

Temperature Measurements of Expansion Products

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**Sandia
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Objective & Goals:

- ***to obtain time-dependent thermal characteristics from the radiative emission of the expansion products resulting from debris generated from metallic and explosive targets shocked to very high pressures and temperatures***

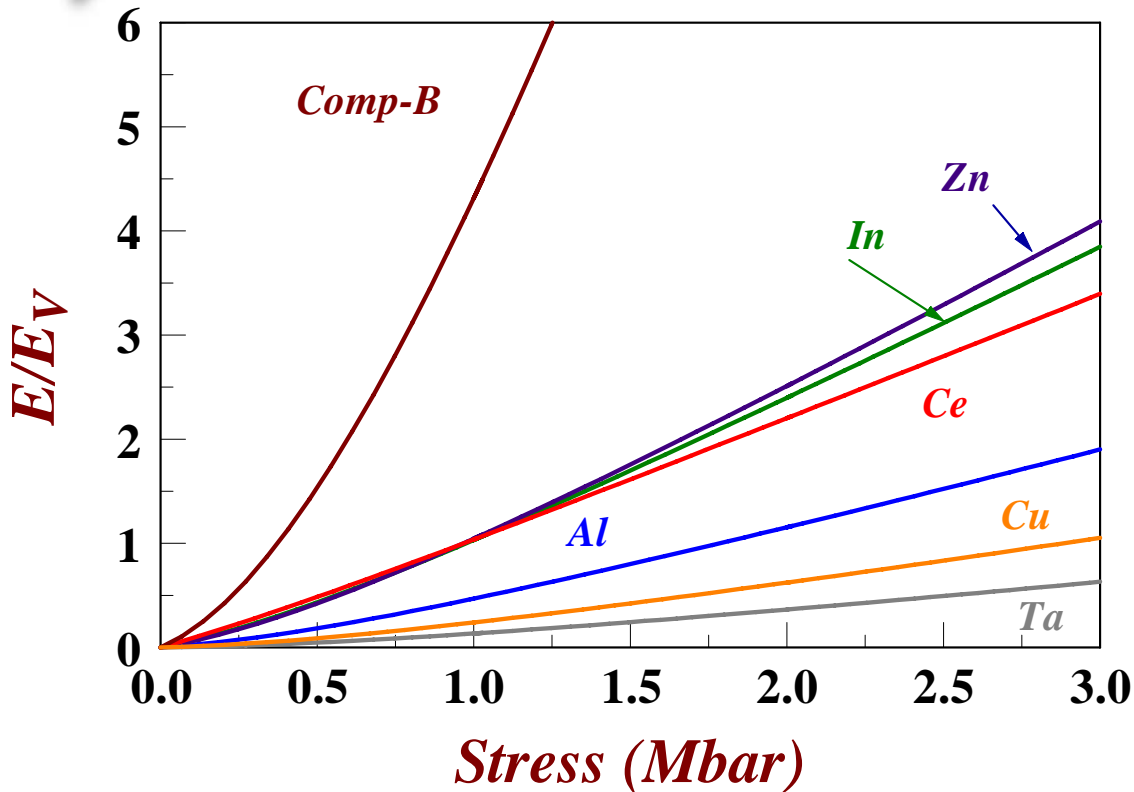


Assumptions:

- ***Assume we have captured the spectrum of a blackbody with grey emissivity***
 - ***Knowing it is not '1', but emissivity is constant over all wavelengths***
- ***Local thermodynamic equilibrium is assumed***
 - ***Not strictly true: spectral lines are a clear indicator of a non-equilibrated source***
 - ***Making spatial average estimates of temperature where gradients may exist: debris cloud at a single temperature***

Even though these approximations/assumptions may be sources for large uncertainties, the calculated temperature estimates appear quite reasonable—as will be seen

