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TANGENTIAL IMAGING SYSTEM

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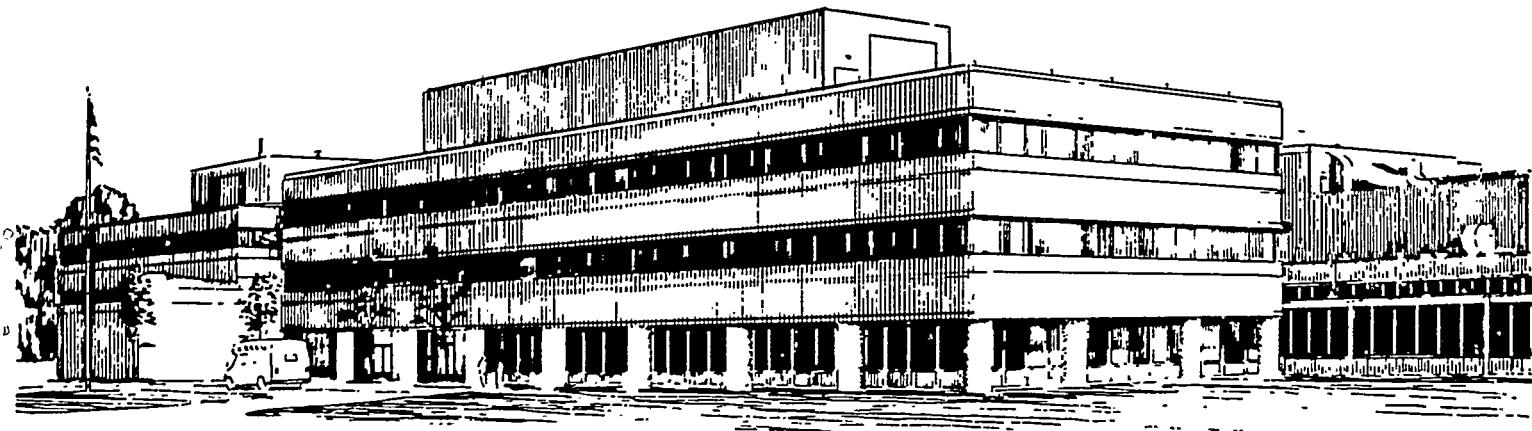
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A Linear Systems Description of the CO₂ Laser Based
Tangential Imaging System

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Abstract

We demonstrate that the phase variation produced by the projection of the density fluctuations onto a laser beam that is aligned tangent to the magnetic field lines in a toroidal plasma, is in fact a convolution of the density fluctuation profile in the tangency plane with a shift-invariant point spread function. Thus a spatial filter can be used to invert the corresponding transfer function to produce an undistorted image of the plasma density fluctuations at the tangency plane. Numerical simulations demonstrate that a spatial filter consisting of a simple and versatile step-function form of a Zernike phase mirror, will recover a reasonably accurate image of the fluctuations. [1]

1. Introduction

A CO₂ Laser Based Tangential Imaging System is being developed on CDX-U, which can, in principle, measure the two dimensional structure of density fluctuations in a toroidal plasma device [1]. The diagnostic overcomes two limitations of standard scattering techniques: good spatial resolution (under .5cm) is achievable for long fluctuation wavelengths; also the evolution in time of the entire spectrum of interest (.5cm - 8cm wavelengths) can be recovered from a spatially resolved image of the fluctuations [2], [3], [4].

A CO₂ laser operating at 10.6 microns is used to form an image of the plasma (electron) density fluctuations. The Gaussian laser beam is expanded to an 8cm diameter by a diverging lens, and then collimated by an off-axis parabolic mirror. The beam is then passed through the CDX-U tokamak along a horizontal chord that is aligned tangent to the magnetic field lines. The phase shifted beam is focussed by an imaging lens onto a phase mirror and then recollimated onto a horizontal detector array located at the image plane. The plane that contains the tangent points between the beam and the field lines is termed the tangency plane. Fig. 1 shows this and the geometry of the beam on CDX-U.

The technique uses the approximate uniformity of fluctuations along the magnetic field lines to recover an image of the plasma density fluctuations localized to the tangency plane [5]. As described by Nazikian and Grek [2], the effect of a projection through a toroidally symmetric plasma is to alias short wavelength features in the fluctuation profile into long wavelength, radial streaks in the projection. Since the nature of the aliasing feature is deterministic, it may be removed by focal plane filters.

