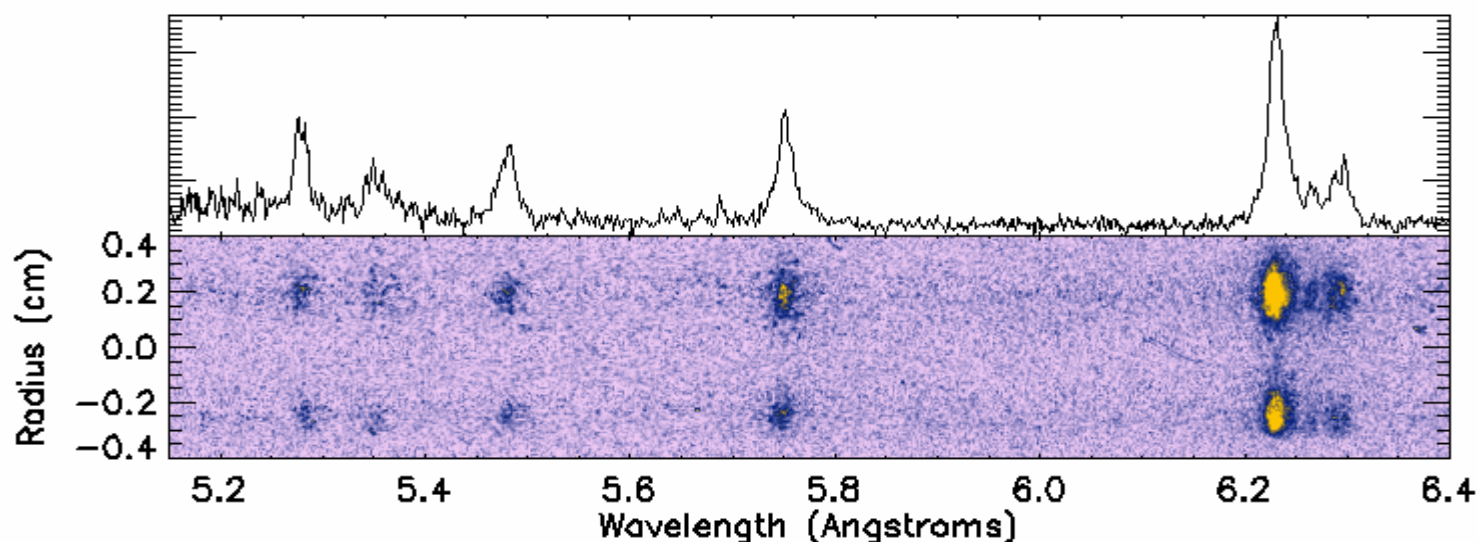


Si Emission Measurements of a Strong Radiating Shock in the Z-pinch Dynamic Hohlraum



Radiative Properties of Hot Dense Matter Workshop

September 11 – 15, 2006

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Sandia is a multiprogram laboratory operated by Sandia Corporation, a Lockheed Martin Company, for the United States Department of Energy under contract DE-AC04-94AL85000.



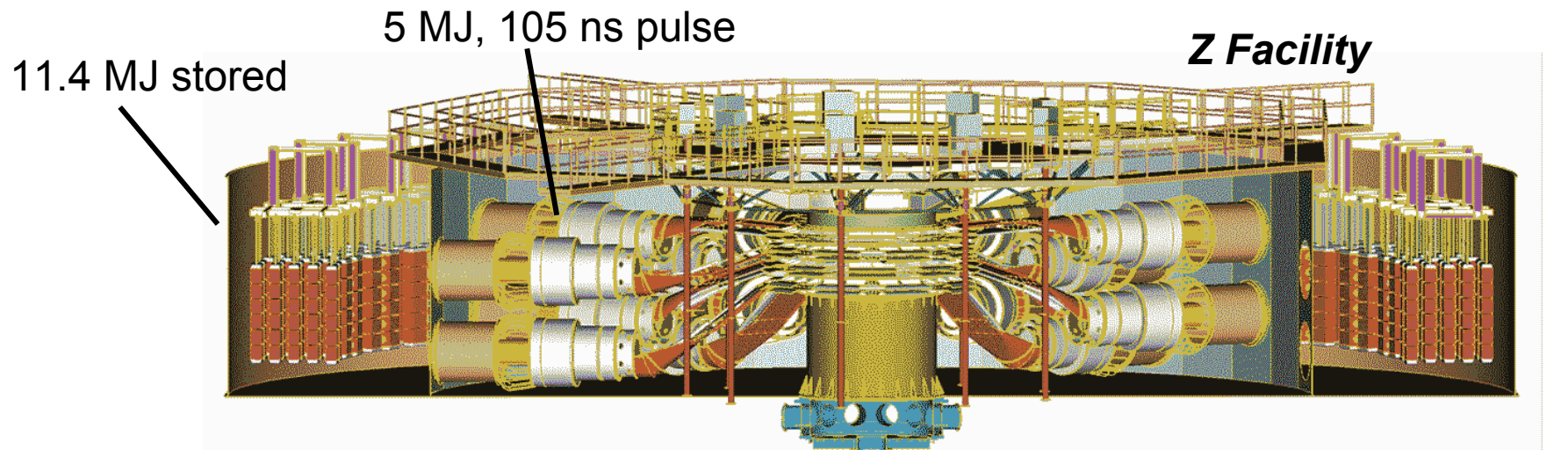


Acknowledgements

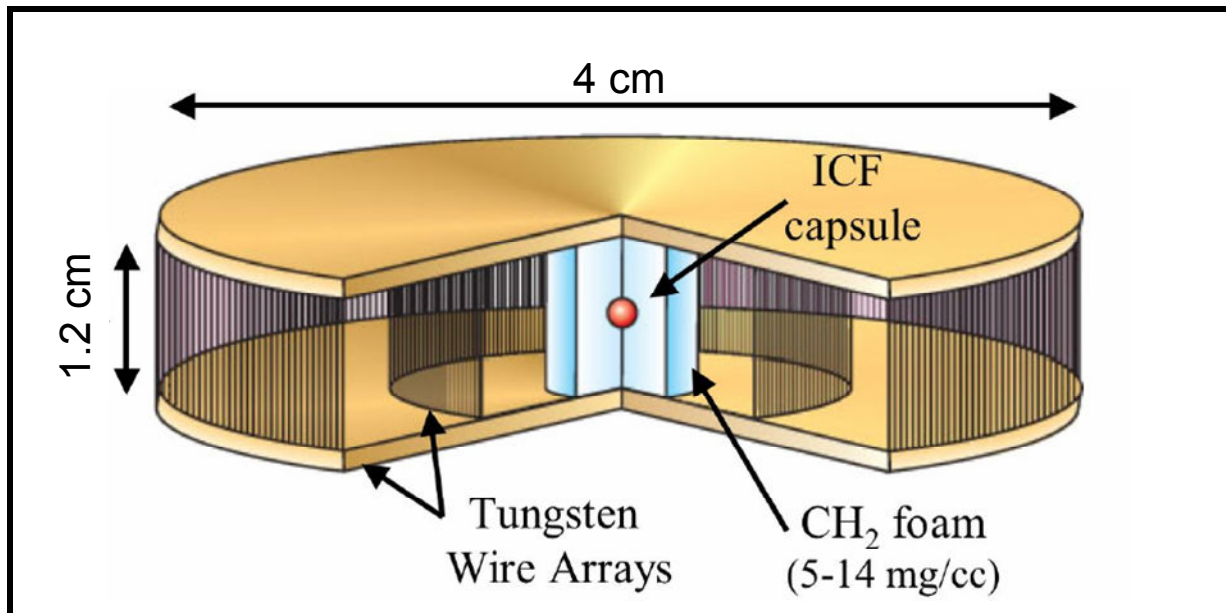
- Jim Bailey, Kyle Peterson, Steve Slutz, Gordon Chandler, Ray Lemke, Diana Schroen, Ray Leeper, Tom Mehlhorn
Sandia National Laboratories, Albuquerque, NM
- Yitzhac Maron, V. Fisher, E. Stambulchik
Weizmann Institute of Science, Israel
- Igor Golovkin, Joe MacFarlane
Prism Computational Sciences, Madison, WI



The Z-pinch dynamic hohlraum reaches internal temperatures >200 eV and peak axial power >10 TW.



~ 2 MJ electrical at Z pinch



Total Radiated X-ray Energy ~ 0.8 MJ

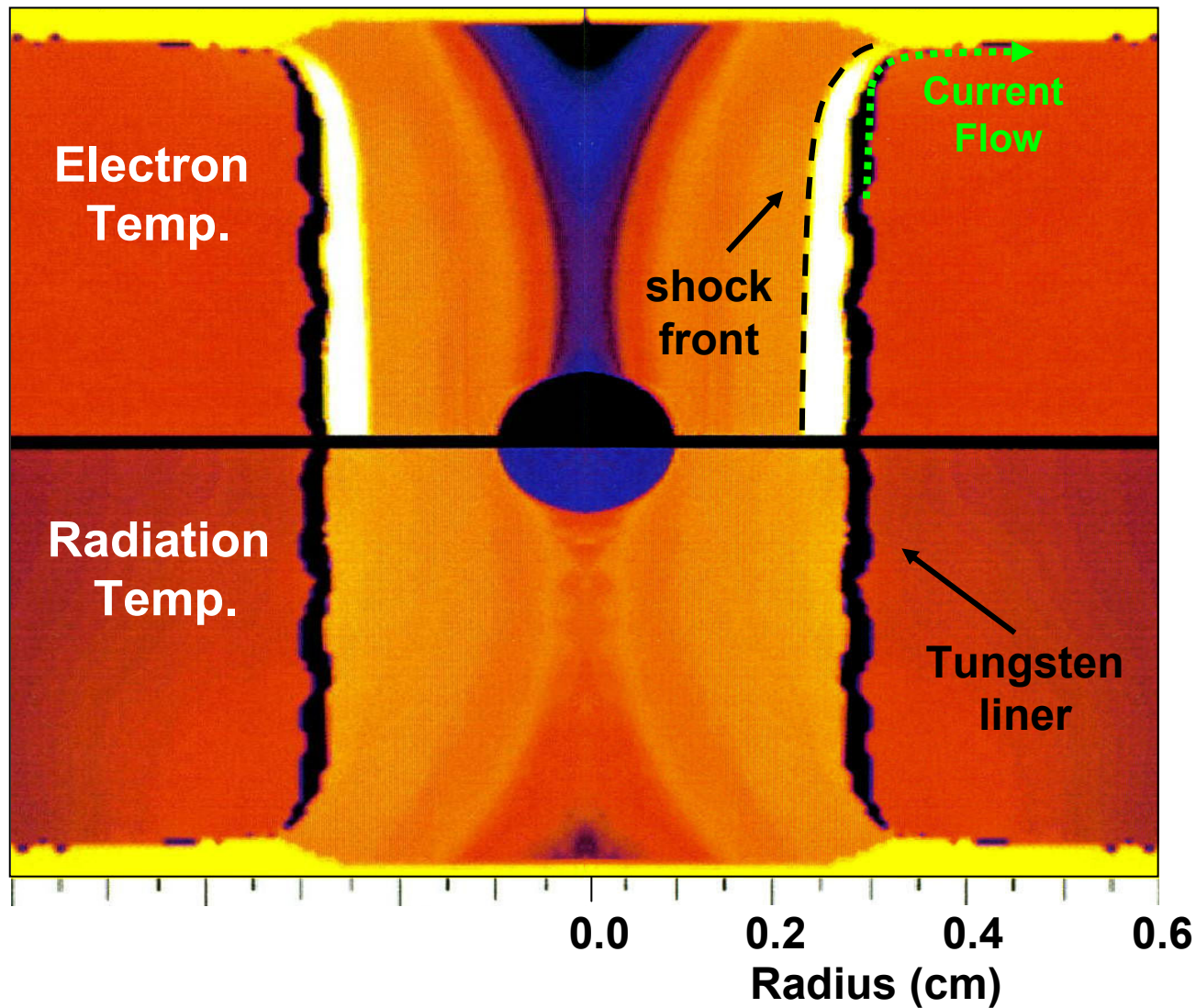
Peak Axial Power > 10 TW

Peak Internal Rad. Temp. > 200 eV



The dynamic hohlraum is formed by an imploding tungsten liner, and heated by a strong radiating shock.

