

# Mix in MagLIF Stagnation

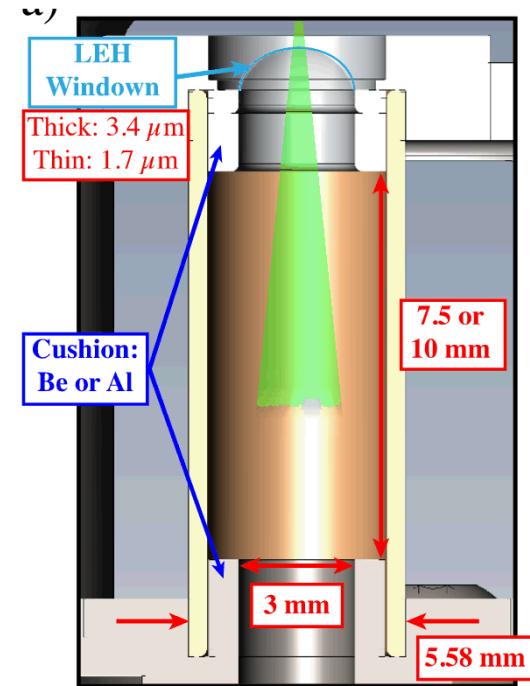
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# Questions to Address

- What do we know today and how do we know it (what measurements, model evidence)?
- How does our measurement knowledge differ from our models?
- What are the big questions we need to address?
- What experiments and/or simulations can we conduct to address these questions?
- What new measurements, diagnostics, and/or models are needed?

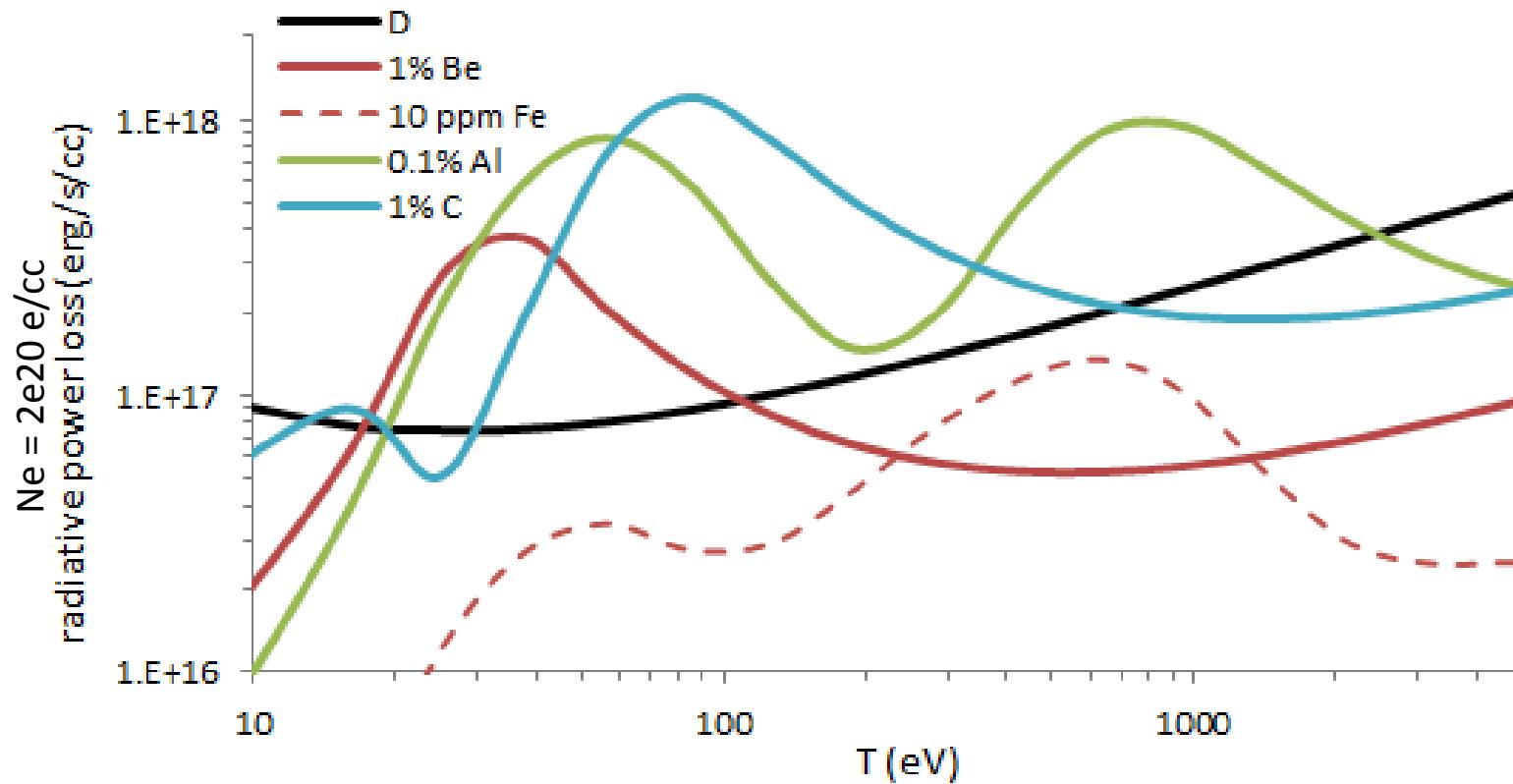
# What do we know today and how do we know it (what measurements, model evidence)?

- Mix from the endcaps affects the yield
  - Be endcaps result in a  $\sim 10x$  higher yield than Al endcaps
- Be from the liner and/or endcaps end up in the hot stagnation plasma.
  - Fe He-like emission is seen in time-integrated spectra
- The intensity ratio of the Fe He-like lines to the continuum is consistent with a few % Be mix.
  - Assumes a lot...
- Early-time mix can significantly affect stag performance
  - 0.1% Ar in the DD duds the yield.
- Mix from the laser-heating phase is probably important
  - Trend of decreasing yield with increasing laser energy
- The LEH window may be a source of mix.
  - Shot with window on bottom of cushion had  $\sim 30x$  less yield
  - Shot using phase-plate had  $\sim 20x$  less yield



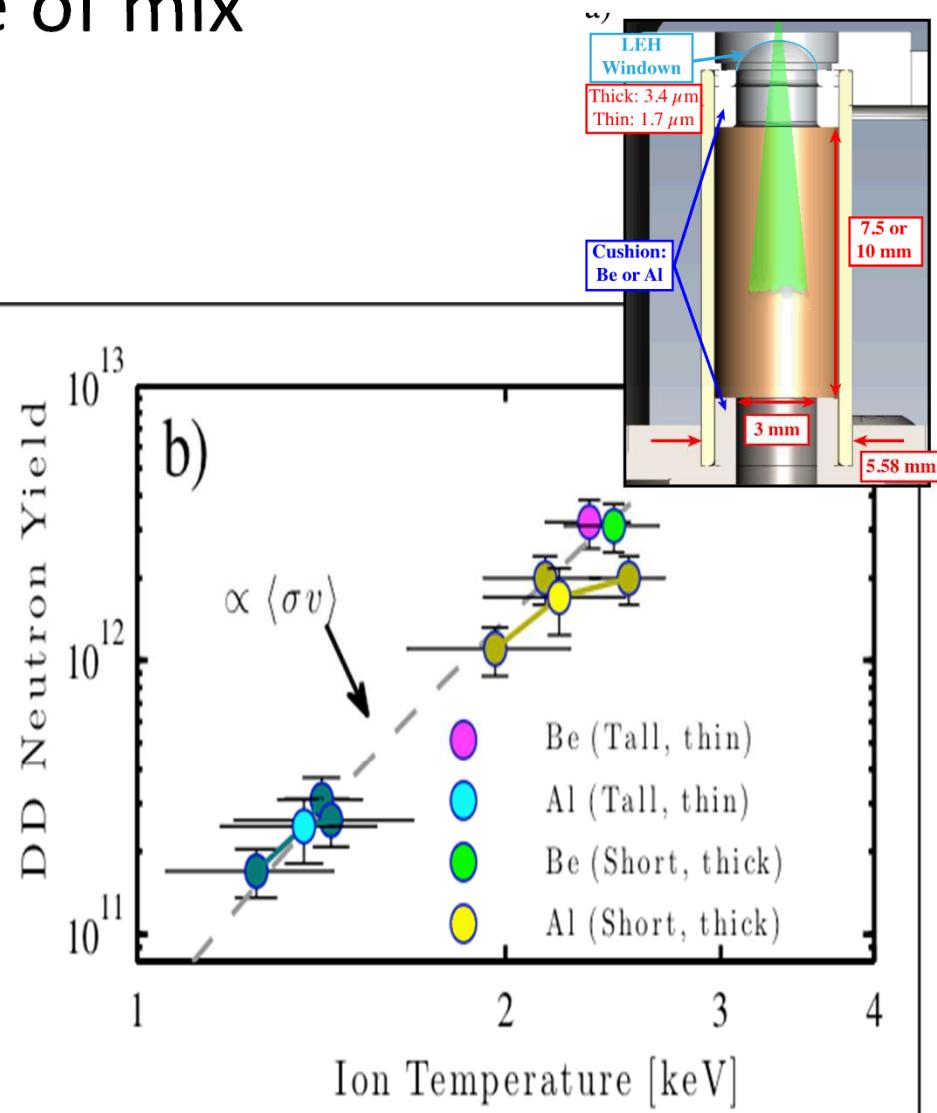
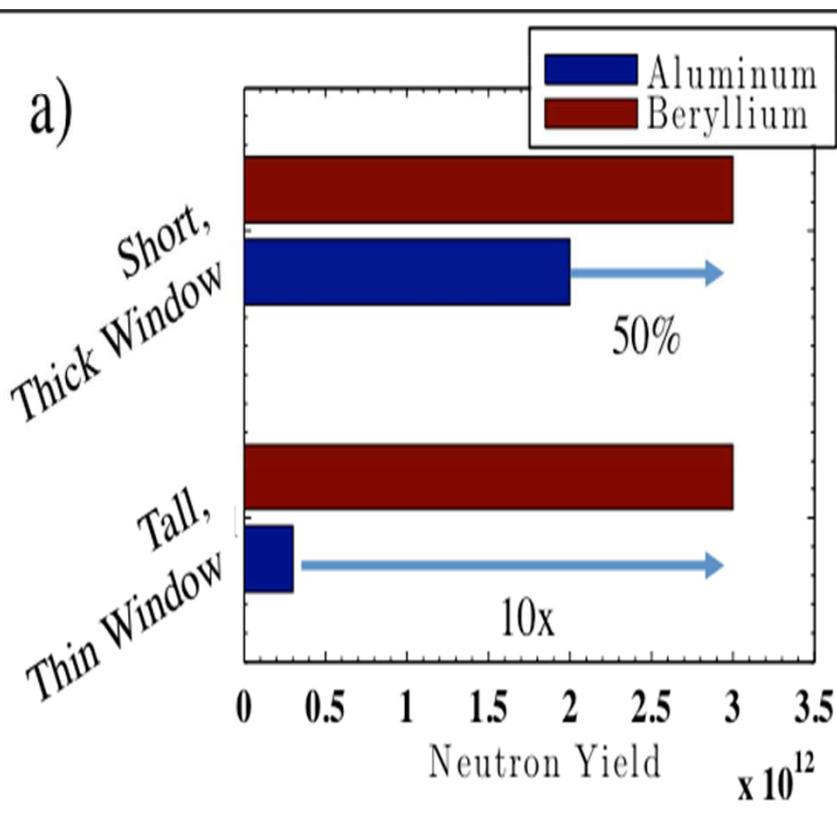
In Activation  
-DD Yield  
Crystal Spectrometers  
-Lines and Continuum  
Crystal Imager  
-Hot plasma distribution

