

# LA-UR-14-23621

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Title: (U) CIELO: Status of  $^{239}\text{Pu}$  Evaluation

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# CIELO: Status of $^{239}\text{Pu}$ Evaluation

M.B. Chadwick, A.C. Kahler, T. Kawano,  
P. Talou, D. Neudecker  
*LANL*

## $^{239}\text{Pu}$ : Some Particular Challenges

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- ◆ Build on the excellent WPEC subgroup 34 work from CEA & ORNL
- ◆ Capture discrepancies. *We're waiting for new DANCE data*
- ◆ New PFNS results coming (IAEA CRP etc), Chi-nu
- ◆ Inelastic scattering discrepancies between evaluations
- ◆ Use of new IAEA Standards, including fission (TPC)
- ◆ Other new data that will impact the evaluation – new PFGS data from DANCE; New FPY data from TUNL (impact esp at 14 MeV)

# Starter File Using ENDF/B-VII.1+ WPEC Sg34 Work

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Kawano has made a new starter file, for testing the impact of SG34 in the RR region:

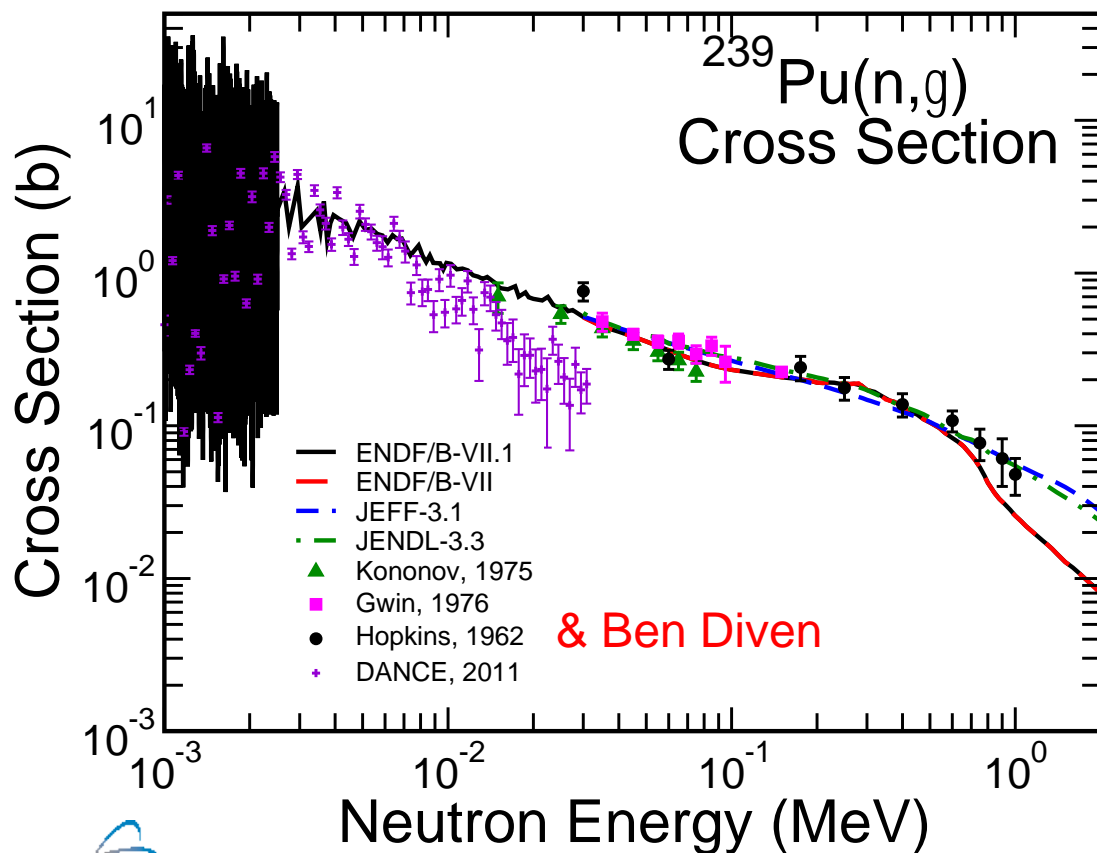
Vii.1 but with:

- sg34 resonances up to 2.5 keV
- file 3 fix for mt18 to zero out cross secs
- jeff3.2 eq jeff3.1.2 (?) Latest nubar up to 650 eV say
- vii.1 chi still

Date Testing Results from Kahler...

# Plutonium Capture: Improvements Are Needed

Existing uncertainties >15%

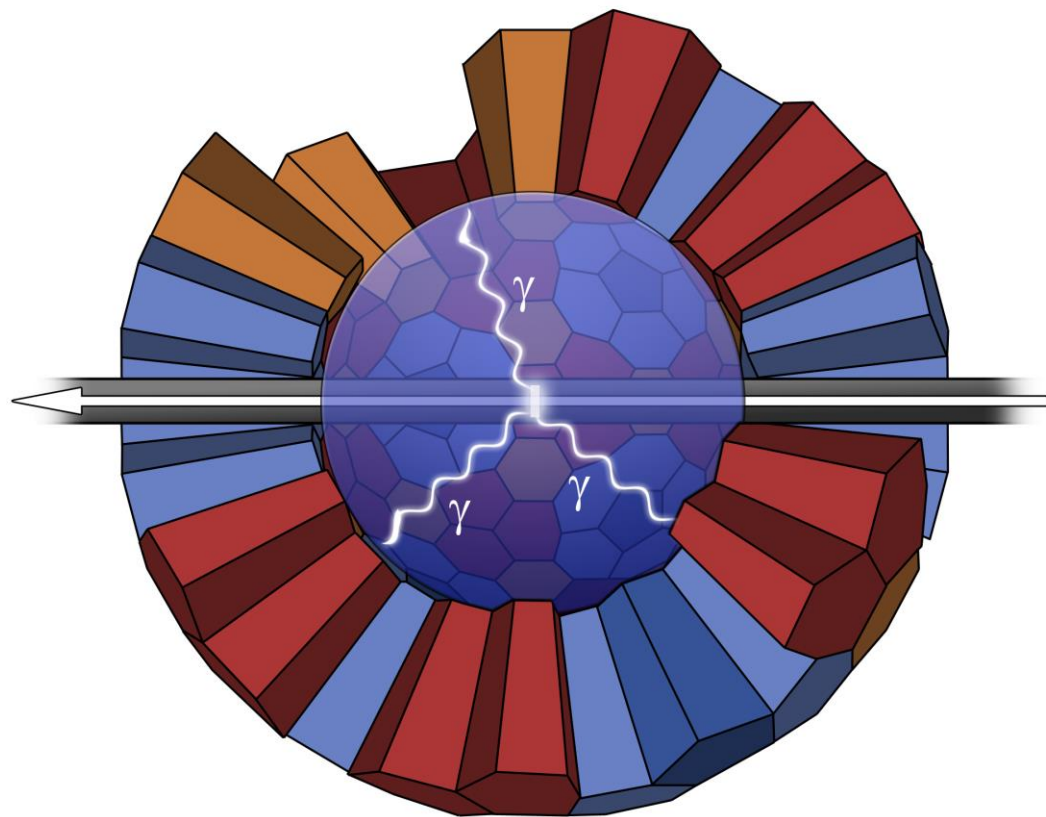


- SG33 & PROFIL (PHENIX)  $^{239}\text{Pu}(n,g)$  integral testing suggests B-VII is ~ 10% low over this fast reactor spectrum. Also, Ishikawa's ADJ work suggests JENDL should be raised 5-10%
- DANCE measurements now being analyzed

