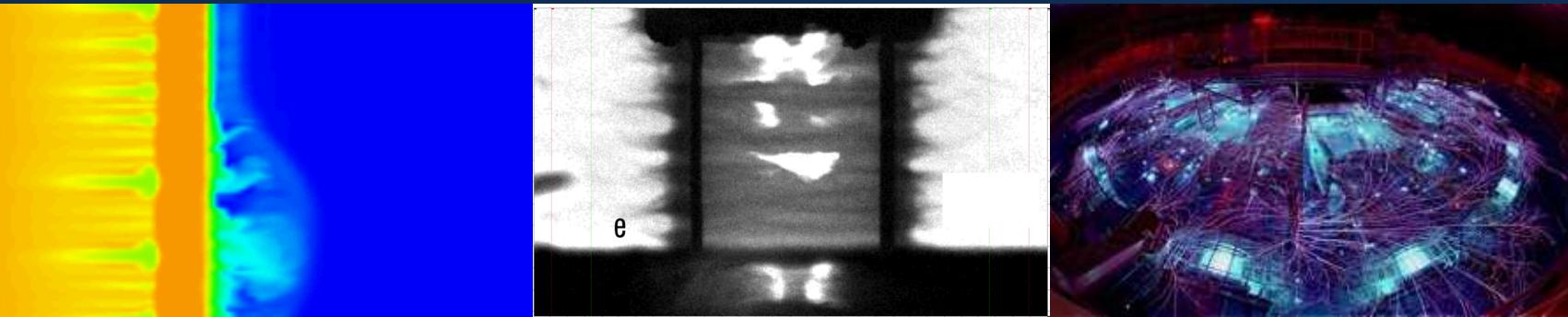


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# Progress in Pulsed Magneto-Inertial Fusion



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**PPC 2015**

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# This presentation represents the work of many scientists and engineers working on the MagLIF project



T.J. Awe, C.J. Bourdon, E.M. Campbell, G.A. Chandler, P.J. Christenson, M.E. Cuneo, M. Geissel, M.R. Gomez, K.D. Hahn, S.B. Hansen, E.C. Harding, A.J. Harvey-Thompson, M.C. Herrmann, M.H. Hess, C.A. Jennings, B. Jones, M. Jones, R.J. Kaye, P.F. Knapp, D.C. Lamppa, J.S. Lash, M.R. Lopez, M.R. Martin, R.D. McBride, L.A. McPherson, T.N. Nagayama, K.J. Peterson, J.L. Porter, G.A. Rochau, D.C. Rovang, C.L. Ruiz, S.E. Rosenthal, M.E. Savage, P.F. Schmit, A.B. Sefkow, D.B. Sinars, S.A. Slutz, I.C. Smith, W.A. Stygar, R.A. Vesey, E.P. Yu

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B.E. Blue, D.G. Schroen, K. Tomlinson, M.S. Wei

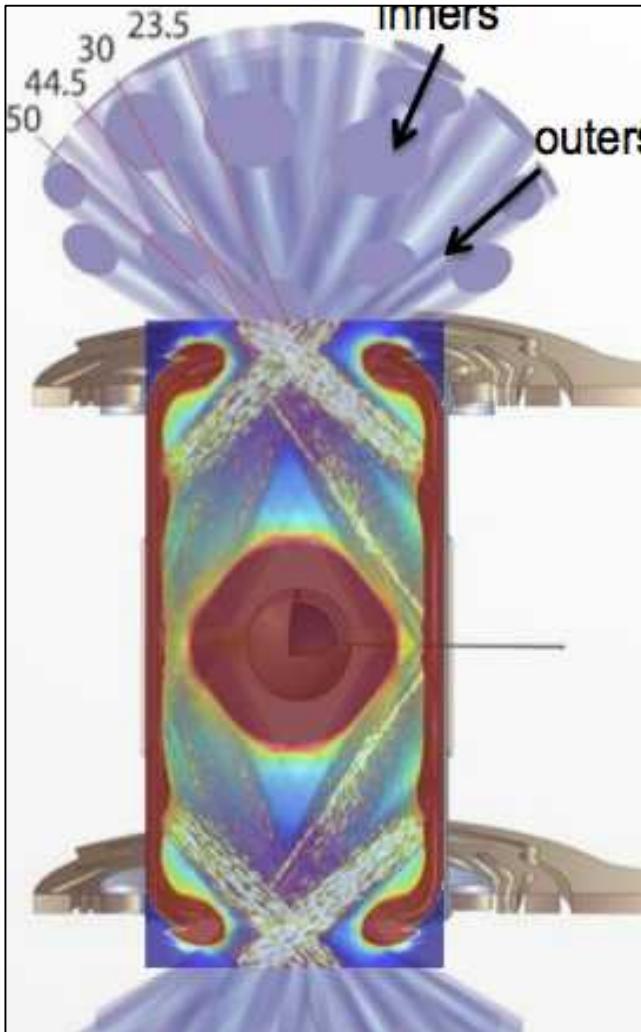
*General Atomics, San Diego, CA 92186 USA*

J. Davies, R. Betti, P.-Y. Chang, G. Fiksel, D. Barnak, V. Glebov

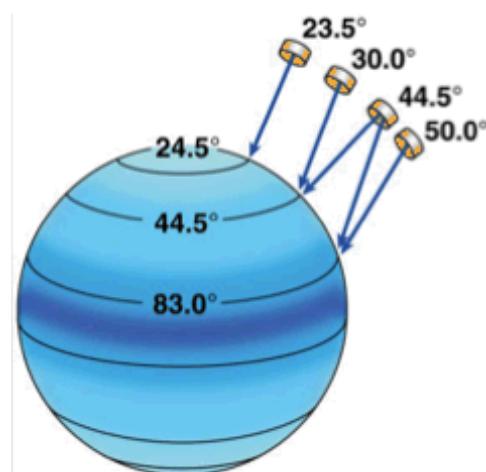
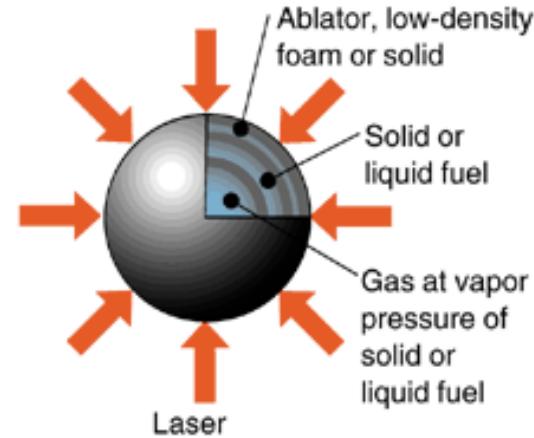
*Laboratory for Laser Energetics, University of Rochester, Rochester, NY, USA*

# The United States ICF program is pursuing three main approaches to ignition

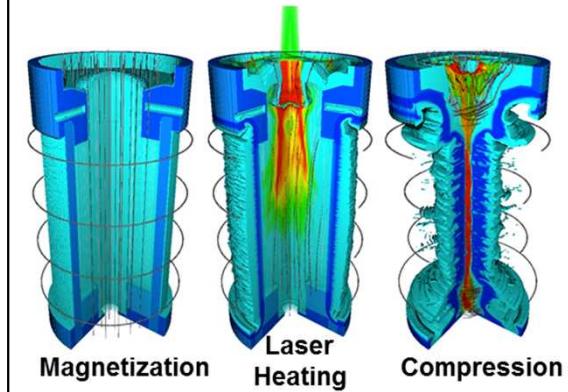
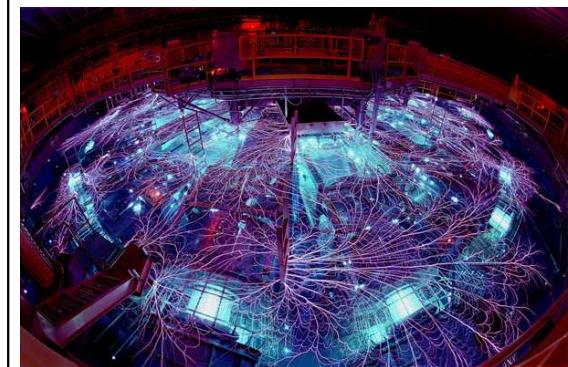
## Radiation-driven implosions



## Laser-driven implosions

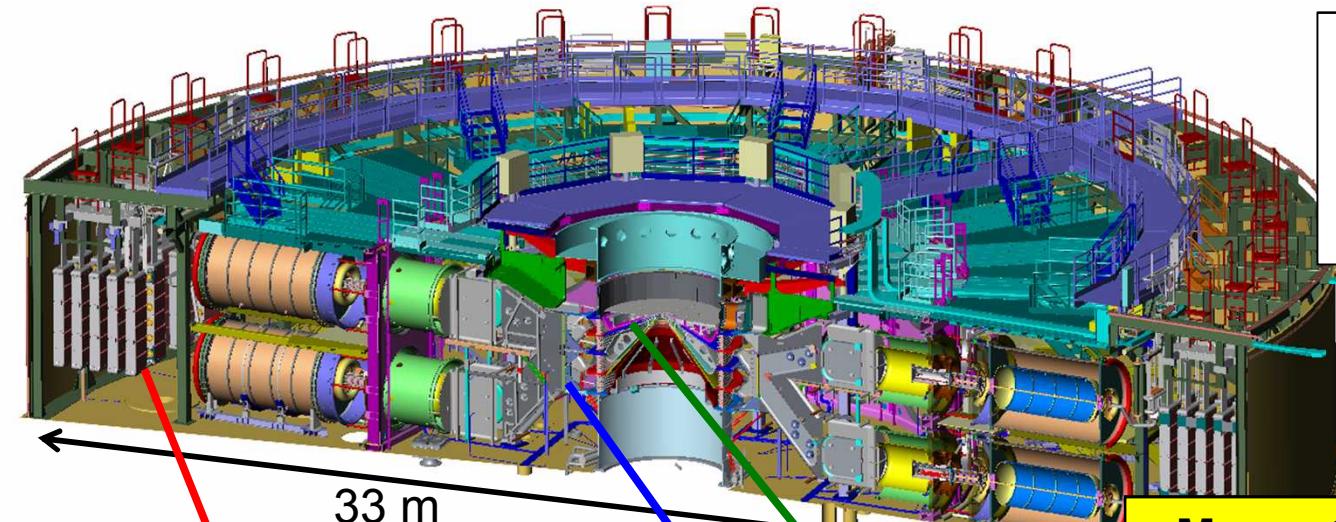


## Magnetically-driven implosions

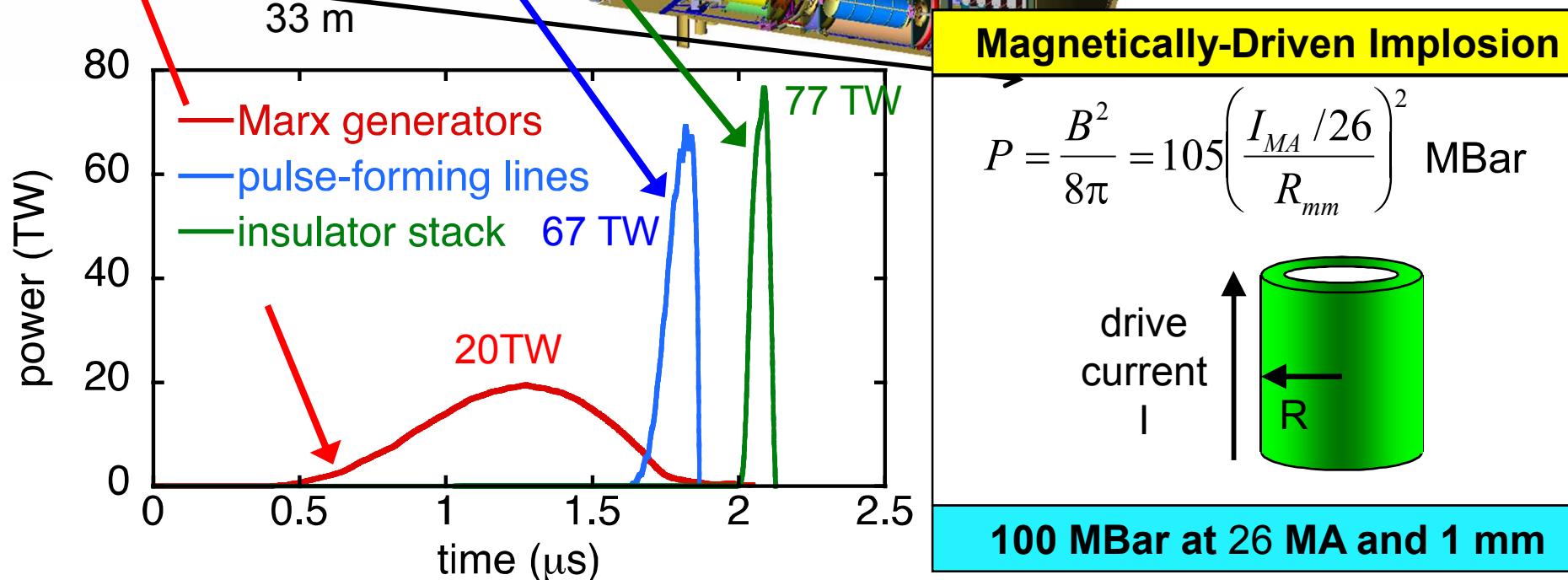


Focus of today's talk

“Magnetic direct drive” is based on the idea that we can efficiently use large currents to create high pressures



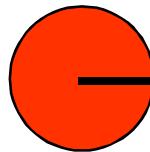
Z today couples  $\sim 0.5$  MJ out of 20 MJ stored to MagLIF target (0.1 MJ in DD fuel).



