

3 Dimensional Magneto-hydrodynamic modeling of liner implosions on the Z Generator

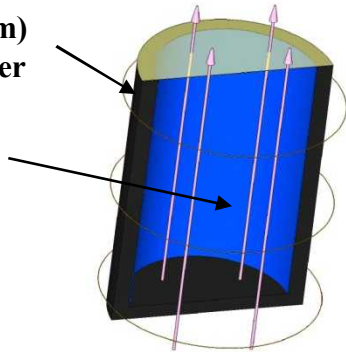
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S.A. Slutz, M.E. Cuneo, M.C. Herrmann,

* Sandia National Laboratories is a multi-program laboratory managed and operated by Sandia Corporation, a wholly owned subsidiary of Lockheed Martin Corporation, for the U.S. Department of Energy's National Nuclear Security Administration under contract DE-AC04-94AL85000.

Magnetized Liner Inertial Fusion (MagLIF)* may be a promising path to high fusion yields on Z

Metal (beryllium)
Cylindrical Liner

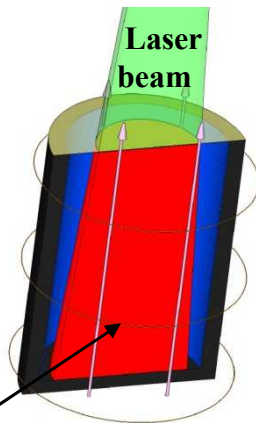
Cold deuterium/tritium
gas (fuel)



**Idea: Directly drive solid liner
containing fusion fuel**

1. An axial magnetic field (B_z) is applied
to inhibit thermal conduction and
enhance alpha particle deposition

Laser
beam

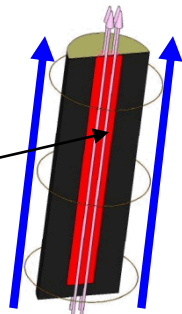


2. Z Beamlet preheats the fuel

Liner Integrity Critical!

Laser
preheated
fuel

Compressed
axial field



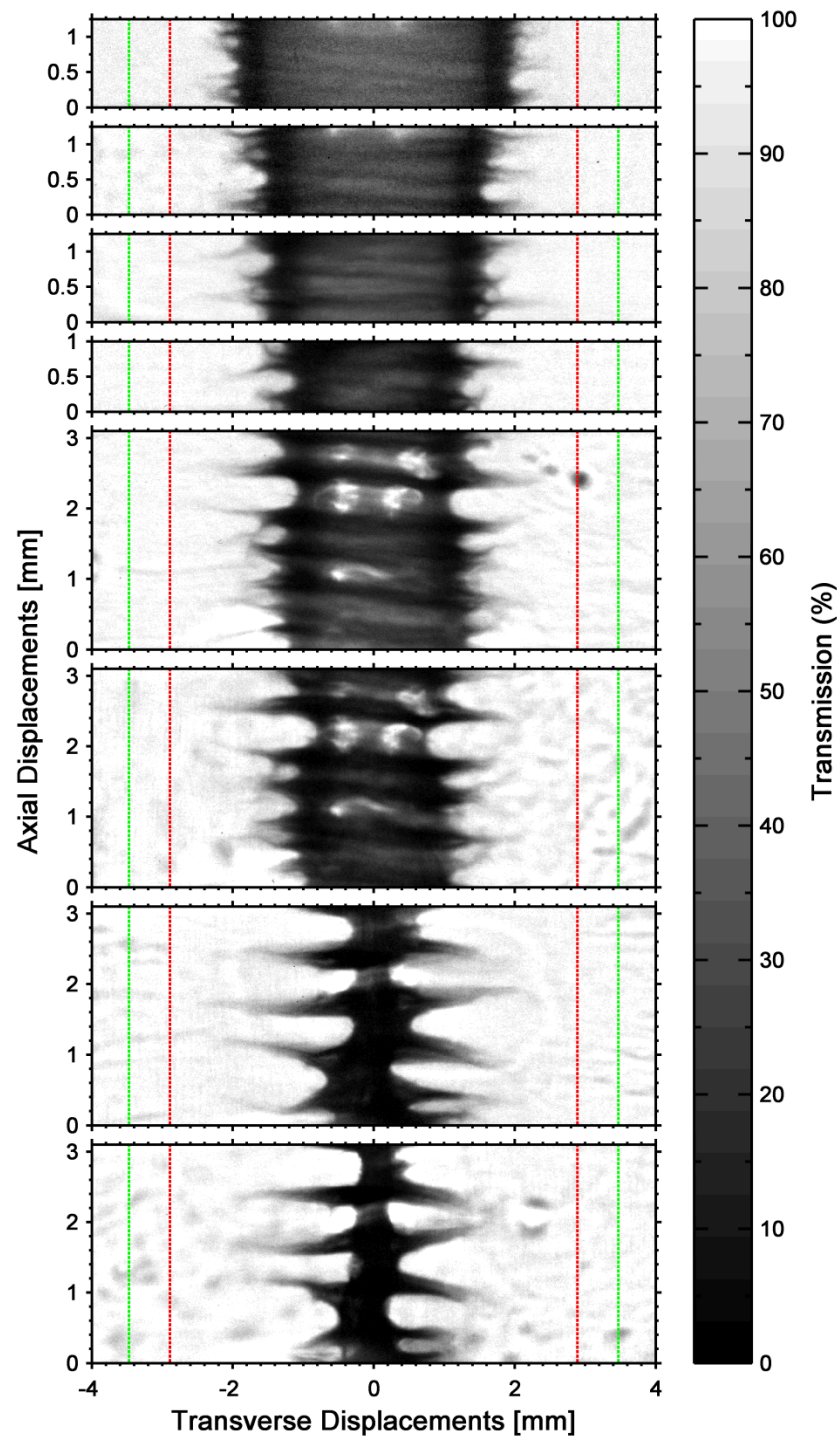
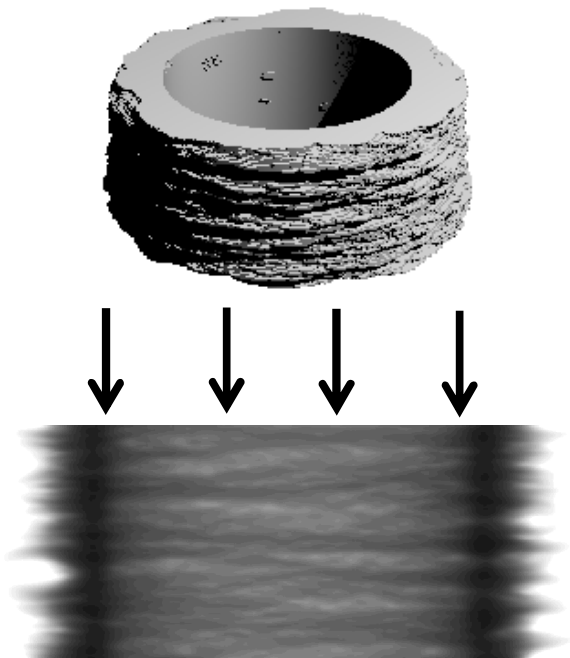
3. The Z accelerator efficiently
drives a z-pinch implosion