In this paper we demonstrate that the projection of an object at the tangency plane through a toroidally symmetric plasma, can be represented to a good approximation by a shift-invariant point spread function acting on the object distribution. Thus a modified Zernike phase mirror must be implemented to recover the undistorted Fourier spectrum of the object. Understanding that the primary effect of the point spread function is to produce

long wavelength radial streaks in the image leads to a simple and versatile step-function design for a practical phase mirror, whose action approximates the required transfer function.

2. Derivation of the Point Spread Function

We assume a toroidally symmetric plasma and also that $1/q = 0$, where q is the safety factor. The beam center at the object or tangency plane is located at $x = -R_0, y = 0, z = 0$, where x is the horizontal coordinate across the beam, y is the vertical coordinate, and z is in the direction of beam propagation.

The plasma density fluctuations are mapped onto a variation in phase on the laser beam by a line integration (i. e. , a projection) through the plasma. In the absence of diffraction effects, the phase of the wave front is given by:

$$\phi(x_1, y) = \left(\frac{\lambda e^2}{c^2 m_e 4\pi \epsilon_0} \right) \int n_e(x = x_1 - R_0, y, z) dz, \quad (1)$$

where $\epsilon_0 = 8.854 \times 10^{-12} F/m$, n_e is the plasma electron density, and the range of integration is over the plasma chord. The beam centered coordinate, ($x_1 = x + R_0$), will be used for simplicity, from this equation on. It is assumed that R_0 is a constant over the beam cross-section.

Assuming the uniformity of fluctuations along the field lines,

$$\phi(x_1, y) \propto \int n_e(-\sqrt{(x_1 - R_0)^2 + z^2} + R_0, y) dz. \quad (2)$$

If a Taylor expansion in x_1 and z about the tangency plane is made (this is essentially a parabolic approximation to the field lines), we obtain:

$$\phi(x_1, y) \propto \int n_e(-\sqrt{(x_1 - R_0)^2 + z^2} + R_0, y) dz \approx \int n_e(x_1 - \frac{z^2}{2R_0}, y) dz. \quad (3)$$

Substituting $u = x_1 - \frac{z^2}{2R_o}$ yields

$$\phi(x_1, y) \propto \sqrt{2R_o} \int \frac{\Theta(x_1 - u)}{\sqrt{x_1 - u}} n(u, y) du, \quad (4)$$

which is now in the form of a convolution integral, with a shift invariant optical point spread function (P.S.F.) given by:

$$h(x_1, u) = h(x_1 - u) = \frac{\Theta(x_1 - u)}{\sqrt{x_1 - u}}, \quad (5)$$

with

$$\Theta(x_1 - u) = \begin{cases} 0 & \text{for } u > x_1 \\ 1 & \text{for } u < x_1. \end{cases}$$

Note that the limits of integration are extended to $\pm\infty$ with $n(x_1, y)$ set to zero outside the laser beam cross-section. The Taylor series expansion and the approximation that R_o is a constant, hold for $x_1/R_o \ll 1$, that is, for an imaging window that is much smaller than the major radius of its location. The first approximation is expected to be valid even though the second order term (i. e. , $z^2/2R_o$), in the expansion becomes sizeable, since the primary contributions to the integral in Eq. (2) come from the regions near the tangency plane where z is small. Fig. 2 shows the projection through the plasma of a Gaussian density disturbance, calculated from a numerical evaluation of Eq. (2), compared with that calculated from the point spread function. The two curves are identical to within a reasonable accuracy. $\Delta x/R$ for this example is $\approx .33$, as compared to only $\approx .23$ for CDX-U, where the approximation is thus expected to be better. The $1/e$ width of the disturbance is about 3cm , which signifies a Δk_x of about $.7\text{cm}^{-1}$. The essential form of the P.S.F. is also portrayed by the solid curve in Fig. 2; the streaking of an object in the tangency plane, in the x-direction is evident.

The integral expression can be rewritten in the form of a two-dimensional convolution

which is more general:

$$\phi(x_1, y) \approx \sqrt{2R_o} \int \int \frac{\Theta(x_1 - u)}{\sqrt{x_1 - u}} \delta(y - v) n(u, v) \, du \, dv. \quad (6)$$

Note that y variations are mapped onto the laser beam phase front with no distortion. Thus the y coordinate will be dropped for simplicity.

