



New Perspectives on the NEO Threat and Other Reasons to Consider Asteroid Deflection/Fragmentation

SAND2008-7238C

Mark Boslough
Sandia National Labs
Albuquerque, NM

Asteroid Deflection Research Symposium 2008 October 23-24, 2008

Red Storm simulation

New perspectives on the impact threat

1. Some nuclear deflection strategies are better:
Multiple low-yield shallow bursts are best.
2. Low-altitude airbursts dominate residual threat:
Mitigation should be for ~hundred m asteroids.
3. Recent controversies in impact science:
Credibility of community requires solid research.
4. Total asteroid threat is diminishing:
Mitigation not justified for NEO protection alone.
5. Better reasons to consider asteroid deflection:
Geo-engineering and space resources.

1. Nuclear deflection

- Best way to couple energy and momentum is buried nuclear burst (direct consequence of scaling laws)
- For a given total yield, multiple bursts are better than single burst (direct consequence of scaling laws)
- Non-weaponizable impact-triggered fission device is possible in principle
- Impact-triggered fusion device would make multiple low-yield shallow bursts feasible

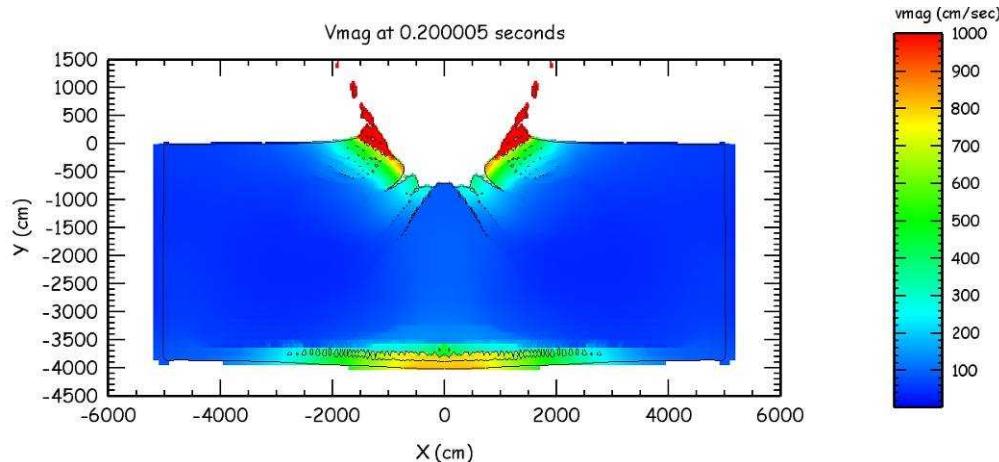
Numerical simulations of momentum coupling

- Three prongs of analysis: Numerical simulations, analytic models, scaled experiments.
- 2D parameter studies, sensitivity analysis, and intuition (require workstation computing).
- 3D models quantify effects of heterogeneities, strength, fracture, and spallation that are not amenable to either scaled experiments or analytical models (require capability computing).

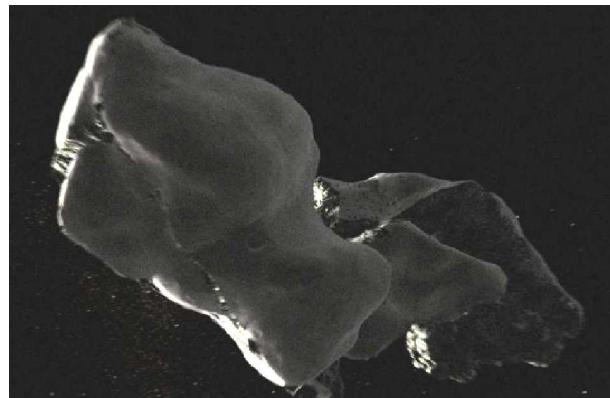
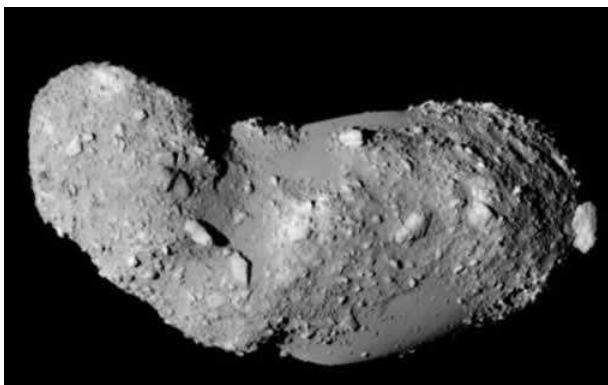
2D workstation-class problem

Example:

1 kiloton surface burst in an intermediate-strength 100-meter diameter dunite disk ($m=10^9$ kg) couples 3×10^8 kg·m/s, but spalls material off the back surface.



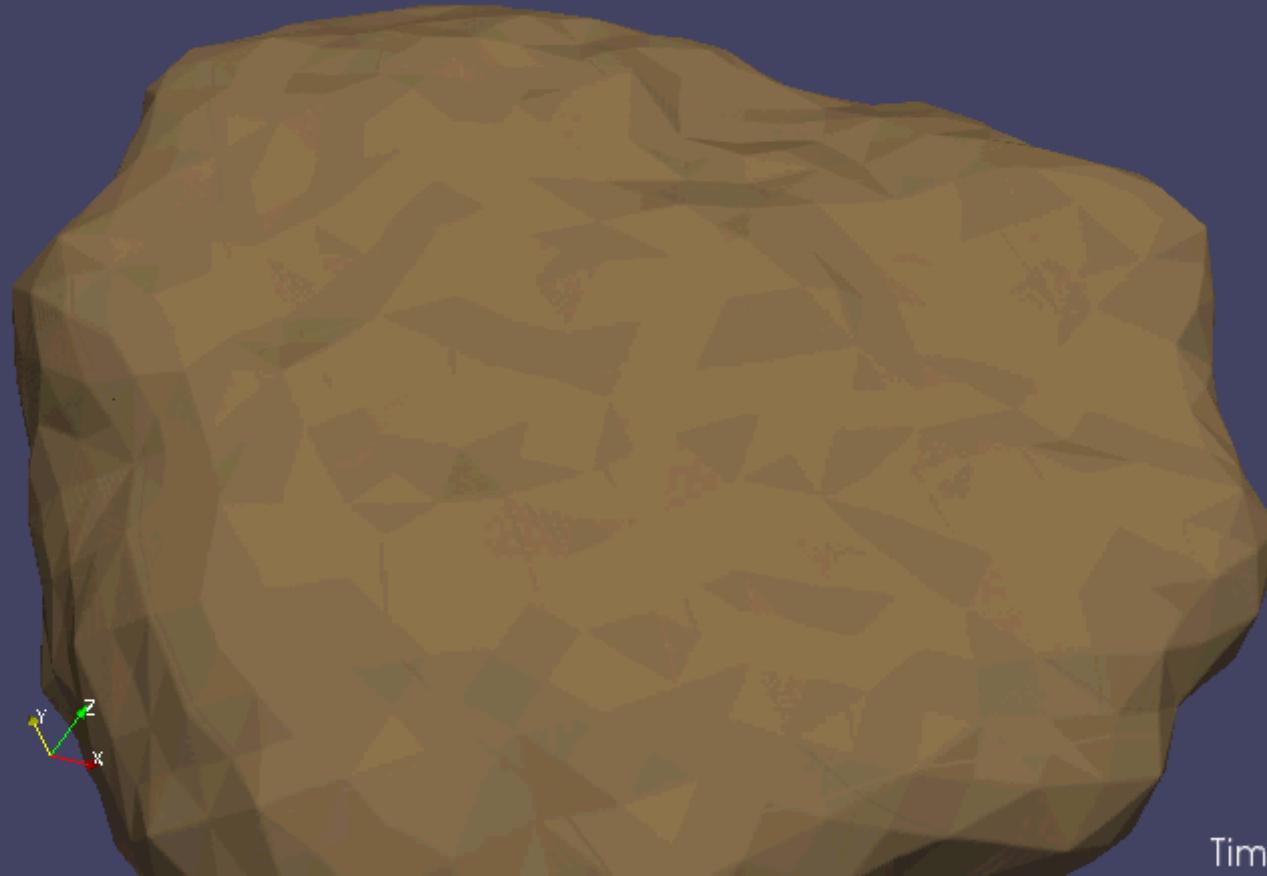
3D capability-class problem



Example:

For known 3D geometry, internal structure of Itokawa must be assumed. Over what range of parameters can momentum be coupled to Itokawa without breaking it?

Golevka deflection: 100 kt explosion, 5 m deep 1 billion computational cells, Red Storm

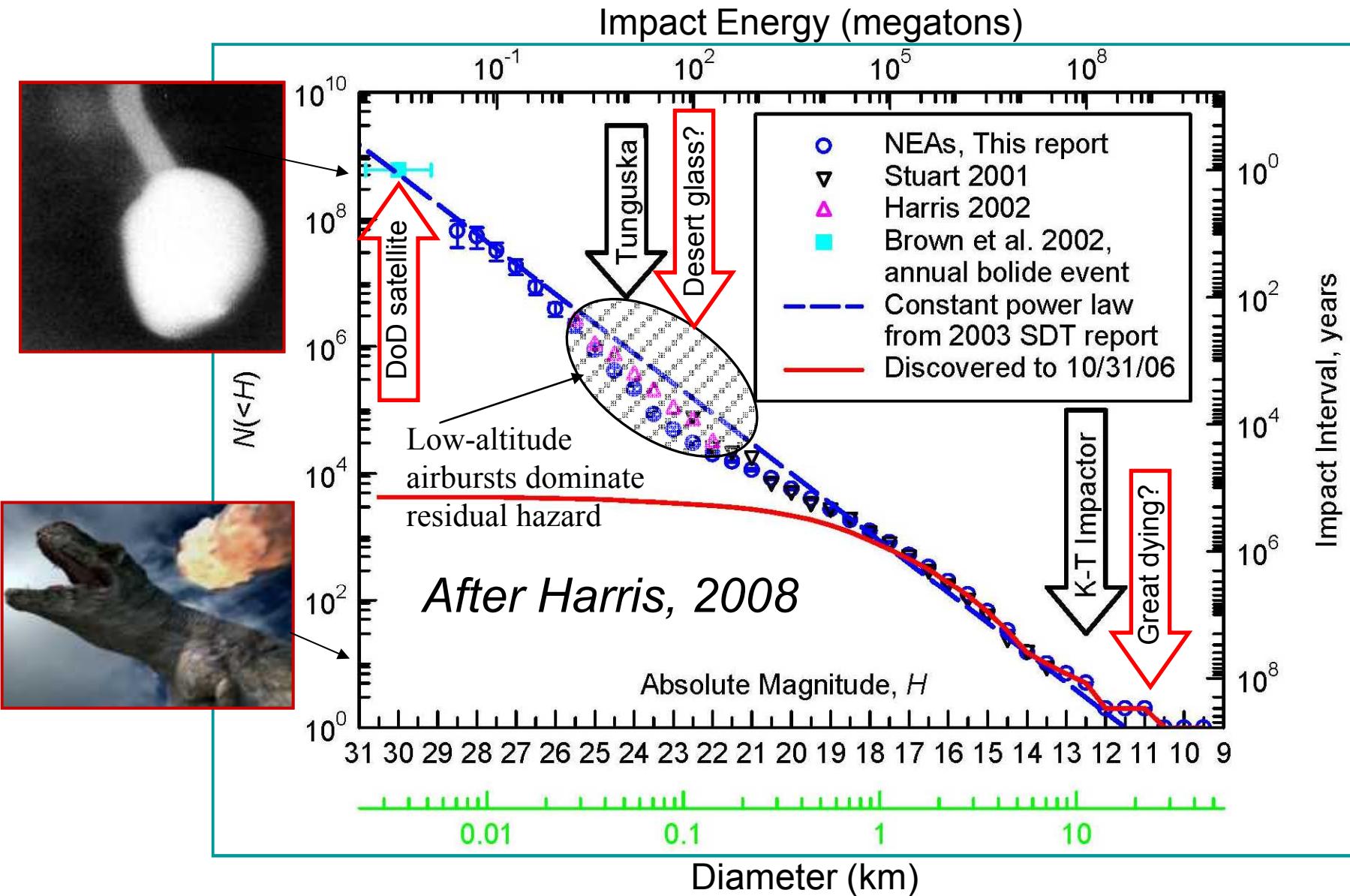


Time = 0.00000

2. Low-altitude airbursts

- The relative threat from low-altitude airbursts (LAAs) is increasing
- Our understanding of LAAs is improving
- The next destructive NEO will almost certainly be a LAA
- “100/100/100 event”: (~100 m, ~100 Mt, ~1/100 chance this century)
- 100/100/100 event will dominate threat after survey is complete
- Time for technology development is similar to threat reduction time
- There are other reasons to deflect/fragment small asteroids
- Technology should be focused small (~few hundred meter) asteroids

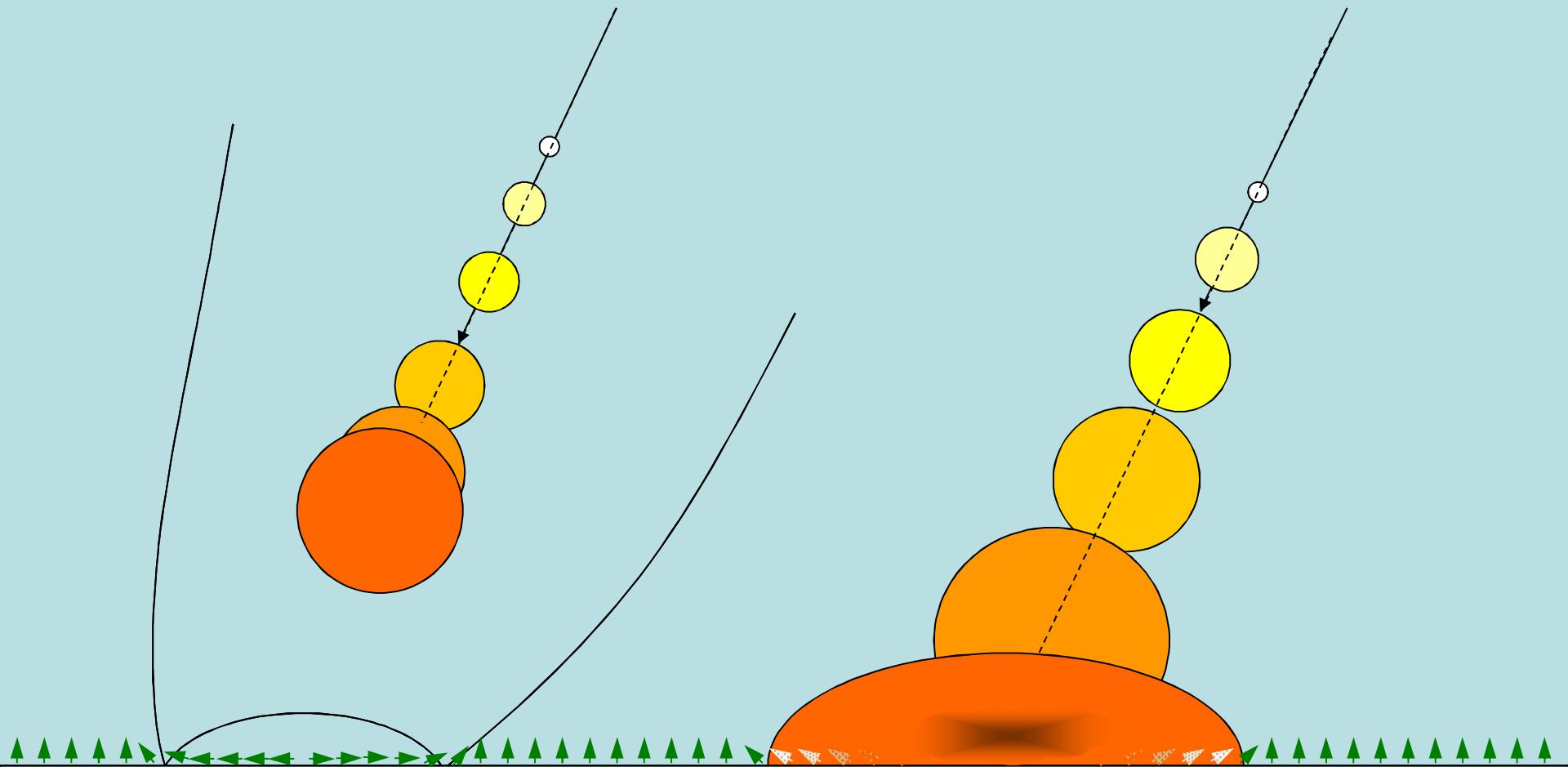
The nature of the impact threat is changing



There are two types of Low-Altitude Airburst

1. The explosion generates a fireball that descends rapidly but does not reach the ground. Most of the damage at ground level is mechanical, due to the blast wave. This occurs for explosions between about 1 and 10 Mt, and may occur on time scales of several hundred years. The only known example is 1908 Tunguska event.
2. The fireball is much larger and descends all the way to the surface. The damage is dominated by high-temperature thermal radiation. The threshold yield is about 10 Mt, depending on other parameters, and they recur on time scales of several thousand years. The best putative example is the Libyan Desert Glass event, 29 million years ago.