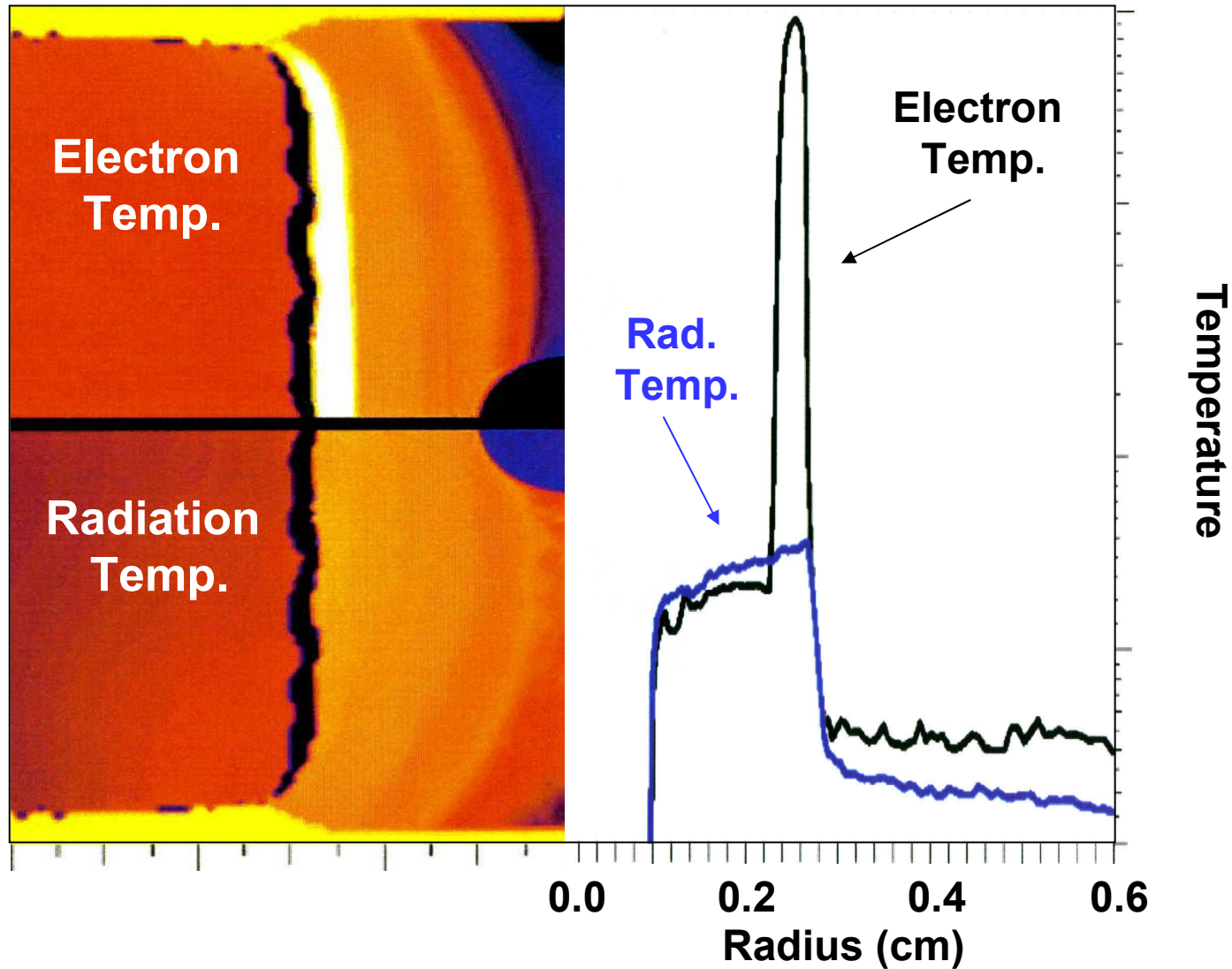
2-D RMHD Simulation





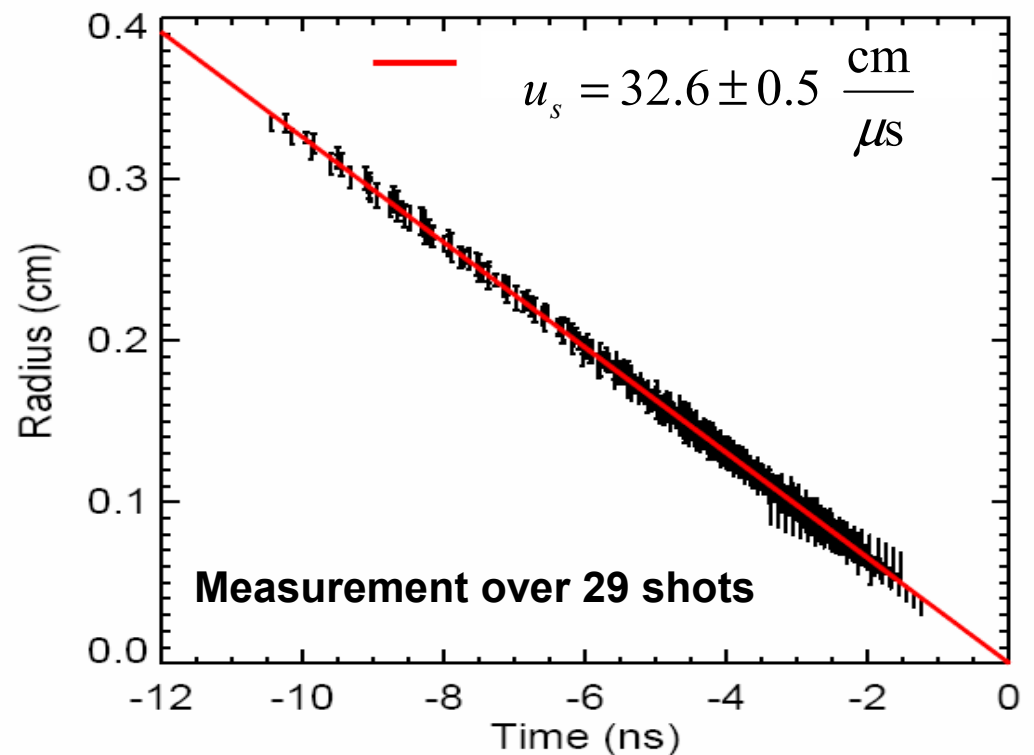
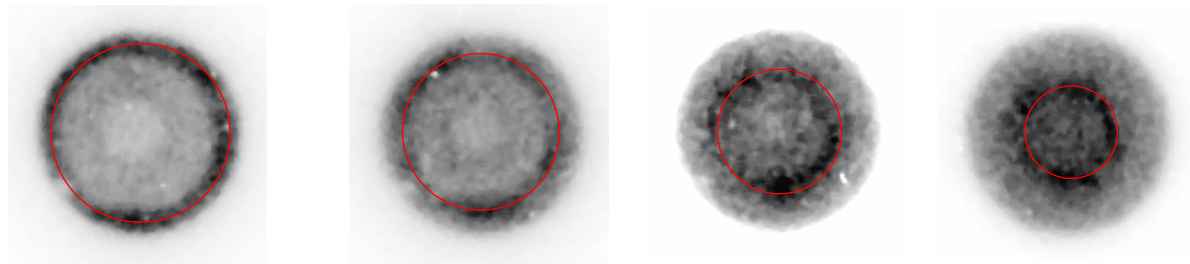
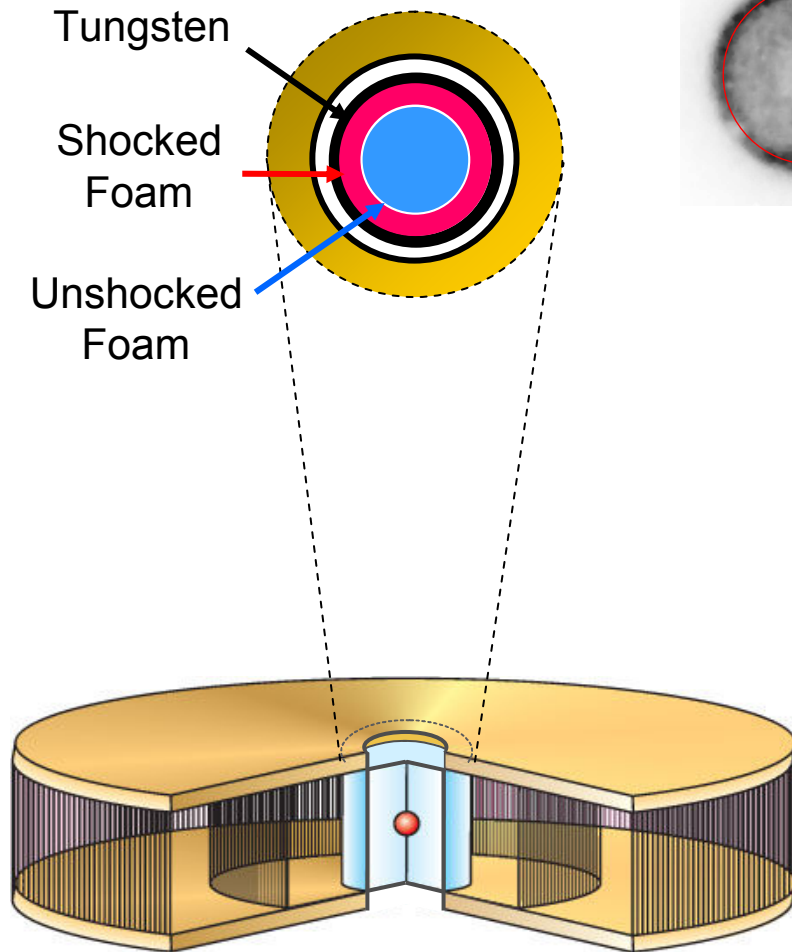
The dynamic hohlraum is formed by an imploding tungsten liner, and heated by a strong radiating shock.

2-D RMHD Simulation





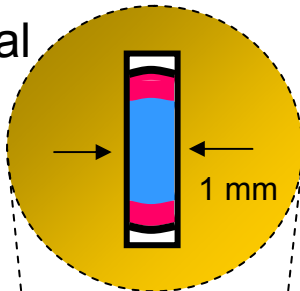
The radiating shock was first evidenced by 2-D broadband x-ray images.





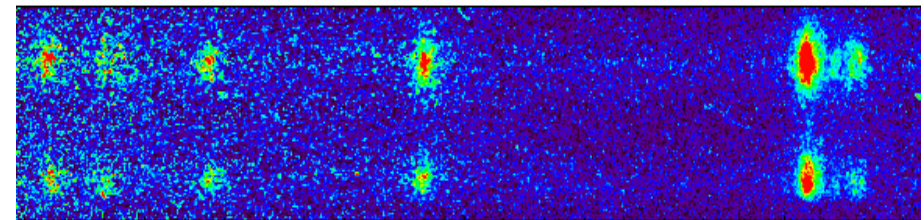
Si emission spectroscopy has been used to probe the local conditions of the DH source shock.

slot aperture
for 1-D spatial
resolution

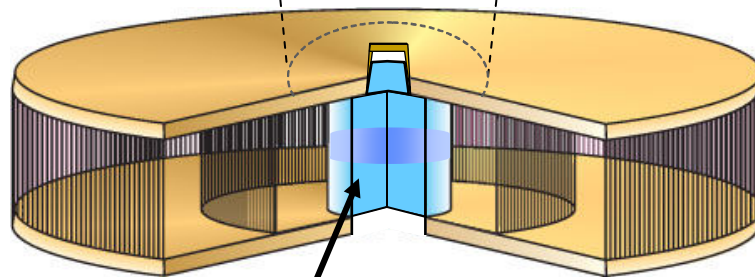


Shocked Si
Emission

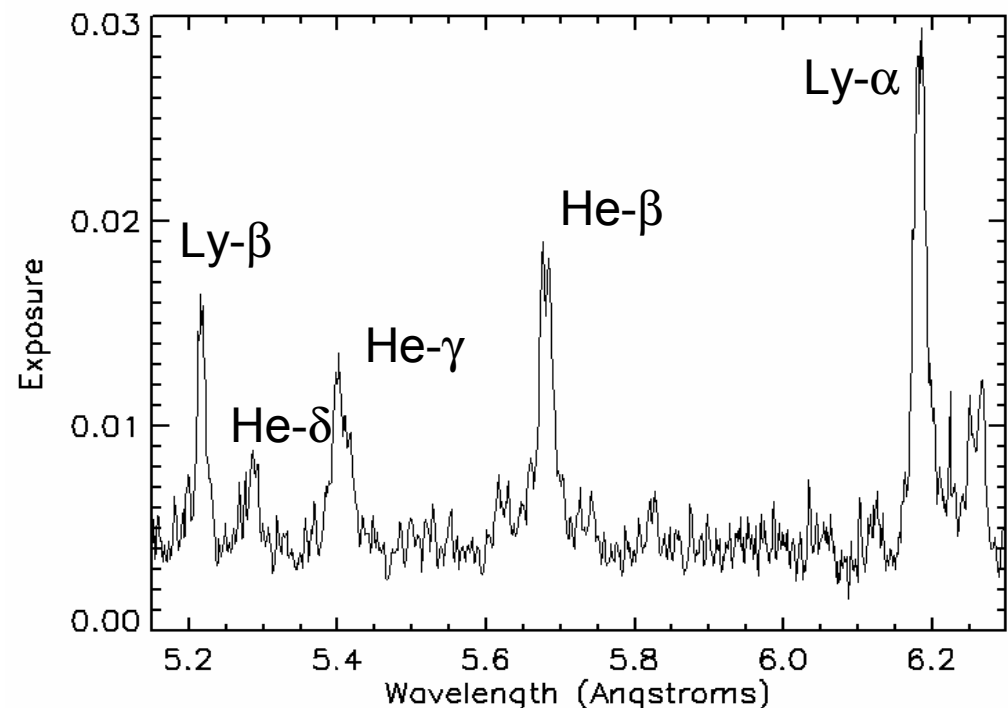
Time- and 1-D Space-Resolved Si Spectra



Wavelength

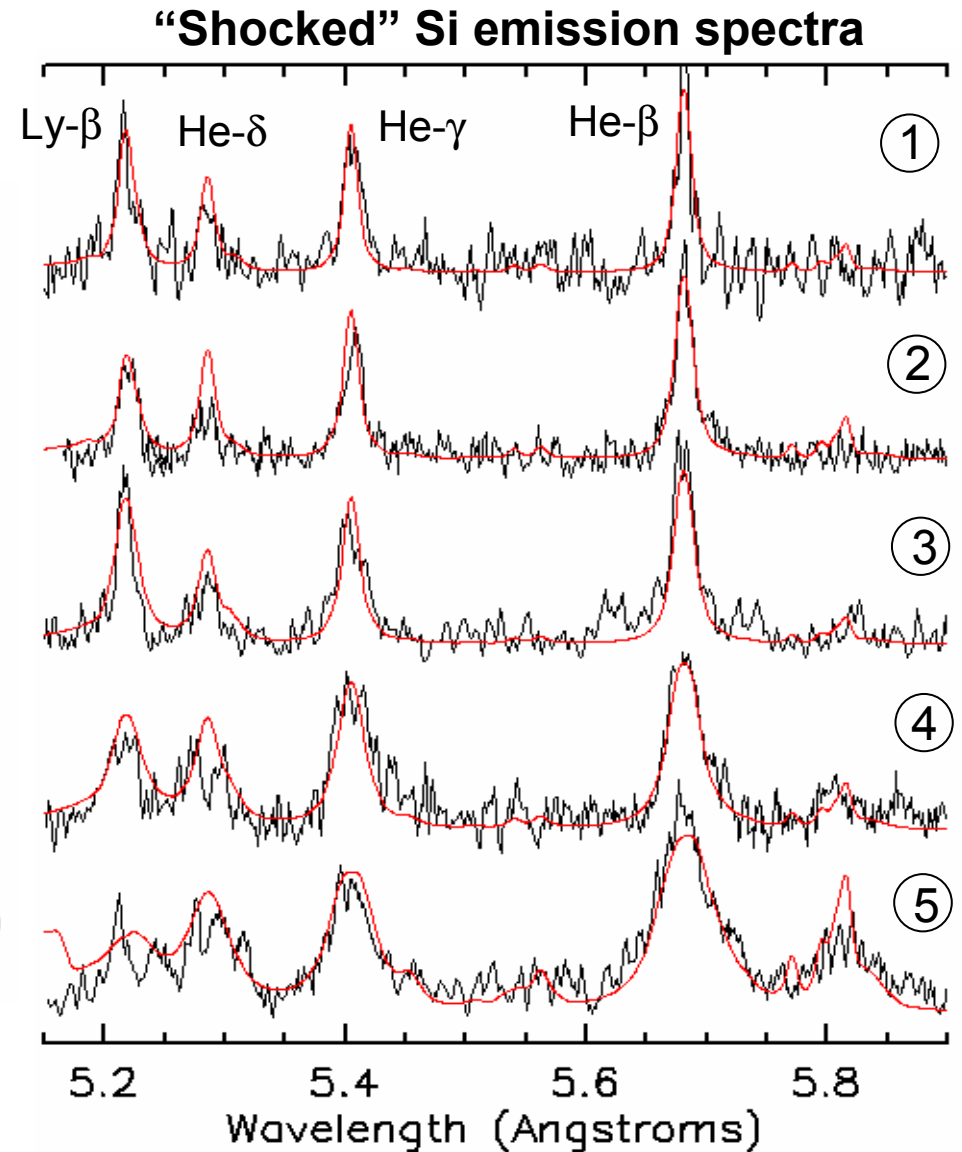
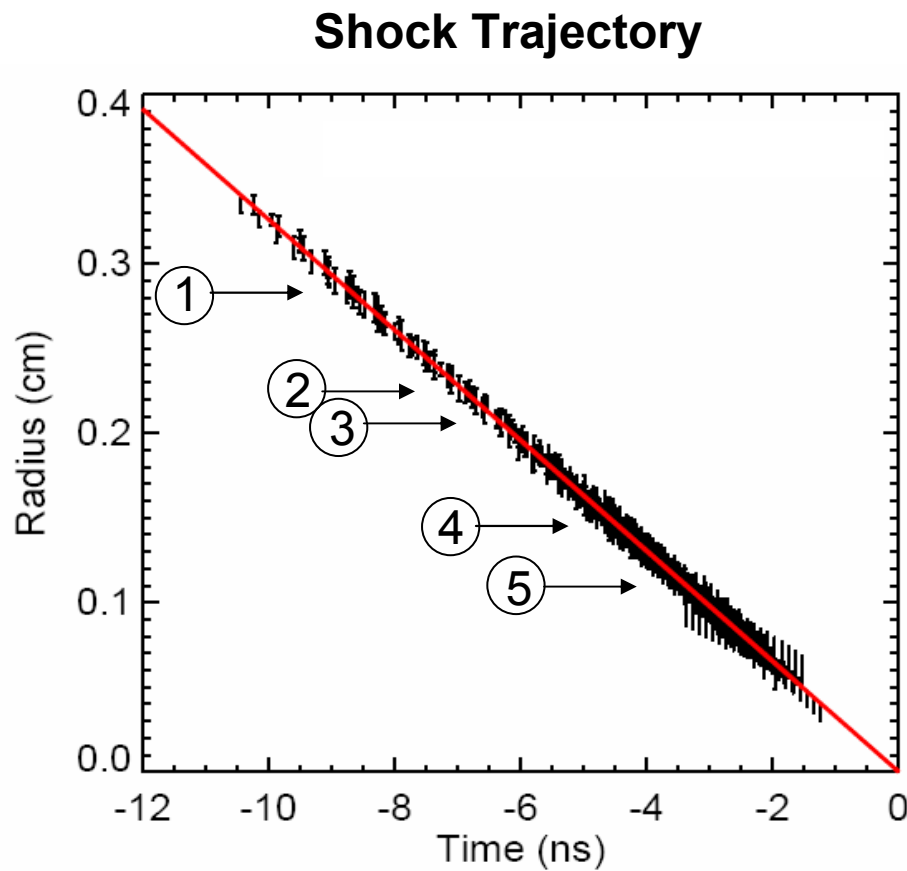


central 3mm height
of the 15 mg/cc CH_2 foam
is doped with 2.5% Si_2O_3





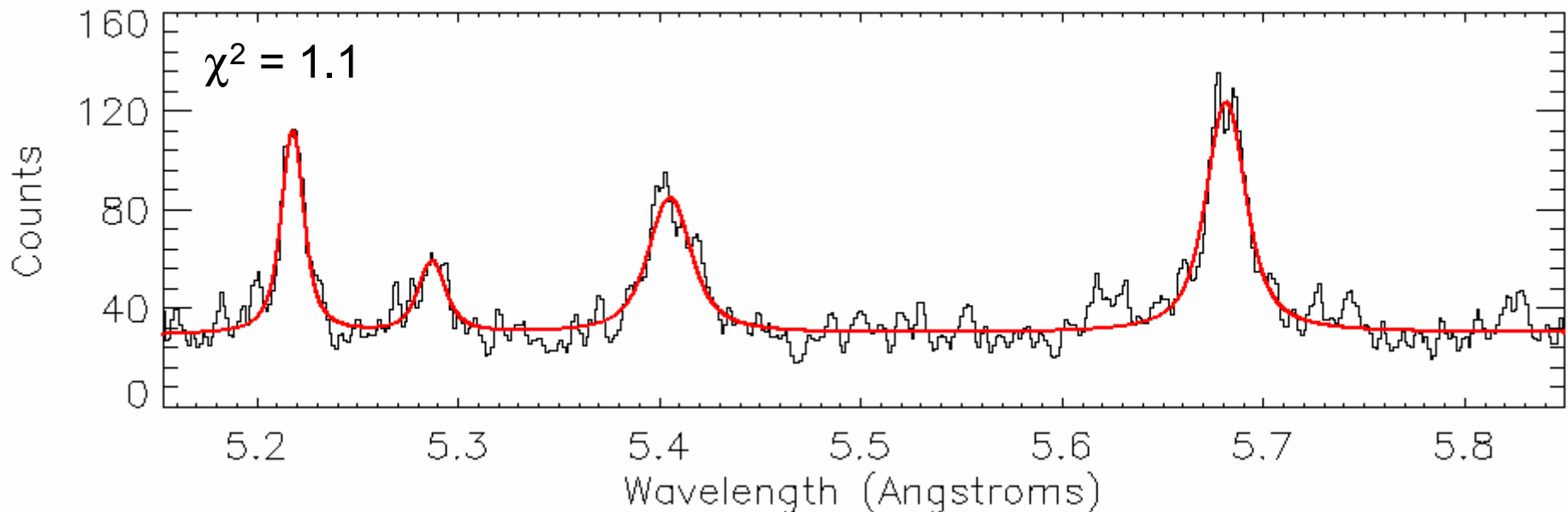
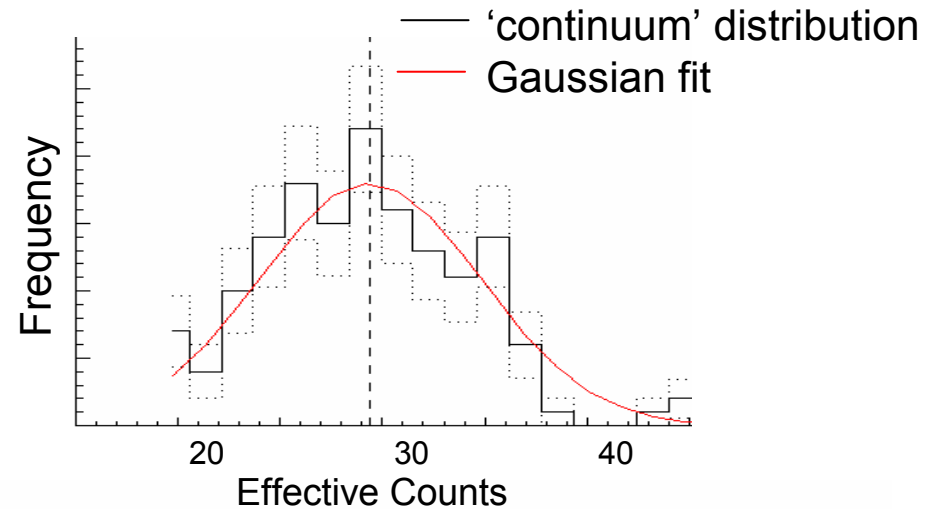
Si emission from the shocked DH foam has been recorded for shock radii from 1.1 – 2.8 mm.