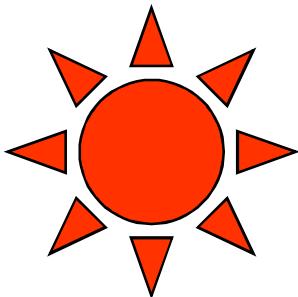
Implosion time  $\sim 50$  ns; stagnation  $\sim 0.1$ -1 ns

(1 atm = 1 bar =  $10^5$  Pascals)

# Under extreme conditions a mass of DT can undergo significant thermonuclear fusion before falling apart



$\rho, R, T$



- Consider a mass of DT with radius  $R$ , density  $\rho$ , and temperature  $T$
- How does the disassembly time compare with the time for thermonuclear burn?

$$\tau_{disassembly} \sim \frac{R}{c_s} \sim \frac{R}{\sqrt{T}}$$

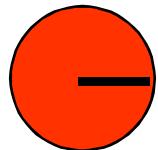
$$\tau_{burn} \sim \frac{1}{n_i \langle \sigma v \rangle} \sim \frac{1}{\rho \langle \sigma v \rangle}$$

- The fractional burn up of the DT (for small burn up) is:

$$f_{burn} \approx \frac{\tau_{disassembly}}{\tau_{burn}} \sim \rho R \frac{\langle \sigma v \rangle}{\sqrt{T}}$$

- At sufficiently high  $\rho R$  and  $T$  the fractional burn up becomes significant and the energy deposited by alpha particles greatly exceeds the initial energy in the fusion fuel (“ignition”)
- Typical conditions are:  $\rho R \approx 0.4 \text{ g/cm}^2$   
 $T \approx 5 \text{ keV (50,000,000 K)}$

# For hot spot ignition fusion fuel must be brought to a pressure of a few hundred billion atmospheres



$\rho, R, T$

For ignition conditions:

$$\left\{ \begin{array}{l} \rho R \approx 0.4 \text{ g/cm}^2 \\ T \approx 5 \text{ keV} \end{array} \right\}$$

$$E_{HS} \propto m_{HS} T_{HS} \propto \rho_{HS} R_{HS}^3 T_{HS} \propto \frac{(\rho_{HS} R_{HS})^3 T_{HS}^3}{P_{HS}^2}$$

$$P_{HS}^2 \sim (CR)^6$$
$$\sim (\text{velocity})^6$$

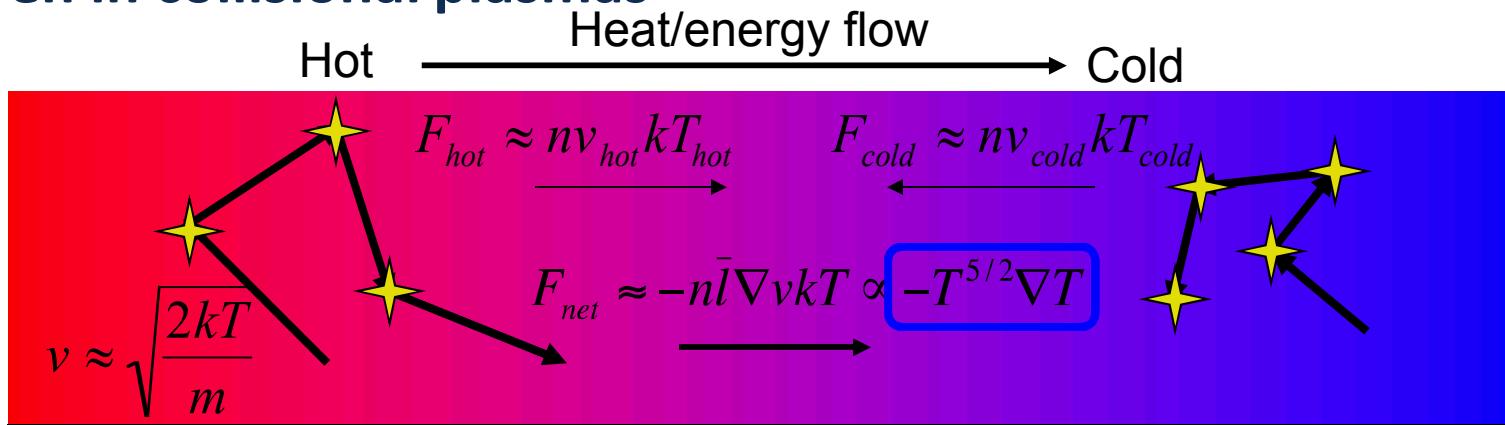
$$E_{NIF} \sim 15 \text{ kJ} \Rightarrow P \sim 400 \text{ GBar} \quad R \sim 30 \mu\text{m} \Rightarrow \text{ and } \rho \sim 130 \text{ g/cm}^3$$

This is consistent with detailed calculations

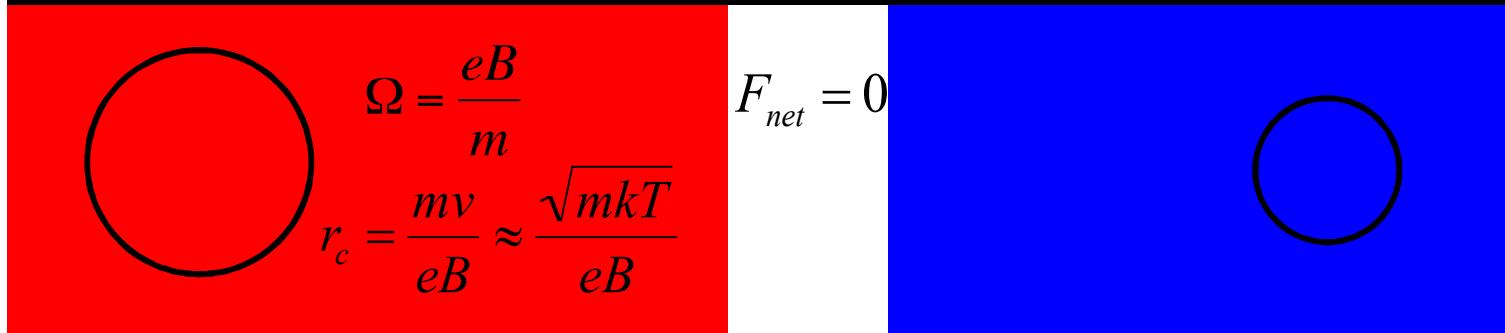
Note: The key challenge for ICF is to make the fuel both **dense** and **hot**. This leads to challenging compression requirements—a NIF capsule has a radial convergence of 35-45x, for a volume compression of ~50,000!

# Magneto-inertial fusion is based on the idea that energy and particle transport can be reduced by strong magnetic fields, even in collisional plasmas

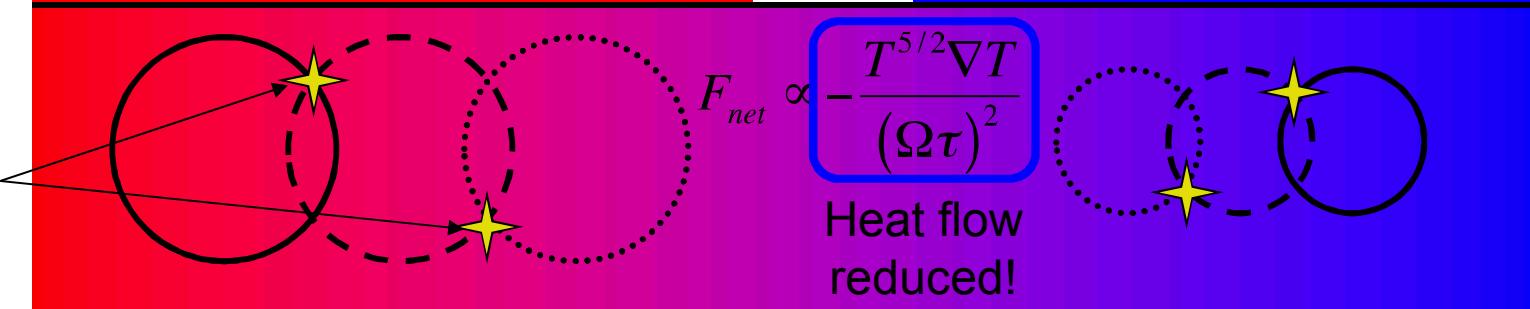
Collisional  
no B



Strong B  
(perpendicular  
to this slide)  
No collisions

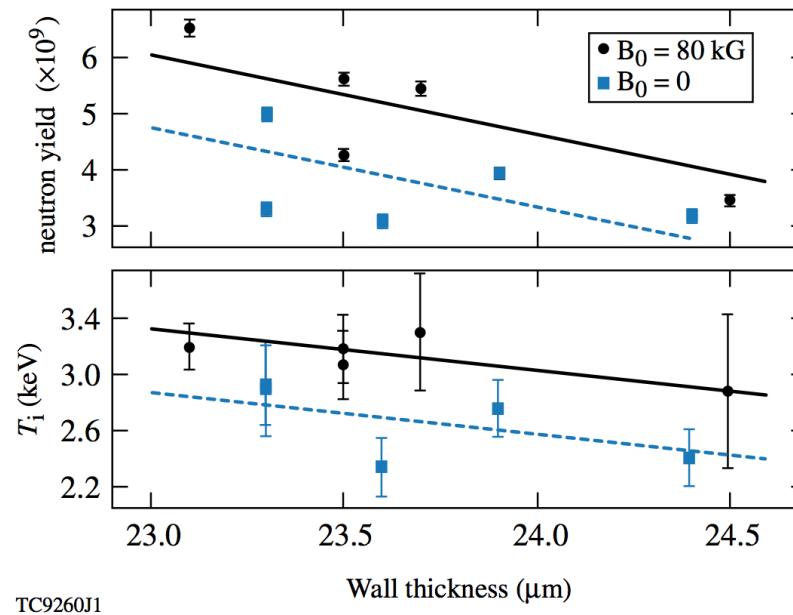
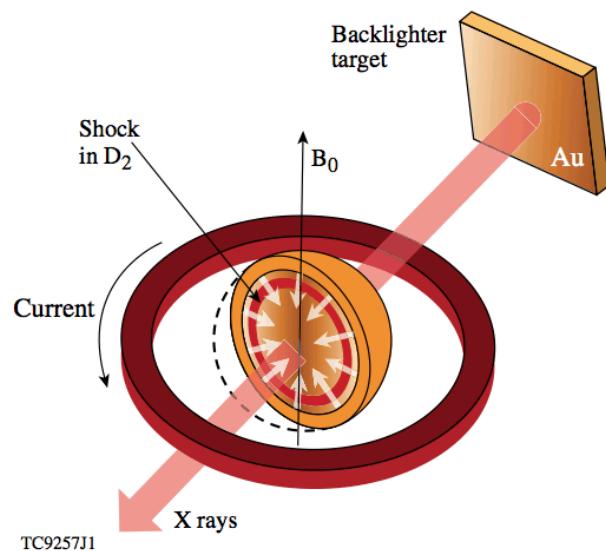


