

# Two-Dimensional Radiation MHD Model Assessment of Designs for Initial Argon Gas Distributions to be Imploded on the Z machine\*



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# Outline

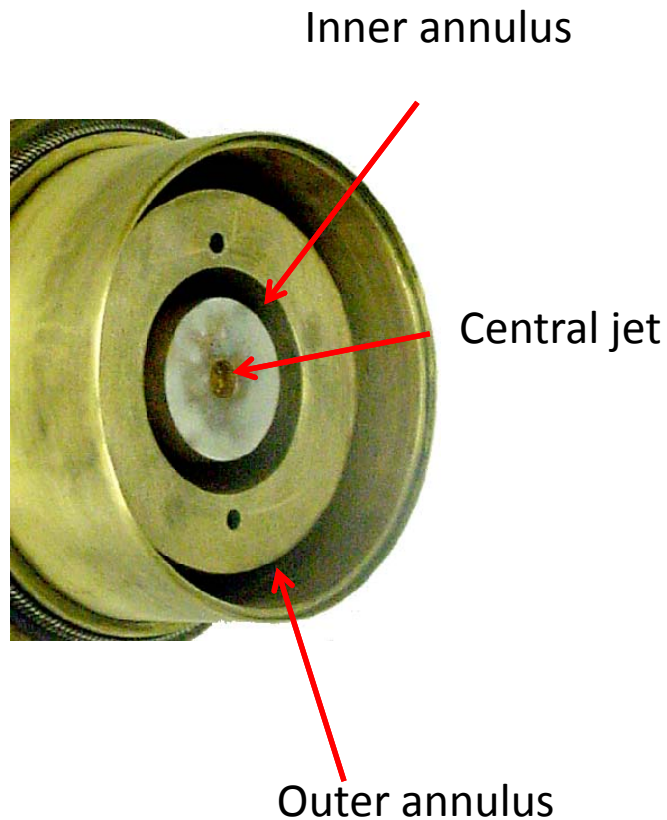
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In the near future scientists at Sandia National Laboratories will be performing, for the first time, argon gas puff experiments on the refurbished Z machine. As part of the effort to determine the initial argon loads to be deployed, we theoretically assess the K-shell emission and stability properties of the gas distributions generated by the new Alameda Applied Sciences (AASC) 8 cm diameter double-annulus nozzle with a central jet that was recently constructed for these experiments .

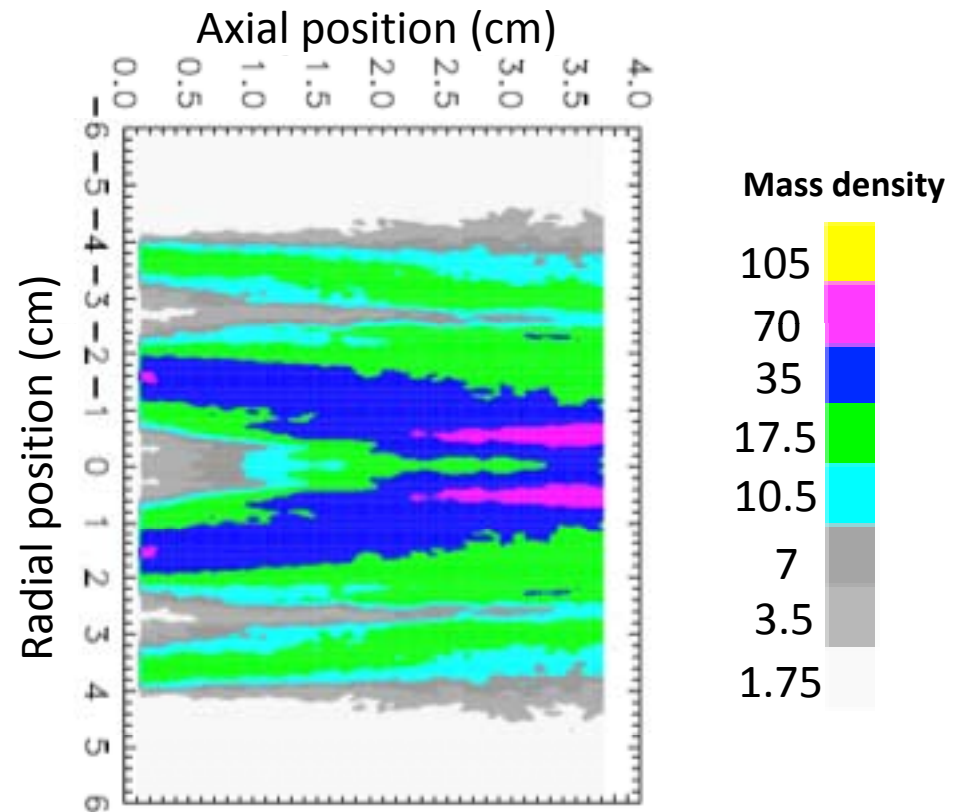
- Review past performance of pre- refurbished Z argon experiments and their implications for future ZR argon experiments.
- Employ 2D rad-MHD model to compare the stability and K-shell emission properties of gas distributions generated by the new nozzle with the distribution used in the pre-refurbished Z experiments (Titan1234 nozzle).
- Use 2D rad-MHD model to determine optimal K-shell producing AASC nozzle gas distribution.
- Summarize results of this argon gas puff assessment for ZR

## Summary of pre-refurbished Z argon experiments

The pre-refurbished Z argon experimental results are well described in [H. Sze, P. L. Coleman, *et. al.*, *Physics of Plasmas Letters*, 8, 3135 (2001).]

