

Low Background Signal Readout Electronics for the Majorana Demonstrator

This content has been downloaded from IOPscience. Please scroll down to see the full text.

[View the table of contents for this issue](#), or go to the [journal homepage](#) for more

Download details:

IP Address: 128.219.49.13

This content was downloaded on 04/01/2017 at 20:34

Please note that [terms and conditions apply](#).

You may also be interested in:

[A zero-suppression algorithm for the readout electronics of the SciFi Tracker for the LHCb detector upgrade](#)

H. Chanal

[Irradiation tests and expected performance of readout electronics of the ATLAS Hadronic Endcap calorimeter for the HL-LHC](#)

F Ahmadov

[Readout electronics and test bench for the CMS Phase I pixel detector](#)

Riccardo Del Burgo

[Low background x-ray detection with Micromegas for axion research](#)

S Aune, J F Castel, T Dafni et al.

[Readout electronics for the Wide Field of view Cherenkov/Fluorescence Telescope Array](#)

J. Zhang, S. Zhang, Y. Zhang et al.

[Caliste-MM: a spectro-polarimeter based on the micromegas concept for soft X-ray astrophysics](#)

P. Serrano, D. Attié, D. Desforge et al.

[Validation of a highly integrated SiPM readout system with a TOF-PET demonstrator](#)

T. Niknejad, S. Setayeshi, S. Tavernier et al.

[A fast embedded readout system for large-area Medipix and Timepix systems](#)

A S Brogna, M Balzer, S Smale et al.

Low Background Signal Readout Electronics for the MAJORANA DEMONSTRATOR

I. Guinn¹, N. Abgrall², F. T. Avignone III^{3,4}, A. S. Barabash⁵,
F. E. Bertrand⁴, V. Brudanin⁶, M. Busch^{7,8}, M. Buuck¹, D. Byram⁹,
A.S. Caldwell¹⁰, Y-D. Chan², C. D. Christofferson¹⁰, C. Cuesta¹,
J. A. Detwiler¹, Yu. Efremenko¹¹, H. Ejiri¹², S. R. Elliott¹³,
A. Galindo-Uribarri⁴, G. K. Giovanetti^{14,8}, J. Goett¹³, M. P. Green⁴,
J. Gruszko¹, V. E. Guiseppe³, R. Henning^{14,8}, E. W. Hoppe¹⁵,
S. Howard¹⁰, M. A. Howe^{14,8}, B. R. Jasinski⁹, K. J. Keeter¹⁶,
M. F. Kidd¹⁷, S. I. Konovalov⁵, R. T. Kouzes¹⁵, B. D. LaFerriere¹⁵,
J. Leon¹, J. MacMullin^{14,8}, R. D. Martin⁹, S. J. Meijer^{14,8},
S. Mertens², J. L. Orrell¹⁵, C. O'Shaughnessy^{14,8}, N. R. Overman¹⁵,
A. W. P. Poon², D. C. Radford⁴, J. Rager^{14,8}, K. Rielage¹³,
R. G. H. Robertson¹, E. Romero-Romero^{11,4}, M. C. Ronquest¹³,
B. Shanks^{14,8}, M. Shirchenko⁶, N. Snyder⁹, A. M. Suriano¹⁰,
D. Tedeschi³, J. E. Trimble^{14,8}, R. L. Varner⁴, S. Vasilyev⁶, K. Vetter²
¹⁸, K. Vorren^{14,8}, B. R. White⁴, J. F. Wilkerson^{14,8,4}, C. Wiseman³,
W. Xu¹³, E. Yakushev⁶, C-H. Yu⁴, and V. Yumatov⁵

The MAJORANA Collaboration

¹Center for Experimental Nuclear Physics and Astrophysics and
Department of Physics, University of Washington, Seattle, WA, USA

²Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA

³Department of Physics and Astronomy, University of South Carolina, Columbia, SC, USA

⁴Oak Ridge National Laboratory, Oak Ridge, TN, USA

⁵Institute for Theoretical and Experimental Physics, Moscow, Russia

⁶Joint Institute for Nuclear Research, Dubna, Russia

⁷Department of Physics, Duke University, Durham, NC, USA

⁸Triangle Universities Nuclear Laboratory, Durham, NC, USA

⁹Department of Physics, University of South Dakota, Vermillion, SD, USA

¹⁰South Dakota School of Mines and Technology, Rapid City, SD, USA

¹¹Department of Physics and Astronomy, University of Tennessee, Knoxville, TN, USA

¹²Research Center for Nuclear Physics and Department of Physics, Osaka University, Ibaraki, Osaka, Japan

¹³Los Alamos National Laboratory, Los Alamos, NM, USA

¹⁴Department of Physics and Astronomy, University of North Carolina, Chapel Hill, NC, USA

¹⁵Pacific Northwest National Laboratory, Richland, WA, USA

¹⁶Department of Physics, Black Hills State University, Spearfish, SD, USA

¹⁷Tennessee Tech University, Cookeville, TN, USA

E-mail: iguinn@uw.edu

¹⁸ Alternate Address: Department of Nuclear Engineering, University of California, Berkeley, CA, USA



Content from this work may be used under the terms of the [Creative Commons Attribution 3.0 licence](https://creativecommons.org/licenses/by/3.0/). Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI.