Strong B  
with collisions



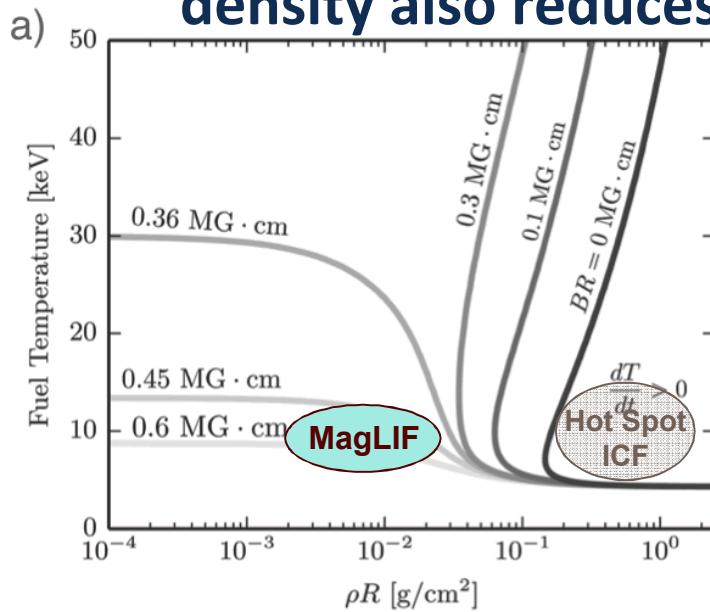
“Anomalous” heat transport can reduce the benefit of magnetic fields (e.g., in tokamaks) but there remains a significant benefit

# Laser-driven spherical capsule implosions at the University of Rochester\* showed clear indicators of higher temperatures (and yields) due to fuel magnetization

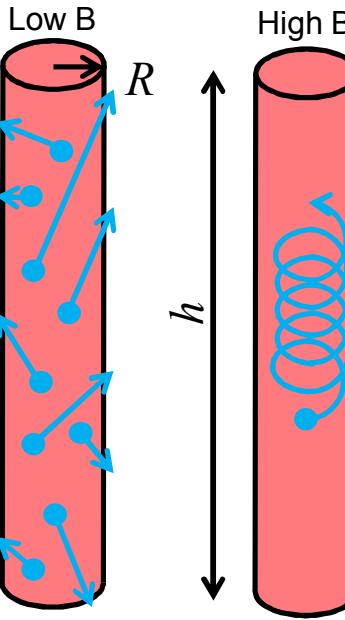
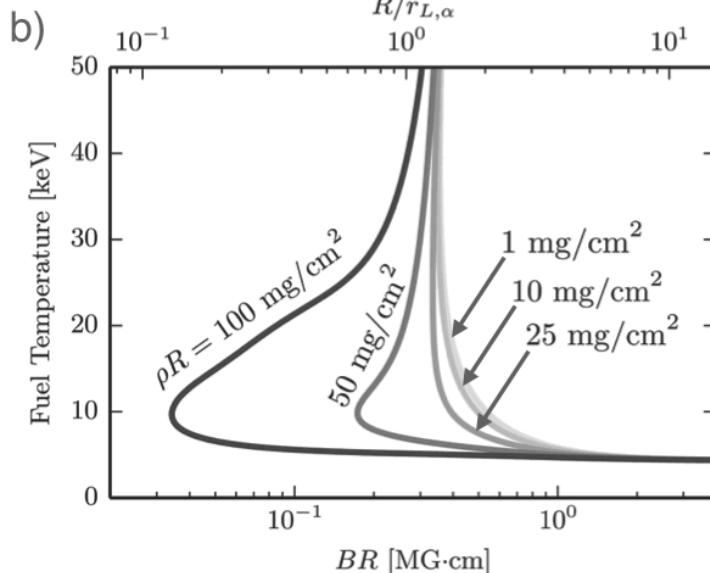


- Simple axial field used in a spherical implosion geometry
- Field suppressed electron heat conduction losses along one direction
- The resulting 30% increase in temperature and 15% increase in yield is consistent with rough estimates for heat loss suppression
- This is an example of success with a target that produced fusion yield without magnetization—can we produce yield in targets that wouldn't produce significant yield otherwise? (and gain benefit from doing so?)

# Magnetization (“BR”) can be used to reduce rho-R requirements and reduce electron heat losses, lower density also reduces bremsstrahlung radiation losses



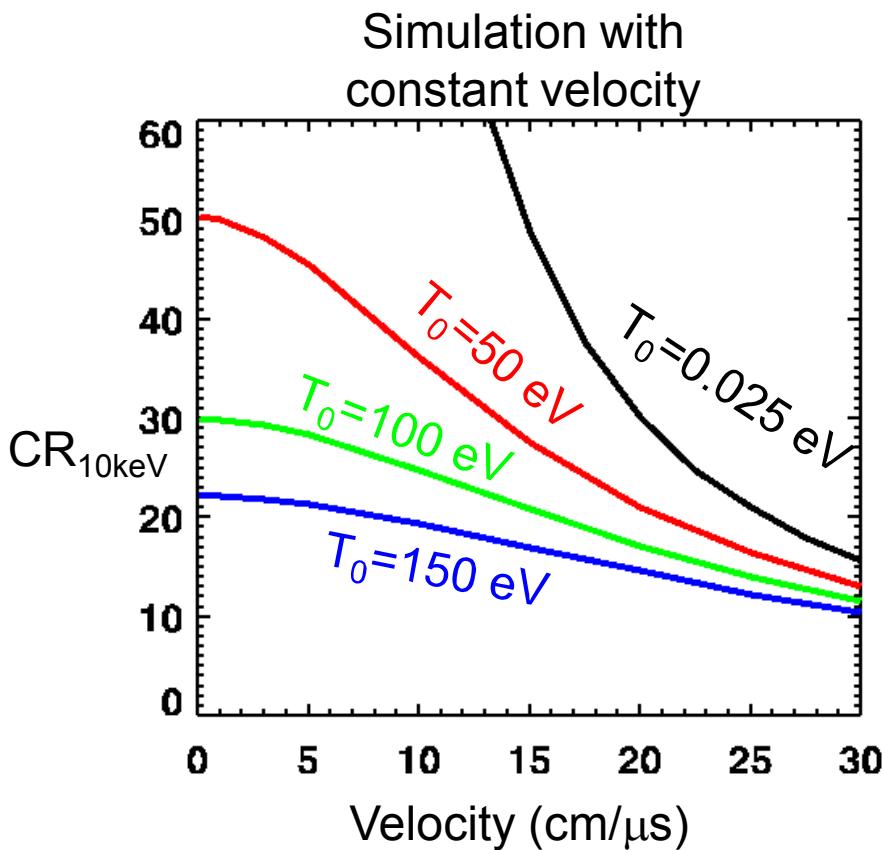
$$\frac{R}{r_\alpha} = \frac{BR [T \cdot \text{cm}]}{26.5} = \frac{BR [G \cdot \text{cm}]}{2.65e5} \approx 4BR [MG \cdot \text{cm}]$$



Fraction of trapped  $\alpha$ 's (tritons) is a function of **BR** only

At  $BR > 0.5 \text{ MG-cm}$  the effects saturate (particles are well confined). Measurements to date suggest  $0.4 \text{ MG-cm}$ !

# Typical ICF implosions need high velocities to reach fusion temperatures—starting the implosion with heated fuel potentially reduces requirements



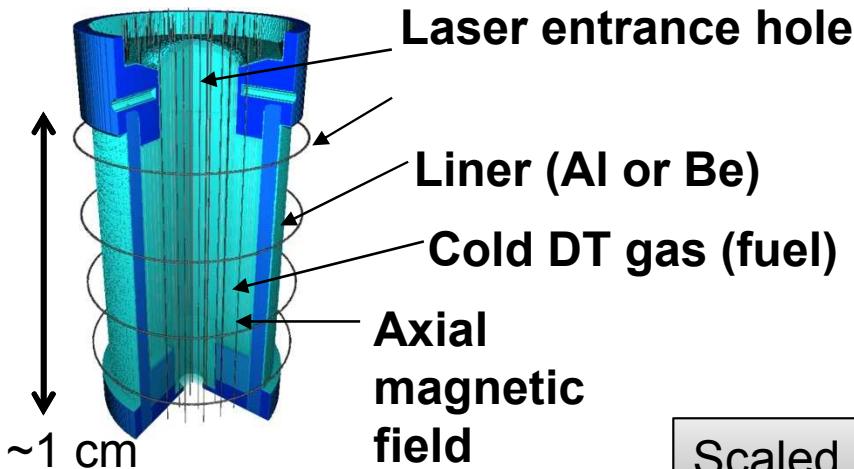
Heating fuel to ignition temperatures is typically done with a high-velocity shock (or series of shocks)

High velocities make it easier to reach fusion temperatures and also reduce the time available for losses (e.g., electron heat conduction or radiation)

Heating the fuel prior to the implosion ***in the absence of losses*** can allow low-velocity, low-convergence implosions to reach ignition temperatures – **magnetization is the key!**

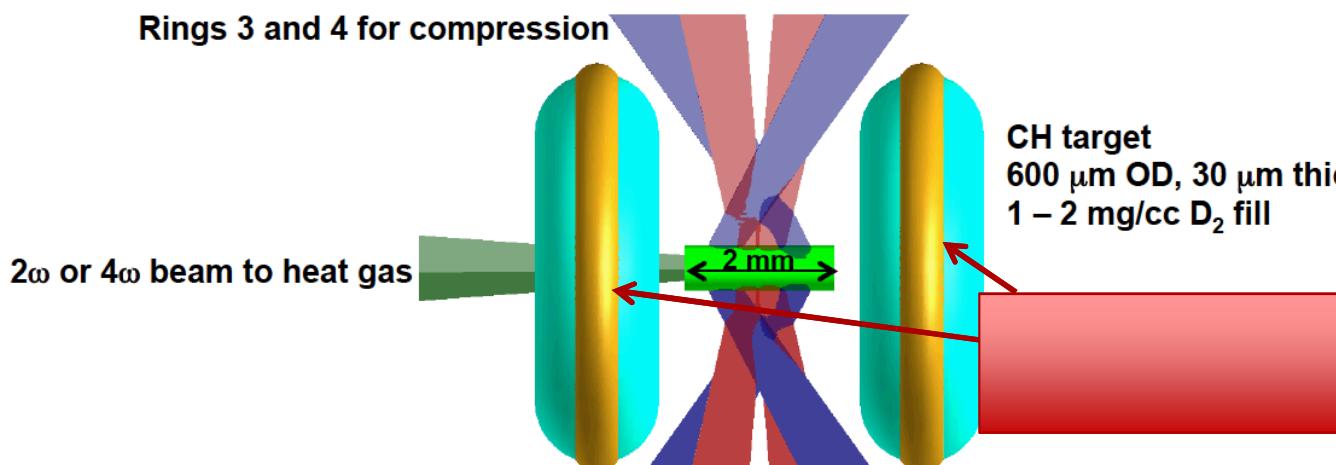
# We are evaluating a Magnetized Liner Inertial Fusion (MagLIF)\* concept that is well suited to pulsed power drivers and that may reduce fusion requirements

MagLIF on Z



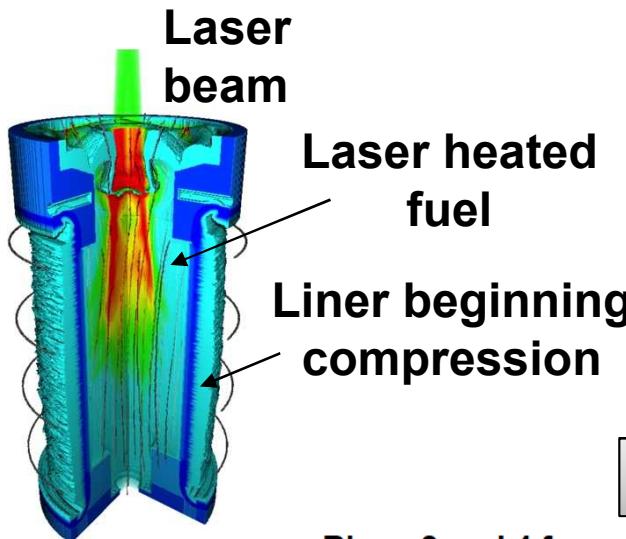
- Axial magnetization of fuel/liner ( $B_{z0} = 10-30$  T)
  - Inhibits thermal conduction losses, may help stabilize liner compression, ions magnetized too ( $\beta: 5 \sim 80$ ;  $\omega\tau > 200$ )