*S. A. Slutz *et al.*, Phys. Plasmas **17**, 056303 (2010).

Liner Implosion

Experiments have been performed on the Z generator to study the evolution instabilities in Be liner implosions.

6 keV Transmission Radiographs taken of imploding liners



GORGON 3D resistive MHD code used to study the development of implosion instabilities

GORGON – 3D Resistive MHD

Fixed square grid finite volume hydrodynamics

Single fluid – separate electron and ion temperatures

Explicit electro-magnetic field solution (wave equation in vacuum / diffusion equation in plasma)

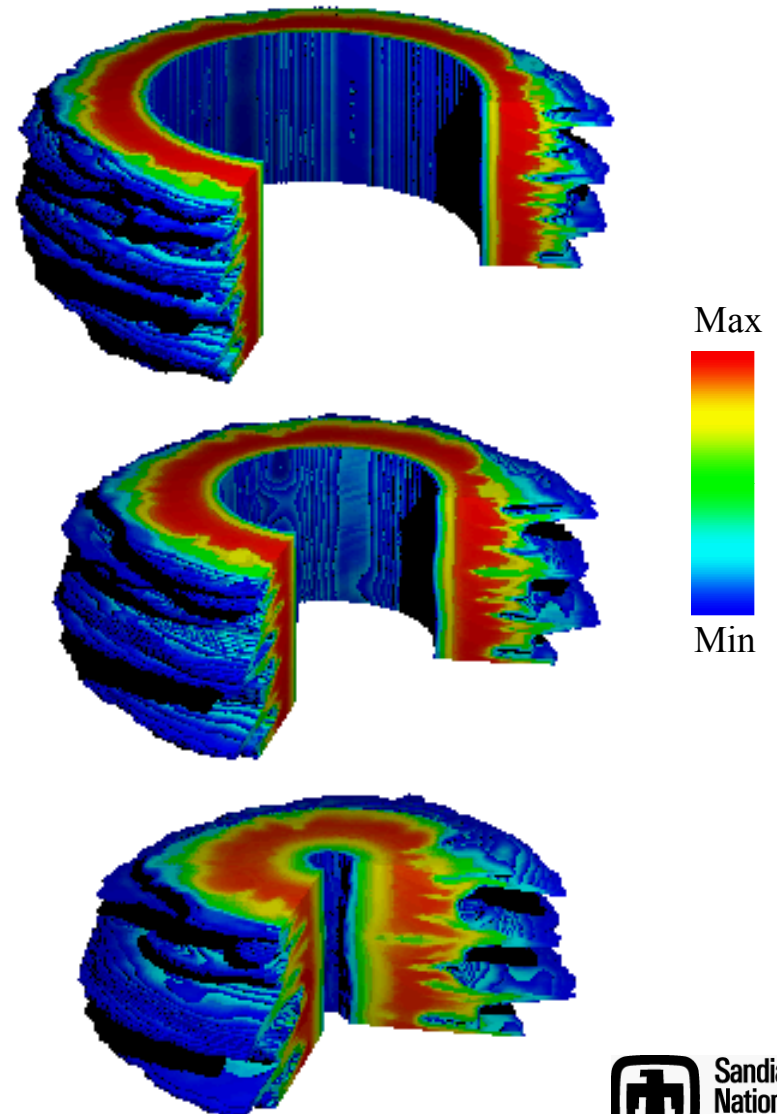
Van Leer Advection

Driven from measured generator current

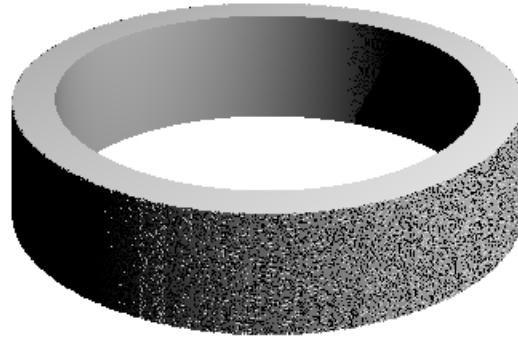


Density used to construct synthetic radiographs for direct comparison with experimental measurements

