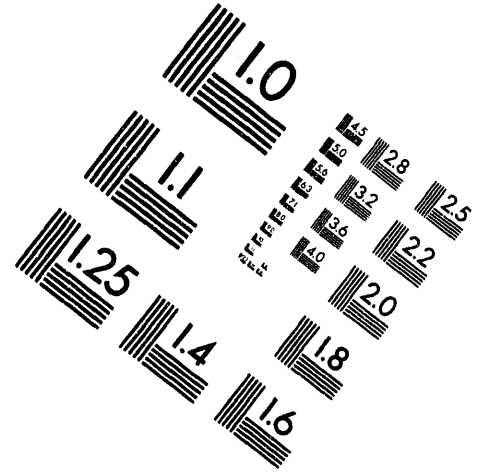
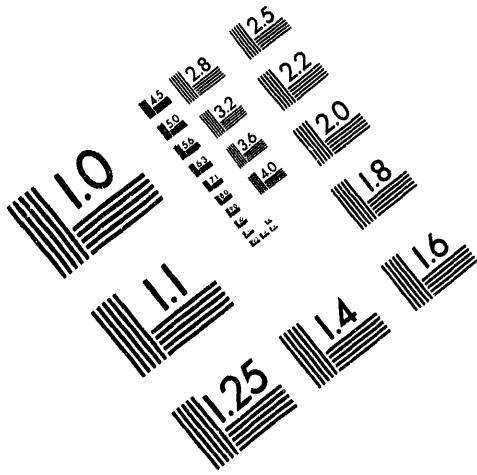




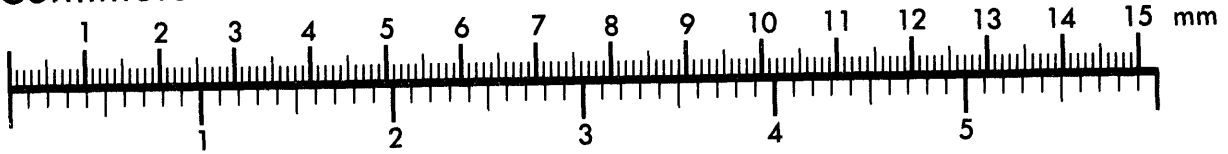
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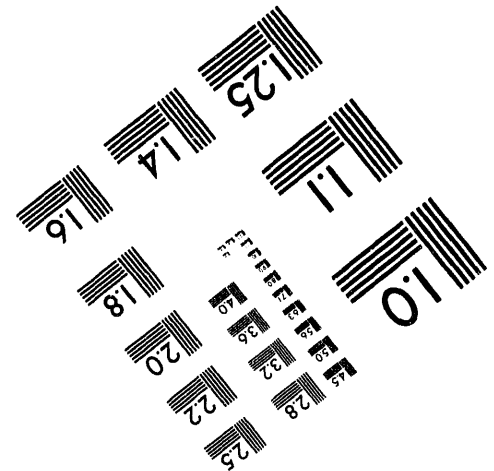
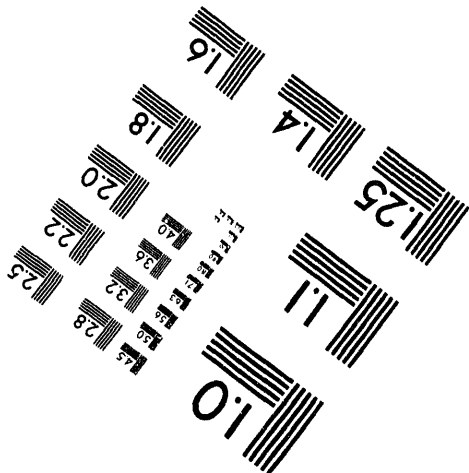
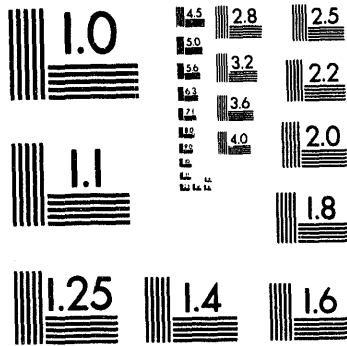
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## ICRF HEATING OF TFTR DEUTERIUM SUPERSHOT PLASMAS IN THE $^3\text{He}$ MINORITY REGIME<sup>+</sup>

G. TAYLOR, J.R. WILSON, R.C. GOLDFINGER\*\*, J.C. HOSEA, D.J. HOFFMAN\*\*,  
R. MAJESKI, C.K. PHILLIPS, D.A. RASMUSSEN\*\*, J.H. ROGERS, G. SCHILLING,  
J.E. STEVENS, M.G. BELL, R.V. BUDNY, C.E. BUSH\*\*, Z. CHANG\*, D. DARROW,  
D.R. ERNST\*\*\*, E. FREDRICKSON, G. HAMMETT, K. HILL, A. JANOS, D. JASSBY,  
D.W. JOHNSON, L.C. JOHNSON, S.S. MEDLEY, H.K. PARK, J. SCHIVELL,  
J.D. STRACHAN, E. SYNAKOWSKI, and S. ZWEBEN

*Princeton Plasma Physics Laboratory, P.O. Box 451, Princeton, New Jersey 08543, USA*

**Abstract --** The increased core electron temperature produced by ICRF heating of TFTR, D-T neutral-beam-heated supershot plasmas is expected to extend the alpha particle slowing down time and hence enhance the central alpha particle pressure. In preparation for the TFTR D-T operational phase, which is due to start in late 1993, a series of experiments were conducted on TFTR to explore the effect of ICRF heating on the performance and stability of low recycling, deuterium supershot plasmas in the  $^3\text{He}$  minority heating regime. The coupling of up to 7.4 MW of 47 MHz ICRF power to full size ( $R \sim 2.62$  m,  $a \sim 0.96$  m),  $^3\text{He}$  minority, deuterium supershots heated with up to 30 MW of deuterium neutral beam injection has resulted in a significant increase in core electron temperature ( $\Delta T_e = 3-4$  keV). Simulations of equivalent D-T supershots predict that such ICRF heating should result in approximately a 60% increase in the alpha particle slowing down time and an enhancement of about 30% in the central alpha pressure. Future experiments to be conducted at ICRF powers up to 12.5 MW during the upcoming TFTR D-T campaign may result in even greater enhancements in core alpha parameters. This paper presents results from experiments performed at an axial toroidal magnetic field of  $\sim 4.8$  T, where the minority resonance was within 0.1-0.15 m of the plasma core. Combined ICRF and neutral beam heating powers in these experiments reached TFTR record levels of over 37 MW, which allowed an exploration of the power loading limits on the carbon limiter tiles. The plasma current was operated at 1.85 and 2.2 MA and sawtooth suppression was observed at the higher plasma current.

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\*University of Wisconsin, Madison, Wisconsin 53706, USA

\*\*Oak Ridge National Laboratory, Oak Ridge, Tennessee 08731, USA

\*\*\* Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

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discussed in Sec. III together with results from a D-T plasma simulation. A summary of our conclusions and plans for future work in this plasma regime are presented in Sec. IV.