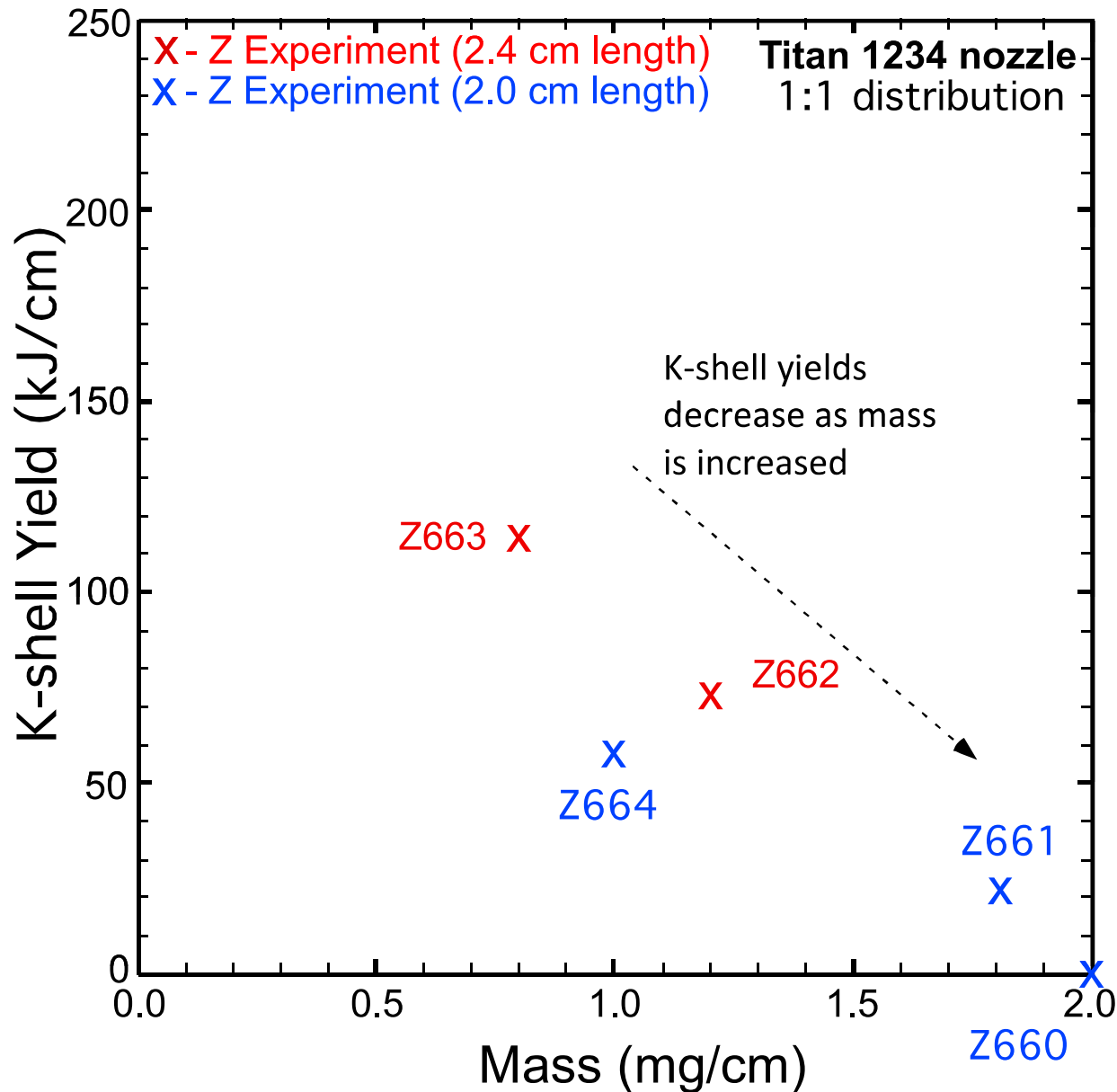


**Generic 8 cm diameter multiple shell nozzle**



8 cm diameter, Titan 1234 Nozzle -  
1:1 gas distribution

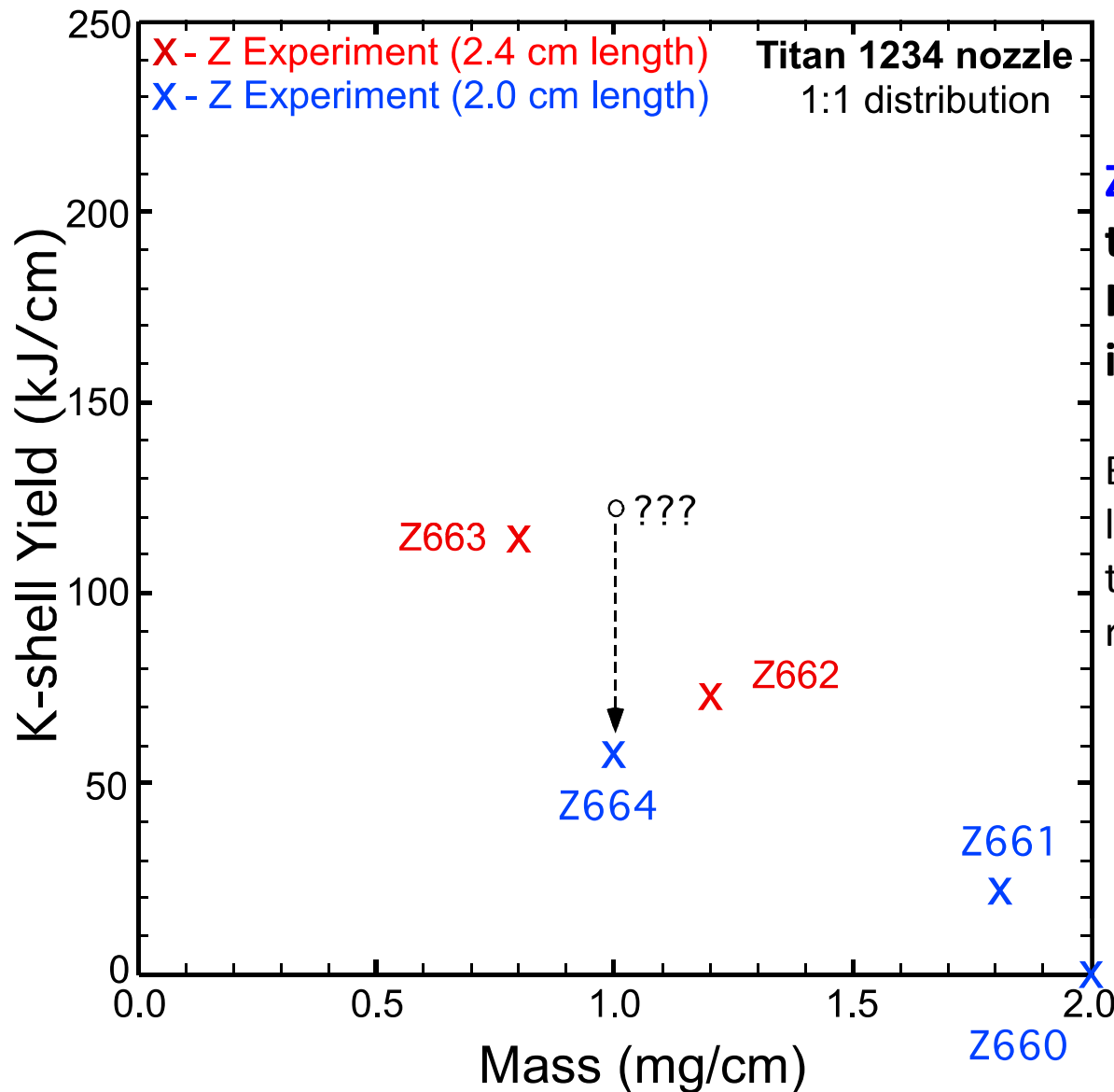
## Summary of pre-refurbished Z argon experiments