Scaled MagLIF on OMEGA



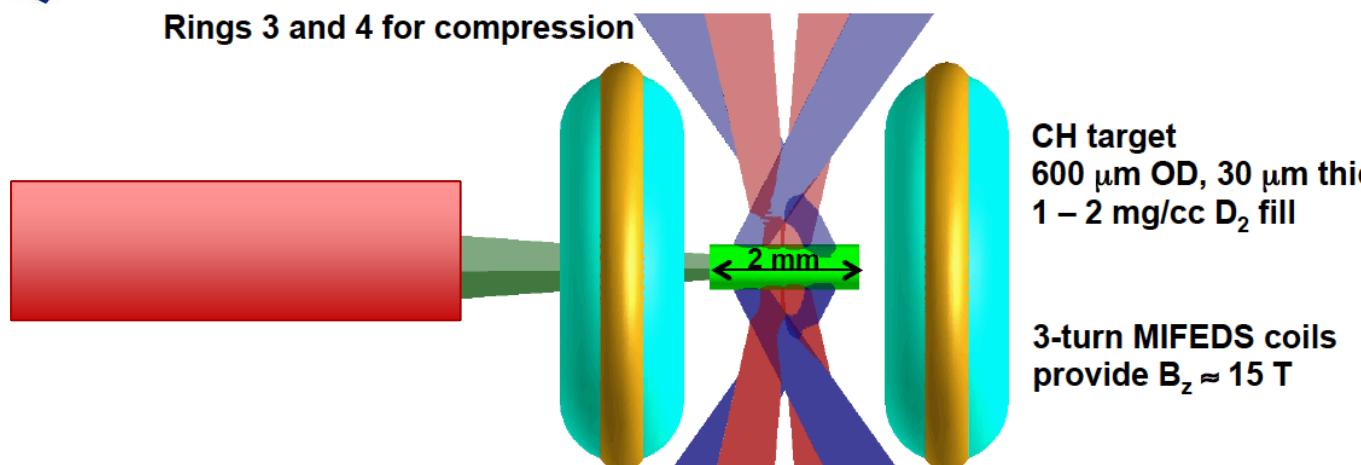
We are evaluating a Magnetized Liner Inertial Fusion (MagLIF)\* concept that is well suited to pulsed power drivers and that may reduce fusion requirements

### MagLIF on Z



- Laser heating of fuel (2-10 kJ)
  - Reduces amount of radial fuel compression needed to reach fusion temperatures ( $R_0/R_f = 23-35$ )

### Scaled MagLIF on OMEGA



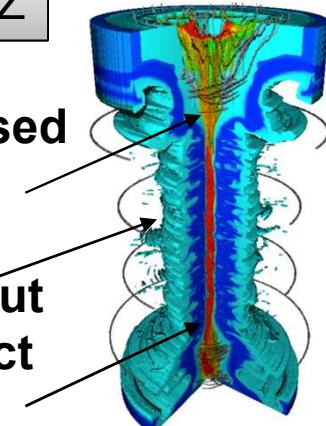
# We are evaluating a Magnetized Liner Inertial Fusion (MagLIF)\* concept that is well suited to pulsed power drivers and that may reduce fusion requirements

## MagLIF on Z

Compressed axial field

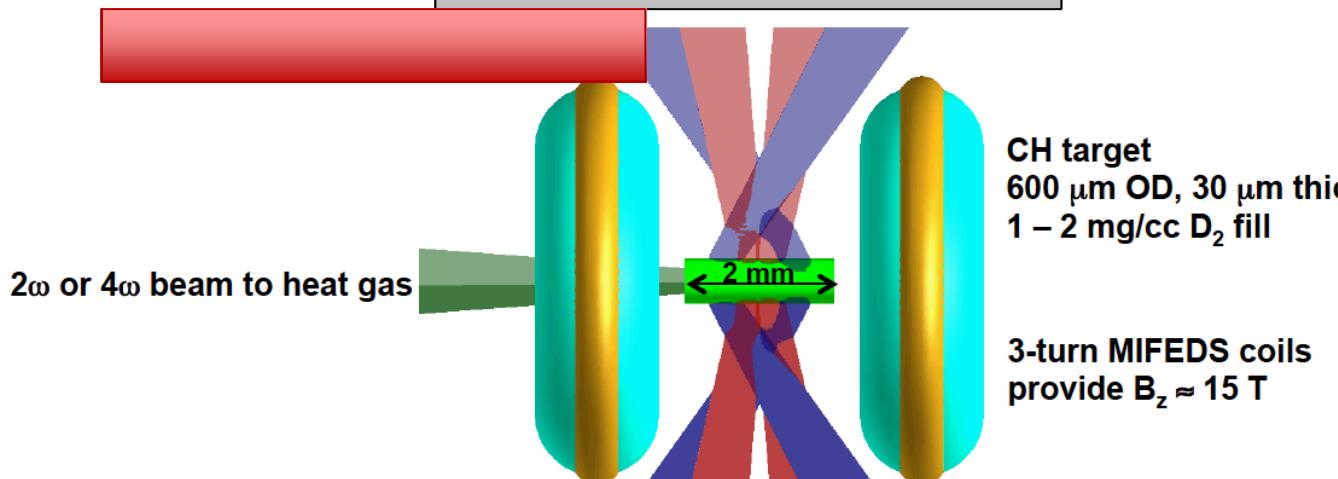
Liner unstable but sufficiently intact

Compressed fuel reaches fusion temperatures

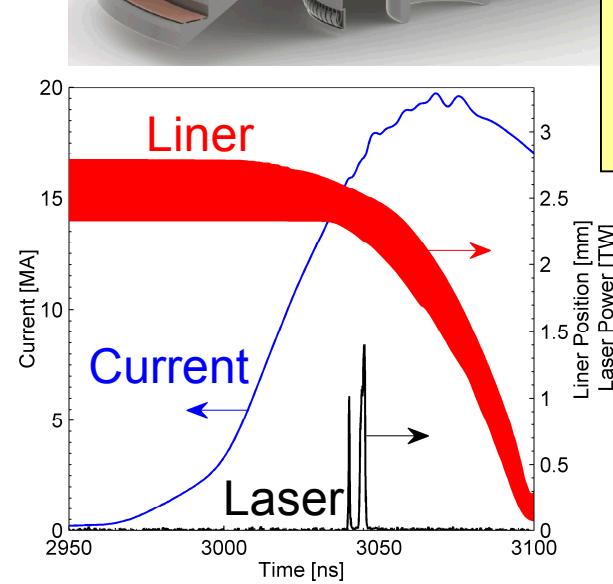
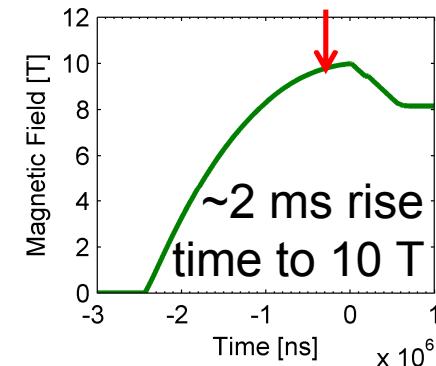
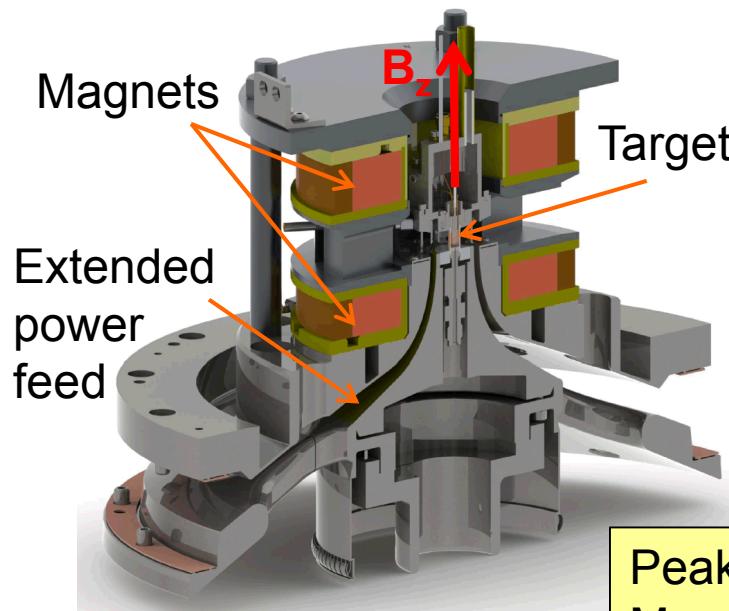
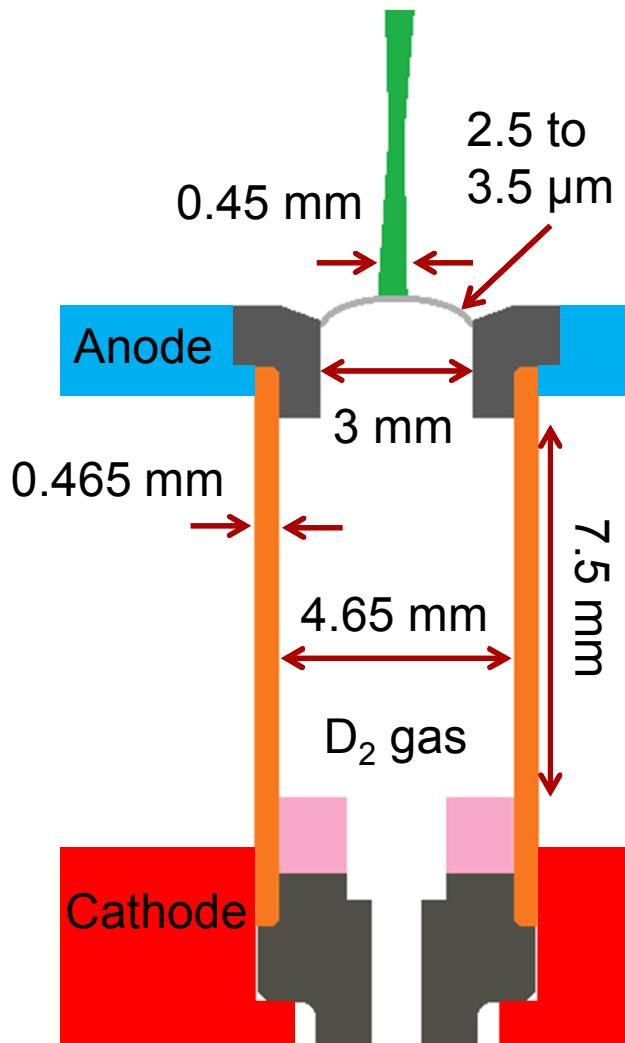


- Liner compression of fuel (70-100 km/s, ~100 ns)
  - “Slow”, quasi-adiabatic compression of fuel
  - Low velocity requirements allow use of thick liners ( $R/\Delta R \sim 6$ ) that are robust to instabilities (need sufficient  $pR$  at stagnation to inertially confine fuel)