In frequency space the transfer function is simply the Fourier transform of the P.S.F.; Fourier transforming Eq. (5) gives:

$$\mathcal{F}\{\phi(x_1)\} \propto \mathcal{F}\{h(x_1) * n(x_1)\} = \mathcal{F}\{h(x_1)\} \mathcal{F}\{n(x_1)\}, \quad (7)$$

where $\mathcal{F}\{\dots\}$ represents the Fourier transform operation in the x direction.

Therefore,

$$\tilde{\Phi}(k_x) \propto H(k_x) \tilde{n}(k_x), \quad (8)$$

where k_x is the spatial wavenumber in the x direction of the plasma at the tangency plane, $\tilde{\Phi}(k_x) \equiv \mathcal{F}\{\phi(x_1)\}$, $\tilde{n}(k_x) \equiv \mathcal{F}\{n(x_1)\}$, and the transfer function is given by

$$H(k_x) \equiv \mathcal{F}\{h(x_1)\} = \begin{cases} \sqrt{\frac{1}{2|k_x|}} e^{+i\frac{\pi}{4}} & \text{for } k_x > 0 \\ \sqrt{\frac{1}{2|k_x|}} e^{-i\frac{\pi}{4}} & \text{for } k_x < 0 \end{cases}, \quad (9)$$

The amplitude and phase of $H(k_x)$ are shown plotted in Fig. 3. An expansion of the transfer function shows that the $k_x = 0$ component is unshifted in phase. Note that the low k_x (long wavelength) components are greatly amplified by the transfer function.

The Fourier modes of the fluctuation profile are shifted by $\pm\pi/4$ and the power spectrum of the projection is given by the simple form:

$$S_{\tilde{\Phi}(k_x)} = |\tilde{\Phi}(k_x)|^2 \propto \frac{1}{2|k_x|} S_{\tilde{n}}(k_x), \quad (10)$$

where

$$S_{\tilde{n}}(k_x) = |\tilde{n}(k_x)|^2. \quad (11)$$

3. Imaging of the Projection with a Spatial Filter

Phase contrast imaging is used to map the phase variations in the projection into a proportional intensity variation at the image plane where the detectors are located. Use of a standard phase contrast phase mirror, however, will result in a distorted image. Spectral analysis (detailed below) of a phase mirror that introduces an asymmetry between the positive and negative spatial frequencies ($\pm k_x$), shows that it is possible to both invert the transfer function inherent in the plasma projection as well as to recover an intensity image. A simple and versatile step-function design for a practical phase mirror will be presented.

For simplicity, consider the complex amplitude distribution of light of the projection, at the image plane, for a sinusoidal phase variation in x (as before, x, y are the coordinates transverse to the beam propagation direction, z . We use 'x' here to replace x_1 from Sec. 2):

$$u(x) = u_o e^{j\phi_o \sin(k_x x + \theta)} \quad (12)$$

$$\approx u_o \{1 + j\phi_o \sin(k_x x + \theta)\}. \quad (13)$$

The phase mirror is located at the focal plane of the imaging lens, and so will act on the angular spectrum (i. e. , the spatial Fourier transform) of the projection. The spatial coordinates x', y' on the mirror which is centered on the focal point, correspond to the spatial frequencies of the projection: $x' = k_x \lambda f$, and $y' = k_y \lambda f$, where λ represents the wavelength of the light, and f represents the focal distance of the imaging lens [6].

Now consider a phase mirror with its $k_x > 0$ half having an amplitude reflection coefficient $b(k_x) \cdot e^{i\beta(k_x)}$, and its $k_x \leq 0$ half having an amplitude reflection coefficient $a(k_x)$. In this paper, a, b , and β will be used to denote $a(k_x), b(k_x)$, and $\beta(k_x)$; a_o will represent $a(k_x = 0)$. The amplitude distribution becomes:

$$u(x) = u_o \left\{ a_o + \frac{\phi_o}{2} [b \cdot e^{i(k_x x + \theta + \beta)} - a \cdot e^{-i(k_x x + \theta)}] \right\}. \quad (14)$$

Thus the $+k$ and $-k$ sidebands of the sine wave variation are reflected differently. The corresponding intensity distribution is:

$$I(x) = |u(x)|^2 \approx |u_o|^2 \left\{ a_o^2 + a_o \cdot \phi_o [b \cdot \cos(k_x x + \theta + \beta) - a \cdot \cos(k_x x + \theta)] \right\}, \quad (15)$$

where $\phi_o \ll 1$ has been assumed so that the intensity varies linearly with ϕ_o .