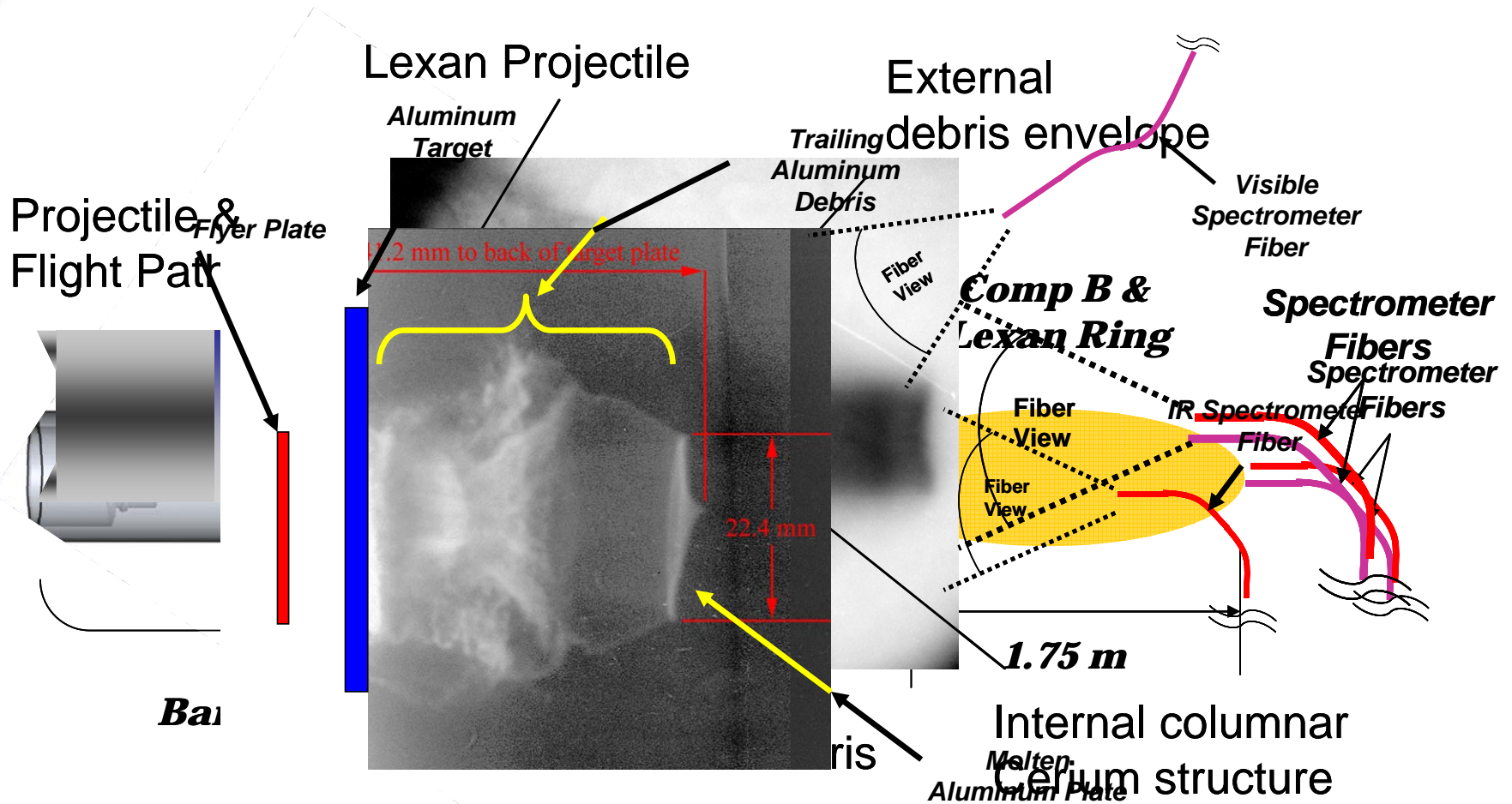
Materials: E/E_v Relation



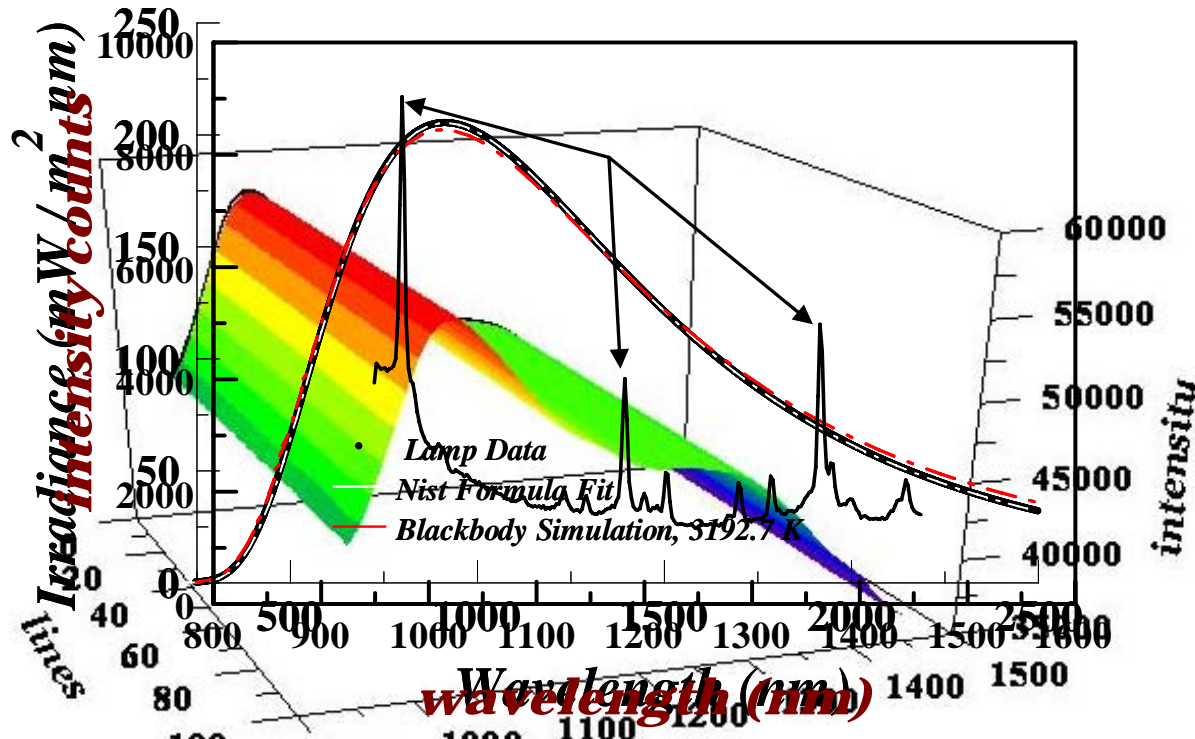
**Experiments
conducted on
aluminum, cerium,
and Composition-B**

- E is the internal energy increase of the shocked material,
- E_v is the specific energy required to vaporize the material

Experimental Method:



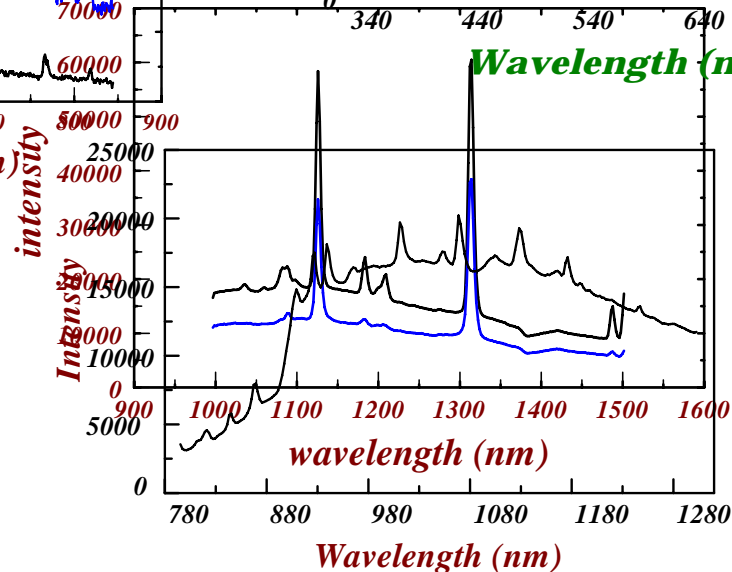
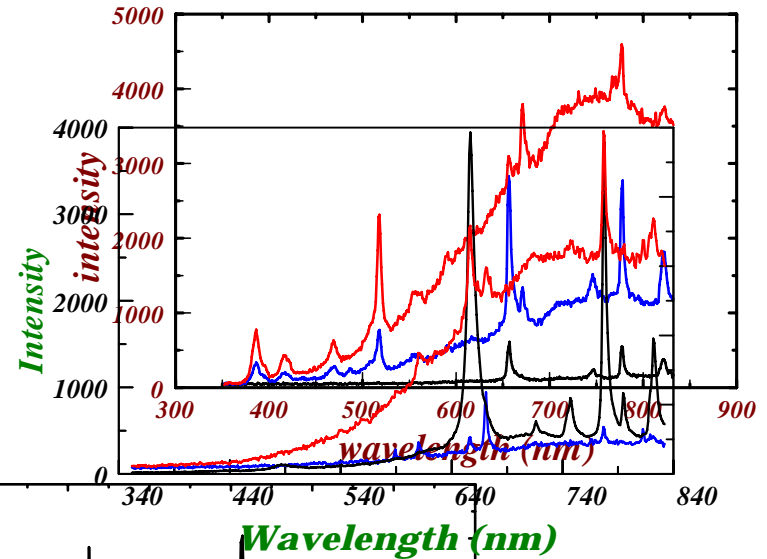
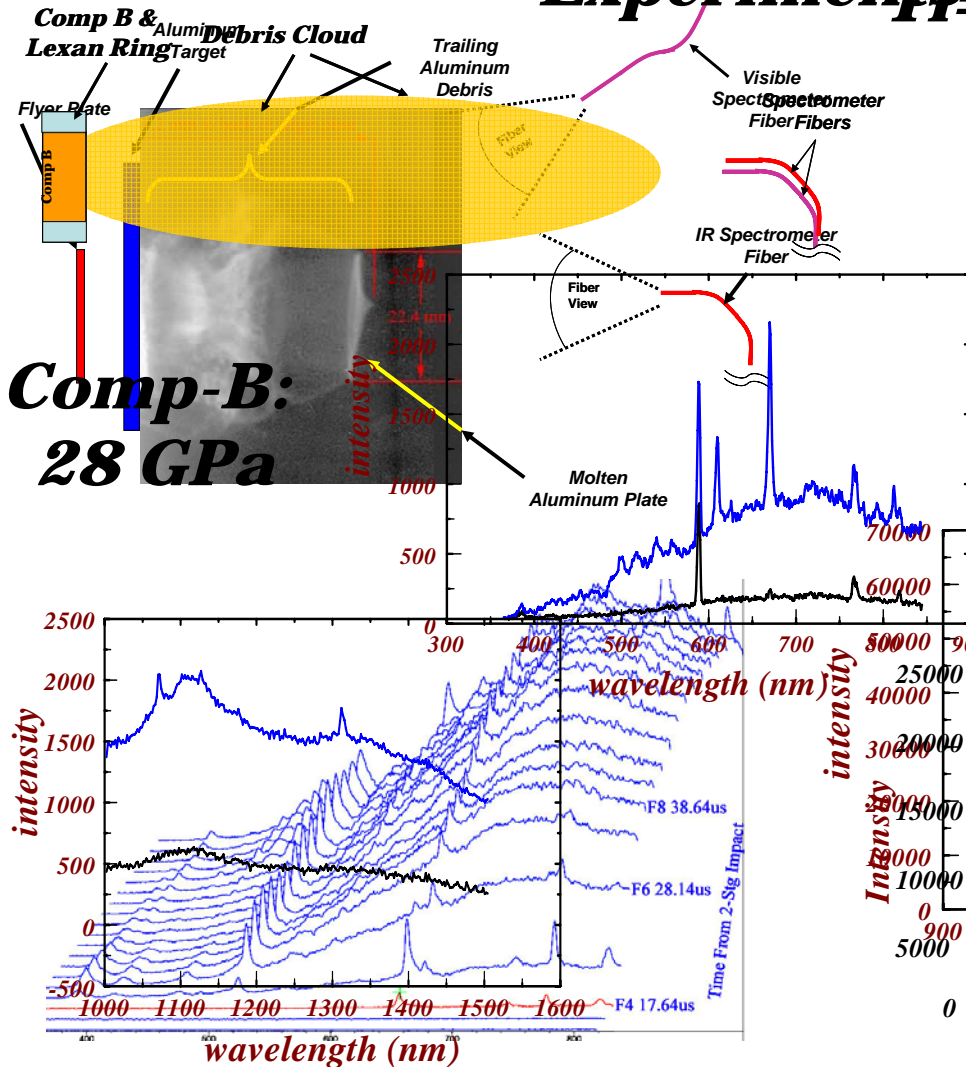
Spectrometer Calibration



Known spectrum of the lamp from NIST Tables, plotted with the measured calibration lamp data. The lamp is recorded through a spectrometer. The lamp is multiplied by a correction term. Determination of amplitude correction arrays are plotted as a function of wavelength and intensity. 950-1500 nm

Experimental Results:

Experiments on Comp-B at 28 GPa



**Comp-B:
130 GPa**



Extracting temperature :

***From our basic premise,
Planck's Law can describe
the spectral radiance as:***

$$I(\lambda, T) = \varepsilon \frac{2hc^2}{\lambda^5} \frac{1}{e^{c_2/\lambda T} - 1},$$

Assume ε , $2hc^2$ is a constant:

$$I(\lambda) = \frac{s}{\lambda^5 \left(e^{c_2/\lambda T} - 1 \right)}.$$

- ***For non-quantitative measurements of intensity, the factor 's' may be used as a scaling value that matches the blackbody shape of the observed signal***
- ***Recorded, corrected spectrum can be fit using non-linear least-squares routine with two adjustable factors: s and T***
- ***Assumptions: grey body, $\varepsilon(\lambda) = \text{constant}$ and thermal equilibrium is assumed***