## II. EXPERIMENT AND RESULTS

### A. DISCHARGE OPTIMIZATION

Figure 1 shows a poloidal cross-section through the TFTR vacuum vessel and indicates schematically the position of the ICRH antennas and limiter positions. Since the inner limiter has considerably more power handling capability than the outer limiters (SEVIER *et al.*, 1983; PELESSONE *et al.*, 1983) the plasmas in these experiments were run so that the majority of the power flow was to the inboard bumper limiter. However, to simultaneously bring the outer edge of the plasma close to the four ICRH antennas on the outboard side it was necessary to develop a "full bore" ( $R \sim 2.62\text{m}$ ,  $a \sim 0.96\text{m}$ ) supershot plasma. Prior to the 1991-92 TFTR run period, supershot plasma performance was optimized at  $R \sim 2.45\text{m}$ ,  $a \sim 0.83\text{m}$  but with the installation of an upgraded RF limiter in 1991 the supershot performance at the larger major radius was significantly improved (HAWRYLUK *et al.*, 1992).

Supershots are characterized by low edge recycling with all the fueling provided by the beam ions and hence they have a low edge density. However, NBI provides a relatively significant increase in edge density, so that after 100-200 ms sufficient density had built-up in front of the antennas to allow efficient coupling of ICRH power (MAJESKI *et al.*, 1993). The evolution of the electron density at the last closed flux surface on the outboard side ( $R \sim 3.6\text{m}$ ) from Abel-inverted multi-channel far-infrared interferometer data (MANSFIELD *et al.*, 1987; PARK, 1989) is shown in Fig. 2. The rapid rise in edge density throughout the NBI pulse from 3 to 4 s (shown lightly shaded in Fig. 2) is particularly challenging for ICRH in this regime since it can result in a factor of 5 to 6 increase in antenna loading impedance. The start of the ICRH pulse (shown by the darker shading in Fig. 2) was delayed  $\sim 200$  ms from the beginning of NBI to allow time for some NBI fueling to occur and the ICRH pulse shape was

longer interacts with the bottom of the RF limiters. Since this in turn increases the connection length of the outer flux surface it will result in increased power loading at the top of the RF limiters. Furthermore, since the scrape-off length increases as the square root of the connection length some of this power may fall behind the front surface, on the edges of the RF limiters. This effect can increase the power loading significantly since the power falling on the limiter edge is no longer at near-tangential incidence.

### C. SUPERSHOT PLASMA PERFORMANCE WITH ICRH

Experiments were performed at plasma currents of 1.85 and 2.2 MA. 22-23 MW of NBI was typically employed in the 1.85 MA plasmas and 28-30 MW of NBI was used in the 2.2 MA plasmas. The plasma current and NBI power were chosen to stay well below the empirical TFTR  $\beta$  limit,  $\beta_n$  was typically  $<1.5$  for these discharges. Up to 5.7 MW and 7.4 MW of 47 MHz ICRH power has been coupled into R~2.62 m, a~0.96 m deuterium discharges at a plasma current of 1.85 MA and 2.2 MA, respectively. The plasmas had an axial toroidal magnetic field of ~4.8 T. The Shafranov shift on axis was 0.2-0.25 m and 0.16-0.2 m for the 1.8 MA and 2.2 MA plasmas, respectively. The  $^3\text{He}$  resonance was within 0.1-0.15 m of the plasma core.

Figure 6 shows a comparison between two plasmas heated with 23 MW of NBI between 3 and 4 seconds. Approximately 60% of the NBI power was in the direction of the plasma current in both cases. One of these discharges (solid line) had 5.7 MW of ICRH coupled into it between 3.2 and 4 seconds, the other plasma (dashed line) had no ICRH power but both plasmas had a similar size  $^3\text{He}$  puff at 2.5 s. The central electron temperature measured by electron cyclotron emission (ECE) (STAUFFER *et al.*, 1985) increased by 40% [Fig.6(b)] and the diamagnetically measured global stored energy increased by 20% [Fig.6(c)] with the addition of ICRH. With the application of ICRH the core ion temperature at 3.7 seconds measured by charge exchange recombination spectroscopy (FONCK *et al.*, 1989) increased from 20 to 24 keV. The confinement time [Fig.6(f)] and central electron density evolution

several years ago. This empirical observation remains an unresolved issue and is the subject of an ongoing study (TAYLOR *et al.*, 1993b). The change in central ion temperature [Fig. 7(b)] was of the same order as the  $\pm 10\%$  measurement errors of the charge exchange recombination spectroscopy diagnostic. There was an  $\sim 4\%$  drop in the core electron density (Fig. 7(c)) and  $\sim 10\%$  drop in density profile peakedness ( $n_e(0)/\langle n_e \rangle$ ) determined from far-infrared interferometry data. The peakedness degradation was due in part to an increase in edge density, this was correlated with a small increase in  $Z_{\text{eff}}$  measured by visible bremsstrahlung (RAMSEY and TURNER, 1987) from approximately 3.1 to 3.4 with the addition of 5.7 MW of ICRH. The decreased core electron density could result from reduced beam penetration due to the higher edge density and/or  $Z_{\text{eff}}$ .  $\tau_E$  decreased 5-10% as the ICRH power was increased [Fig. 7(d)], although this is comparable to the degradation expected for Goldston empirical L-mode scaling (GOLDSTON *et al.*, 1984) it is also of the same order as the uncertainty in the  $\tau_E$  measurement. The relatively large scatter in  $\tau_E$  with no ICRH is due to variations in wall conditioning. The enhancement over the Goldston empirical L-mode was typically 1.7-2.3 for the 1.85 MA dataset. The measured plasma reactivity [Fig. 7(e)] remained unchanged with the addition of ICRH. Figure 7(f) shows the projected core alpha energy slowing down time calculated for D-T equivalent plasmas, with the same central density and temperature measured during the D-D experiments. A 60% enhancement in the core alpha particle energy slowing down time was calculated for equivalent D-T plasmas at the highest ICRH powers.