**Current loss was observed in most of these experiments, especially for Z660 the most massive load.**

H. Sze's speculation – current loss due to UV light from implosion illuminating the convolute power feed

## Load Instability likely played a role in pre-refurbished Z experiments



**Z664** had more energy/cm than **Z662**, why was K-yield/cm less? H. Sze suspects it is due to axial non-uniformity.

By using a reduced pinch length for Z664 they eliminated the region of the pinch that had radiated best in Z662 (near anode)

Based on earlier Z argon experiments one expects for ZR:

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- 1) Current loss could be an issue, especially at large mass.
- 2) K-shell yields of **~274 kJ** obtained on pre-refurbished **Z** were in  $I^2$  scaling regime, well designed loads on ZR should also be in this regime. (Estimate ~35% more energy into a **ZR** load than for Z, which is **~ 370 kJ** of K-shell emission).
- 3) Load instability likely played a role in earlier Z experiments and is likely to do so on ZR.
- 4) Since the AASC nozzle is new and never been tested, we don't know its stability and K-shell behavior.

## Modeling the K-shell emission and stability properties of the AASC nozzle gas distributions

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- To theoretically model the non-linear growth of instabilities and multi-dimensional plasma motion that is present in gas puff implosions we employ *Numerex's* Mach2 – two dimensional magneto-hydrodynamics code.
- To account for the non-LTE kinetics, opacity , and non-local transport of radiation that affect the atomic populations of these high temperature K-shell emitting plasmas we incorporated into the Mach2 code a self-consistent EOS calculation that models this physics. It is called the **tabular collisional radiative equilibrium model -TCRE**

Mach2 reference – R. E. Peterkin, M. H. Frese, and C. Sovienc, J. Comput. Phys. **140**,148 (1998).

TCRE reference – J. W. Thornhill, J. P. Apruzese , *et. al.*, Phys. Plasmas **8**, 3480 (2001).

## Modeling the K-shell emission and stability properties of the AASC nozzle gas distributions

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There are three measured gas distributions analyzed to characterize the stability and K-shell emission properties of the AASC nozzle



1) AASC [31\_30\_250]

2) AASC [22\_30\_250]

3) AASC [31\_58\_0] -----No central jet

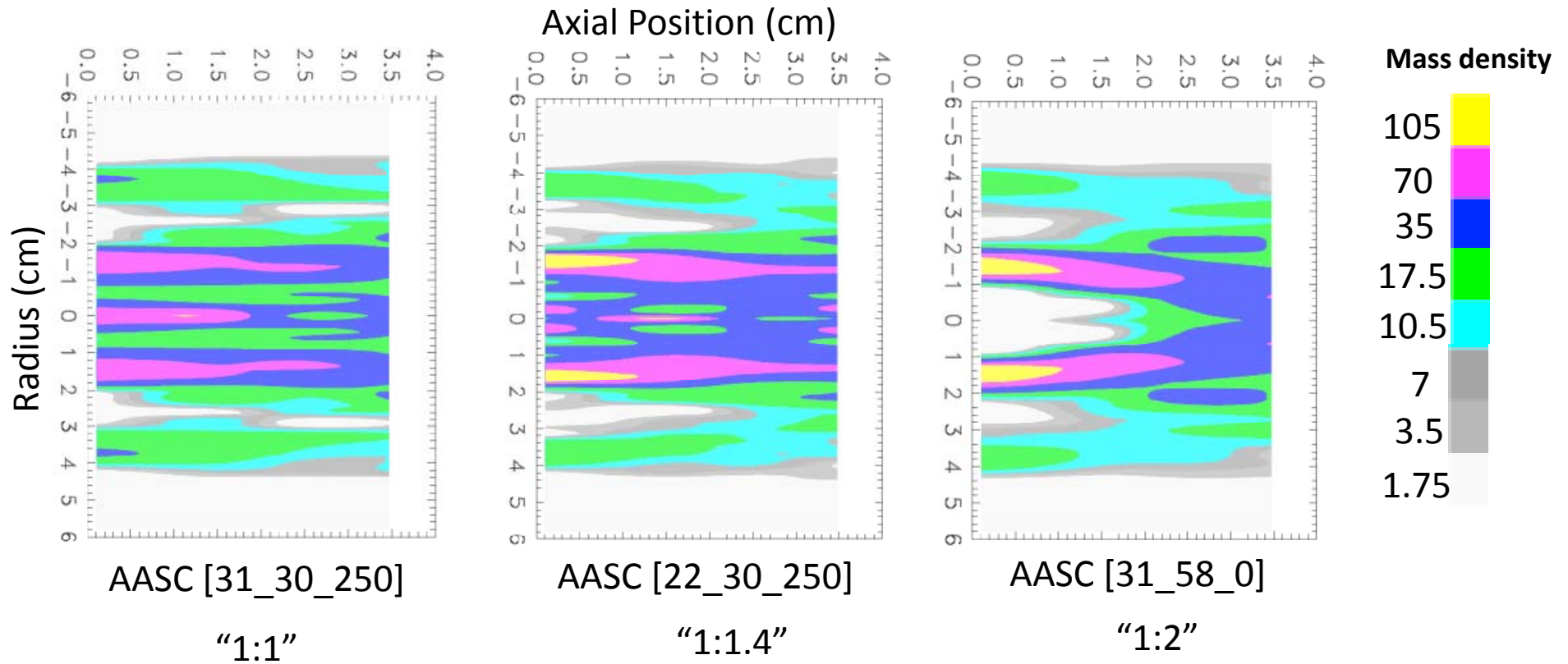
Pressure (psi) in  
**outer** annulus

Pressure (psi) in  
**inner** annulus

Pressure in  
**central jet**



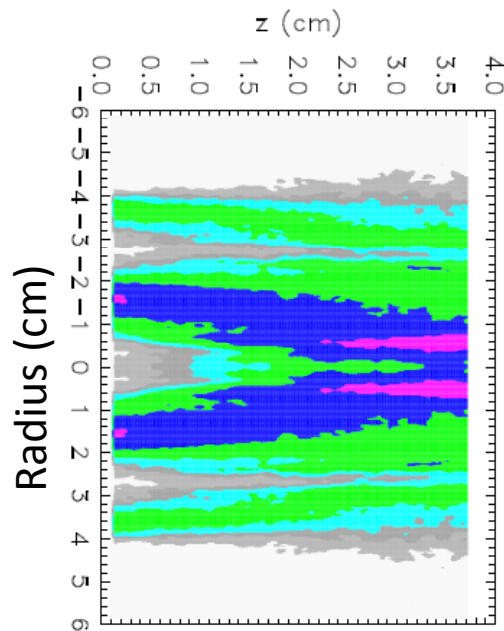
# Modeling the K-shell emission and stability properties of the AASC nozzle gas distributions



- Distributions measured using Fiber Optic Interferometry
- Measurements taken along 4 axial slices (0.5, 1.5, 2.5 and 3.5 cm) @ ~20 radial positions
- Bicubic spline used to interpolate between data points
- The load mass is varied in the calculations by renormalizing each of the above three distributions to the mass of interest.