Extracting temperature : Fitting functions

Wien's approximation for fitting function: 's' and 'T' will be used to estimate starting values

$$I \cong s e^{-c_2 / \lambda T} / \lambda^5 \cdot$$

's' and 'T' : choosing two distinct signal and wavelength values, 'I' and 'λ'

$$T \cong \frac{c_2 (1/\lambda_2 - 1/\lambda_1)}{\ln(I_1 / I_2) + \ln(\lambda_1^5 / \lambda_2^5)}$$

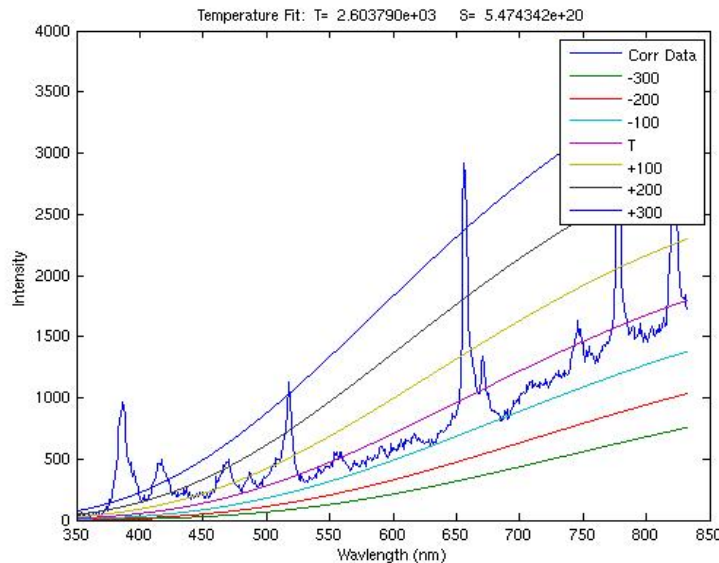
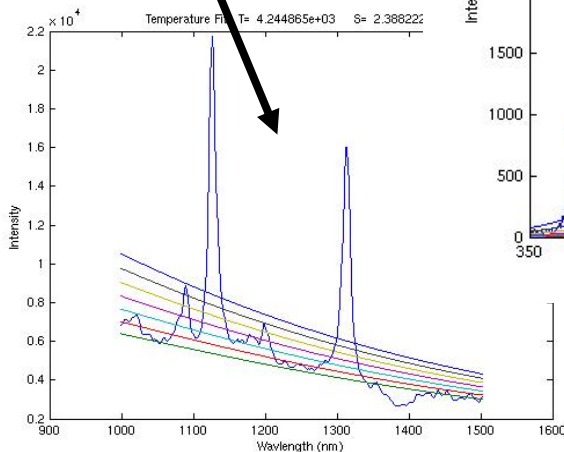
$$s \cong I_1 \lambda_1^5 e^{c_2 / \lambda_1 T}.$$

- ***These curves give way of visualizing the sensitivity of the fit:***
 - > ***If the curves easily envelop the corrected data—fit is sensitive—estimate is probably reasonable***
 - > ***If the curves do not encompass the spectral lines—fit is insensitive and will not capture a good temperature estimate.***
- ***One can interpret these curves as error bars: 300K for reasonable approximations***

Extracting temperature : fitting the data

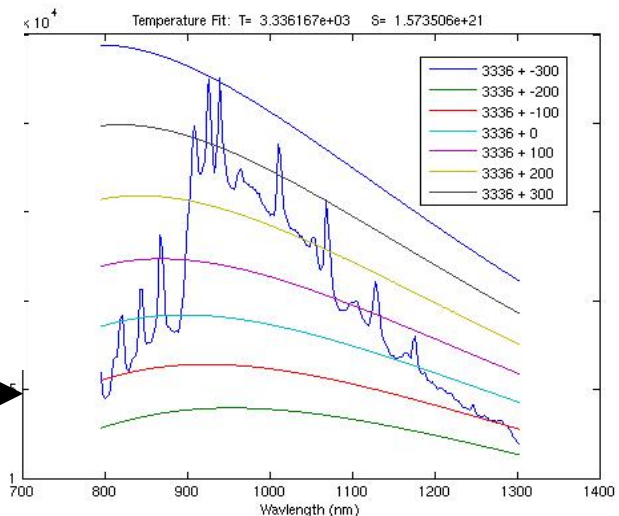
- These curves give way of visualizing the sensitivity of the fit:
 - > If the curves easily envelop the corrected data—fit is sensitive—estimate is probably reasonable
 - > If the curves do not encompass the spectral lines—fit is insensitive and will not capture a good temperature estimate.
- One can interpret these curves as error bars: 300K for reasonable approximations