# Radiative losses vary with material and charge state

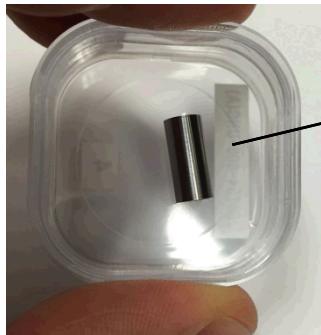


- Mix levels of  $\sim 1\%$  Be are tolerable ( $\sim 2\times$  degradation)
- 0.1% Aluminum is devastating ( $\sim 10\times$  degradation)
- 1% carbon is devastating ( $\sim 10\times$  degradation)  
Note that uniform mix of half thin (thick) window mass gives  $\sim 3\%$  (8%) C
- Mid-Z impurities in Be don't hurt at all, even at 10x detected levels

# Yield and Ti are consistent with endcaps being a source of mix

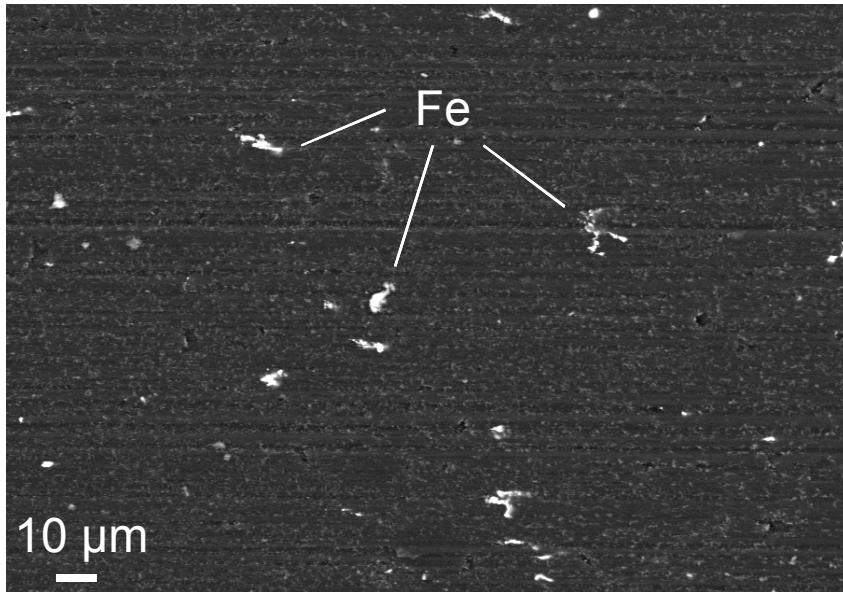


Fe occurs “naturally” in our S-65 grade Be liners and endcaps.



Machined  
Be liner

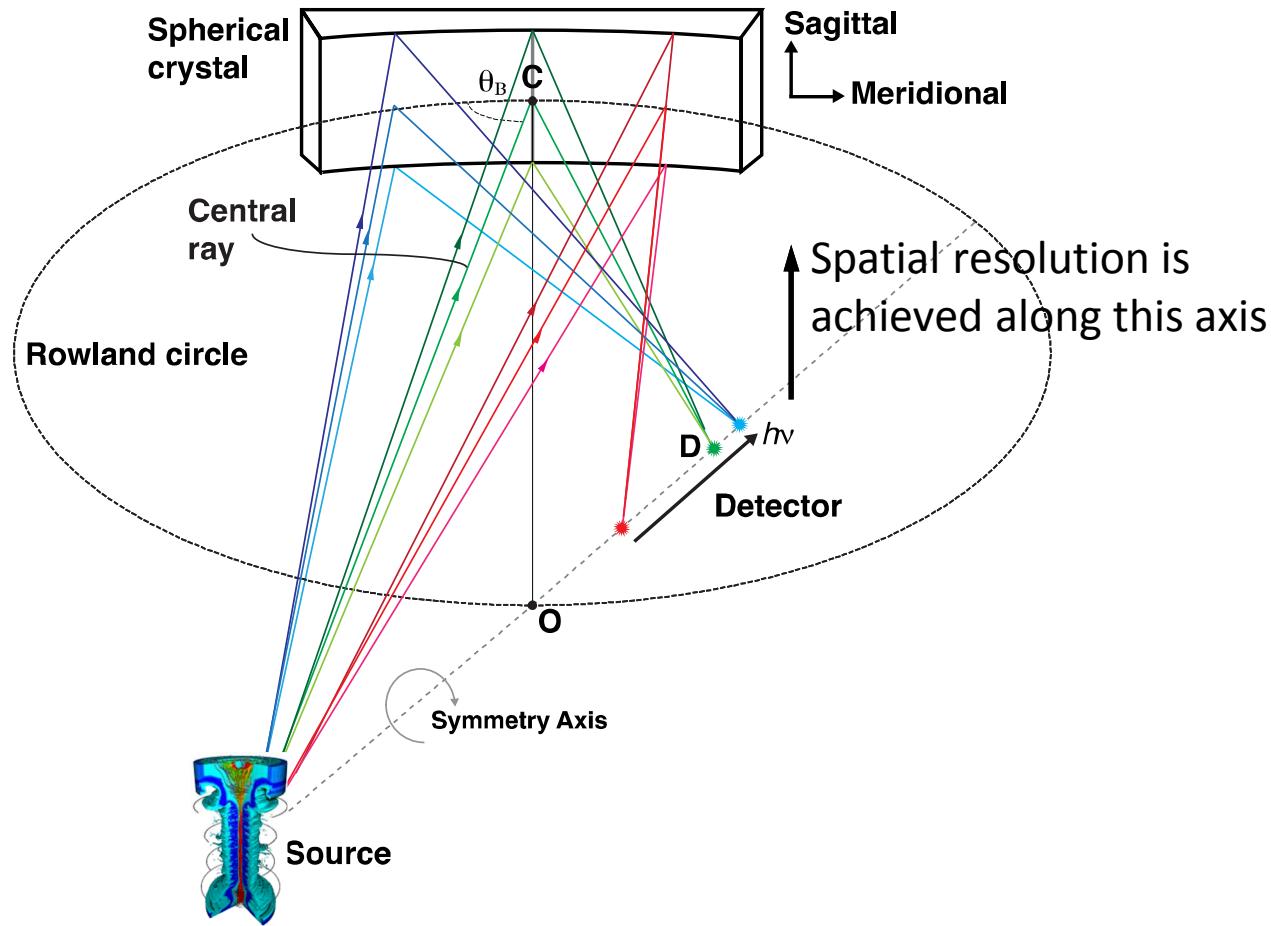
SEM image of the Be liner outer surface



- The Be stock material contains  $\sim$ 100 ppm (0.01% atomic) of Fe that originate from the manufacturing process.
- The Fe appears as micron-sized impurities that are uniformly distributed on the visible surfaces. We are assuming it is uniform in the bulk.

To resolve the Fe emission generated at stagnation we use a spherically-bent crystal spectrometer.