There are two types of Low-Altitude Airburst



Type 1: Tunguska

Type 1: Libyan Desert

Type 1 airburst simulation: 5 megaton

Type 2 airburst simulation: 15 megaton

Type 1 LAA: “Tunguska-Type”



Type 1 LAA: “Tunguska-Type”



Consequences of Type 1 airburst



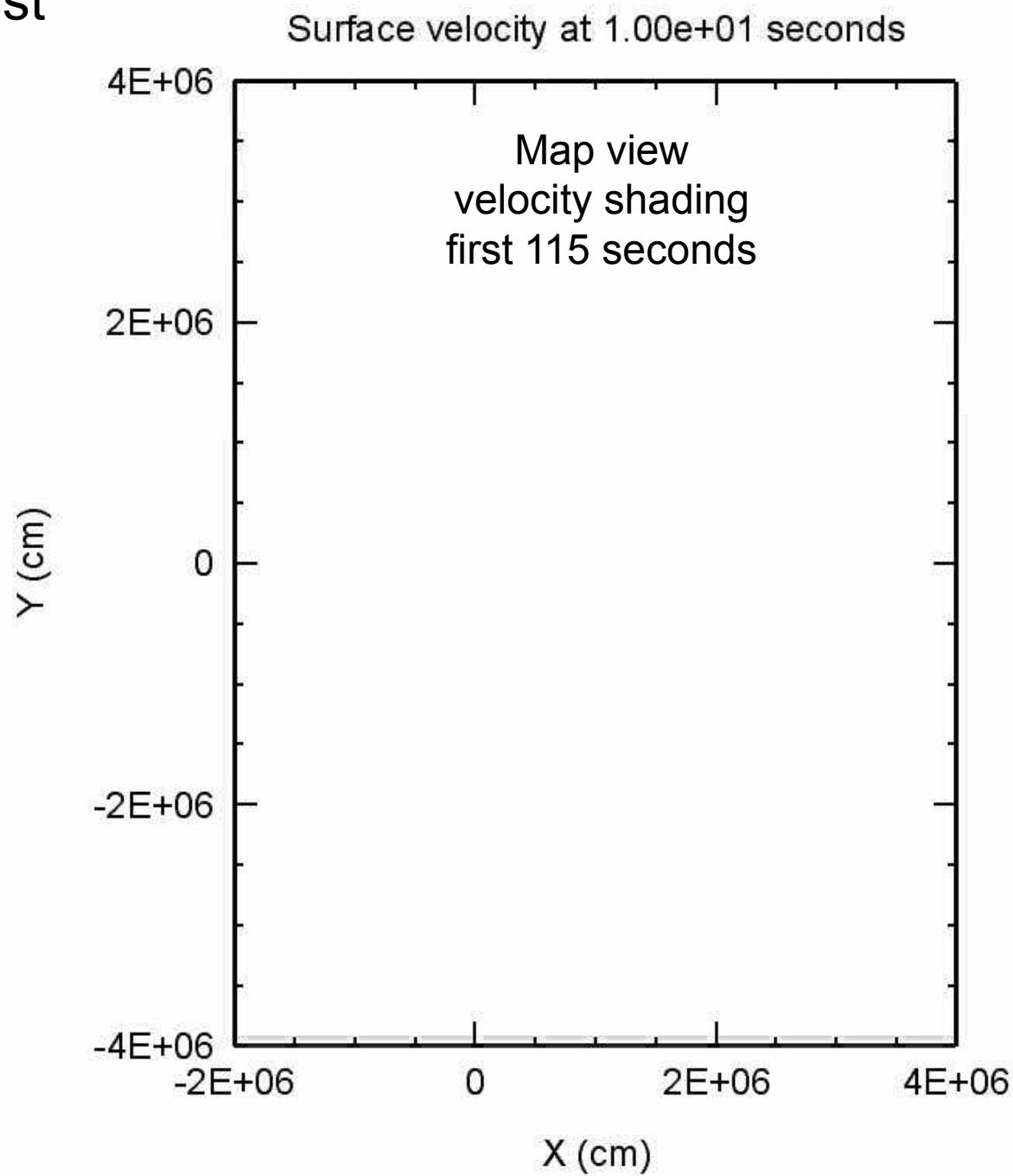
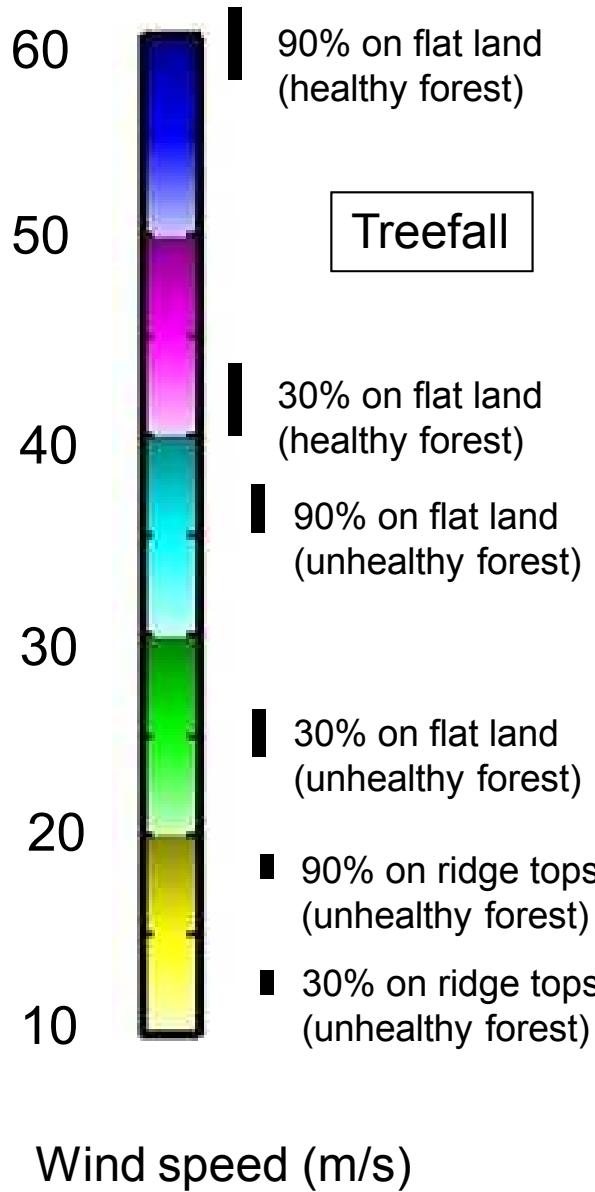
Krinov, 1963



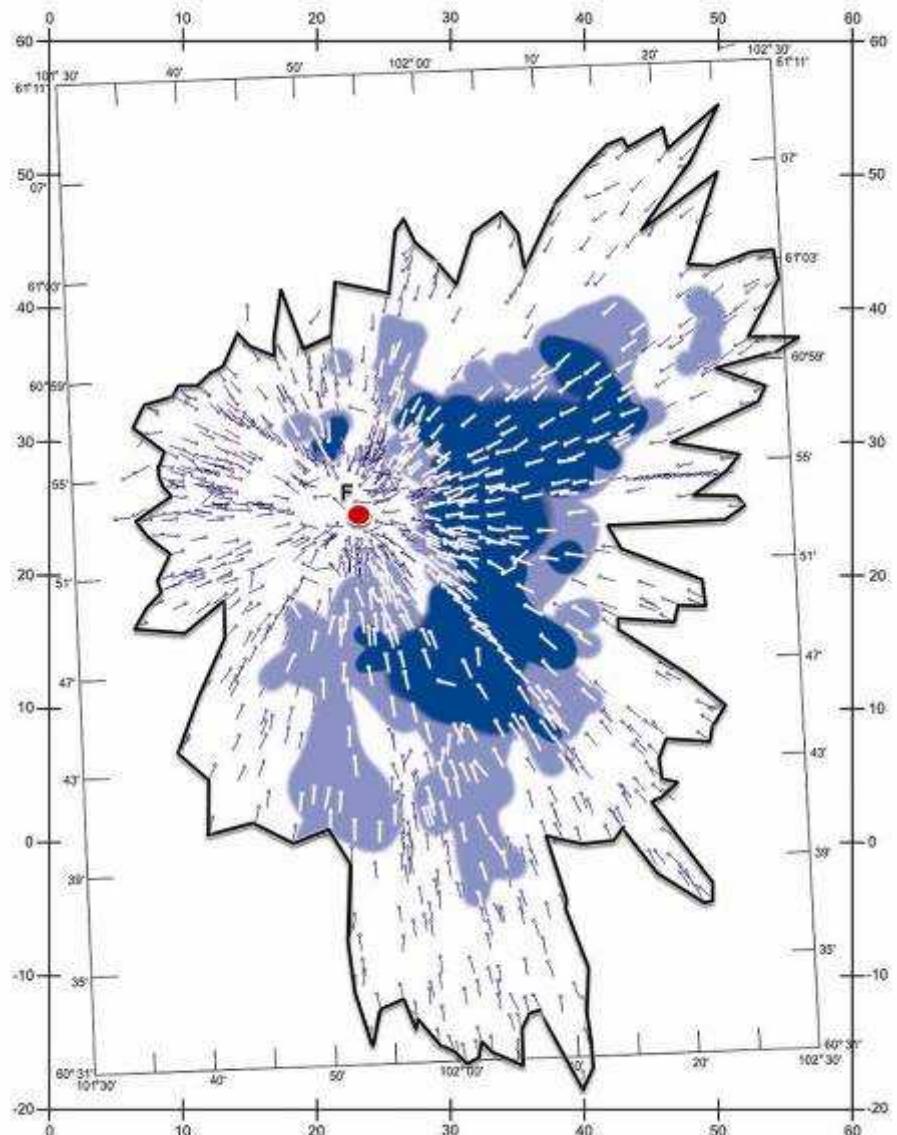
Fig. 485 — THE SIBERIAN FOREST DEVASTATED BY THE BLAST FROM THE METEORITE OF 30 JUNE 1908.



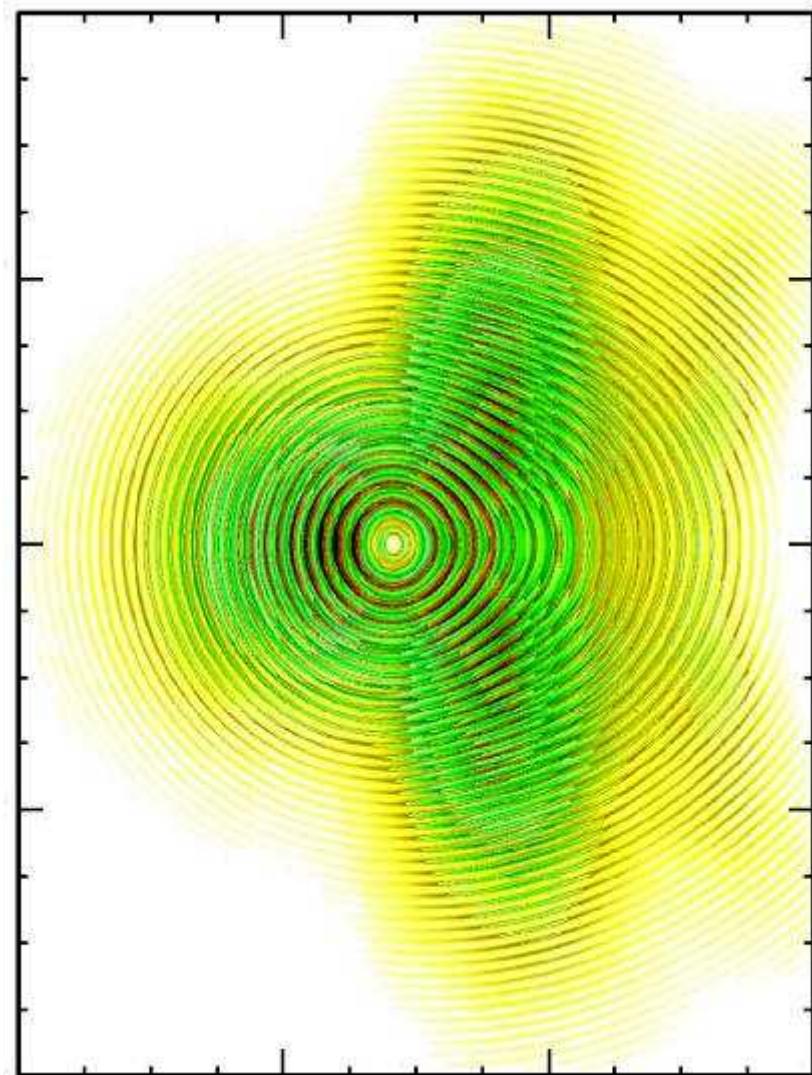
Modeling Type 1 airburst



5 Mt explosion at 12 km above surface, 35° entry angle



Tunguska treefall map (Longo et al, 2005)



Wind speed map (this study)

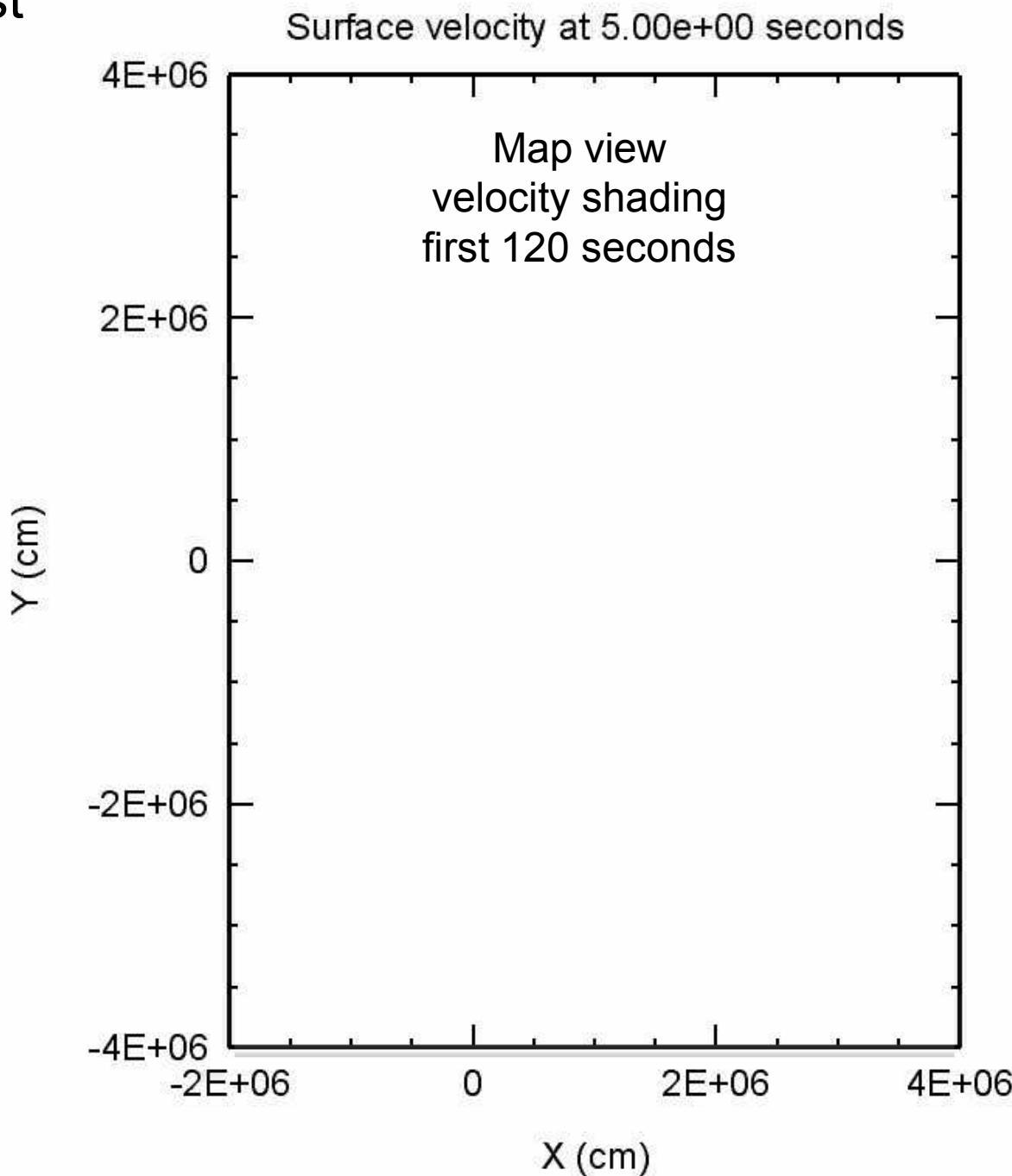
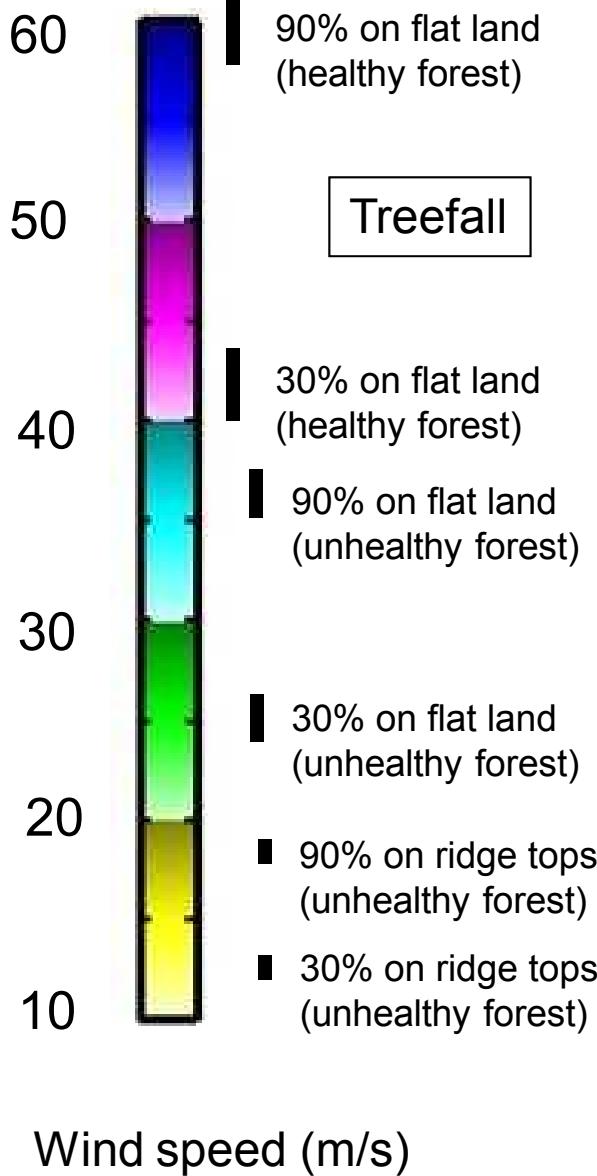
Type 2 LAA: “Libyan-Desert-Type”