Log Density

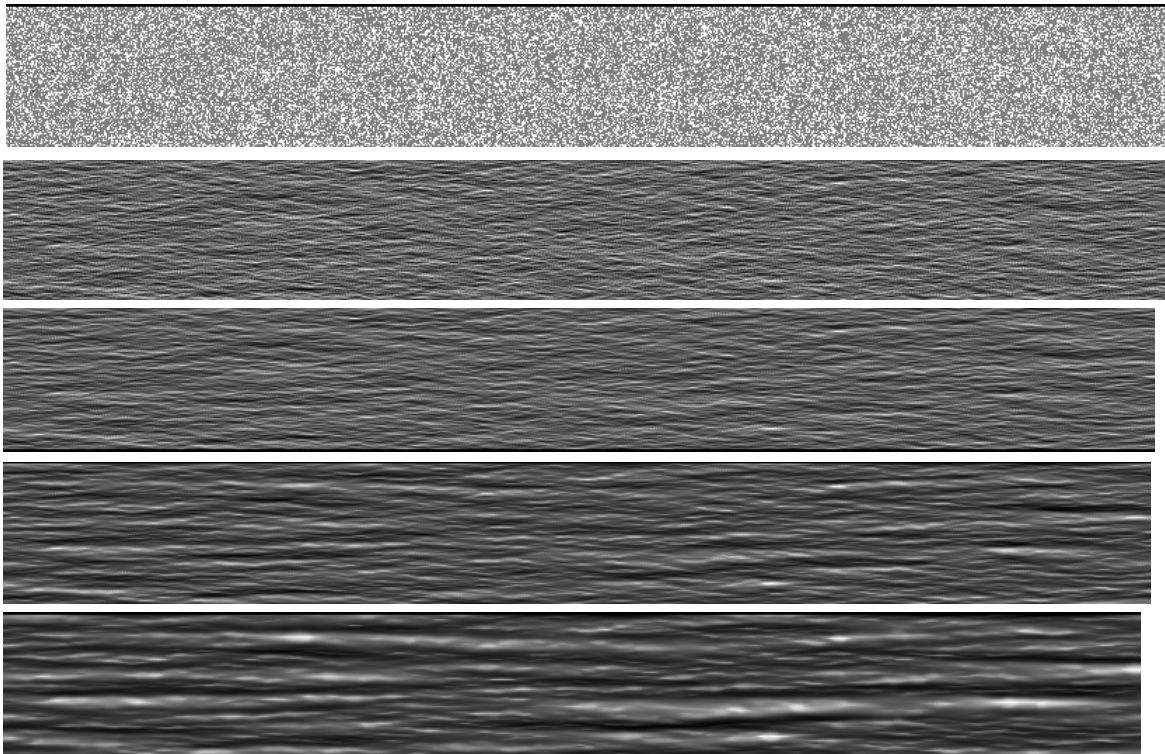


Calculations are initialized with a randomized 20 micron surface roughness

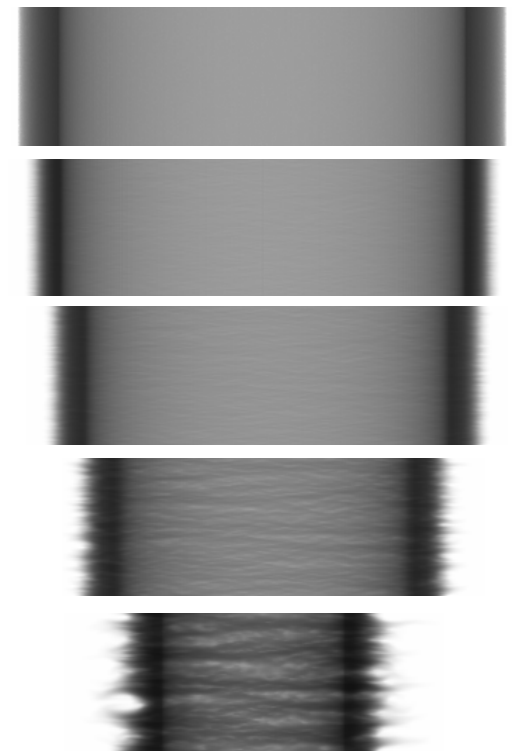


This initialization is not intended to study how different surface structures develop, and does not attempt to reproduce initial liner surface

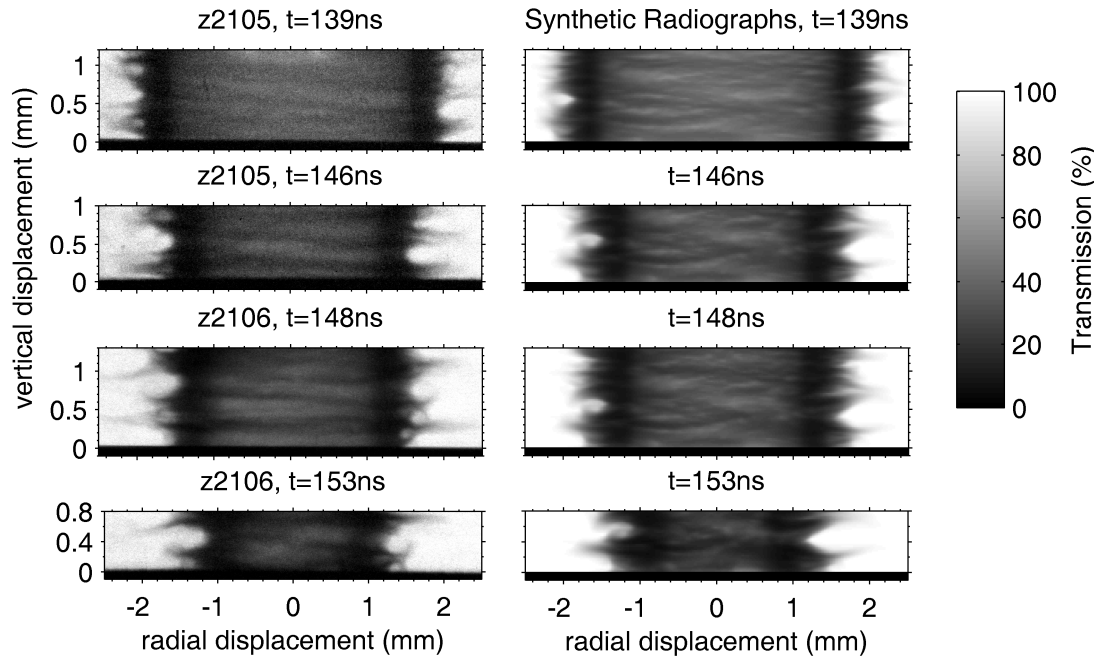
Liner radiograph
unwrapped to study
growth of surface
structures