Spherical crystal spectrometer<sup>1</sup>



<sup>1</sup>E.C.Harding et. al., RSI (2015)

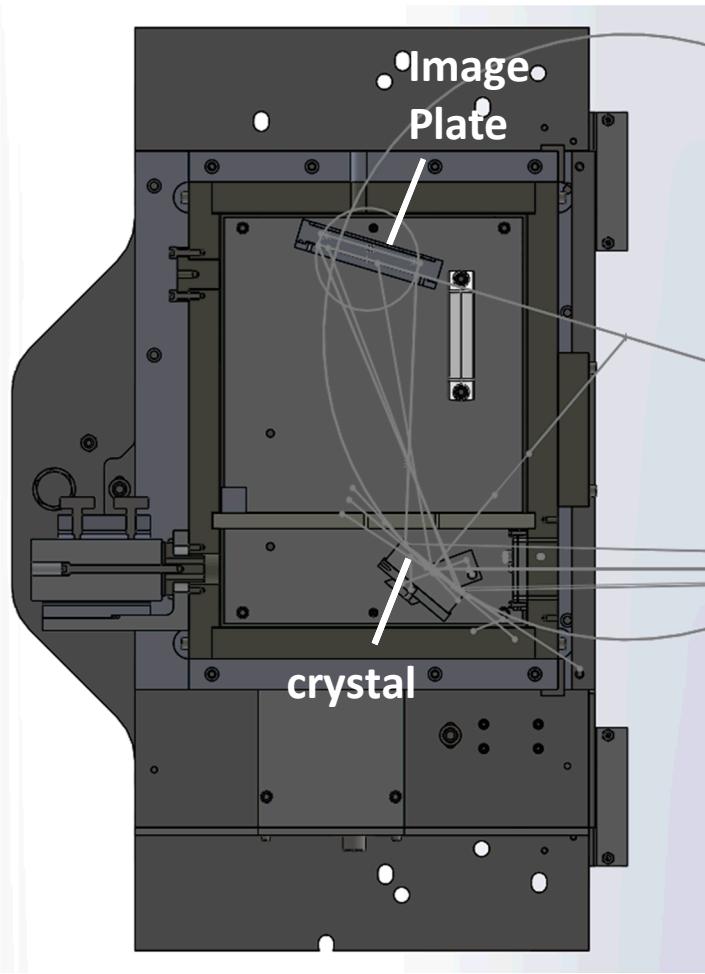
D. Sinars et. al. JSQRT (2006)

FSSR used on dynamic hohlraum capsule implosions

# X-Ray Scattering Spherical Spectrometer (XRS3) Spectrometer

## Spectrometer setup for He-like Fe emission

Crystal	Q20-23 (2d = 2.749 Å)
Source-to-crystal	800 mm
Crystal-to-detector	256.92 mm
Crystal Radius	250 mm
Center Bragg Angle	40°
Crystal size <sup>1</sup>	60 x 36 mm
<b>Spectral Range<sup>2</sup></b>	6328 - 7977 eV
<b>Spatial Mag. (M<sub>sag</sub>)</b>	0.30x
<b>Spectral Resolution<sup>3</sup></b>	2 eV
<b>Spatial Resolution<sup>3</sup></b>	210 μm
Throughput	1.9e-7 steradians

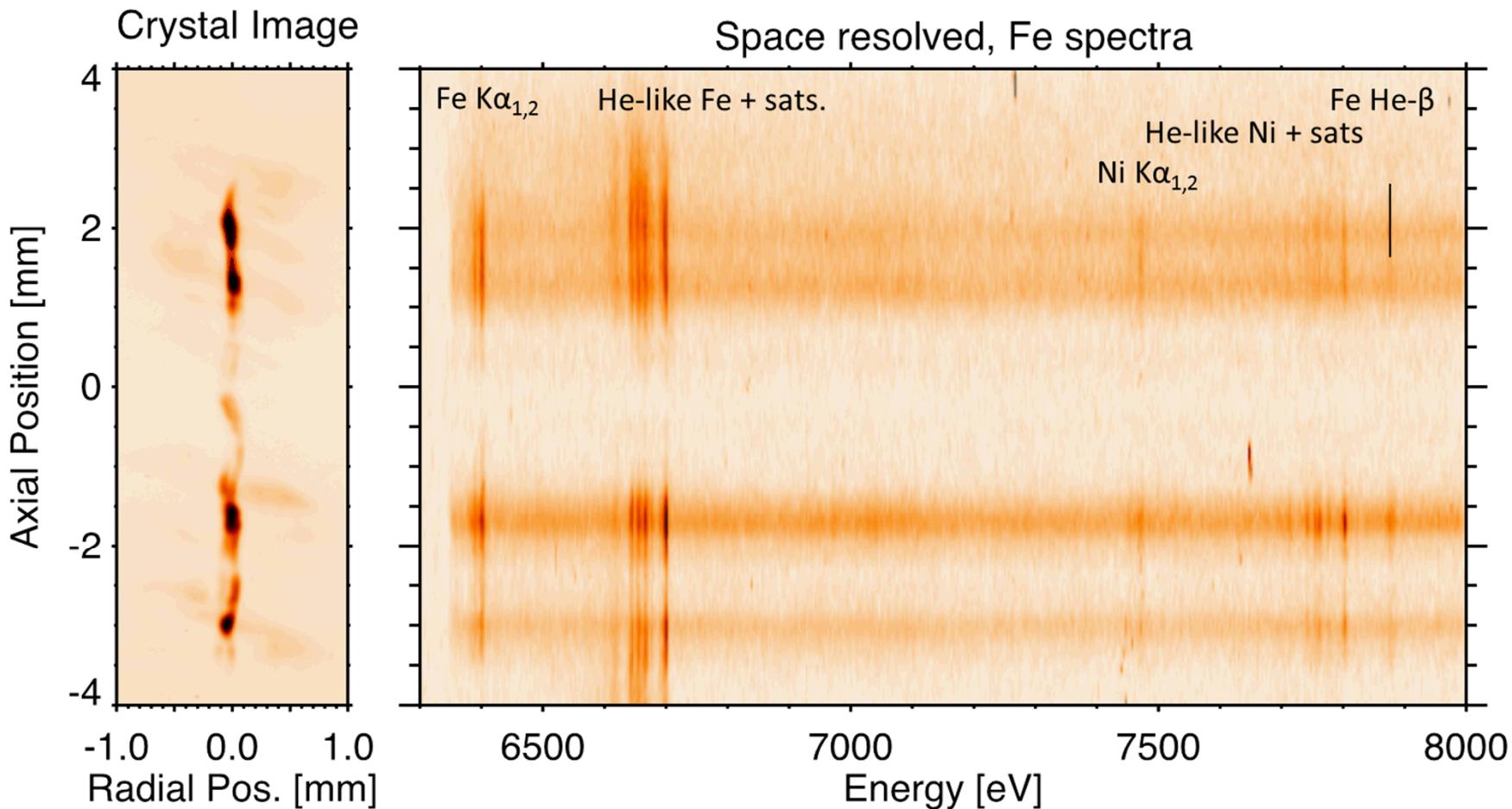


<sup>1</sup>This is a tiled crystal consisting of 2 strips, each one is 60 x 18 mm

<sup>2</sup>Detector length must be 85 mm to capture entire spectral range.

<sup>3</sup>Limited by the Image Plate resolution of 63 microns.

We believe we are observing He-like Fe emission from stagnation.



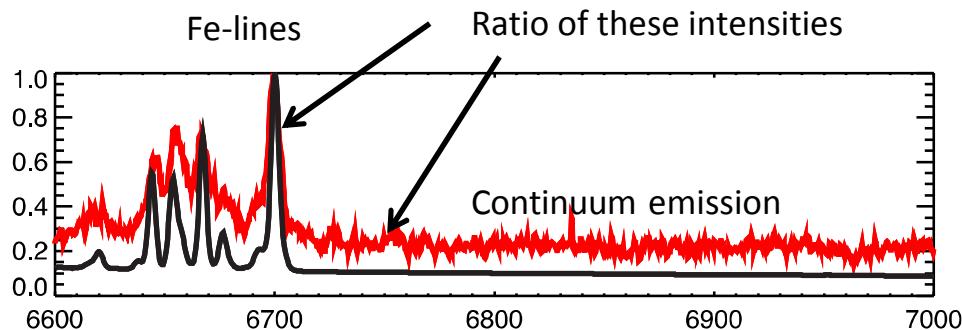
Z2850

**Target:** 7.5 mm tall Be AR6 liner with Be cushions and a 3.5 um LEH window

**Yield:** 3E12 DD (highest performing MagLIF shot to date)

# Ratio of Fe line emission and nearby continuum emission is sensitive to mix fraction.

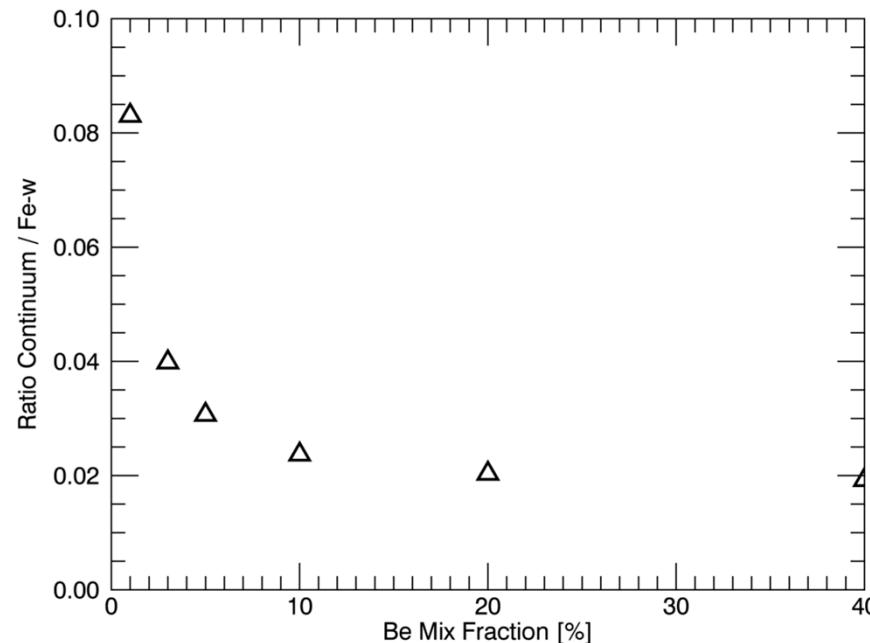
## XRS3 data and PrismSPECT



### Key Assumptions:

- Lines and continuum are from the same plasma at the same time
  - If continuum has other contributions, then this is a lower bound on mix
- Fe/Be ratio is fixed throughout the volume
- Others?