Abstract. The MAJORANA DEMONSTRATOR is a planned 40 kg array of Germanium detectors intended to demonstrate the feasibility of constructing a tonne-scale experiment that will seek neutrinoless double beta decay ($0\nu\beta\beta$) in ^{76}Ge . Such an experiment would require backgrounds of less than 1 count/tonne-year in the 4 keV region of interest around the 2039 keV Q-value of the $\beta\beta$ decay. Designing low-noise electronics, which must be placed in close proximity to the detectors, presents a challenge to reaching this background target. This paper will discuss the MAJORANA collaboration's solutions to some of these challenges.

1. Introduction to the MAJORANA DEMONSTRATOR

The MAJORANA DEMONSTRATOR (MJD) is an array of p-type point contact (PPC) high purity Germanium (HPGe) detectors intended to search for neutrinoless double beta decay ($0\nu\beta\beta$ decay) in ^{76}Ge [1][2][3][4]. MJD will consist of 40 kg of detectors, 30 kg of which will be isotopically enriched to 87% ^{76}Ge . The array will consist of 14 strings of four or five detectors placed in two separate cryostats. One of the main goals of the experiment is to demonstrate the feasibility of building a tonne-scale array of detectors to search for $0\nu\beta\beta$ decay with a much higher sensitivity. This involves achieving backgrounds in the 4 keV region of interest (ROI) around the 2039 keV Q-value of the $\beta\beta$ decay of less than 1 count/ROI-t-y. Because many backgrounds will not directly scale with detector mass, the specific background goal of MJD is less than 3 counts/ROI-t-y.

MJD uses a wide variety of background reduction techniques. The PPC geometry allows discrimination between multi-site events, consisting mostly of Compton-scattered gamma backgrounds, and single-site events, which includes $0\nu\beta\beta$ decay. The array is housed in passive shielding of copper, lead and high density polyethylene, along with an active muon veto system. Furthermore, the experiment is located 4850 ft underground at the Sanford Underground Research Facility (SURF), with 4260 mwe overburden to avoid cosmic rays. All materials used inside of the shielding are made of highly radiopure materials; in particular, the copper parts are made out of ultra-pure electroformed copper (EFCu) that is grown underground at SURF. An extensive radio-assay campaign verifies the purity of all materials used in the experiment. This assay data is used in a detailed model of the expected backgrounds of the DEMONSTRATOR[5]. In October 2014, the upper limit on background projected by the model was 3.1 counts in the ROI. A breakdown of the contributions from each background source in the ROI is shown in figure 1, with the largest contributions being the cables, electrical connectors and front end electronics inside of the cryostat. These background predictions are expected to shrink as assay limits improve.

2. Signal Readout Electronics

The signal readout electronics chain is responsible for integrating the current pulses from the Germanium detectors, amplifying that signal and carrying it out of the shielding and to a digitizer. The first stage of integration is done directly above each detector by low-noise, low-mass front-end boards (LMFEs). The signal is carried by bundles of coaxial cables along the strings of detectors and the thermosyphon crossarm to a feedthrough flange, where it is fed into a preamplifier. Because each end of the cable bundles are not easily accessible, signal connectors are placed above the cold plate to allow connection and disconnection of cables. Since all of these components are inside of the cryostat, it is important to design them with radiopure materials and with very low masses to minimize backgrounds, without compromising noise characteristics. The components must further be robust under vacuum and thermal cycling to liquid nitrogen temperature, and must not break when handled inside of a glove box with reduced dexterity.

