


Effect of Laser-Matter Interaction on Molten Pool Flow and Keyhole Dynamics

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Laser-induced keyholing occurs in additive manufacturing and welding processes, but the keyhole dynamics have not been well understood. A multiphase and multiphysics numerical model is used to predict the keyhole shapes recorded in the experimental observations and to predict transient and nonuniform distributions of laser absorption, temperature, and flow velocity in the process. When compared against data from a state-of-the-art dynamic x-ray radiography technique, good agreement is found for the keyhole shapes and fluctuation of the gas-liquid interface, thereby validating the simulation method. A detailed discussion is then given to elucidate the effects of laser absorption on the dynamic behavior of the front and rear keyhole walls. A quantitative comparison of different driving forces on the keyhole is also given to evaluate their significance to the keyhole dynamics.

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I. INTRODUCTION

The keyholing phenomenon is frequently found in laser-based metal additive manufacturing and welding processes. As shown in Fig. 1, a keyhole starts with a laser of an adequately high power intensity that induces not only melting but also surface evaporation of the metals. The high-speed flow of evaporated metal (denoted by the purple arrows) exerts a recoil pressure on the molten pool surface, which pushes the molten metal aside to create a keyhole. The gas-liquid interface of the keyhole usually exhibits strong fluctuations, which is found to be the major reason for the formation of spatters and pores [1–4]. The detailed fluctuation dynamics of laser-induced keyholes is complex and not fully understood. Extensive experimental and modeling efforts have focused on investigating this issue.

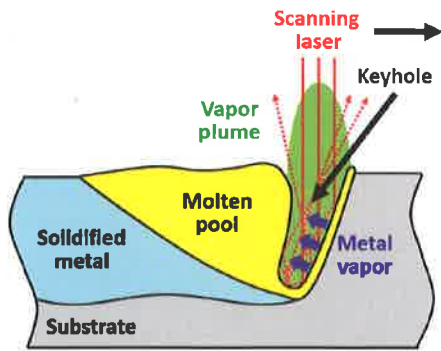
The experimental investigations of keyhole dynamics are primarily carried out through visualization of the keyhole shape. Optical imaging has been used most frequently to observe the keyhole dynamics from the top view [3,5–8] and side view [9–11]. Alternatively, x-ray imaging provides a unique capability to penetrate nondestructively opaque metals for *in situ* observation of keyhole dynamics without disturbing the process [1,2,4,12–17]. These visualization-based works have collectively disclosed that the keyhole adopts different shapes as a function of power

and speed and its internal surface generally fluctuates noticeably. The shape and fluctuation magnitude and frequency are affected by a variety of factors, including laser parameters (laser power, scanning speed, pulse width, duty cycle, etc.); the protective gas environment; and properties of substrate metals.

Since the absorbed laser light is the dominant energy source that induces keyhole formation, it is critical to measure the laser absorption by the keyhole. The measurements using calorimetry [18,19] or the integrated sphere method [20] have clearly revealed that the total laser absorption is noticeably increased when a keyhole is formed and that the increment in absorptivity is highly dependent on the keyhole geometry. This originates in the multiple-reflection phenomenon of the laser light inside the keyhole. The integrated sphere method can also capture the dynamic fluctuation of laser absorption [20], which is attributed to the dynamic fluctuation of keyhole geometry.

Despite the significant advancements in experimental observation of keyhole dynamics, important information regarding keyhole dynamics is still missing. First, the current techniques cannot measure the spatial distribution of the laser absorption on the keyhole wall, which is expected to be nonuniform and has decisive effects on the keyhole dynamics. Furthermore, the keyhole evolution is closely coupled with the metal evaporation and thermofluidic flows of the liquid/vapor metals, both of which are difficult to measure through experimentation, especially when the metals are in a high-temperature and dynamic

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Q4 F1:1 FIG. 1. Schematic of laser-induced keyhole.

74 state. Without the above information, it is infeasible to
 75 obtain a complete understanding of the keyhole dynamics.
 76 Numerical modeling complements experiments and
 77 enhances the understanding of keyhole dynamics by pro-
 78 viding information that is difficult to measure with exper-
 79 iments. Ray-tracing models are widely used to simulate
 80 the multiple reflections and partial absorption of laser light
 81 inside the keyhole. The models can successfully predict
 82 the nonuniform laser absorption in the keyhole of different
 83 shapes [21–26] and have well demonstrated that the laser
 84 absorption is highly dependent on the keyhole shape. In the
 85 more recent multiphysics modeling of keyhole dynamics
 86 [27–39], the ray-tracing model is coupled with a compu-
 87 tational thermofluidic dynamics model and a free surface
 88 tracking model (based on level-set, volume-of-fluid, or
 89 arbitrary Lagrangian-Eulerian methods). The ray-tracing
 90 model can predict the distribution of laser absorption inten-
 91 sity on the keyhole wall, which is used in the thermofluidic
 92 model to drive the variations of temperature field and fluid
 93 flow in both the molten pool and surrounding gaseous
 94 region. The predicted flow velocity will inform the surface
 95 tracking model, which updates the location of the metal

surface, a part of which is the keyhole wall. While these
 multiphysics models have already been used to investigate
 various issues regarding the formation and dynamic evo-
 lution of laser-induced keyholes, their predictions of the
 keyhole shape and fluctuation have rarely been validated
 against physical measurements of real keyholes. Hence,
 the quantitative accuracy of the simulations is still not
 clear.

In this paper, we present a combination of the state-
 of-the-art dynamic x-ray radiography (DXR) experimental
 measurements with an in-house multiphysics numerical
 model to investigate the dynamic phenomena of laser-
 induced keyholes in stainless steel (SS) 304. The DXR
 system at the Advanced Photon Source (APS) is first used
 to perform *in situ* observation of the keyhole dynamics
 for a wide range of laser powers and scanning speeds.
 The multiphysics model is then used to simulate the three-
 dimensional (3D) dynamic keyhole phenomenon. The
 model can well replicate the keyhole shape and fluctuation
 observed in the DXR experiments. The simulation results
 are further leveraged to understand the effects of nonuni-
 form laser absorption on temperature and flow fields of
 the liquid/vapor metals as well as the dynamic keyhole
 behavior.