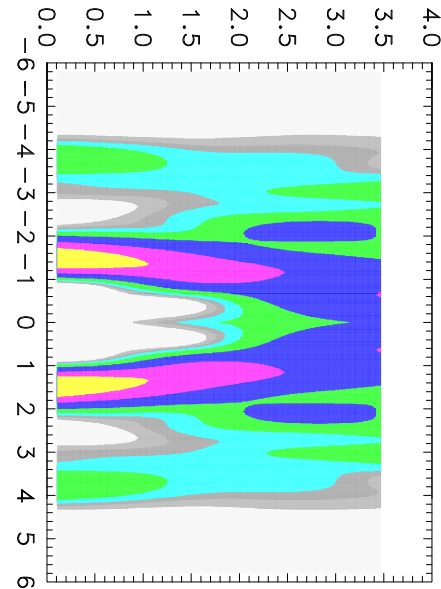
# Comparison of the Titan 1234 and AASC gas distribution's K-shell and stability properties

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**Titan 1:1**

Laser induced fluorescence  
(120 axial x 540 radial )



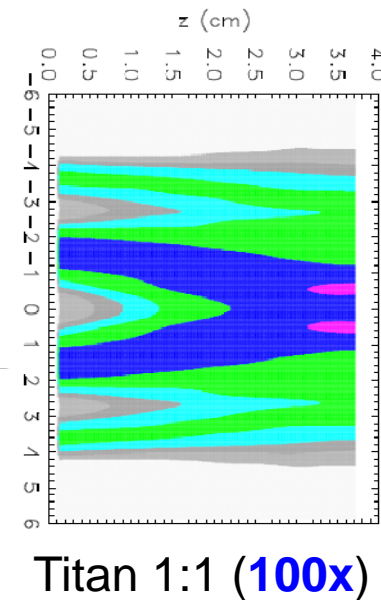
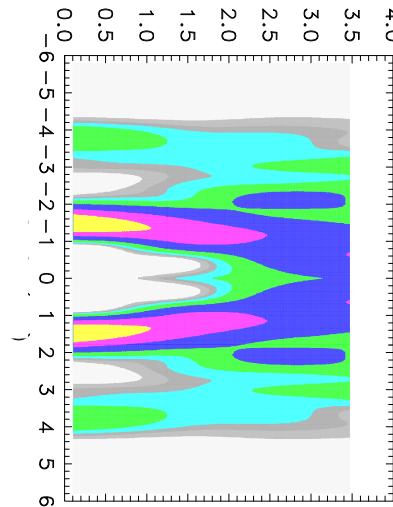
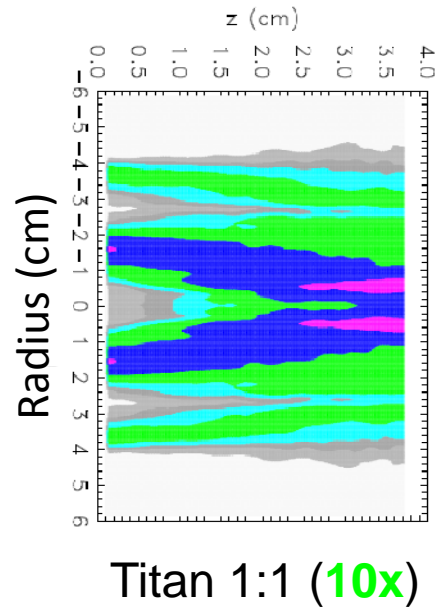
**AASC 1:2**

Fiber optic interferometry  
(4 axial x 20 radial)

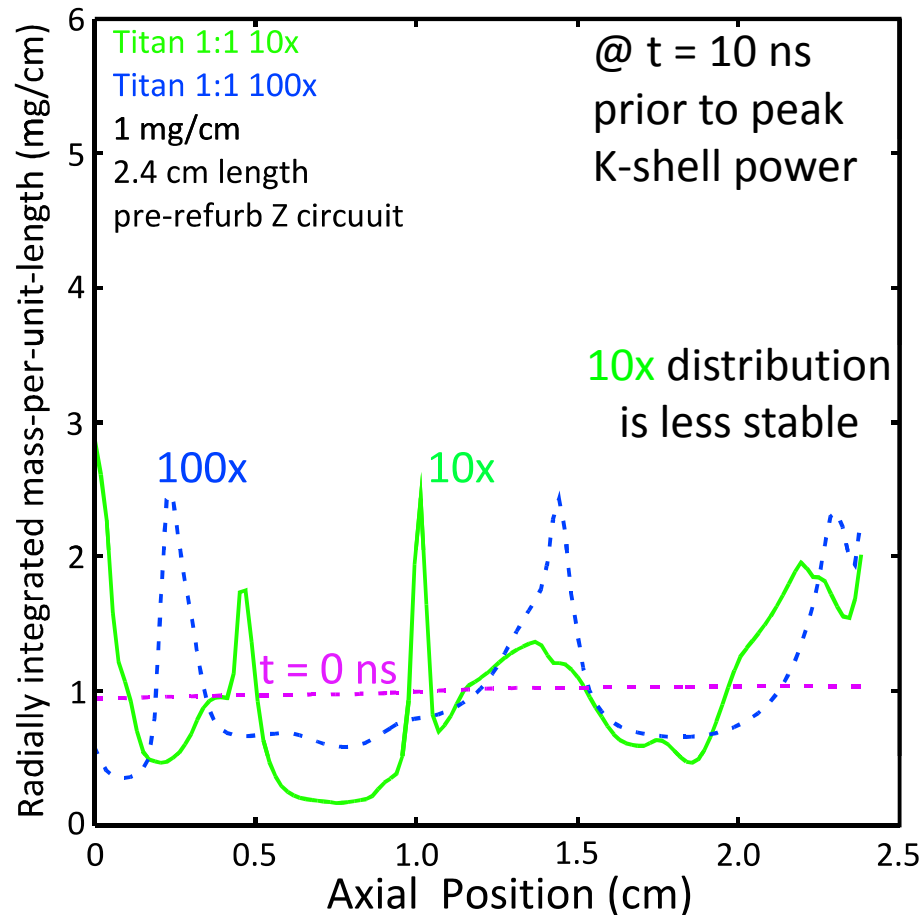
The **AASC** gas distribution is inherently “smoothed” due to bicubic spline interpolation of data