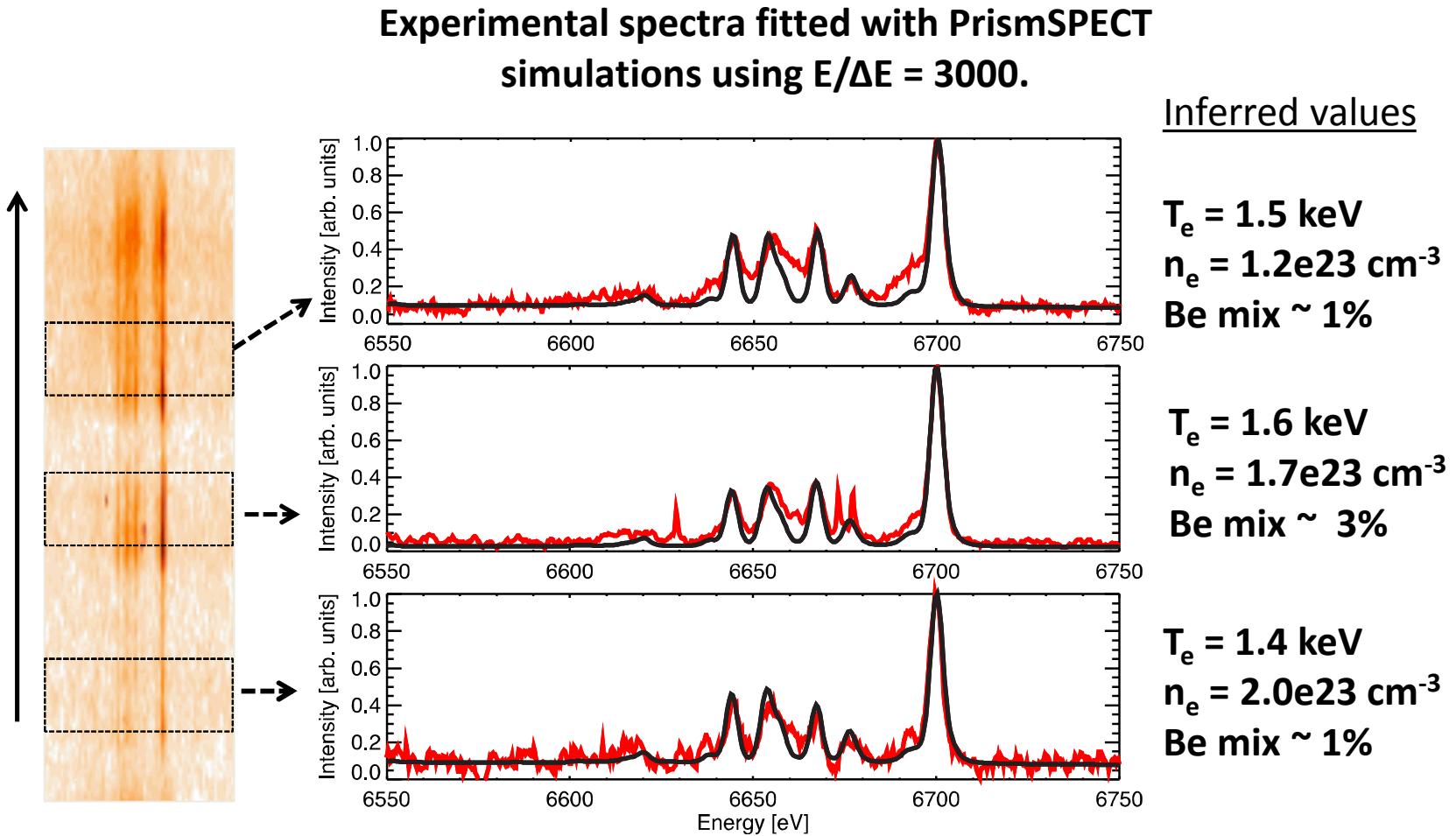
## Intensities Ratios vs. Be mix



PrismSPECT runs use  
 $D + Be + Fe$

Relative amounts of each are varied  
w/ fixed total mass density and  $T_e$ .

# Example: (Z2839) Be mix fraction appears to be 1-3%



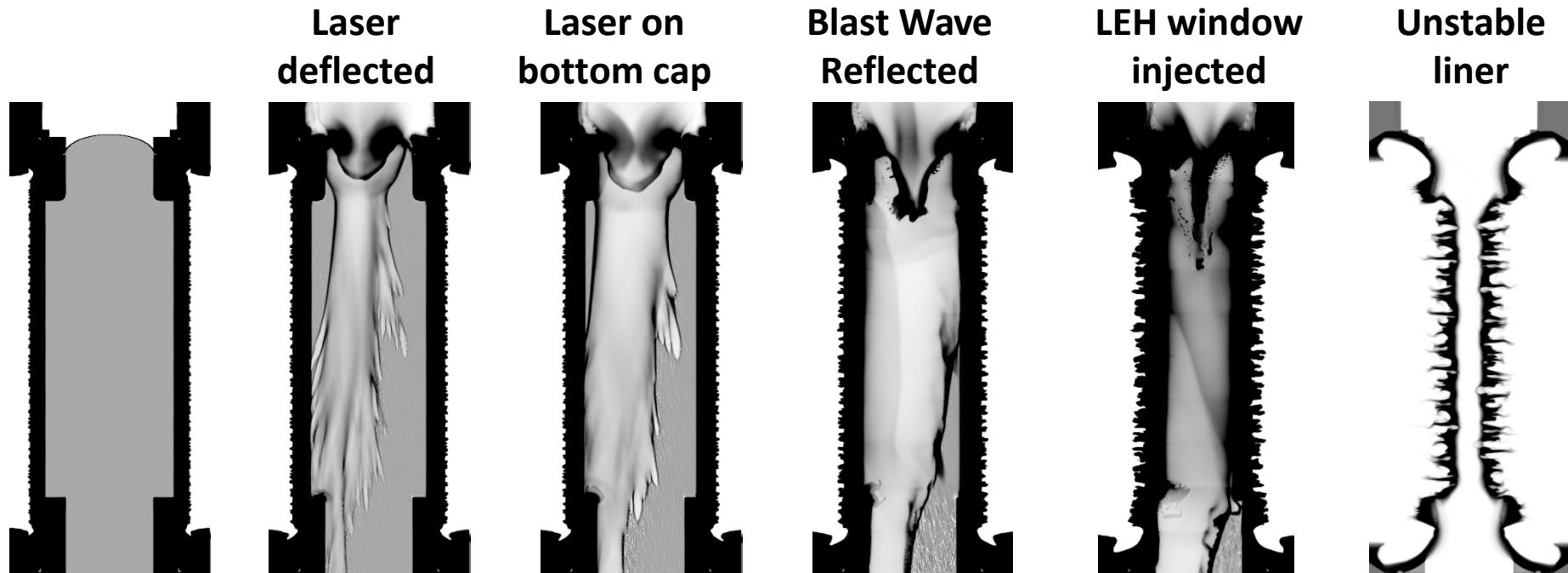
Z2839

**Target:** 10 mm tall Be AR6 liner with Be cushions and a 1.5  $\mu\text{m}$  LEH window

**Yield:**  $3 \times 10^{12}$  DD (highest performing MagLIF shot to date)

# How does our measurement knowledge differ from our models?

- Models indicate mix can occur from multiple origins:
  - Blast wave from laser preheat causes blowoff from liner wall and endcaps
  - Laser can pass through the gas and cause blowoff from the bottom end cap
  - Laser can deflect through LEH plasma and hit the liner/endcap causing blowoff
  - The exploded LEH window can mix into the gas
  - The liner is RT unstable

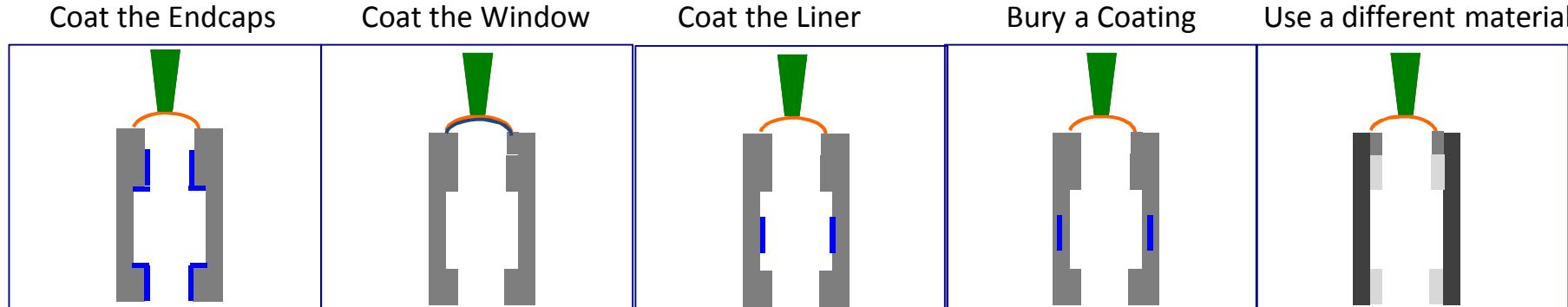


# What are the big questions we need to address?

- What are the true origins of the Be mix and at what level of relative importance?
  - Present data is ambiguous on the source and simulations indicate it may come from multiple places.
- How is the Be mix distributed in space and time?
  - Present interpretation is that Fe emission all comes during stagnation
  - Spectrometers integrate over at least 2 spatial dimensions
- Is the LEH window a significant contributor to mix?
- How can we reduce the amount of mix from preheat where it has the largest impact?