Operation at 2.2 MA was attempted at higher NBI power ( $\sim 28-30$  MW) and higher ICRH power (up to 7.4 MW). At these higher input powers, plasma performance was relatively poor and unreliable as has already been indicated in the previous section on power handling limits. Global energy confinement times were 15-20% lower in the 2.2 MA plasmas compared to the 1.85 MA discharges. The enhancement over the Goldston empirical L-mode scaling was only 1.2-1.7 for the 2.2 MA plasmas. Figure 8 provides an overview of the general features of these discharges; the evolution of major plasma parameters for two 2.2 MA plasmas is shown. Discharges that could be directly compared with and without ICRF were difficult to obtain due to variations in plasma conditioning. The two shots of Fig. 8 are generally similar and have the

by an ECE Grating polychromator (CAVALLO *et al.*, 1988). Measurements of MeV ion loss (ZWEBEN *et al.*, 1990) were also made for these discharges. Escaping MeV ions were measured by detectors at  $90^\circ$ ,  $60^\circ$  and  $45^\circ$  below the outer midplane just outside the limiter radius. The results are shown in Fig.11, where the time at which the ICRH power first steps up is indicated by the vertical dashed line. Although the evolution of the D-D neutron rates for these two plasmas are very similar [Fig.11(b)] the fast ion loss for all three detector locations is considerably enhanced for the plasma with ICRH (solid line) and in particular the signal at the  $45^\circ$  detector during ICRH is approximately five times that at the  $90^\circ$  detector [Fig.11(e)]. These observations are qualitatively similar to earlier MeV ion loss measurements during  $^3\text{He}$  minority ICRH experiments (ZWEBEN *et al.*, 1992). The time evolution of the signal at all three detectors was examined for evidence of a 10 kHz modulation coincident with the  $m=1$ ,  $n=1$  MHD mode seen on the ECE signal, but none was seen. It would appear therefore that the enhancement in fast ion losses is not due predominantly to the presence of the MHD mode. The enhancement in the fast ion loss at the  $90^\circ$  detector for the shot with ICRH occurs coincident with the turn-on of the ICRH consistent with the first-orbit loss of D- $^3\text{He}$  fusion product alpha particles, however the enhancement appears to be delayed by 50-100 ms at the  $45^\circ$  detector, consistent with either ICRH-induced loss of previously confined D-D or D- $^3\text{He}$  fusion products, or direct ICRH tail ion losses. Figure 12 shows the dependence of the measured MeV ion loss on the coupled ICRH power for the plasmas shown in Fig. 7. The signal at the  $90^\circ$  detector rises relatively little with increasing ICRH power, but the signal at the  $45^\circ$  detector increases rapidly above 2-3 MW. The  $45^\circ$  detector signal with 6MW of ICRH is about an order of magnitude greater than the D-D fusion product signal seen with NBI alone and is probably only due to  $<1\%$  of the ICRH tail ion population being lost to the walls. For the same ICRH power scan D-D neutron collimator measurements (JOHNSON, 1992) indicate no apparent broadening of the neutron emissivity profile with increasing ICRH power, consistent with relatively little RF-induced beam ion spreading.

0.8-1MeV with 5.7 MW of ICRH. In the SNAP computer model the  $^3\text{He}$  is assumed to have a flat radial concentration profile so that its profile shape is similar to the electron density profile, however recent charge exchange recombination spectroscopy data analysis of  $^4\text{He}$  transport in supershots (SYNAKOWSKI *et al.*, 1993) suggest a more peaked  $^3\text{He}$  concentration profile may exist in these discharges. In some (~10%) of shots analyzed by SNAP with a 2%  $^3\text{He}/n_e$  concentration, there was insufficient measured electron density to satisfy quasi-neutrality in the calculation. This problem was obviated for most of these shots by reducing the  $^3\text{He}/n_e$  concentration in the model to 1.5%. This reduced  $^3\text{He}/n_e$  concentration had little effect on the calculated stored energy and increased the predicted neutron production rate by only 1-2%. The anisotropy in the diamagnetically measured stored energy as a function of ICRH power is shown in Fig. 14(b). The energy distribution has a residual parallel component due to the beam ion population which is ~600 kJ without ICRH. As the ICRH power is increased to ~6 MW the parallel anisotropy drops to ~200 kJ. If this behavior is interpreted as being due entirely to the presence of the perpendicular energy in the RF tail then it implies a tail stored energy, which is consistent with the model results. (The uncertainty in the diamagnetic determination of the tail stored energy is typically 50-100 kJ.) Figure 14(c) shows the calculated fraction of RF power to the electrons as a function of applied RF power. At 5.7 MW just under 50% of the RF power is calculated to go to electron heating. As the  $^3\text{He}$  tail slows down it increasingly damps on the ion population near the plasma core, this process accounts for a further ~30% of the power coupled to the plasma at 5.7MW. The code indicates that the remainder of the RF power (~20%) is deposited along the fundamental deuterium and carbon resonance layer on the small major radius side of the discharge, with a distribution which moderately peaks where these resonances intersect the shear Alfvén mode conversion layer. At lower ICRH powers ( $\leq 2$  MW) this damping mechanism is even more significant and the calculated ICRH power fraction to the fundamental deuterium and carbon resonance layer rises to ~ 60%. Since the region where these layers intersect is comparable to the poloidal resolution limits in the code, there is some concern that large errors in the predicted RF damping may result. Furthermore, the accuracy of the reduced order approximation which forms the basis for the RF model in the

#### IV. SUMMARY AND CONCLUSIONS

Coupling of up to 7.4 MW of ICRH to deuterium supershot plasmas in the  $^3\text{He}$  minority regime has resulted in significant core electron heating ( $\Delta T_e \sim 40\%$ ). The enhancement in projected core alpha slowing down time for the 1.85 MA ICRH power scan appears linear with increasing ICRH power [Fig.7(f)] up to the ICRH power levels explored so far ( $\sim 5.7$  MW). The projected alpha particle slowing down time was enhanced  $\sim 60\%$  as a result of the increased core electron temperature. A time dependent D-T simulation indicated the need for longer heating pulses ( $\sim 1.4$ s) than were used in these experiments, this would provide sufficient time for the increased alpha slowing down time to result in enhanced alpha pressure.

It was apparent that the plasma performance was being degraded as the ICRH power was increased, largely as a result of increased edge density and  $Z_{\text{eff}}$ . It may be possible to further improve performance with additional plasma conditioning. Operation at 2.2 MA, at higher total heating powers ( $\sim 37$  MW) proved problematic; plasma performance was found to be critically dependent on the vertical position of the plasma, however relatively low radiated power fraction ( $< 15\%$ ) and carbon in-flux was obtained with up to 35 MW of heating power.

Calculated RF tail energies were found to be consistent with diamagnetic measurements of the residual anisotropy in the energy distribution. RF tail temperatures were calculated to be 0.8-1 MeV near the magnetic axis. There was an order of magnitude enhancement in MeV ion losses with the addition of 5.7 MW of ICRH, although this represents  $< 1\%$  of the tail ion population there is some concern that with increased ICRH power fast ion losses could become significant. Although the fraction of RF power calculated to go to the core electron population via  $^3\text{He}$  tail damping and direct electron heating increased to  $\sim 45\%$  as the RF power was increased there appears to be a saturation in the fraction of RF power to the electrons above  $\sim 4$  MW. Also, modeling indicates that  $\sim 20\%$  of the RF power is deposited on the high field side, along the fundamental deuterium and carbon resonance. Because of concerns regarding

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## FIGURE CAPTIONS

**Fig.1** Schematic diagram showing a poloidal cross-section of the TFTR vacuum vessel.

**Fig.2** Evolution of the density at the outer closed flux surface. NBI from 3-4 s (lighter shading) results in a rapid rise in edge density. The ICRH pulse (darker shading) was delayed from the start of NBI to allow time for density to build in from of the antennas. Also a modest step and ramp were programmed into the start of the ICRH pulse in order to avoid antenna arcs.