# Comparison of the Titan 1234 and AASC gas distribution's K-shell and stability properties

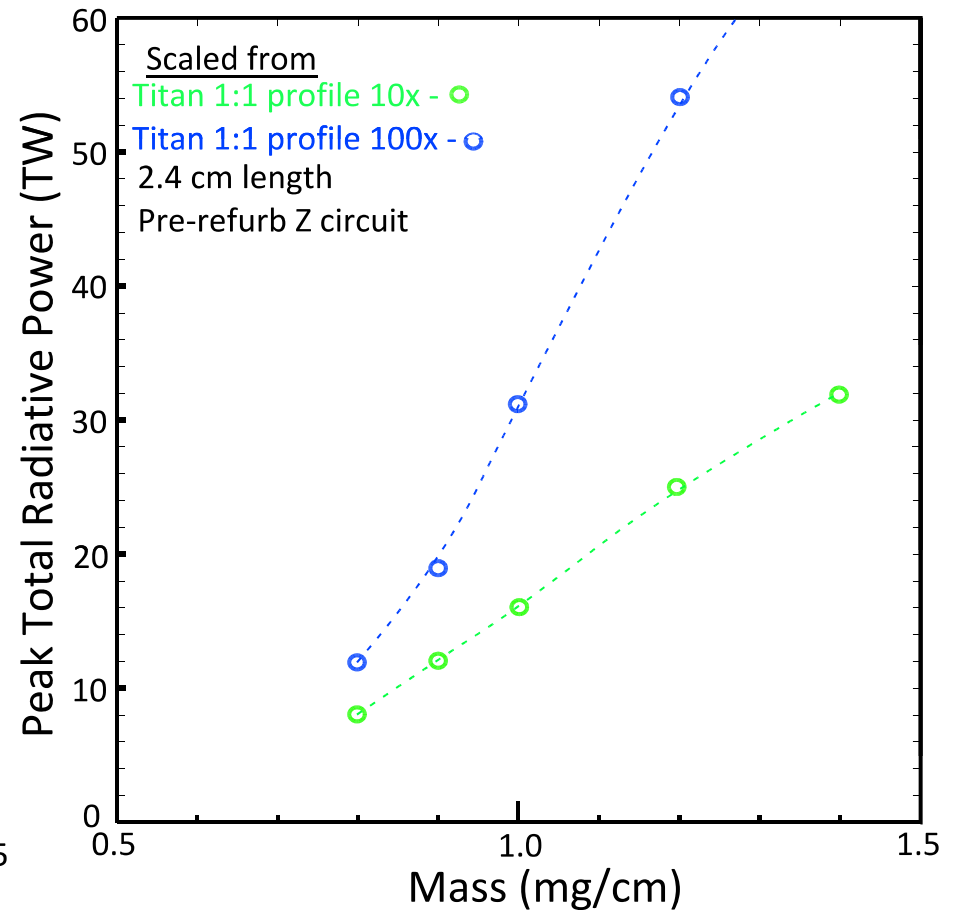
To make for a “fair” comparison with the inherently “smoothed” AASC distributions, the Titan gas distribution is “smoothed” by averaging over nearest neighbor data points successive times. Calculations are performed using a **10x** and **100x** Titan gas distribution.



Greater stability of Titan **100x** distribution gives rise to higher radiative powers than **10x** distribution