Now, the expression may be rewritten for a sinusoidal density variation, $n = n_o \sin(k_x x + \theta)$, using Eq. (9) for the plasma transfer function:

$$\frac{I(x)}{|u_o|^2} \approx a_o^2 + a_o \cdot \frac{\kappa n_o}{\sqrt{2|k_x|}} \left[b \cdot \cos(k_x x + \theta + \beta + \frac{\pi}{4}) - a \cdot \cos(k_x x + \theta + \frac{\pi}{4}) \right] \quad (16)$$

$$\begin{aligned} &= a_o^2 + a_o \cdot \frac{\kappa n_o}{\sqrt{2|k_x|}} \left[\cos(k_x x + \theta) (b \cdot \cos(\beta + \frac{\pi}{4}) - \frac{a}{\sqrt{2}}) \right. \\ &\quad \left. + \sin(k_x x + \theta) (-b \cdot \sin(\beta + \frac{\pi}{4}) + \frac{a}{\sqrt{2}}) \right], \end{aligned} \quad (17)$$

where κ is a constant of proportionality. (Note that in the preceding analysis, $k_x > 0$ always, since it is the wavenumber of a sine wave spectrum; the 'negative k ' components produced by the plasma scattering are denoted by $-k_x$). One condition to recover an intensity variation that is proportional to the density fluctuations at the tangency plane, is to phase shift one of the sidebands by $\beta = \arccos(\frac{a}{b\sqrt{2}}) - \frac{\pi}{4}$, and to set $a \propto b \propto \sqrt{2|k|}$. A narrow strip of finite reflectance, a_o is required to reflect the main beam located at the $k = 0$ position. See Fig. 4 for a diagram of the phase mirror.

The equation for the intensity variation at the detector plane shows that an exact image of the density fluctuations is theoretically recovered:

$$\frac{I(x)}{a_o^2 |u_o|^2} = 1 + \frac{\kappa n_o}{\sqrt{2} a_o} \frac{a}{\sqrt{2|k_x|}} \left(1 - \sqrt{\frac{2b^2}{a^2} - 1} \right) \sin(k_x x + \theta). \quad (18)$$

Hence, $I(k_x) \propto n(k_x)$, with a 'contrast enhancement factor' equal to $1 - \sqrt{\frac{2b^2}{a^2} - 1}$.

Since a phase mirror with such reflectance properties is impractically expensive to manufacture, a much simpler yet effective alternative design is considered for the experiment: the reflectivities a, b are set to constant and finite values, with β remaining at

$\arccos(\frac{a}{b\sqrt{2}}) - \frac{\pi}{4}$. The reflectance of the mirror thus takes on a step-function form. By displacing the center of the phase mirror in x' , both sidebands of a long wavelength variation (ie. $\sin(k_x x + \theta)$ for $k_x < k_{min}$), are passed through the 'a' half so that they interfere and cancel at the image plane. One can adjust the x' displacement to select any cutoff wavelength (k_{min}) to optimize or adjust aspects of the recovered image.

The resultant intensity variation produced by the phase mirror is:

$$\frac{I(x)}{a^2|u_o|^2} = \begin{cases} 1 + \frac{\kappa n_o}{\sqrt{2}|k_x|} \frac{1}{\sqrt{2}} \left(1 - \sqrt{\frac{2b^2}{a^2} - 1} \right) \sin(k_x + \theta), & \text{for } |k_x| > k_{min} \\ 0, & \text{for } |k_x| < k_{min} \end{cases} . \quad (19)$$

Thus the amplitude factor of the plasma transfer function is approximately inverted in that the plasma primarily amplifies the long wavelength components of the density fluctuation profile, whilst the phase mirror eliminates the long wavelength components altogether. Phase distortion caused by the plasma is completely removed, however. The quality of the recovered image using this simpler phase mirror design is reasonably good according to numerical simulations. Fig. 5 shows the image recovered from a density fluctuation profile of random numbers that have been convolved with a Gaussian function of width $w_{1/e} \approx 3\text{cm}$, or $\Delta k \approx .7\text{cm}^{-1}$. We conclude from the simulation results that the step- function phase mirror design in principle, adequately removes the distortion in the projection, caused by propagation through the plasma.

4. Conclusion

We establish the somewhat counter-intuitive result that a tangential imaging beam will have a phase variation that is equal to the density fluctuation profile at the tangency plane, convolved by a shift-invariant point spread function, (for an imaging window that is much smaller than the major radius). It is shown that a modified Zernike phase mirror can be used to produce a deconvolved image of the profile. A practical, step-function phase

mirror design based on the knowledge that a projection through the plasma produces long wavelength radial streaks in the image, is shown to provide adequate image recovery in numerical simulations. Hence, it is demonstrated that the localized, two dimensional structure of density fluctuations can, in principle, be imaged.

