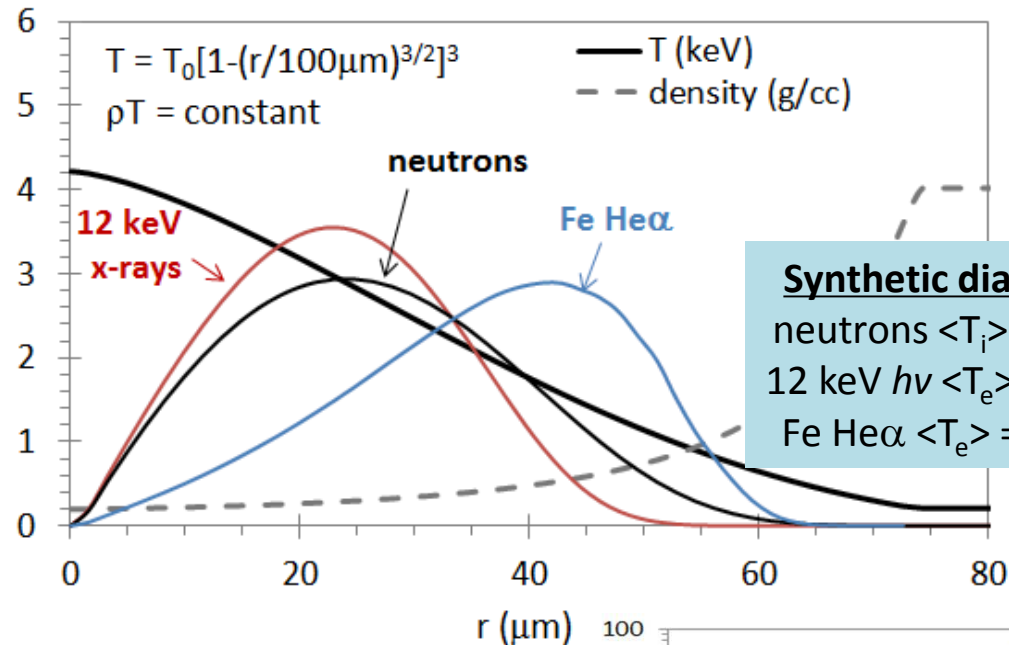
“cartoon” stagnation distilled
from x-ray images & continuur
spectra



**Inferences from
high-res line spectra**

$\rho_D = 0.33 \text{ g/cm}^3$
 $T_e = 1.8 \text{ keV}$
 $T_i = 2.3 \text{ keV}$
 $\rho R_{\text{liner}} = 1.1 \text{ g/cm}^2$

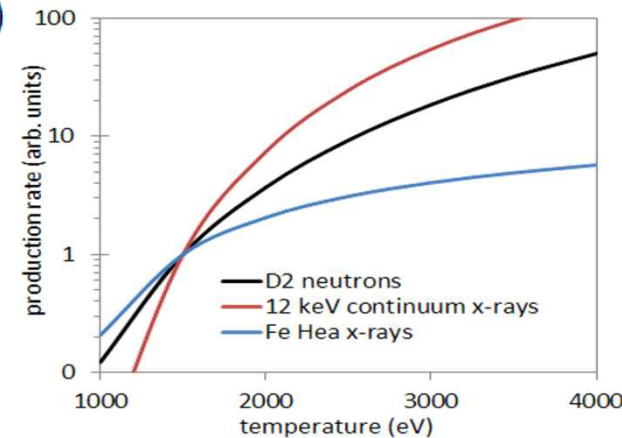
Isobaric model* fits diagnosed Ti/Te



Synthetic diagnostics

neutrons $\langle T_i \rangle = 2.5 \text{ keV}$
 12 keV $h\nu \langle T_e \rangle = 3.1 \text{ keV}$
 Fe Heα $\langle T_e \rangle = 1.7 \text{ keV}$

High-energy (12-15 keV) continuum
x-rays sample hotter portions of
the plasma than Fe line emission
(and even D-D neutrons)



*Springer et al., *EPJ Web of Conferences* **59**, 04001 (2013), and S. Hansen et al., *Phys Plasmas* (2014).