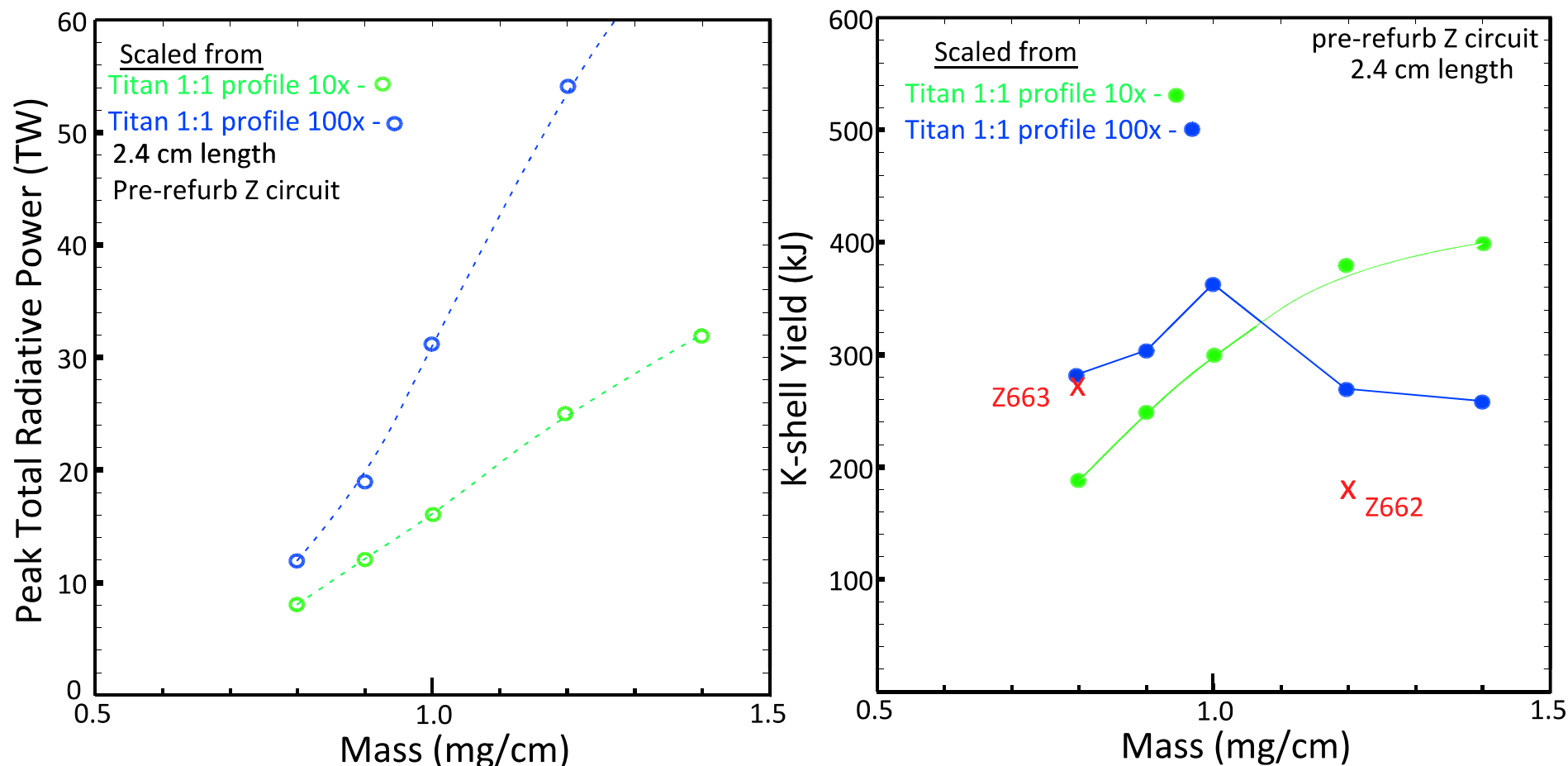


At each axial location, the mass is radially integrated. Large fluctuations denote an unstable implosion.



Lower powers of **10 x** distribution are due to increased axial non-uniformity

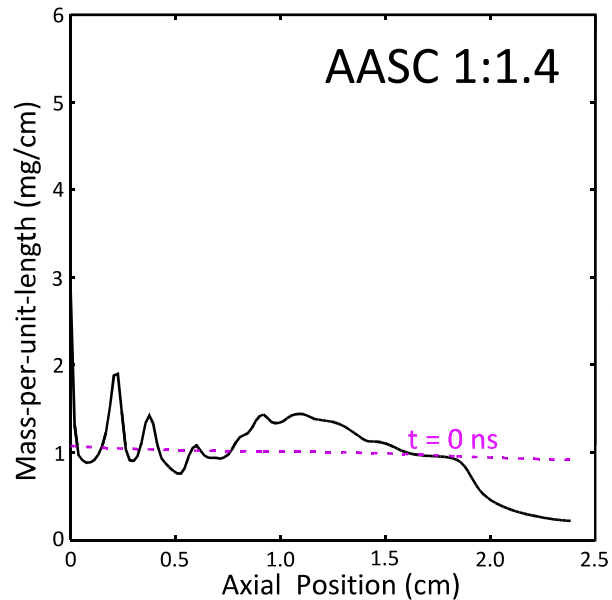
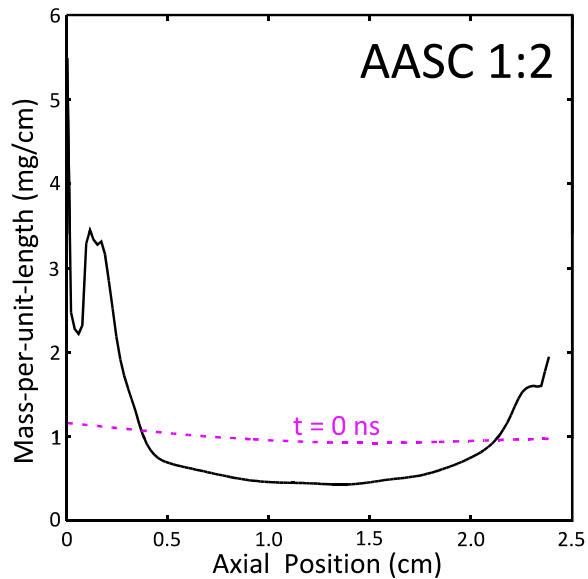
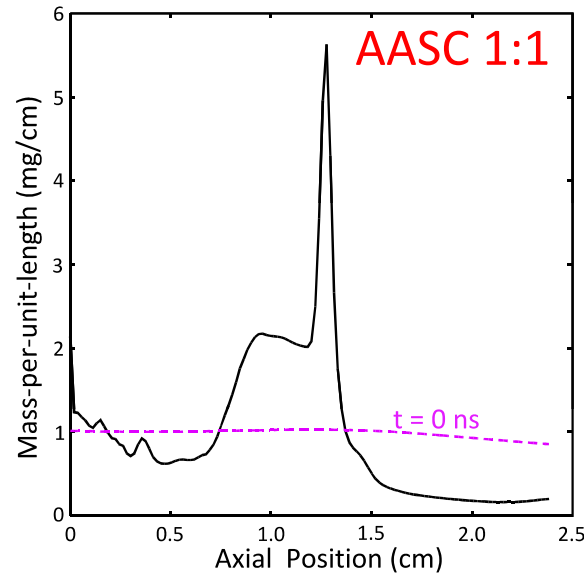
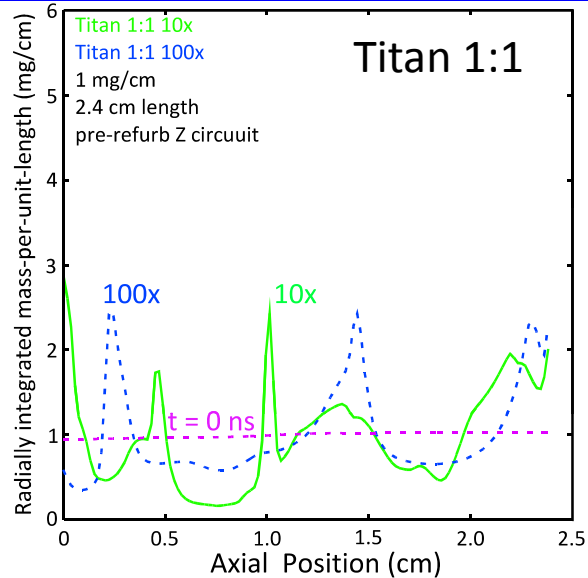
## 2D Calculated K-yields and total radiative powers for Titan 1234 10x and 100x smoothed gas distribution :Pre-refurbished Z circuit



**K-shell yield curve turns over when total power is sufficiently large that the plasma cools too rapidly to sustain K-shell emission**