**Fig.3** Comparison of the evolution of (a) the diamagnetically determined stored energy and (b) neutron reactivity for two similar plasmas heated with 23 MW of NBI. One plasma (thicker line) had a 1 torr.liter/s, 100 ms  $^3\text{He}$  puff at 2.5 s, the other plasma had no  $^3\text{He}$  puff. The helium-3 puff was observed to degrade the stored energy by ~5% and the neutron reactivity by ~10%.

**Fig.4** Radiated power fraction versus total heating power for  $^3\text{He}$  minority, deuterium supershots at 1.85 and 2.2 MA. Data are at the time of maximum stored energy. For most plasmas the radiated power fraction was 10-15%.

**Fig.5** Evolution of plasmas parameters for two plasmas, one centered vertically (thin line) the other positioned ~2 mm higher (thicker line). Both plasmas had ~36 MW of input heating power, with similar NBI and ICRH power waveforms. The time dependence of (a) the NBI power, (b) the ICRH power, (c) the plasma stored energy, (d) the total radiated power, (e) the C-II light and (f) the  $D_\alpha$  light is shown. A significant increase in radiation and a carbon "bloom" is seen for the plasma with the 2mm vertical shift.

**Fig.9** Plasma performance during an ICRH power scan of 2.2 MA D-D supershots with 28-30 MW of NBI. Data were obtained at the time of maximum stored energy. (a) the central electron temperature, (b) the neutron production rate, (c) the global energy confinement time, and (d) the projected alpha energy slowing down time for an equivalent D-T plasma are plotted versus coupled ICRH power. A linear fit has been made to the data for each plot, the error bars represent typical uncertainties on the data points. Although the observed trends are similar to the 1.85 MA data there is considerable variability in plasma performance. (The open square and circle indicate the shots used for the discharge evolution comparison of Fig. 8.)

**Fig.10** The amplitude of Mirnov coil MHD fluctuation data is shown for the two plasmas in Fig. 6. One plasma (solid line) had 5.7 MW of ICRH coupled into it at the time indicated the other (dashed line) had no ICRH. Data are shown for (a) the  $m=2$  monitor signal and the amplitude of fluctuations with frequencies in the range (b) 5-15 kHz, (c) 15-25 kHz, (d) 25-35 kHz, (e) 35-45 kHz, (f) 45-55 kHz, (g) 55-65 kHz and (h) 65-75 kHz.

**Fig.11** Time evolution of the MeV ion losses for the two plasmas shown in Fig.6. (a) The timing of the ICRH and NBI pulses, (b) the D-D fusion rate, (c) the  $90^\circ$  fast ion loss, (d) the  $60^\circ$  fast ion loss and (e) the  $45^\circ$  fast ion loss. The plasma with ICRH coupled is indicated by the solid line. The neutron production rate is similar for the two discharges however the fast ion losses are enhanced significantly in the plasma with ICRH. The  $45^\circ$  detector signal shows a delayed turn on relative to the initial step in ICRH power (vertical dashed line).