“Noisy” data requires a careful statistical analysis to determine line-widths and line-strengths with error bars.

- Determine frequency distribution of the data continuum and assume Poisson statistics to get weights.
- Perform a least-squares fit to the spectra using ROBFIT¹
 - assume Voigt line profiles ($\eta = 1$)
 - simultaneously fit background and emission lines

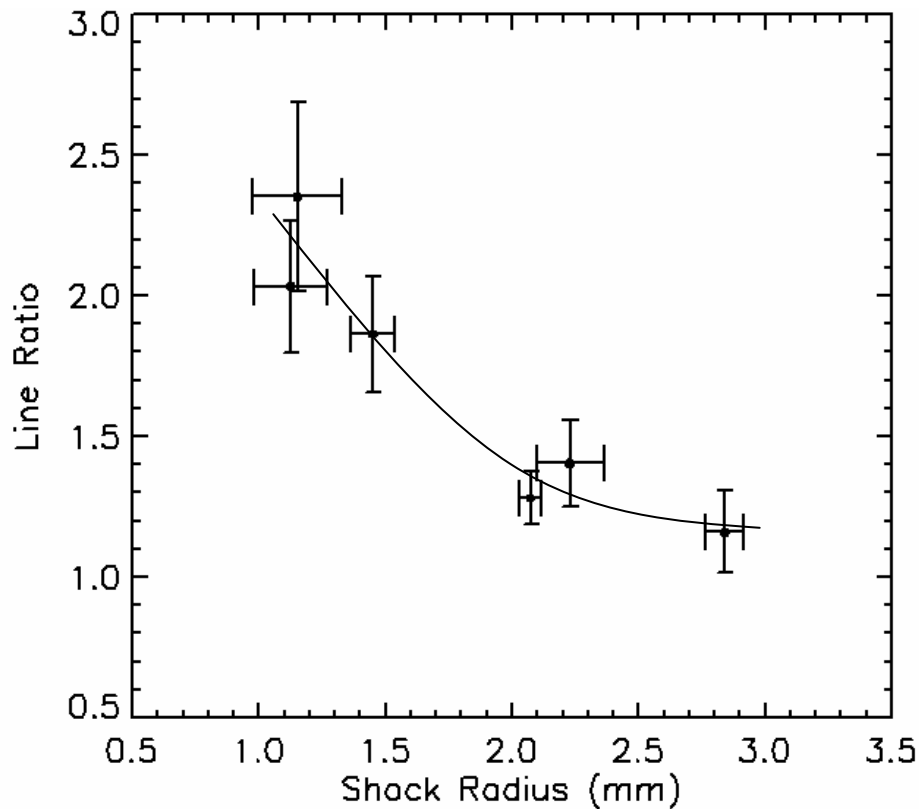


¹R.L. Coldwell and G.J. Bamford, *The Theory and Operation of Spectral Analysis using ROBFIT*, (AIP, 1991).

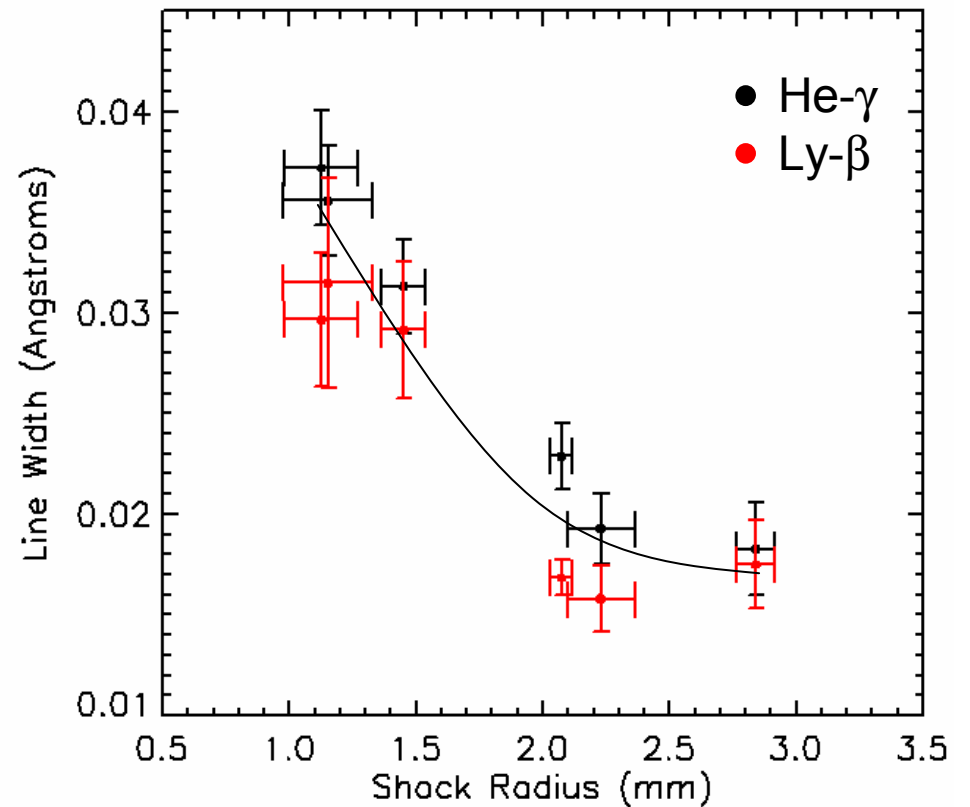


The ROBFIT analysis shows He / H line ratios and line widths that increase with decreasing radius.

He- γ / Ly- β Ratio



Line Widths

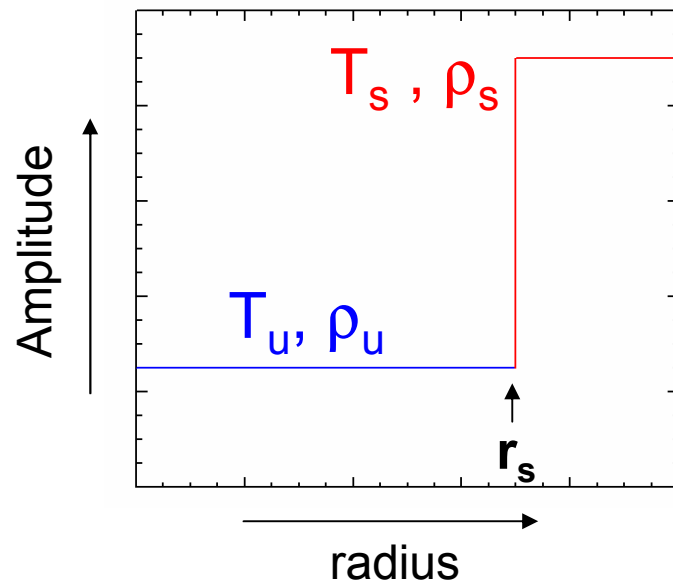




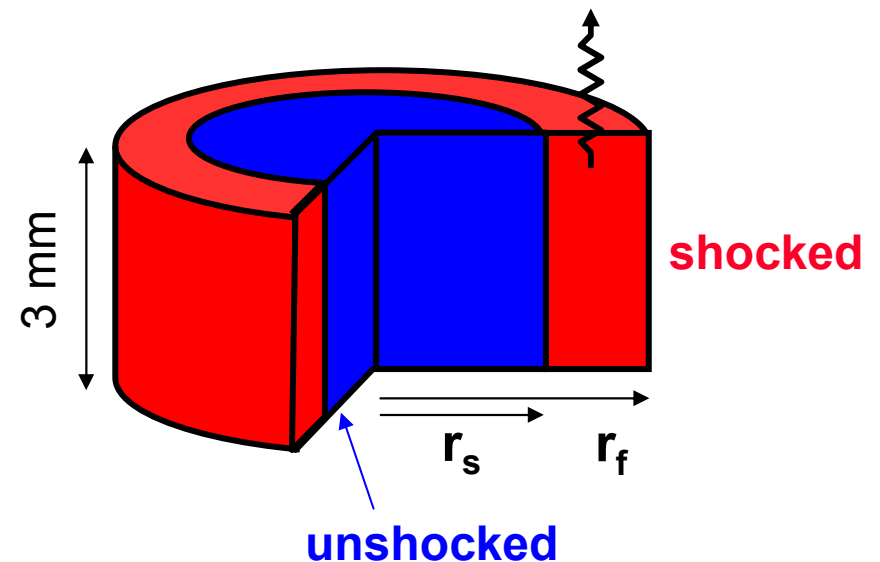
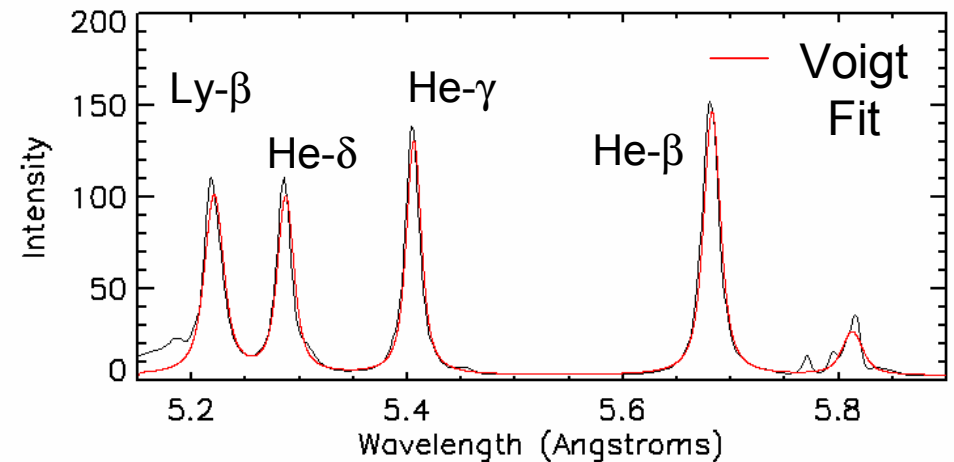
Shock conditions are determined from a 2-D model with non-LTE population kinetics and non-local radiation.

SPECT3D modeling

- C-R radiation transport
- Non-LTE population kinetics
- >4000 atomic level transitions
- Fix T_u, ρ_u -- vary T_s, ρ_s
- Fit each simulated spectra with ROBFIT
- Tabulate line ratios / widths vs. T_s, ρ_s



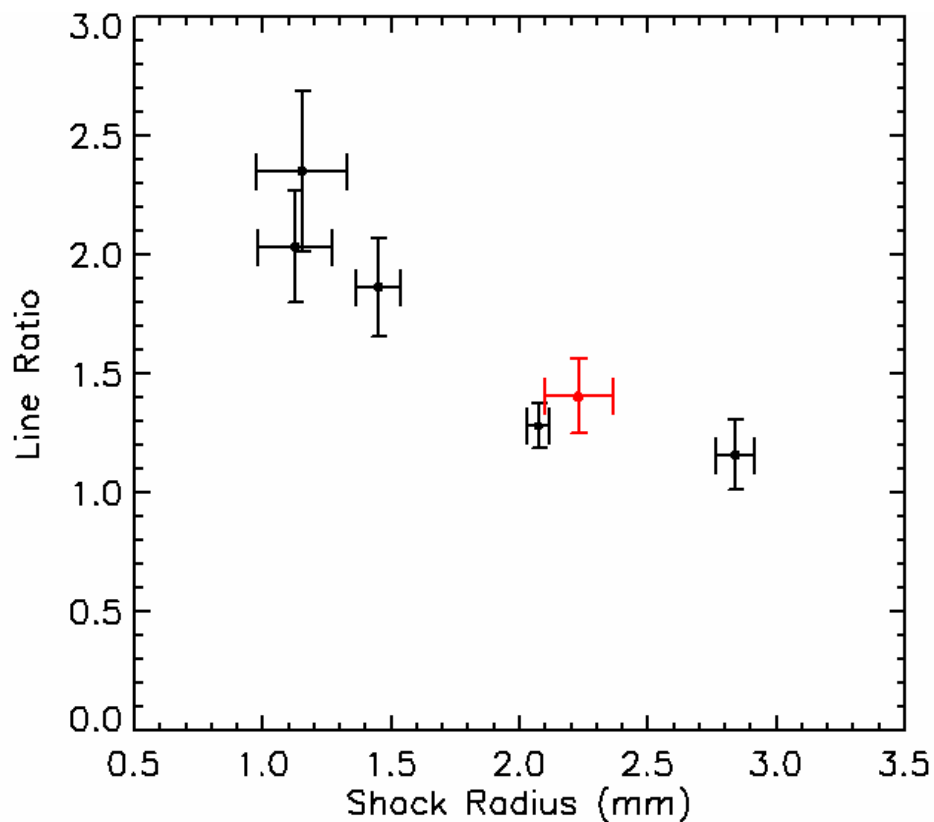
Simulated Si Emission Spectrum



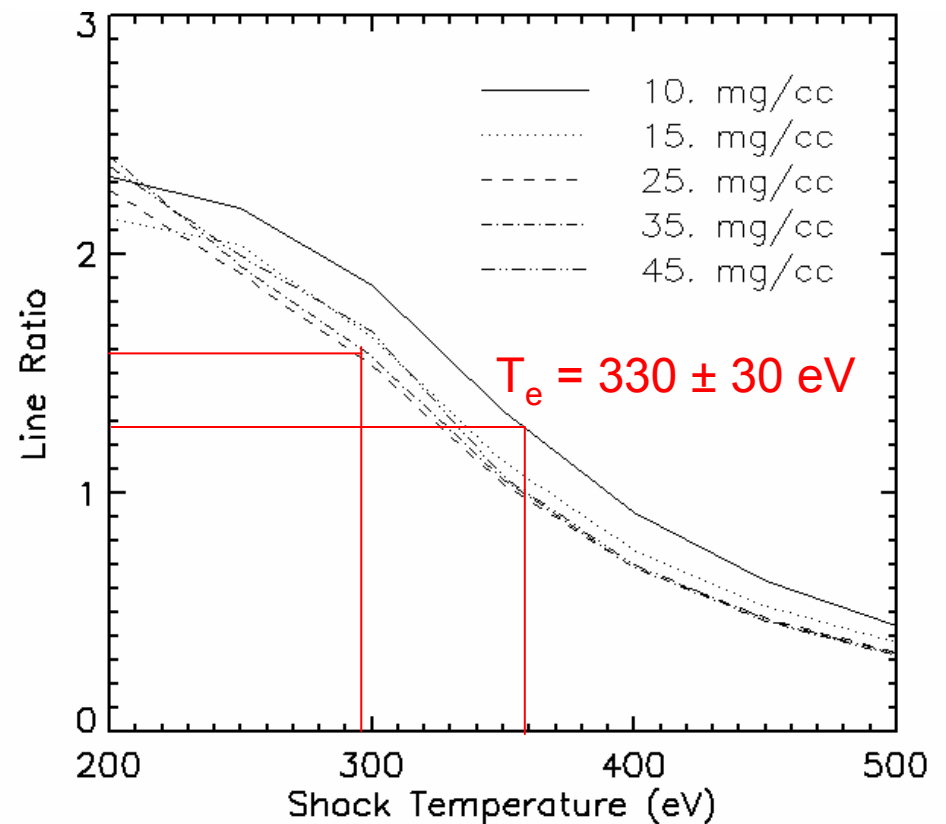


The shock temperature is determined from the He-like to H-like line ratios.