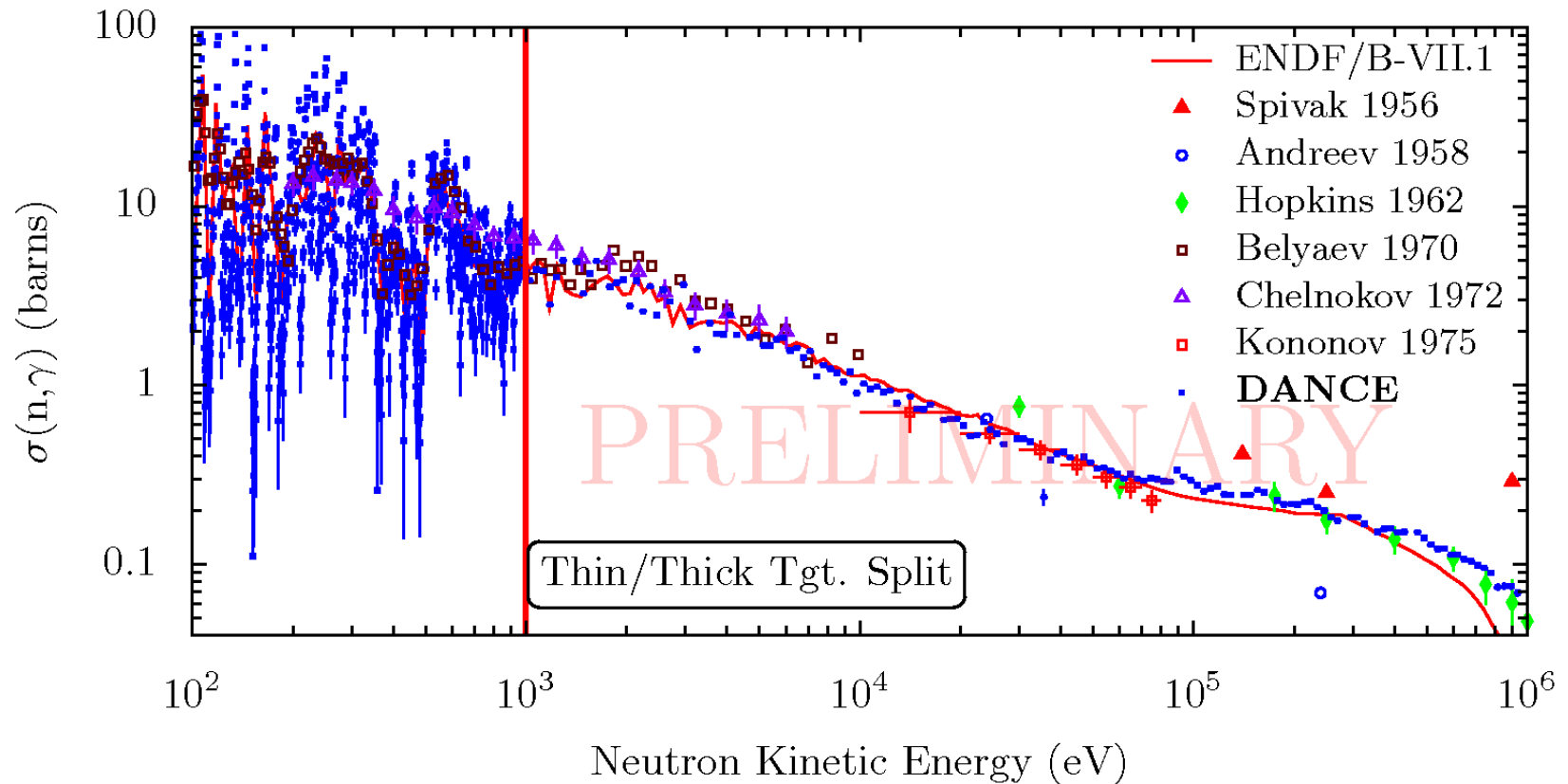
# New Data – Capture Experiments with DANCE (Shea Mosby, Marian Jandel, et al.)

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- 160 detector crystals
- Digital acquisition system
- 85% efficient
- Radioactive/rare targets
  - 5 ug target run this year
- Sophisticated Analysis
  - Energy / multiplicity
  - Unique Q-value for capture
  - Ratio-to-fission method like M. Jandel PRL 2012



## Preliminary Results for $^{239}\text{Pu}$

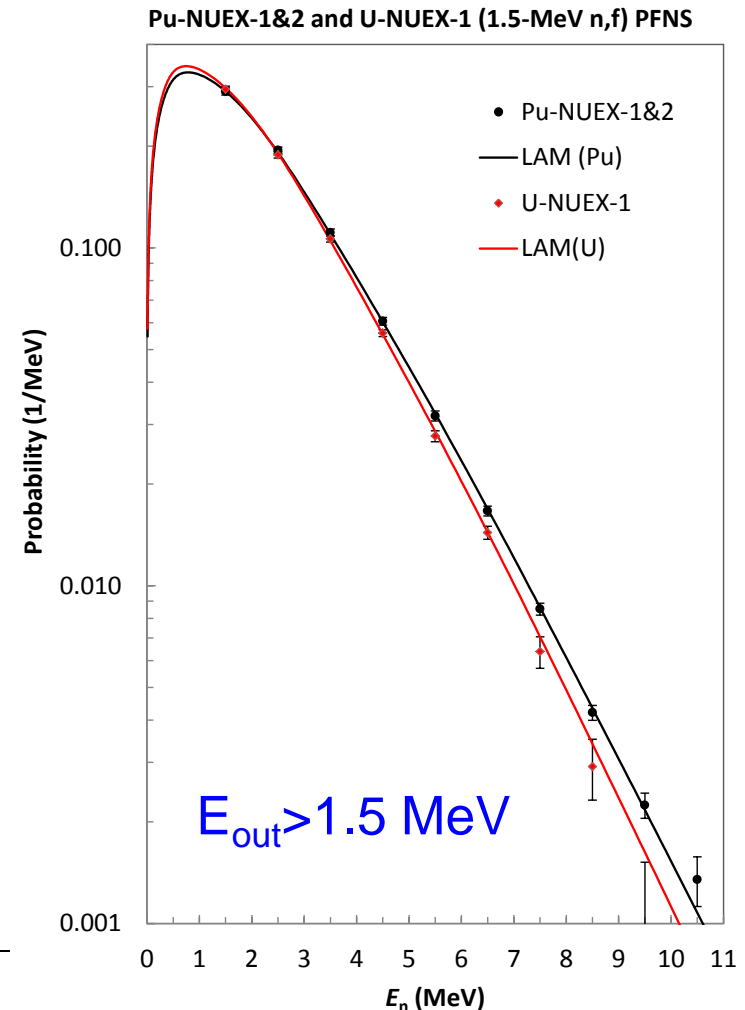
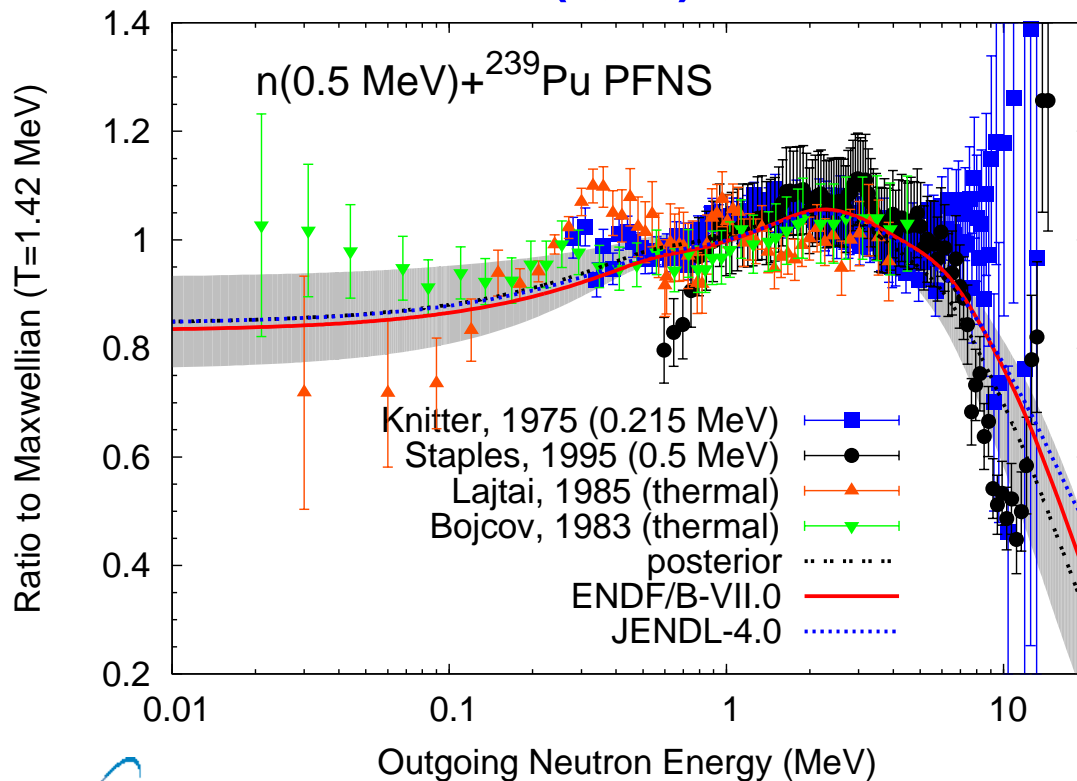


- **Backgrounds still being investigated**
- **Expect 4% uncertainty at 10 keV similar to Jandel work on U5**

# Determining the Prompt Fission Neutron Spectrum (Chi): One of Our Highest Priorities & an IAEA CRP

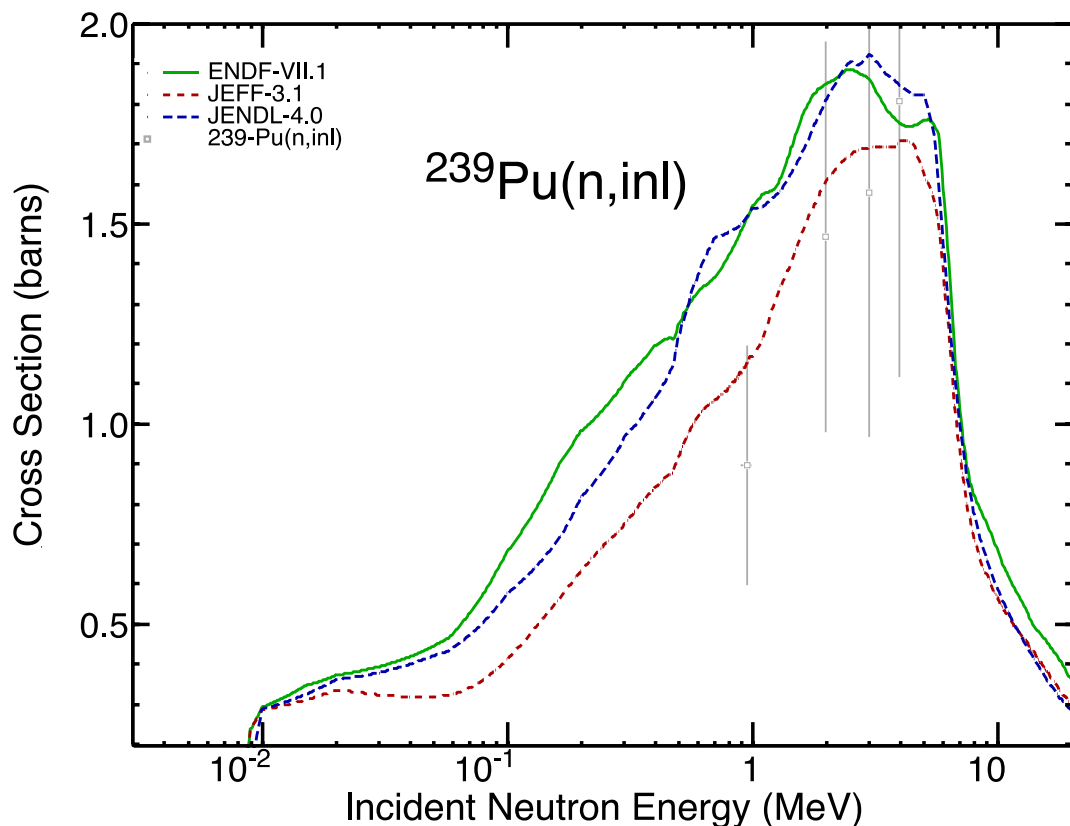
Large uncertainties below 1 MeV and  
above 5 MeV impact criticality  
calculations and (n,2n) transmutations