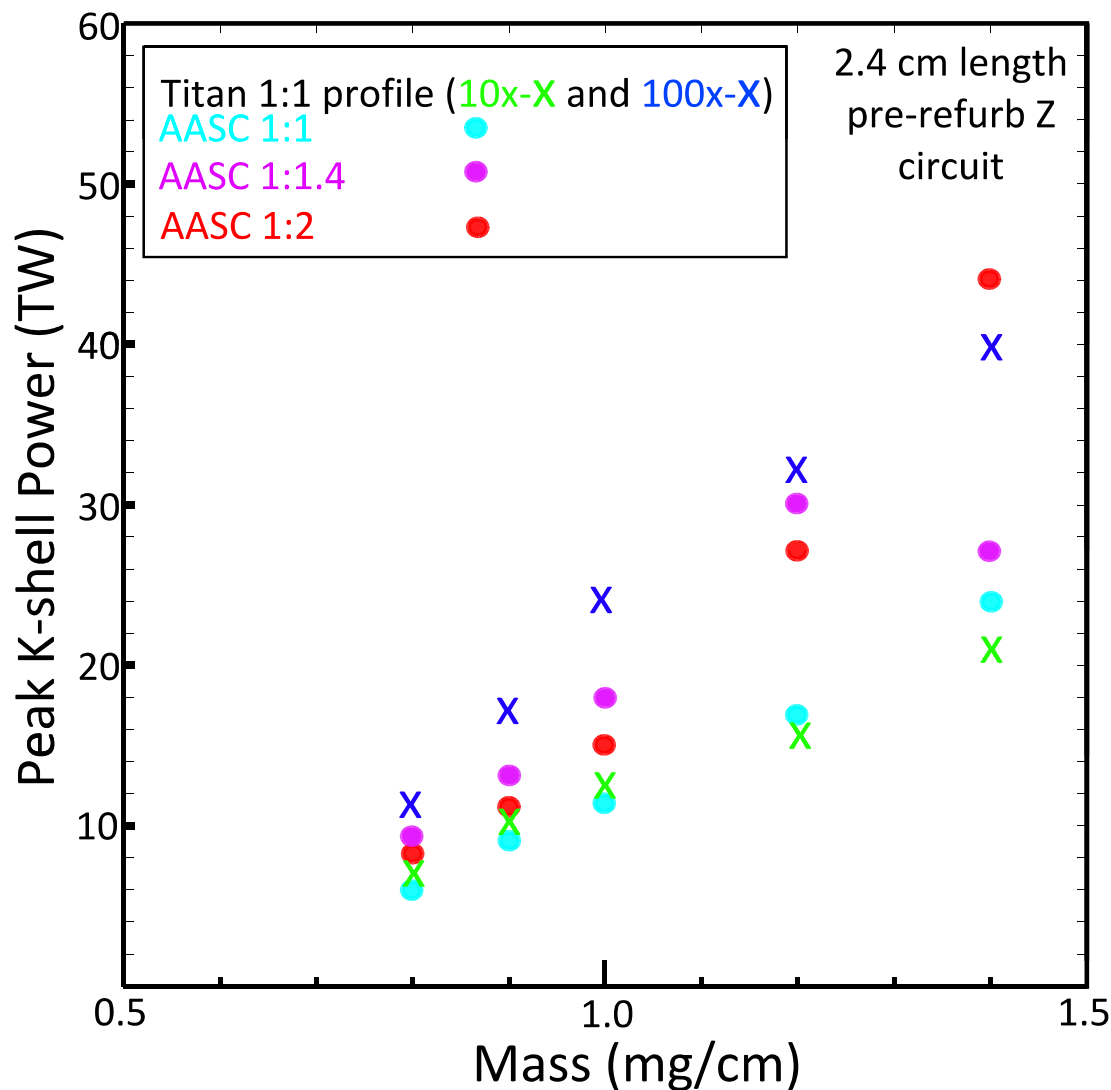
- 100 x profile best matches pre-refurbished Z experiments.
- 10 x profile has higher yield at large mass because of less radiative cooling.

## The AASC 1:1 gas distribution is the least stable



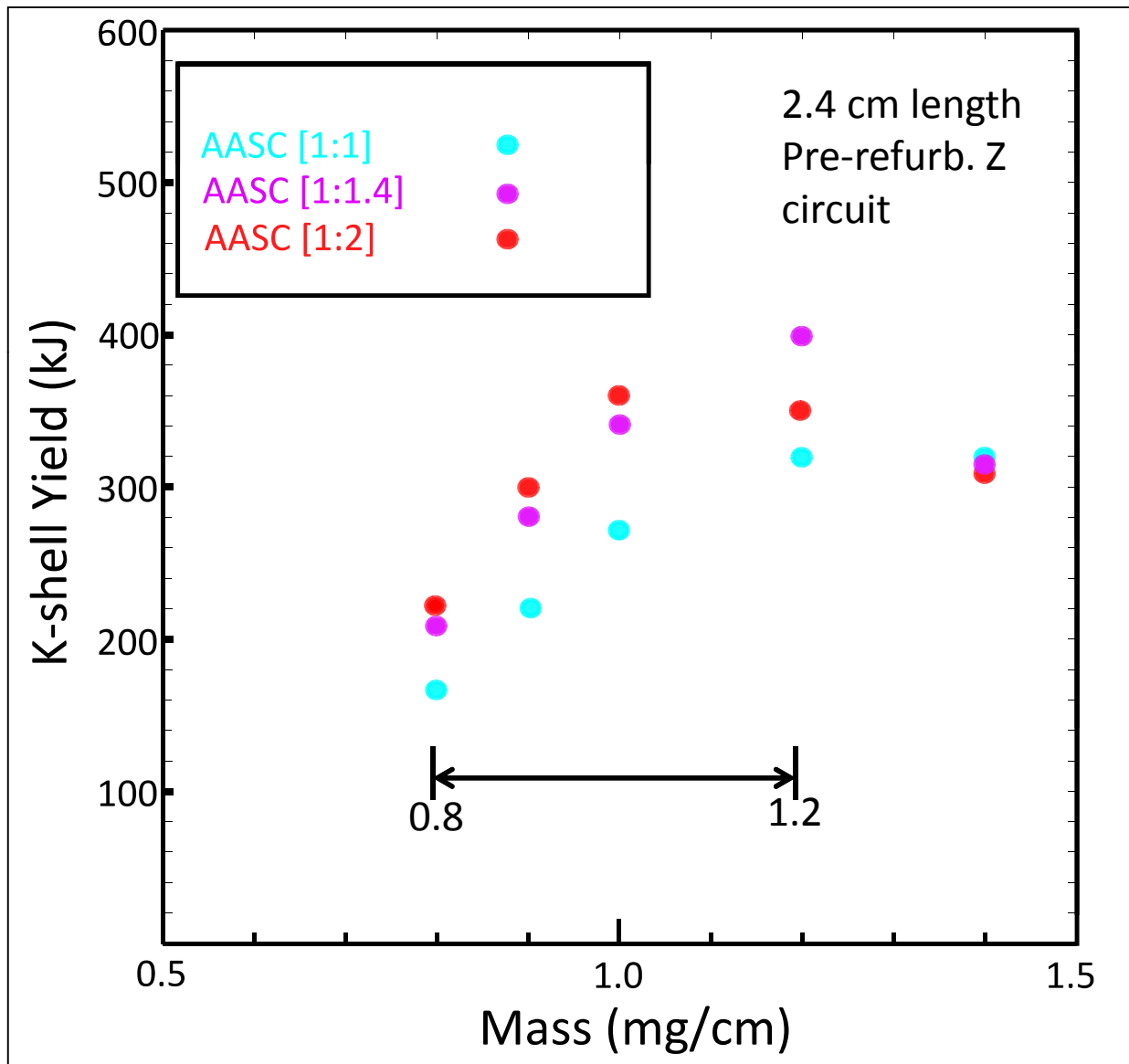
The radially integrated mass-per-unit-length as a function of axial position – 10 ns prior to stagnation. Large fluctuations denote an unstable implosion. These calculations were for a 2.4 cm length, 1 mg/cm load driven by the pre-refurbished Z circuit.