Measured
He- γ / Ly- β Line Ratios



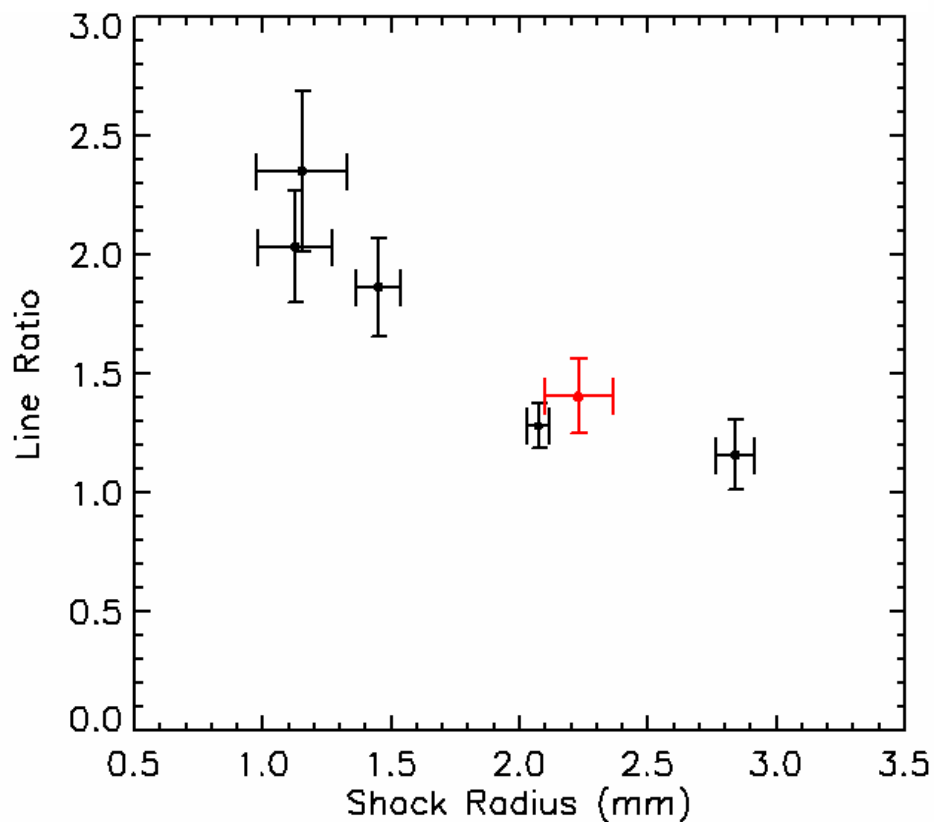
Calculated He- γ / Ly- β
Line Ratios vs. Shock Temperature



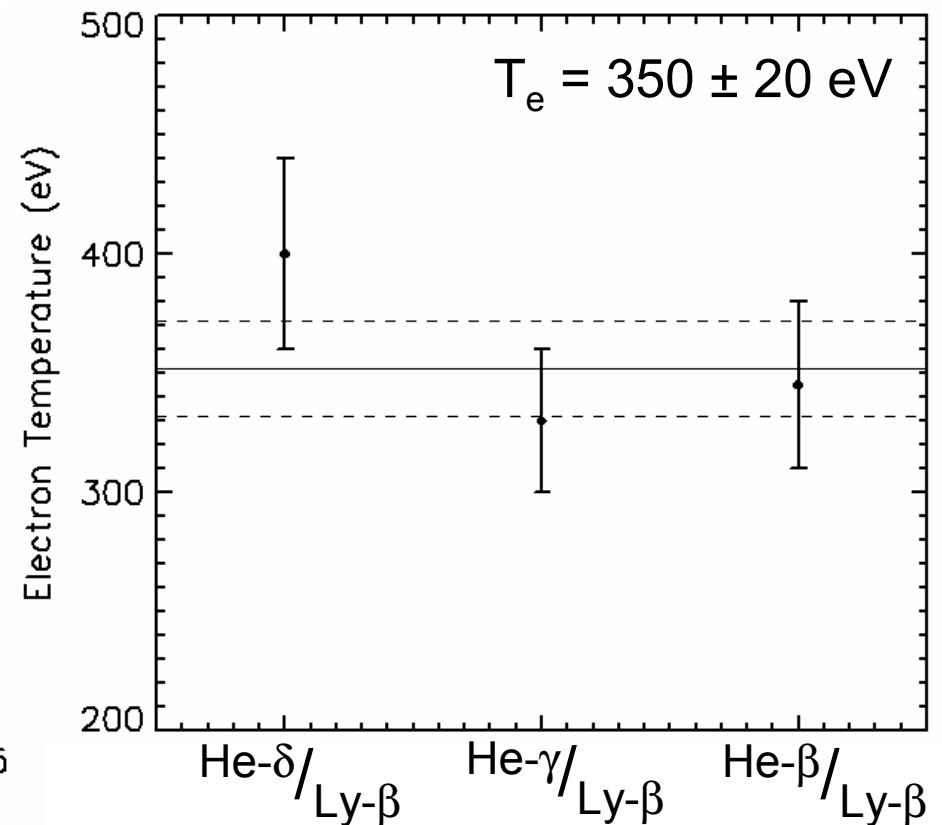


The shock temperature is determined from the He-like to H-like line ratios.

Measured
He- γ / Ly- β Line Ratios



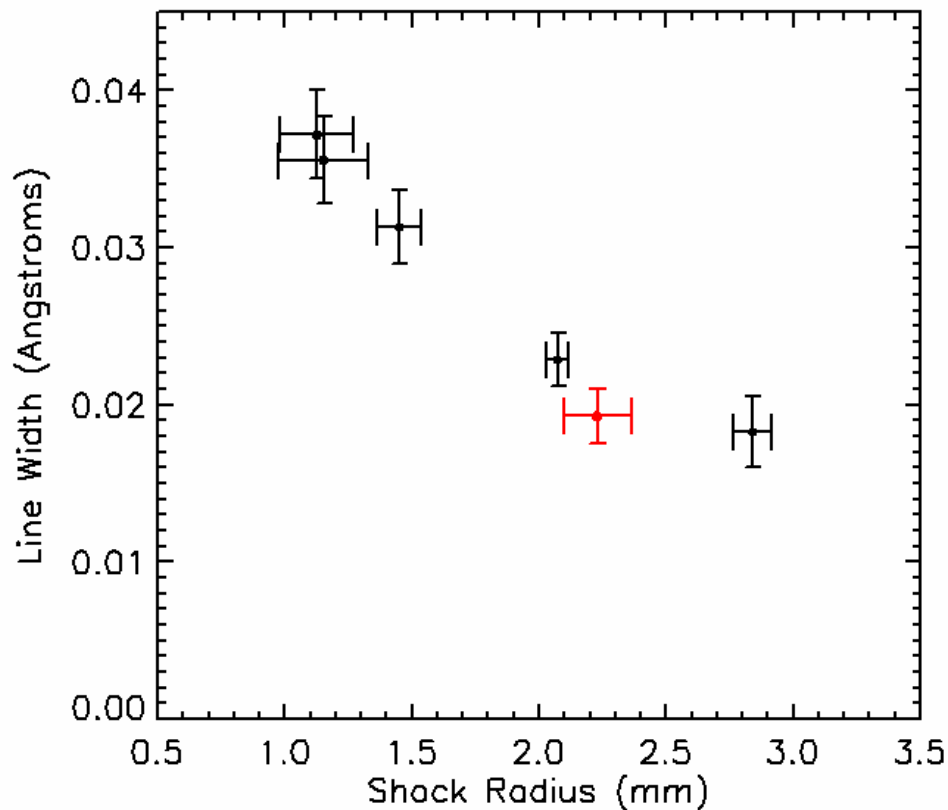
Temperatures Inferred from
Calculated Line Ratios at $r_s = 2.2$ mm



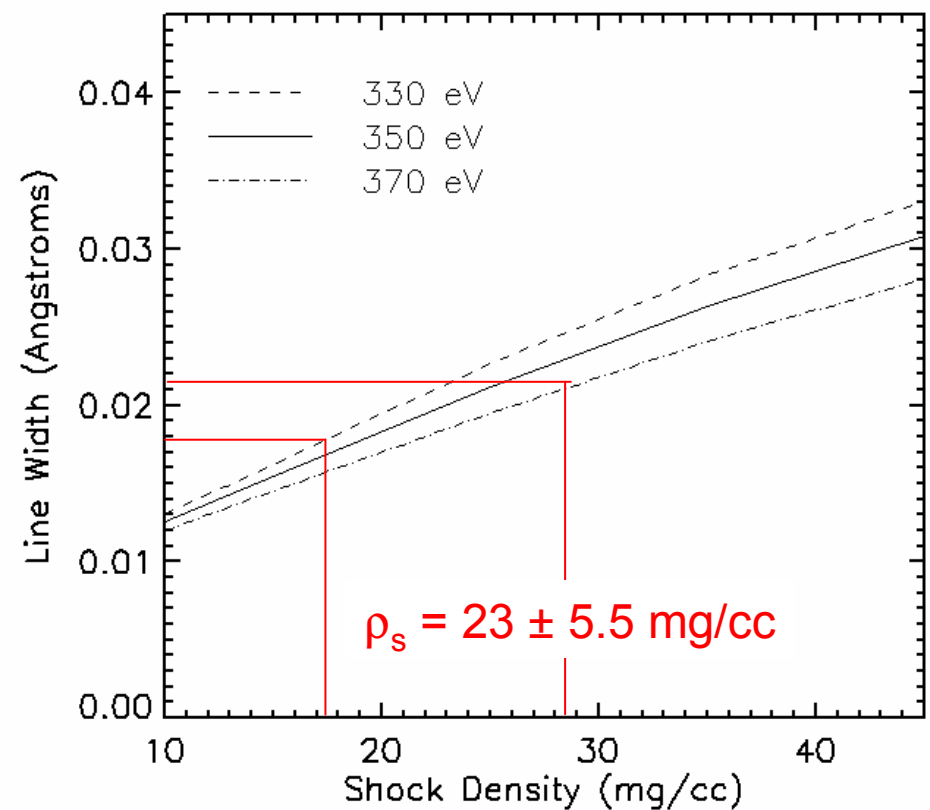


Once the temperature is known, the shock density is determined from the opacity & stark broadened line widths.

Measured
He- γ Line Widths



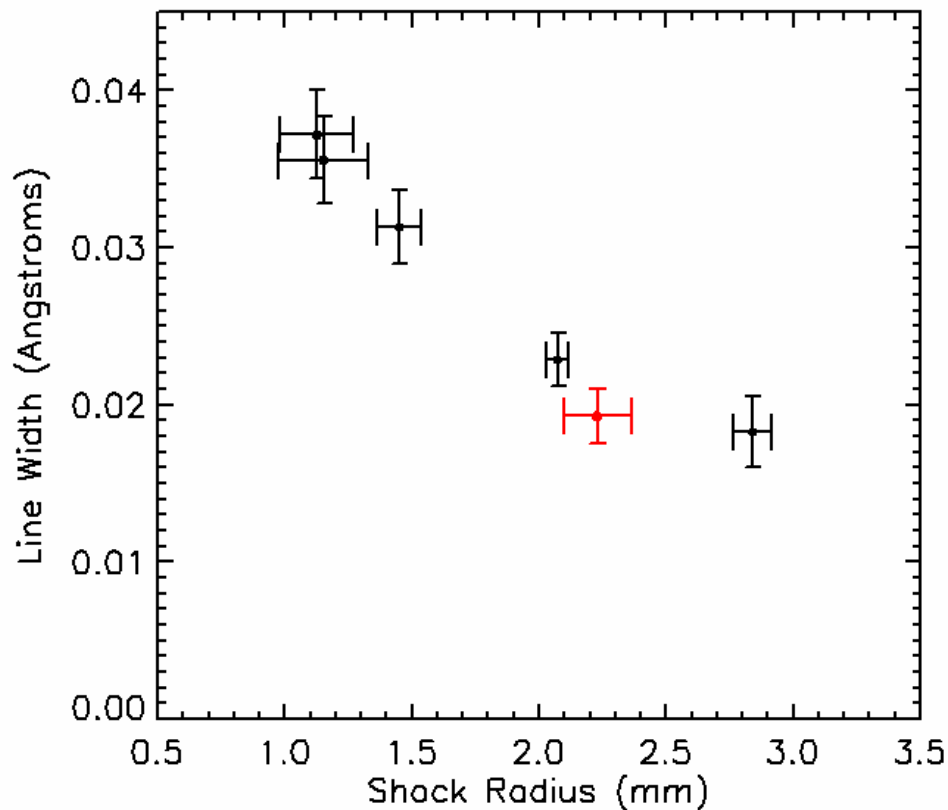
Calculated He- γ Line
Width vs. Shock Density



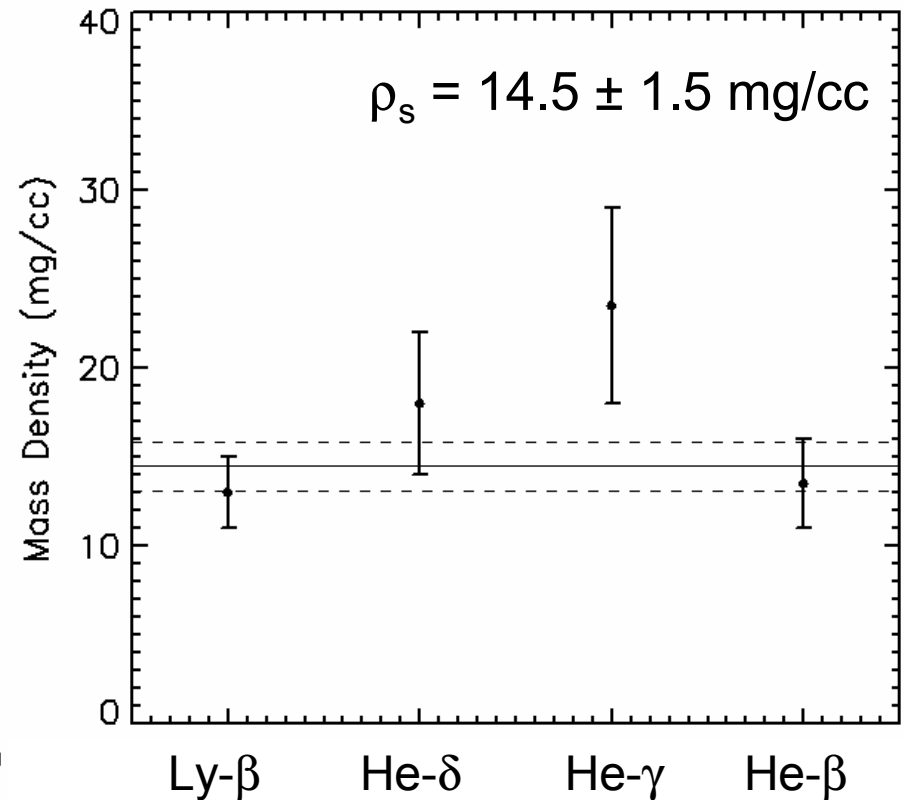


Once the temperature is known, the shock density is determined from the opacity & stark broadened line widths.