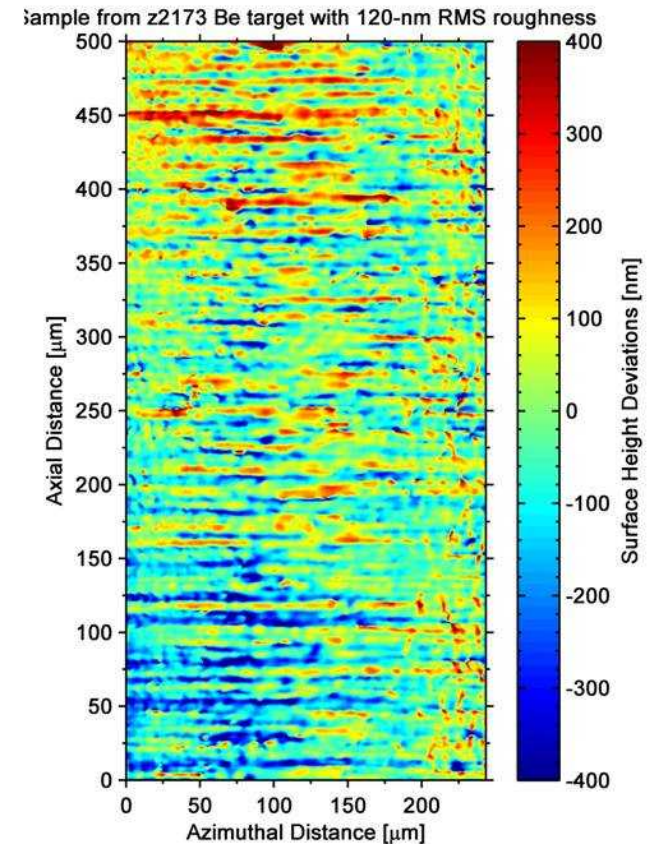
Synthetic Radiographs



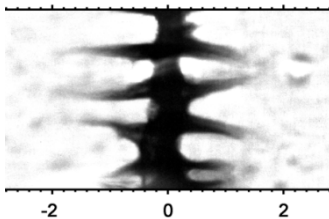
Reasonable early time agreement with measured radiographs, but late time discrepancies



Measured surface roughness is not random, but has some azimuthal correlation



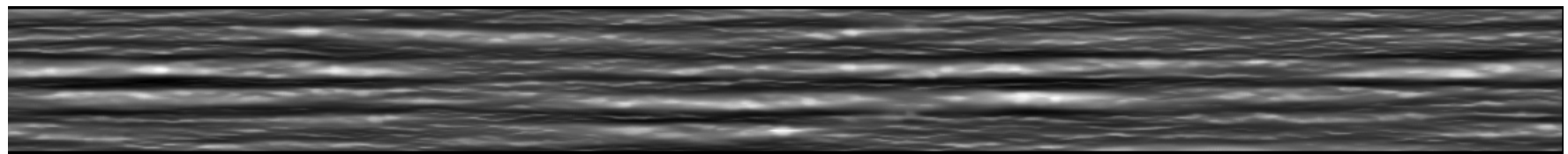
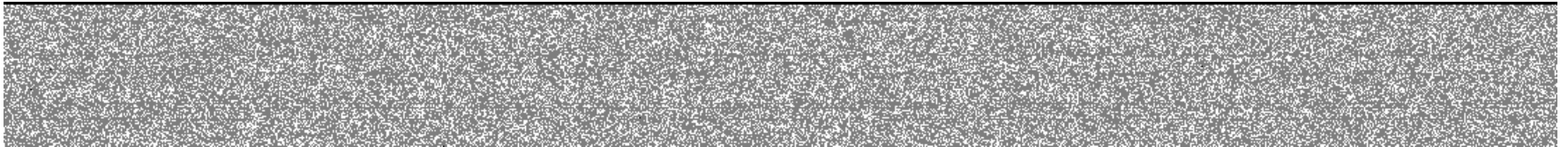
Level of disruption different at late times





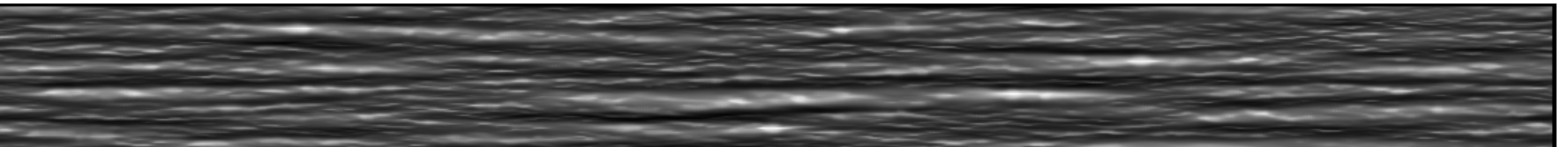
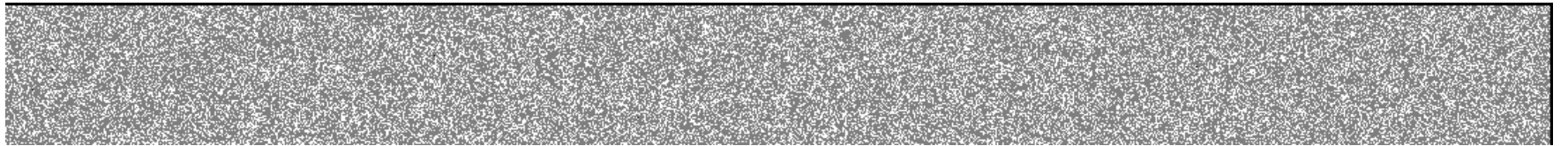
A small degree of azimuthal correlation is able to persist late in time