II. RESULTS

A. Dynamic x-ray radiography experiments

In our DXR experiments, a high-energy, high-flux x-ray
 beam penetrates through the opaque metal substrate to
 image the keyhole shape with high spatial and temporal
 resolutions (see Appendix A for the details). The typi-
 cal keyhole shapes generated by different laser powers
 and scanning speeds are shown in Fig. 2. To quantify the
 keyhole shape, we schematically define the local keyhole
 depth d_k , local keyhole width w_k , and local front keyhole
 wall inclination α in the inset of Fig. 2. Here, we define the

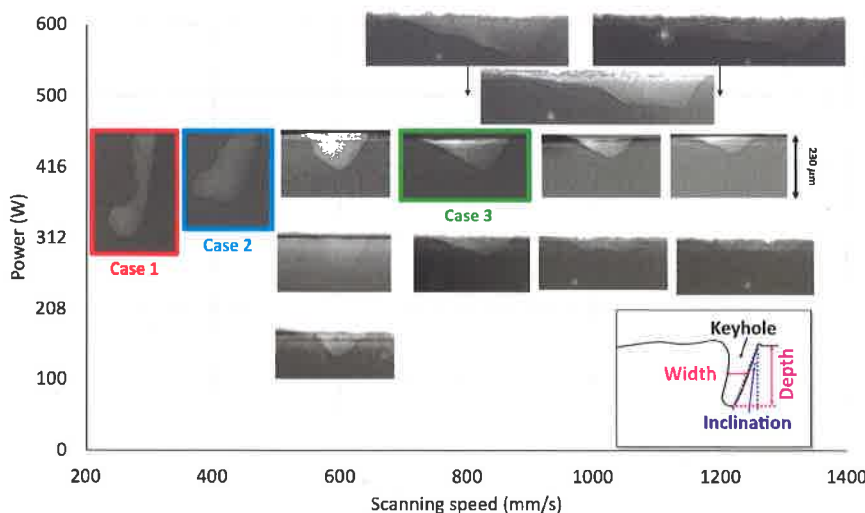
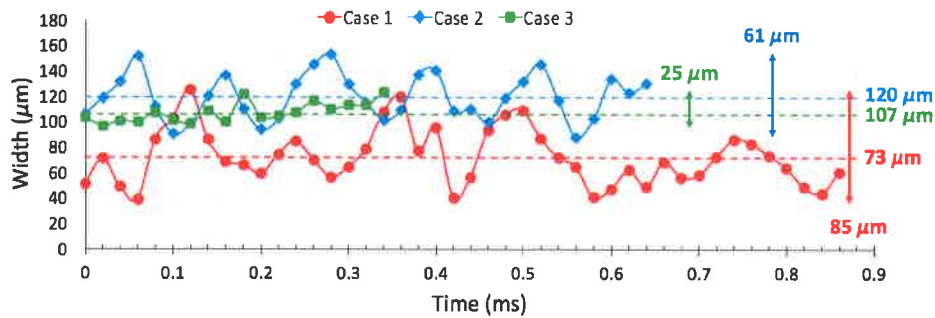


FIG. 2. Keyhole processing map for SS304 with different laser powers and scanning speeds. The colored boxes highlight the three selected cases of typical keyhole shapes for further investigation. For all three cases, the experiments are performed at the same laser power $P = 416$ W and varying scanning speeds: case 1 at $V = 300$ mm/s; case 2 at $V = 400$ mm/s; case at 3 $V = 800$ mm/s. The inset sketch defines the geometrical parameters used in the course of the study.



F3:1 FIG. 3. Temporal variation of keyhole width in the three selected cases: case 1 at $P = 416$ W, $V = 300$ mm/s; case 2 at $P = 416$ W,
 F3:2 $V = 400$ mm/s; case 3 at $P = 416$ W, $V = 800$ mm/s. The solid dotted lines show the instantaneous width, the dashed lines show the
 F3:3 average width, and the double-headed arrows show the magnitudes of width fluctuation.

131 keyhole width as the horizontal distance between the front
 132 and rear keyhole wall at mid-depth.

133 The effects of laser scanning speed are investigated
 134 as a series of experiments at constant laser power $P =$
 135 416 W. With increasing laser scanning speed ($V = 300 -$
 136 1200 mm/s), the keyhole depth decreases from around
 137 370 μm to around 118 μm , and the front wall inclina-
 138 tion α increases from around 9° to 47° . These trends are
 139 consistent with those reported in Ref. [5]. The keyhole
 140 width varies along the depth direction in all cases. When
 141 the laser scanning speed is ≤ 400 mm/s, the keyhole is
 142 generally wider at the lower portion than on the upper por-
 143 tion. As the scanning speed is increased to be ≥ 600 mm/s,
 144 the upper portion of the keyhole becomes wider than the
 145 lower portion. Three typical cases are selected from this
 146 series (case 1 in the red box, case 2 in the blue box, and
 147 case 3 in the green box in Fig. 2) for direct comparison
 148 with numerical simulation. In case 1, with the combina-
 149 tion of high laser power of $P = 416$ W and low scanning
 150 speed of $V = 300$ mm/s, the keyhole presents the tradi-
 151 tional deep and slender shape. As the scanning speed
 152 increases in case 2 to $V = 400$ mm/s, the keyhole becomes
 153 shallower and wider, presenting a quasiquadrilateral shape.
 154 A further increase of the scanning speed to $V = 800$ mm/s
 155 in the green case 3 changes the keyhole to a triangular
 156 shape.

157 Varying the laser power is investigated at a constant
 158 scanning speed of $V = 600$ mm/s. As the laser power
 159 increases from 208 to 416 W, the keyhole depth increases
 160 from about 55 to 129 μm , the front wall inclination
 161 decreases from around 45° to 34° , the keyhole width at half
 162 depth increases from around 68 to 155 μm , and similarly
 163 the keyhole width at the top surface increases from about
 164 102 to 220 μm .

165 While the front keyhole wall is generally stable, which
 166 leads to a relatively constant keyhole depth and front-wall
 167 inclination, the rear keyhole wall continuously changes its
 168 shape and relative location with respect to the front wall.
 169 Since the keyhole width is not uniform along the keyhole

170 depth direction, we quantify the general keyhole width
 171 with its value at half of the keyhole depth and plot, in
 172 Fig. 3, the temporal variation of the keyhole width for the
 173 three typical cases. The average keyhole width is the mini-
 174 mum for case 1 of $P = 416$ W and $V = 300$ mm/s (73 μm),
 175 indicating that the keyhole is rather narrow when the laser
 176 scanning speed is low. With the scanning speed increased
 177 to $V = 400$ mm/s and $V = 800$ mm/s, the average keyhole
 178 width is noticeably increased in case 2 (120 μm) and case
 179 3 (107 μm).