*Poor
Estimate*



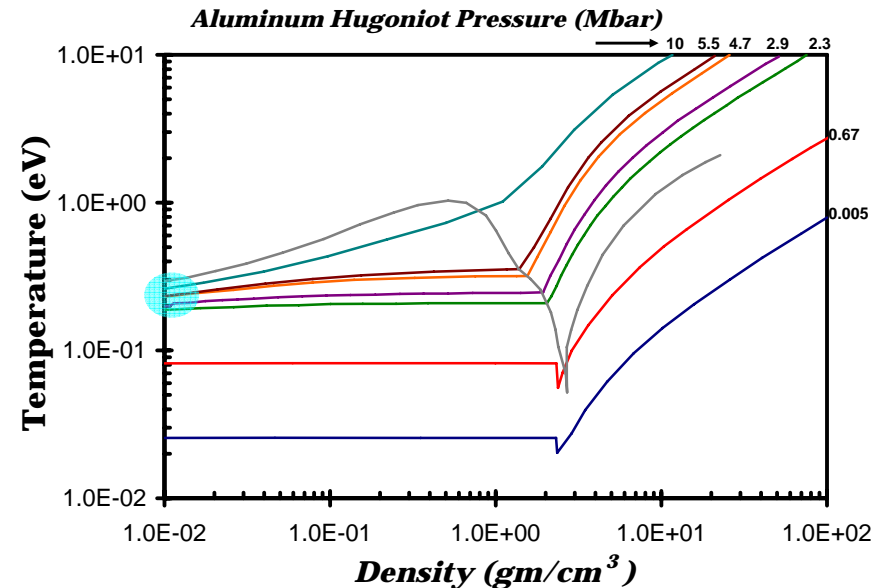
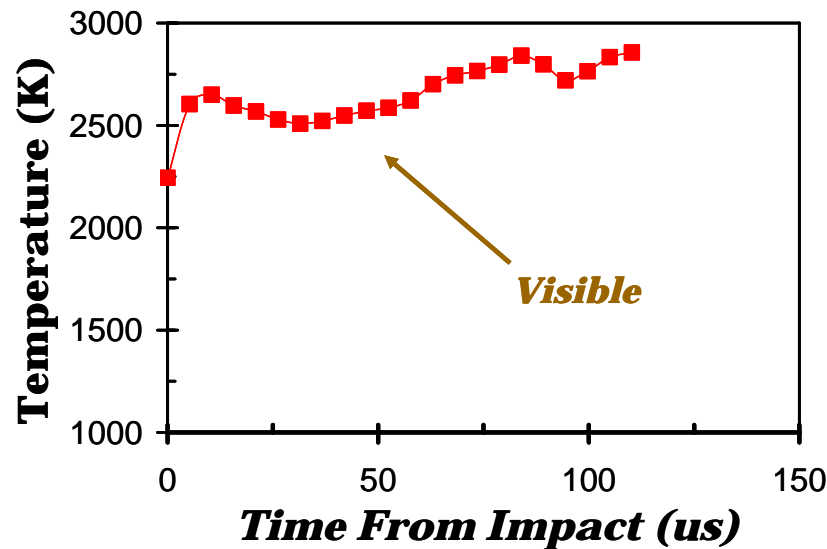
*Comp-B:
130 GPa*

*Al:
200 GPa*



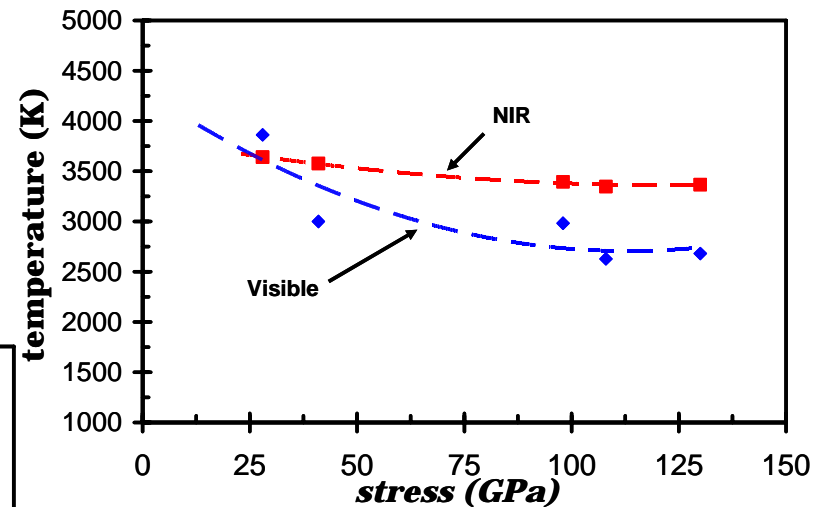
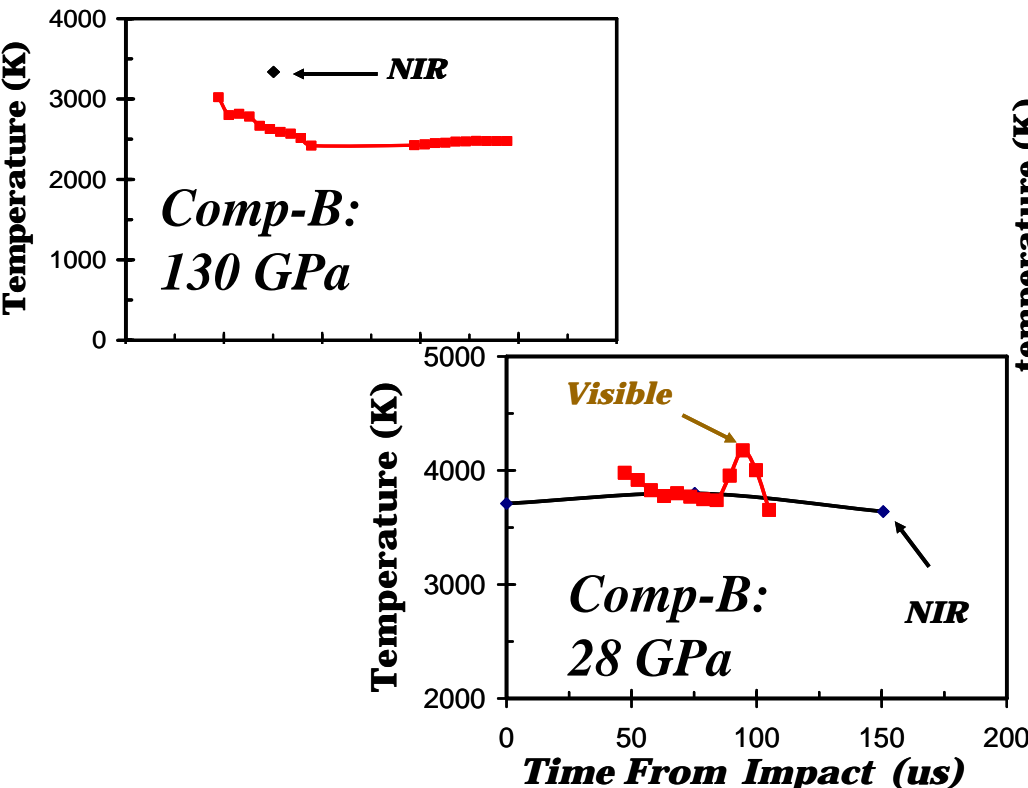
Temperature Estimates: Aluminum