20 % random number generator bias correlation at 6 random heights

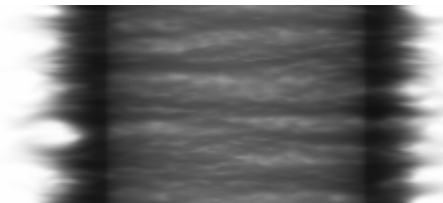


3110ns

No Correlation



No correlation

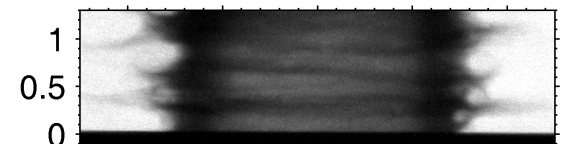


Correlation



Experiment

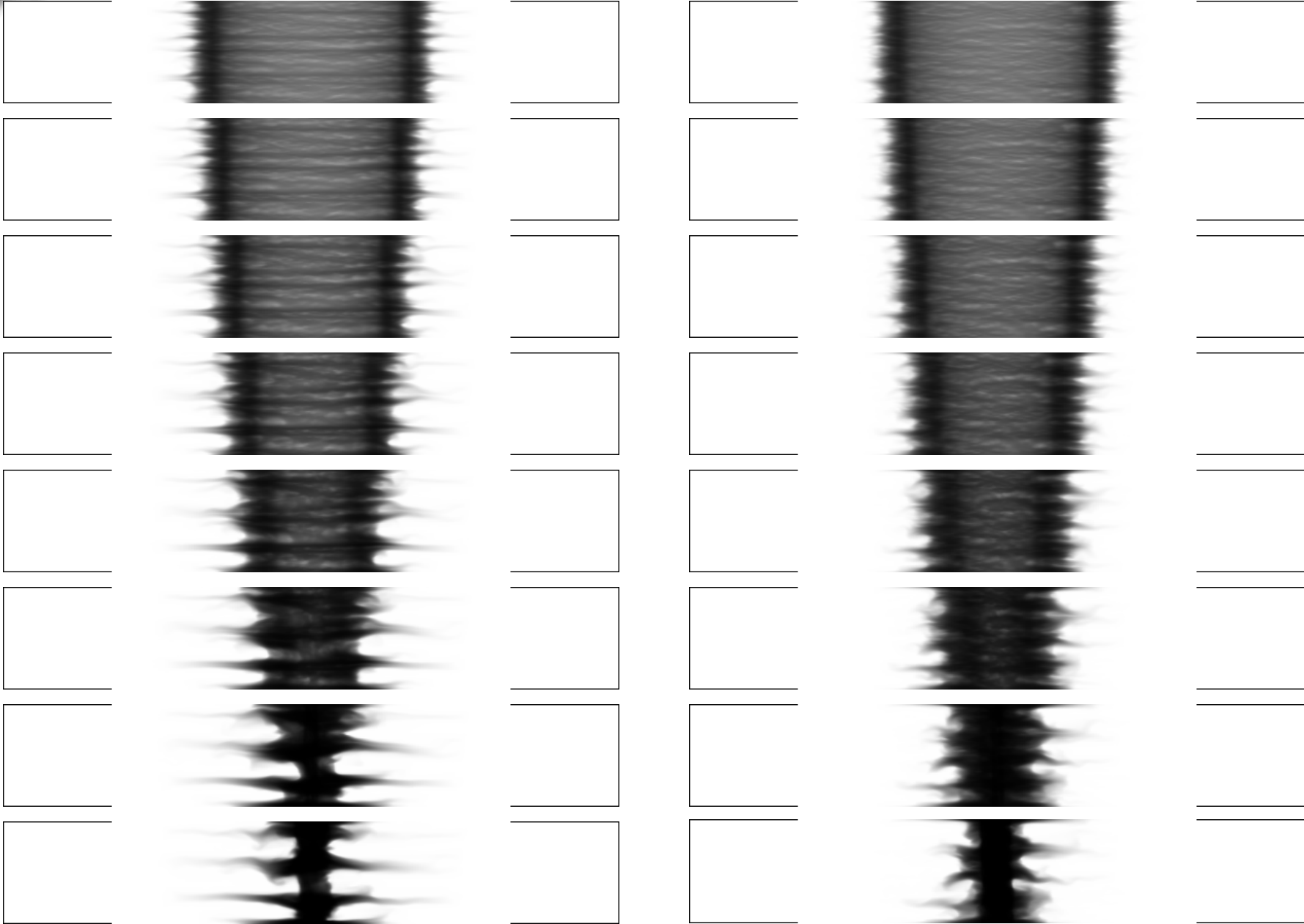
z2106, t=148ns



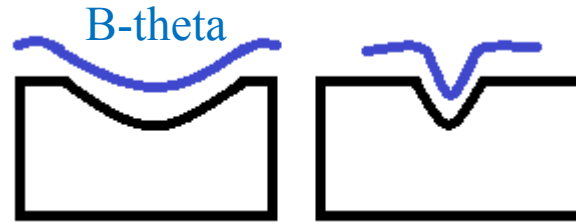
Effect of correlation more evident later in time

Azimuthal Correlate pert.