5. Acknowledgment

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Figure Captions

Figure 1: Aerial view of the imaging geometry.

Figure 2: Comparison between a numerical simulation of (solid line), and an analytic approximation to (dashed line), Eq. (2) for a Gaussian test object with a $1/e$ width of 3cm , located at $x_1 = 0$.

Figure 3: The amplitude and phase of the plasma transfer function plotted against wavenumber (k_x).

Figure 4: Model phase mirror with three reflection zones, corresponding to $k_x < 0$, $k_x = 0$, and $k_x > 0$.

Figure 5: Numerical simulation of a density fluctuation profile in the tangency plane (solid line), compared with the filtered projection (dashed line), for a model phase plate with $a = a_o \ll 1$, $b = 1$, $\beta \approx +\pi/4$.

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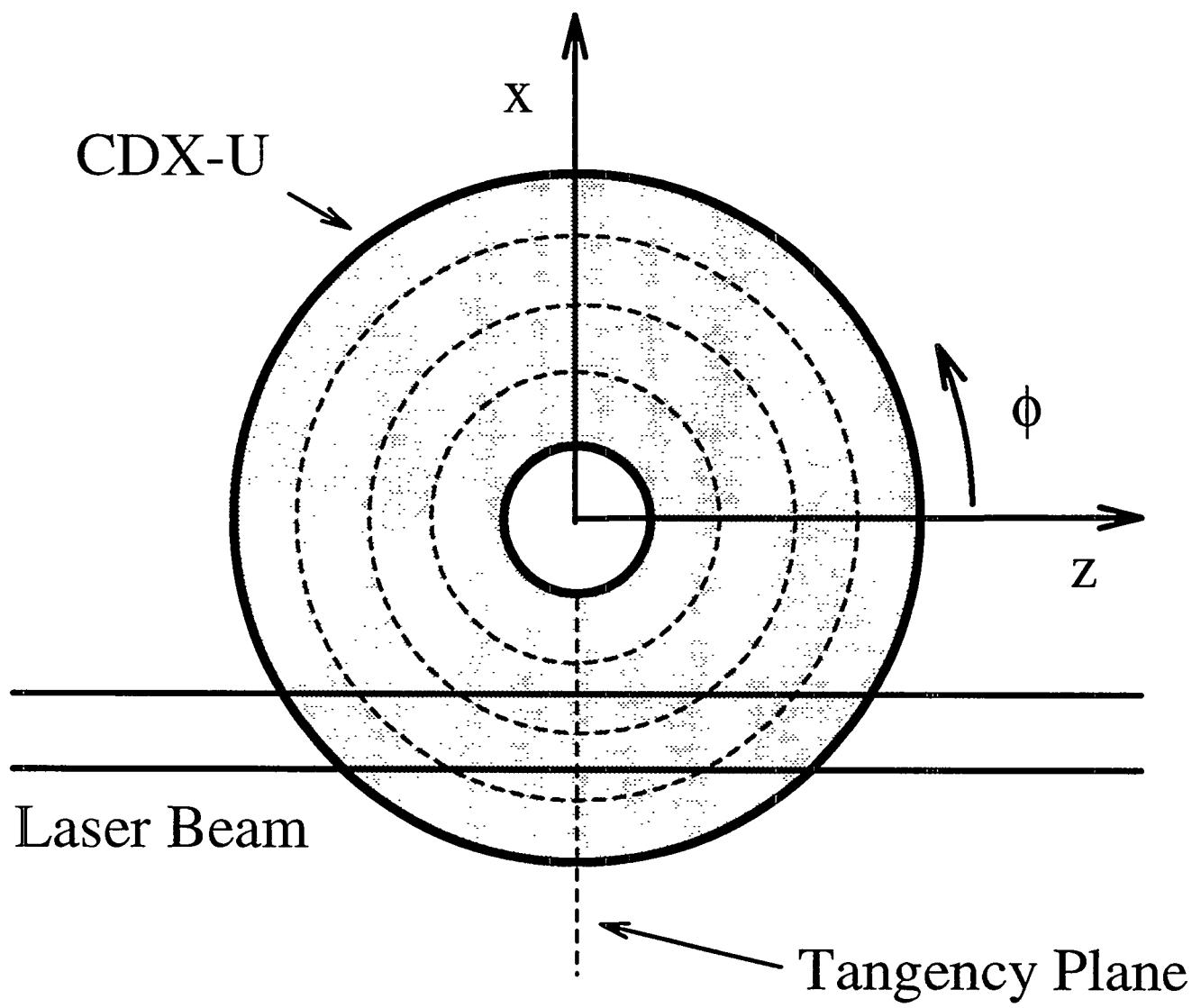