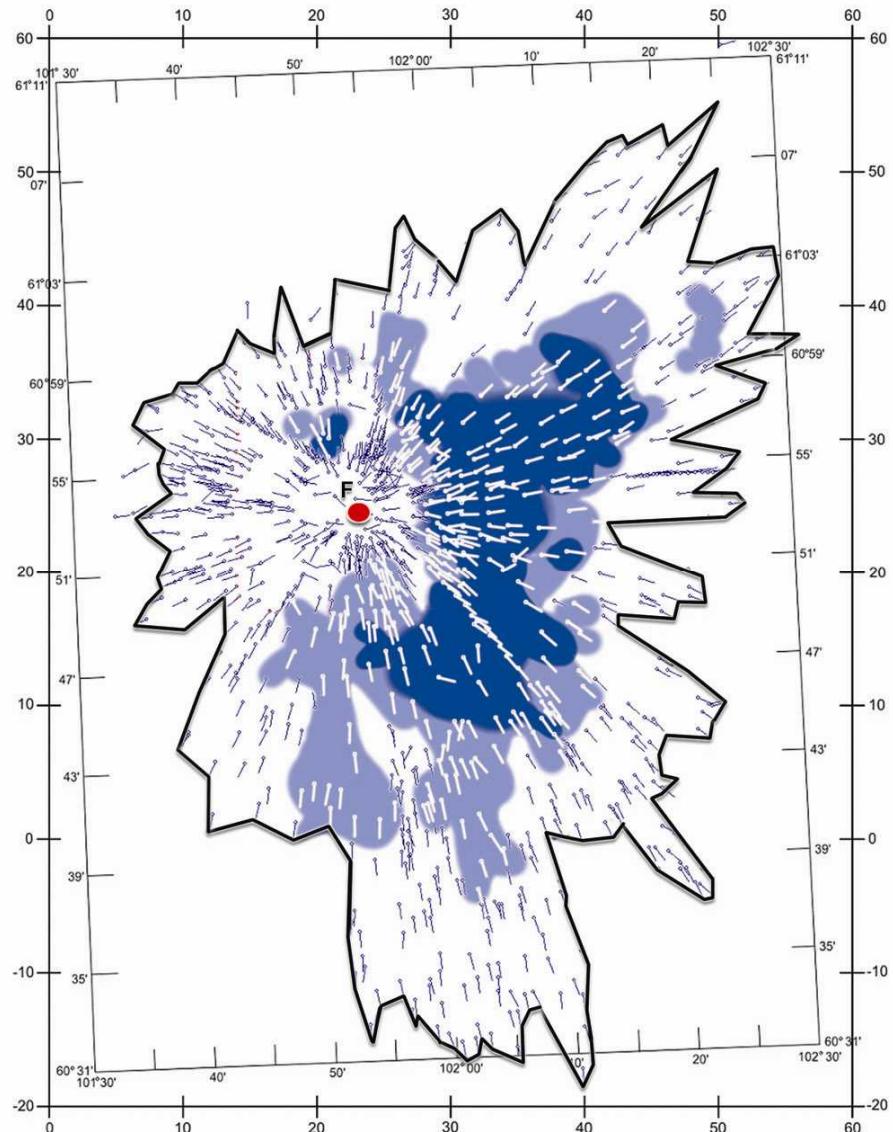
Consequences of Type 2 airburst



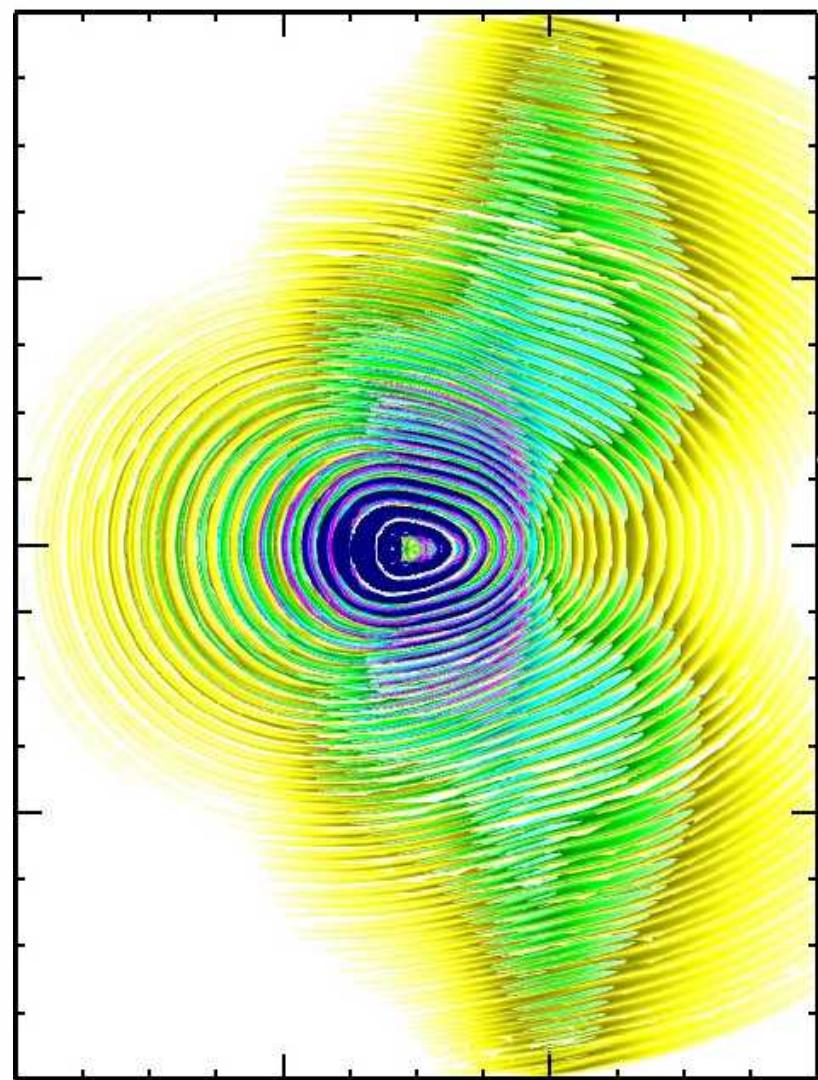
Modeling Type 2 airburst



15 Mt explosion at 18 km above surface, 35° entry angle



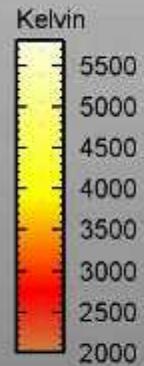
Tunguska treefall map (Longo et al, 2005)



Wind speed map (this study)

100 Mt explosion, 90° entry angle

Temperature at 3.56 seconds



1 km

Temperature shading

100 Mt explosion, 90° entry angle

Wind Speed at 300 seconds



Wind speed shading

Animated version for National Geographic
documentary “Ancient Asteroid”

Tutankhamun's Fireball (Ancient Asteroid)

Meteor Crater impact was a low-altitude airburst

Less than 30% of the initial energy made it to the surface (Melosh, 2005)