**Fig.12** Dependence of the relative fast ion loss versus coupled ICRH power for the 1.85 MA ICRH power scan of Fig.7.

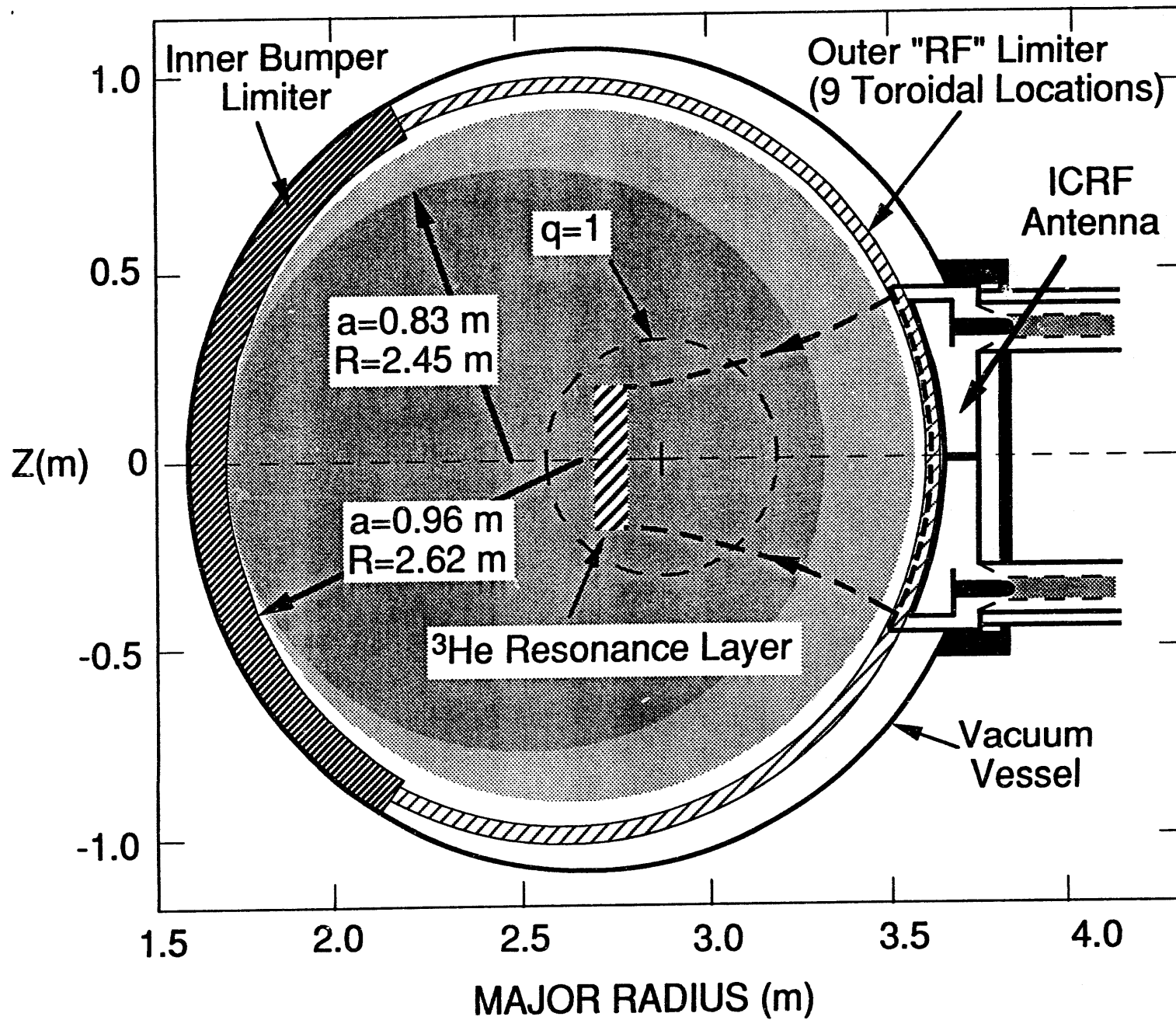


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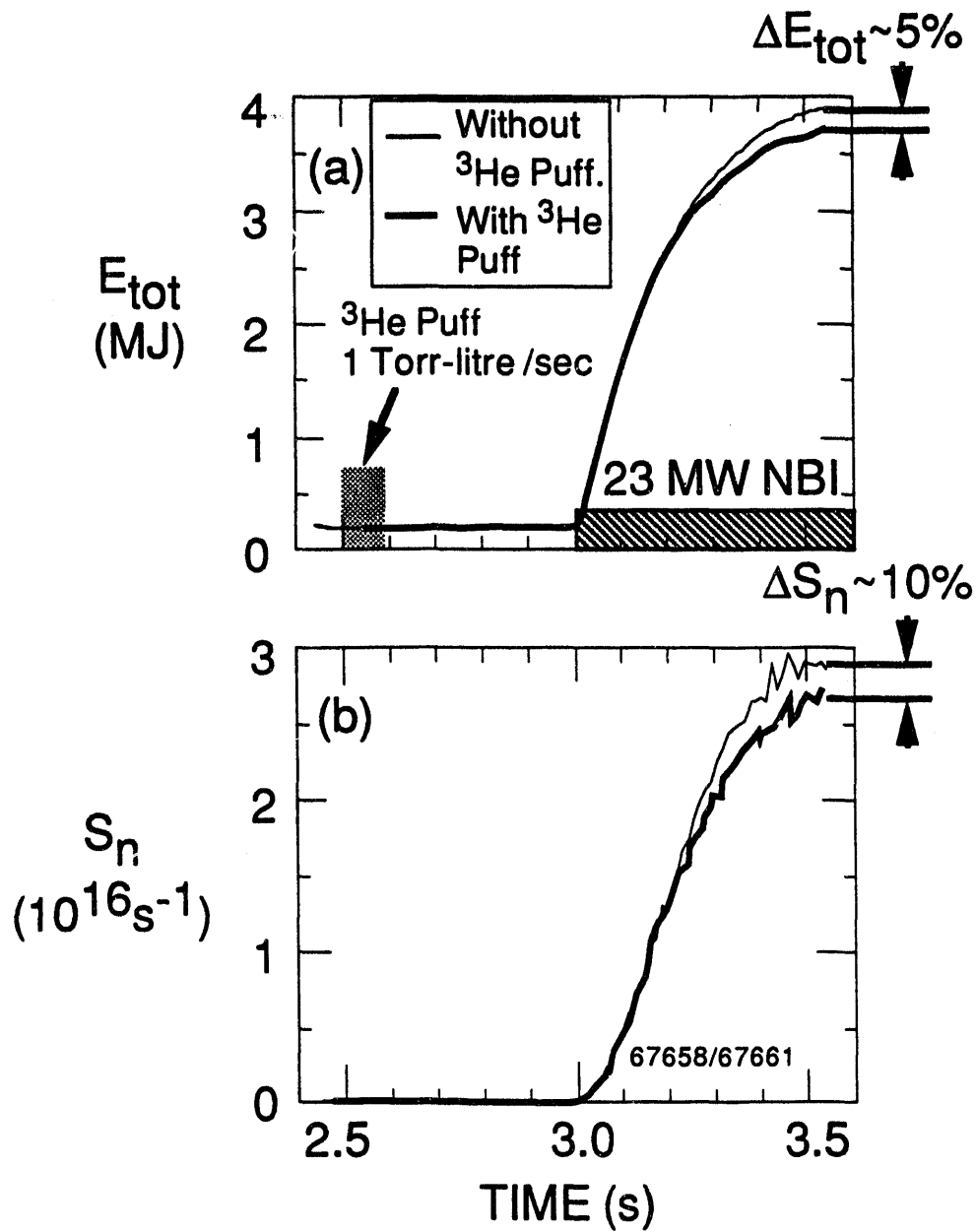