180 As the keyhole continuously changes its width around
 181 the average value, we quantify the fluctuation of keyhole
 182 width by its magnitude and period. The fluctuation mag-
 183 nitude is defined as the difference between the maximum
 184 and minimum width in each case and is found to decrease
 185 as the laser scanning speed increases from case 1 to case 3
 186 (85, 61, and 25 μm , as delineated by the colored double-
 187 headed arrows in Fig. 3). The fluctuation period is defined
 188 as the time difference between two adjacent moments when
 189 the instantaneous keyhole width decreases from above
 190 to below the average value and is found to be around
 191 100 μs in all three cases. This corresponds to a fluctuation
 192 frequency of around 10 kHz, which is of the same mag-
 193 nitude as the previous experimental measurements based
 194 on acoustic and optical signals [40–42]. Note that the fre-
 195 quency of 10 kHz primarily describes the fluctuation of
 196 the entire keyhole. Local fluctuations with higher frequen-
 197 cies are found at different locations and instances on the
 198 keyhole wall but the length scale and magnitude of those
 199 fluctuations are small and do not significantly affect the
 200 overall shape of the keyhole.

B. Multiphysics numerical simulations

201 Our simulations use a multiphysics model to predict
 202 the transient evolution of the 3D keyhole shape and the
 203 distributions of temperature, pressure, and flow velocity
 204 in the substrate, molten pool, keyhole, and surrounding
 205 air. The details of the model are given in Appendix B.
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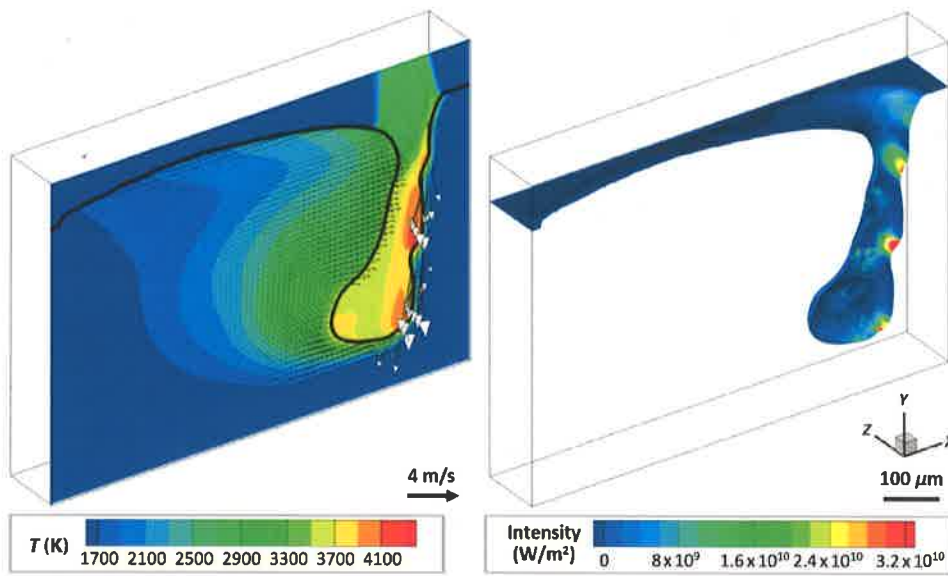


FIG. 4. Typical results of 3D simulation of laser-induced keyhole: (a) predicted temperature and velocity on the 2D center plane (velocity in the gas region is hidden for image clarity); (b) distribution of laser absorption on the 3D keyhole wall.

207 Figure 4 gives an example of typical simulation results, 208 with panel (a) showing the predicted temperature and 209 velocity distributions on the center plane of the track and 210 panel (b) showing the distribution of laser absorption on 211 the 3D keyhole wall.

212 The three representative cases shown in red (case 1), 213 blue (case 2), and green (case 3) in Fig. 2 are simulated 214 with our multiphysics model. The typical predicted 215 keyhole shapes for the three cases are shown and compared 216 with the DXR observations in Fig. 5. We compare the average 217 keyhole depth (denoted as \bar{d}_k), average keyhole width at 218 mid-depth (denoted as \bar{w}_k), and average front wall 219 inclination (denoted as $\bar{\alpha}$) from the DXR images and modeling 220 results for the three cases in Table I, which shows a reasonably 221 good agreement. Here, the average values are calculated based 222 on the measurements of

corresponding variables at multiple moments from either 223 DXR or simulation results. Furthermore, the model also 224 successfully captures the dynamic keyhole fluctuation. For 225 example, multiple snapshots from different moments in the 226 simulation of case 1 are shown in Figs. 5(f)–5(h), which 227 closely resemble the different keyhole shapes captured by 228 the DXR in Figs. 5(a)–5(c).

Beyond the capability to predict keyhole shape and fluctuation, 230 the model also predicts the laser absorption by 231 the keyhole, as shown in Figs. 5(f)–5(j). By integrating 232 all the absorbed laser power over the entire keyhole, we 233 find the total laser absorption to be approximately 322, 234 292, and ~ 186 W for the three respective cases, as given 235 in Table I. With the incident laser power as 416 W, the total 236 laser absorptivity of the keyhole is 80%, 70%, and 48% for 237 the three cases, respectively. These values are very close to 238

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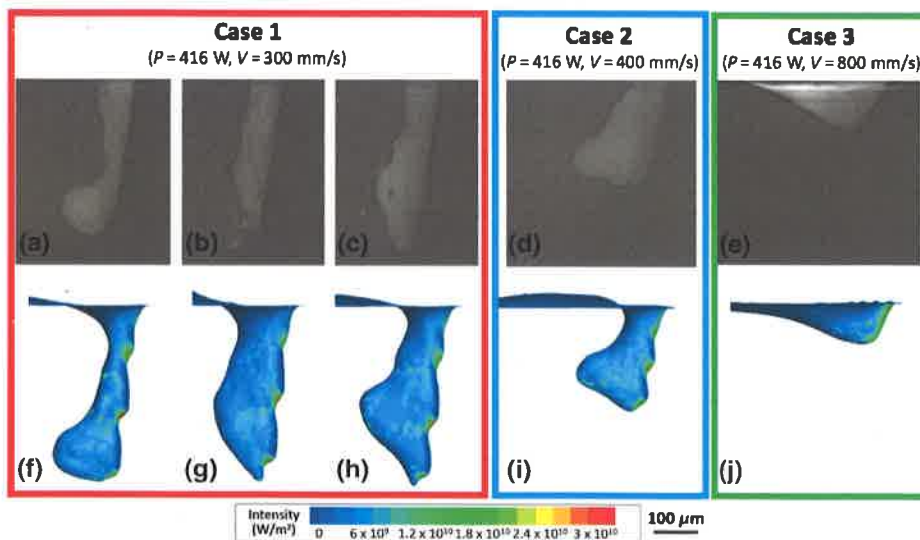


FIG. 5. Comparison of keyhole shapes from DXR observation and simulation results for the three selected cases. (a)–(e) DXR images of the keyhole shapes. (f)–(j) Simulated keyhole shapes and distributions of laser absorption. For illustrative purposes, different instances are shown for case 1, which exhibits the most fluctuation out of the three cases. (a)–(c) Case 1 at times t_0 , $t_0 + 140$ ms, and $t_0 + 340$ ms. (f)–(h) Predictions at three selected instances from the simulation.