Figure 1

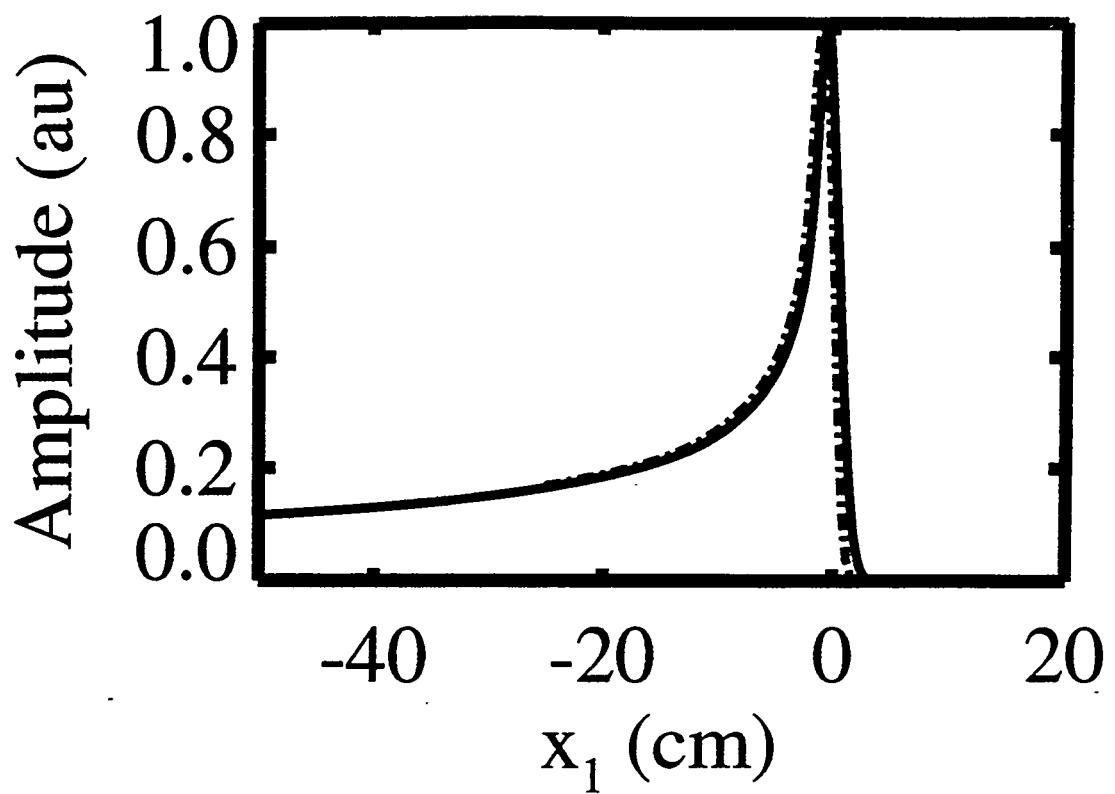


Figure 2

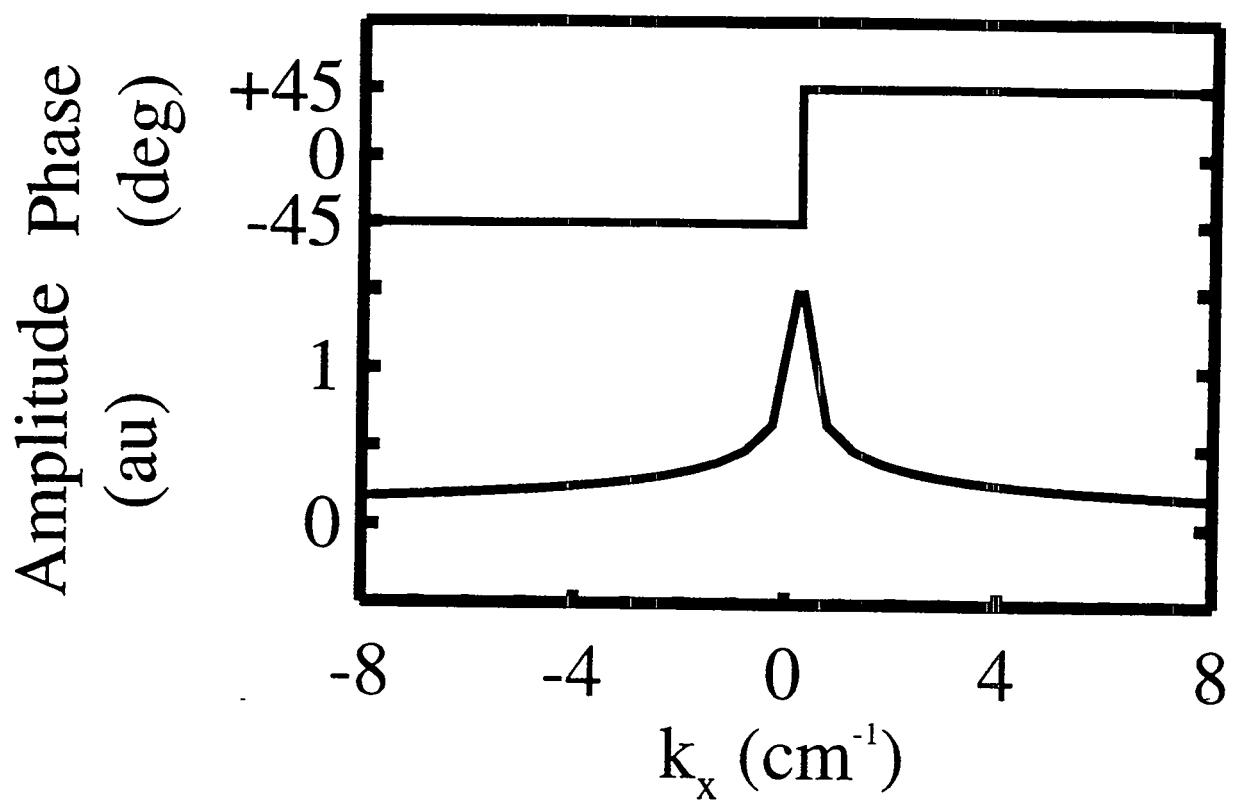