Fig. 3

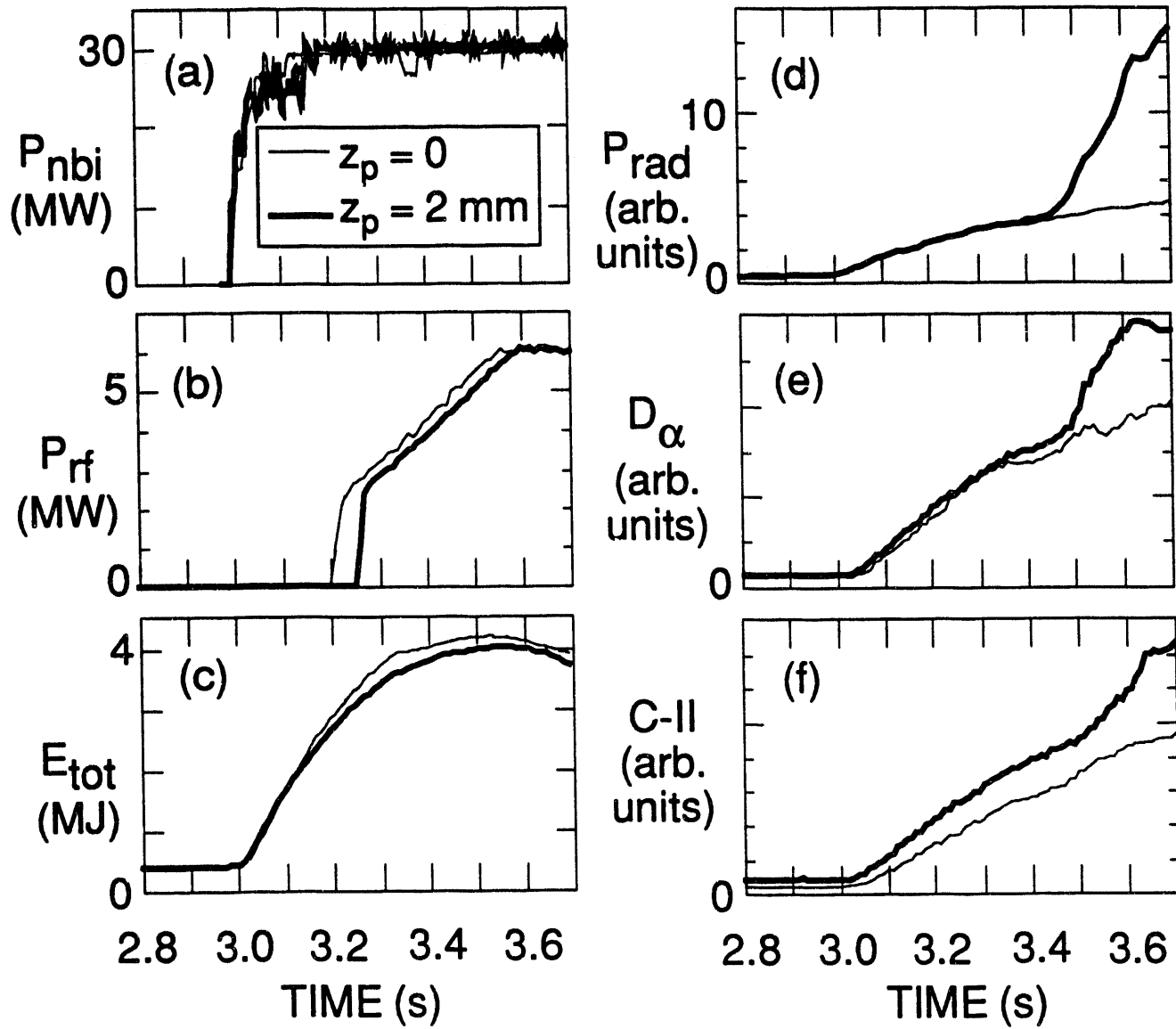
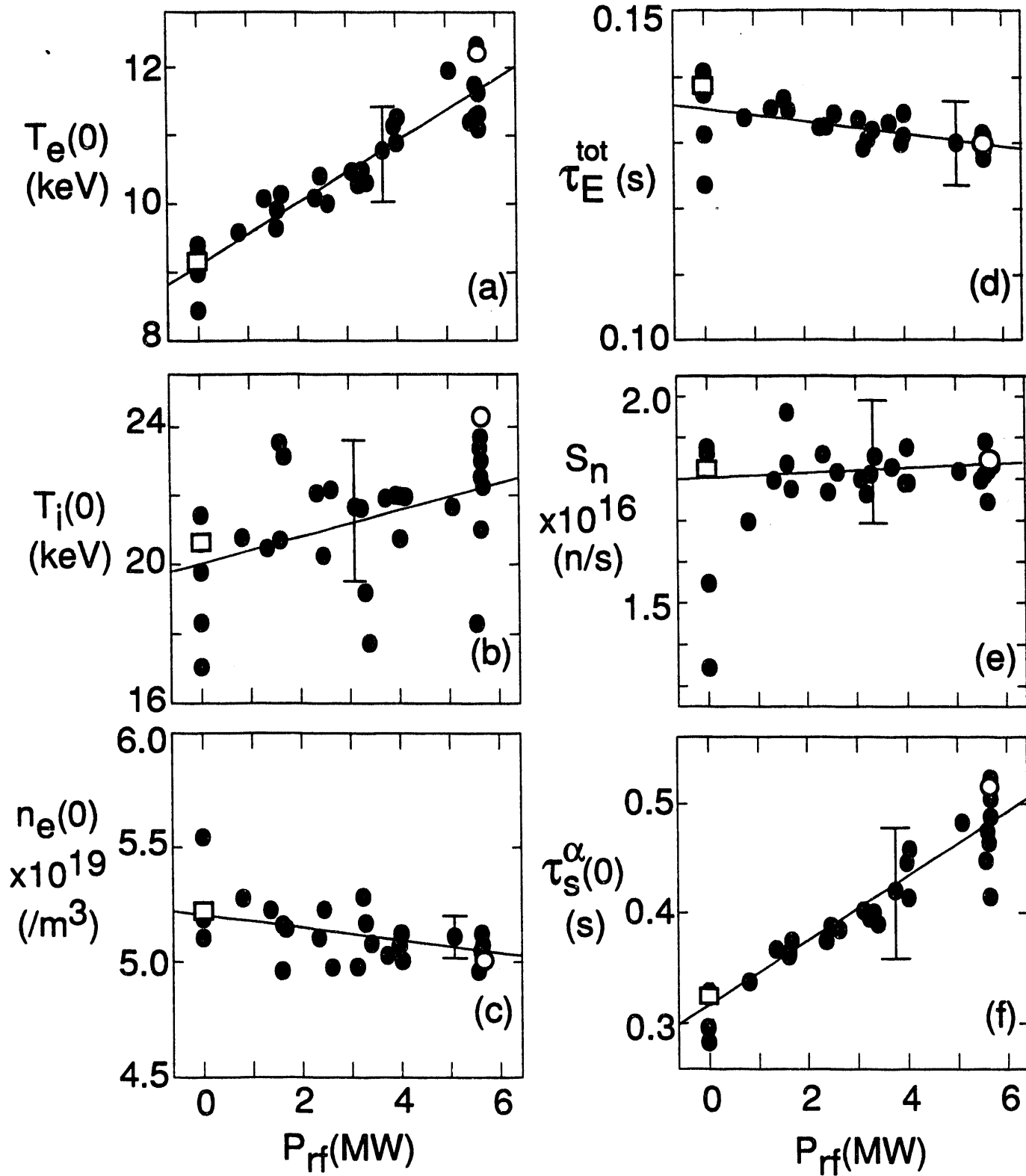