9.0 Mt initial KE, 2.5 Mt formed crater, 6.5 Mt explosion in air



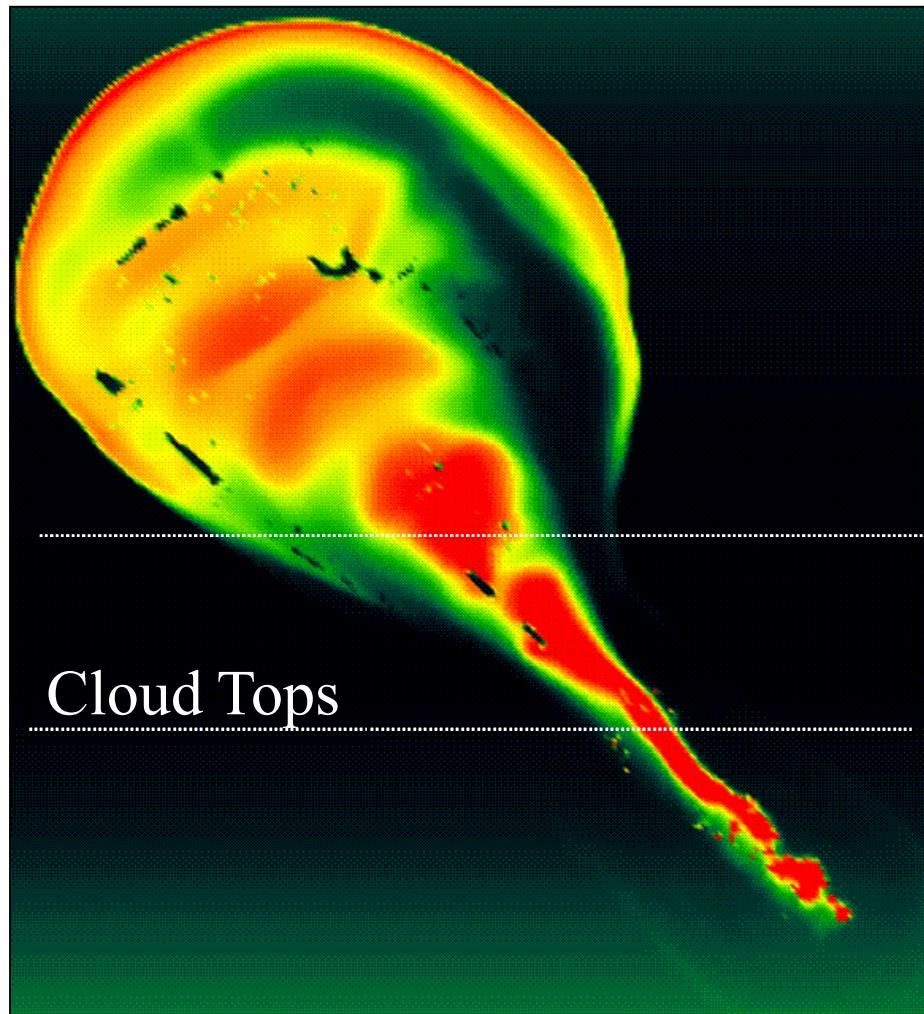
The physics of airbursts



© 1993 Michael Carroll

Shoemaker-Levy 9 comet crash: Jupiter, 1994

Shoemaker-Levy 9 impact, July 1994



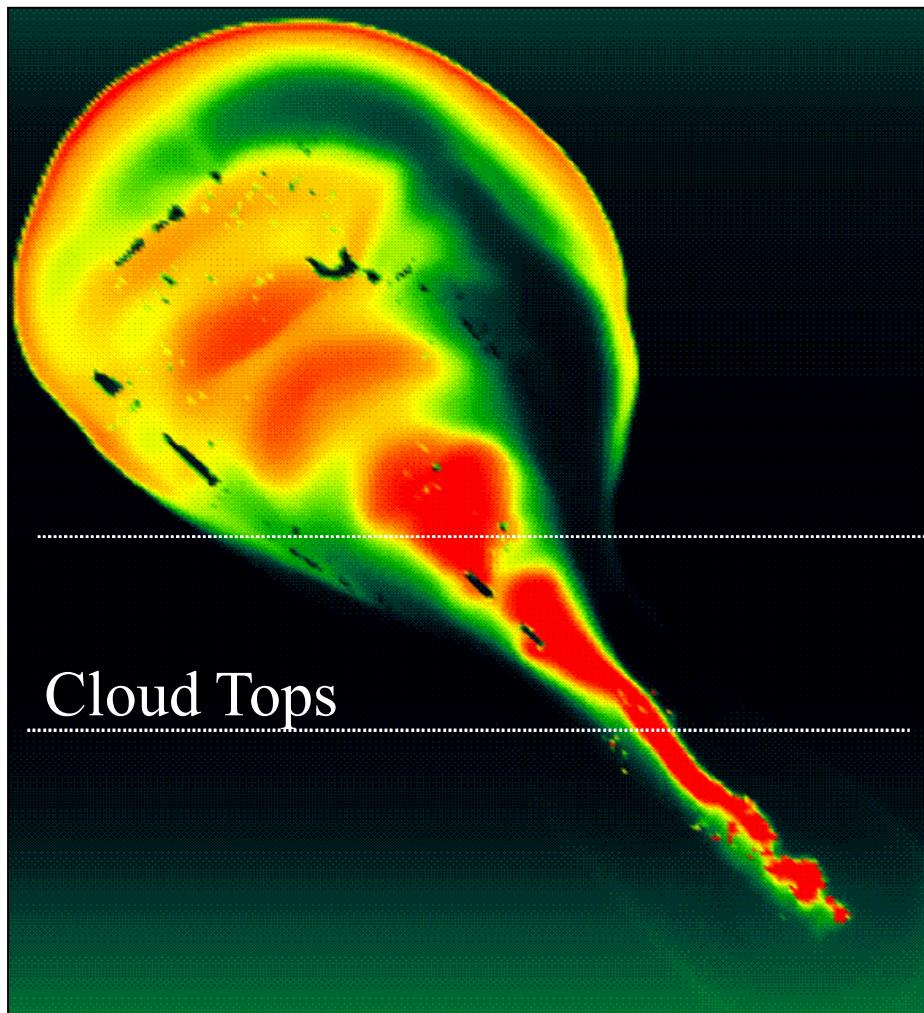
Visible From Earth



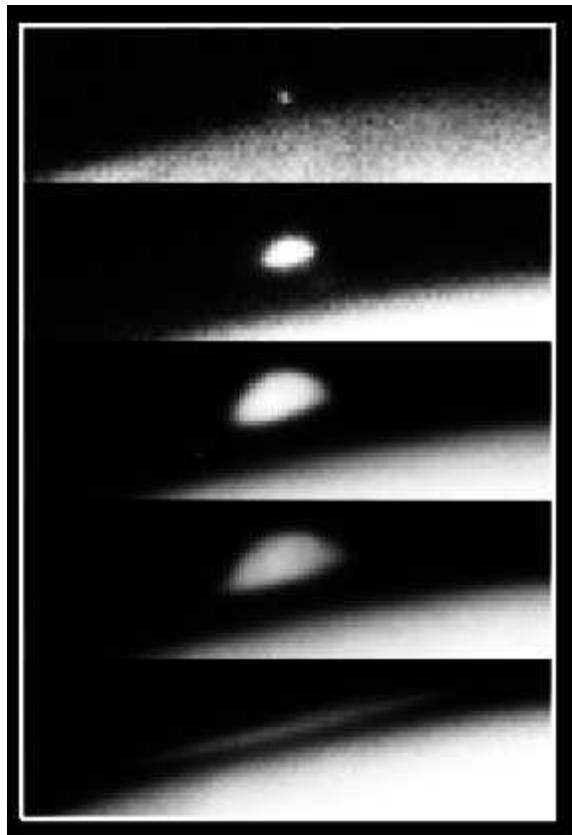
Behind Jupiter

This is a robust result

Shoemaker-Levy 9 impact, July 1994

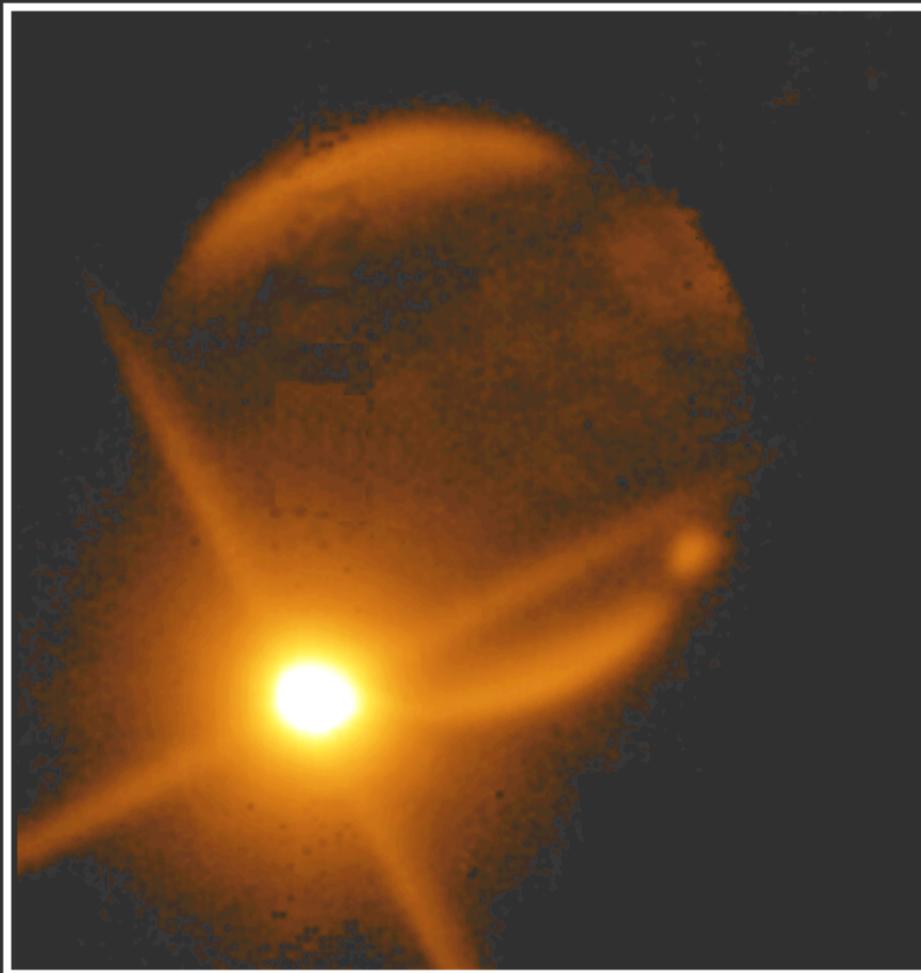


Impact G



This is a *validated* result

View from Earth



Shoemaker-Levy 9 comet crash: Jupiter, 1994

Plumes and line explosions on Earth



Annals of
the New York Academy
of Sciences
Volume 822

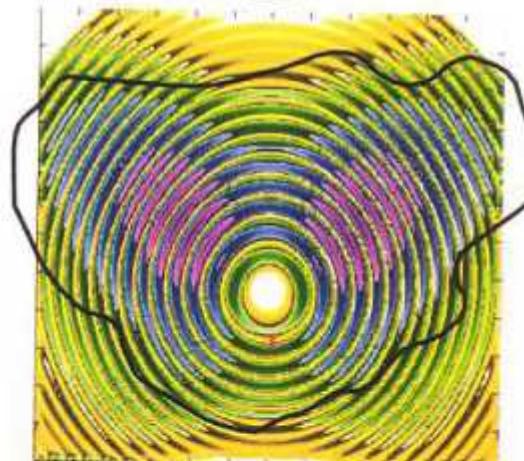
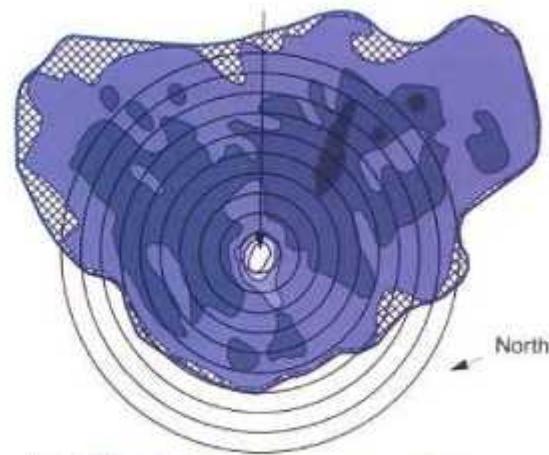
Near-Earth Objects

The
United Nations
International
Conference

Editor
John L. Remo



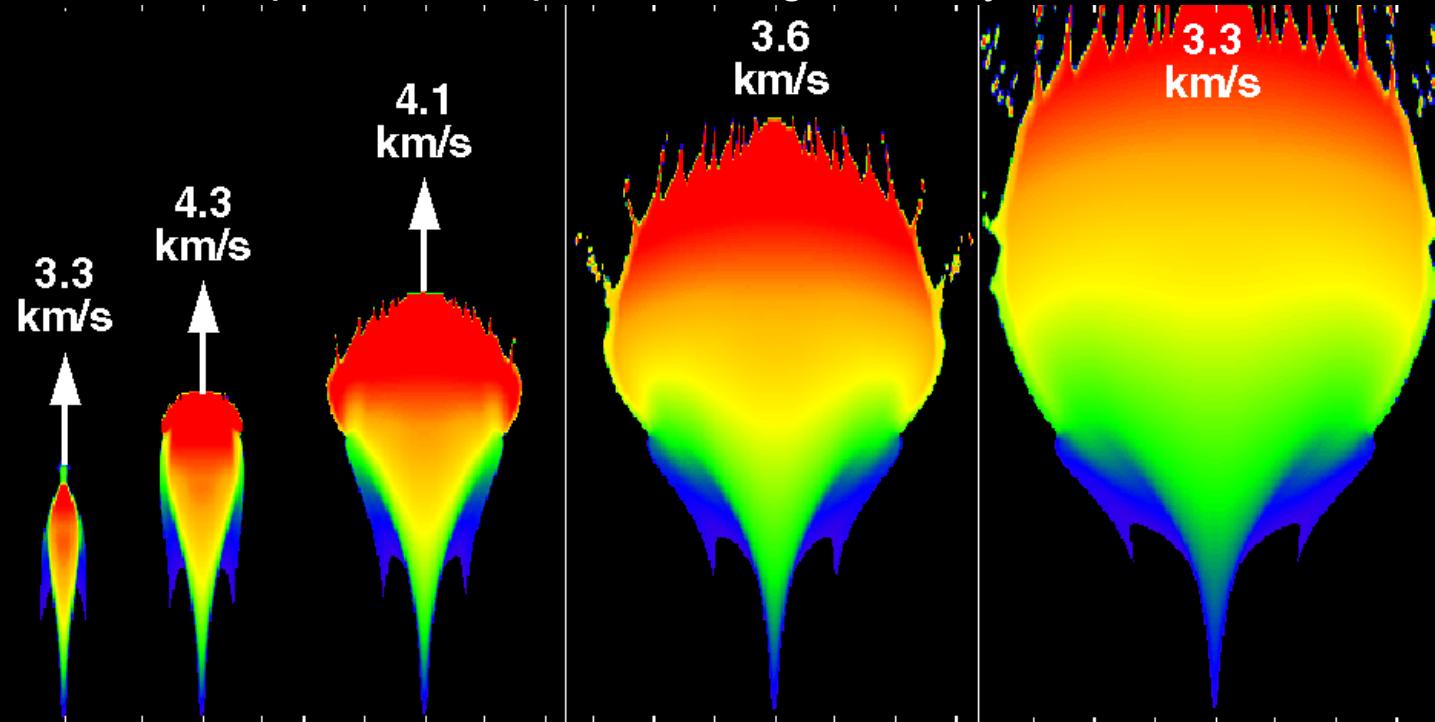
Treefall pattern at Tunguska



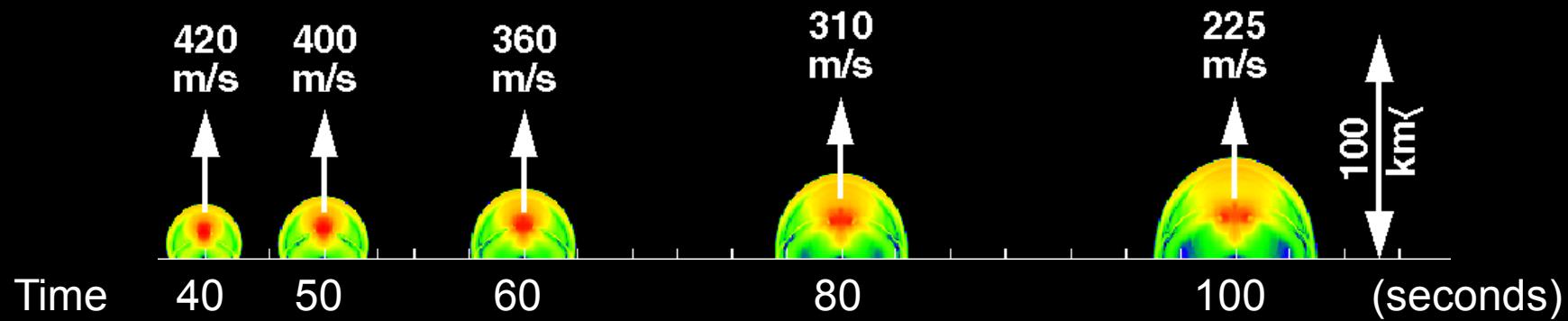
12 Mt impact at 55°

FIGURE 16. Comparison of butterfly-shaped treefall pattern at Tunguska [17] to map of peak surface windspeeds calculated from the collision of a 12 megaton object entering the atmosphere at an oblique angle, 55° from the zenith.

3 MT impact source produces high velocity “ballistic fireball”



3 MT explosion source produces high velocity “buoyant fireball”



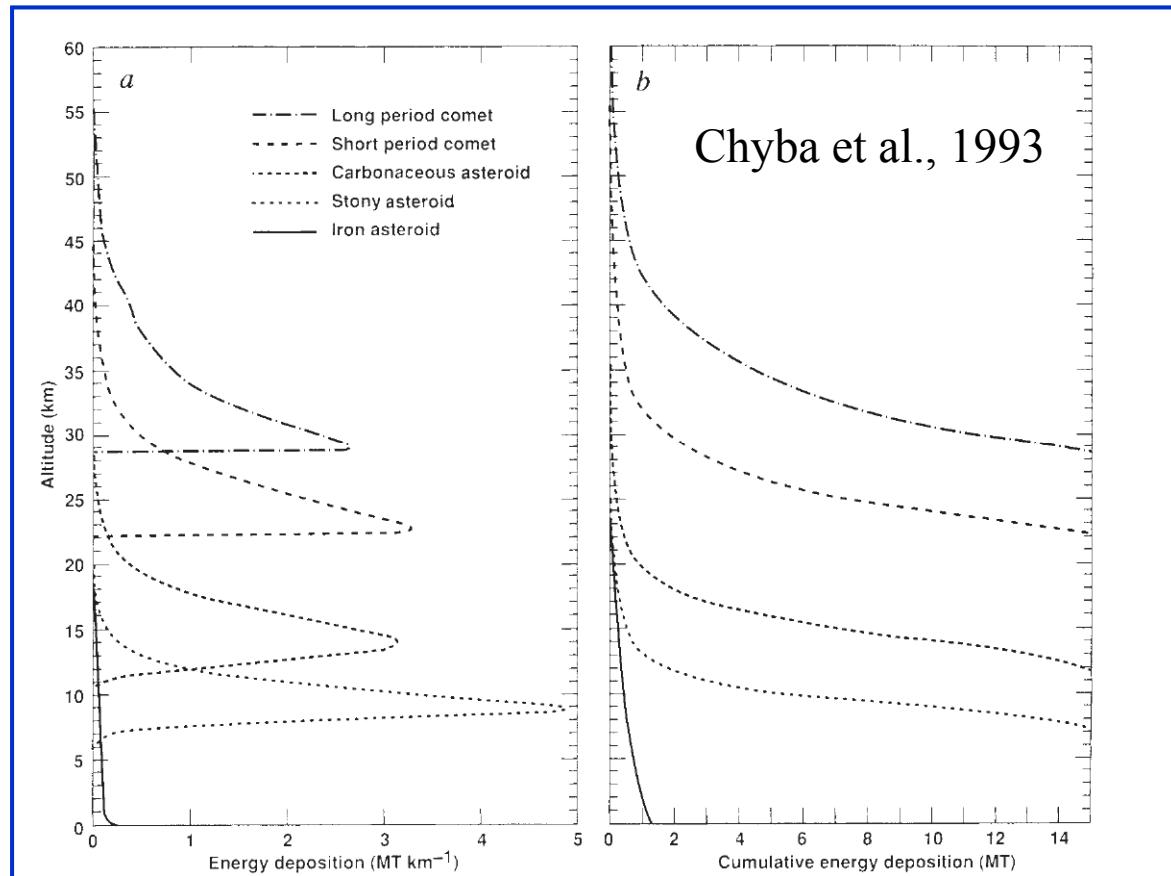
Pancake model: Earth's atmosphere protects us from low-altitude airbursts

Chyba et al. (1993), "The 1908 Tunguska explosion: atmospheric disruption of a stony asteroid" *Science*.