Measured
He- γ Line Widths

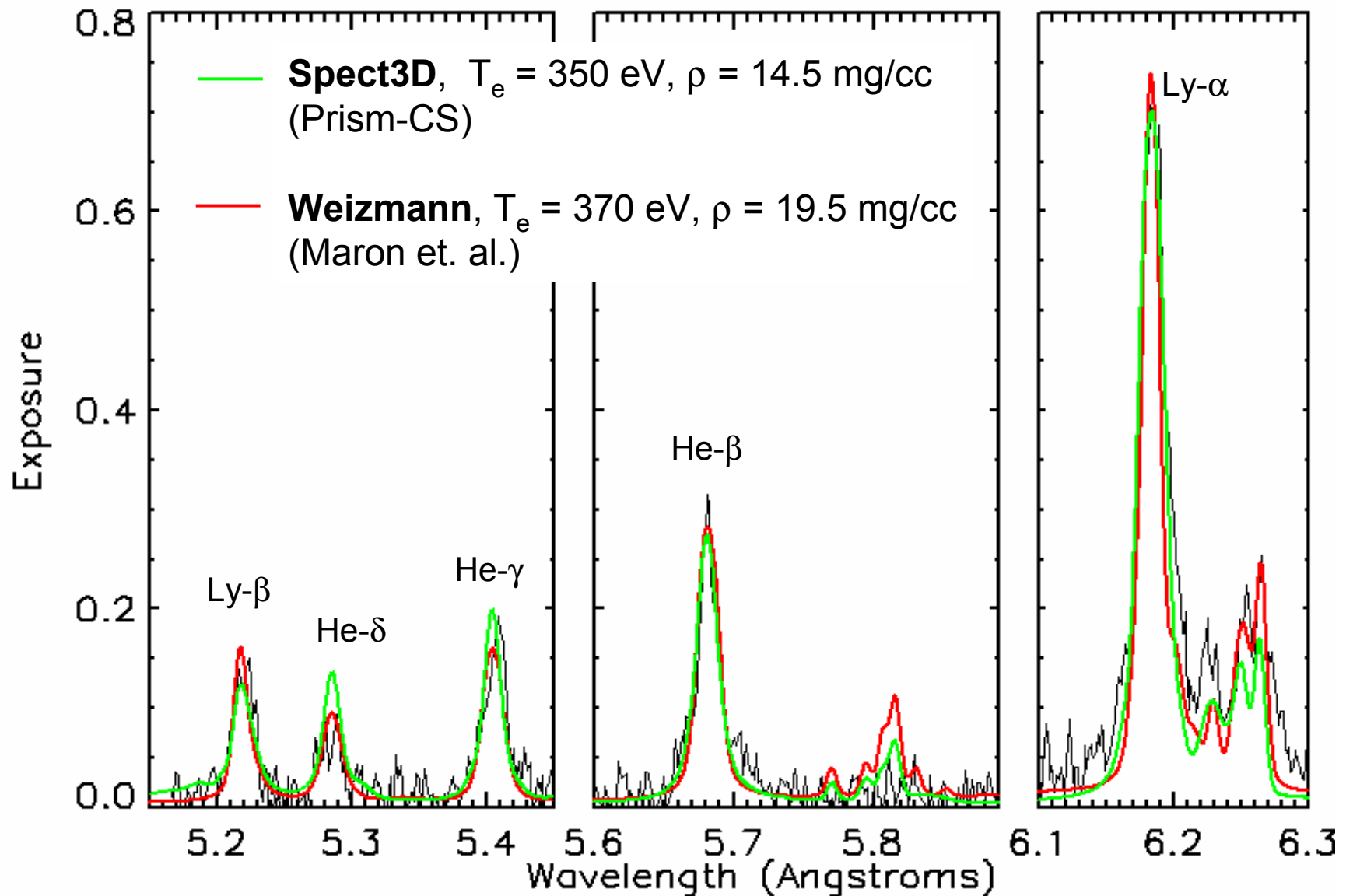


Densities Inferred from Calculated
Line Widths at $r_s = 2.2$ mm



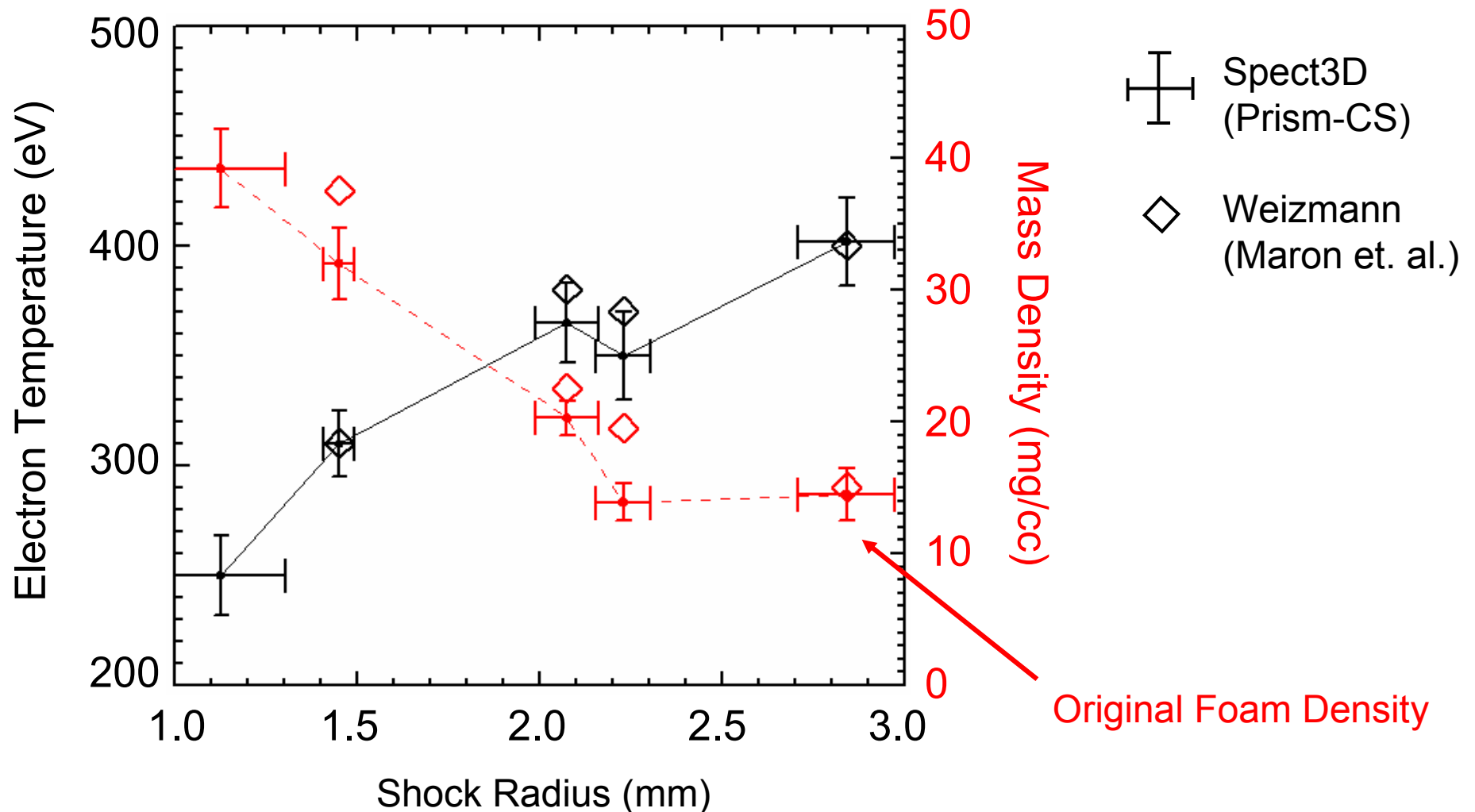


The 'Best-Fit' plasma conditions are then tested for consistency across the measured range.



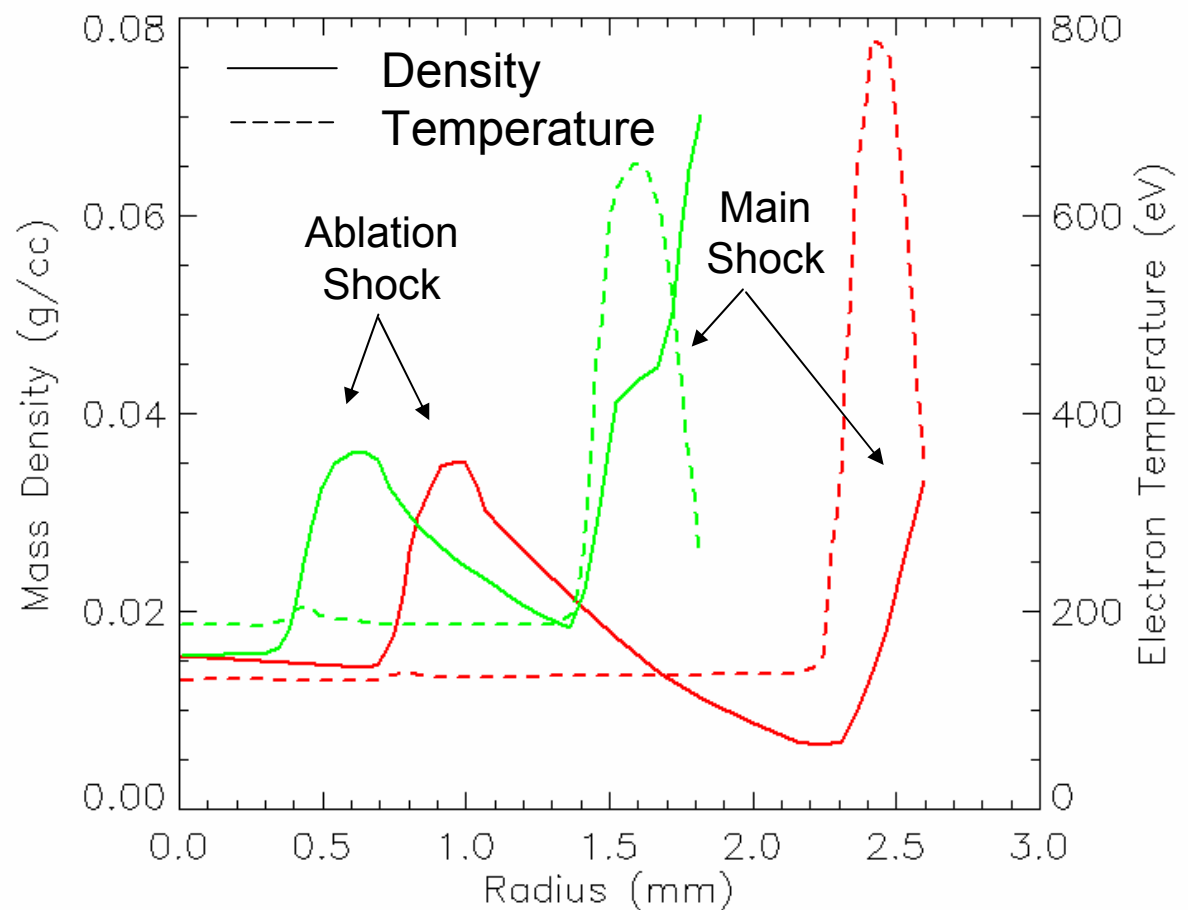
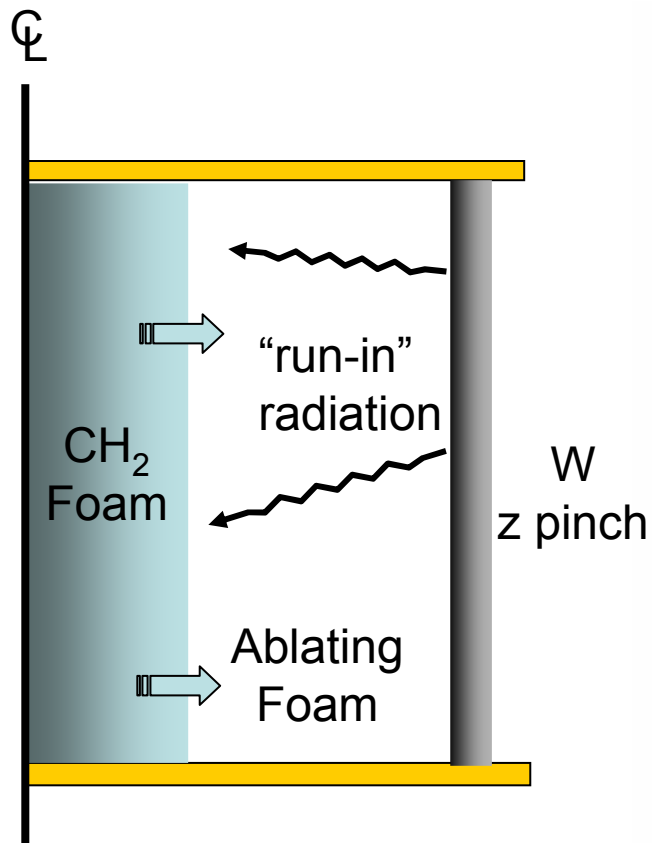


The measured shock temperature decreases and the shock density increases with decreasing shock radius.





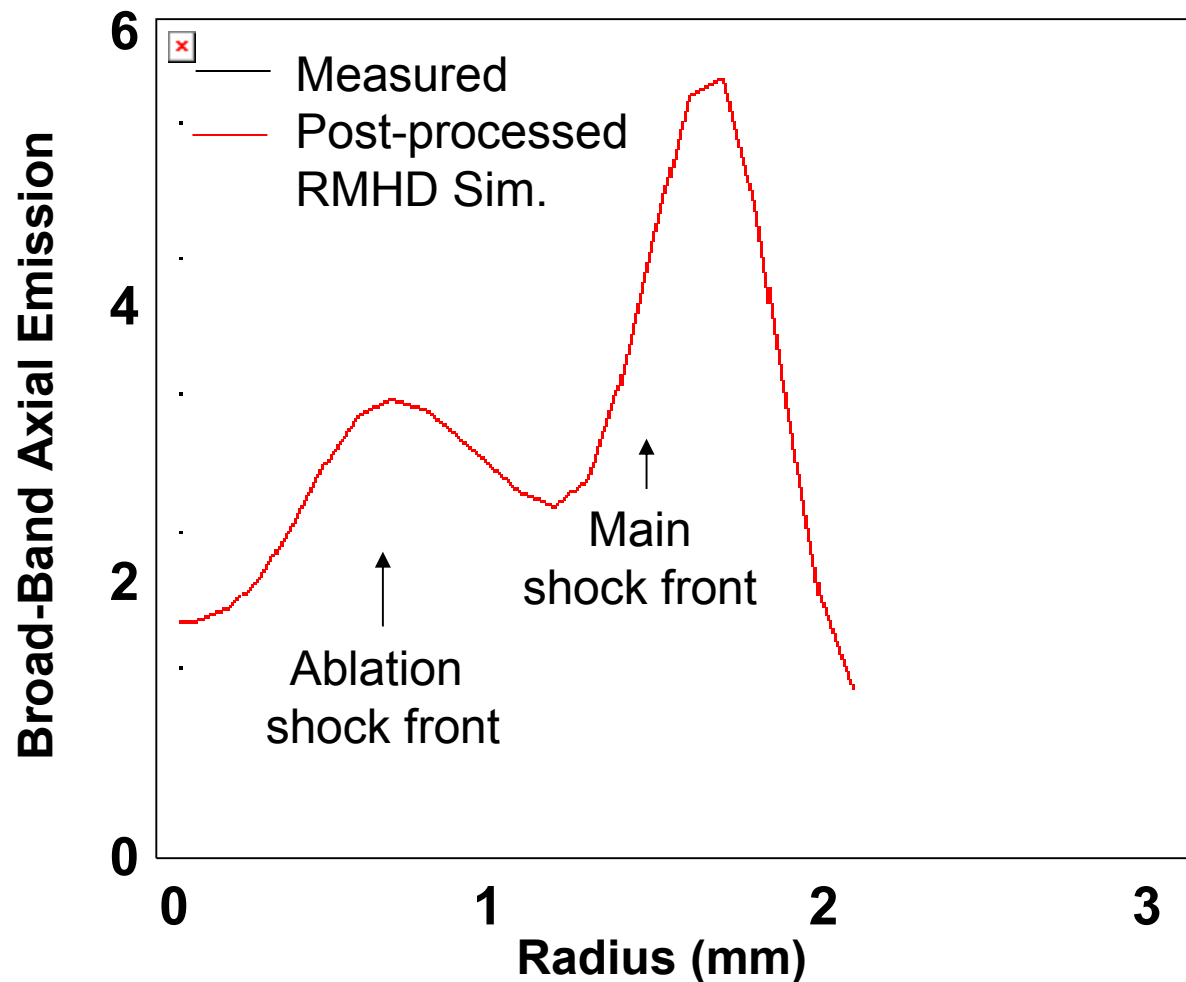
2-D RMHD calculations predict the existence of an ablatively driven shock ahead of the main compression shock.



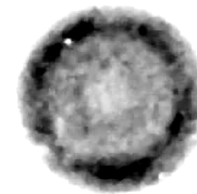


2-D shock emission data support the presence of an ablatively driven shock ahead of the main shock.

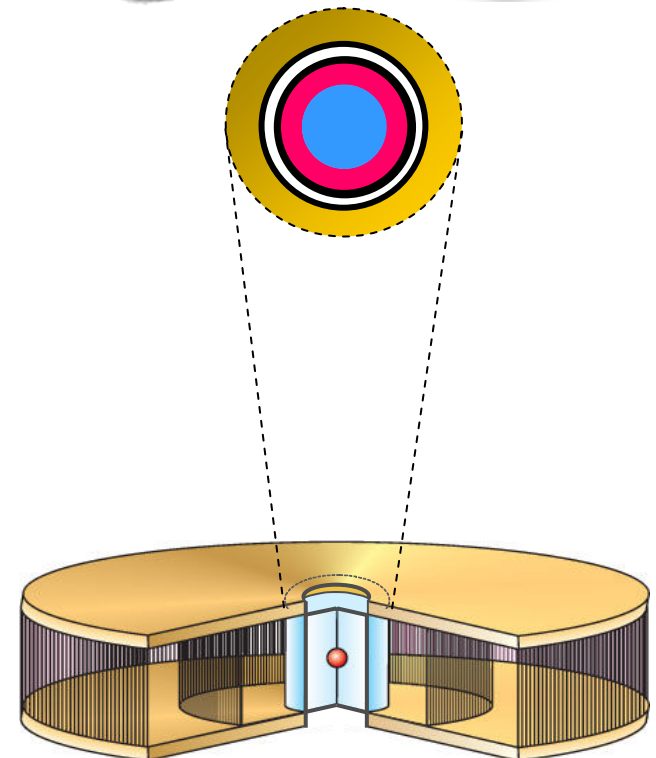
Azimuthally Averaged Radial Lineout of Shock Emission at $h\nu \sim 200\text{-}300\text{eV}$



Measured Image



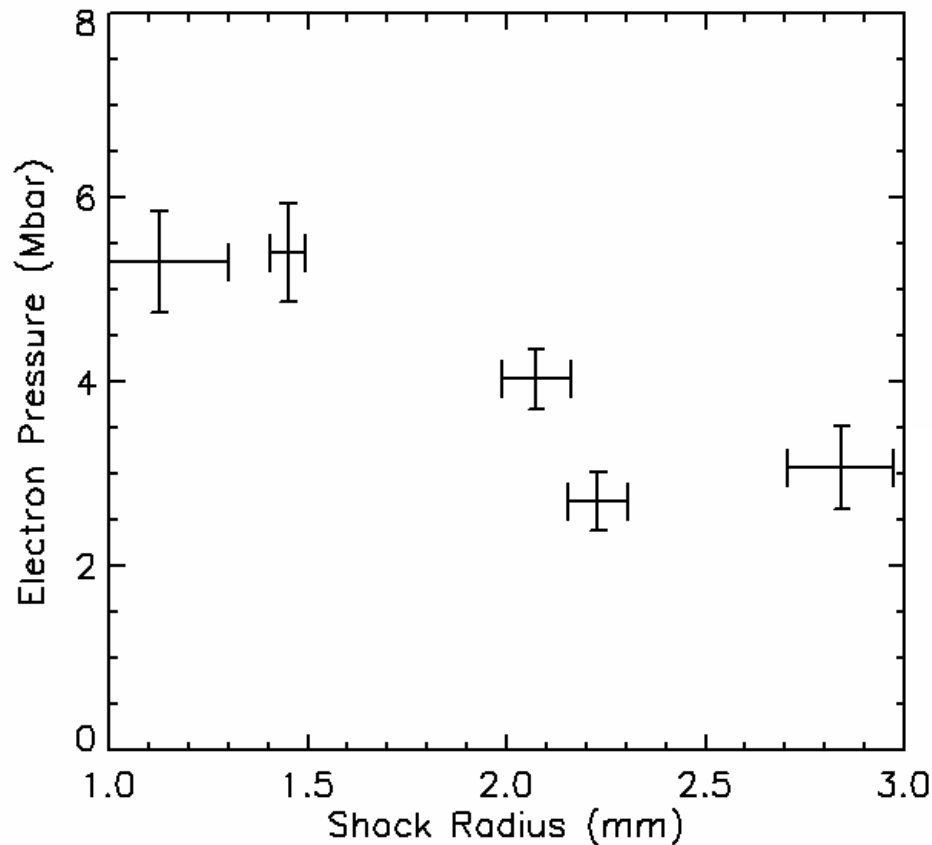
Post-Processed Image



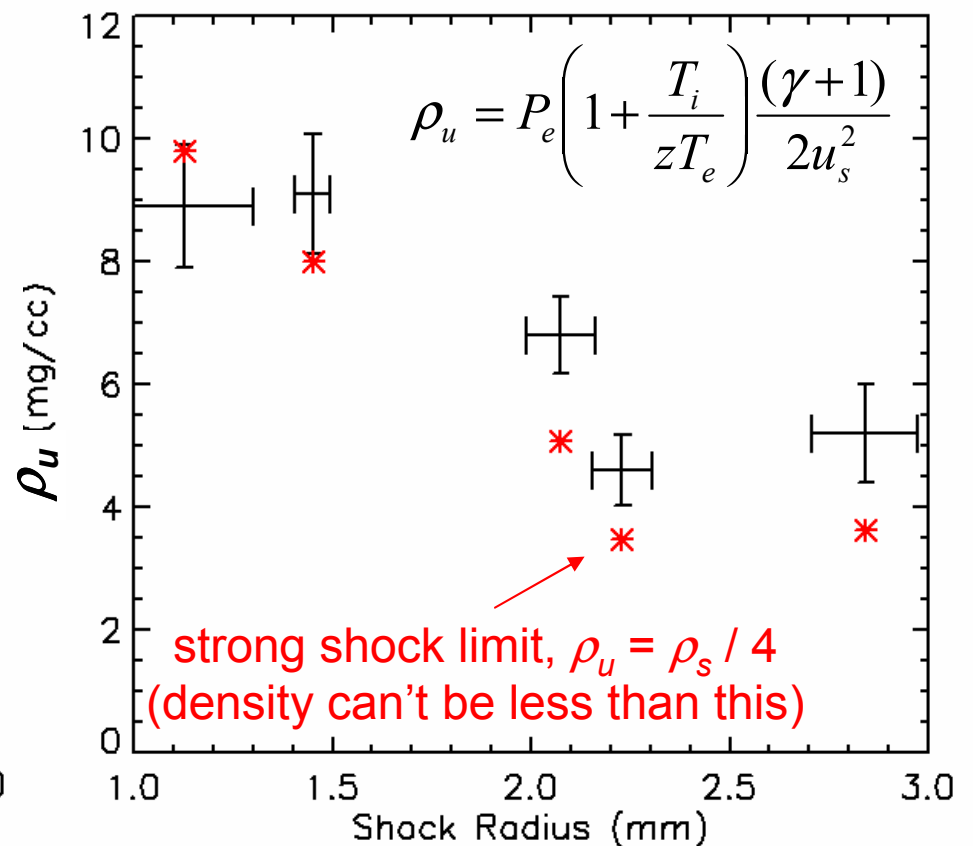


The electron pressure and shock velocity provide a **LOWER BOUND** on the density ahead of the main shock front.