TABLE I. Keyhole dimensions and laser absorption in the three selected cases. The keyhole dimensions are quantified by the keyhole depth, width at mid-depth, and front wall inclination. The laser absorption is quantified by the total absorption, fresh ray absorption, and reflected ray absorption.

		Case 1 $P = 416 \text{ W}$ $V = 300 \text{ mm/s}$	Case 2 $P = 416 \text{ W}$ $V = 400 \text{ mm/s}$	Case 3 $P = 416 \text{ W}$ $V = 800 \text{ mm/s}$
Average depth \bar{d}_k (μm)	DXR	~ 370	~ 250	~ 118
	Model	~ 405	~ 257	~ 103
Average width at mid-depth \bar{w}_k (μm)	DXR	~ 73	~ 120	~ 103
	Model	~ 80	~ 110	~ 140
Average front wall inclination $\bar{\alpha}$ (degrees)	DXR	~ 9	~ 12	~ 37
	Model	~ 9	~ 13	~ 26
Total absorption (W)	Model	332	292	186
Fresh ray absorption (W)	Model	136	135	141
Reflected ray absorption (W)	Model	196	157	45
Reflected ray absorption on front wall (W)	Model	86	70	0
Reflected ray absorption on rear wall (W)	Model	110	87	45

239 the *in situ* measurement of the laser absorption of a keyhole
240 in SS316 reported in Ref. [19].

241 III. DISCUSSION

242 In this section, we leverage the modeling results to dis-
243 cuss the physics of keyhole formation and fluctuation.
244 As explained below, the distribution of laser absorption
245 directly affects the temperature, surface evaporation, and
246 fluid flow in the process, all of which affect the geom-
247 etry of the front and rear keyhole walls. The key-
248 hole shape, in turn, alters the distribution of laser
249 absorption.

250 A. Laser absorption in dynamic keyhole

251 To the best of our knowledge, it is still challenging to
252 measure through experimentation the spatial distribution
253 and temporal variation of laser absorption by a moving
254 keyhole. Our multiphysics model provides quantitative
255 information regarding this issue.

256 1. Spatial distribution of laser absorption

257 Figures 6(a)–6(c) show the spatial distribution of laser
258 absorption of all laser rays. We break down the absorption
259 of all laser light into fresh ray absorption and reflected ray
260 absorption. Here, the fresh rays refer to those entering the
261 keyhole directly from the laser source that have not been
262 absorbed by the keyhole wall (as denoted by the solid red
263 lines in Fig. 1). After the fresh rays are partially absorbed
264 by the keyhole wall for the first time, they become reflected
265 rays (as denoted by the dashed red arrows in Fig. 1)
266 and stay inside the keyhole, where they experience multi-
267 ple additional impingements and partial absorption before
268 they escape from the keyhole. The variations in fresh ray
269 absorption intensity and reflected ray absorption are shown

in Figs. 6(d)–6(f) and 6(g)–6(i), respectively. The powers
for fresh ray absorption and reflected ray absorption are
given in Table I.

The fresh ray absorption takes place solely on the front
keyhole wall and the absorption power is very similar in

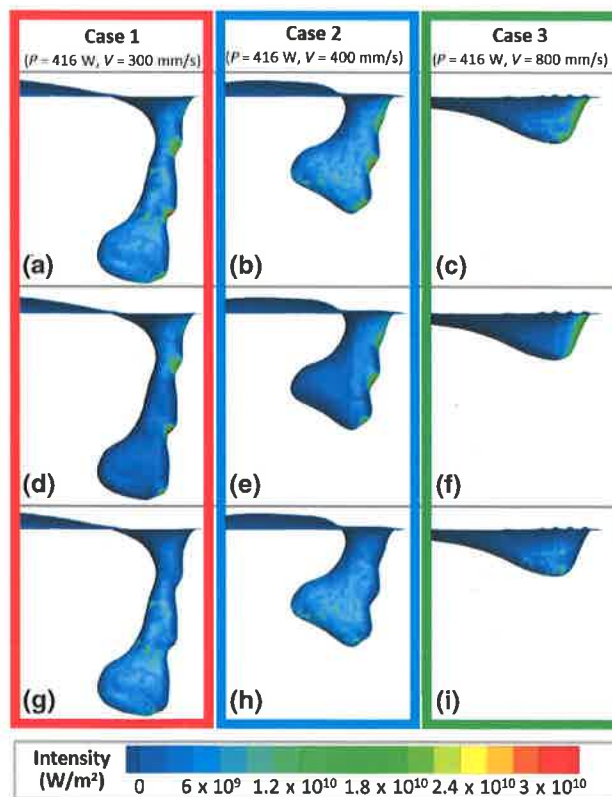


FIG. 6. Distribution of laser absorption intensity: (a)–(c) for all
rays, (d)–(f) for fresh rays, and (g)–(i) for reflected rays in the
three selected cases.

270
271
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F6:1
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275 the three cases (136, 135, and 141 W). With the inci-
 276 dent power being 416 W, the average absorptivity for the
 277 fresh ray on the front keyhole wall is approximately 33%.
 278 The spatial distributions of fresh ray absorption are quite
 279 different in the three cases. In cases 1 and 2, noticeable pro-
 280 trusions can be found on the front keyhole wall. The fresh
 281 ray absorption is concentrated on the up side of the protru-
 282 sions and generates very high local absorption intensities
 283 ($> 3 \times 10^{10}$ W/m²). In case 3, the front keyhole wall is flat
 284 and the absorption intensity is fairly uniform throughout
 285 the front wall (approximately 2.5×10^{10} W/m²).

286 The distribution of reflected ray absorption is dependent
 287 on the keyhole shape. In case 1, the keyhole has a larger
 288 width at the bottom than near the top, so the reflected rays
 289 can get trapped in the keyhole to experience multiple (usu-
 290 ally more than four) reflections before they escape. The
 291 reflected ray absorption is 196 W, much higher than the
 292 fresh ray absorption (136 W). The absorption is spread
 293 over the entire keyhole in a nonuniform manner, with a
 294 slightly larger fraction on the rear wall (110 W) than on the
 295 front wall (86 W). The reflected rays can generate absorp-
 296 tion intensity as high as 6×10^9 W/m² on the keyhole
 297 wall. While reflected rays completely determine the key-
 298 hole wall temperature on the rear keyhole wall, they just
 299 have minor effects on the front wall compared with the
 300 high absorption intensities caused by fresh ray absorption
 301 on the up side of the protrusion. In case 2, the absorption
 302 of reflected rays is distributed in a similar manner as in
 303 case 1. In case 3, the keyhole is wide compared to the
 304 spot size on top, which offers a large exit for the rays to
 305 escape. The reflected rays can be absorbed by the rear key-
 306 hole wall at most one time before they leave the keyhole.
 307 These rays are not reflected back to the front wall, so there
 308 is no reflected ray absorption on the front wall, which low-
 309 ers the total absorptivity. The decreased total absorptivity
 310 in the cases of higher scanning speed and lower molten
 311 pool depth is consistent with the experimental observation
 312 in Ref. [19].