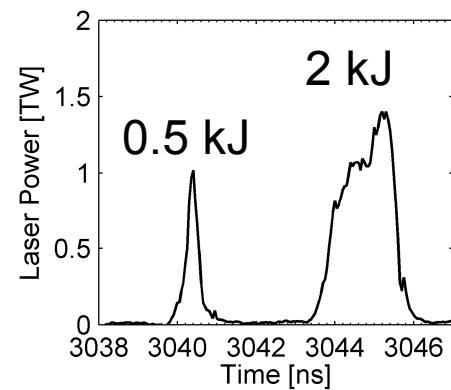
## Scaled MagLIF on OMEGA



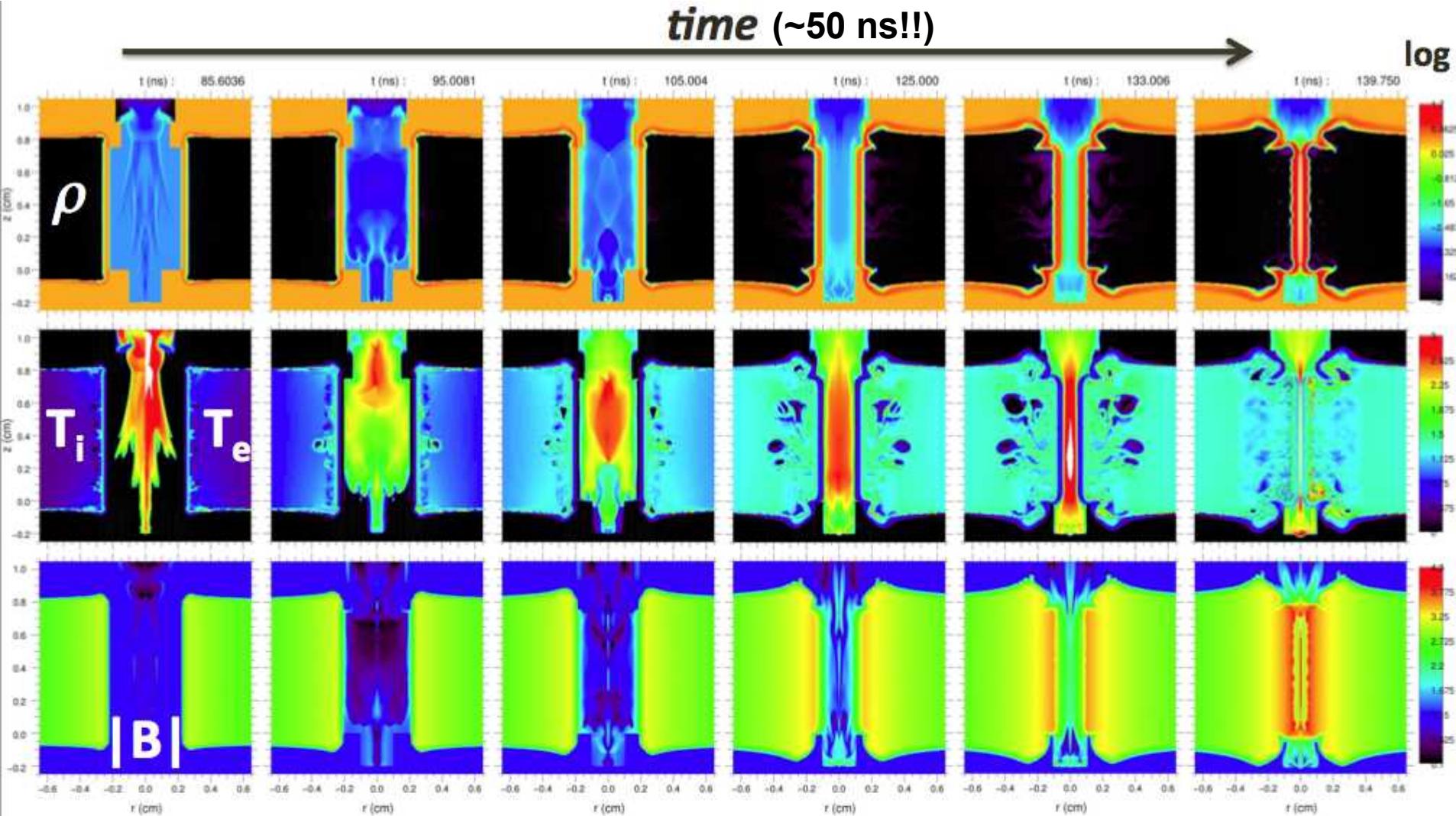
# The initial experiments used 10 T, 2.5 kJ laser energy, and 19-20 MA current to drive a $D_2$ filled (0.7 mg/cc) Be liner



Peak current is 19 MA  
Magnetic field is 10 T  
Total laser energy is 2.5 kJ

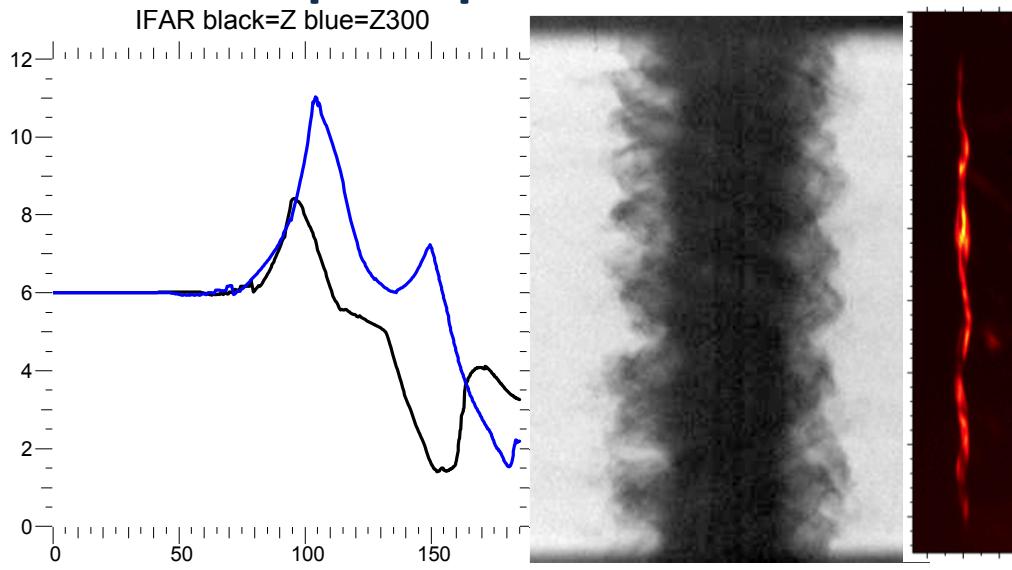


# Example fully-integrated HYDRA calculations of near-term Z experiments (19 MA, 10 T, 2 kJ) illustrate the stages of a MagLIF implosion



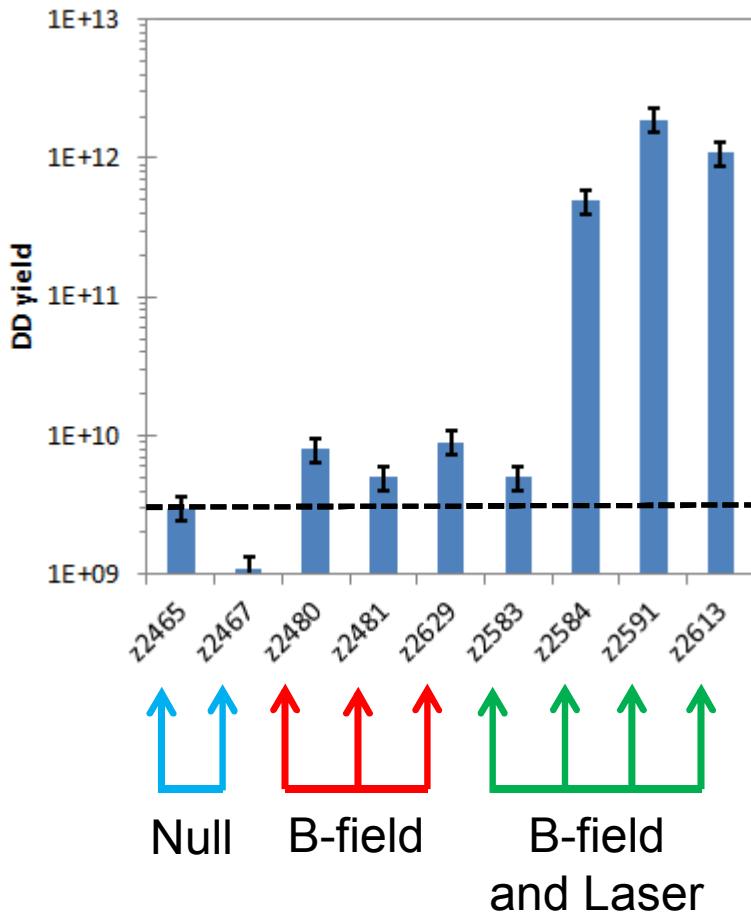
# Relative to the primary ICF approach, MagLIF uses a very different (conservative?) fuel compression method and largely untested magneto-inertial fusion principles

Metric	X-ray Drive on NIF	100 kJ MagLIF on Z
P	~140-160 Mbar	26 MA at 1 mm is 100 Mbar
P vs. R	Goes as $R^2$	Goes as $1/R^2$
Drive nature	Surface-like	Can be significant redistribution from diffusion & low-density plasma
Peak velocity	350-380 km/s	70-100 km/s
IFAR	17 (high foot) to 20 (low foot)	8.5
Hot spot CR	35 (high foot) to 45 (low foot)	25
Volume Change	42875x to 91125x (high & low foot)	625x
Fuel rho-R	>0.3 g/cm <sup>2</sup>	~0.003 g/cm <sup>2</sup>
Liner rho-R	n/a	>0.3 g/cm <sup>2</sup>
BR	n/a	>0.5 MG-cm
Burn time	~0.02 ns	~2 ns
T <sub>ion</sub>	>4 keV	>4 keV



- By traditional ICF implosion metrics MagLIF is very conservative, though different P vs. R
- Reaching fusion conditions relies on largely untested MIF principles
  - Long stagnation time (2 ns) → more susceptible to high-Z contamination
  - Magnetic suppression of heat transport

# Z shots producing DD yields in excess of $10^{12}$ were only observed in experiments with laser and B-field



- High yields were only observed on experiments incorporating **both** applied magnetic field and laser heating
- Experiments without laser and/or B-field produced yields at the background level of the measurement

**1D Simulation**

Bz	Preheat Energy	Peak Fuel $\langle T_i \rangle$
0	0	0.5 keV
0	2 kJ	0.92 keV
10 T	0	0.33 keV
10 T	2 kJ	3.00 keV

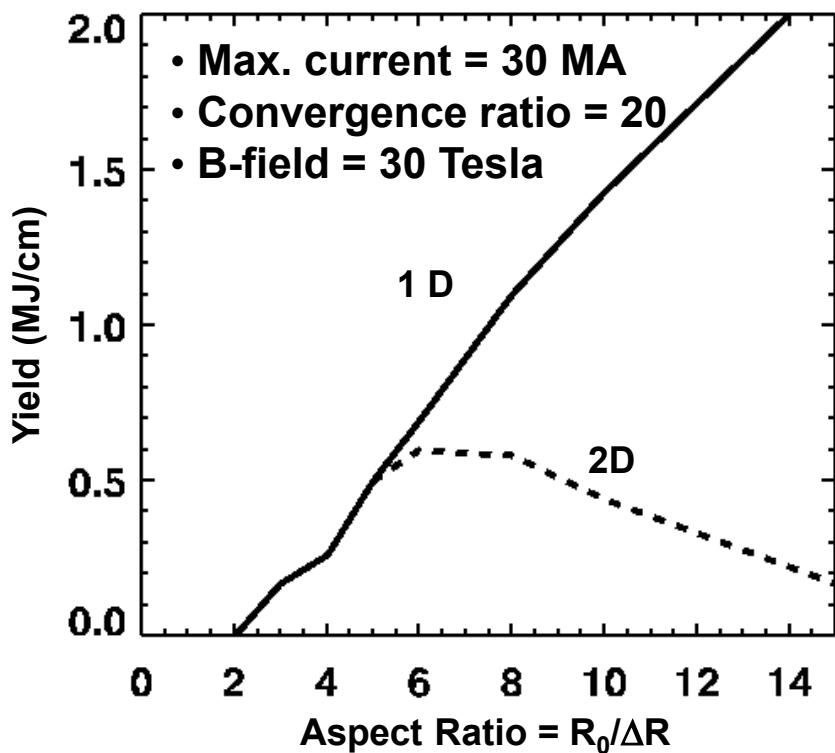
# Backups

# Comparison of 1D and 2D HYDRA calculations of near-term Z experiments (19 MA, 10 T, 2 kJ)

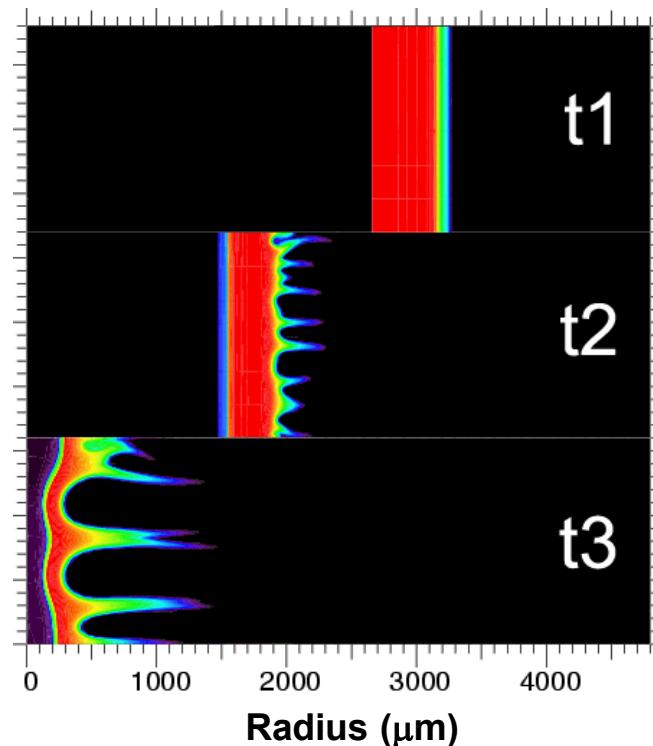
Parameter	1D ideal	2D integrated
• $E_{\text{gas}}^{\text{abs}}$	2.20 kJ	1.74 kJ
• $m_{\text{loss}}$	0%	43%
• $\Phi_{\text{loss}}$	36%	38%
• $\text{CR}_{\text{2D}}$	28 ( $r_{\text{stag}}$ 84 $\mu\text{m}$ )	37 ( $r_{\text{stag}}$ 63 $\mu\text{m}$ )
• $T_i^{\text{peak}}$	5.0 keV	6.5 keV
• $\langle T_i \rangle^{\text{DD}}$	2.9 keV	3.2 keV
• $\rho_{\text{gas}}^{\text{stag}}$	0.6 g $\text{cm}^{-3}$	0.5 g $\text{cm}^{-3}$
• $\rho R_{\text{liner}}^{\text{stag}}$	1.0 g $\text{cm}^{-2}$	0.9 g $\text{cm}^{-2}$
• $P^{\text{stag}}$	2.5 Gbar	2.2 Gbar (peak in bottle)
• $B_z^f r_{\text{stag}}$	4.1e5 G cm ( $r_{\text{stag}}/r_{\alpha}$ 1.5)	5.3e5 G cm ( $r_{\text{stag}}/r_{\alpha}$ 2.0)
• $Y_n^{\text{DD}}$	2.6e14 (in 7.5mm)	6.1e13 (24% of 1D)
• $Y_n^{\text{DD}}/Y_n^{\text{DT}}$	23	44
• $t_{\text{burn}}^{\text{FWHM}}$	3.2 ns	2.1 ns

Note: A unique property of magnetic drive is increasing pressure with decreasing radius. If less energy is coupled to fuel, target converges farther in simulations until plasma pressure is sufficient to stop the implosion.