Lestone's talk: **accurate** underground  
NUEX data released by Los Alamos:





# Major Advance (4): Actinide Elastic & Inelastic Scattering: Large Discrepancies are Starting to be Resolved



Insights from advanced coupled-channel scattering theory needed

**<sup>238</sup>U**: Improvements have been accomplished by Capote, Trkov, using Danon's new RPI data

Dietrich, Thompson & Kawano determined that convergence of CC solutions is slow

Iwamoto, Romain, & Kawano have all made notable advances

# Inelastic Scattering Discrepancy

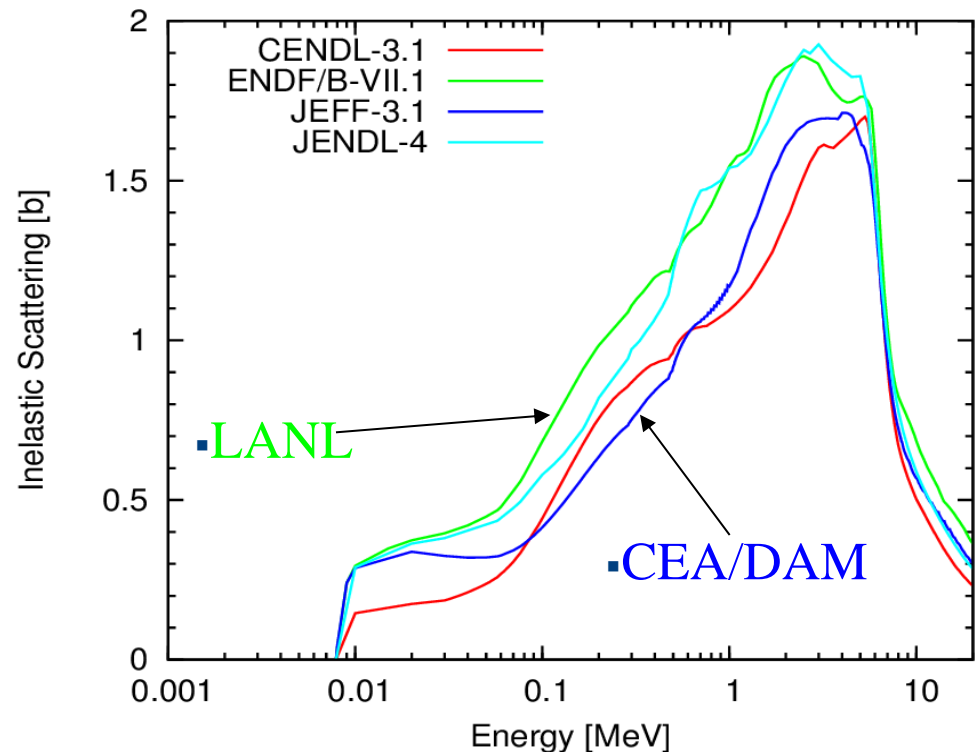
- **IAEA Technical Meeting on Model Calculation for Major Actinides**

- Summary report published: INDC(NDS)-0597, R. Capote, et al.

- These two files equally work for Jezebel keff prediction.

- Probably, the difference in the inelastic scattering comes from the optical potential parameters adopted in each library

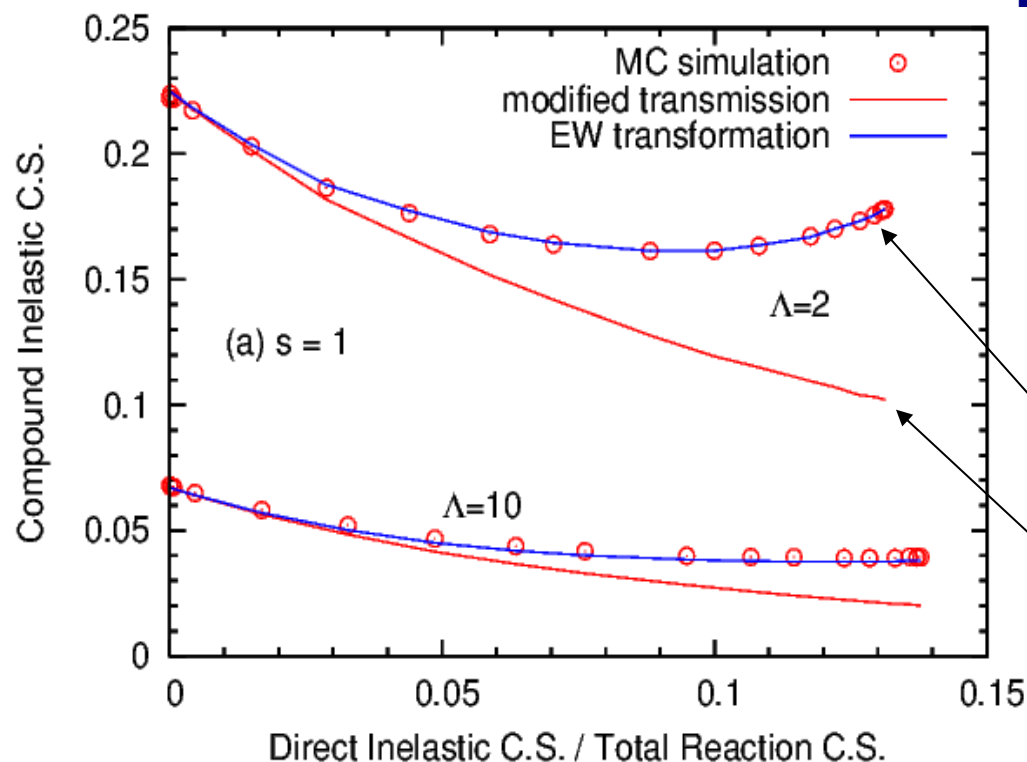
- CEA total cross section is higher than ENDF in the 30keV - 500keV range
    - total and absorption cross sections anti-correlated



## Compound Inelastic Scattering When Direct Reaction Exists (Kawano et al.)

- **Direct reactions in the Hauser-Feshbach codes are treated in rather approximated way**

- Calculated inelastic scattering cross sections by all the Hauser-Feshbach codes, like **Recent development by applying the**



- Recent development by applying the Gaussian-Orthogonal Ensemble (GOE) at LANL demonstrates a **significant under estimation** of inelastic scattering, when strongly coupled channels exist