313 **2. Temporal variation of laser absorption**

314 As the keyhole shape changes dynamically during the
 315 process, the laser absorption constantly deviates by a small
 316 amount from the results shown in Fig. 6 and Table I. Here,
 317 we select three scalar variables to quantify the spatial dis-
 318 tribution at each moment: total laser absorption, fresh ray
 319 absorption, and keyhole wall area above critical absorption
 320 intensity (denoted as $S_{critical}$). The term $S_{critical}$ is the total
 321 area of the keyhole wall regions where the local absorp-
 322 tion is above the critical power intensity for evaporation,
 323 which is estimated to be 1×10^9 W/m² [4]. Evaporation is
 324 expected to occur in these regions. Figure 7(a) shows larger
 325 fluctuations of total laser absorption (approximately 40 W)
 326 in cases 1 and 2 than in case 3 (< 10 W). This indicates
 327 a less stable keyhole shape in the first two cases, which

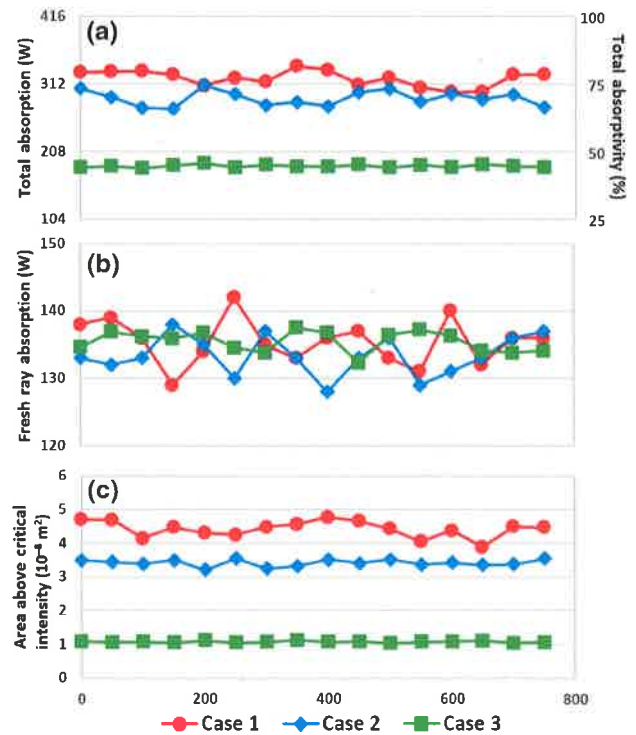


FIG. 7. Temporal variations of (a) total laser absorption and
 corresponding absorptivity (with respect to the input laser power
 416 W), (b) fresh ray absorption, and (c) keyhole wall area with
 absorption intensity above a critical value. For all three cases,
 the experiments are performed at the same laser power $P = 416$ W
 and varying scanning speeds: case 1 at $V = 300$ mm/s; case 2 at
 $V = 400$ mm/s; case 3 at $V = 800$ mm/s.

is consistent with the experimental observation of keyhole
 fluctuation in Fig. 3. Figure 7(b) shows very similar fresh
 ray absorption (between 130 and 140 W) and only minor
 fluctuations for all three cases. This suggests that fresh ray
 absorption is not highly dependent on the front wall geom-
 etry. This also suggests that the fluctuations of the total
 absorption in cases 1 and 2 are because of the randomness
 in the reflected ray absorption. Figure 7(c) shows that the
 total keyhole wall areas of approximately 4.5×10^{-8} m²
 in case 1 and approximately 3.5×10^{-8} m² in case 2
 (both with fluctuations of $> 3 \times 10^{-9}$ m²) have sufficient
 laser absorption to trigger local evaporation. The regions
 include the up side of the protrusions on the front keyhole
 wall and the scattered locations on the rear keyhole wall.
 In case 3, the laser absorption of around 1×10^{-8} m² (with
 a fluctuation of $< 1 \times 10^{-9}$ m²) of the keyhole wall can be
 above the critical intensity. The regions include the entire
 front keyhole wall and scattered regions on the rear wall.

In the following two subsections, we discuss the effects
 of laser absorption on the shape and fluctuation of both the
 front and rear keyhole wall.

B. Dynamics of the front keyhole wall

The front keyhole wall is primarily controlled by the fresh ray absorption. The power absorbed of fresh rays is very similar for the three cases, but the front keyhole shape and spatial distribution of fresh ray absorption intensity are quite different in the three cases. In cases 1 and 2, the front keyhole wall is less inclined with noticeable protrusions on the wall. The fresh ray absorption is concentrated on the up side of the protrusions and locally generates very high absorption intensity ($> 3 \times 10^{10}$ W/m²). In case 3, the front keyhole wall is more inclined and smooth, and the absorption intensity is fairly uniform (2.5×10^{10} W/m²). The spatial distribution of fresh ray absorption decisively affects the geometry of the front keyhole wall, including the front wall inclination and the protrusion generation on the front wall. The front wall geometry, in turn, alters the distribution of fresh ray absorption.

1. Inclination of front keyhole wall

A series of analytical models has already been proposed to estimate the front wall inclination and keyhole depth based on processing parameters and thermophysical properties of the materials [5,6,43,44]. Inspired by these works, here we present a similar analysis based on our numerical simulation results, which include more information regarding the dynamic thermofluidic phenomena during the laser-induced keyholing process.

The front wall inclination is determined by the combination of the laser scanning speed V_l and the drilling velocity V_d , as estimated by Eq. (1):

$$\bar{\alpha} = \arctan\left(\frac{V_l}{V_d}\right). \quad (1)$$

Here, V_l is 300, 400, and 800 mm/s for the three cases and V_d can be considered to be the vertical component of the liquid flow velocity on the front keyhole wall. The average V_d is found from the simulations to be 6.5, 3.4, and 2.2 m/s for the three cases. If we plug the V_l and V_d values into Eq. (1), the average inclination angles of the three cases are found to be 3°, 7°, and 20°, reasonably close to the measurement results from the DXR observations and simulations, as already given in Table I.