$$m \frac{dv}{dt} = -\frac{1}{2} C_D \rho A v^2$$

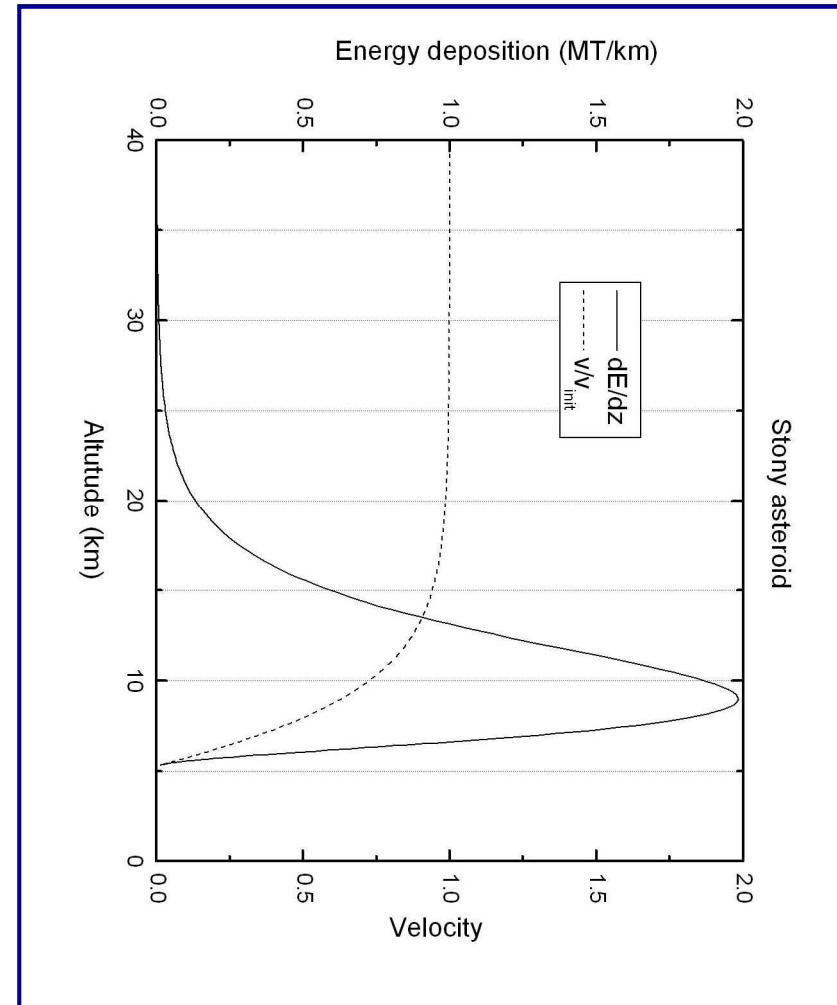
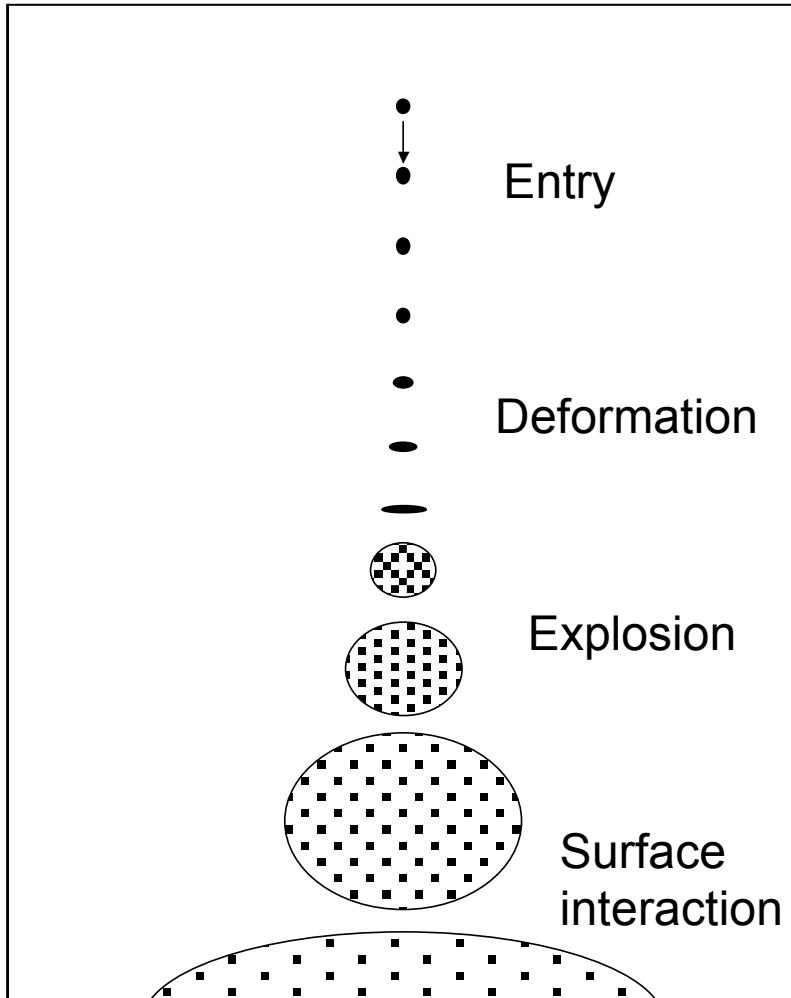
$$q \frac{dm}{dt} = -\frac{1}{2} C_H \rho A v^3$$

$$r \frac{d^2r}{dt^2} = -\frac{1}{2} C_D \frac{\rho}{\rho_m} v^2$$



A stony asteroid deposits essentially all of its kinetic energy above 7 km. In this model the energy deposition curve is sharply peaked because of the mutually-reinforcing effects of atmospheric drag and deformation. Subsequent modeling has been based on point-source explosions and nuclear weapons effects.

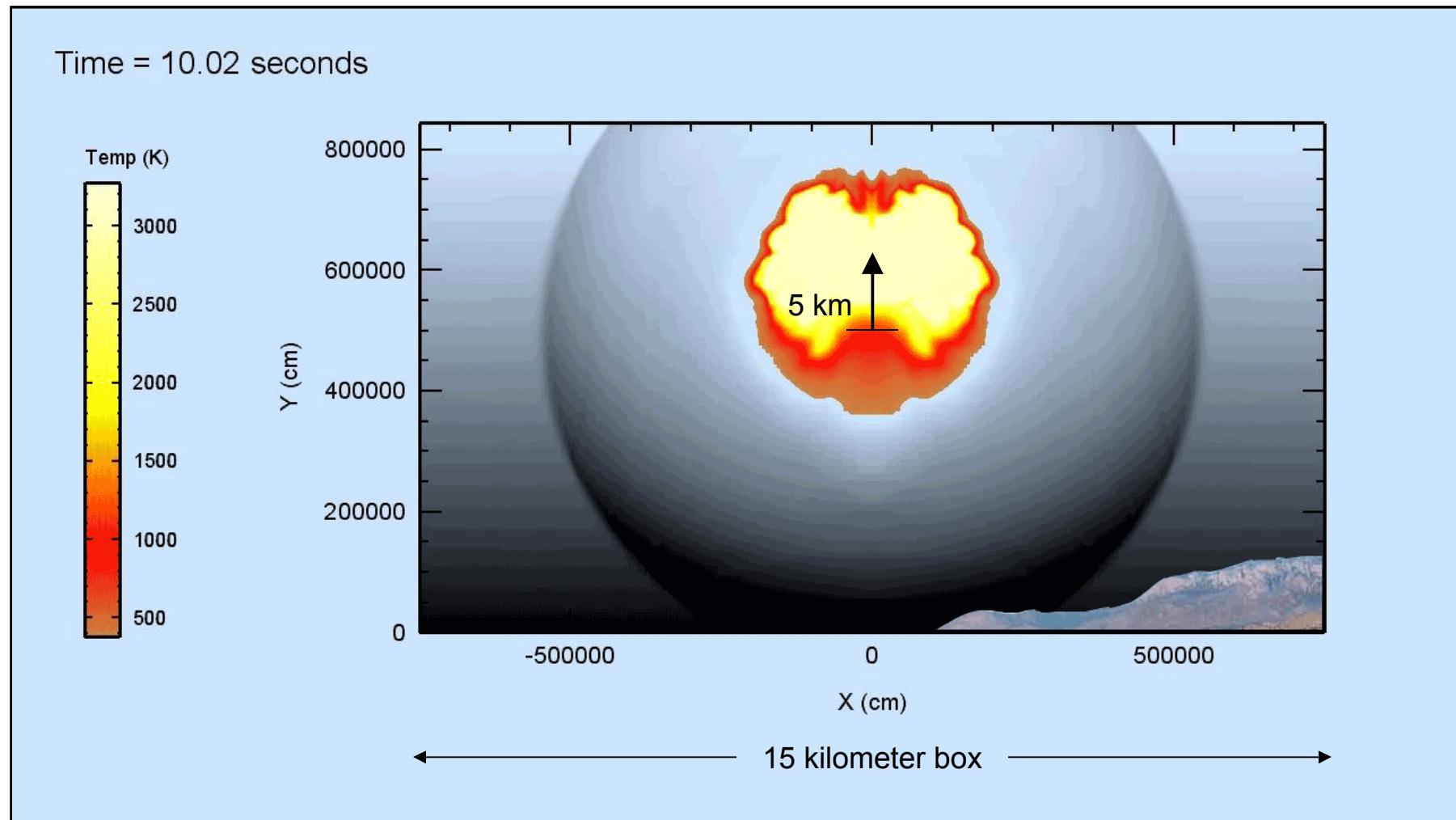
Pancake model revisited: Earth's atmosphere is actually penetrated by hot vapor jet



The “point source explosion” model is a poor approximation.

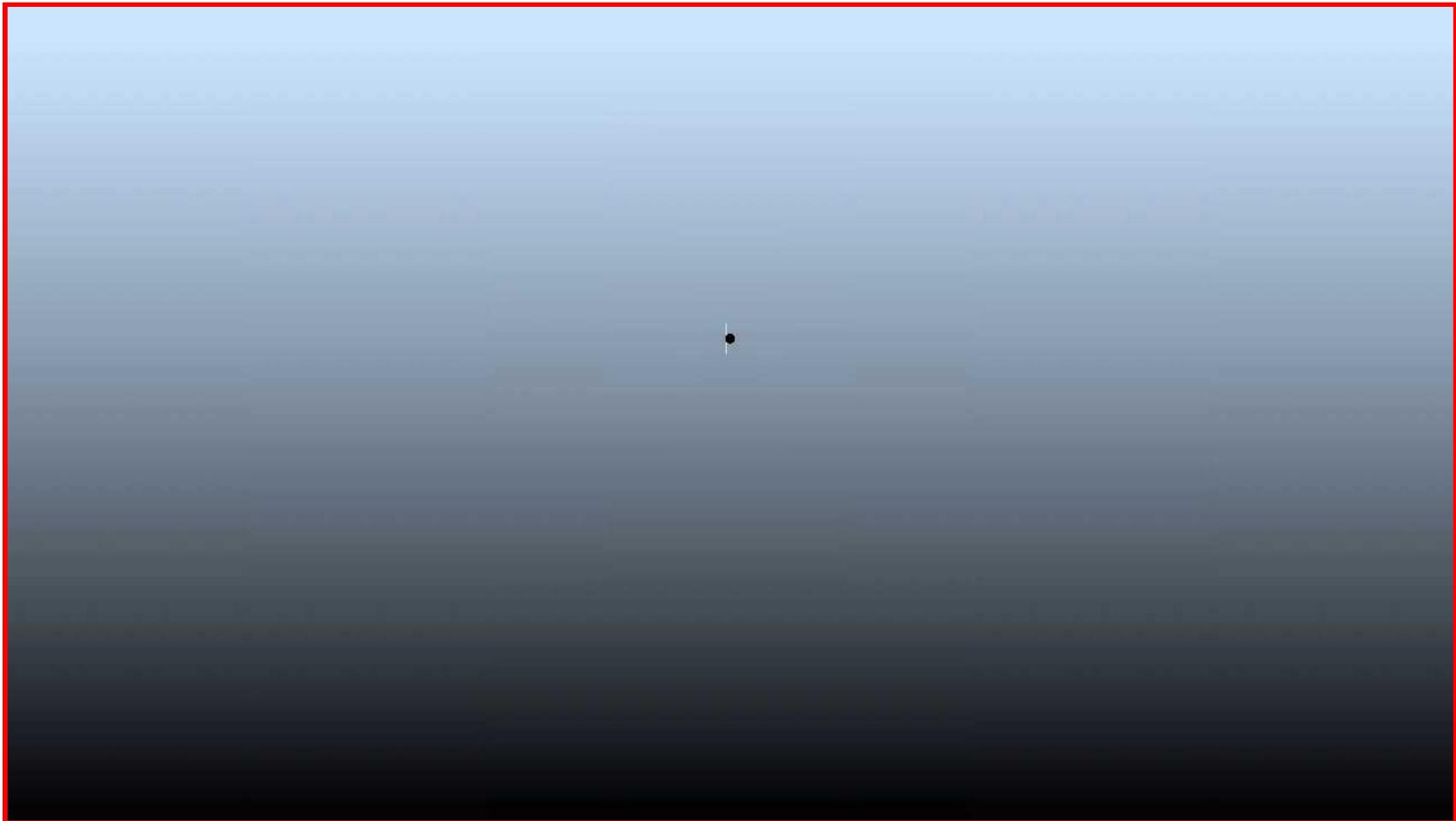
Movies: Difference between explosion and impact

5 megaton point explosion at 5 km altitude: first 20 seconds



Movies: Difference between explosion and impact

5 megaton point explosion at 5 km altitude: first 20 seconds

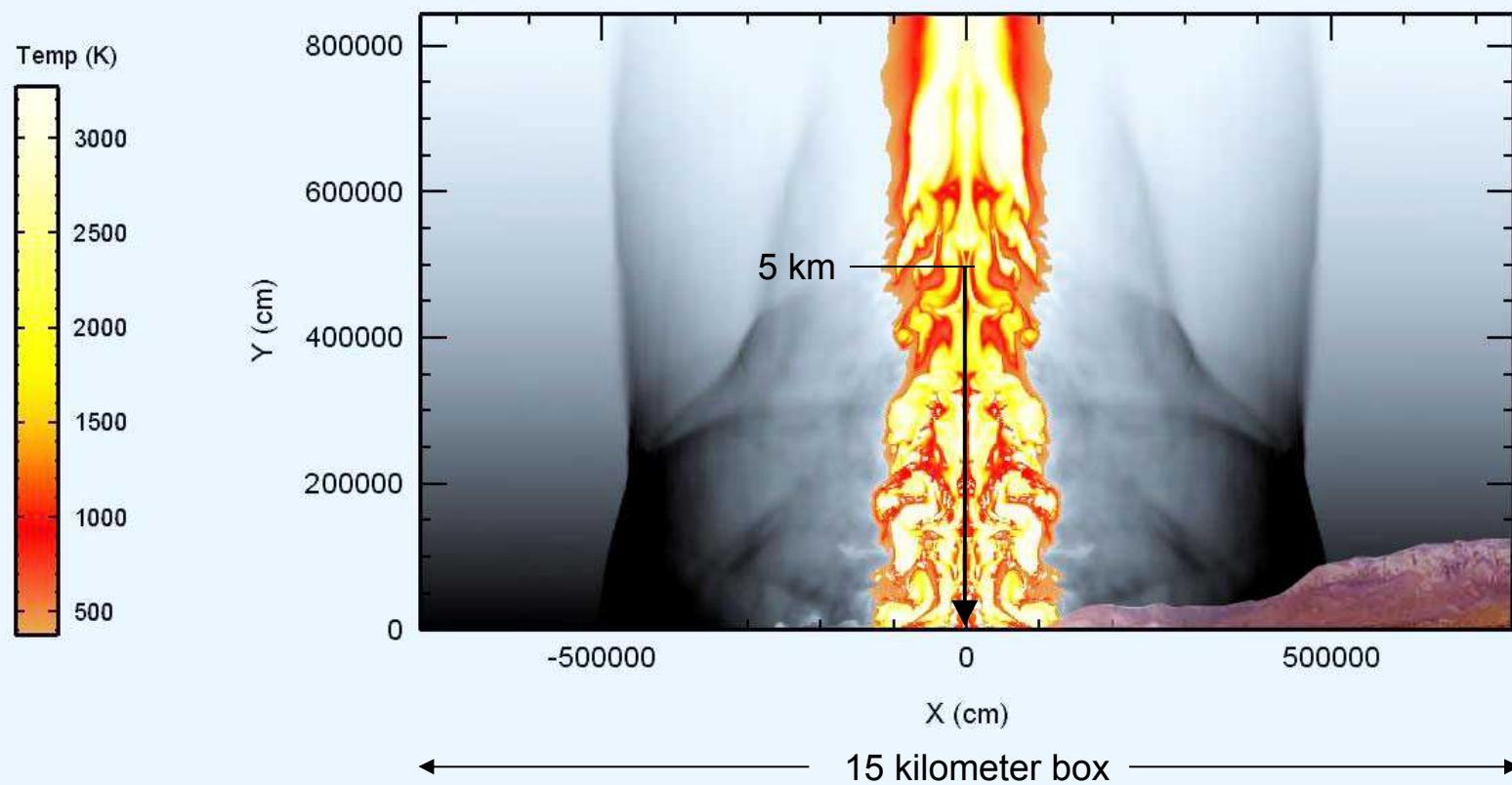


Box dimensions: 8.4 x 15 km

Movies: Difference between explosion and impact

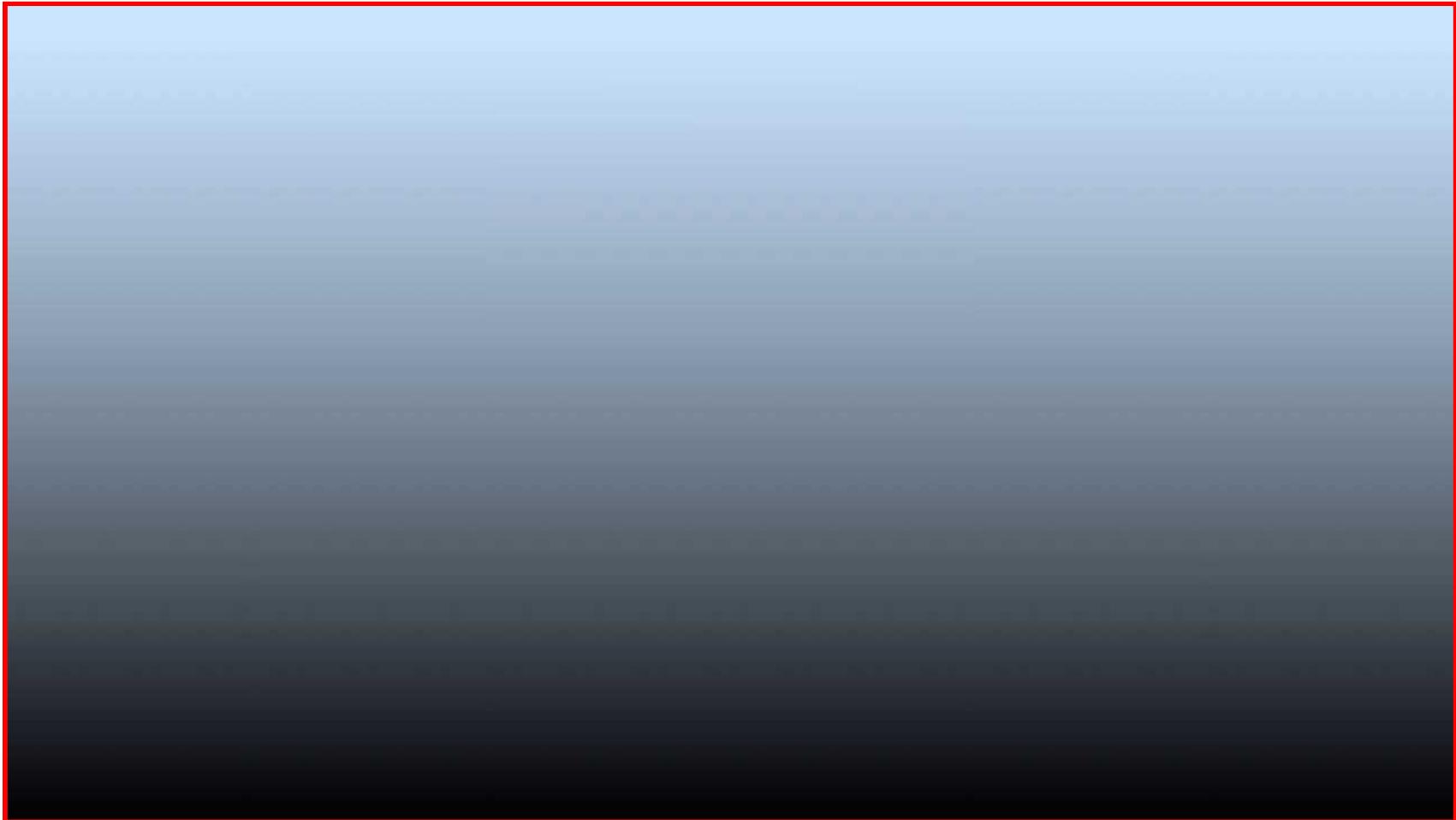
5 megaton impact airburst at 5 km altitude: first 20 seconds