# What experiments and/or simulations can we conduct to address these questions?

- Use localized coatings/tracers to assess the origins of the mix.
  - Liner (Co), endcap (Co), window (Co, Cl)
  - This is our baseline approach; ~1 week per year of dedicated development on Z
  - Should use OMEGA-EP and Pecos (ZBL) to develop techniques for window coatings
- Use different liner/endcap materials and assess impact on yield and/or intensity of line emission.
  - Li (target development in progress), high purity Be
- Assess mix with different laser preheating.
- Assess yield with various LEH window positions.
  - A coupled problem

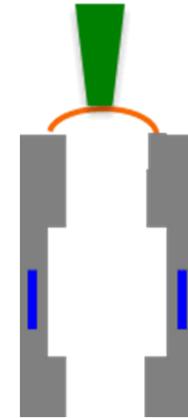


# Future experiments and diagnostics

- Time-integrated x-ray emission lines from dopant, impurities, and tracers can unambiguously track material fractions at stagnation
  - Co coating on endcaps (and windows) will emit in range of XRS3
  - Ultra-pure and high-impurity Be can help isolate mix source
- In general, x-ray diagnostics must dance along multiple fine lines:
  - AR-6 liner is opaque to  $h\nu < 3$  keV initially (and  $h\nu < 6$  keV at stagnation)  $\rightarrow Z \geq 18$  (26)
  - Detectable emission  $\rightarrow T \geq \sim h\nu/3$  and  $\sim 100$  ppm at preheat ( $\sim$ few ppm at stagnation)
  - Concentrations  $> 100$  ppm of  $Z = 18$  (or  $\sim 10$  ppm of  $Z = 36$ ) degrade neutron yields by 2-3x and decrease (or eliminate) x-ray signals  $\rightarrow$
  - Fuel dopants can be precisely fixed but are restricted to gases (Ar/Kr)
  - There are many good options for endcap/liner tracers (Ti/Co/Zn) but since we lack a predictive simulation capability for dynamic mix, we must rely on empirical data
  - Window tracers can be volumetric (Cl) or coated, but must not interfere with laser energy deposition
- With 500 J of preheat energy and AR12 liners, time-dependent K-shell spectroscopy of 100 ppm Ar in fuel and 1% Cl window could provide temporal evolution of fuel temperature and mix

# What new measurements, diagnostics, targets, and/or models are needed?

- Time-resolved spectroscopy of tracers
  - Requires single line-of-sight gating...
- 2-D spatial imaging of tracers...
- Be liners with buried tracer layers
  - Assess depth of the liner that penetrates stag column
- High purity Be Liners...
- Li liners (in development)
  - Reduce impact of liner mix
  - Oxygen contaminants need to be controlled
- Cryo DD layers?

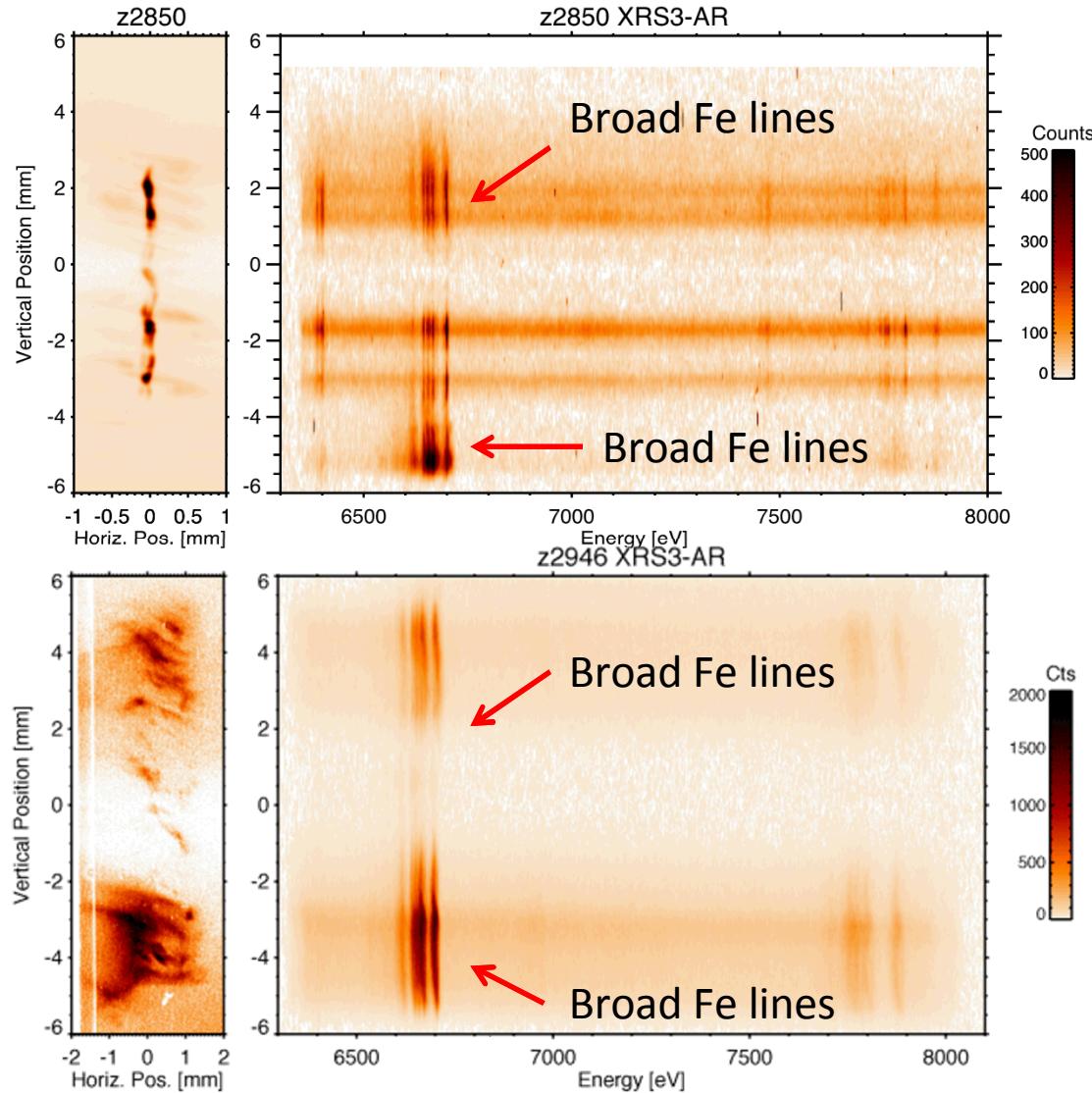


Materion's UHP9999 material is nearly a single crystalline piece of Be. This is the highest purity Be we can get.



Initial discussions with Don Hashiguchi and Jerry Holman (Materion Elmore, OH): UHP 9999 is fragile. Machining will likely require a careful EDM process and possibly diamond turning. EDM only (i.e., no SPDT) may yield Ra ~ 400nm. 1 liner is ~\$8k.

# Fe lines can come from somewhere other than the stagnation column

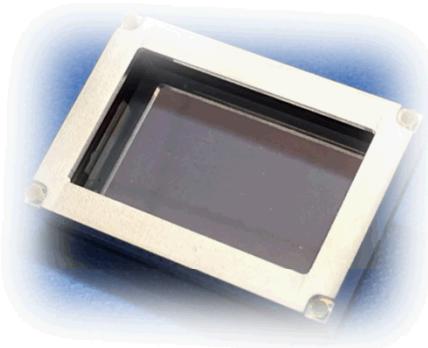
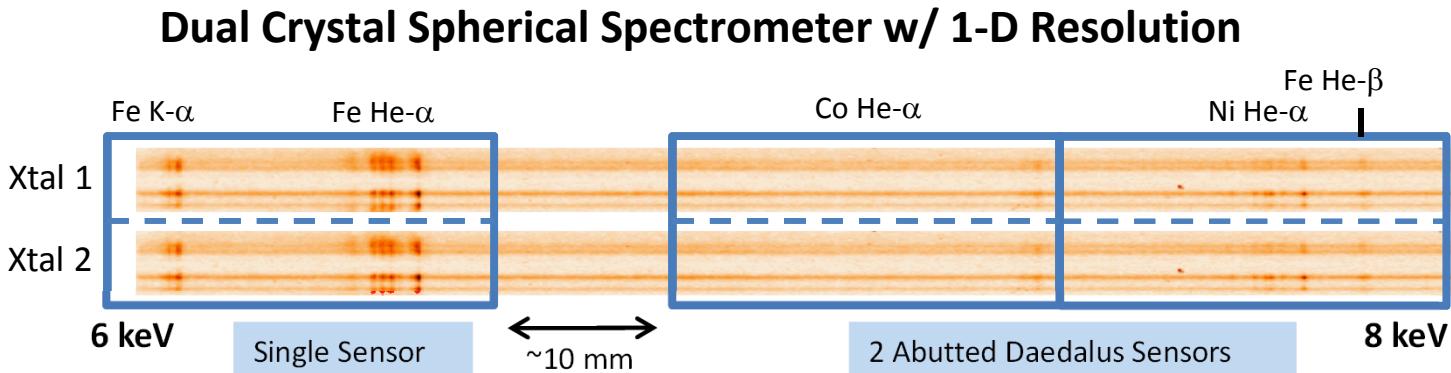
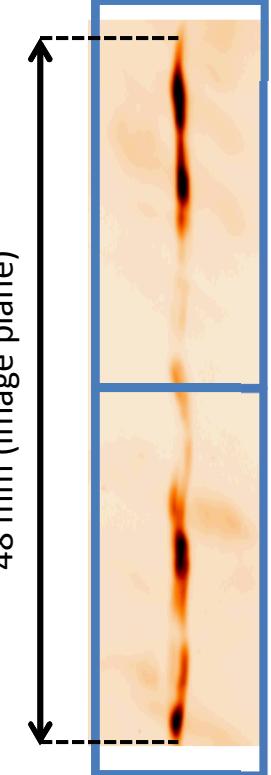


**Full MagLIF shot**  
**(z2850),  $Y_{DD} = 3e12$ :**  
Fe spectral lines near the top and bottom of the target appear broadened.

**Implosion only shot (z2946), no laser heating:**  
No stagnation column but there are strong Fe signals appearing near the ends. This spatially broad source of Fe emission maybe contaminating other shots like z2850.