Note that approximately the same amount of power (130 W) is deposited on the front keyhole wall due to the fresh ray absorption in the three cases, but the resultant V_d is different. The term V_d can be estimated as

$$V_d = \int_{t_{\text{dwell}}} a_d dt. \quad (2)$$

Here, a_d is the acceleration for drilling primarily due to the recoil pressure. According to the simulations, the average acceleration due to the recoil pressure is approximately 2.8×10^4 , 2.7×10^4 , and 2.3×10^4 m/s² for the

three cases. The term t_{dwell} is the dwell time of fresh rays illuminating the front keyhole wall, which is calculated according to Eq. (3):

$$t_{\text{dwell}} = \frac{d_l}{V_l}. \quad (3)$$

Here, d_l is the diameter of the laser spot, which is 100 μm for our work, and V_l is the laser scanning speed. The fresh ray dwell time is found to be 3.33×10^{-4} , 2.50×10^{-4} , and 1.25×10^{-4} s for the three cases.

In case 1, multiple protrusions on the front wall lead to fresh ray absorption localization on the regions of very small surface area on the up side of those protrusions, and the resulting absorption intensity is rather high (3.2×10^{10} W/m²). The local temperature becomes very high (around 4200 K) and induces intense evaporation. A strong recoil pressure (as high as 1×10^6 Pa) is available to generate the highest acceleration for drilling (i.e., a_d). Meanwhile, the time for acceleration is the longest (i.e., $t_{\text{dwell}} = 3.33 \times 10^{-4}$ s). The drilling velocity on the front wall is therefore the highest (6.5 m/s).

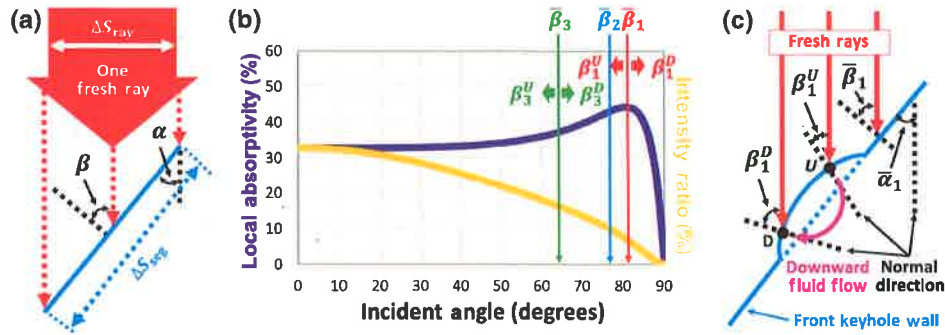
In case 2, protrusions are also present on the front keyhole wall. In this case, the absorption intensity (3.1×10^{10} W/m²), surface temperature (4200 K), and recoil pressure (1×10^6 Pa) on the front wall are similar to those in case 1. However, a shorter fresh ray dwell time (2.50×10^{-4} s) is available to accelerate the drilling velocity, leading to a decrease in V_d (3.4 m/s).

In case 3, the front wall is rather flat and the fresh ray absorption is uniformly distributed over the entire front wall, leading to a relatively low absorption intensity (2.5×10^{10} W/m²) and, hence, the surface temperature is relatively low (around 4000 K). A lower recoil pressure (6×10^5 Pa) is available for an even shorter fresh ray dwell time (1.25×10^{-4} s) to accelerate the drilling velocity. Moreover, since the front keyhole wall is flat, the pressure gradient due to the recoil pressure is along the normal direction of the front wall. Not the entire pressure gradient, but only its vertical component, is used to accelerate the drilling velocity. As a result, the drilling velocity is the lowest (2.2 m/s) in case 3.

2. Protrusion dynamics on the front keyhole wall

Protrusions have been found on the front keyhole wall in cases 1 and 2, and they facilitate the localization of fresh ray absorption to induce high absorption intensity and strong evaporation. The simulations well demonstrate the generation and movement of these protrusions.

For any segment on the front keyhole wall, e.g., the blue segment in Fig. 8(a), if the local inclination angle is α , the incident angle β of the incident fresh ray on this segment can be calculated as $\beta = 90^\circ - \alpha$. The absorption intensity



F8:1 FIG. 8. Laser absorption on the keyhole wall. (a) The absorption of one fresh laser ray by a linear segment of the front keyhole wall.
 F8:2 (b) Laser absorption by keyhole wall. The purple curve shows the local absorptivity as a function of incident angle. The orange curve
 F8:3 shows the ratio of the absorbed laser power intensity on the front keyhole wall over the incident laser power intensity from the fresh
 F8:4 laser ray. The red (β_1^U, β_1^D); blue (β_2); and green (β_3^U, β_3^D) terms mark out the incident angles of the fresh ray on the front
 F8:5 keyhole wall for cases 1, 2, and 3, respectively, and will be used in the discussion of Sec. III B. (c) Nonuniform absorption of fresh
 F8:6 rays across a protrusion on the front keyhole wall. Note that the inclination angle and protrusion size are exaggerated in this figure for
 F8:7 clarity of illustration.

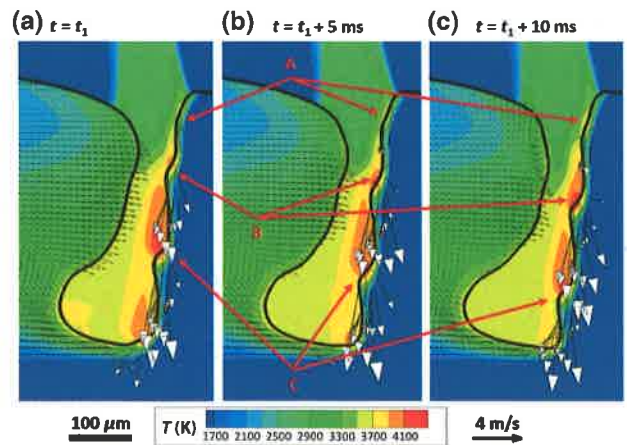
446 due to the fresh ray can be calculated as

$$\begin{aligned}
 I_{abs,fresh} &= \frac{I_{fresh} A(\beta) \Delta S_{ray}}{\Delta S_{seg}} \\
 &= I_{fresh} A(\beta) \cos \beta \\
 &= I_{fresh} R_{intensity}.
 \end{aligned}
 \tag{4}$$

450 Here, $I_{abs,fresh}$ is the absorption intensity on this segment
 451 due to the incident fresh ray, the power intensity of which
 452 is I_{fresh} . The term $A(\beta)$ is the local laser absorptivity as
 453 a function of the incident angle β . The dependence is
 454 shown graphically by the purple curve in Fig. 8(b), which
 455 is calculated with the Fresnel equation. The term ΔS_{ray} is the
 456 cross-section area of the ray. The term ΔS_{seg} is the
 457 surface area of the segment, which can be calculated as
 458 $\Delta S_{ray} = \Delta S_{seg} \cos \beta$. The multiplication term $A(\beta) \cos \beta$
 459 calculates the ratio of absorption intensity over incident
 460 laser intensity. This intensity ratio, $R_{intensity}$, as a function
 461 of β is shown by the orange curve in Fig. 8(b).