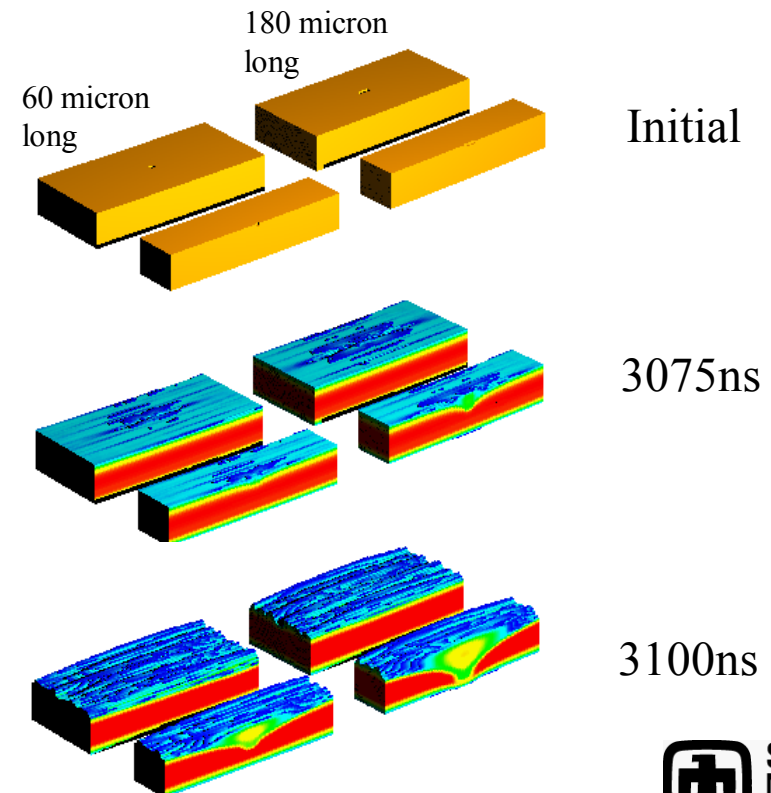
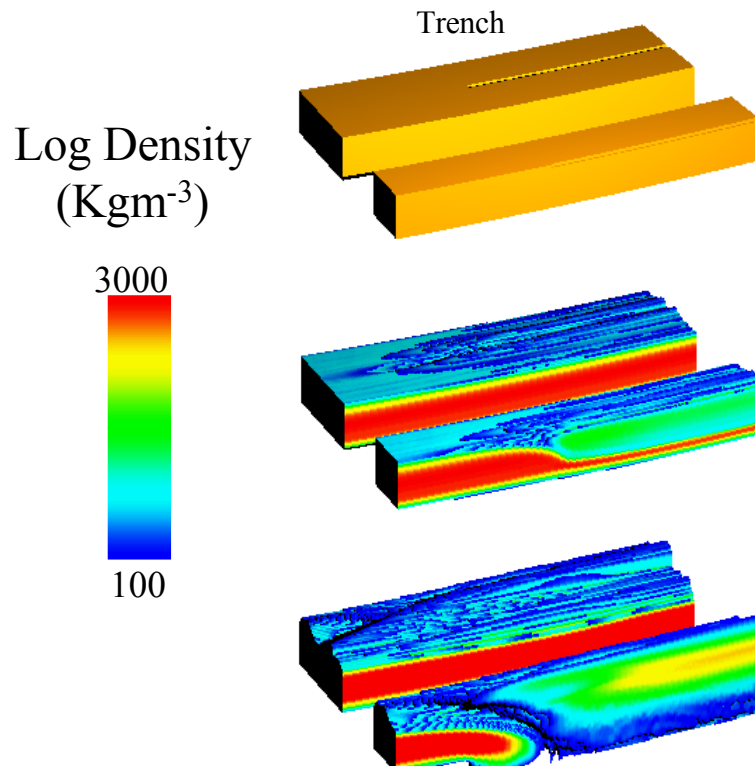
random pert.



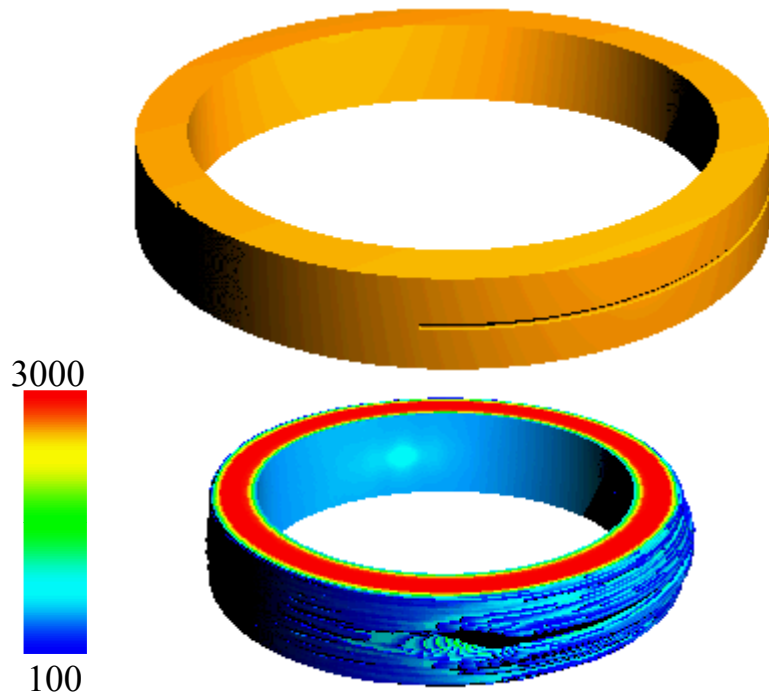
The degree of azimuthal correlation of an instability affects how much it can penetrate and disrupt an imploding shell