- **Aluminum:** not trivial, greater uncertainty
 - > **Large magnitude of spectral radiance**
 - > **Presence of large atomic emission lines**
- **Temperature history suggests constant temperature**
 - > **Consistent with a solid—cooling is not as rapid**
- **Reasonable results when compared to phase diagram**



Temperature Estimates: *Comp-B*

- Using dissociation energy (E/E_v): 0.5 – 5 for tests
 - > Tests represent experiments that undergo detonation, and those where detonation is overdriven (or delayed)
- Lower pressure experiment produced higher temperature
 - > Same trend for tests at 41, 98, and 108 GPa
- Released from higher pressure: velocity is faster





Conclusions:

- ***First temperature measurements/estimates of expansion products from materials shocked to high temperatures***
 - > ***Ascertained temperatures are estimates, results appear reasonable***
- ***Lower pressure experiments produced higher temperature***
 - > ***Encouraging results based on simplistic assumptions***
- ***Further research is necessary to validate assumptions.***
- ***Experiments providing early time measurements: shocked temperatures***
- ***Combination of pressure, density measurements of shocked and expanded materials: potential to determine phase diagrams.***



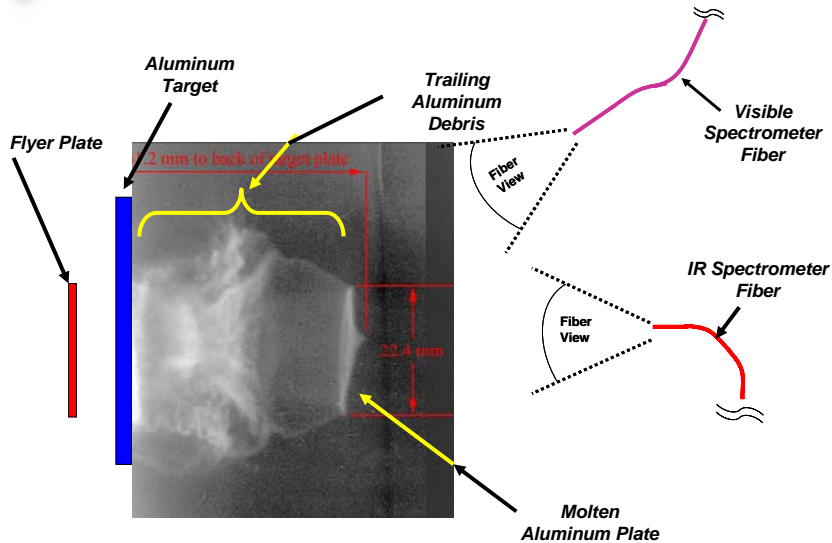
Acknowledgements:

- **William G. Breiland: Sandia National Laboratories**
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- **Tom F. Thornhill: Ktech Corporaton**
- **John R. Martinez: Ktech Corporation**

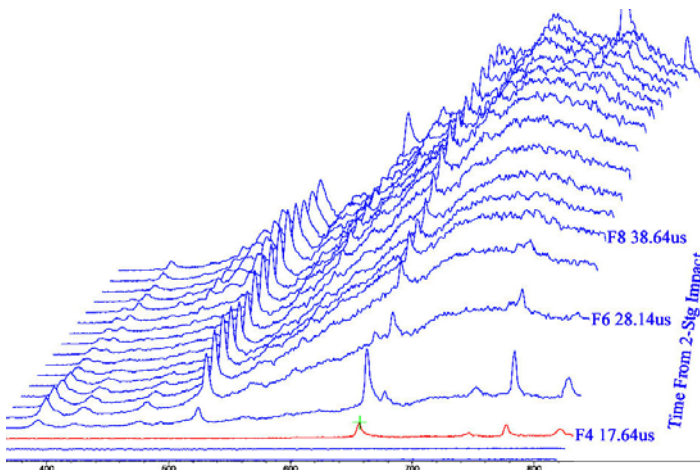
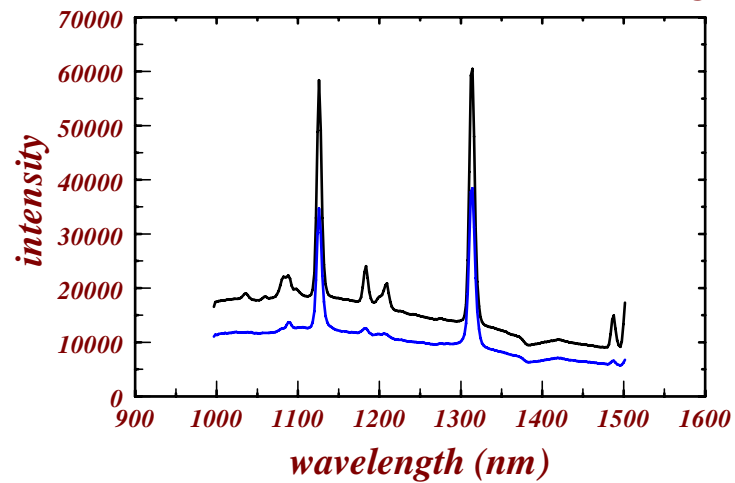
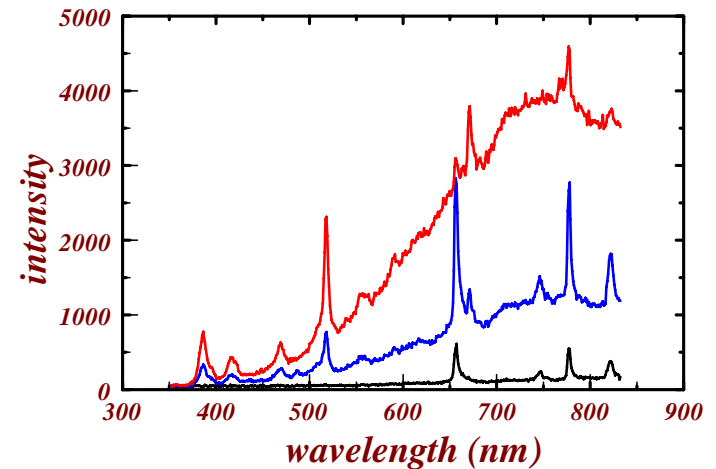