462 In case 1, the protrusions usually start near the top of
 463 the front keyhole wall on the convex surface because of
 464 the natural transition from the horizontal top surface ahead
 465 of the keyhole to the inclined front keyhole wall (see
 466 protrusion A in Fig. 9) and these protrusions flow down
 467 along the front keyhole wall until they reach the keyhole
 468 bottom (see protrusions B and C in Fig. 9). The distribu-
 469 tion of fresh ray absorption intensity on these protrusions
 470 is not uniform, as illustrated in Fig. 8(c). The average front
 471 keyhole wall inclination is $\bar{\alpha}_1 = 9^\circ$ and, correspondingly,
 472 the average incident angle of the fresh laser rays on the
 473 front wall is $\bar{\beta}_1 = 90^\circ - \bar{\alpha}_1 = 81^\circ$. For any point on the
 474 up side of the protrusion (e.g., point U), we can draw
 475 the local incident angle of a fresh ray on point U [i.e.,
 476 β_1^U in Fig. 8(c)] and find $\beta_1^U < \bar{\beta}_1$. Similarly, if we can

477 find one point from the down side of the protrusion (e.g.,
 478 point D) and draw the incident angle of a fresh ray on this
 479 point [i.e., β_1^D in Fig. 8(c)], we should find $\beta_1^D > \bar{\beta}_1$. By
 480 mapping the values of $\bar{\beta}_1, \beta_1^U$, and β_1^D in Fig. 8(b), it is
 481 evident that the intensity ratio on the up side of the pro-
 482 trusion ($R_{intensity}^U$) is generally higher than that on the down
 483 side of the protrusion ($R_{intensity}^D$), i.e., $R_{intensity}^U > R_{intensity}^D$.
 484 With similar power intensity I_{fresh} for the fresh rays inci-
 485 dent upon points U and D, Eq. (4) gives $I_{abs,fresh}^U > I_{abs,fresh}^D$,
 486 i.e., the fresh ray absorption intensity is higher on the up
 487 side than on the down side of the protrusion. As shown in



F9:1 FIG. 9. Three consecutive instants in time showing the generation and downward motion of protrusions on the front keyhole
 F9:2 wall in case 1. Protrusion A starts to form on the convex surface in the transition zone between the horizontal top surface
 F9:3 and inclined front keyhole wall. Protrusions B and C are moving
 F9:4 downward along the front keyhole wall (notice the high velocity
 F9:5 of the fluid flow on the front keyhole wall).
 F9:6
 F9:7

488 Fig. 6(d), the absorbed intensity is $> 3 \times 10^{10}$ W/m² on
 489 the up side of the protrusions and $< 2 \times 10^9$ W/m² on the
 490 down side of the protrusions. In certain scenarios, a por-
 491 tion of the down-side area can even be completely shaded
 492 by the protrusion and receive zero illumination by fresh
 493 rays. Thanks to the nonuniform distribution of fresh ray
 494 absorption, the temperature is considerably higher on the
 495 up side than the down side on the protrusion. Figure 5(d)
 496 shows the temperature on the up side of the protrusion
 497 at around 4200 K, which is well above the boiling point
 498 of the metal (3375 K) and causes strong evaporation and
 499 recoil pressure (1×10^6 Pa). On the other hand, the tem-
 500 perature on the down side is around 2900 K; hence, no
 501 evaporation and recoil pressure are induced. A remarkable
 502 pressure difference is created between the two sides of the
 503 protrusion, which induces a downward liquid flow, as indi-
 504 cated by the pink arrow in Fig. 8(c). This downward flow
 505 causes the protrusions to flow downward along the front
 506 keyhole wall as the protrusions grow in size. The intermit-
 507 tent generation of protrusion at the top of the front keyhole
 508 wall and their downward motion mean that the front wall
 509 appears to be wavy, which is consistent with the DXR
 510 observation for case 1.

511 In case 2, the average front keyhole wall inclination is
 512 $\bar{\alpha}_2 = 13^\circ$ and, correspondingly, the average incident angle
 513 of the fresh laser rays on the front wall is $\bar{\beta}_2 = 77^\circ$. Pro-
 514 trusions are also found in the DXR observations and in
 515 the numerical simulations for this case for reasons iden-
 516 tical to those described earlier in case 1. In case 3, the
 517 front keyhole wall becomes very flat and no protrusions
 518 are observed in the DXR measurements and the simulation.
 519 This suggests that the front wall becomes more resistant to
 520 protrusion generation compared to cases 1 and 2. The aver-
 521 age front keyhole wall inclination for case 3 is $\bar{\alpha}_3 = 26^\circ$
 522 and, correspondingly, the average incident angle of the
 523 fresh laser rays on the front wall is $\bar{\beta}_3 = 64^\circ$. Given a
 524 random protrusion on the front wall [similar to the one
 525 in Fig. 8(c)], the incident angle of fresh rays on the up
 526 side and down side of the protrusion (denoted as β_3^U and

527 β_3^D) are slightly different than $\bar{\beta}_3$. As shown in Fig. 8(b),
 528 the intensity ratio for both β_3^U and β_3^D are reasonably high,
 529 indicating that both the up side and down side of the protru-
 530 sion receive sufficient intensity to induce evaporation and
 531 recoil pressure. In this case, the liquid metal is driven away
 532 from both the up side and down side of the protrusion, and
 533 the protrusion is suppressed so that the front wall remains
 534 flat. The driven liquid flows around the side of the key-
 535 hole and enters the molten pool trailing the keyhole, as
 536 described in detail in the next subsection.

C. Dynamics of the rear keyhole wall

537
 538 The rear keyhole wall is found to have quite differ-
 539 ent shapes in the three cases. The keyhole width in the
 540 upper part of the keyhole is relatively small in cases 1
 541 and 2 but fairly large in case 3. The general shape of the
 542 rear keyhole wall is found to be independent of the laser
 543 absorption on the rear wall. Instead, it is closely related to
 544 the fluid dynamics initiated on the front keyhole wall. Fur-
 545 thermore, the general shape of the rear wall significantly
 546 alters the distribution of reflected ray absorption, which
 547 strongly affects the fluctuation of the rear wall from its
 548 average position.