# Reducing the implosion velocity requirements through fuel heating and magnetization allows us to use thicker, more massive liners to compress the fuel that are more stable



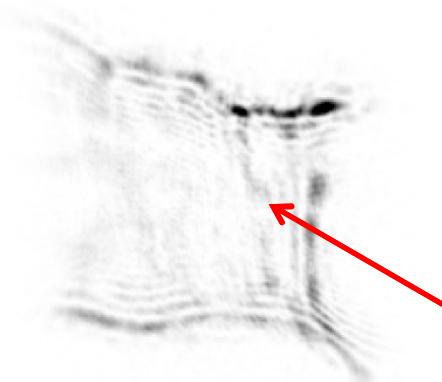
- The Magneto-Rayleigh-Taylor instability degrades the yield as the aspect ratio is increased (due to decreased liner  $\rho R$ )



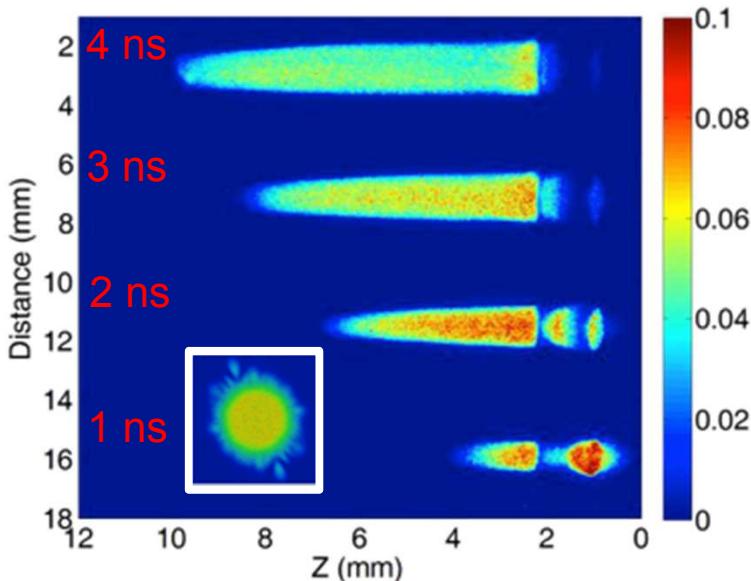
- Simulations of  $\text{AR}=6$  Be liner show reasonably uniform fuel compression and sufficient liner  $\rho R$  at stagnation to inertially confine the fuel—important because fuel density is low!

# Poor beam spot quality may be one reason that we are struggling to couple well to the fusion fuel

Z-Beamlet currently does not use any beam smoothing techniques adopted by the laser community

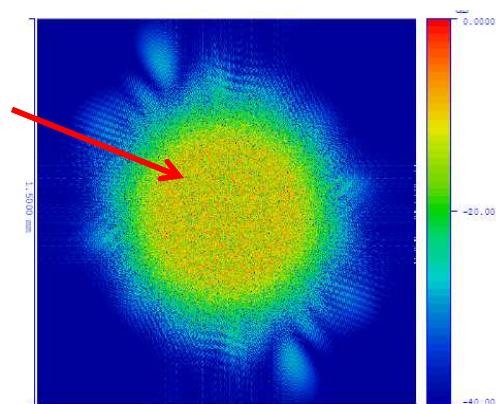


4 ns/3.1 kJ, 2  $\mu$ m LEH, no prepulse with DPP (SNL Omega-EP data)

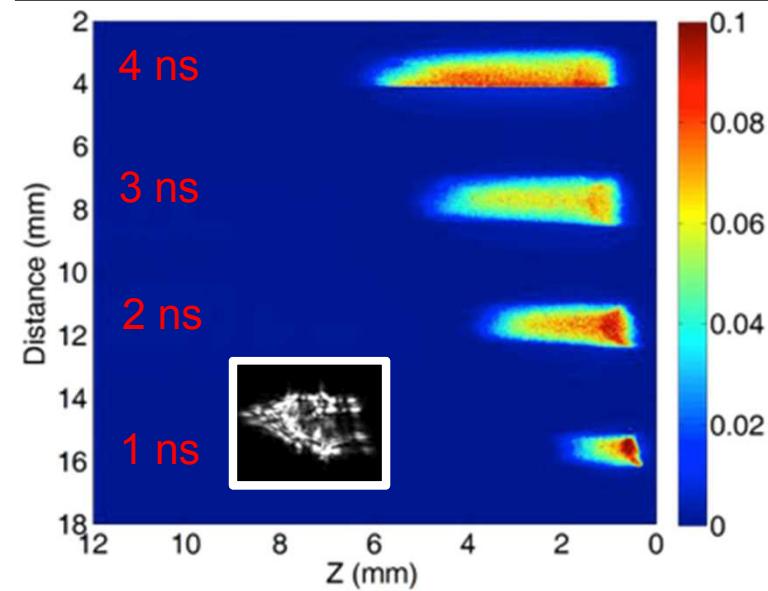


OMEGA-EP  
750um DPP

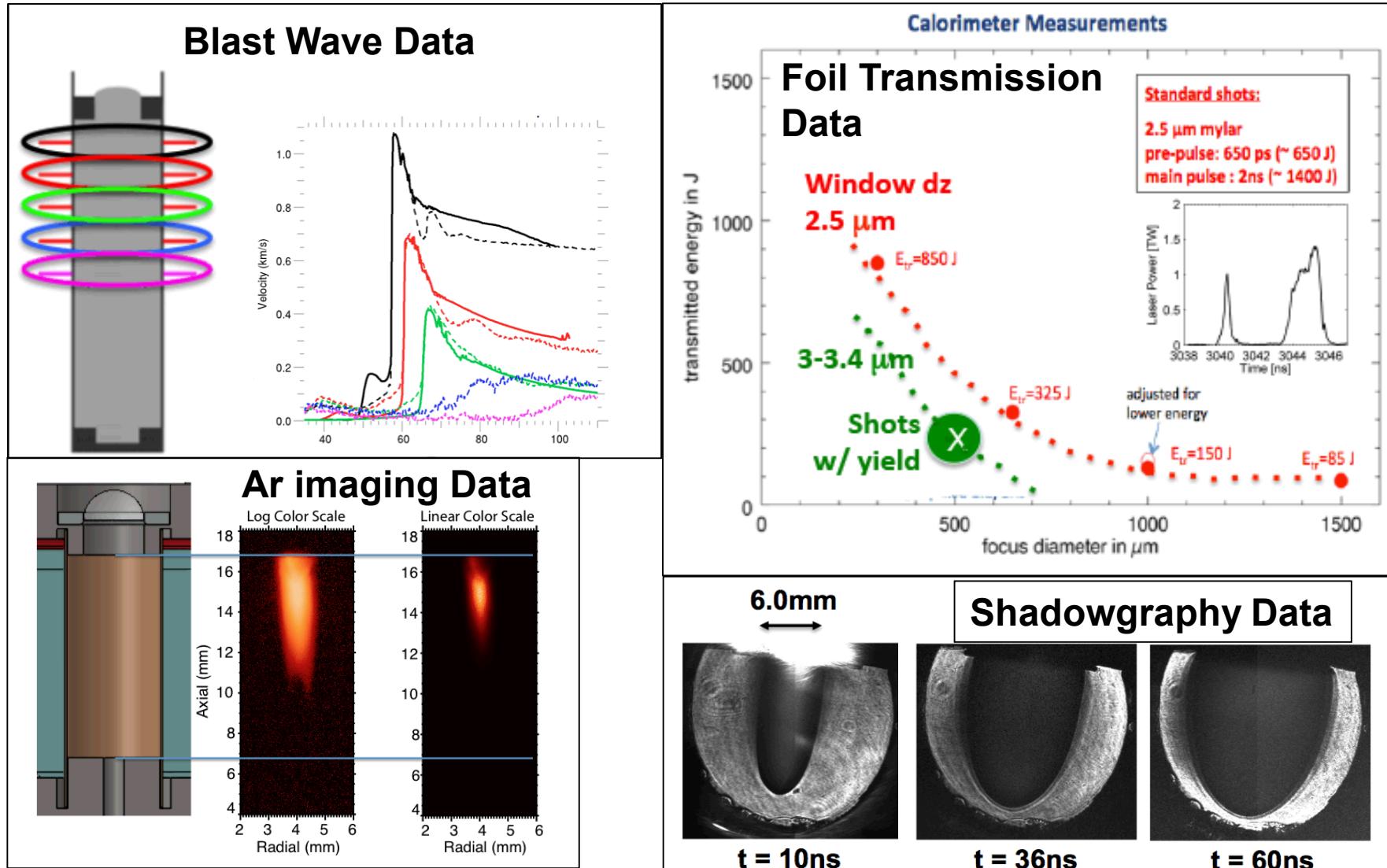
ZBL: No DPP  
(representative)



4 ns/2.93 kJ, 2  $\mu$ m LEH, no prepulse without DPP (SNL Omega-EP data)