- A typical case is U-238(n,n') in the fast energy range

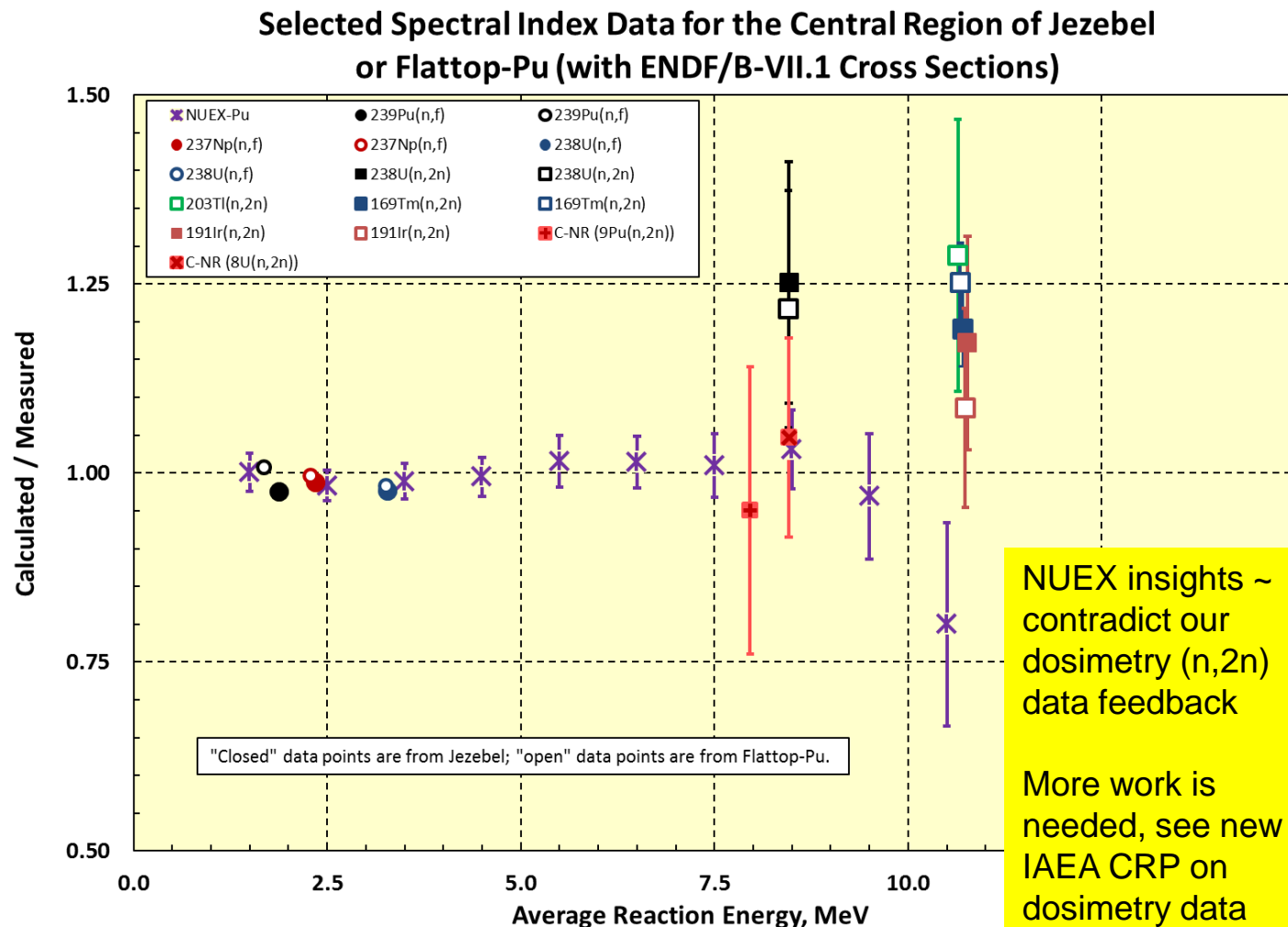
- exact calculation

- approximation in H-F codes

# Reaction Rates in Fast Critical Assemblies Provide Integral Test of Prompt Fission Neutron Spectrum & (n,2n) Cross Sections - *Plutonium-239 PFNS Data*

**$^{239}\text{Pu}$**

**With NUEX  
data added  
(Lestone)**



**Backup: In Case You Didn't Think We Have Lots of Work .....**

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	A	B	C	D
1	CIELO: Summary of tasks to address:			
2				
3	<b>Actinides: 239Pu, 235U, and 238U - specific issues for each nuclide are noted</b>			
4				
5	<b>Fast Region (keVs and above to 20 MeV) - fission listed separately</b>			
6				
7	Review Overall Goals, as embodied in this document and in LAUR CIELO document			
8				
9	Inelastic and elastic scattering - below a few MeV (eg 7)			
10	Review existing discrepancies between evaluations			
11	Collect all available experimental data			
12	Review various theoretical approaches, as embodied in codes (including HF, Coupled Channels, KKM, ....)			
13	Discuss and review optical model options			
14	238U: dispersive coupled-channels OM developed at IAEA			
15	Seek consensus on best evaluated representation of data			
16	238U: 238U Elastic and inelastic scattering data from RPI. Quasi differential available (mainly inelastic) from RPI from from 0.5 MeV up to 20 MeV. - EN			
17	235U: New (n,xng) data to be published in PRC by Kerveno et al. (IPHC, Strasbourg (F)) could be useful to model inelastic scattering on first levels, see			
18	Understand implications from integral data testing on changes in inelastic scattering - especially k-eff and reaction rates (spectral indices for 85/5f etc)			
19	Assess covariances and implement in ENDF format			
20	Create ENDF formatted files			
21	Inelastic and elastic scattering - 7-20 MeV			
22	Review existing discrepancies between evaluations, data, and models (including preequilibrium)			
23	Collect all available experimental data - including Kammerdiener's data and Baba's (U8) data			
24	Review various theoretical approaches, as embodied in codes (including preeq, HF, Coupled Channels, KKM, PFNS background, ...)			
25	Discuss and review optical model options			
26	Seek consensus on best evaluated representation of data - including possible continued use of pseudostates			
27	Understand implications from integral data testing on changes in inelastic scattering -especially 14 MeV pulsed spheres/transmission data			
28	Assess covariances and implement in ENDF format			
29	Create ENDF formatted files			
30	Neutron Capture			
31	239Pu: Review discrepancies between evaluations, which exceed 10% at the higher energies			
32	235U: Review discrepancies between evaluations, which exceed 25% near 1 KeV (Japan's higher result) and 10% at the higher energies			
33	238U: Consider adopting 238U capture from standards - ENDF/B-VII used this, but with some small differences. Study implications from data testing of			
34	238U: Monitor Standards results for any changes, based on new measurements from DANCE, nTOF, Geel			
35	239Pu: Review data (very few measurements, especially above 100 keV there is just the LANL Hopkins data); See if DANCE data is available in time			
36	235U: Review new DANCE data and RPI data, that appear to corroborate JENDL changes near 1 keV, but point to higher energy changes too			
37	Review guidance from integral PROFIL data (suggests PU9 and (maybe) U5 from ENDF should be higher), and Wallner AMS data at 25 keV and 420 keV			
38	Assess model calculations predictions (consistent with above inelastic scattering HF/CC/OM calculations)			
39	Seek consensus on best evaluated representation of data			
40	Understand implications from integral data testing on changes in capture - especially k-eff and reaction rates (spectral indices for 85/5f etc)			
41	Assess covariances and implement in ENDF format			
42	Create ENDF formatted files			
43				
44	n2n			
45	Discuss data, including discrepancies in rise from threshold, and differences near 14 MeV			
46	Review existing evaluations (including "GEANIE evaluation" for 239Pu), data, and calculation predictions			
47	235U: New (n,xng) data to be published in PRC by Kerveno et al. (IPHC, Strasbourg (F)) could be useful to model n2n scattering, see prelim results in f			
48	239Pu: Carefully note insights on n2n making 238Pu from LANL, and discuss contradictory feedback from PROFIL			
49	Validate any changes against n,2n reaction rates in critical assemblies, eg Fig 57 in NDS112,(2012) ENDF			
50	Create ENDF file and covariances			



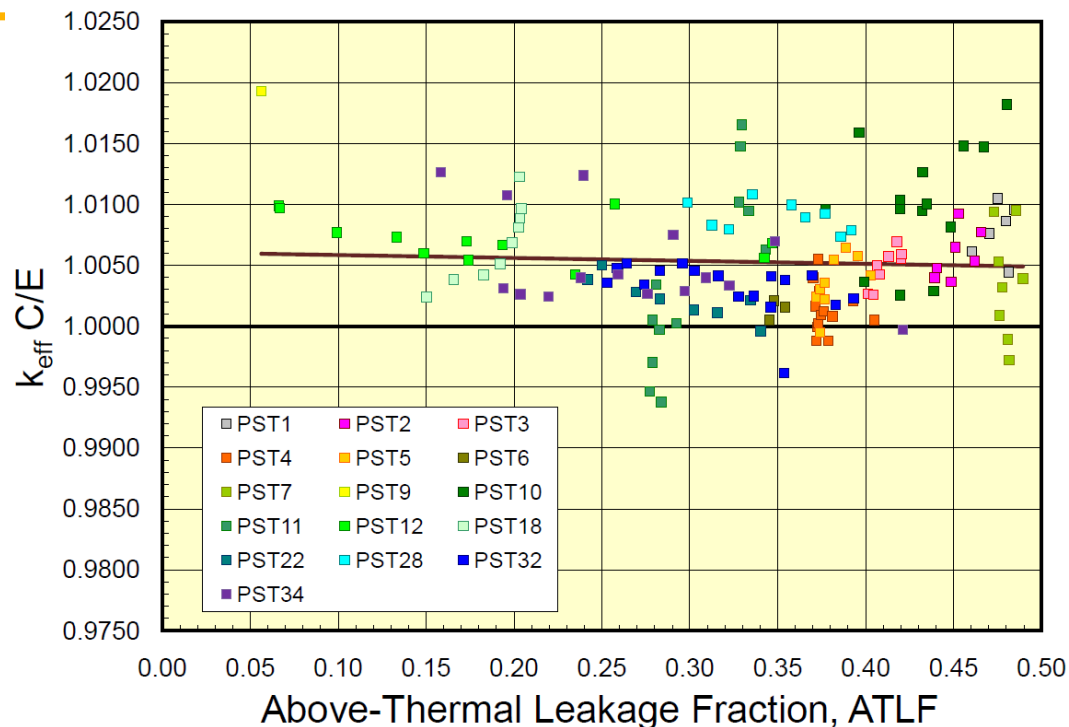
71	Fission (all energies), cross sections, nubar and spectra for n,g		
72			
73	Review Overall Goals, as embodied in this document and in LAUR CIELO document		
74	Fission Cross Section		
75			
76	Seek consensus that we adopt the fission cross section standard from the IAEA group		
77	Assess implications of adopting standard fission cross section on integral testing		
78	If IAEA standards team updates their value, use it; this would include any recent/forthcoming fission measurements, eg nTOF, RPI, TPC		
79	Modeling of fission would occur as part of the above inelastic/capture/n2n activities, but seek consensus that we do not use calculations in the		
80			
81	238U:Subthreshold fission for 238U – discrepancies between different evaluations. Lead spectrometer measurements near 70 keV suggest a p		
82	prompt nubar		
83	Review existing evaluations and experimental data, & review various theoretical approaches; 238U low energy interp fix needed in ENDF		
84	Seek to use an "unadjusted" nbar in a final evaluation, avoiding the ENDF "tweak" near an MeV that was adopted to better match Jezebel, Go		
85	Study Koning-Rochman nubar near thermal, from their optimization search (but it's 3 SD below the standards constants value)		
86	Develop a new evaluation based on a covariance analysis of the data		
87	Understand implications from integral data testing on changes in nubar - especially k-eff		
88	Create ENDF formatted files, including covariances		
89			
90	PFNS		
91	Review work of IAEA CRP on PFNS		
92	Aim to adopt the CRP's recommendation		
93	Seek consensus on using LANL high-accuracy NUEX Pu9 and U5 data, as published in Dec NDS2011 to help define high-energy spectrum		
94	Use new PFNS measurements, especially below MeV, coming from LANSCE/Chi-nu in the coming years		
95	Use guidance on high energy tail of spectrum from dosimetry reactions (new IAEA IRDFF CRP), eg from LANL crits, Russian fast reactor, & CE		
96	As part of IAEA CRP, advance our theoretical models, and use incorporate other data (new and existing)		
97	Understand implications from integral data testing on changes in inelastic scattering - especially k-eff and reaction rates in assemblies		
98	Create ENDF formatted files, including covariances		
99			
100	PFGS		
101	Review existing evaluations and experimental data, and various theoretical approaches		
102	Represent fission gammas separately at all energies, including above 1.09 MeV for U5 and Pu9 (an ENDF drawback), & use new data available		
103	Update PFGS spectra to use modern measurements from DANCE, as well as multiplicity distribution if possible		
104	Create ENDF formatted files, including covariances		
105			
106	Delayed data		
107	Review differences in present evaluation		
108	Develop plan for work needed		
109			
110	Energy Release		
111	Compare energy release data in evaluations, for prompt n, g, fission fragments; and delayed energy release		
112	Update as necessary - eg ~ MeV level changes are implied for 239Pu from Jandel's DANCE data for 239Pu (but 235U looks good)		
113	Consider updating energy release incident-energy-dependence based on Lestone's work		
114			