Stagnation data can constrain preheat energy and mix

- Neutron yields are highly sensitive to initial preheat energy (Slutz, Sefkow, Peterson, McBride) and radiation losses from mix (Slutz/McBride)
- Spectroscopy constrains mix (Be from Fe on XRS3; Al from Zn on CRITR)

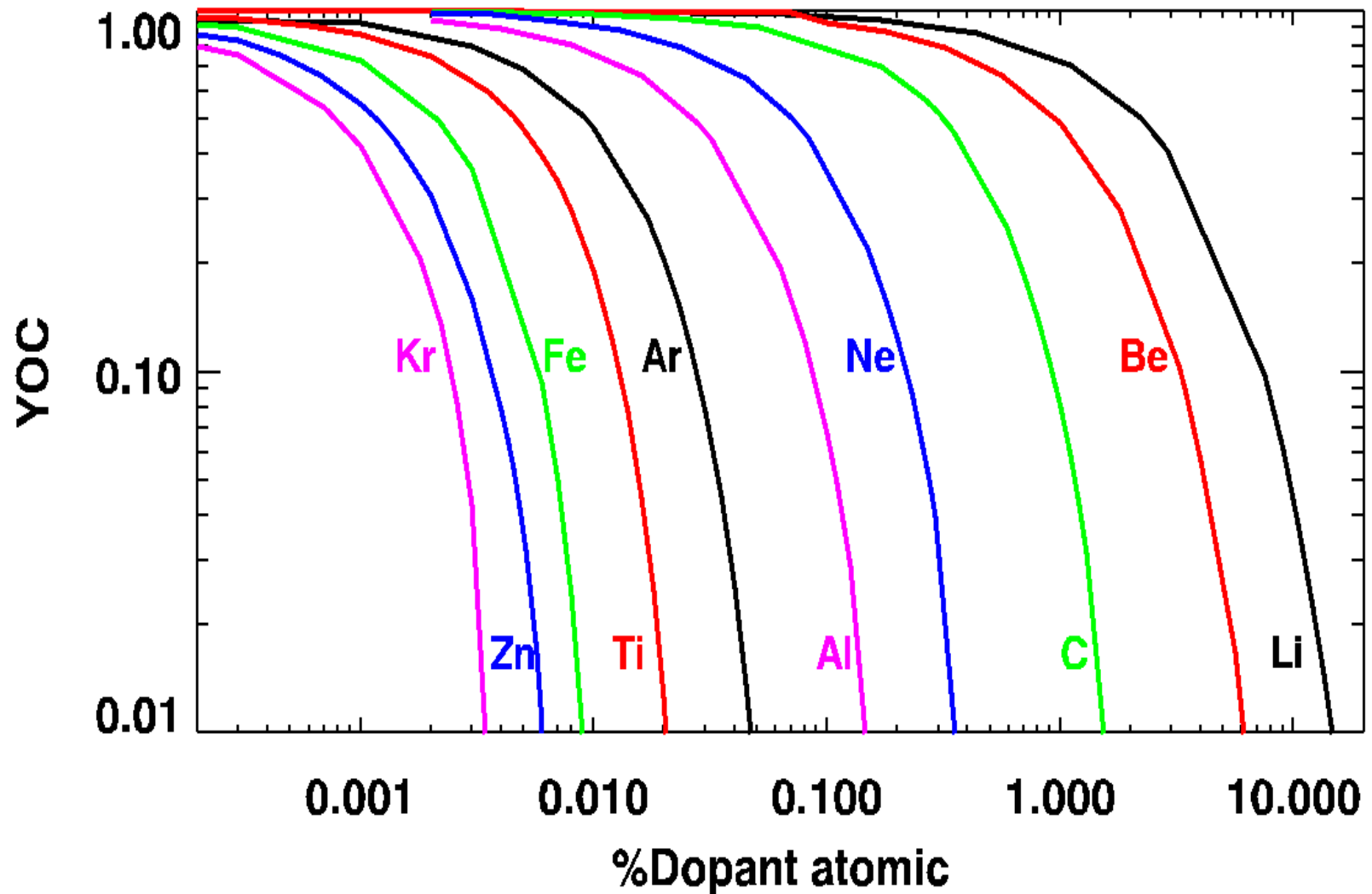
Window thickness (μm)	Preheat energy (J)	Endcap mix (%)	Window mix (%)	Liner mix (%)
3.5	150	0.05	0%	1%
3.5	300	0.05	5%	1%
1.5	150	0.10	0%	1%
1.5	300	0.15	5%	1%

Diagram illustrating the relationship between parameters and their constraints:

- Preheat energy (J) vs. Endcap mix (%):** A blue arrow labeled "increase" points from 150 J to 300 J, and an orange arrow labeled "Same?" points from 0.05% to 0.10%.
- Endcap mix (%) vs. Window mix (%):** A blue arrow labeled "increase" points from 0.05% to 0.10%, and an orange arrow labeled "Same?" points from 0% to 5%.
- Window mix (%) vs. Liner mix (%):** A blue arrow labeled "shock -> brems?" points from 0% to 5%, and an orange arrow labeled "Same?" points from 1% to 1%.