**Electron Pressure Behind Main Shock
(determined from Si measurements)**

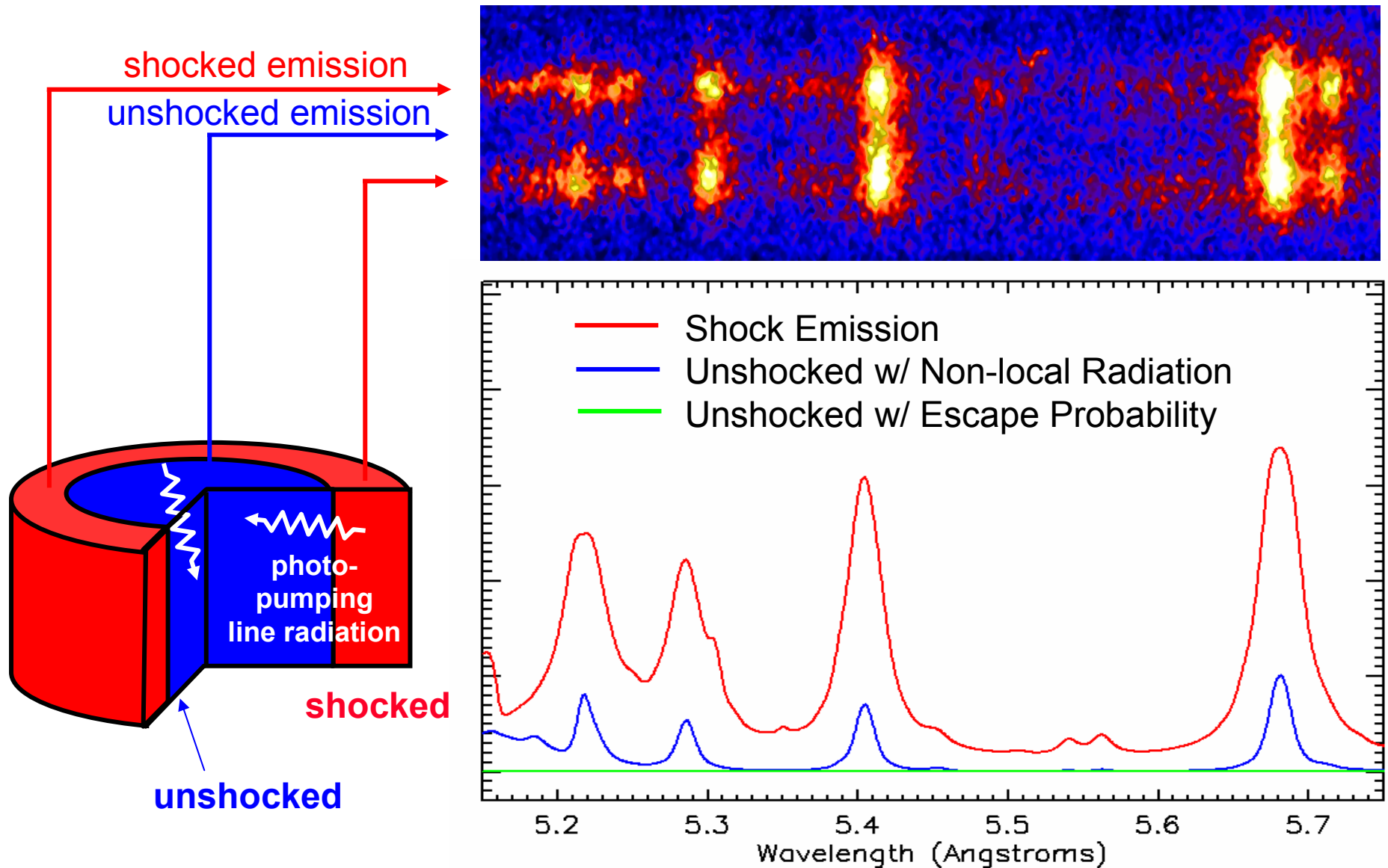


**Density Ahead of Main Shock Front
Assuming $T_i = T_e$**





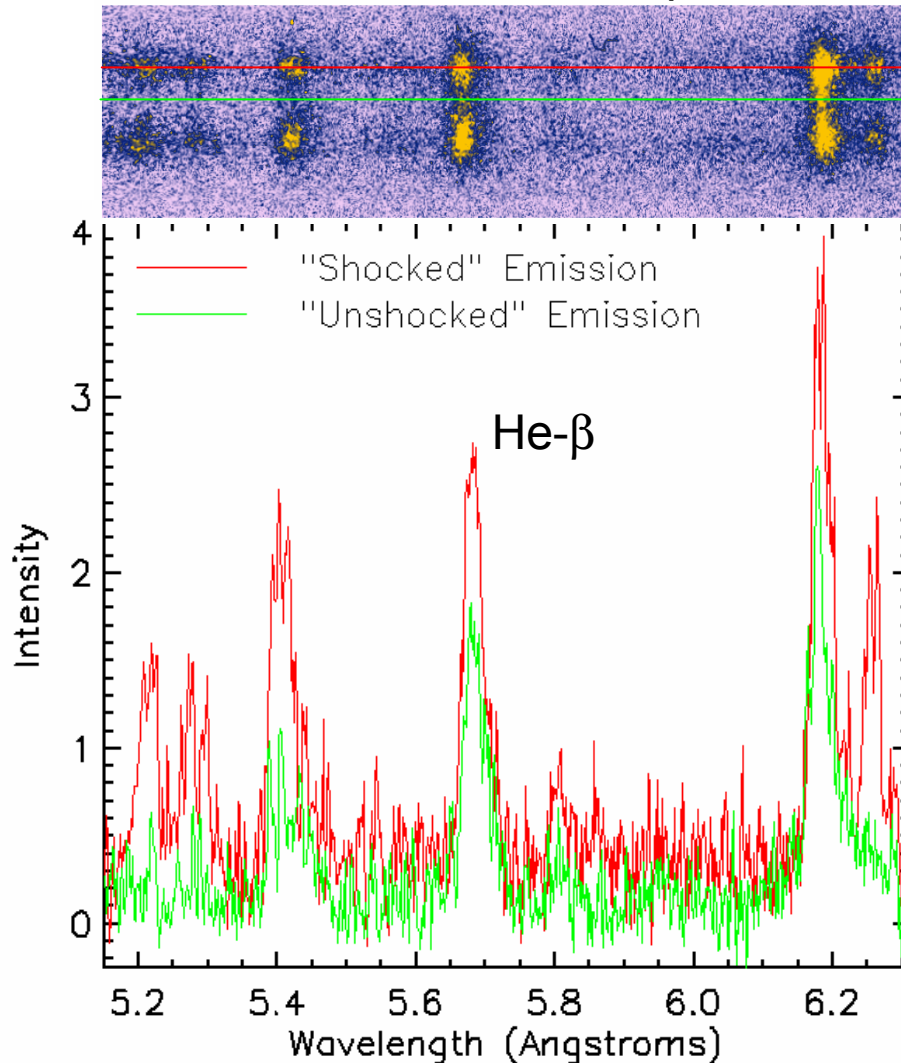
At late times ($r_s < 1.5$ mm), photo-pumped line emission can be used to determine conditions ahead of the shock.



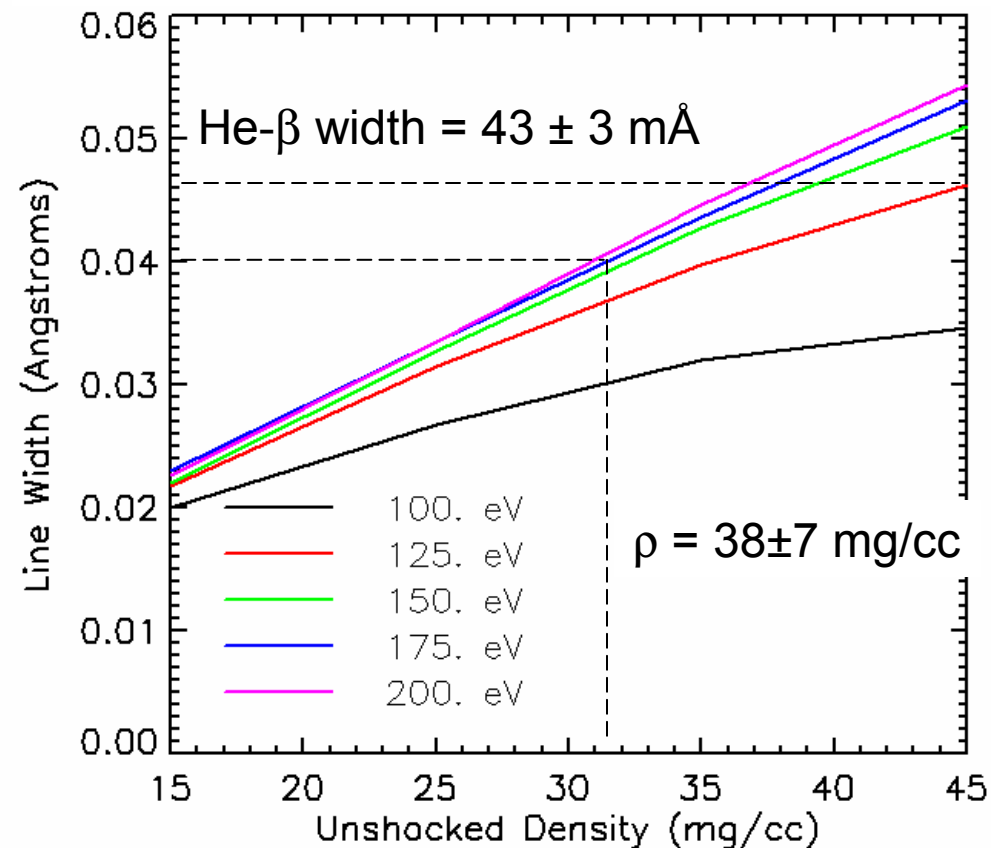


At $r_s \sim 1.5$ mm, the measured emission ahead of the main shock indicates a colder plasma of similar density.

Measured Si Emission Spectrum



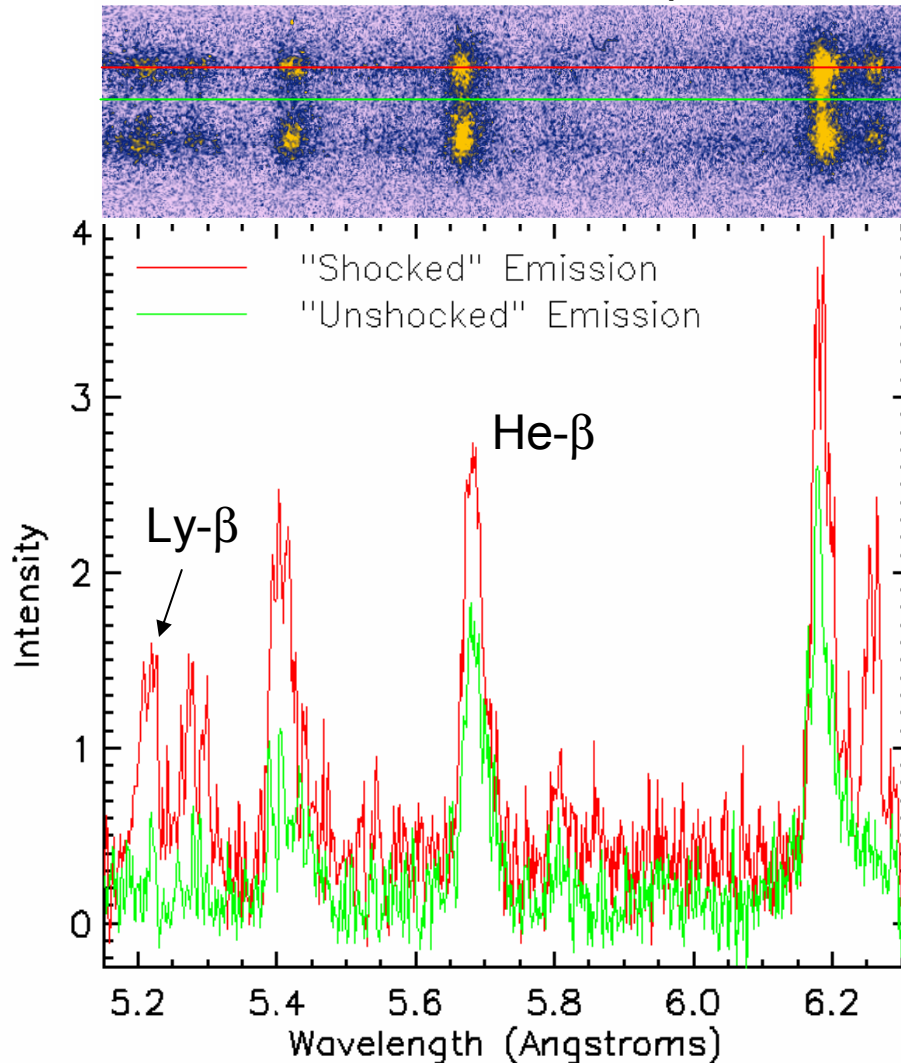
Spect3D Simulation of Photo-Pumped
He- β Line Width vs. Density
at $T_s = 310$ eV, $\rho_s = 32$ mg/cc



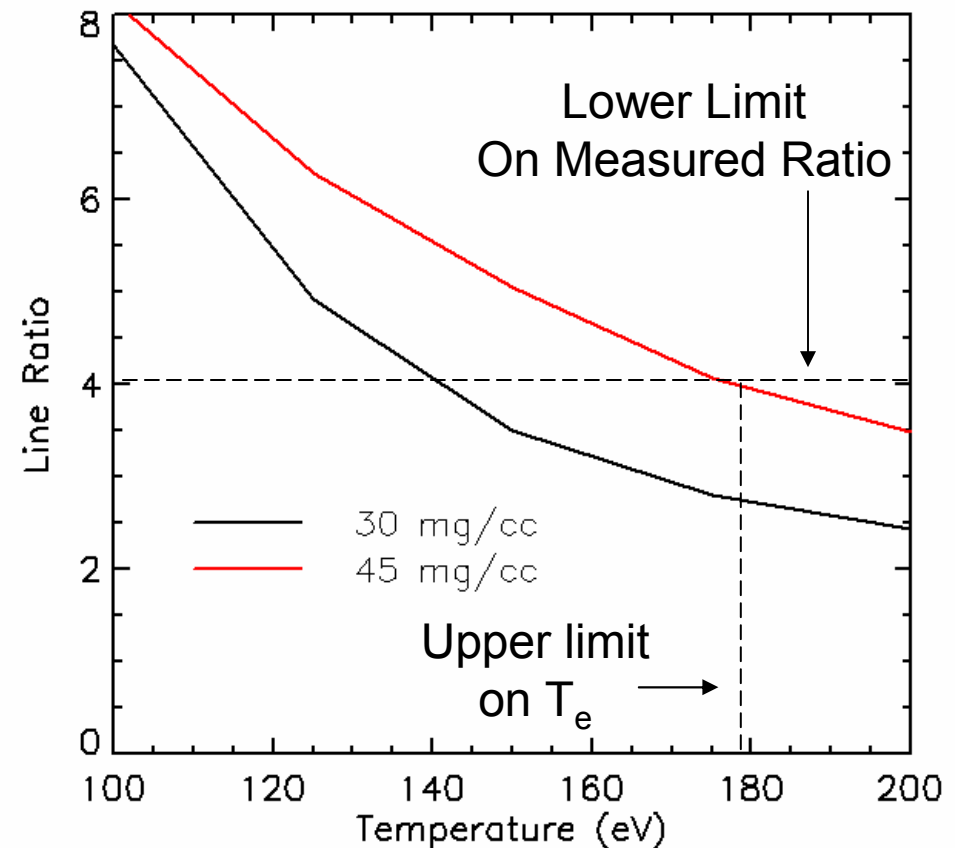


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Measured Si Emission Spectrum