116	Integral Data Testing and Validation			
117				
118	Review Overall Goals, as embodied in this document and in LAUR CIELO document			
119	Define suite of critical assembly, reactor, transmission, etc experiments to use in validation assessments, and observables (k-eff, rates, spectral indices)			
120	238U: selection of 12 ICSBEP criticality benchmarks sensitive to elastic scattering is available from JSI/IAEA (Trkov, Capote)			
121	Seek to ensure good performance in data testing, which includes:			
122	Fast, intermediate, and thermal assemblies, k-eff			
123	239Pu: Aim for (Partial?) improvement of longstanding overprediction of thermal Pu solutions			
124	Modeling spectral indices well in various systems (Incl fast), 8f/5f, 9f/5f, 237np-f/5f, 233u-f/5f etc, see Table XXXVIII in VII.1 NDS 2011 paper			
125	Modeling of post irradiation experiments (PIE) such as PROFIL (CEA) and MANTRA (INL)			
126	Modeling MOX experiments for mock up of LWR, eg in EOLE, Cadarache			
127	See if PFNS improvements give improved n2n detector responses in fast crits, eg through a softer PFNS spec above 10 MeV			
128	nubar validation using multiplication subcritical measurements			
129	LLNL pulsed spheres			
130	Can we obtain improved predictions of intermediate assemblies, eg ZPR at Argonne			
131	Aim to maintain good prediction of crits, including new as-built high-resolution 3D MCNP Jezebel model?			
132	Use sensitivity methodologies for assessing changes/improvements by reaction and energy range			
133				
134				
135				
136				
137				



# Pu-SOL-THERM Benchmarks – I. Prelim LANL testing of new Subgroup 34 resonance results

- A ~500 pcm bias in calculated PST reactivity is a long-standing issue.
- WPEC Sub-Group 34 was tasked with defining a new (better?) set of resolved resonance parameters for  $^{239}\text{Pu}$  in an attempt to resolve this issue.
- Can define a sub-set of these 150 benchmarks to test revised data files.



- Consider benchmark attributes such as (i) ATLF; (ii)  $^{239}\text{Pu}$  atom-% in Pu; (iii) Above-Thermal Fission Fraction (ATFF); (iv) H/Pu number density (or gPu per liter) to define this sub-set.

## Pu-SOL-THERM Benchmarks – II. Prelim LANL testing of new Subgroup 34 resonance results

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- **A set of seven Pu-SOL-THERM benchmarks have been extracted from the larger set.**
  - PST1.4 & PST12.13 span the ATLF space;
  - PST12.10 & PST34.15 span the ATFF space;
  - PST4.1 & PST18.6 span the  $^{239}\text{Pu}$  atom percent space;
  - PST12.10 & PST34.4 span the g Pu per liter space.
  
- **All benchmark experiments are performed in simple geometry**
  - PST1.4 & PST4.1 are a water-reflected spheres;
  - PST18.6, PST34.4 & PST34.15 are water-reflected cylinders;
  - PST12.10 & PST12.13 are a water-reflected slabs;

# Pu-SOL-THERM Benchmarks – III. Prelim LANL testing of new Subgroup 34 resonance results

• The E71 1.00576  $k_{\text{calc}}$  average demonstrates that the 7 benchmark subset reflects the larger population.

• Data revisions in the “Leal7a”  $^{239}\text{Pu}$  evaluated file have eliminated ~50% of the long-standing  $k_{\text{calc}}$  bias.

Calculated Eigenvalues<sup>(a)</sup> for a Selection of PST Assemblies  
Using Various  $^{239}\text{Pu}$  Cross Sections

Assembly	ENDF/B-VII.1	JEFF-3.1.2 <sup>(b)</sup>	JENDL-4.0 <sup>(b)</sup>	Leal7a <sup>(c)</sup> + e71	Leal7a (RR, nu, pfns only) + e71
PST1.4	1.00448	1.00127	1.00588	1.00199	1.00202
PST4.1	1.00383	0.99907	1.00482	1.00044	1.00044
PST9	1.01939	1.01367	1.02510	1.01543	1.01546
PST12.10	1.00412	0.99973	1.00498	1.00083	1.00080
PST12.13	1.00955	1.00468	1.01069	1.00611	1.00620
PST18.6	1.00472	1.00153	1.00557	1.00202	1.00208
PST34.4	1.00258	0.99999	1.00417	0.99922	0.99937
PST34.15	0.99742	0.99563	0.99844	0.99679	0.99707
<b>Average</b>	<b>1.00576</b>	<b>1.00195</b>	<b>1.00746</b>	<b>1.00285</b>	<b>1.00293</b>
a) MCNP calculations are for 250M histories; stochastic uncertainty is ~5 pcm. b) JEFF-3.1.2 and JENDL-4.0 $^{239}\text{Pu}$ only; remaining nuclides are ENDF/B-VII.1 c) “LEAL7a” evaluation provides revised resolved resonance parameters coupled to a joint ORNL/CEA evaluated $^{239}\text{Pu}$ file; the “LEAL7a (RR,nu,pfns)” file couples just these data to the existing ENDF/B-VII.1 $^{239}\text{Pu}$ file.					

# Time-line

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## ■ May 2013: CIELO WPEC Subgroup initiated

- Teams identified

## ■ Nov 2013: NEMEA7-CIELO: Main collaboration kick-off

- Refine scope of work, collaborators who will work on tasks
- Will result in detailed work plans, time line goals, for each nucleus