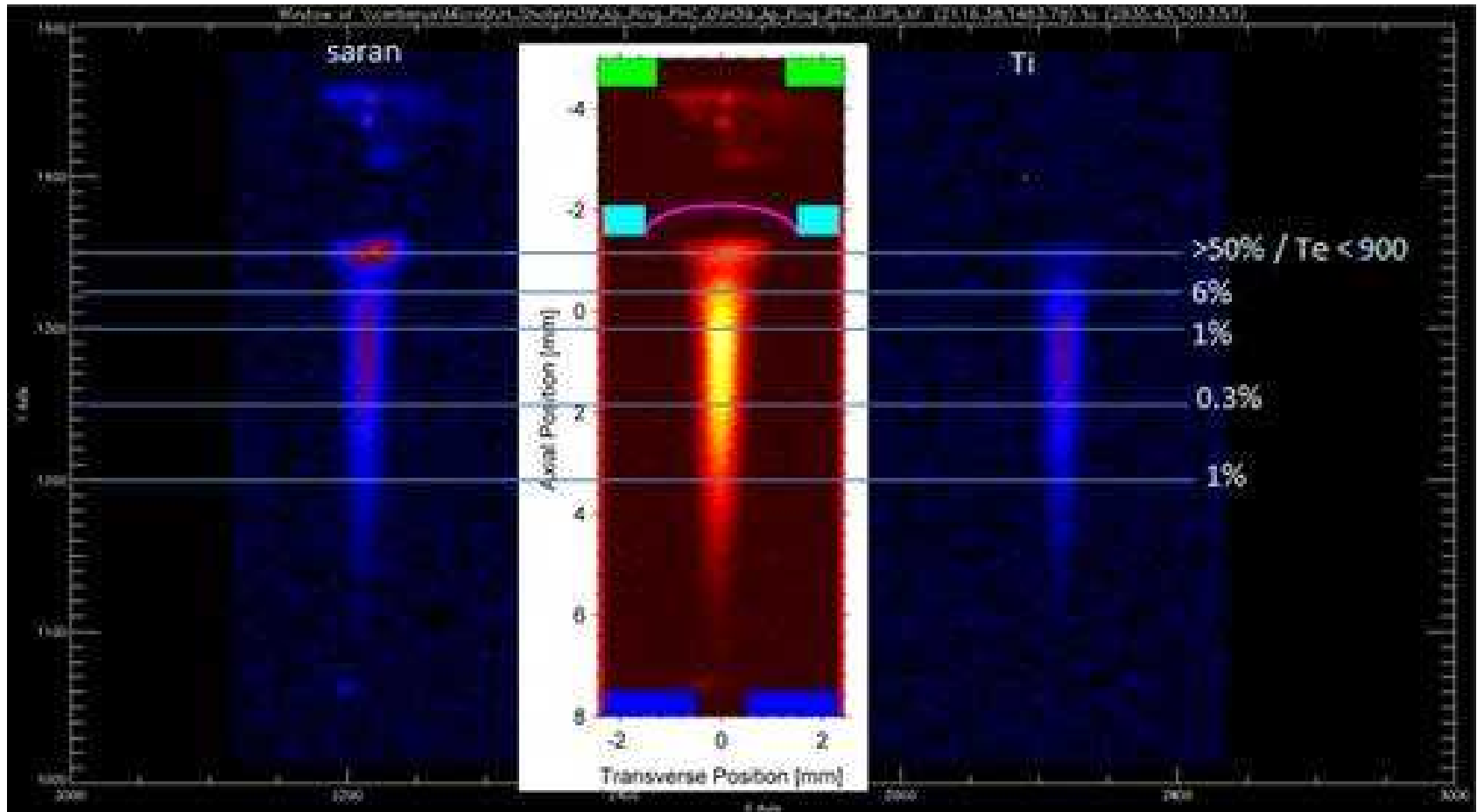
- Fitting yields from a very simple model to data from shots with Al & Be endcaps (and unconditioned beams) constrains preheat energy and mix
- Assumes that for a given laser energy and window thickness, the preheat energies and mix fractions must be the same regardless of endcap material
- Endcap mix fractions consistent with measured yields increase with laser energy under several plausible scenarios

Yield degradation from mix (LTE?)



Steve Slutz

Evidence for window mix from laser-only shot (H39)



- The ratio of TIPC intensities behind saran/Ti filters is sensitive to carbon fraction (which increases continuum over Ar line emission)

How can we measure window mix?

- Characteristic X-rays may be the only way we can track material
- Flash-coat window/endcaps? (Harding)
- Late-time measurements seem possible even with 160 J of preheat energy; earlier time require dedicated targets
- With 500 J of preheat, temperatures support high-energy emission that could escpar liner