Fig. 5



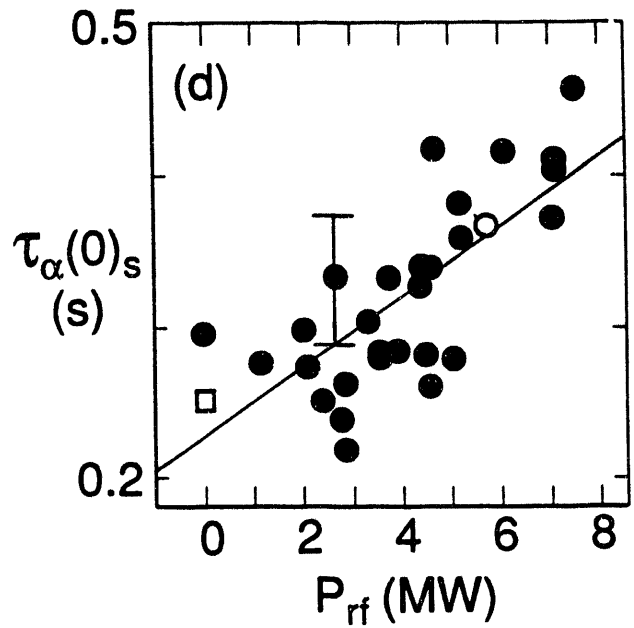
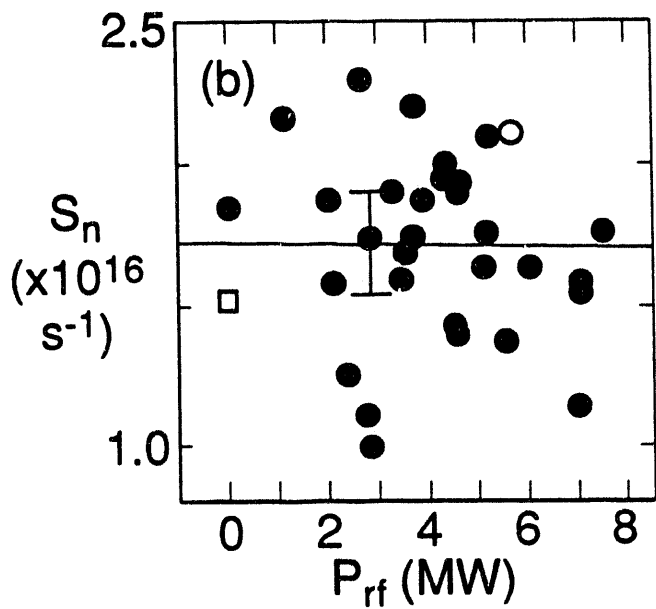
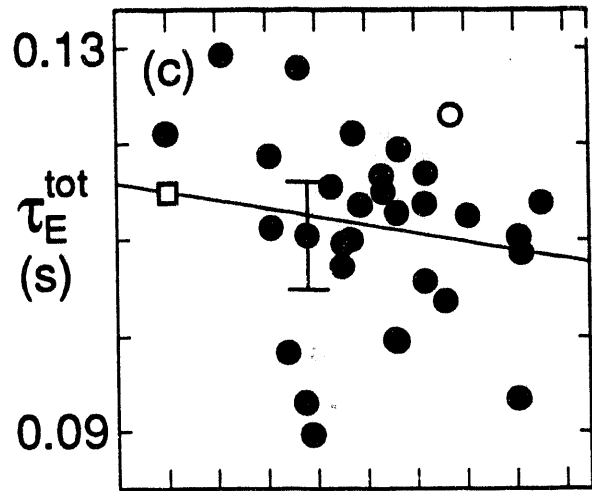
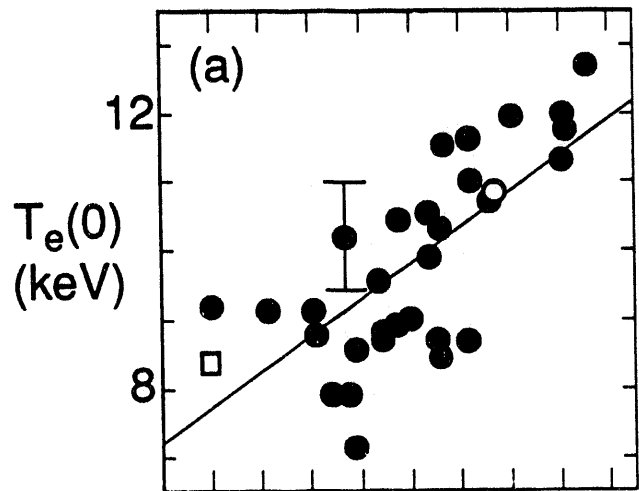
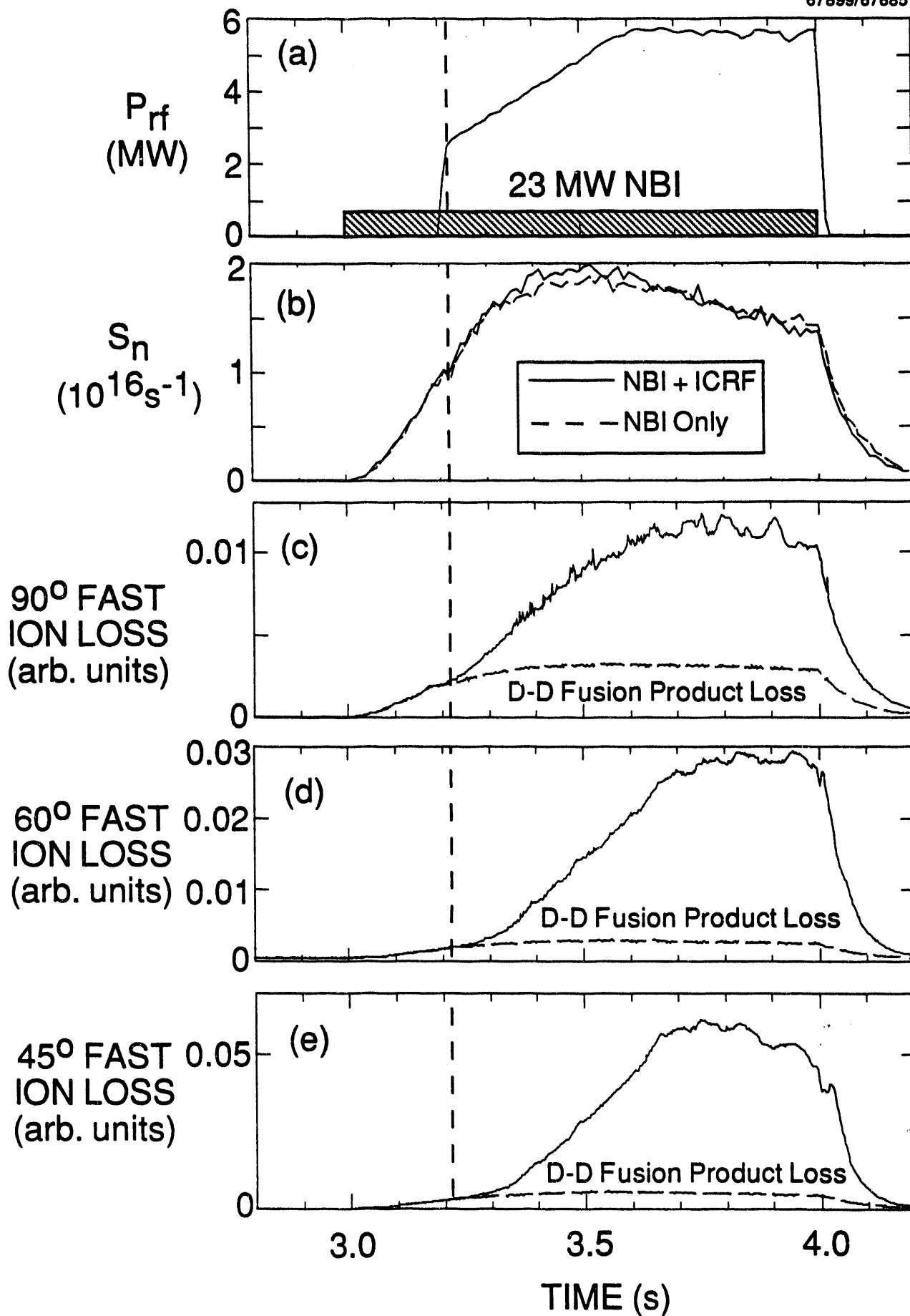