## ■ Next 2.5 Years:

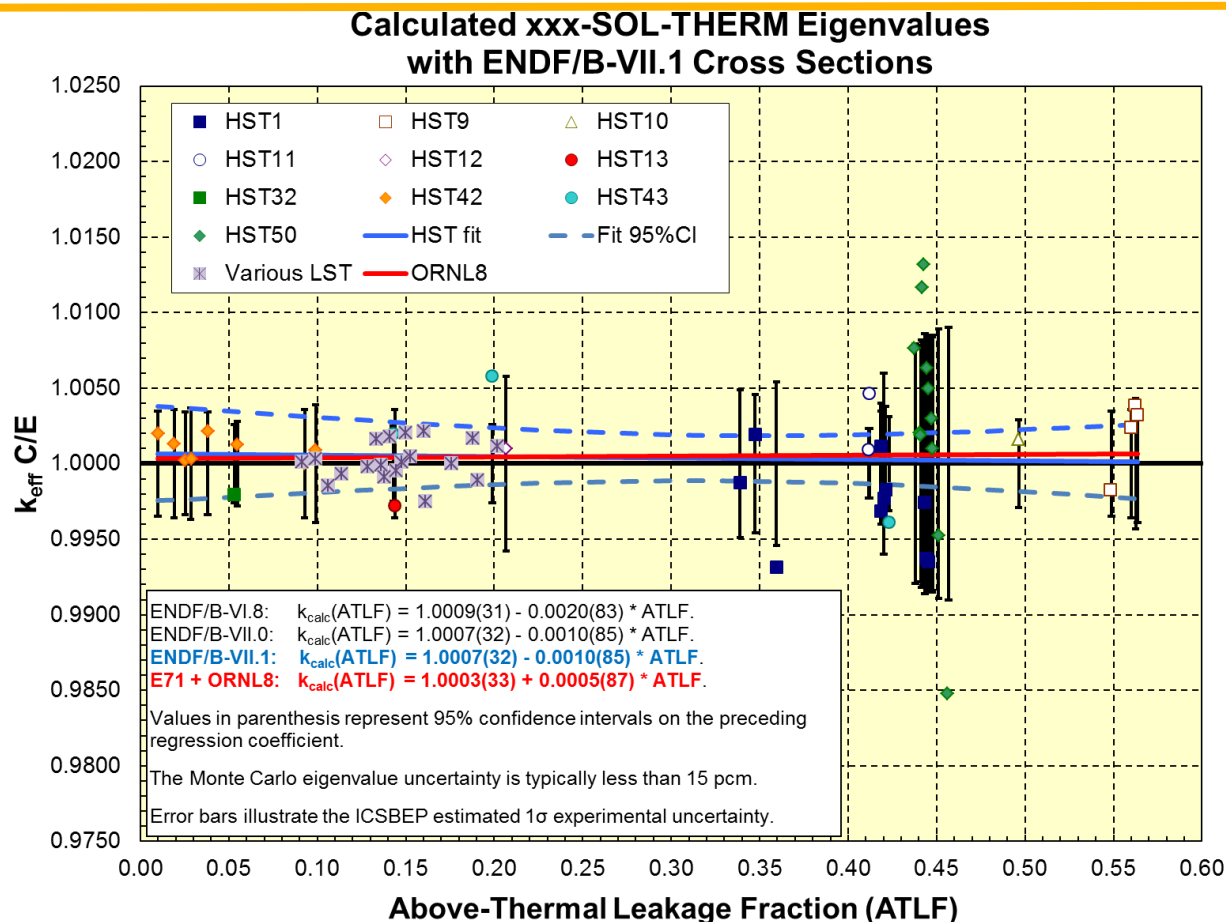
- Various collaboration meetings, continual email collaborative exchanges
- engagement with validation data testers continually
- Incorporate new IAEA standards results (fission, capture, scattering, ...)
- Explore interdependencies on criticality from the 6 CIELO nuclides

## ■ May 2016:

- Document conclusions from CIELO collaborations in WPEC report (& NDS paper?)
- Create formatted files that embody CIELO's initial conclusions

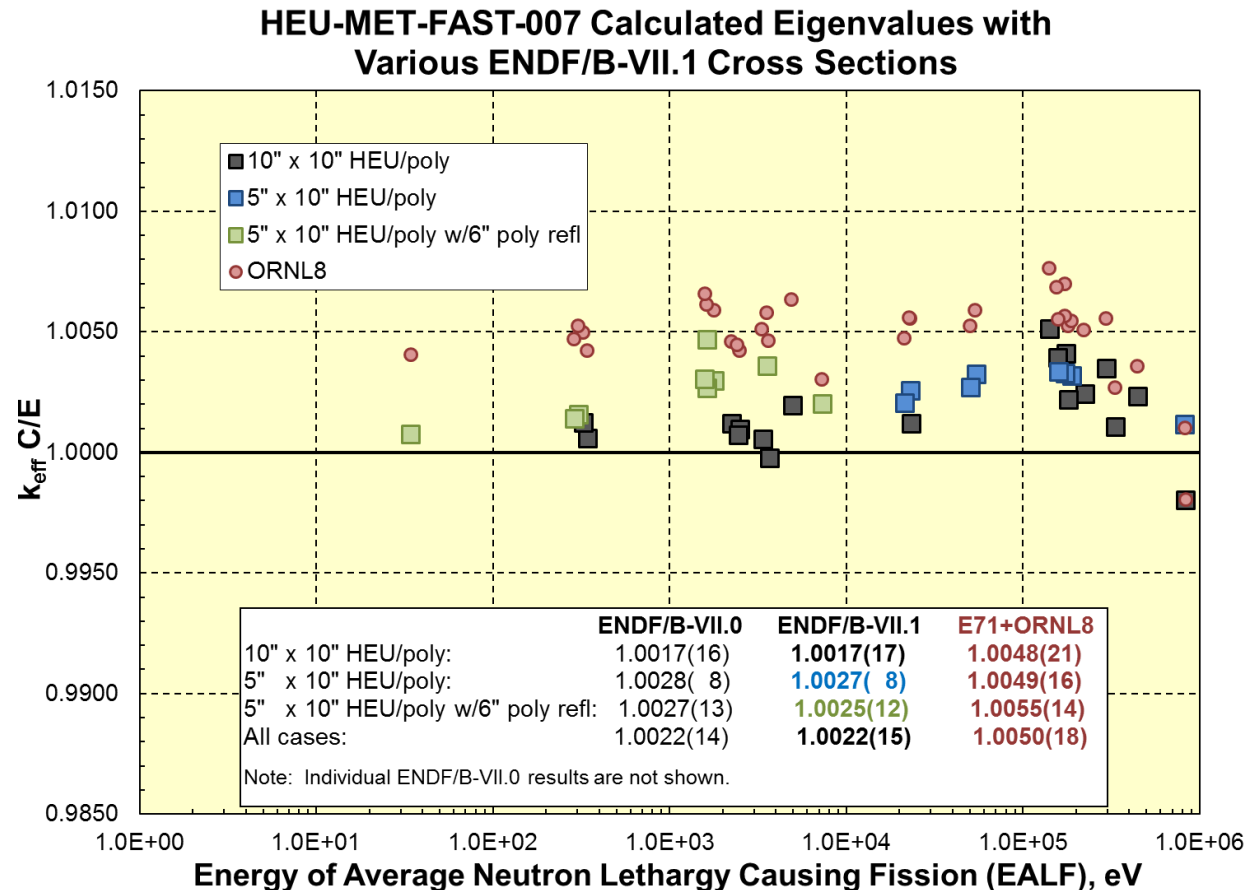
# HST Benchmarks - LANL testing of prelim Res file

- Regression fit to HST benchmarks versus ATLF has been excellent since ENDF/B-VI.3 (Lubitz).
- This excellent fit is retained with the latest **(ORNL8)**  $^{235}\text{U}$  resolved resonance file.



# HMF7 (HEU + CH<sub>2</sub>) : LANL testing of prelim Res file

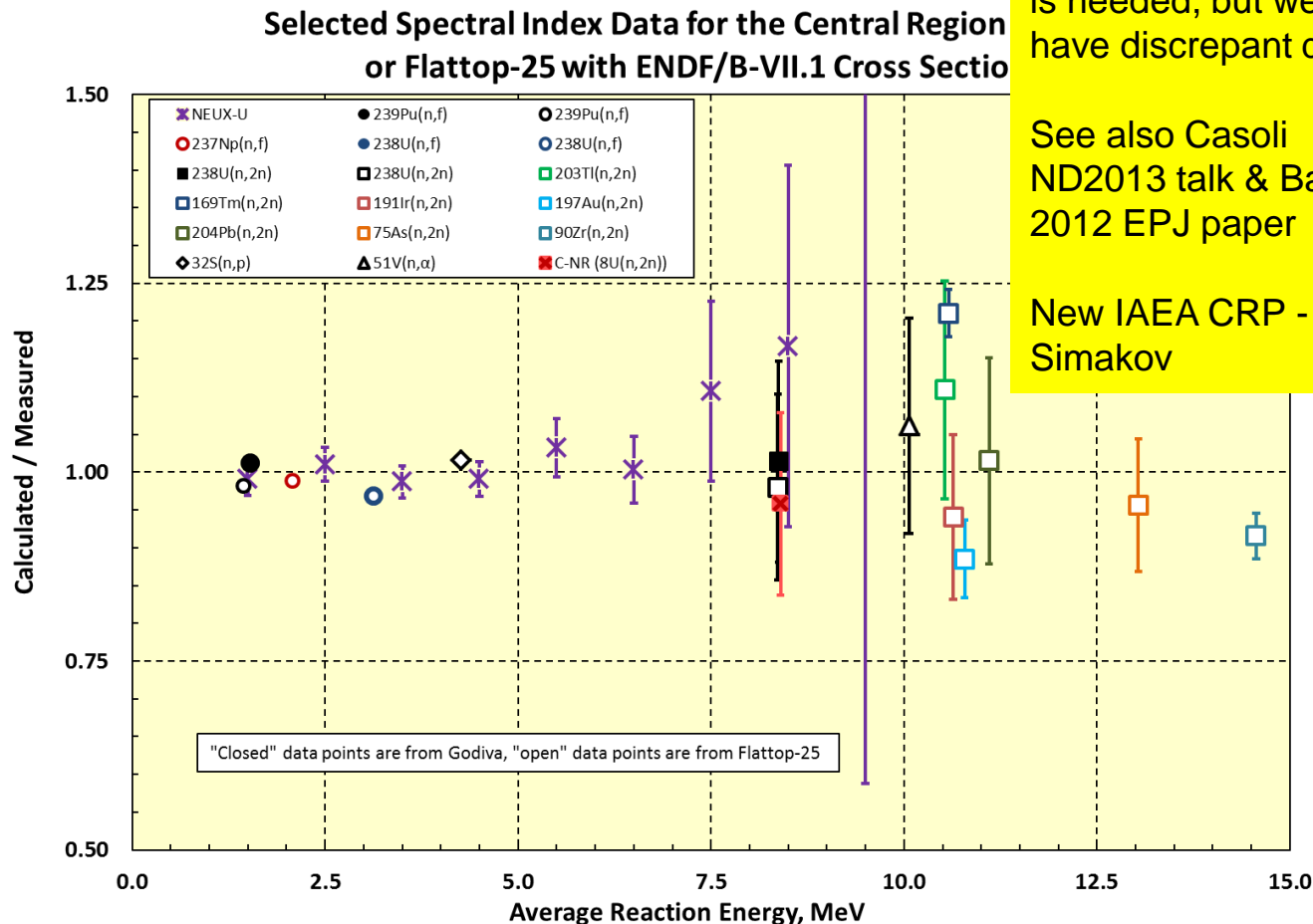
- HEU + poly system tests xs data over several orders of magnitude.
- E70 & E71 results are near unity at either energy extreme but are biased high in the intermediate energy range.
- This bias is worsened with the latest ORNL8 <sup>235</sup>U evaluated file.



