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ABSTRACT

An extraordinary number of Hubble constant measurements challenges physicists with selection of the best numerical value. The standard U.S. Nuclear Data Program (USNDP) codes and procedures have been applied to resolve this issue. The nuclear data approach has produced the most probable or recommended Hubble constant value of 67.00(770) (km/sec)/Mpc. This recommended value is based on the last 25 years of experimental research and includes contributions from different types of measurements. The present result implies $(14.6 \pm 1.7) \times 10^9$ years as a rough estimate for the age of the Universe. The complete list of recommended results is given and possible implications are discussed.

Subject headings: cosmology: observations — methods: data analysis

1. Introduction

The Hubble constant and its present-day numerical value play an important role in modern astrophysics and cosmology (Boyd 2008; Dolgov et al. 1988). For many years, researchers have been improving the accuracy of the constant (Livio, Riess 2013). Fig. 1 shows the time evolution of Hubble research in the last 100 years, and the original effort can be traced as early as 1916 (Huchra et al. 2015). The large number of measurements creates a certain degree of confusion about Hubble constant numerical value, and scientists often rely on recently-published results (Particle Data Group 2014). The precision of Hubble constant measurements has improved enormously over the years; however, it is not always prudent to reject older results in favor of the latest findings. Consequently, it makes perfect sense to analyze all available results, evaluate the data, and extract the recommended value.

2. Hubble Constant Evaluation

The volume of Hubble constant measurements far exceeds other experimental quantities in physics (Pritychenko 2015). Previously, similar situations have been encountered in nuclear and particle physics and resolved with data evaluations. Nuclear data evaluations and their policies

are well described in literature (Pritychenko et al. 2012; ENSDF 2015). Frequently, the evaluations are completely based on or adjusted to available experimental data. These adjustments and specialized mathematical statistics techniques can be applied for nuclear, particle, or any other data sets.

In this work, I would follow standard nuclear data evaluation procedures to deduce the recommended value. Current evaluation input data are mostly based on the NASA/HST Key Project on the Extragalactic Distance Scale compilation (Huchra et al. 2015) and recent results. A visual inspection of historical Hubble Constant measurements, as shown in Fig. 1, is instrumental in the data analysis. It suggests that one may safely reject all measurements prior to 1970. It is common knowledge that Hubble constant measurements heavily rely on the accuracy of astrometric distance determination. Older results, such as those reported by A. Sandage and G. de Vaucouleurs, suffered from inaccurate measurements (Livio, Riess 2013). Therefore, the rejection of all results prior to 1990 could provide a complementary benchmark value of the Hubble constant.

In the present data analysis, the experimental data have been separated into two groups, with 1970 and 1990 time cuts, and further reduced us-

ing the following policies:

- Rejection of repeated results (multiple publication of the same result)
- Rejection of model-dependent results
- Rejection of potential outliers using Chauvenet’s criterion (Birch, Singh 2014)

Common data evaluation practices indicate that recommended value should be based on a large statistical sample that includes different types of measurements. The 1970 and 1990 redacted data sets of ~ 370 and ~ 300 data points, respectively, provide such samples. These large samples create the possibility of deducing Hubble constant value for each method of observation besides the combined value that is based on all measurements. The current data collections were further subdivided using a NASA/HST Key Project on the Extragalactic Distance Scale classification of experimental methods (Huchra et al. 2015):

- S = Type Ia supernovae (SNIa)
- 2 = Type II supernovae (SNII)
- G = Globular cluster luminosity function
- L = Lens
- r = Red Giants
- B = Baryonic Tully-Fisher
- R = Inverse Tully-Fisher (ITF, RTF)
- H = Infrared Tully-Fisher (or IRTF)
- F = Fluctuations
- D = Dn-Sig/Fund Plane
- A = Global Summary
- Z = Sunyaev-Zeldovich
- T = Tully-Fisher
- O = Other
- N = Novae
- P = Planetary nebula luminosity function

- C = CMB Fit

These experimental data sets have been processed with the latest version of the visual averaging library (Birch, Singh 2014). The library program includes limitation of relative statistical weight (LWM), normalized residual (NRM), Rajeval technique (RT), and the Expected Value (EVM) statistical methods to calculate averages of experimental data with uncertainties. The experimental data sets were processed, and evaluated values with reduced $\chi^2 < 2$ were typically accepted as reasonable data fits. The current evaluation incorporates statistical methods based on the inverse squared value of the quoted uncertainties, a procedure that is consistent with the general methodology used in treatment of data for the ENSDF database (ENSDF 2015) and Particle Data Group (Particle Data Group 2014).

3. Results and Discussion

Two sets of recommended values are displayed in Fig. 2, and the combined numerical values are shown in Table 1. The Hubble constant combined central values extracted by means of different mathematical techniques are in good agreement, while uncertainties need further discussion. Visual inspection of the numerical values shown in the Fig. 2 indicates the Best Representation mathematical procedure (depicted as “BR”) provides a best fit for different experimental methods with reasonable uncertainties. Other procedures such as the Bootstrap technique (depicted as “Bs”) result in rather small uncertainties. These small uncertainties are due to specifics of data analysis procedures such as assignment of different statistical weights to results with different uncertainties. In light of this disclosure, it is prudent to select the Best Representation results as the recommended value. The current results are consistent with the recent Particle Data Group publication (Particle Data Group 2014).

Finally, two different time cuts of 1970 and 1990 for Hubble’s data have yielded two recommended values of 67.00(880) and 67.00(770) (km/sec)/Mpc, respectively. The agreement between these values partially reflects the fact that the majority of the Hubble’s constant measurements has been performed in the last 25 years, and a small number of potential outliers has been

rejected. More accurate recent observations imply a preference for the 1990 time cut value of 67.00(770) (km/sec)/Mpc. The last result agrees well with the Planck’s Mission and WMAP values (Wikipedia 2015; Particle Data Group 2014).