Field line tension will not allow us to simply drill a hole through the liner

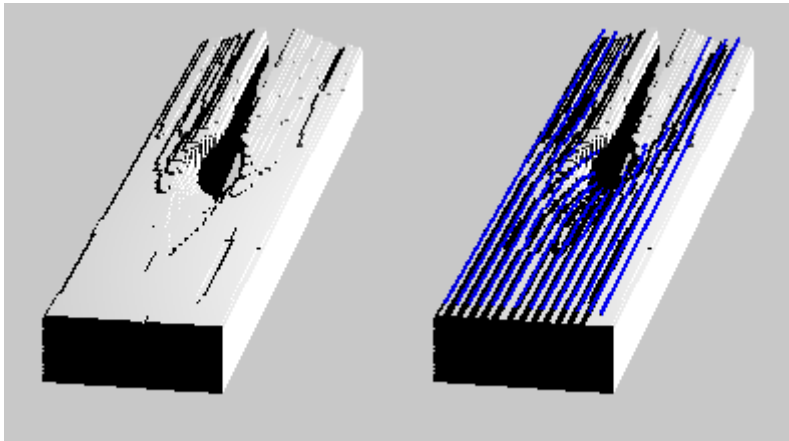
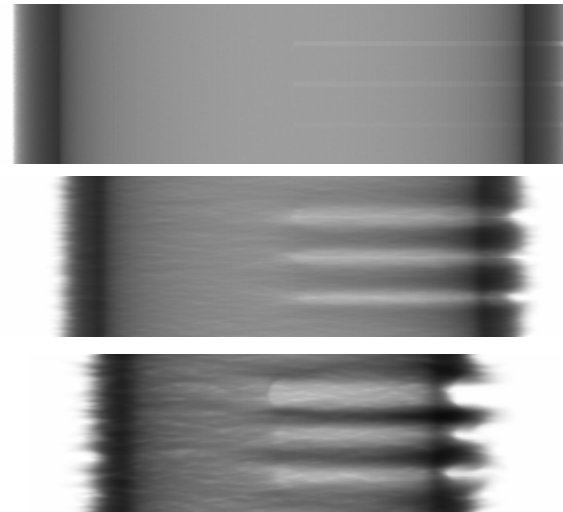


To propagate the instability azimuthally, an instability has to plough its way through a lot of dense material



The ability of established features to propagate azimuthally is quite limited, so gain azimuthal extent and disrupt the shell requires the amalgamation of adjacent features. So trouble develops if initial conditions correlated them to start with.

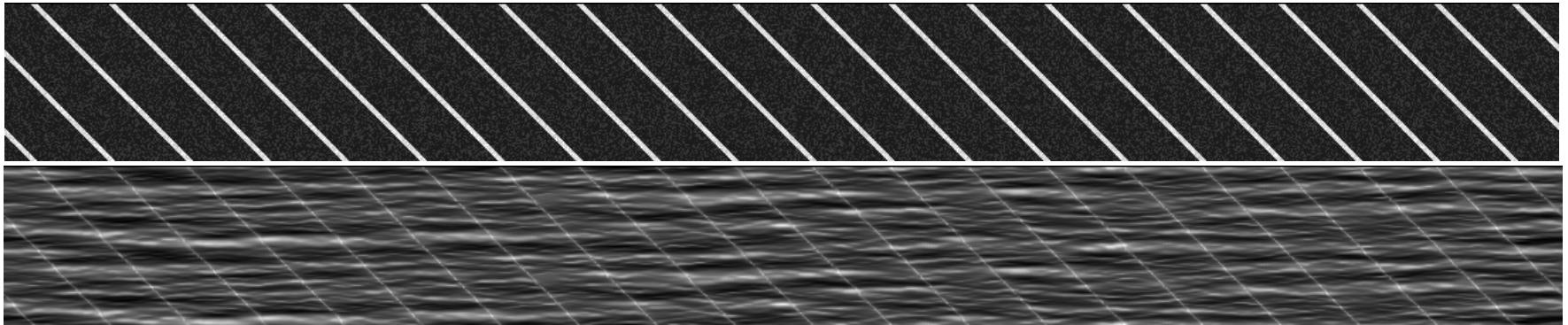
Synthetic Radiographs



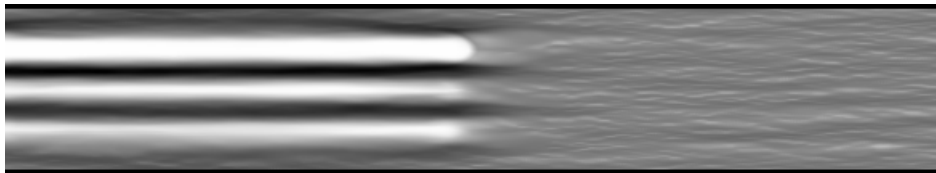
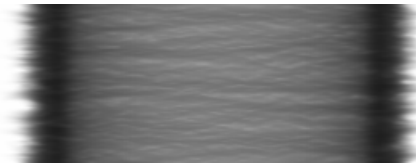
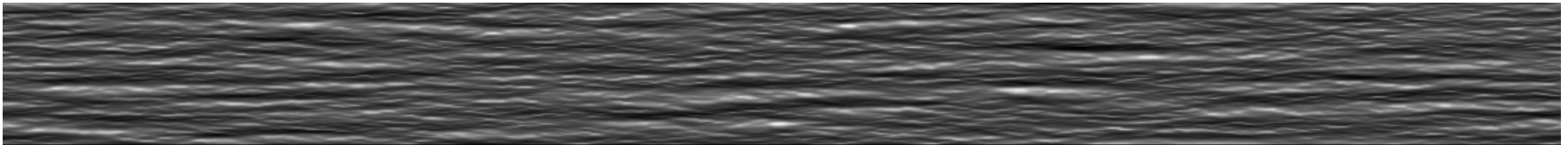