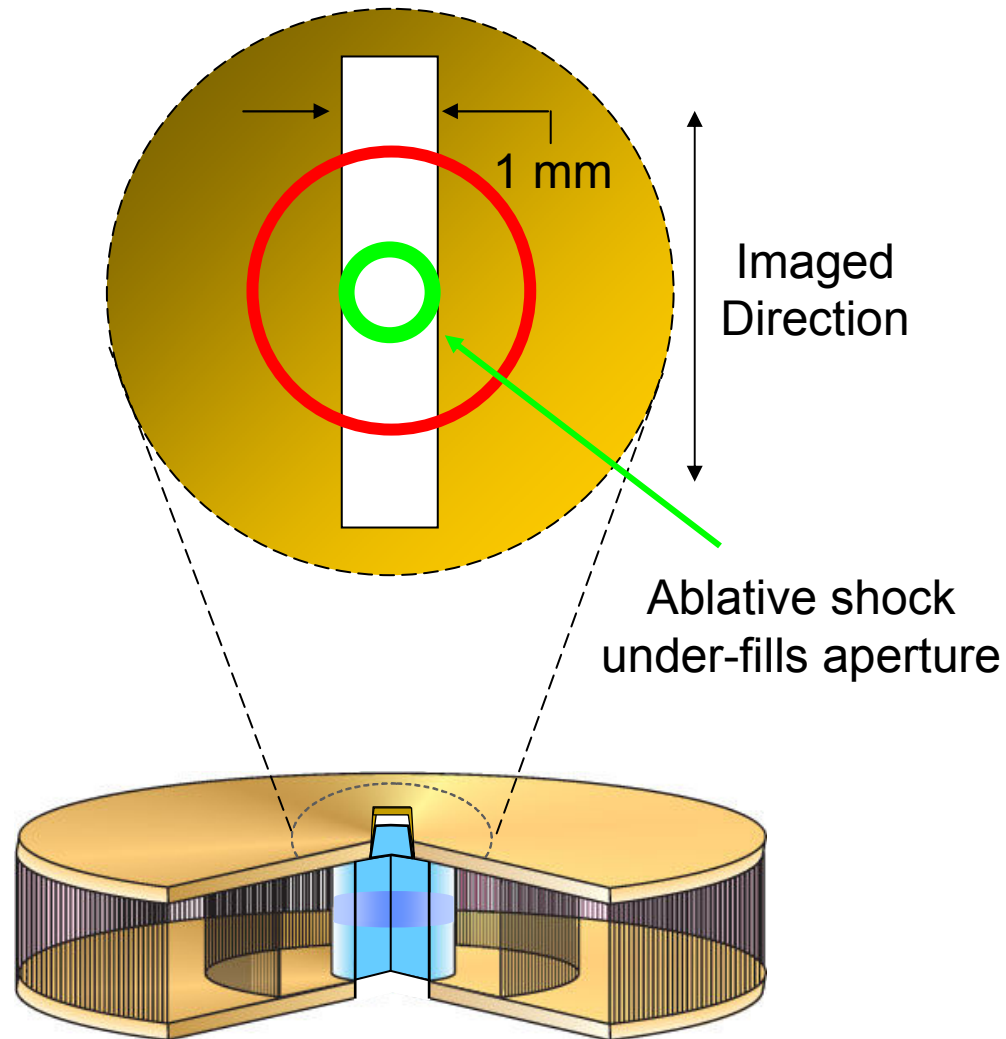
Spect3D Simulation of Photo-Pumped He- β / Ly- β Line Ratio vs. Temperature
at $T_s = 310$ eV, $\rho_s = 32$ mg/cc



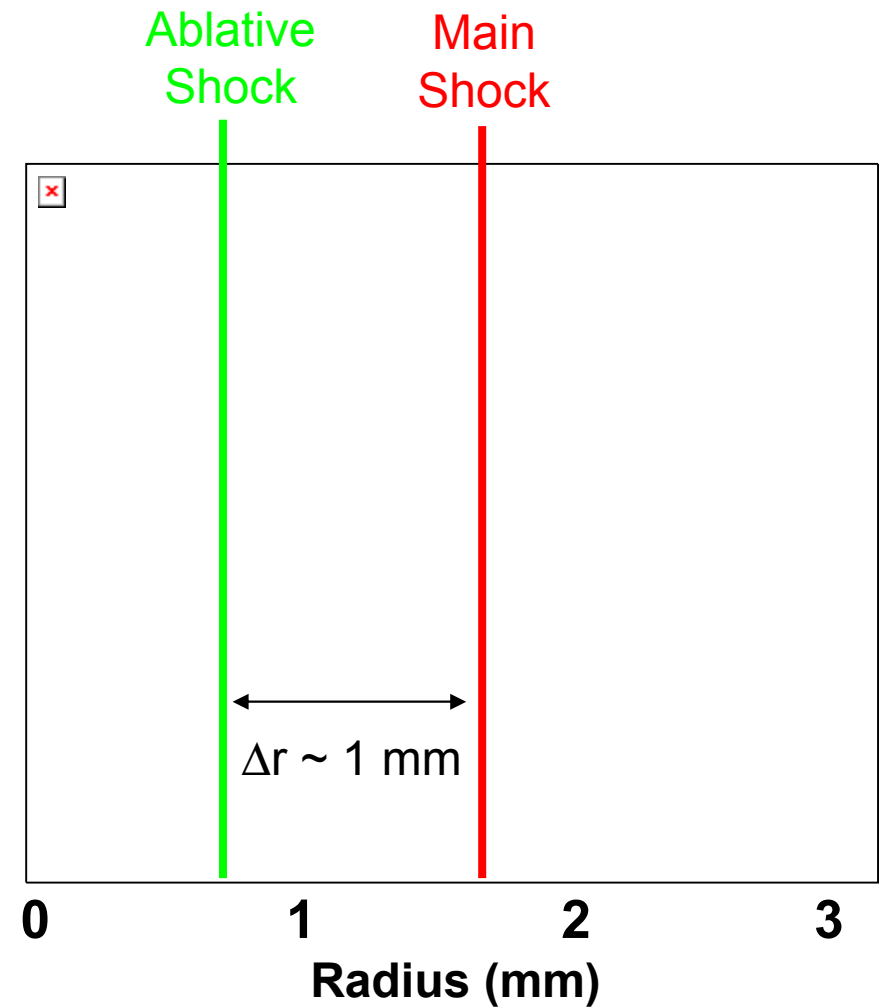


At $r_s < 1.5$ mm, Si emission ahead of the main shock is likely dominated by ablatively shocked plasma.

Spectrometer View

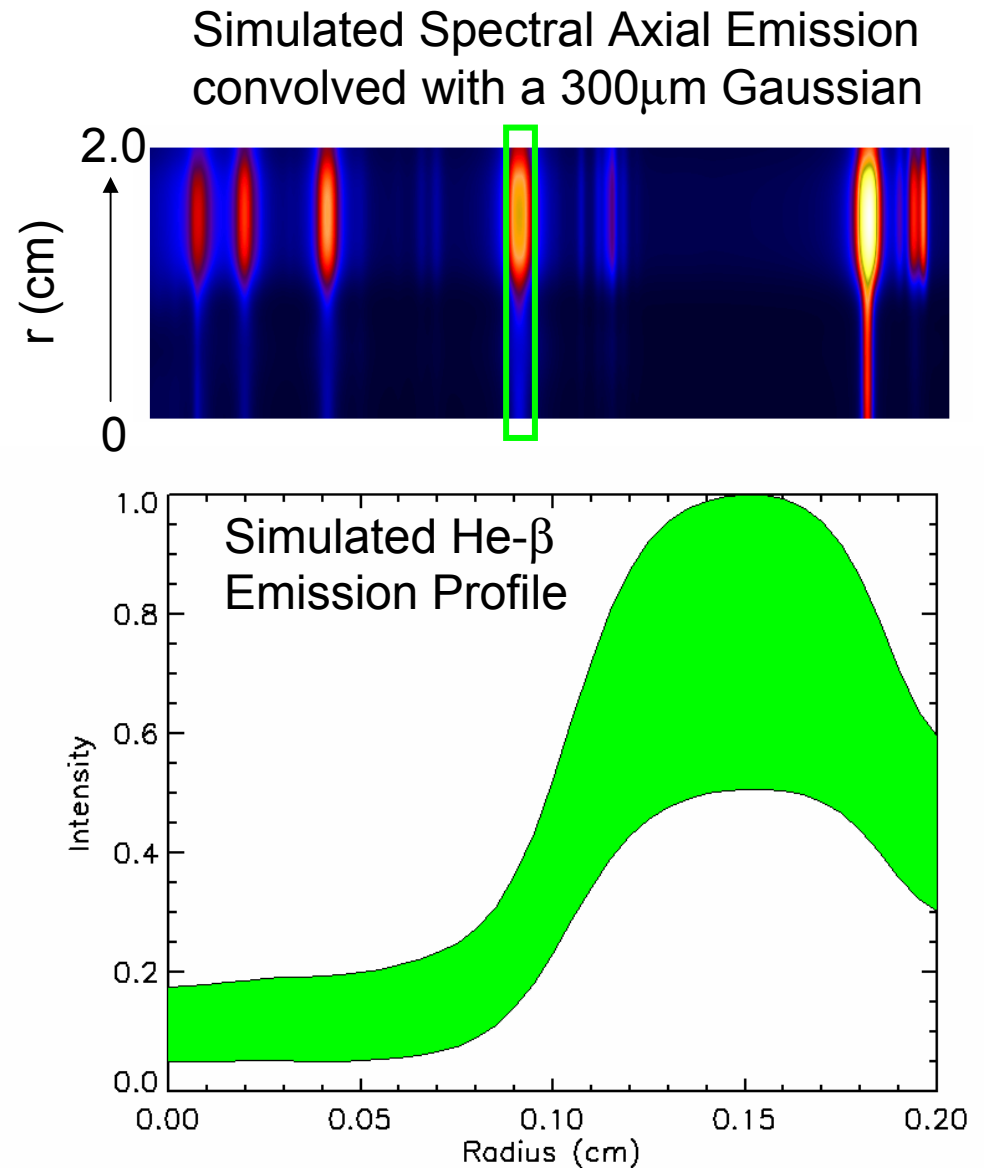
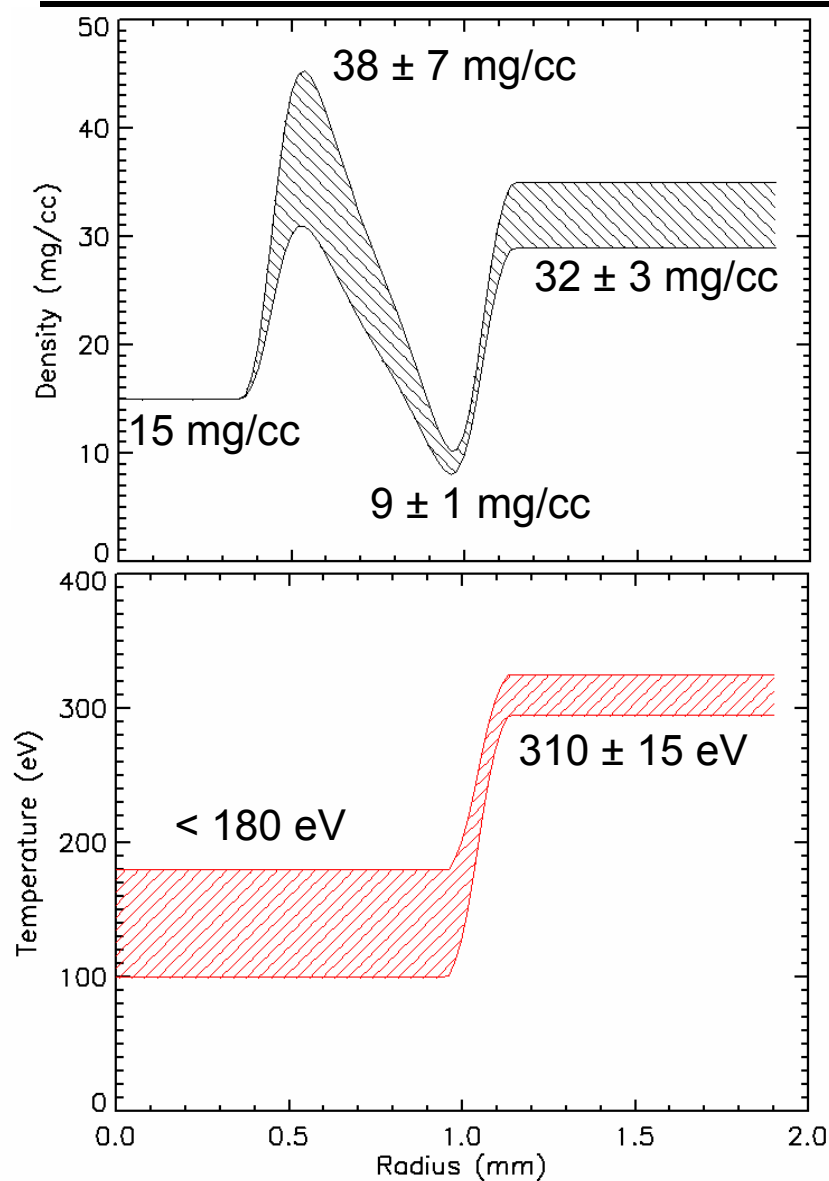


2-D imaged intensity profile





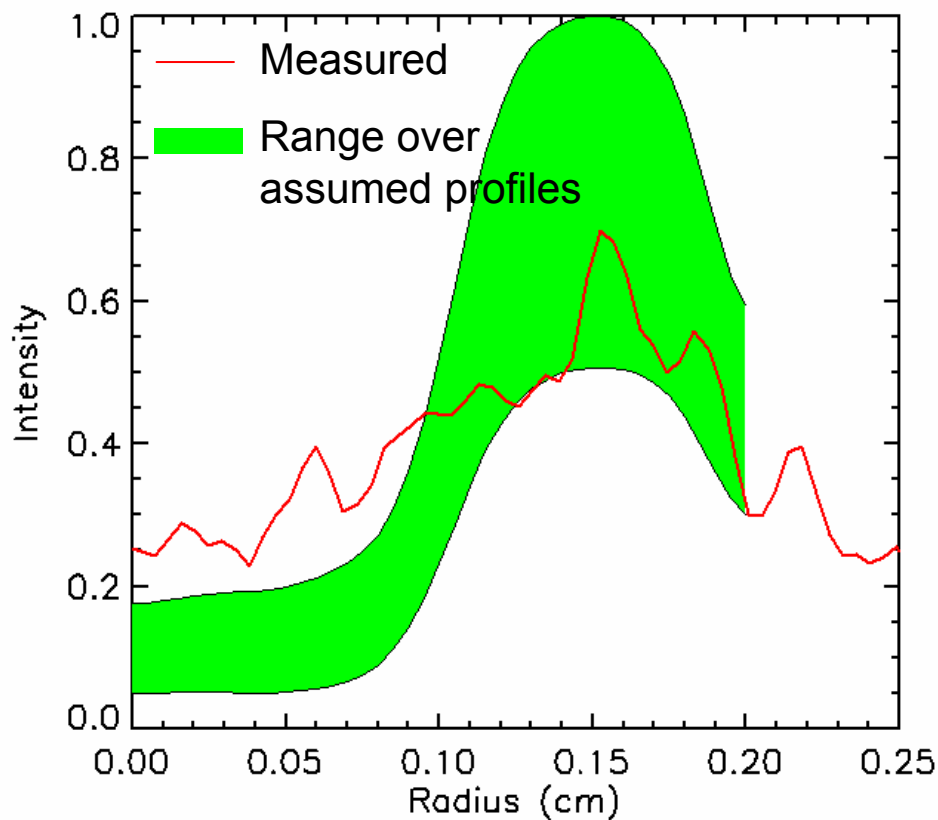
All inferred quantities can be used to create density and temperature profiles to simulate the spatial emission distribution.



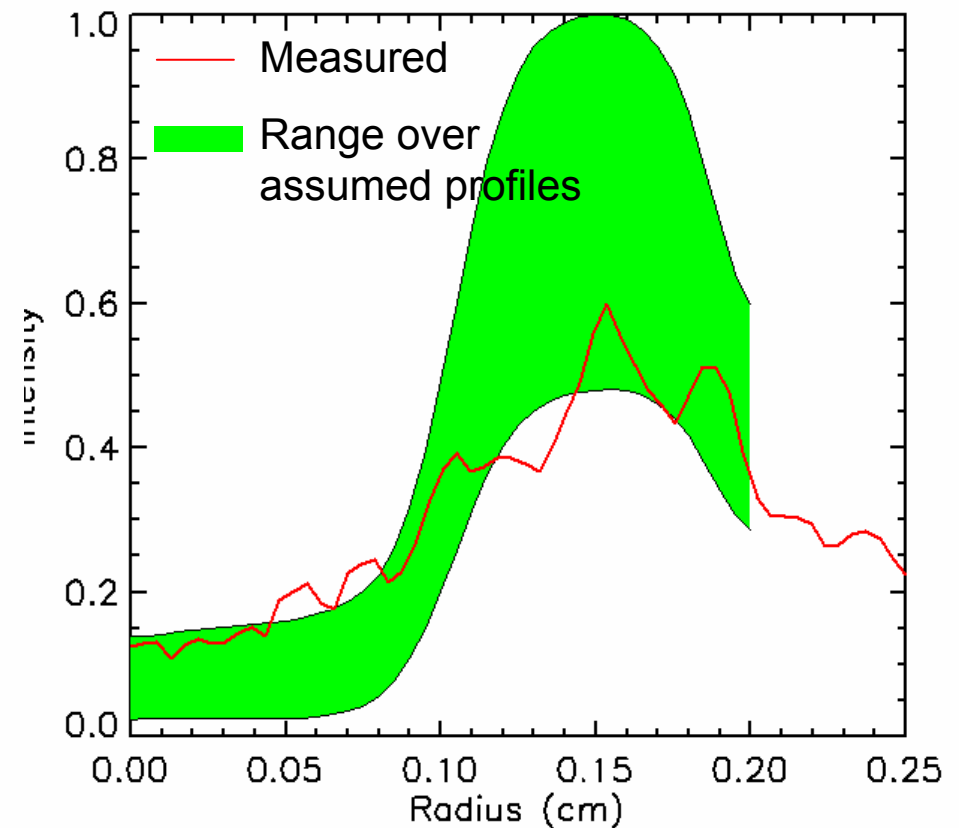


The photo-pumping efficiency is slightly under-predicted for the He- β line, but in fair agreement for the He- γ .