Time = 10.02 seconds



Movies: Difference between explosion and impact

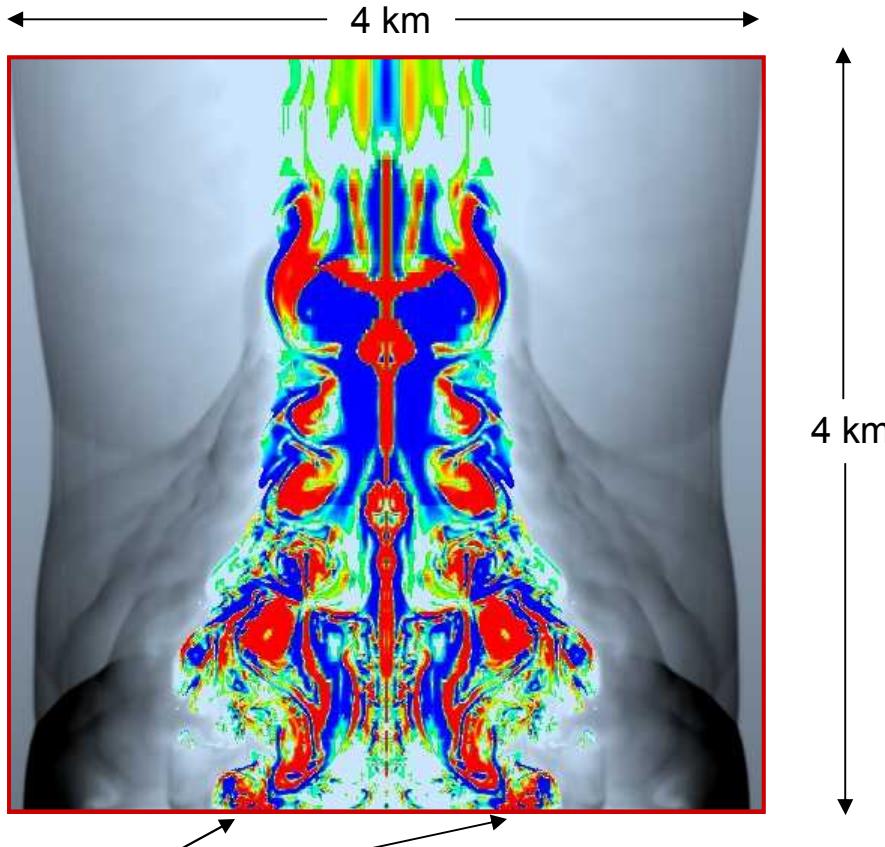
5 megaton impact airburst at 5 km altitude: first 20 seconds



Box dimensions: 8.4 x 15 km

Impact-Induced Vortex Rings

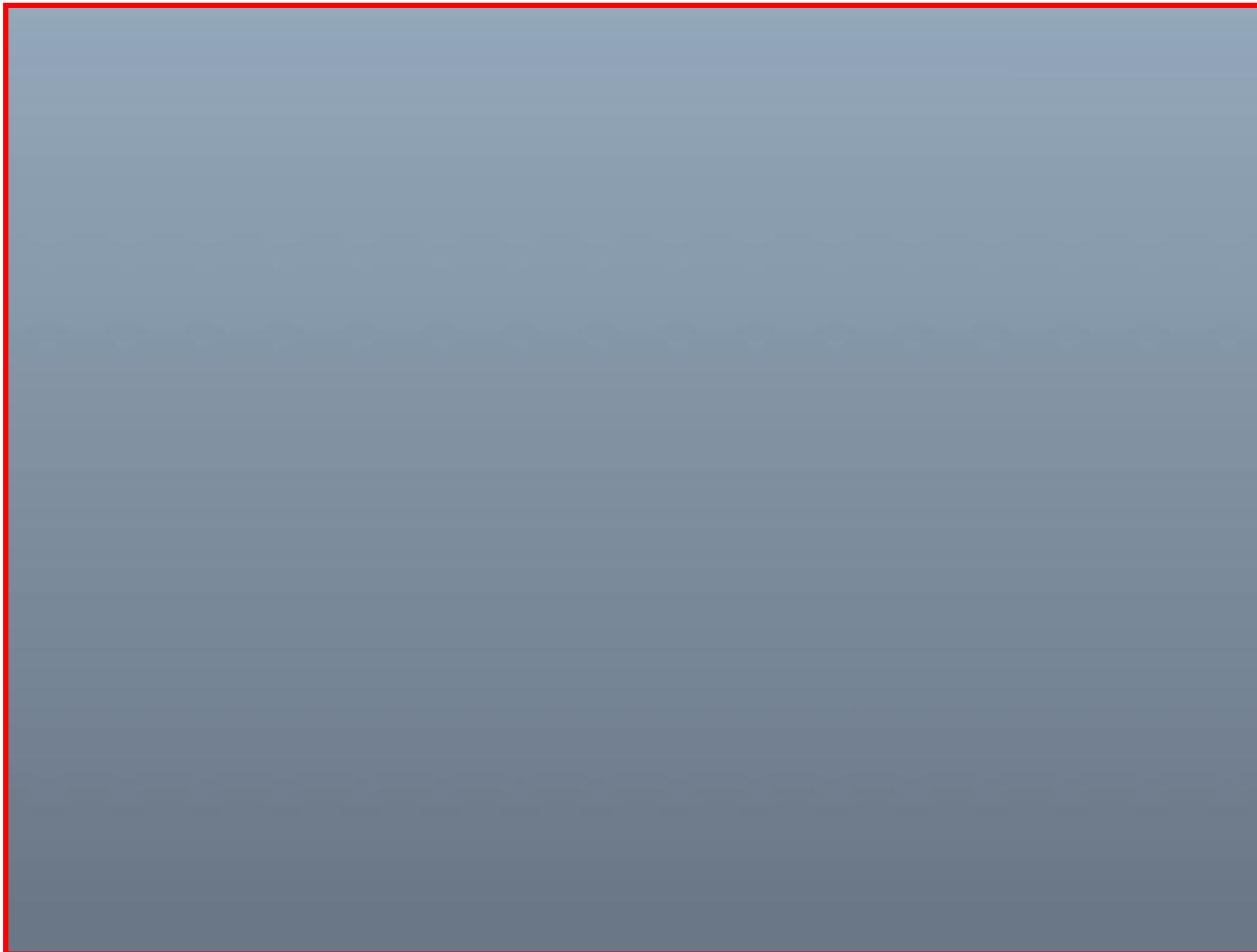
5 megaton impact airburst at 5 km altitude
generates stack of supersonic white-hot “mega-tornado” rings



Red: inward rotation
Blue: outward rotation

Impact-Induced Vortex Rings

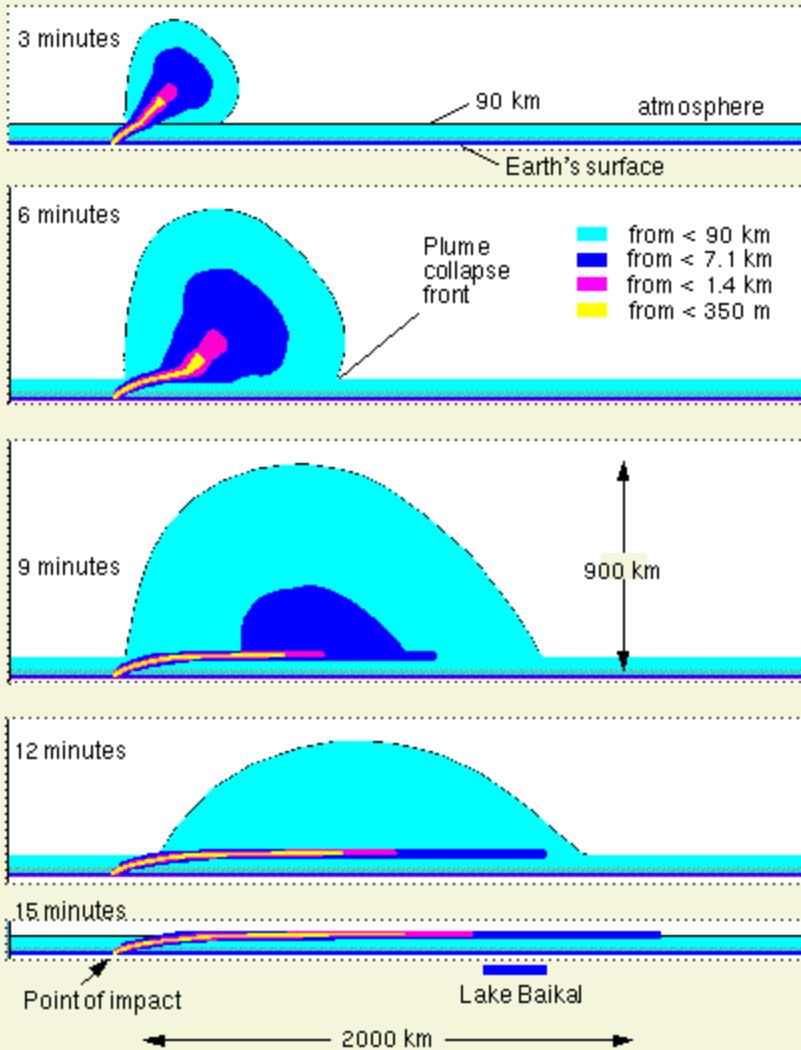
5 megaton impact airburst at 5 km altitude



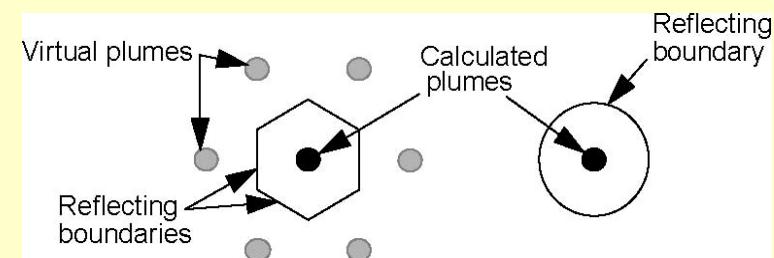
Box dimensions: 3 x 4 km

Bologna, Italy 1996

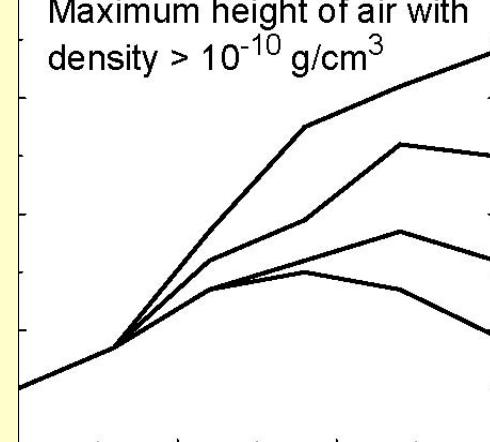
Tunguska96



Silica96



Maximum height of air with
density $> 10^{-10} \text{ g/cm}^3$



Animated version for National Geographic
documentary “Ancient Asteroid”

This video was shown as justification for
Younger-Dryas extinction impact
hypothesis at 2007 Spring AGU, Acapulco.

Tutankhamun's Fireball (Ancient Asteroid)

3. Recent impact controversies

Two new papers

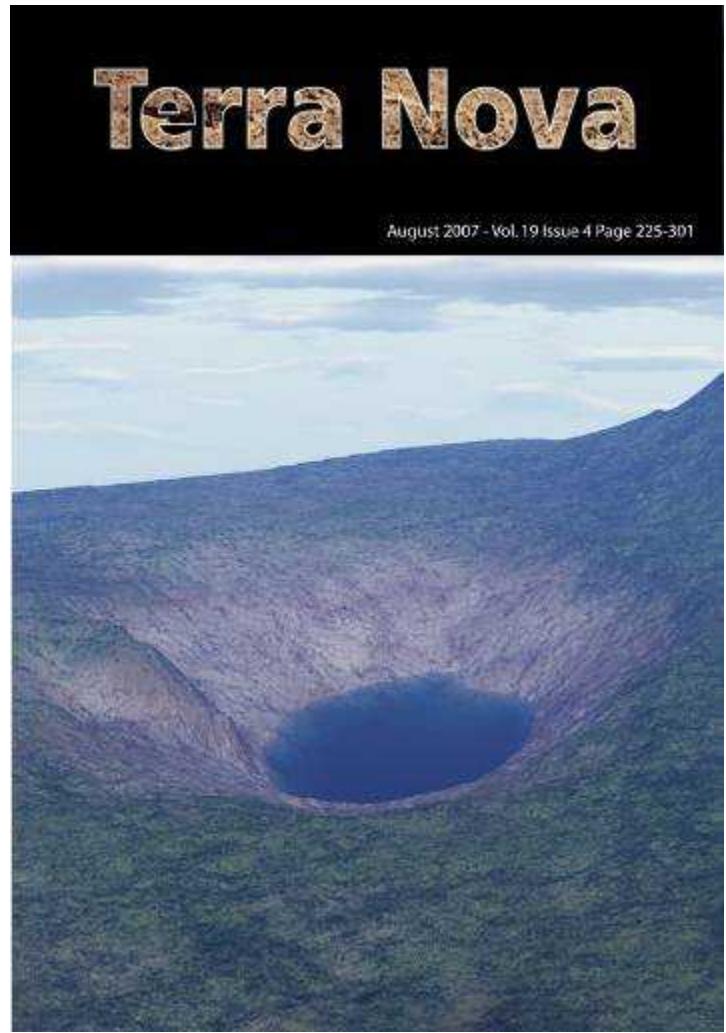
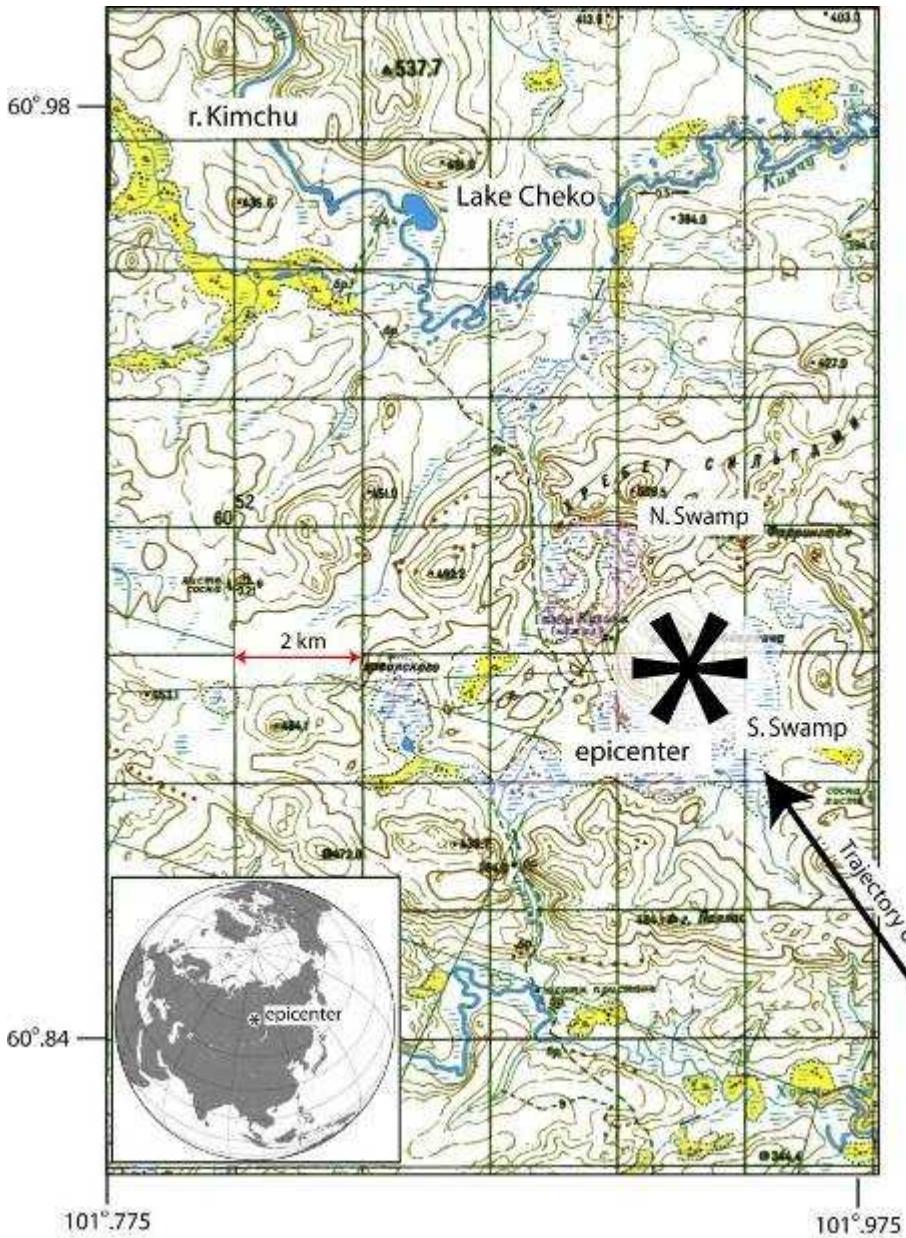
Gasperini et al. (2007) “*A possible impact crater for the 1908 Tunguska Event*”