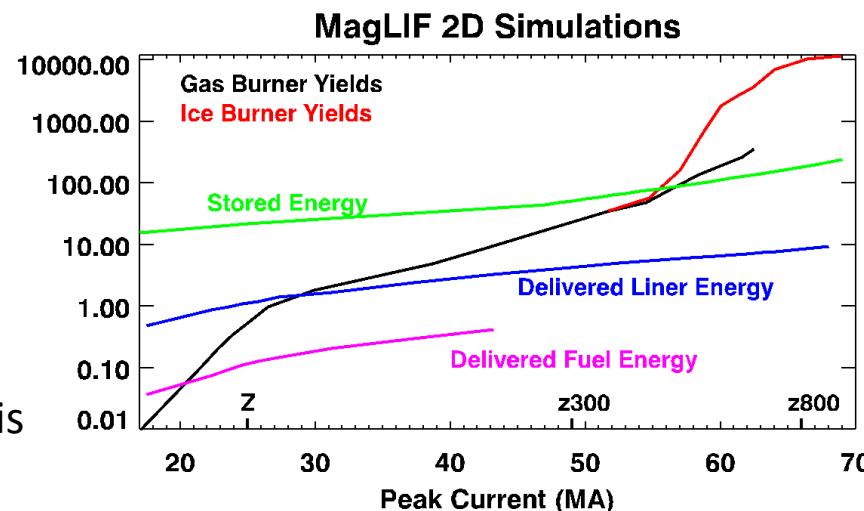
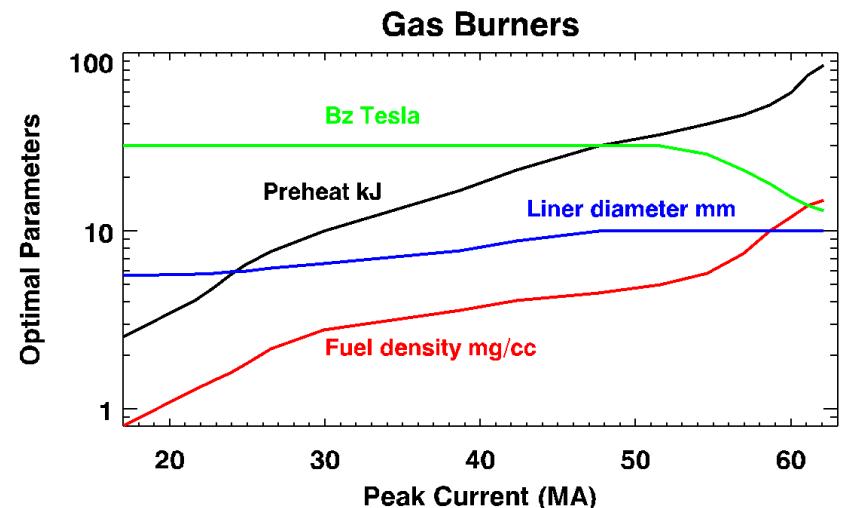
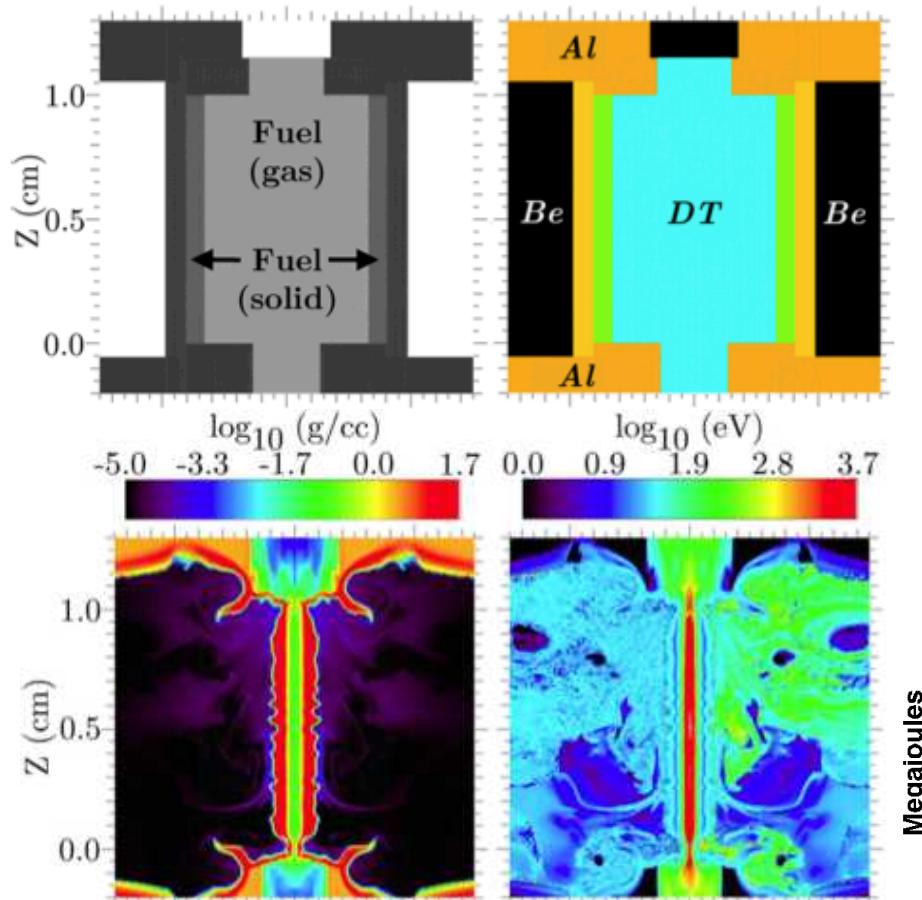
# Laser-only experiments appear to confirm that laser-fuel coupling is a concern: Multiple measurements are consistent with 200-600 J in heated gas out of >2000 J



See also M.R. Gomez, Talk 1B-2

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It may be possible to achieve  $\sim 100$  kJ yields on Z. Achieving alpha heating and ignition may be possible on a future facility. A cryogenic DT layer could enable up to  $\sim 1$  GJ yield.



An intermediate regime exists wherein the  $B_z$  field is

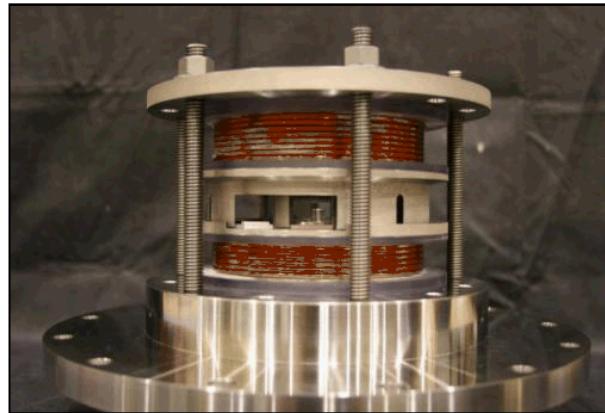
- *strong enough* to reduce conduction losses, but
- *weak enough* not to inhibit the  $\alpha$  deflagration wave

# We have successfully implemented 10 T axial fields over a several cm<sup>3</sup> volume for MagLIF and the capacitor bank is capable of driving 30 T field coils under development

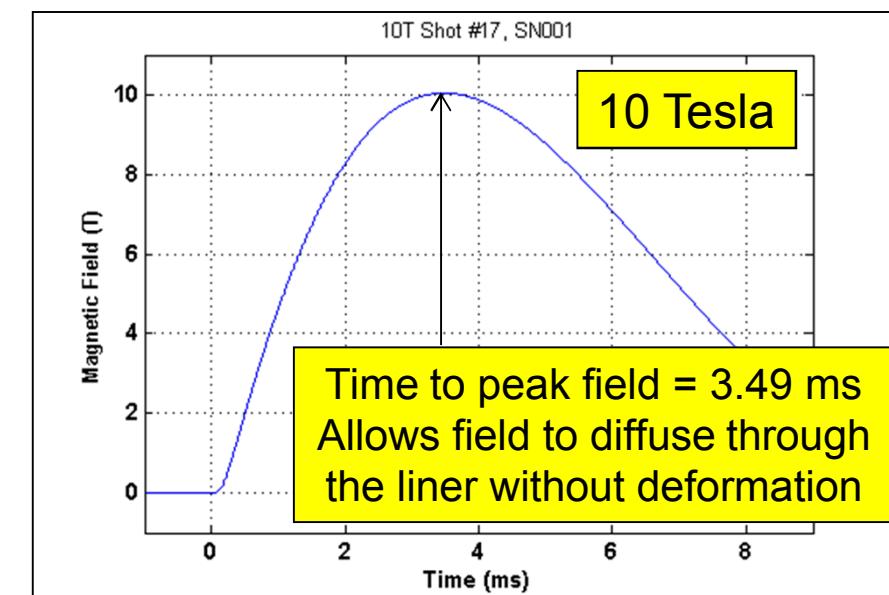
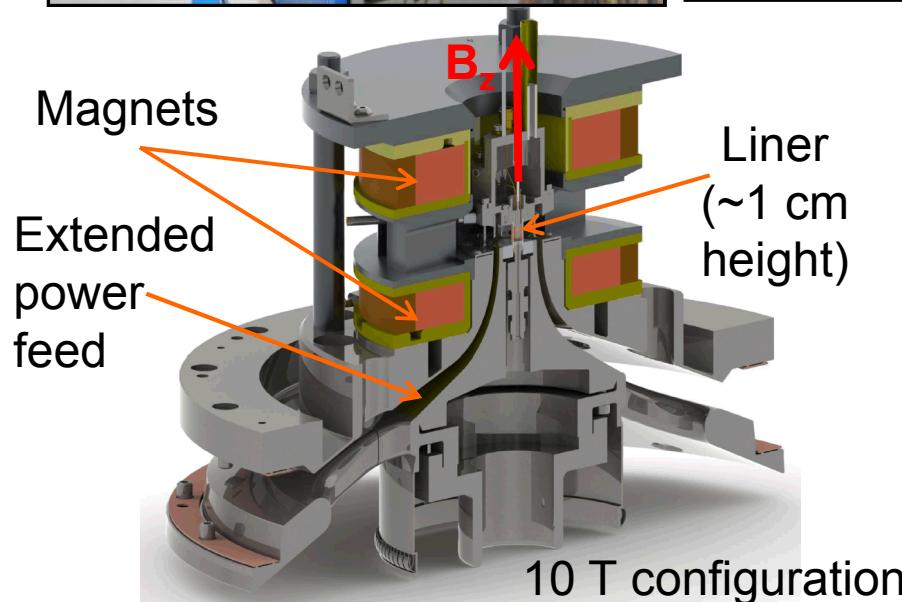
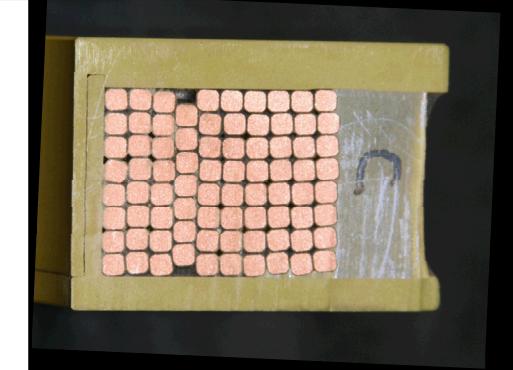
Capacitor bank system on Z  
900 kJ, 8 mF, 15 kV (Feb. 2013)



Example MagLIF coil assembly  
with copper windings visible

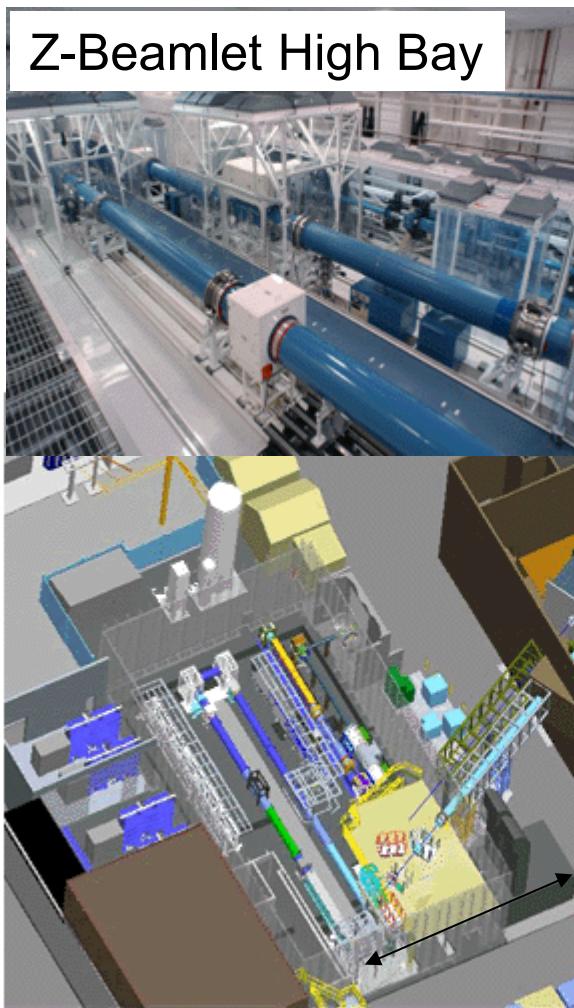


Cross section of coil showing  
Cu wire, Torlon housing, and  
Zylon/epoxy reinforcement

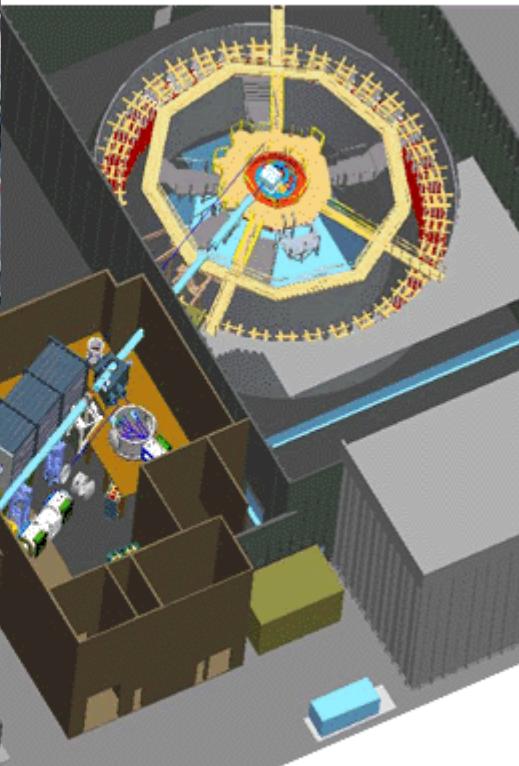


# The Z-Beamlet laser at Sandia\* can be used to radiograph liner targets and heat fusion fuel

Z-Beamlet High Bay



Z facility



Z-Beamlet and Z-Petawatt lasers

Z-Beamlet (ZBL) is routinely used to deliver  $\sim 2.4$  kJ of  $2\omega$  light in 2 pulses for backlighting experiments on Z