# Reaction Rates in Fast Critical Assemblies Provide Integral Test of Prompt Fission Neutron Spectrum & (n,2n) Cross Sections - *Uranium-235 data*

$^{235}\text{U}$

With NUEX  
data added  
(Lestone)



Perhaps suggests a softer ENDF spectrum is needed, but we have discrepant data.

See also Casoli ND2013 talk & Bauge 2012 EPJ paper

New IAEA CRP - Simakov

Similar analysis for plutonium systems – see our ND2013 proceedings paper

# More Data Testing on Preliminary 235U ORNL Res Rile

**$k_{calc}$  Summary for Various Benchmarks and Cross Section Data Sets**

GODIVA (HEU-MET-FAST-001)		
Cross Section Set	Benchmark $k_{eff}$	$k_{calc}$
ENDF/B-VII.1	1.000(1)	0.99983(3)
E71 + ORNL8		0.99985(2)
J4 + ORNL8		0.99757(2)
CEA + ORNL8		0.99957(2)
Flattop-25 (HEU-MET-FAST-028)		
ENDF/B-VII.1	1.0000(16)	1.00285(2)
JENDL-4.0		0.99779(9)
E71 + ORNL8		1.00300(13)
J4 + ORNL8		0.99899(13)
CEA + ORNL8		1.00040(13)
Big-10 (IMF7, detailed model)		
ENDF/B-VII.1	1.0045(7)	1.00443(2)
JENDL-4.0		0.99710(7)
E71 + ORNL8		1.00471(8)
J4 + ORNL8		0.99764(11)
CEA + ORNL8		0.99901(11)



# More Data Testing on Preliminary 235U ORNL Res Rile

3

	HMI6.1	HMF6.2	HMI6.3	HMF72.3	HMI6.4	HMF72.1	HMF73
Benchmark $k_{eff}$	0.9977(8)	1.0001(8)	1.0015(9)	1.0016(69)	1.0016(8)	0.9991(24)	1.0004(16)
<u>endf/b-vii.1 ealf</u>	4.93 keV	10.1 keV	23.5 keV	40.8 keV	79.8 keV	223 keV	416 keV
	$k_{calc}$						
ENDF/B-VII.1	0.99293(2)	0.99690(2)	1.00076(2)	1.01236(2)	1.00730(2)	1.00852(1)	1.00807(2)
ENDF/B-VII.1 + e5 <sup>nat</sup> Cu	0.99264(2)	0.99723(2)	1.00168(2)	1.00762(10)	1.00767(2)		0.99663(2)
ENDF/B-VII.1 + <u>mit/ornl</u> <sup>63,65</sup> Cu	0.99304(15)	0.99709(15)	1.00086(15)	1.01254(10)	1.00791(15)		1.00720(14)
JENDL-4.0	0.99810(11)	1.00197(11)	1.00428(11)		1.00569(10)		1.00267(9)
E71 + ORNL8	1.00188(2)	1.00616(2)	1.00929(2)	1.01744(10)	1.01196(2)	1.00921(9)	1.00809(1)
J4 + ORNL8	0.99629(2)	0.99987(2)	1.00226(2)		1.00451(2)		1.00276(2)
CEA + ORNL8	0.99578(2)	0.99922(2)	1.00149(2)		1.00390(2)		1.00361(1)

JENDL-4.0 is <sup>iso</sup>U only; remaining cross sections are endf/b-vii.1.

J4+ORNL8 is the <sup>235</sup>U data set; remaining cross sections are endf/b-vii.1.

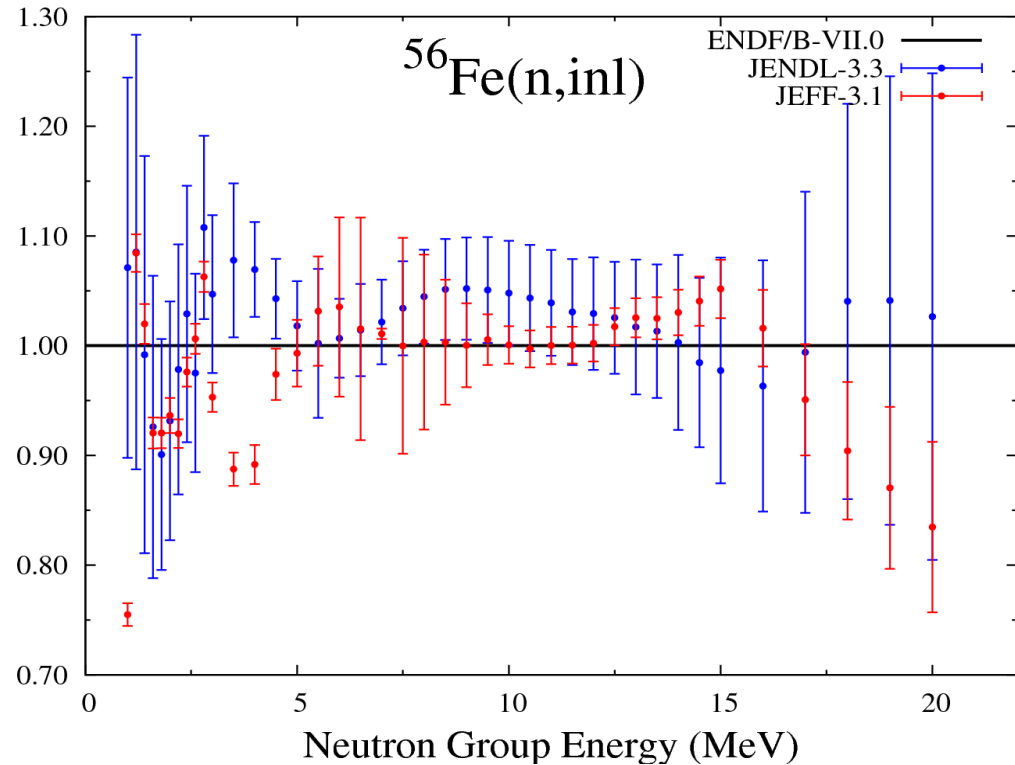
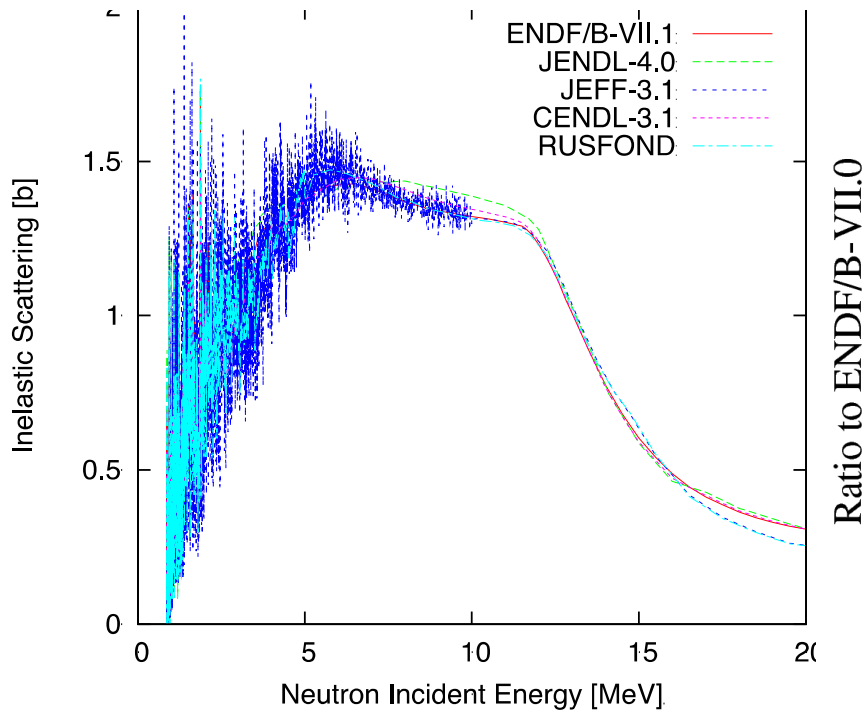
$k_{calc}$  values with a ~2 pcm uncertainty were run for 2 billion histories and include detailed multigroup tallies.

HMI6 has varying amounts of interstitial carbon; HMF72.1 has interstitial carbon steel (Fe); HMF72.3 has interstitial carbon steel (Fe) and polyethylene; HMF73 is HEU only ... all assemblies are surrounded by a thick copper reflector (i.e., HMI6, HMF72 and HMF73 are different flavors of ZEUS).