The knowledge of Hubble constant value has multiple implications in science. As an example, a rough estimate of the age of the Universe can be deduced using the standard methodic (Wikipedia 2015). The adopted value of 67.00(770) (km/sec)/Mpc implies $(14.6 \pm 1.7) \times 10^9$ years estimated value for the age of Universe. The last result is consistent with the recently published value of $(13.798 \pm 0.037) \times 10^9$ years (Planck 2014).

4. Conclusions

The analysis of Hubble constant measurements has been performed using standard USNDP codes and procedures. An evaluated data set of most probable values of Hubble constant has been deduced and shown in the Table 1. These values are consistent with other available results. An accurate constant value is instrumental for many potential applications. The recommended value of the constant is completely based on experimental measurements, and further, more precise observations, would lead to more accurate determination of it.

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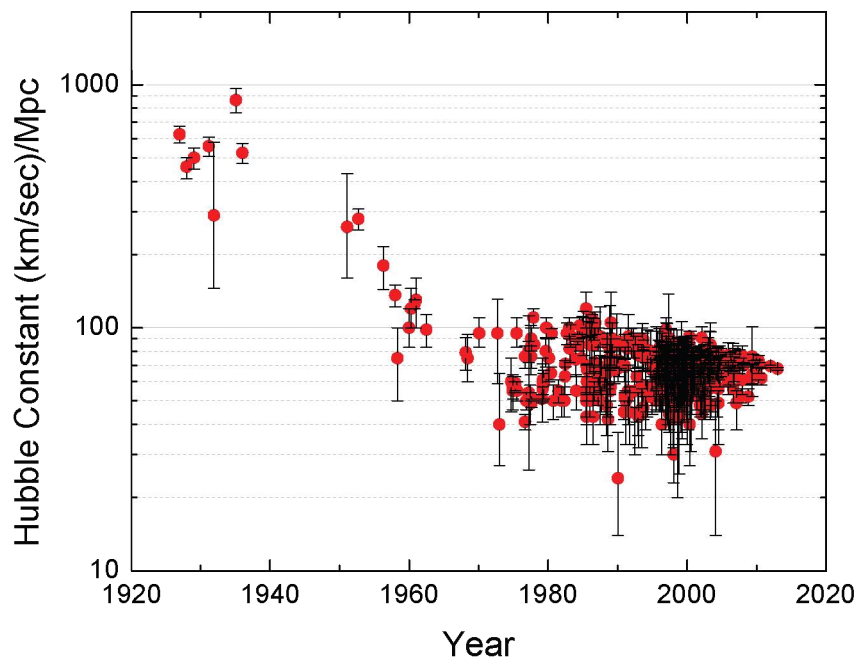


Fig. 1.— Historical evolution of Hubble constant measurements.

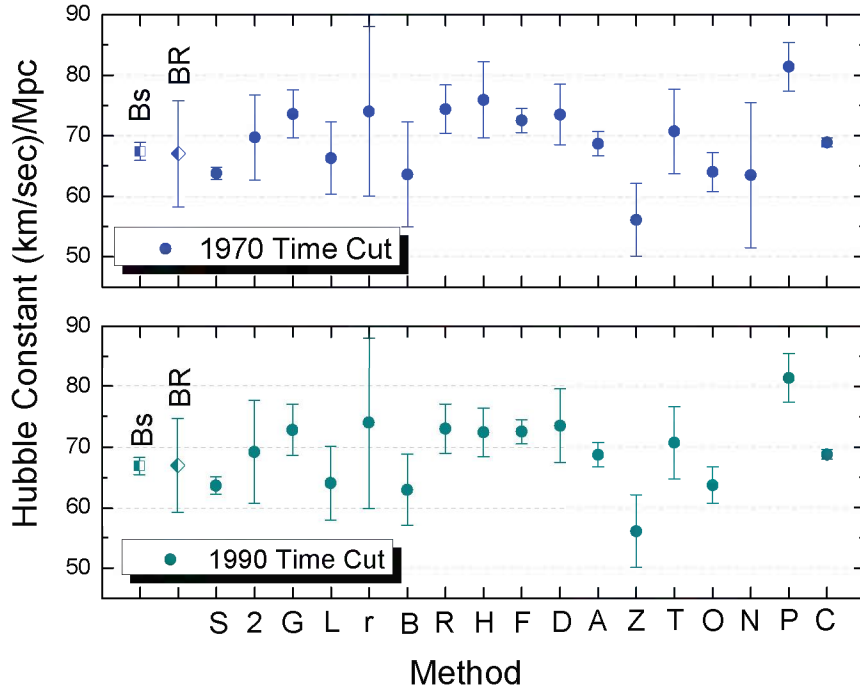


Fig. 2.— Evaluated values of Hubble constant using 1970 and 1990 time cuts for experimental data. These plots include the evaluated or recommended values for combined observations, using the Best Representation (BR) and Bootstrap (Bs) data analysis techniques.

Table 1: Results of the Hubble Constant evaluation for all observations using 1970 and 1990 time cuts.

Method	Time Cut: 1970 (km/sec)/Mpc	Time Cut: 1990 (km/sec)/Mpc
Unweighted Average	68.08(68)	66.71(62)
Weighted Average	65.27(50)	65.68(45)
LWM	67.84(64)	66.85(60)
Normalized Residual	66.04(43)	66.09(42)
Rajeval Technique	67.08(37)	66.53(38)
Best Representation (BR)	67.00(880)	67.00(770)
Bootstrap (BS)	67.40(150)	66.90(130)
Mandel-Paule	67.30(900)	66.20(630)