End:

Dup. slide 7 :



$Ti \Rightarrow Al$, > 11 km/s, 200 GPa



Dup. slide 7 ::

Experiments on Comp-B up to 130 GPa

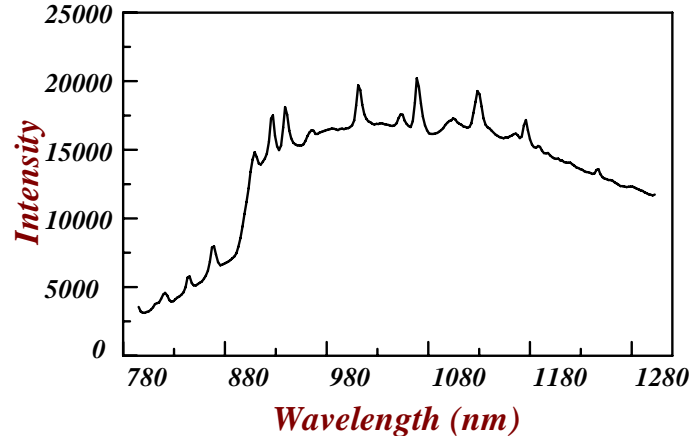
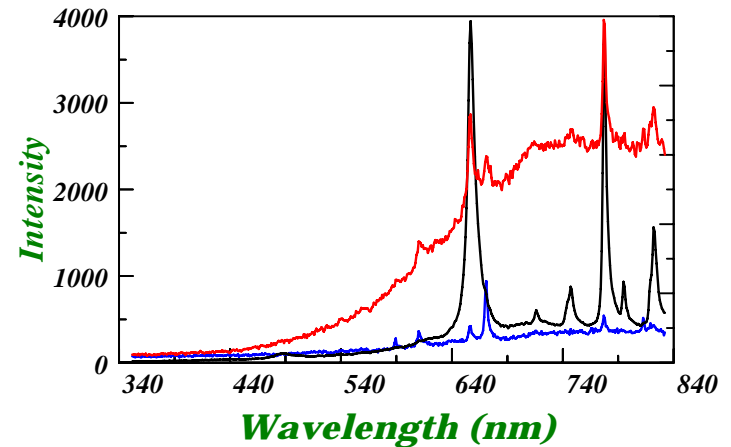
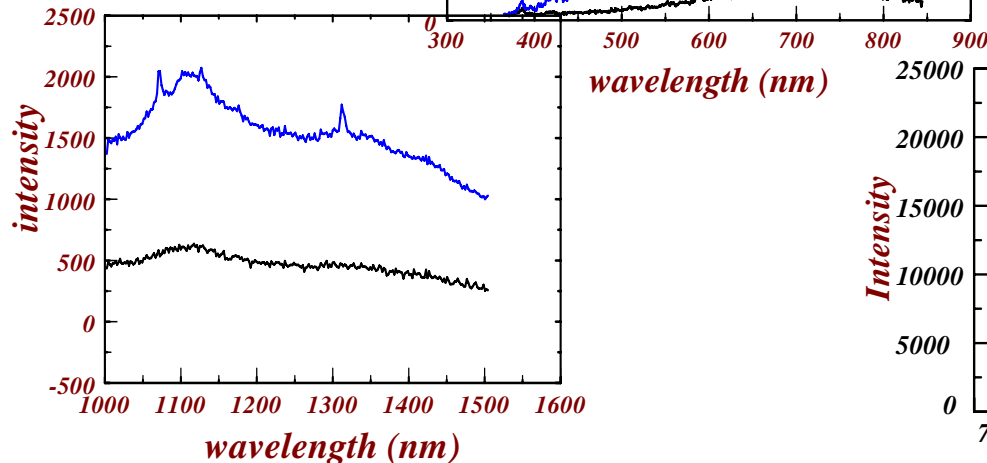
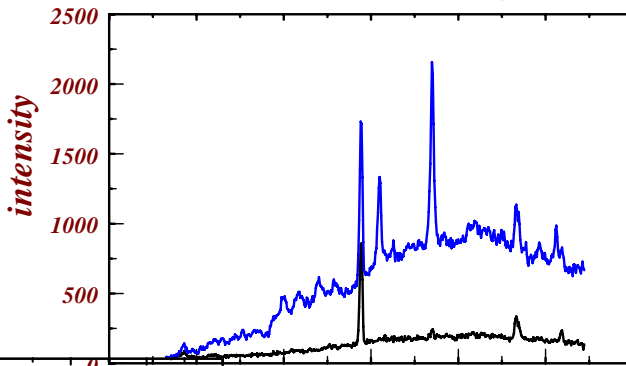
Comp B &
Lexan Ring