He- β intensity profile



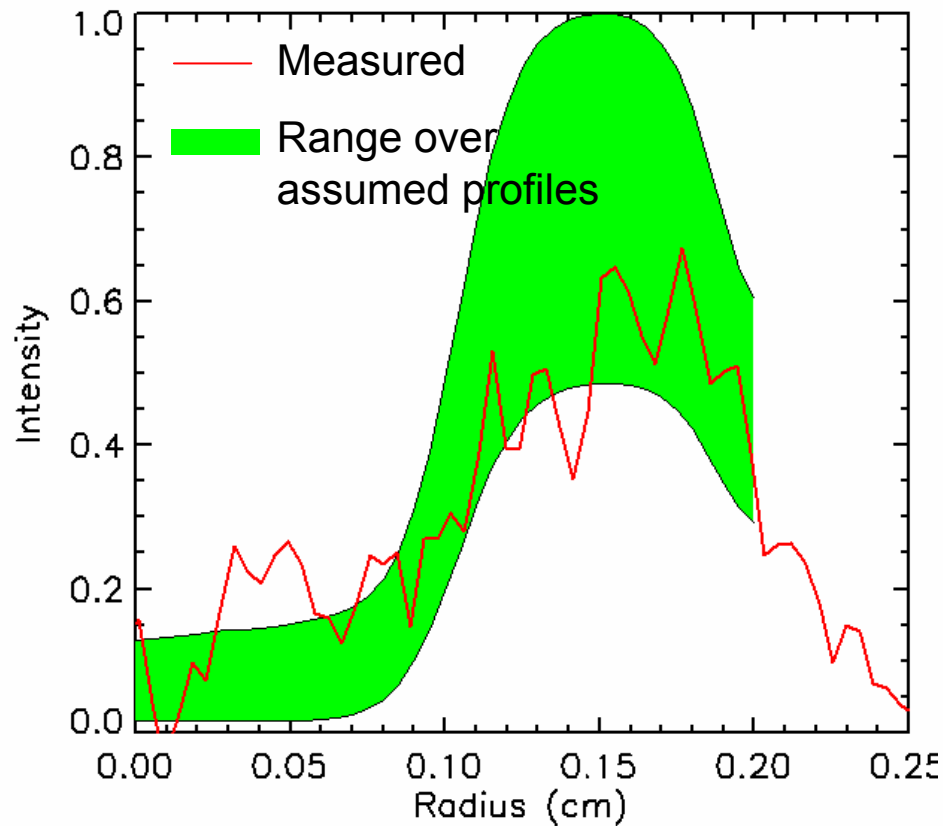
He- γ intensity profile



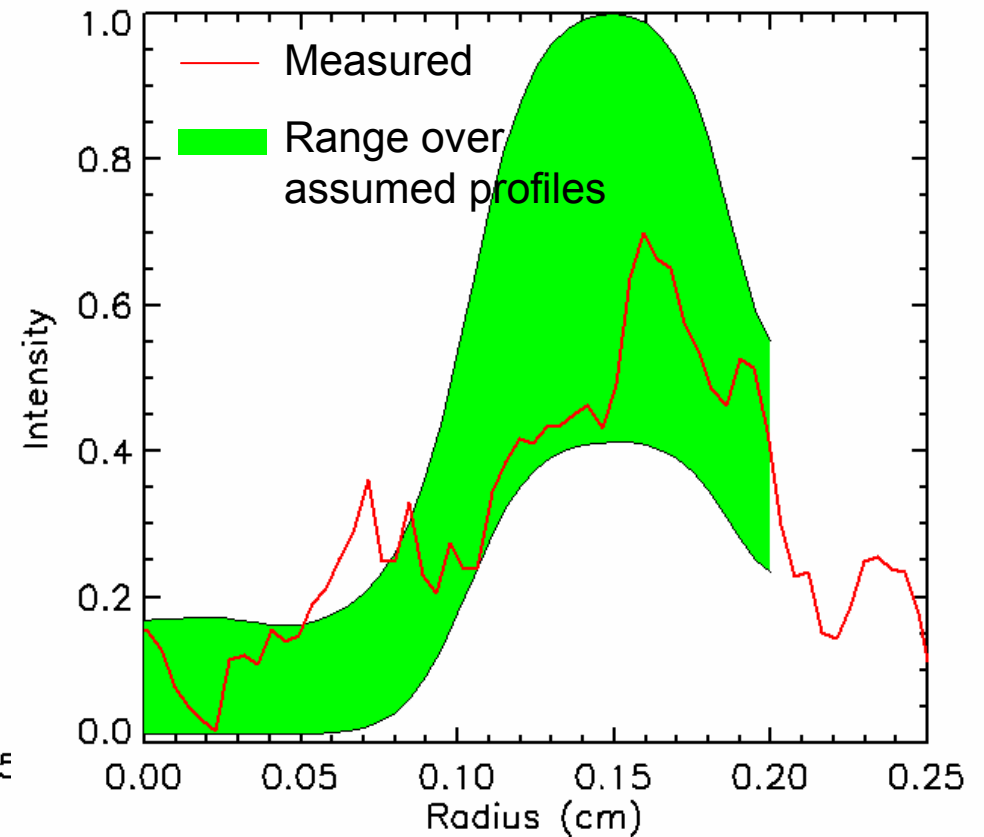


The photo-pumping efficiency is in fair agreement for both the He- δ and the Ly- β .

He- δ intensity profile

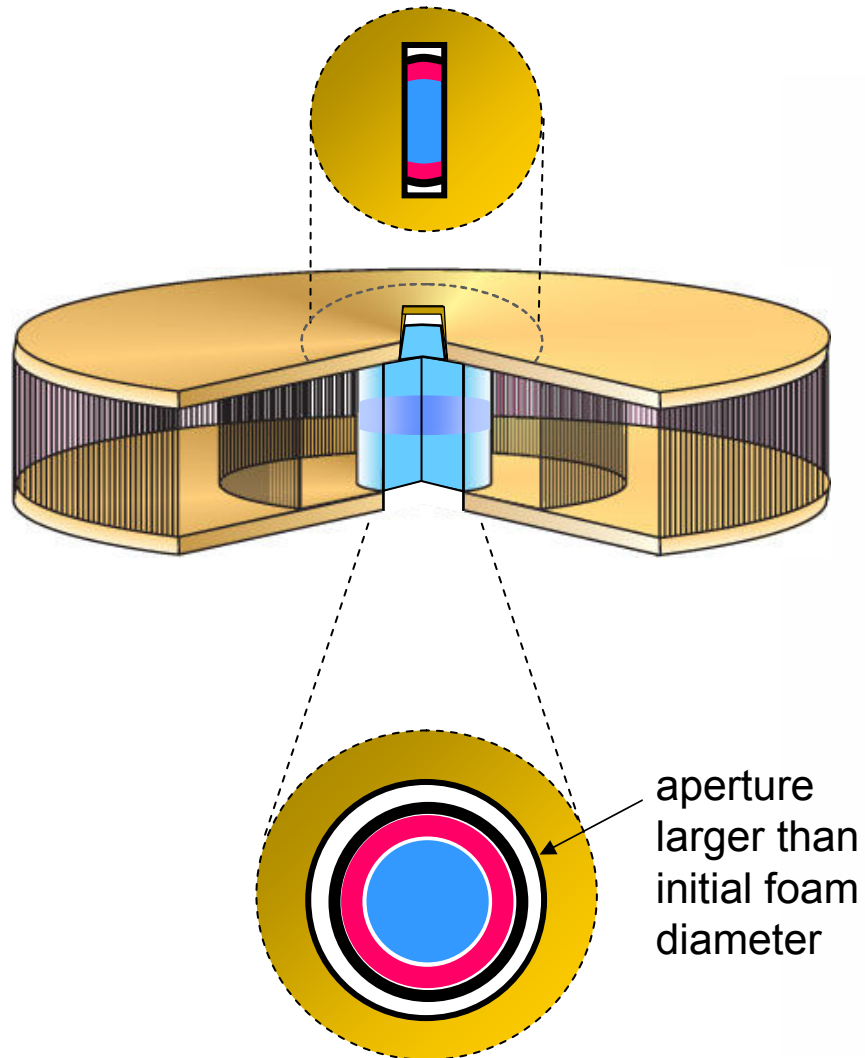


Ly- β intensity profile

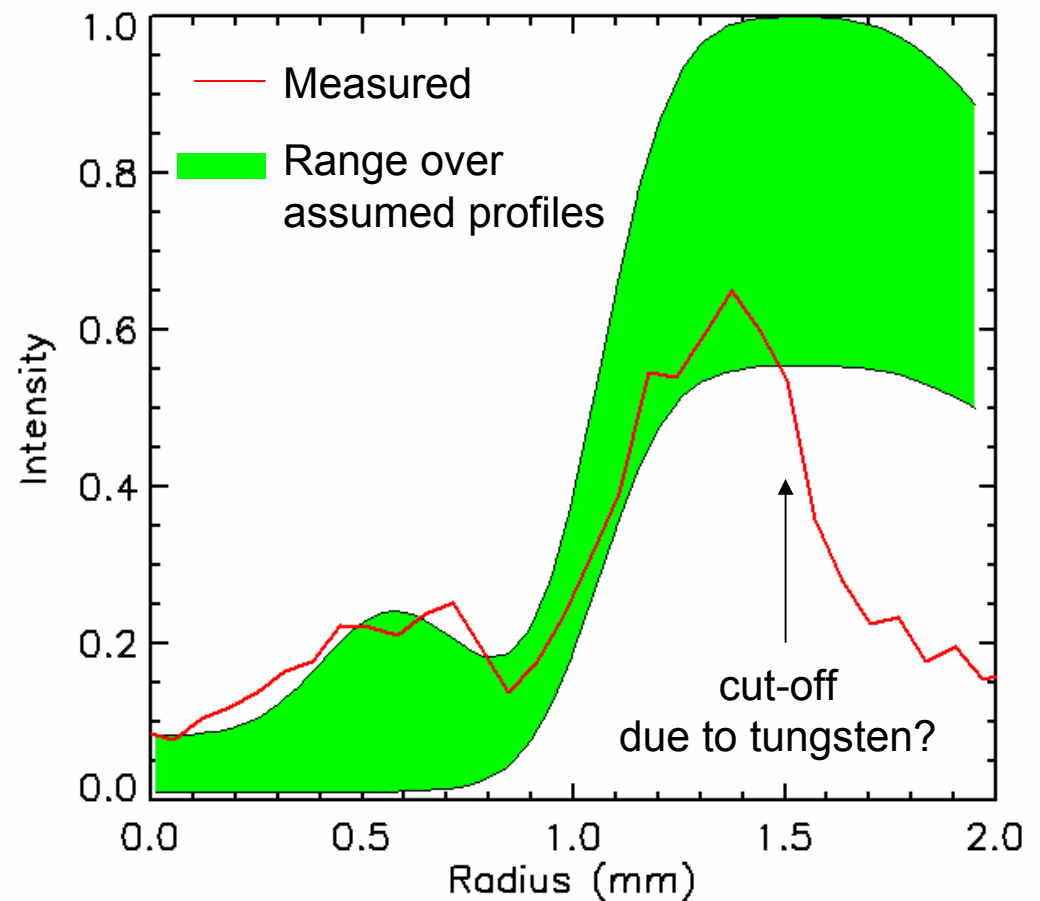




The broadband emission profile from x-ray pinhole camera data is in good agreement up to the emission peak.



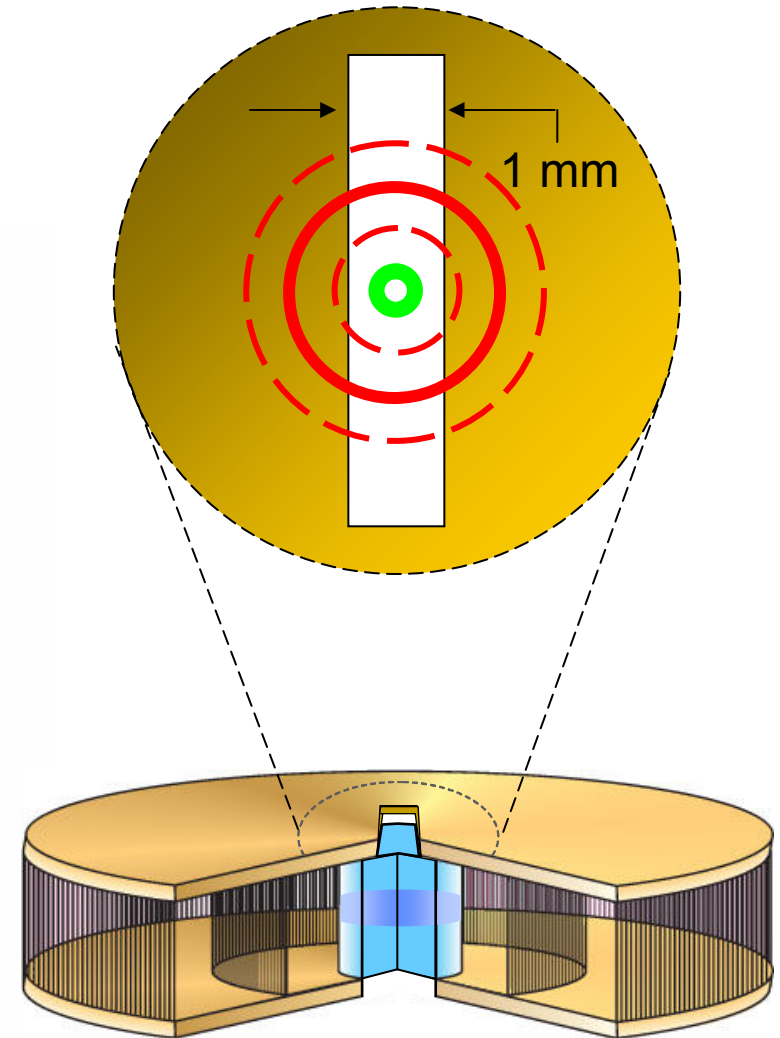
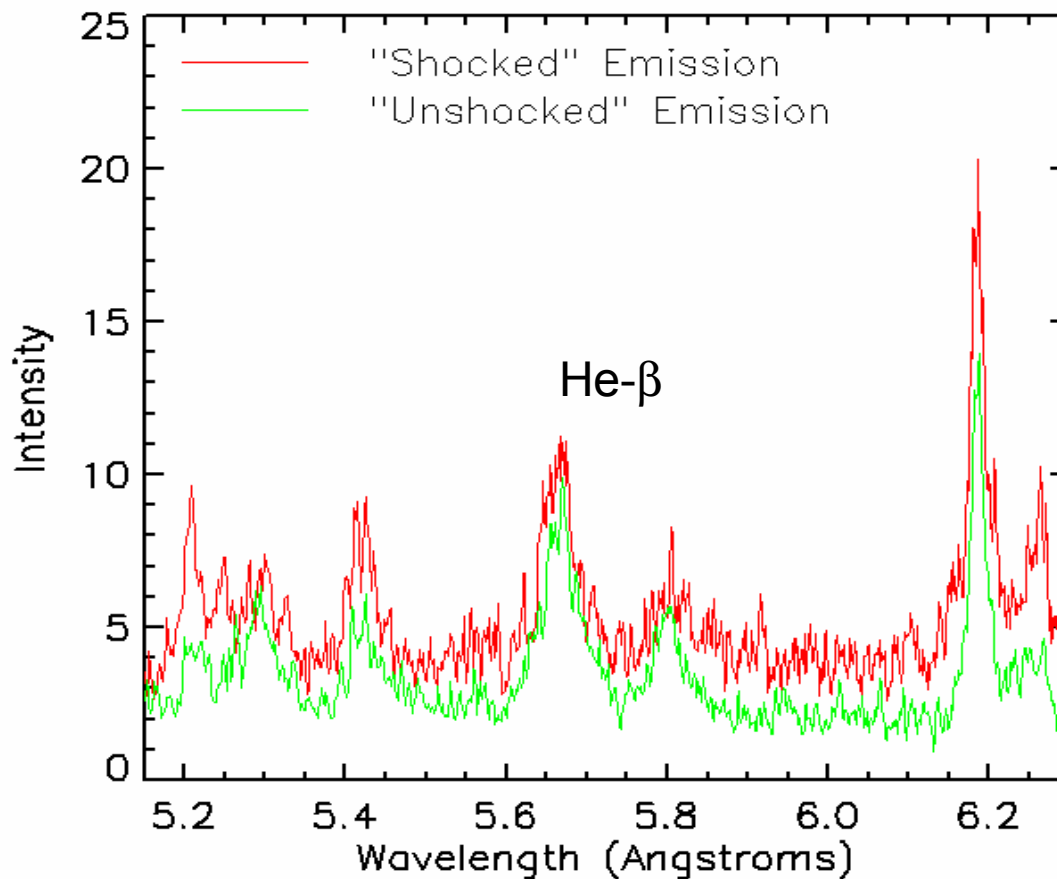
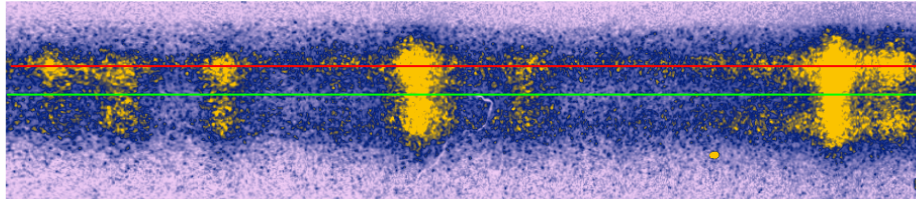
Broadband Axial Emission Profile
 $h\nu > 1000$ eV





At $r_s \sim 1.1$ mm, the spectra are nearly homogeneous across the measured geometry.

Measured Si Emission Spectrum





Summary

- A strong radiating shock is the source of thermal energy for the Z-pinch Dynamic Hohlraum.
- The temperature and density of this shock have been measured through the interpretation of time- and space-resolved Si emission spectra from Si atoms doped across the central 3 mm height of the 12 mm tall DH.
- These measurements indicate a shock temperature decreasing from 400 - 250 eV, a shock density increasing from 15 - 40 mg/cc, and a shock pressure increasing from 4 – 8 Mbar.
- Once the main shock conditions are determined, interpretation of the photo-pumped Si emission spectra ahead of the main shock provides information on the density and temperature of the ablative shock.
- The measured conditions from both shocks can be used to infer the radial profile of the temperature and density conditions in the interior of the DH.