Fig 9



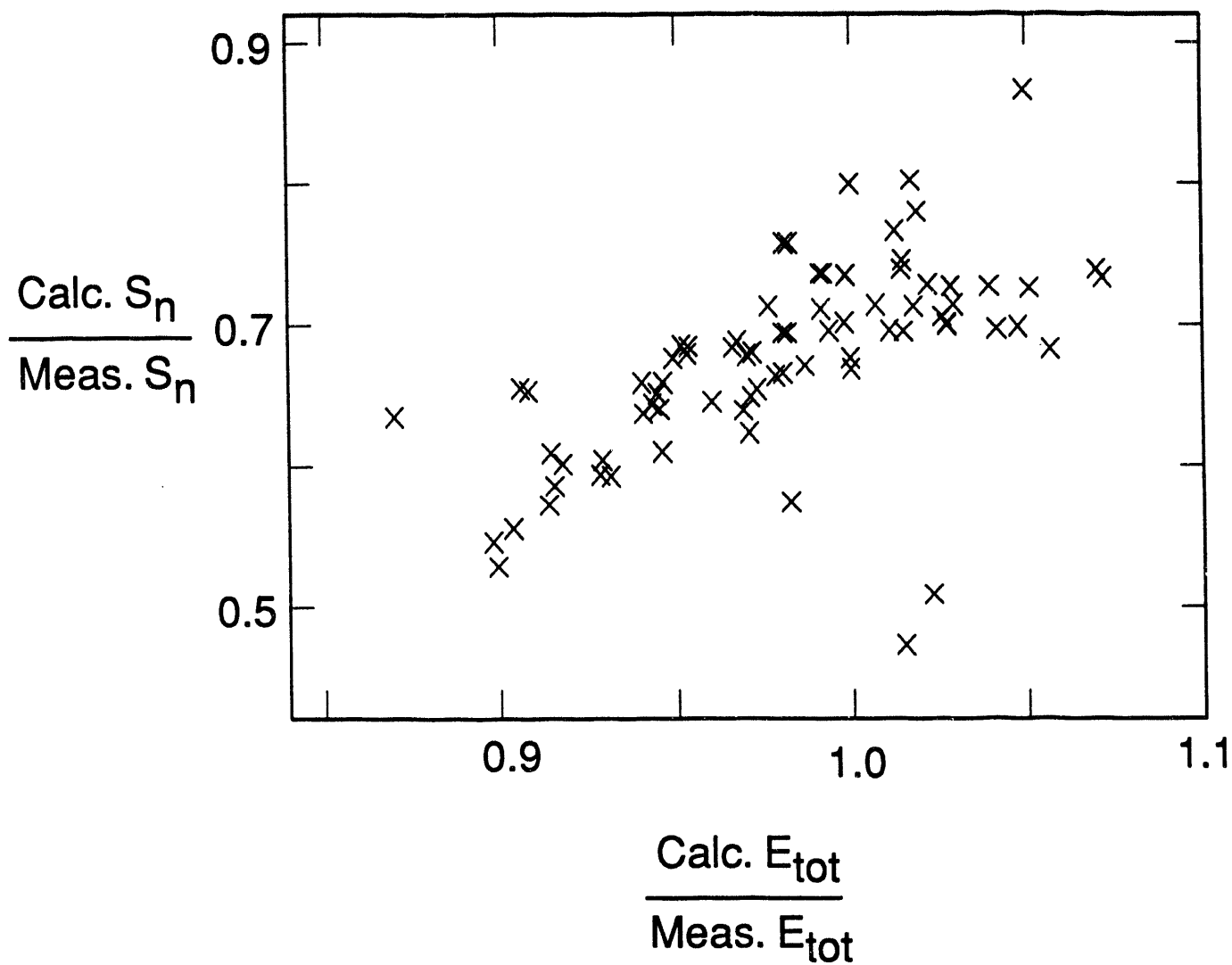
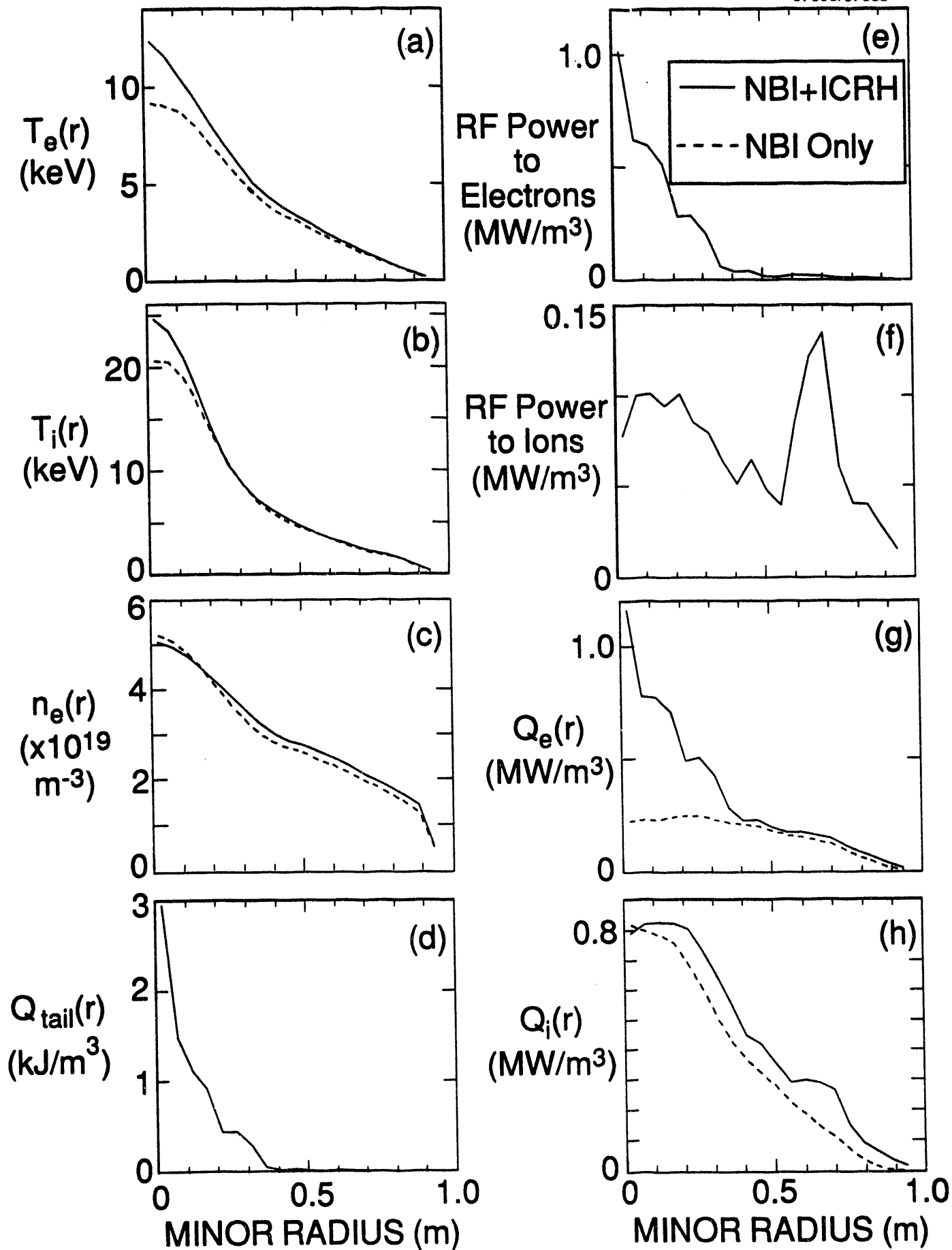


Fig. 13



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