## 2D Calculated K-shell Power for Titan 1234 and AASC Nozzles on Pre-refurbished Z



The Titan **100x**, and AASC **1:1.4** and **1:2** distributions are the most stable and they have the highest K-shell and total powers.

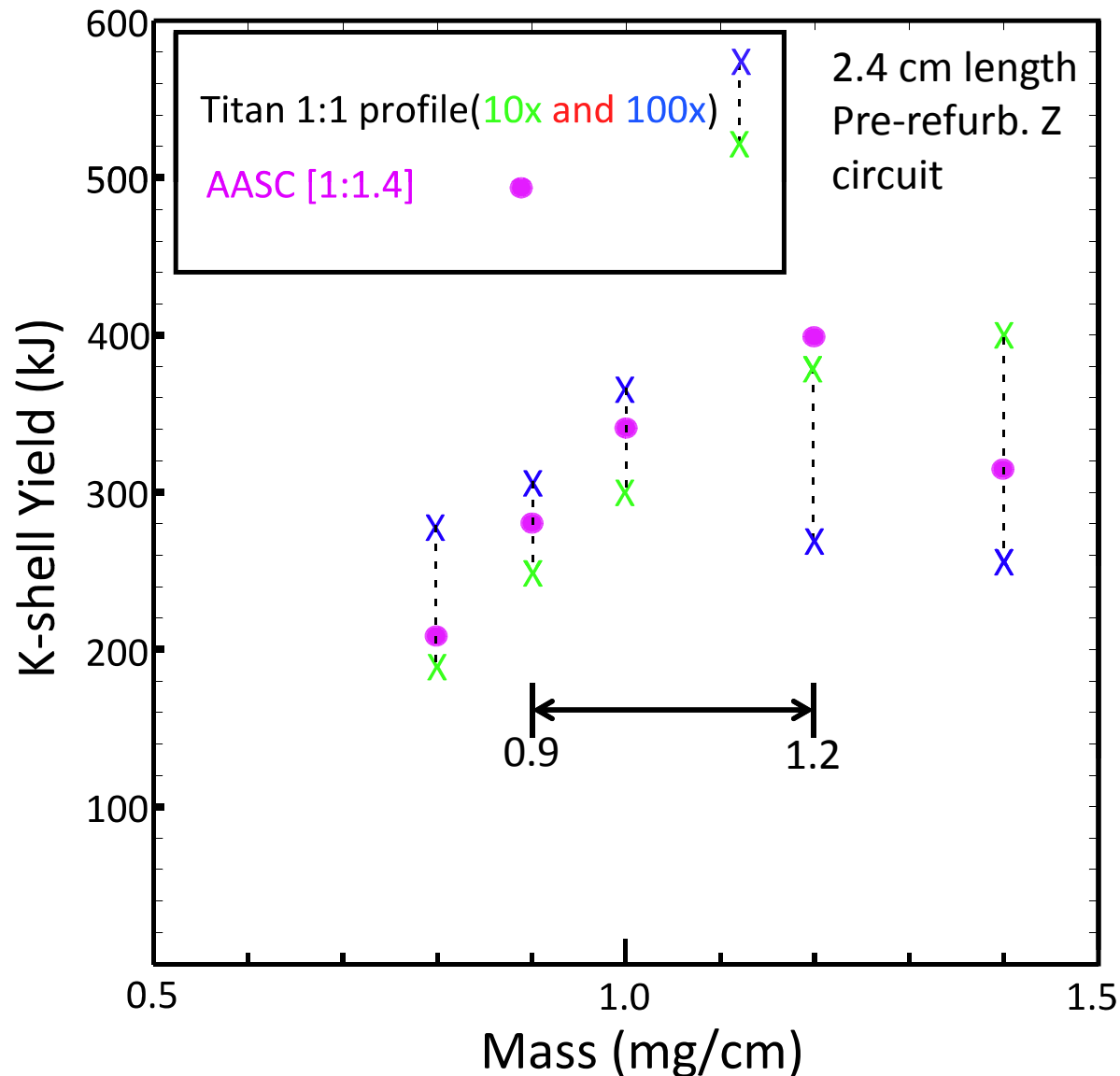
## 2D Calculated K-yields for Titan 1234 and AASC Nozzles on Pre-refurbished Z



The **AASC 1:1.4** and **1:2** loads outperform the **1:1** load – especially in the range **0.8 – 1.2 mg/cm**, where ZR experiments will likely take place.



## 2D Calculated K-yields for Titan 1234 and AASC Nozzles on Pre-refurbished Z



The **AASC 1:1.4** and **1.2** loads perform as well as the Titan load in the range **0.9 – 1.2 mg/cm**.

The issue of severe current losses due to UV irradiance of the convolute power feed is still under investigation.

Stability properties of the AASC 1:1 profile were the worst, resulting in lower calculated K-shell yields and powers than obtained with the AASC 1:1.4, 1:2, and Titan distributions.

The “True” AASC gas distributions are probably not as stable as the modeled bicubic spline Interpolated (“smoothed”) AASC gas distributions analyzed here – would lead to lower calculated yields than shown in this analysis.

Modeled stability of the AASC 1:1.4 and 1.2 distributions was as good as that of the modeled Titan distribution – instabilities no more of a limiting factor of K-yield than they were in the pre-refurbished Z experiments with the Titan nozzle. Therefore, expect  $I^2$  K-shell scaling to be valid for AASC 1:1.4 and AASC 1:2 loads (~ 370 kJ of K-shell emission for ~ 1 mg/cm load).