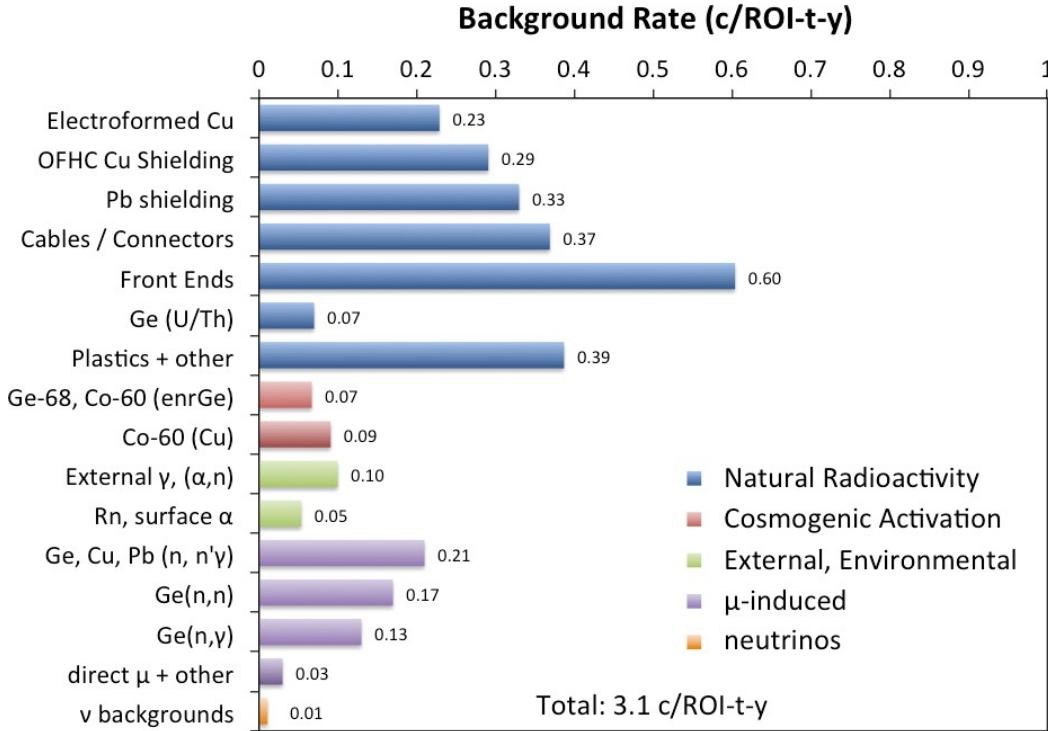


Figure 1. A summary of predicted background for the MAJORANA DEMONSTRATOR, produced in September 2014. The net backgrounds are 3.1 counts/ROI-t-y, with some of the largest contributions coming from signal readout electronics.

2.1. Low-Mass Front-End Boards

The LMFEs, shown in figure 2.1 are resistive feedback charge sensitive amplifiers that act as the first stage of amplification for the detectors [6]. The circuit consists of gold traces on a titanium adhesion layer printed via photolithography on a fused silica substrate. The feedback resistor consists of amorphous Germanium, which has much lower background than more commonly used ceramic resistors. The capacitance is provided by the geometric configuration of the traces. A JFET is attached to the circuit using low background silver epoxy, with the drain and source pads wirebonded to the circuit traces. The circuit also has a pulser line that allows external charge injection that can be used to characterize the circuit. The circuit is housed in an EFCu spring clip, which, when tensioned, presses the circuit against a contact pin at the p^+ point contact of the detector. An ICP-MS assay of an LMFE board and spring clip found backgrounds from Uranium and Thorium at $1.4 \mu\text{Bq/LMFE}$. The net background from the front-end boards is expected to be 0.60 counts/ROI-t-y.

2.2. Signal Cables

A bundle of four coaxial cables connects each LMFE board to a feedthrough flange. In order to reduce backgrounds, coaxial cables with an outer diameter of 0.4 mm and a mass density of 0.4 g/m are used, custom-made by Axon' Cable S.A.S. in France. A photograph of a bundle of 4 cables is shown, with a BNC coaxial cable for size comparison, in figure 3. Due to this small

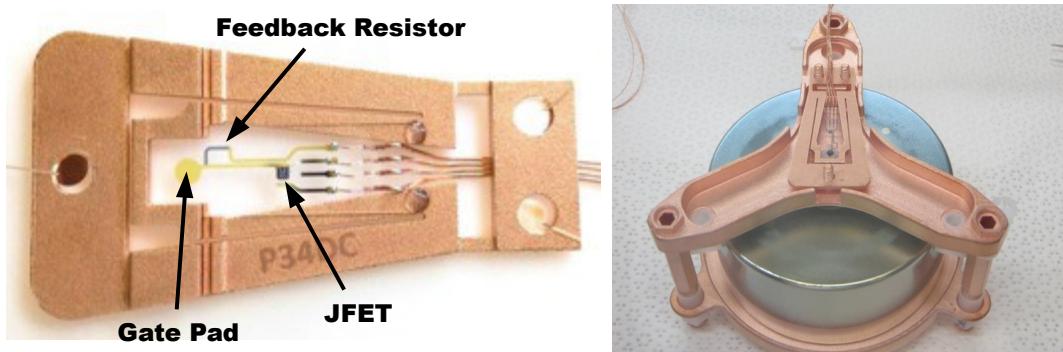


Figure 2. Left: An LMFE board mounted on a spring clip.
Right: an LMFE mounted on a detector unit. The contact pin is pressed between the gate pad and the point contact of the detector.

size, commercial clean copper is used rather than EFCu without heavily impacting backgrounds. The cables have an impedance of 50Ω over up to 85 in. ICP-MS assay of a short length of cable has found Uranium and Thorium backgrounds of $0.059\text{ }\mu\text{Bq}/\text{m}$, providing a total of 0.085 counts/ROI-t-y.