Debris Cloud

Comp B

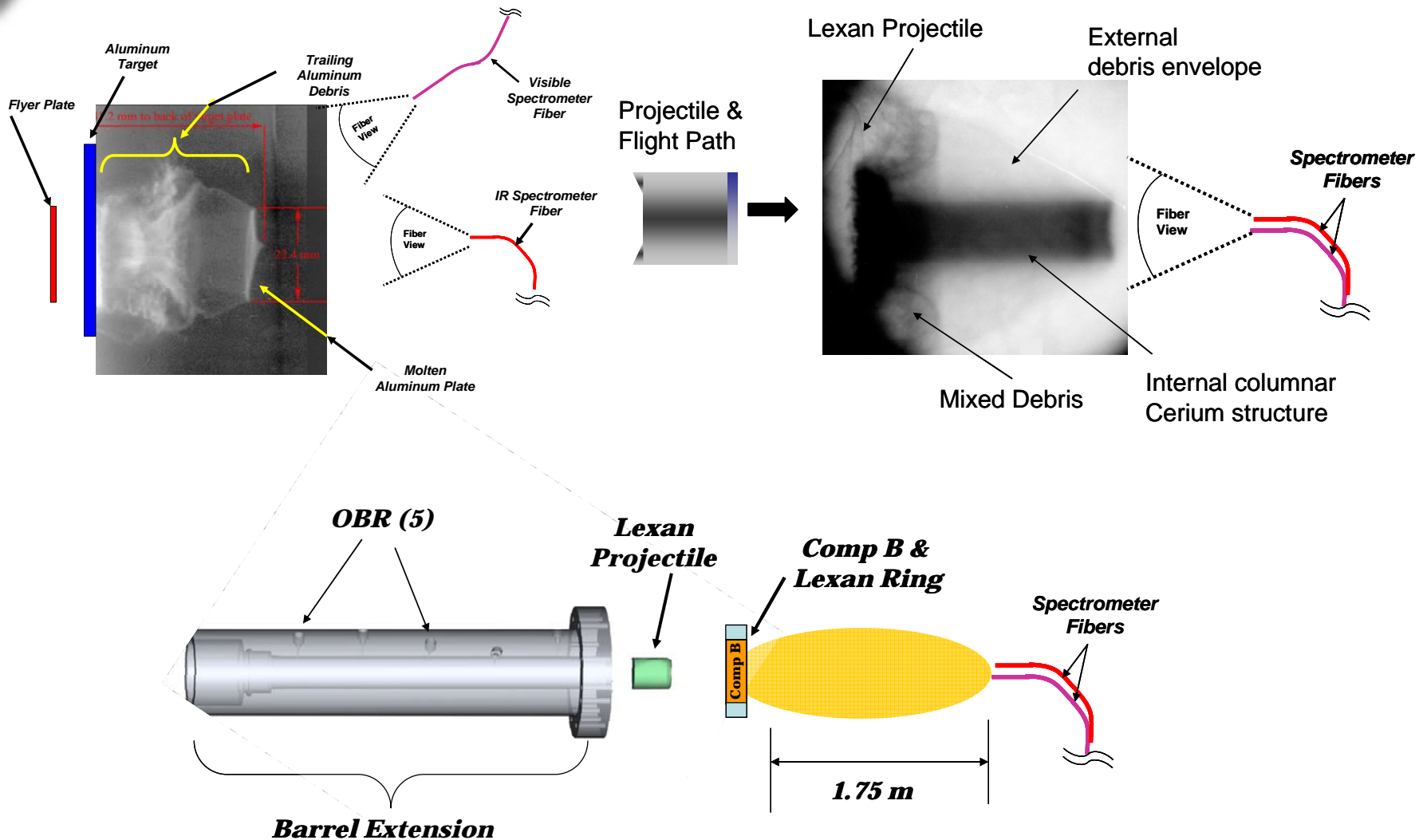
Spectrometer
Fibers

**Comp-B:
28 GPa**

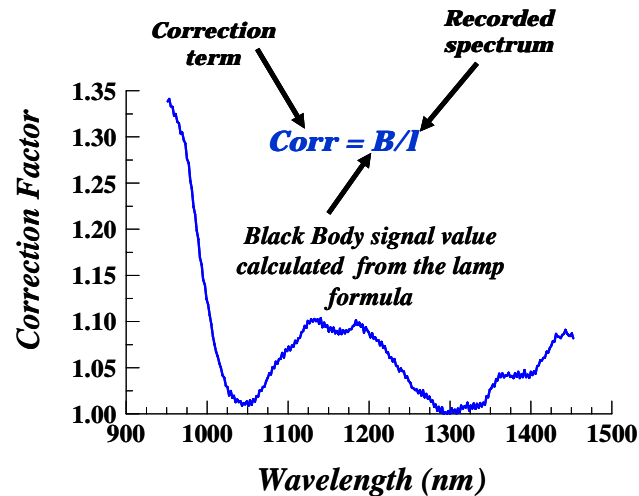


**Comp-B:
130 GPa**

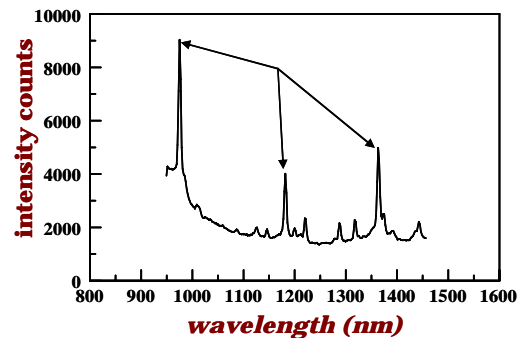
Duplicate slide



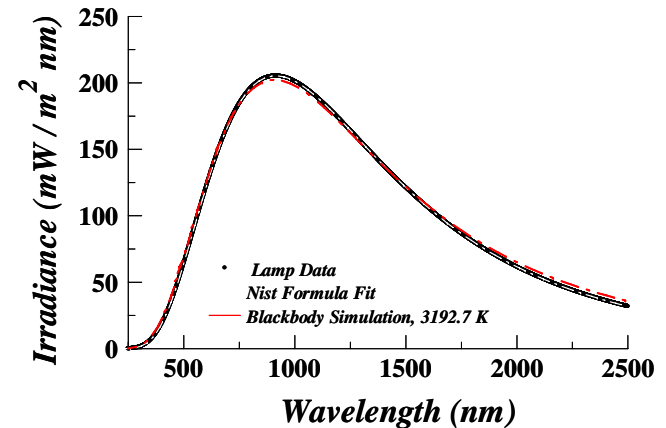
Dup. slide 6 :



Determination of amplitude correction term from any spectrum obtained from 950-1500 nm



A krypton lamp source provides three distinct, relatively high intensity lines, (for a specific spectrometer grating),



Known spectrum of the lamp from NIST Tables, plotted with the measured calibration lamp spectrum multiplied by correction term.
Note that if the measured calibration lamp spectrum is multiplied by the correction term, we simply recover the blackbody curve, as well we should.