Figure 3

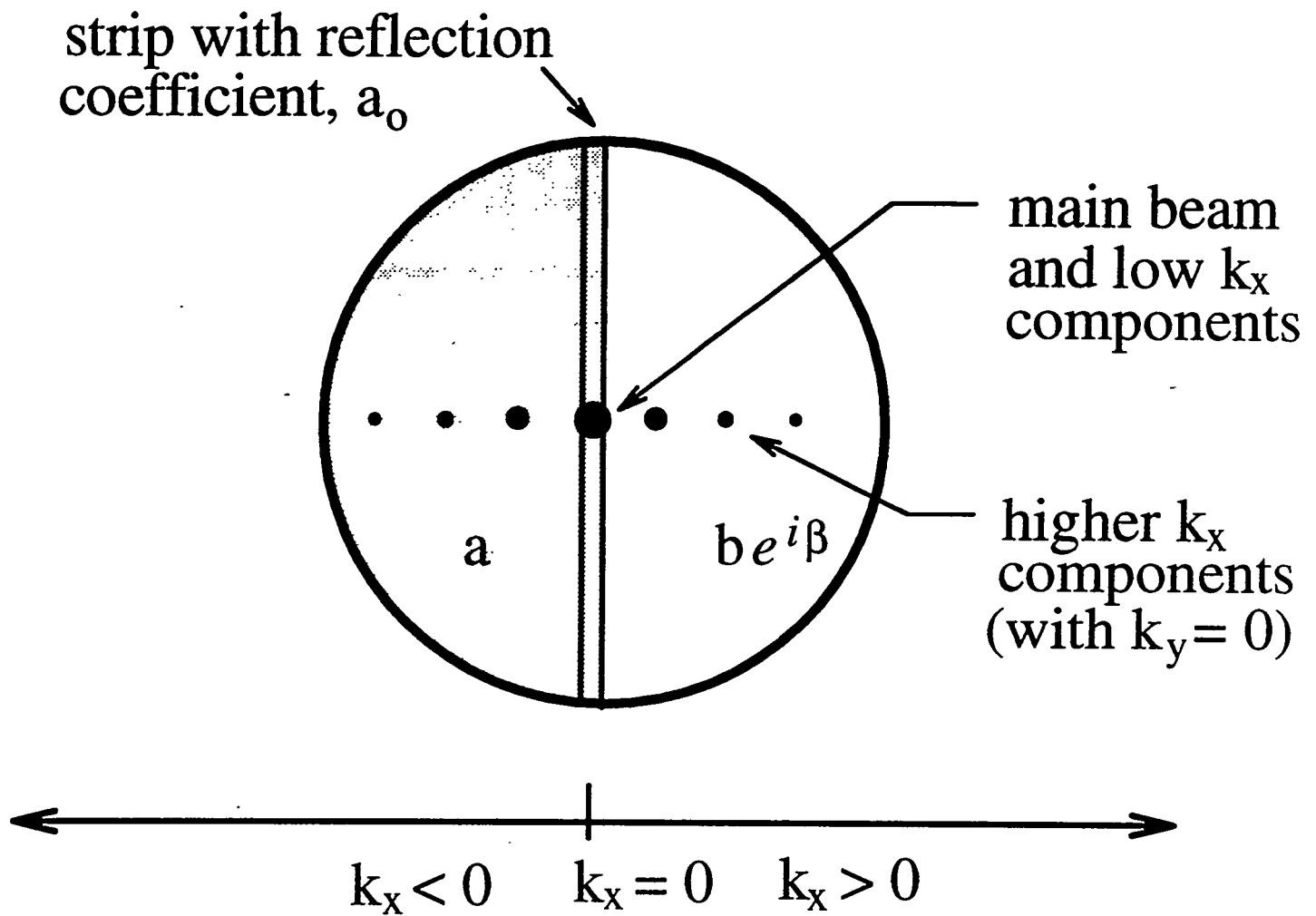


Figure 4