# Fast gated CMOS sensors will be used to separate stagnation from other events

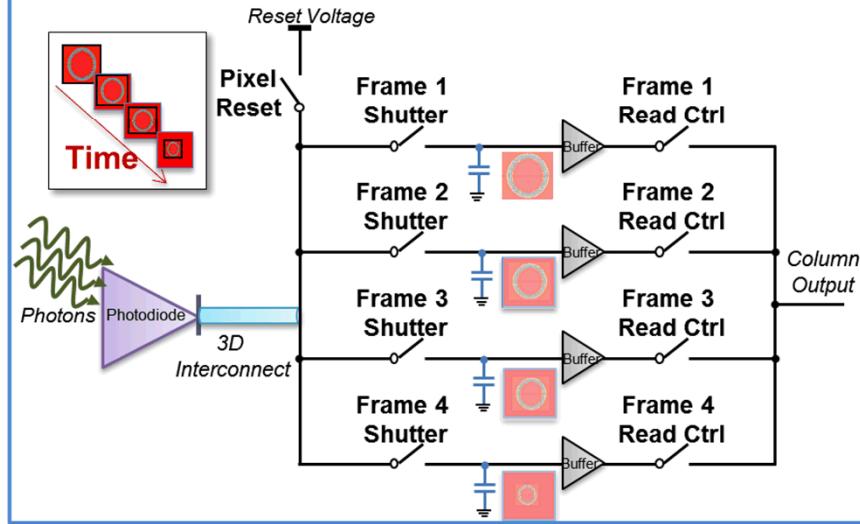
**Spherical Crystal  
2-D Imager  
(Mag 6)**



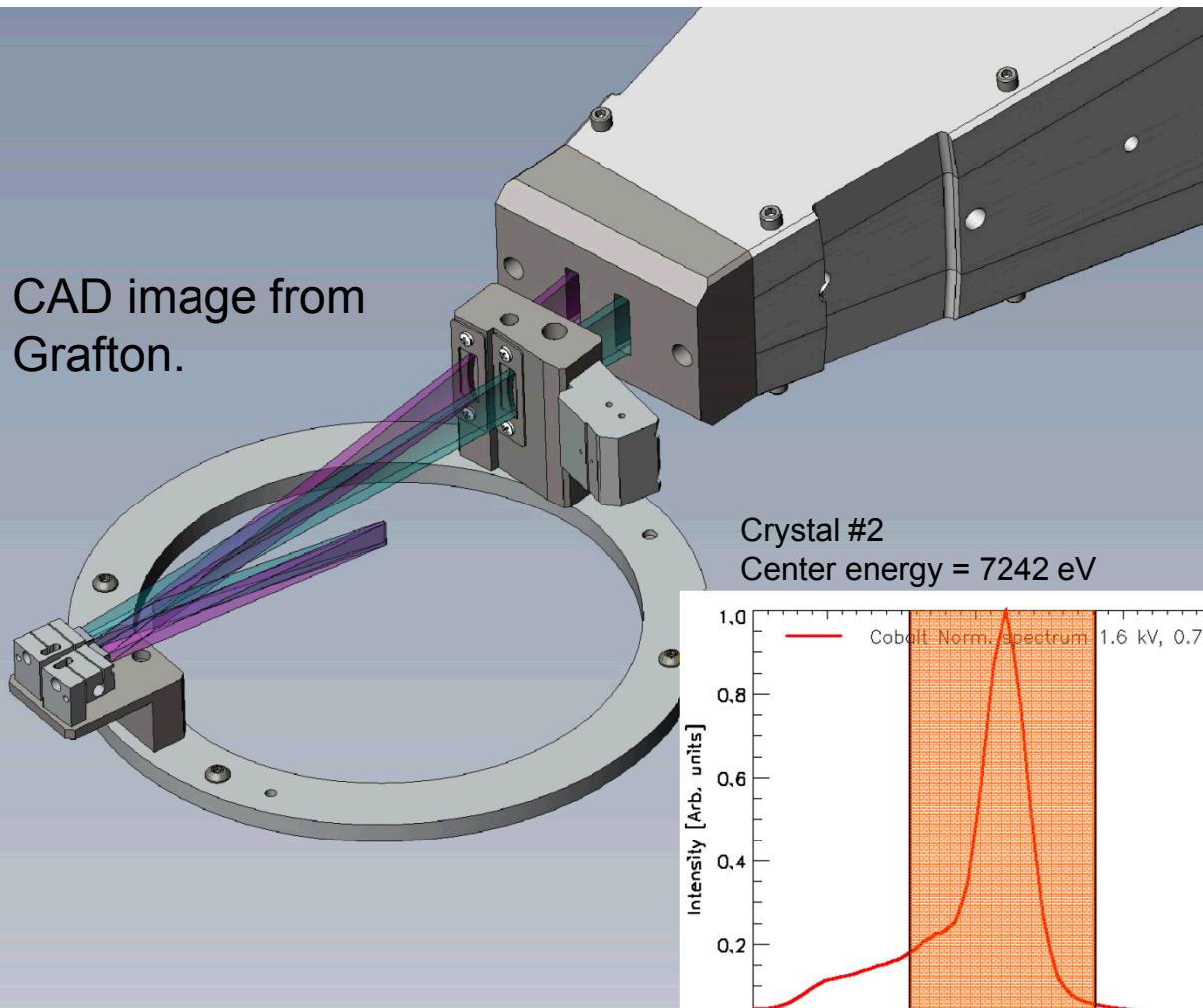
**Next Generation Sensor (FY18)**

- 3 – 9 frames per sensor depending on resolution requirements
- $\sim 1$  ns gate times
- Good dynamic range up to  $\sim 10$  keV

Each Of The 448 x 1024 Pixels Has This Four Sample, Hold & Read-Out Circuit

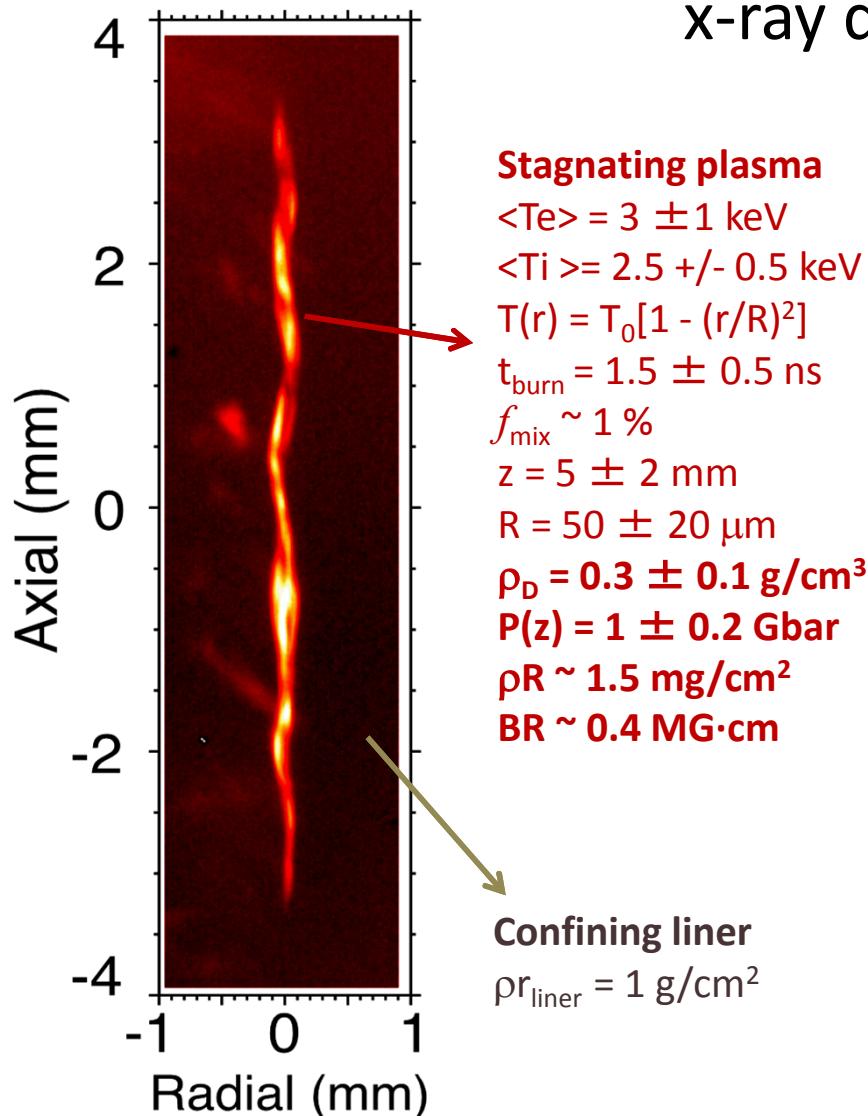


The new Dual crystal imager will capture two images of stagnation by using two side-by-side Ge 335 crystals.