2.3. Signal Connectors

Signal connectors were built to connect two signal bundles together. Because most electrical connectors use electrical contact springs containing Beryllium Copper, which are difficult to make radiopure, no commercially available connectors have low enough backgrounds for MJD. To avoid use of springs, connectors were built that use Mill-Max® gold-plated brass pins with the contact springs removed. The pins and sockets are misaligned slightly, forcing the pin to bend. The restoring force from the bending pin does the job of the spring, maintaining a strong electrical connection. The pins are held in a Vespel® housing, which is machined underground. The signal cables are soldered, using a low background tin-silver eutectic, to the pins and sockets. PTFE Shrink tubing provides strain relief for the cables. Because of the high precision machining needed to properly align the pins and sockets, extensive quality control is necessary to ensure strong connections. A photograph of a signal plug pair is shown in figure 3. ICP-MS assay of the materials in the connectors has found Uranium and Thorium backgrounds at $1.45\text{ }\mu\text{Bq}/\text{connector}$, or $0.18\text{ }\mu\text{Bq}/\text{connection}$. The total contribution to backgrounds from these connectors is expected to be less than 0.28 counts/ROI-t-y. Using pins with BeCu would increase the background contribution from U and Th to above 10 counts/ROI-t-y.

3. Acknowledgements

This material is based upon work supported by the U.S. Department of Energy, Office of Science, Office of Nuclear Physics. We acknowledge support from the Particle Astrophysics Program of the National Science Foundation. This research uses these US DOE Office of Science User Facilities: the National Energy Research Scientific Computing Center and the Oak Ridge Leadership Computing Facility. We acknowledge support from the Russian Foundation for Basic Research. We thank our hosts and colleagues at the Sanford Underground Research Facility for their support.

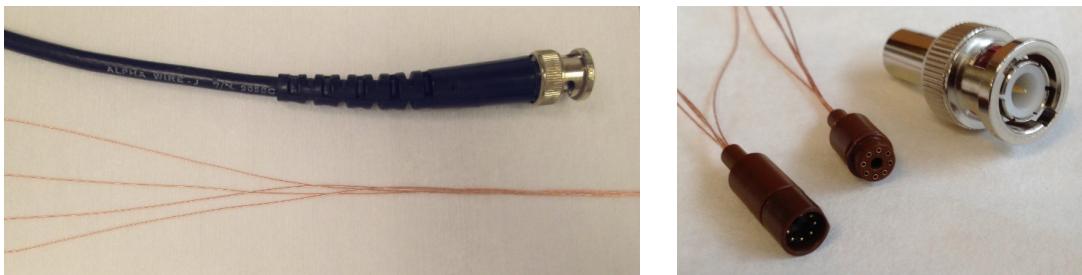


Figure 3. Left: A cable bundle with a BNC cable for comparison.
Right: A male-female signal connector pair, with a BNC connector for comparison.

References

- [1] N. Abgrall *et al.*, 2014 The MAJORANA DEMONSTRATOR Neutrinoless Double Beta Decay Experiment *Adv High Energy Phys.* 2014:365432
- [2] F.T. Avignone III, S.R. Elliot, and J. Engel, 2008 Double beta decay, Majorana neutrinos, and neutrino mass *Rev. Mod. Phys.*, 80 481
- [3] P.S. Barbeau, J.I. Collar, and O. Tench, 2007 Large-Mass Ultra-Low Noise Germanium Detectors: Performance and Applications in Neutrino and Astroparticle Physics *Journal of Cosmology and Astroparticle Physics*, 09 009
- [4] W. Xu *et al.*, 2015 The MAJORANA DEMONSTRATOR: A Search for Neutrinoless Double-beta Decay of ^{76}Ge , Proceedings for Workshop on Germanium-based Detectors and Technologies, 2014 IOP
- [5] C. Cuesta *et al.*, 2014 Background Model of the MAJORANA DEMONSTRATOR *Preprint* arXiv:1405.1370
- [6] P. Barton *et al.*, 2011 Low-noise low-mass front end electronics for low-background physics experiments using germanium detectors, proceedings for Nuclear Science Symposium and Medical Imaging Conference (NSS/MIC), 2011 IEEE