Modifications adding bandwidth to the laser enable us to reach 4-4.5 kJ in 4 ns today

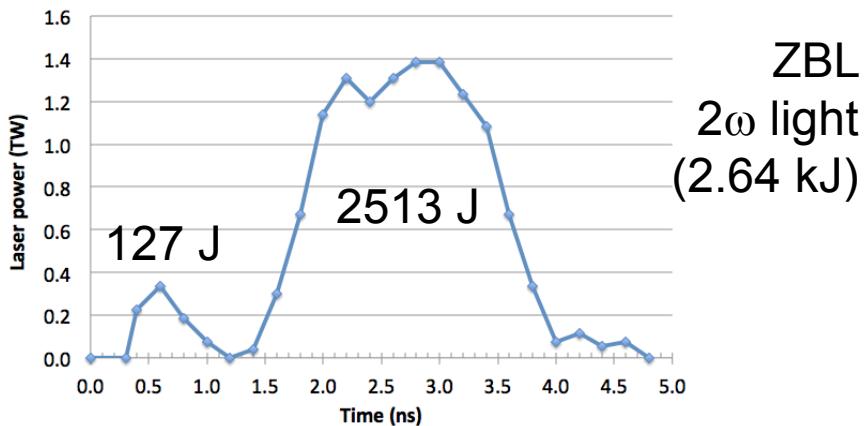
Filling out the booster amps would enable longer pulses (5-7 ns) which would extract up to 6 kJ of  $1\omega$ , for 4.2 kJ of  $2\omega$ .

Typical MagLIF initial fuel densities correspond to 0.10 to  $0.30 \times$  critical density for  $2\omega$

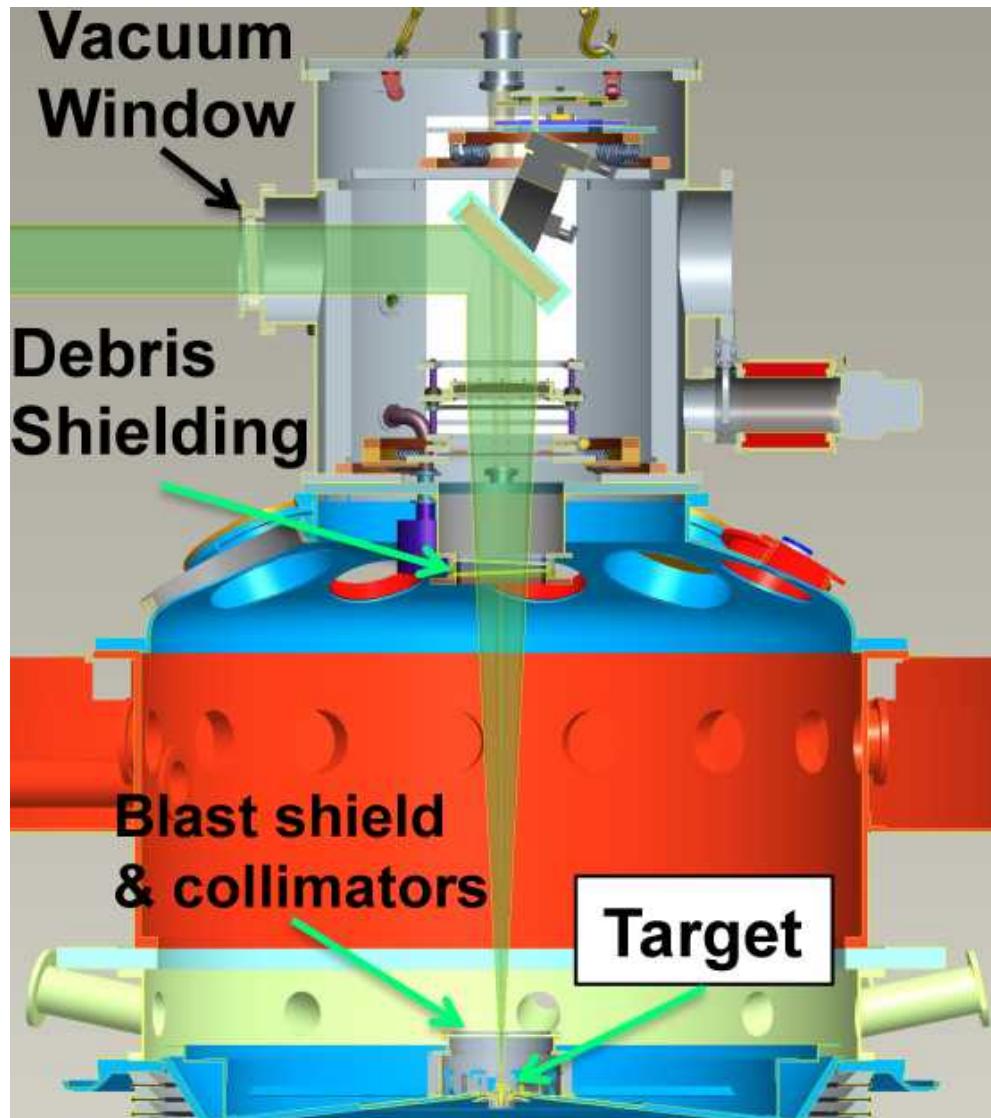
In August 2013 we commissioned a new vacuum final optics assembly to safely enable 2 kJ of on-axis laser heating of fuel



Example pulse measurement

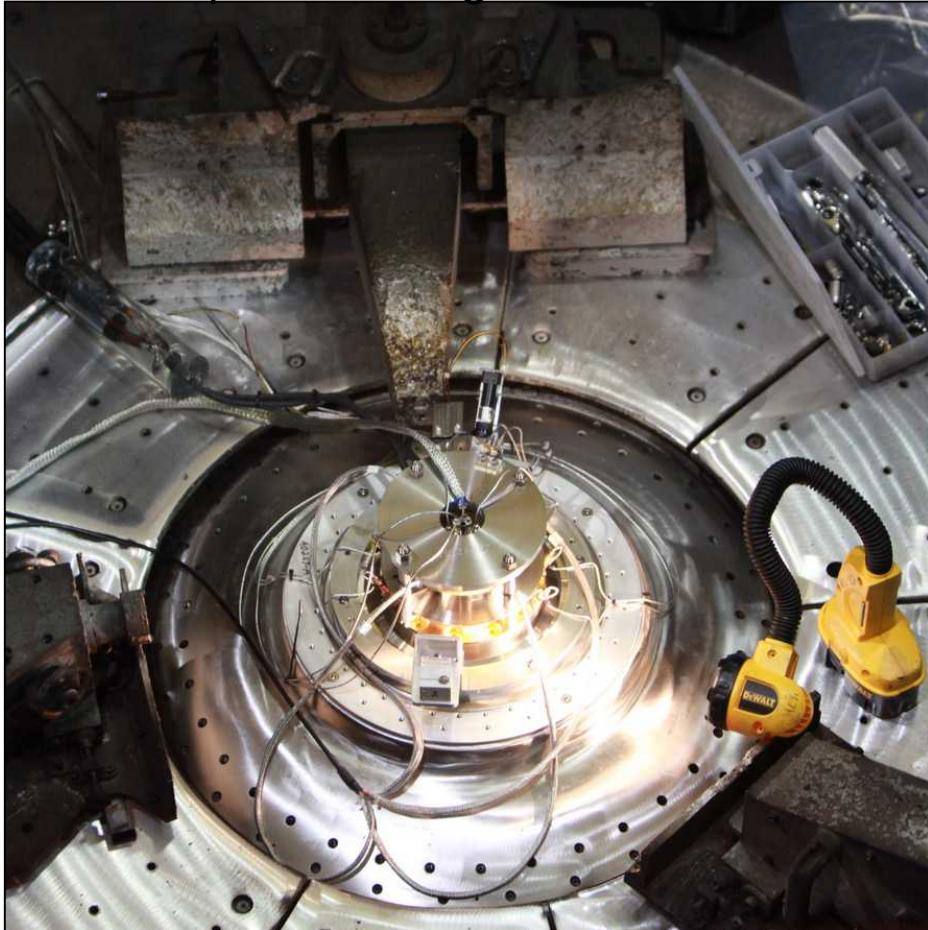


Prepulse vaporizes gas-containing foil; main pulse couples to DD fuel



Z couples several MJ of energy to the load hardware,  
~equivalent to a stick of dynamite, making diagnostic  
measurements and laser coupling challenging

Pre-shot photo of MagLIF load hardware

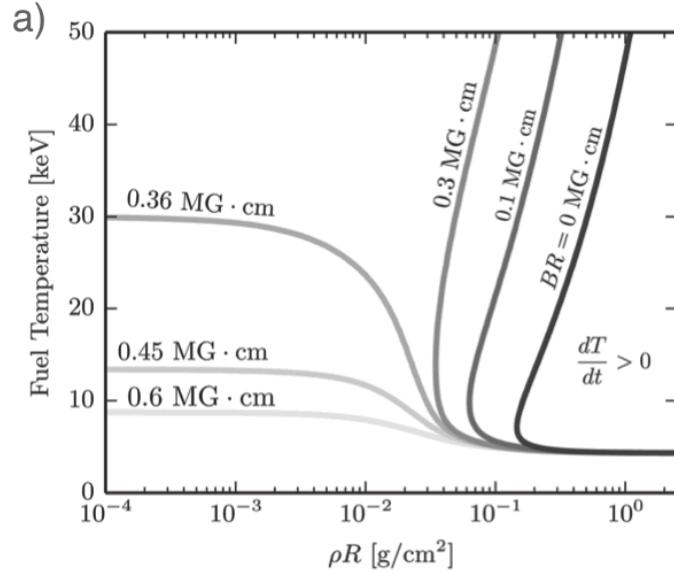


Damage to FOA  
debris shielding



Post-shot  
photo

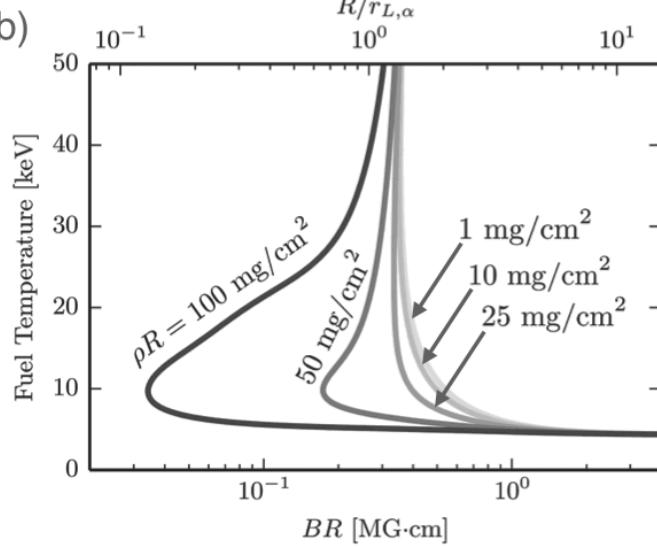
# Magneto-inertial fusion seeks to compress heated fuel, using low fuel density and magnetization to minimize radiation and electron thermal conduction losses, respectively



The  $\square r$  needed for ignition can be significantly reduced by the presence of a strong magnetic field largely through inhibiting electron conduction

Lower  $\square r$  reduces the required final fuel density (e.g.,  $\sim 1 \text{ g/cc} \ll 100 \text{ g/cc}$ ), reducing radiation loss

This means the stagnation plasma pressure at ignition temperatures is significantly reduced (e.g.,  $\sim 5 \text{ Gbar} \ll \sim 500 \text{ Gbar}$  for hot spot ignition)



Large values of  $BR$  are needed and therefore large values of  $B$  are needed,  $B \sim 10,000 \text{ Tesla}$  (Earth's  $B$ -field is  $\sim 0.00003 \text{ Tesla}$ )

This field significantly exceeds pulsed coil technology ( $B_0 \sim 10-30 \text{ T}$ ), therefore flux compression is needed