# Backup

# MagLIF stagnation is diagnosed with extensive neutron & x-ray diagnostics



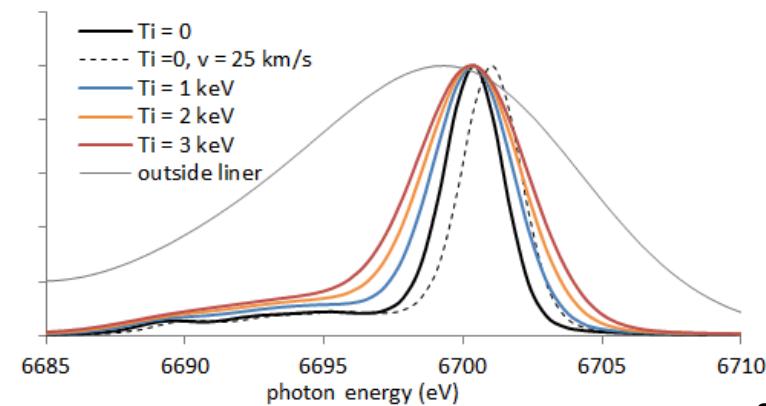
X-ray image of MagLIF plasma

Magnetized Liner Inertial Fusion (MagLIF) has the potential to produce high fusion yields by exploiting:

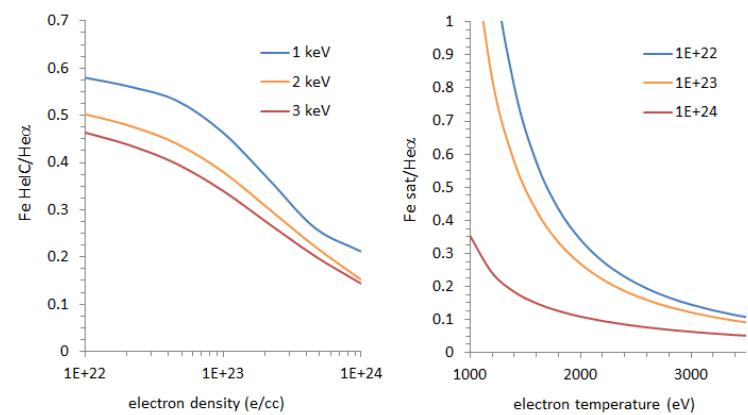
- 1) magnetic confinement that relaxes required pressures 100x (present experiments trap ~40% of fast fusion products)
- 2) a highly efficient driver delivering ~1% of its stored energy *to the fuel*
- 3) Symmetric drive and slow, low-convergence implosions that are robust against instabilities
- 4) Preheat and stagnation stages are both highly sensitive to radiative losses from impurities (mix)

# High-resolution, axially resolving spectrometer provides a wealth of information

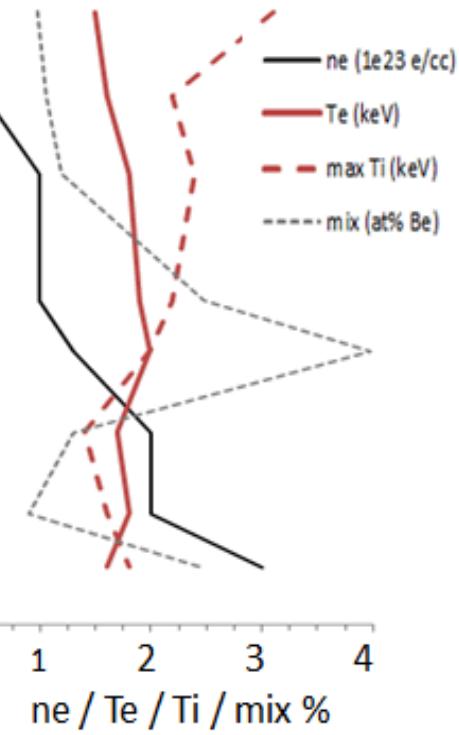
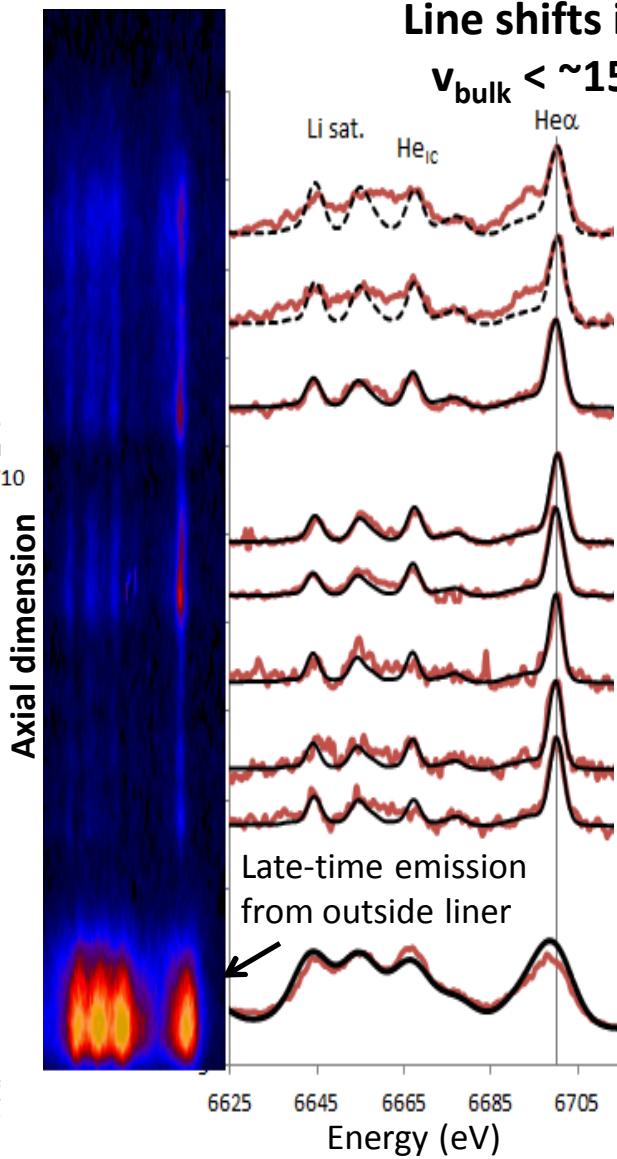
Line shapes  $\rightarrow T_{\text{ion}}$ ,  $v_{\text{bulk}}$ , &  $r_{\text{source}}$



Intensity ratios  $\rightarrow T_{\text{e}}, n_{\text{e}}$ , & mix

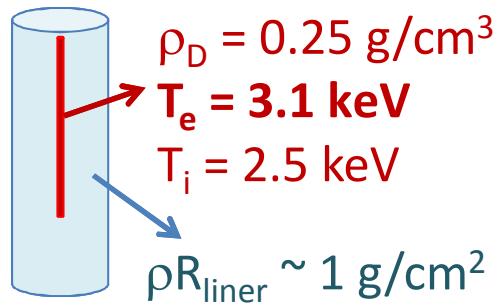


Line shifts indicate  
 $v_{\text{bulk}} < \sim 15 \text{ km/s}$



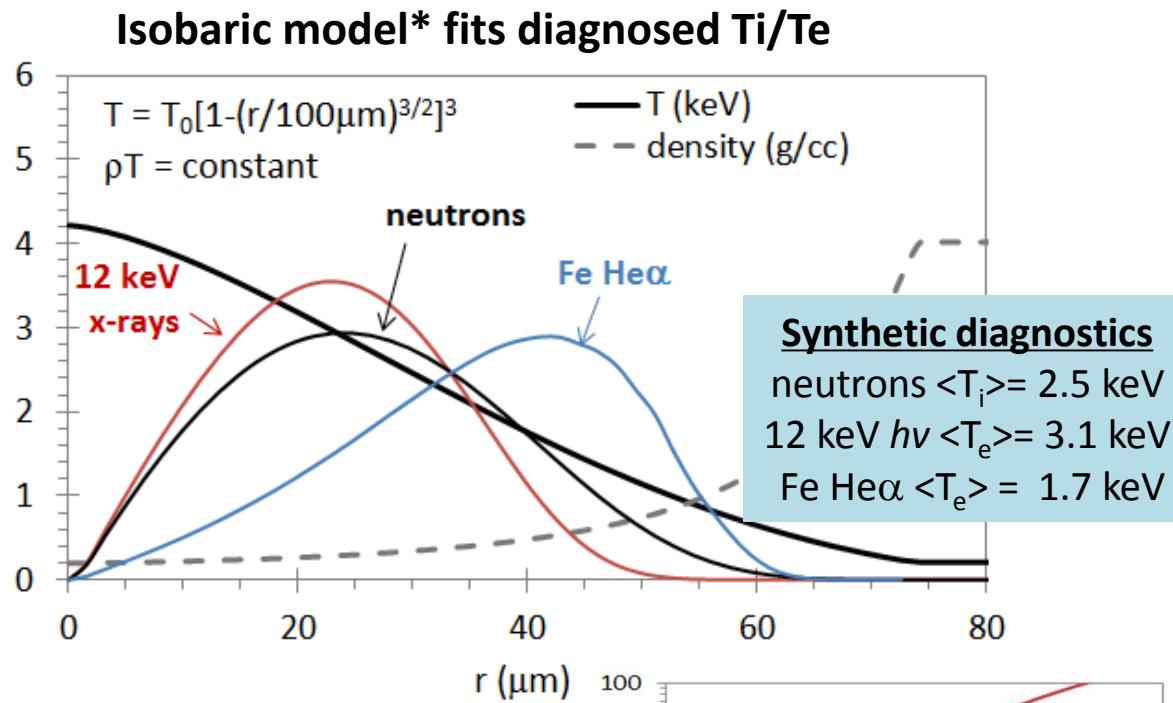
# Data constrains some gradients as well as burn averages

“cartoon” stagnation distilled from x-ray images & continuu spectra

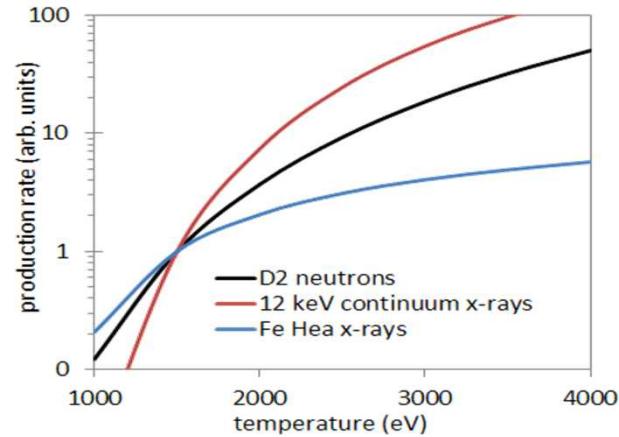


Inferences from high-res line spectra

$\rho_D = 0.33 \text{ g/cm}^3$   
 $T_e = 1.8 \text{ keV}$   
 $T_i = 2.3 \text{ keV}$   
 $\rho R_{\text{liner}} = 1.1 \text{ g/cm}^2$