- Viable hypothesis
- Only evidence is circumstantial
- Have not found any theoretical support

Firestone et al. (2007), “*Evidence for an extraterrestrial impact 12,900 years ago that contributed to the megafaunal extinctions and the Younger Dryas cooling*”

- Not a viable hypothesis
- Evidence is very weak
- Completely inconsistent with physics of impacts and airbursts

The Lake Cheko Controversy



"A possible impact crater for the 1908 Tunguska Event" Gasperini et al., 2007

The Lake Cheko Controversy

Lake Cheko photography:
Discovery Channel

Impact crater or just a lake?

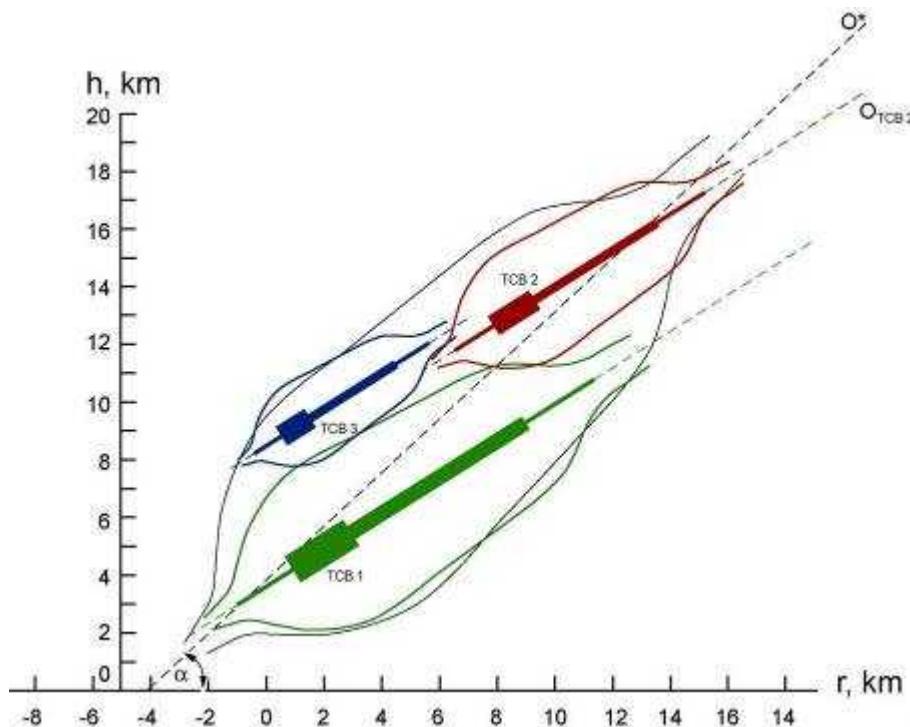


First post-expedition conference: Vanavara

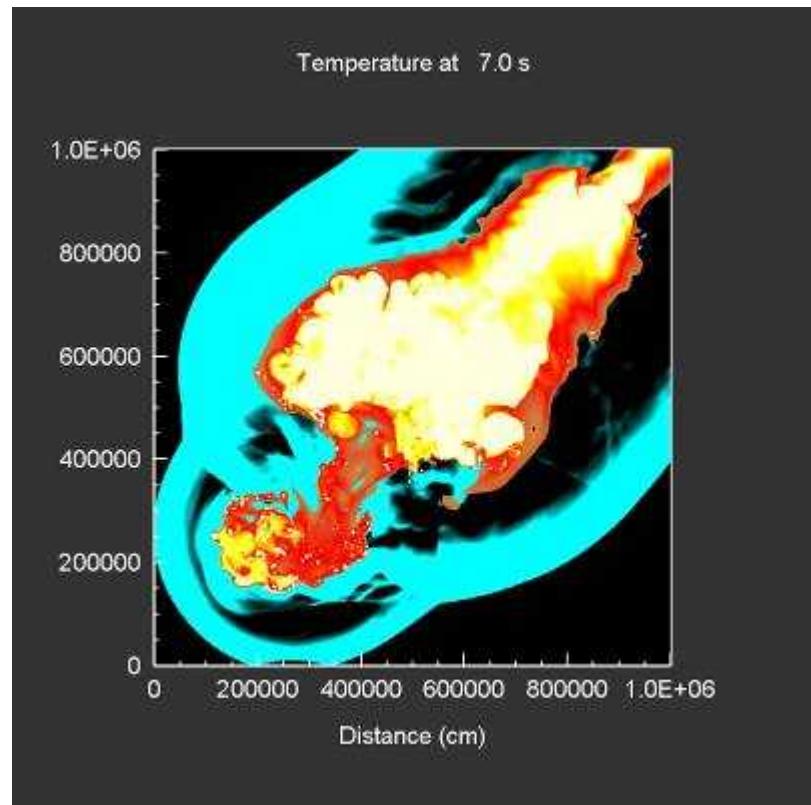


First look: Submersible camera videos!

The Lake Cheko Controversy



“Anfinogenov spindles” (Anfinogenov, 1966)



Multiple explosions (Boslough, 2008)



The Holocene Impact Rate Controversy

June 2008 Atlantic Monthly

“The odds that a potentially devastating space rock will hit Earth this century may be as high as one in 10. So why isn’t NASA trying harder to prevent catastrophe?”

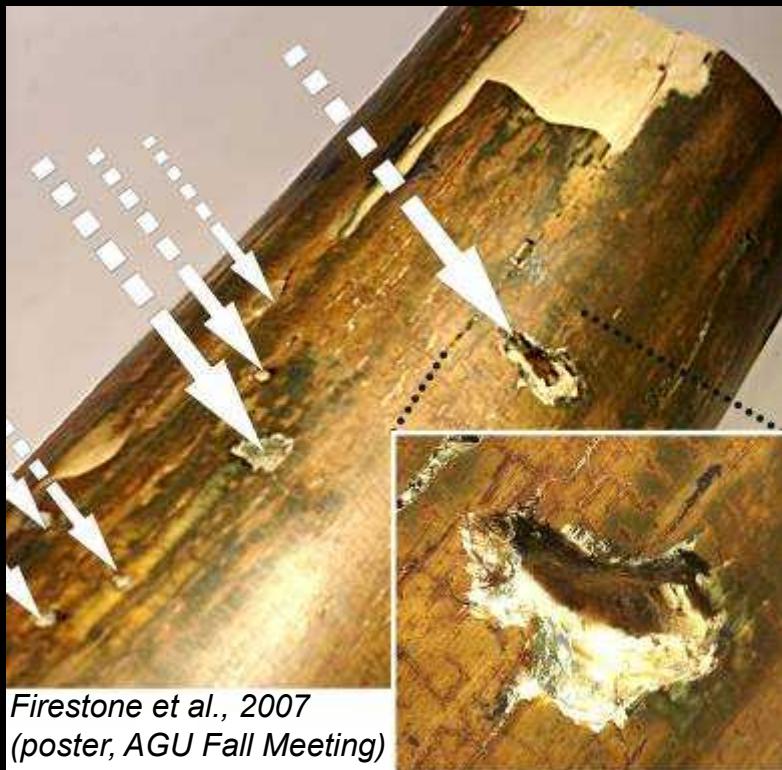
By Gregg Easterbrook

A team of researchers led by [Richard Firestone](#), of the Lawrence Berkeley National Laboratory, in California, recently announced the discovery of evidence that one or two huge space rocks, each perhaps several kilometers across, exploded high above Canada 12,900 years ago.

If, as [Boslough](#) thinks, most asteroids and comets explode before reaching the ground, then this is another reason to fear that the conventional thinking seriously underestimates the frequency of space-rock strikes—the small number of craters may be lulling us into complacency.

Given the scientific findings, shouldn’t space rocks be one of NASA’s priorities? You’d think so, but [Dallas Abbott](#) says NASA has shown no interest in her group’s work: “The NASA people don’t want to believe me. They won’t even listen.”

The Megafauna Extinction Controversy

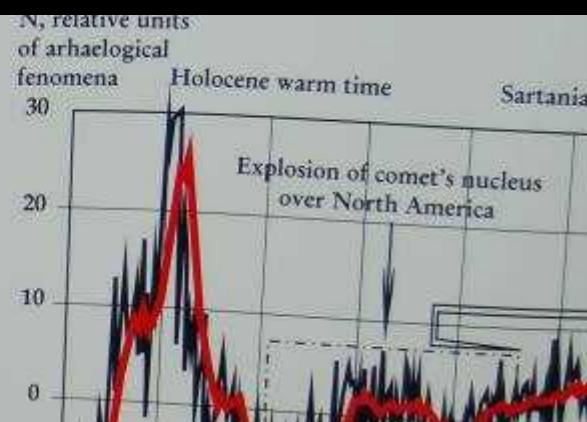


R.B. Firestone et al. (2007),
"Evidence for an extraterrestrial impact 12,900 years ago that contributed to the megafaunal extinctions and the Younger Dryas cooling"



next figure abnormal splash fire wood (Weiss, 1993) 4000-10000 years ago.

Heroes of tragedy



Russian poster at
Tunguska 2008
conference, Moscow

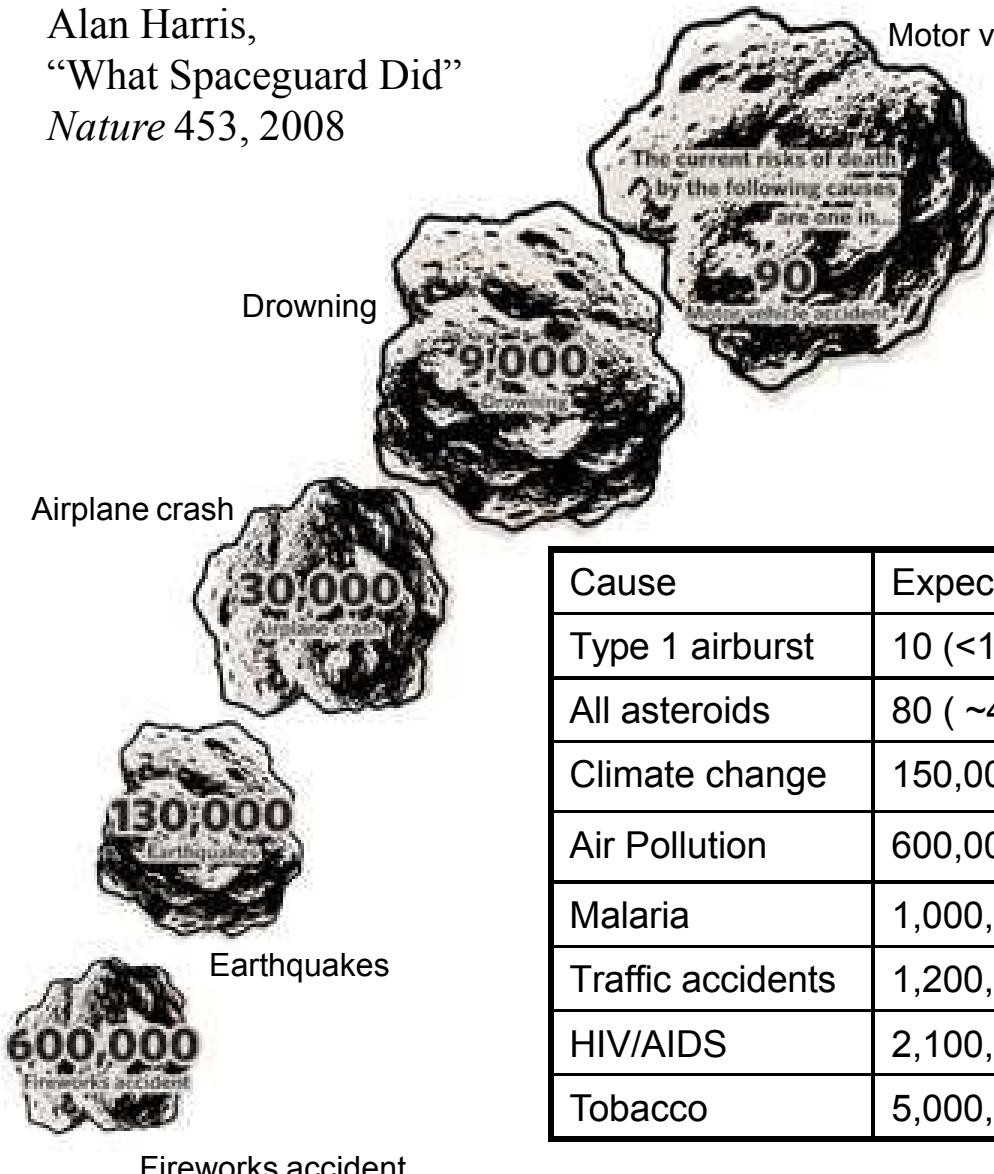
4. The asteroid threat in perspective



© Don Davis

The asteroid threat in perspective

Alan Harris,
“What Spaceguard Did”
Nature 453, 2008



Cause	Expected deaths per year*
Type 1 airburst	10 (<1 per year in US)
All asteroids	80 (~4 per year in US)
Climate change	150,000 – 300,000
Air Pollution	600,000
Malaria	1,000,000
Traffic accidents	1,200,000
HIV/AIDS	2,100,000
Tobacco	5,000,000

*Asteroid estimates from Harris (2008).
All other estimates from the World Health Organization.