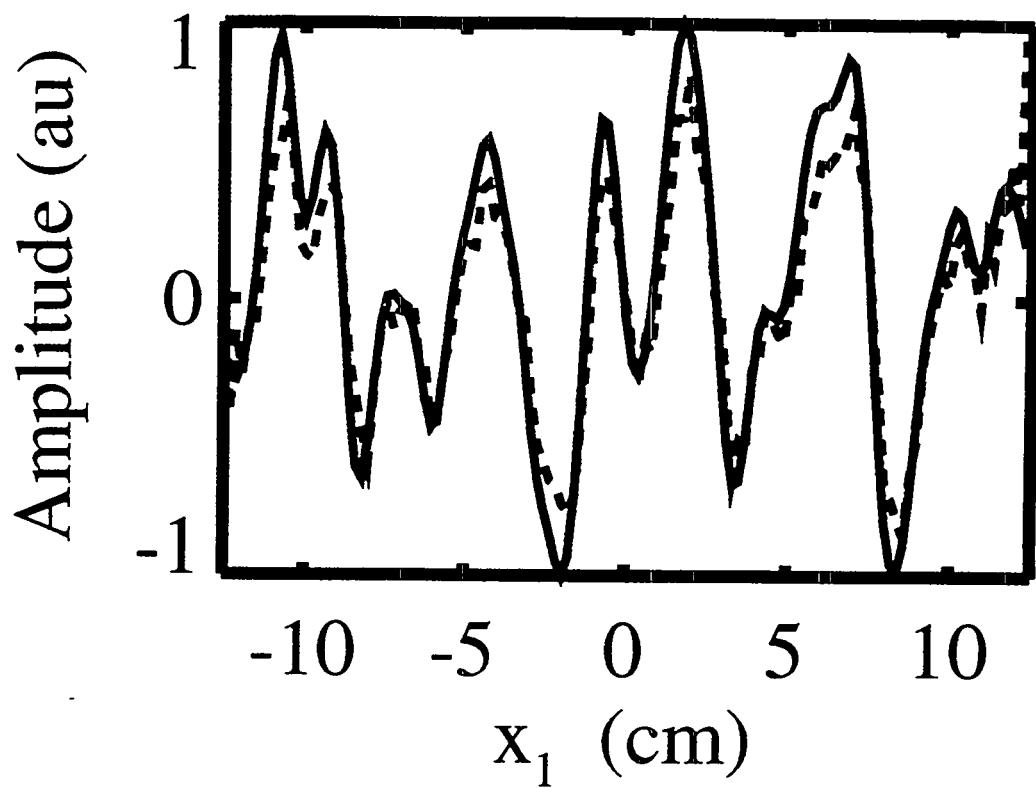


Figure 5

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