High-energy (12-15 keV) continuum x-rays sample hotter portions of the plasma than Fe line emission (and even D-D neutrons)



# Stagnation data can constrain preheat energy and mix

- Neutron yields are highly sensitive to initial preheat energy (Slutz, Sefkow, Peterson, McBride) and radiation losses from mix (Slutz/McBride)
- Spectroscopy constrains mix (Be from Fe on XRS3; Al from Zn on CRITR)

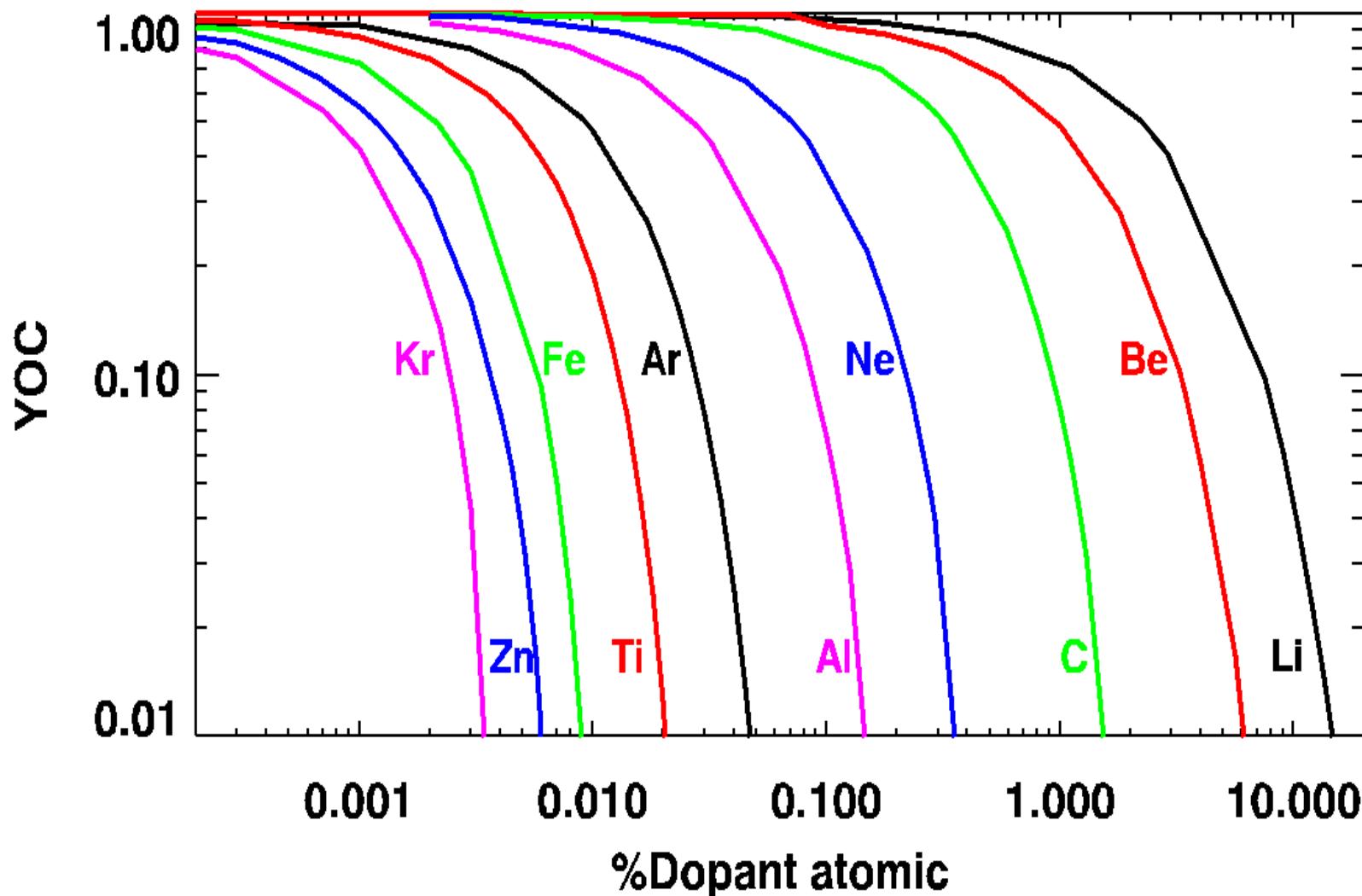
Window thickness ( $\mu\text{m}$ )	Preheat energy (J)	Endcap mix (%)	Window mix (%)	Liner mix (%)
3.5	150	0.05	0%	1%
3.5	300	0.05	5%	1%
1.5	150	0.10	0%	1%
1.5	300	0.15	5%	1%



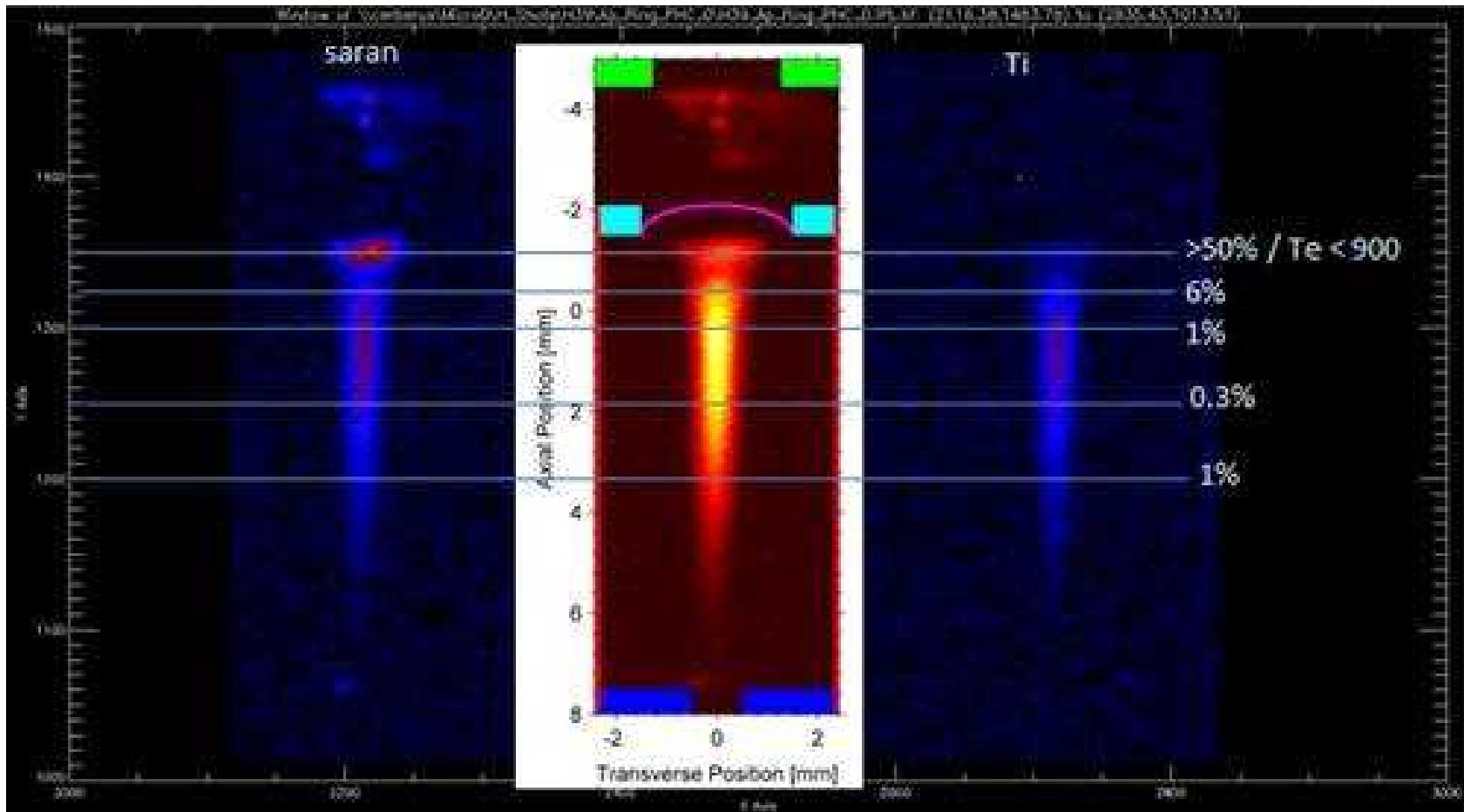
The diagram illustrates the interaction of a shock wave with a target. It shows a blue shock wave front moving towards a target, with an orange arrow indicating energy flow. The target is represented by a blue circle. Arrows point from the shock wave to the target, with labels 'Same?' and 'increase' indicating that the mix fractions are the same for different preheat energies and that they increase with energy. The shock wave is labeled 'shock -> brems'.

- Fitting yields from a very simple model to data from shots with Al & Be endcaps (and unconditioned beams) constrains preheat energy and mix
- Assumes that for a given laser energy and window thickness, the preheat energies and mix fractions must be the same regardless of endcap material
- Endcap mix fractions consistent with measured yields increase with laser energy under several plausible scenarios

# Yield degradation from mix (LTE?)



# Evidence for window mix from laser-only shot (H39)



- The ratio of TIPC intensities behind saran/Ti filters is sensitive to carbon fraction (which increases continuum over Ar line emission)

# How can we measure window mix?

- Characteristic X-rays may be the only way we can track material
- Flash-coat window/endcaps? (Harding)
- Late-time measurements seem possible even with 160 J of preheat energy; earlier time require dedicated targets
- With 500 J of preheat, temperatures support high-energy emission that could escpar liner