4



# $^{56}\text{Fe}$ : Advances Needed in Inelastic Scattering



Herman, Palmiotti

New measurements (IRMM) & SAMMY analyses in resonance region; new Hauser-Feshbach analyses at higher energies

Cecil Lubitz:

*“After several “preliminary” months on CIELO it’s clear that we have bitten off a big chunk. Get ready to chew.”*

160

## General

Intercompare evaluations, and identify goals for a new evaluation

JENDL is a new work (though adopts ENDF n,a); ENDF (JEFF uses ENDF) is a hybrid of KAPL work < 3.2 MeV, LANL (Hale et al) > 3.2 MeV - assess value of  
The 2005 ORNL work generated a resonance analysis for 16O, full R-matrix. Included angular distributions, n,alpha, and it has never been tested. Needed L

## Total, Elastic and Inelastic scattering

Compare existing evaluations and R-matrix analysis, and define path forward

At low energies, assess whether evaluations of elastic scattering indeed need to be lowered by ~3%, as proposed by Plompen, Lubitz, Roubtsov etc  
covariances for muon: Need reliable anisotropic 16O scattering uncertainties. Palmiotti thinks Gerry's present uncertainties are too small on muon.

## Capture

ENDF adopted JENDL's capture cross section to include resonance contribution - establish consensus to use this

## (n,a)

Review different evaluations (all largely same as ENDF)

Review previous data, and agree on scales - eg Bair & Haas had renorm their original data down by ~20%; Are Johnson data the same as these?

Review new data - Georginis (Geel), Khryachkov (IPPE) - contact physicists working on 13C(a,n) for astrophysics

The above new data approx confirm ENDF below 6 MeV but point to changes above

Intercompare R-matrix calcs (Hale, Kunieda, Leal)

Seek to understand why the above R-matrix evaluations, influenced by total cross sec data, suggest ~30% higher (n,a) than most measurements

Define an evaluation strategy... If theory contradicts these data, do we use data instead? Or do we conclude theory is right and measurements had a scale error

Assess whether evaluations (all now based on ENDF) above ~ 6 MeV need changing, if it is concluded new Geel data are more accurate than old Davis data

## Integral

Establish suite of integral validation tests, including k-eff, transmission, etc

2 benchmarks sensitive to oxygen data (+11 more benchmarks with water) are available in the SINBAD database

Broomstick experiment

Following WPEC SG? , With the existing (n,a) evaluations perform well, for the most part, on LEU solutions, Can the new eval perform well too

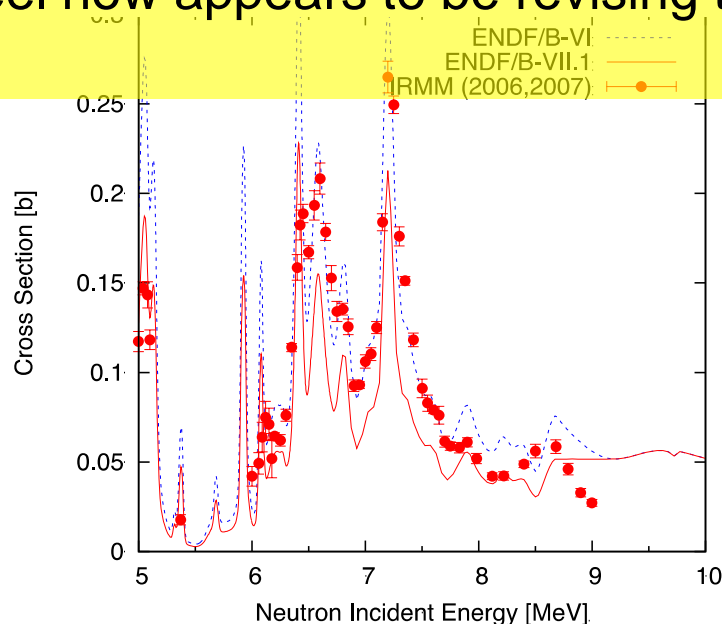
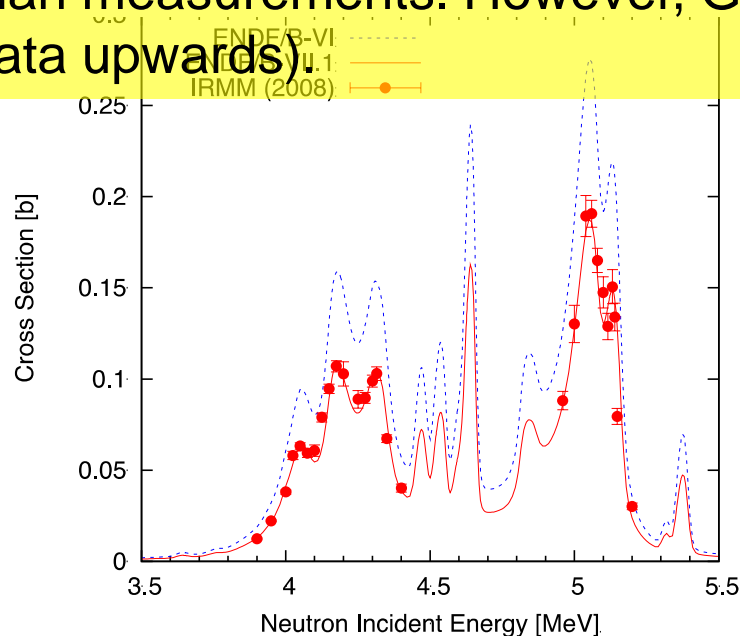
(n,a) impact at higher energies: Does this higher energy >6 MeV region impact any applications significantly (maybe medical applications)? Carlson notes M

check astrophysics constraints on 13C(a,n) reaction rate

# <sup>16</sup>O. Work is Needed to Reconcile R-Matrix Theory & Data & Maintain Criticality and Transmission Performance

An interesting case of **seemingly discrepant** information coming from *theory* versus *experiment*!

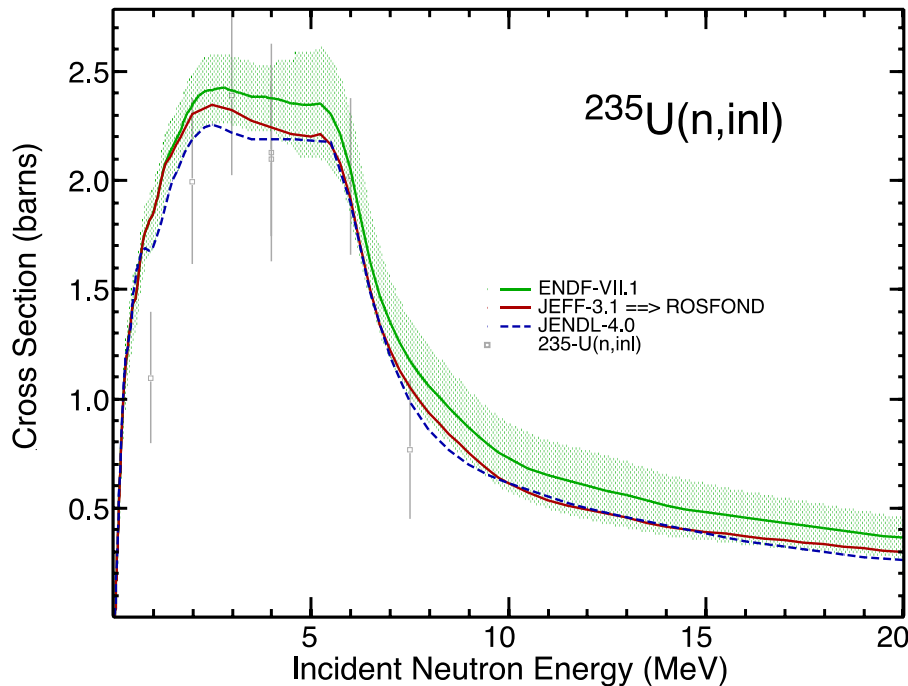
(R-matrix theory + total cross section data seems to suggest a higher (n,a) than measurements. However, Geel now appears to be revising their (n,a) data upwards)



Progress being made by Plompen, Lubitz, Roubstov, Hale, Kunieda, Leal, Moxon, Kopecky ... **Could TPC measure?**



# Uncertainties & Covariances for CIELO



Talou, with US covariance effort led by Oblozinsky, Smith, Herman, Hoblit:

Talou, Young, Kawano, Rising, Chadwick, DS  
112, 3054 (2011)

Covariances are now available in the major evaluated libraries. This allows us to:

- Focus experiment & theory efforts
- Calculate uncertainties on integral neutronic performance
- Provide feedback on cross section updates, via “adjustment” projects (SG33) or “assimilation”
- We’ll work with the new WPEC subgroup 39

## Summary ...

Join our CIELO collaboration

Thanks to IRMM & NEA/IAEA for support!





# BACKUP SLIDES

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# MCNP6 Production release, 2013

- MCNP6 = MCNP5 + MCNPX + several new features
- 2 DVD set will contain 5,X & 6 + ENDF 7.1 and > 1 Gbyte of documents
- MCNP 5/X/6 Beta 2 had 2,452 copies sent out in FY12 and more than 11,000 in the last 11 years!
- See “Initial MCNP6 Release Overview” Nuclear Technology, Dec 2012