Reorientation the perturbation away from azimuthal prevents it from growing

3095ns Periodic boundary conditions in Z to study helical perturbation growth



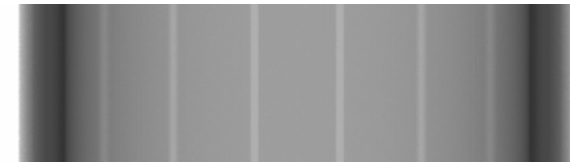
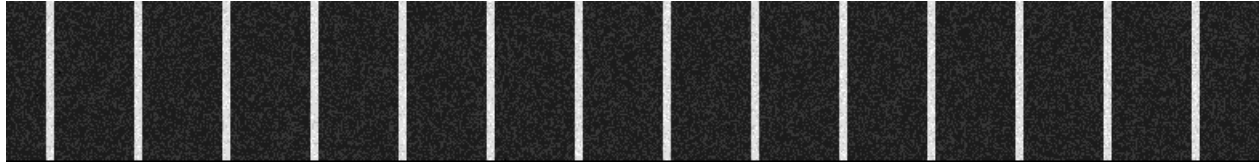
Random seed



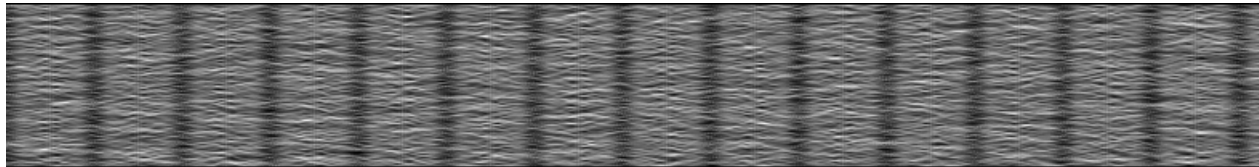
Stark contrast to
disruption from an
azimuthal groove

Vertical features do not grow at all

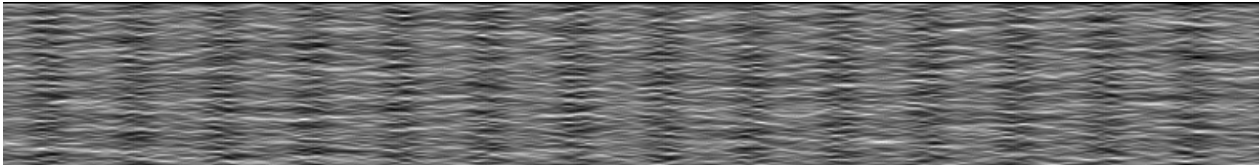
Initial



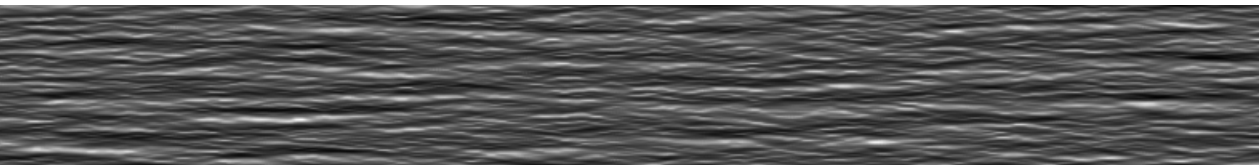
3050



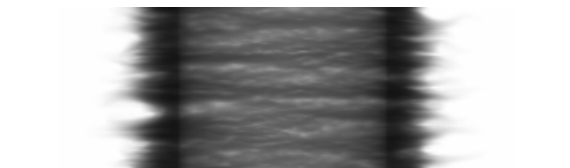
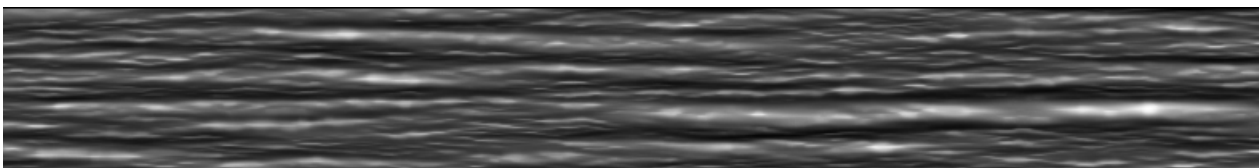
3070



3090



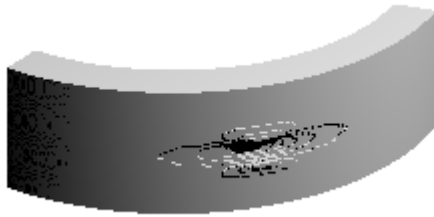
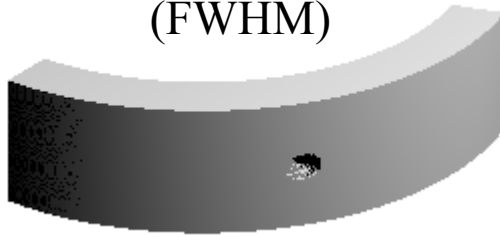
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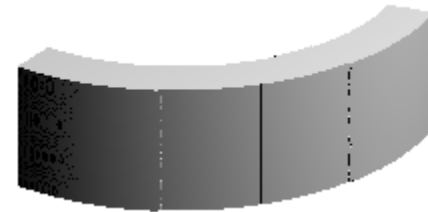
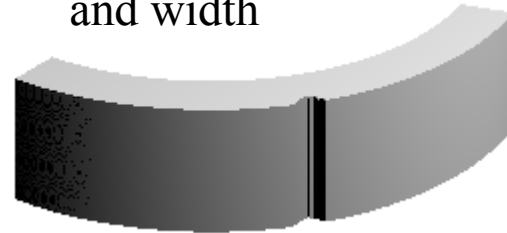
Initial 80 micron deep groove

For a given perturbation, a vertical groove is actually preferable to surface hole of same width and depth

~100 micron hole
(FWHM)



Groove of same depth
and width



Conclusions

If surface perturbations are seeding late time instabilities
then:

- To disrupt an imploding shell a perturbation needs some azimuthal extent.
- Perturbations grow and correlate quite slowly in the azimuth
- Real damage can be done if the initial seed perturbation has any azimuthal correlation
- If an infinitely smooth liner is an unrealistic ideal, then there are surface structures we can pick to leave behind that are better than others