Asteroid impact (all sizes)

1,600,000
Asteroid impact (global)

2,800,000
Regional impact (global)

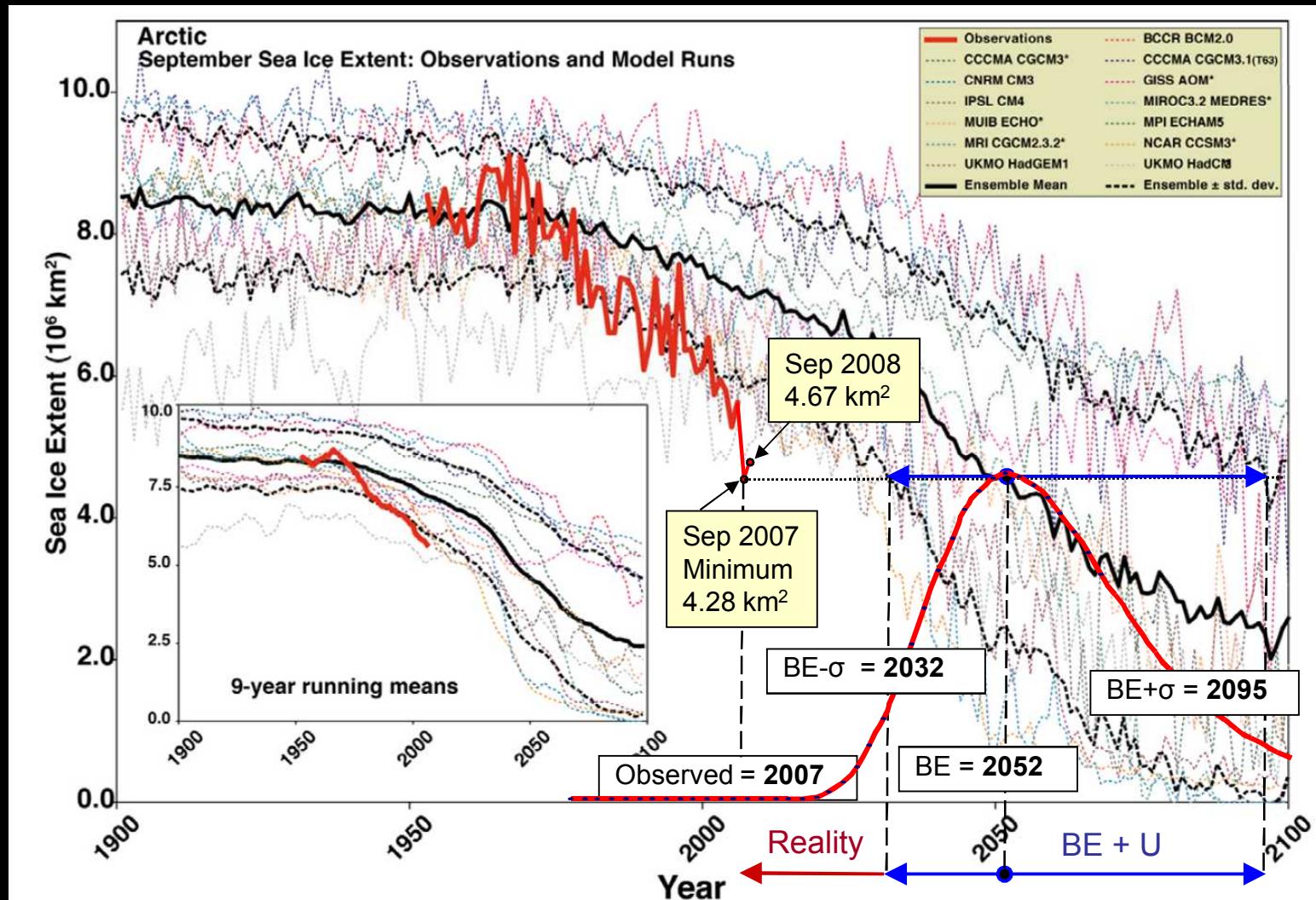
3,000,000
Food poisoning by botulism
Food poisoning by botulism

4,300,000
Impact mass extinction
Impact mass extinction

6,000,000
Tunguska
Tunguska

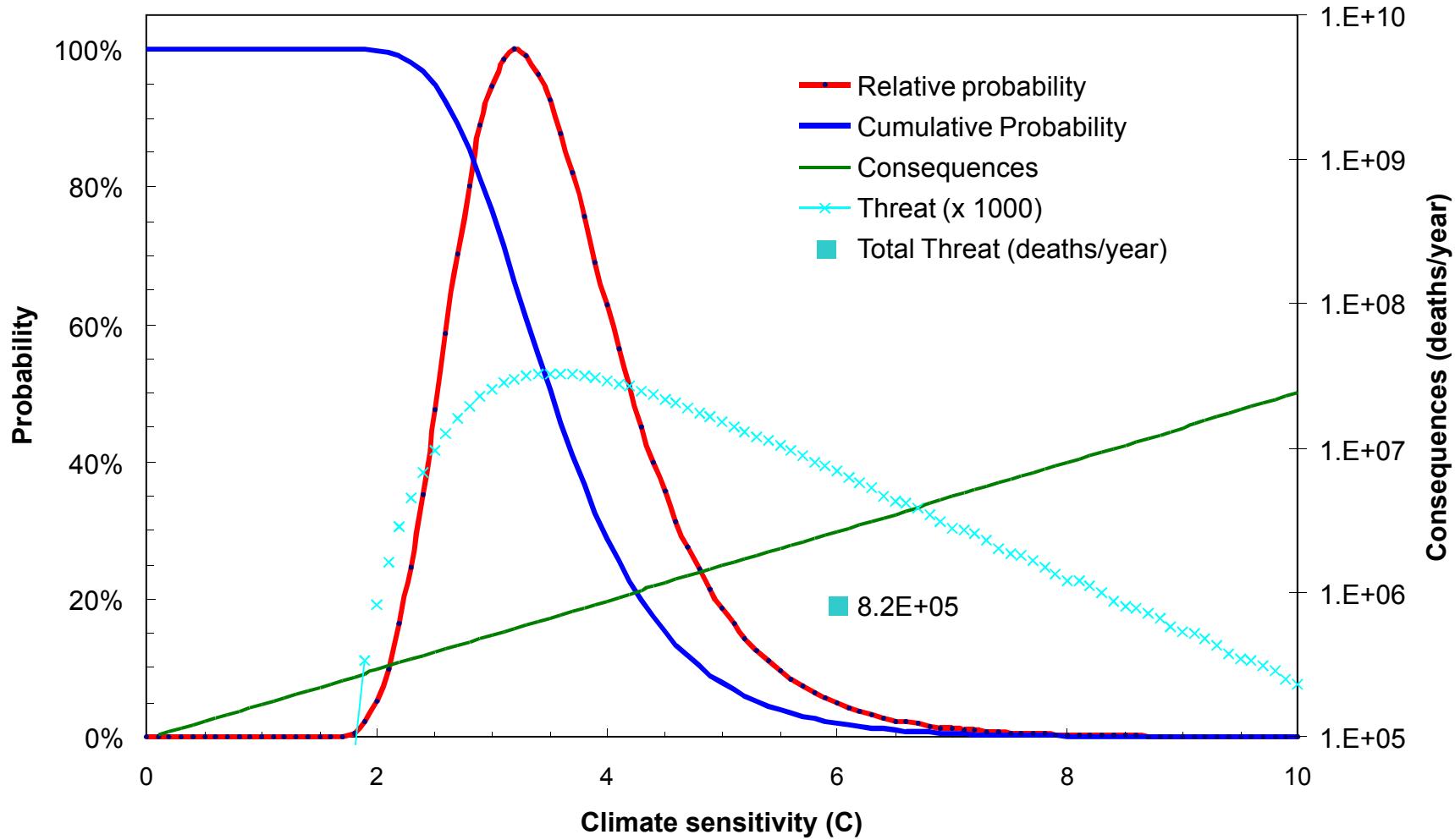
8,000,000
Shark attack
Shark attack

5. Better reasons for NEO deflection



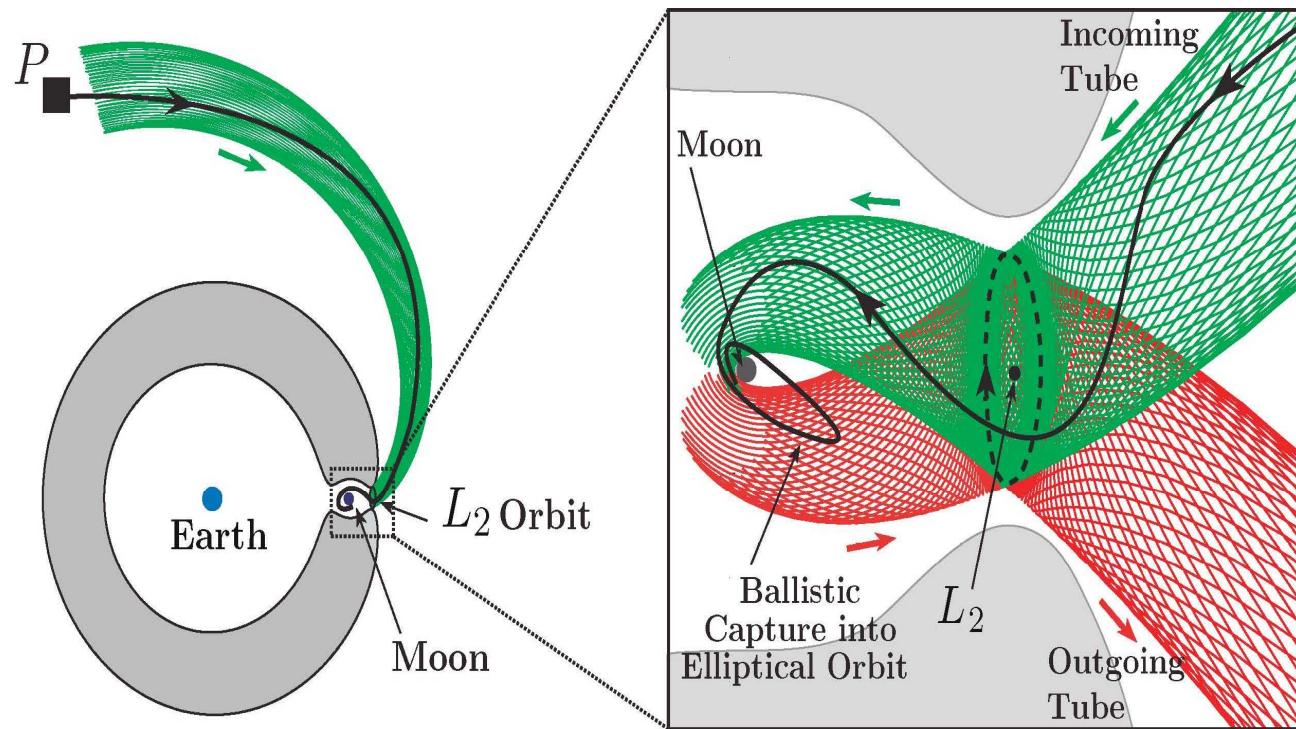
The rate of climate change is exceeding the most extreme expectations

Climate change threat is $\sim 100,000$ times greater



Threat = Probability \cdot Consequences

NEO capture for space resources or geo-engineering



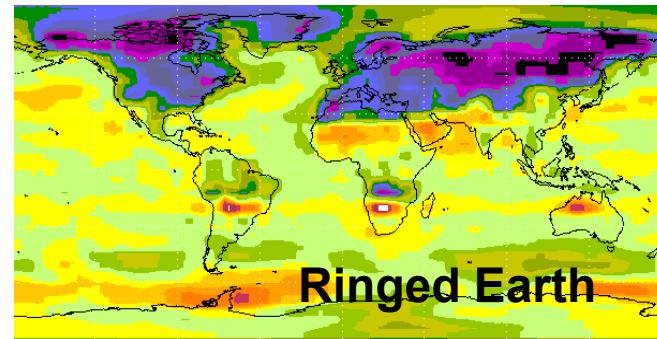
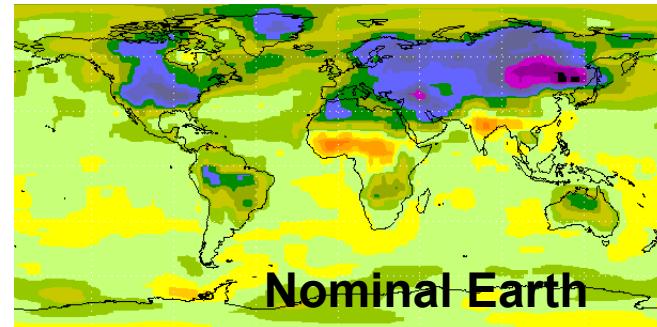
Low-energy pathways allow capture of object into Earth orbit (Marsden and Ross, 2005)

Global climate change mitigation: More urgent than impact mitigation by $\sim 10^5$

- Captured asteroids can be fragmented and used to create engineered orbital debris ring.
- Changes in incoming solar radiation and atmospheric opacity modify Earth's climate.
- Rapid changes in Earth's climate are observable in geologic record.



Climate simulations

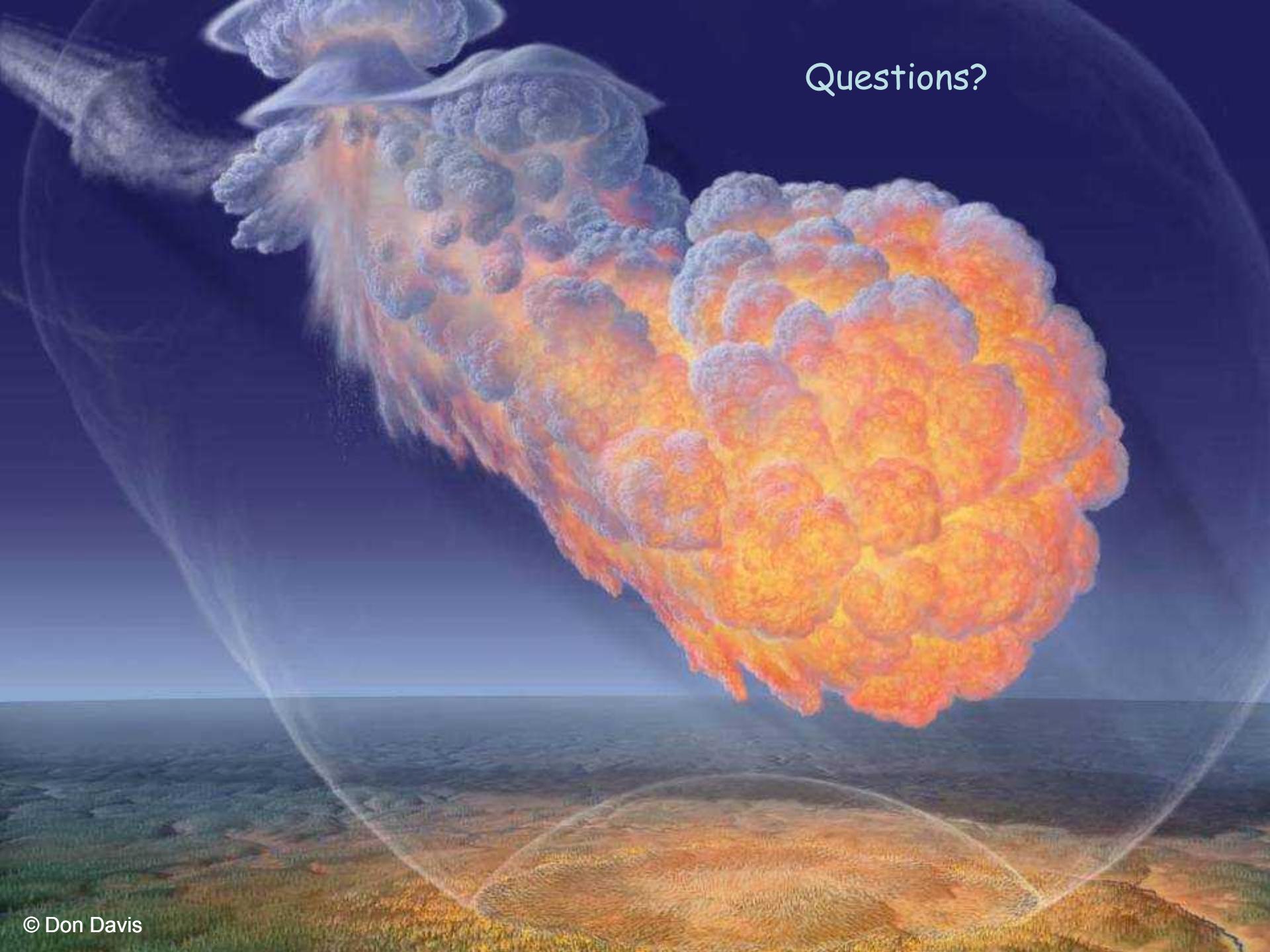


Difference in mean temperature
(low vs. high obliquity)

Fawcett and Boslough (2002), "Climatic effects of an impact-induced equatorial debris ring",
JOURNAL OF GEOPHYSICAL RESEARCH, VOL. 107, NO. D15, 10.1029/2001JD001230, 2002

Conclusions

1. Multiple low-yield shallow bursts are most effective nuclear deflection method.
2. Primary mitigation goal should be for Type 2 airburst prevention.
3. Credibility of community requires skepticism and solid research.
4. Deflection not justified for NEO protection alone.
5. Geo-engineering and space resources are vastly better reasons to develop deflection technology.

A large, multi-colored atomic bomb explosion with a fireball and mushroom cloud over a landscape.

Questions?