1. Molten pool fluid dynamics and rear keyhole wall shape

549
 550 The predicted flow patterns in the molten pool trailing
 551 the keyhole are shown in Fig. 10. The pattern is most
 552 obvious in case 1. Two vortices are evident on the cen-
 553 ter plane, indicated by A and B in Fig. 10(a). Vortex A
 554 is located in the upper part of the molten pool with a
 555 counterclockwise flow, which is primarily driven by the
 556 thermocapillary force on the top surface of the molten pool.
 557 Vortex B is located in the lower part of the molten pool
 558 with a clockwise flow direction. This vortex is initiated on
 559 the front keyhole wall. Driven by the recoil pressure, the
 560 molten metal flows around the keyhole from the left or
 561 right sides and the bottom of the keyhole and enters the
 562

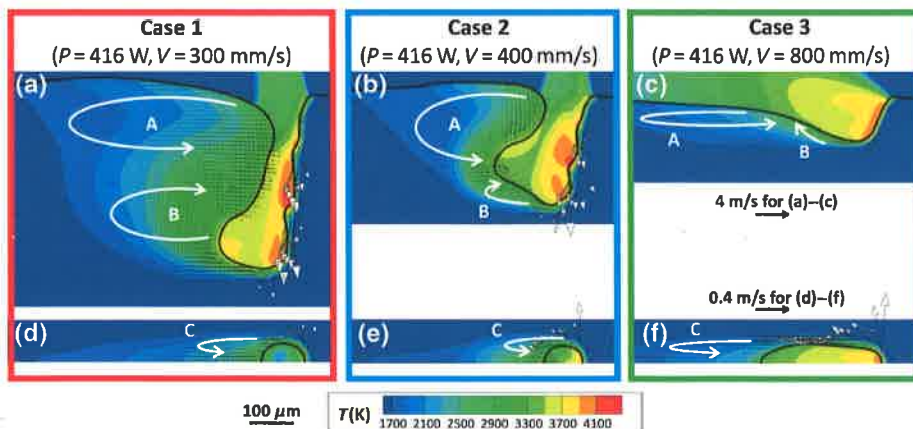


FIG. 10. Simulation results of F10:1 the temperature and velocity fields F10:2 in the three selected cases. (a)–(c) F10:3 Temperature and velocity fields on F10:4 the center plane from the simu- F10:5 lations. (d)–(f) Horizontal planes F10:6 looking from the top of the key- F10:7 hole near the surface of the sub- F10:8 strate from the simulations. Note F10:9 that the velocity field for vapor is F10:10 not shown. The vortices labeled F10:11 A, B, and C are discussed in Sec. F10:12 III C 1. F10:13

563 trailing molten pool. This flow moves toward the tail of
 564 the molten pool bottom and gradually turns upward along
 565 the fusion line. The upward flow collides with the down-
 566 ward flow in the upper vortex, and these two flows both
 567 turn to flow toward the rear keyhole wall. In a horizontal
 568 plane in the upper part of the molten pool close to the top
 569 surface, one vortex can be found in one half of the molten
 570 pool, as indicated by C in Fig. 10(d). The flow starts on the
 571 front keyhole wall, goes around the keyhole from the side,
 572 and moves backward toward the molten pool tail. Then the
 573 flow direction reverses to be toward the rear keyhole wall.
 574 The flow patterns in cases 2 and 3 are essentially similar to
 575 the one in case 1, except that vortex B becomes very small
 576 because the molten pool is shallow.

577 Note that the metal flows in all three vortices, either on
 578 the center plane, i.e., vortices A and B in Figs. 10(a)–10(c),
 579 or the horizontal plane, i.e., vortex C in Figs. 10(d)–10(f),
 580 start with a motion toward the molten pool tail, decelerate,
 581 and turn around to flow toward the rear keyhole wall. The
 582 liquid pressure near the rear keyhole wall is relatively low
 583 (primarily because of the local convex surface and hence
 584 the negative surface tension of the rear keyhole wall), so
 585 there is usually a pressure gradient toward the rear keyhole
 586 wall. The metal flows are accelerated by this pressure gra-
 587 dient as they flow toward the rear keyhole wall. When the
 588 metal arrives at the rear keyhole wall, the horizontal com-
 589 ponent of the flow speed is accelerated to equal V_l . This
 590 ensures that the rear keyhole wall can catch up with the
 591 front keyhole wall so that the distance between the front
 592 and rear keyhole walls (i.e., the keyhole width) remains rel-
 593 atively constant (with some fluctuation around the average
 594 width).

595 In case 1, as protrusions are generated on the front key-
 596 hole wall, the recoil pressure primarily drives the molten
 597 metal to flow downward so that its downward compo-
 598 nent (i.e., the drilling velocity, $V_d = 6.5$ m/s) dominates
 599 the resultant flow, as shown in Fig. 10(a). In a horizon-
 600 tal plane in the upper part of the molten pool, close to its
 601 top surface, as shown in Fig. 10(d), the backward com-
 602 ponent of the flow velocity V_b is correspondingly very
 603 small ($V_b = 300$ mm/s). This slow backward flow is soon
 604 reversed and accelerated to $V_l = 300$ mm/s toward the rear
 605 keyhole wall. Since V_b and V_l are both small compared to
 606 the scan speed, the velocity reversal process can be com-
 607 pleted in a very short time and travel distance so that the
 608 rear keyhole wall is very close to the front keyhole wall;
 609 i.e., the keyhole width is very small in the upper part of
 610 the keyhole. In case 2, the flow velocity in the horizontal
 611 plane as shown in Fig. 10(e) is similar to that in case 1 and,
 612 hence, the keyhole width is also small in the upper part of
 613 the keyhole.

614 In case 3, however, the front keyhole is more inclined
 615 and flatter, and the recoil pressure is more uniformly dis-
 616 tributed on the front keyhole wall, which pushes the molten
 617 metal around the keyhole from the sides. As shown in

Fig. 10(f), the backward flow becomes very fast ($V_b =$ 618
 700 mm/s), and it takes a relatively long time and travel 619
 distance for the flow to reverse its direction and acceler- 620
 ate to the high laser scanning speed of $V_l = 800$ mm/s in 621
 this case. By the time the flow reaches the rear keyhole 622
 wall, the front keyhole wall has traveled forward for quite 623
 a long distance. Hence, the distance between the front and 624
 rear keyhole wall, i.e., the keyhole width, becomes rather 625
 large in the upper part of the keyhole. 626

627 With the keyhole width in cases 1 and 2 being small near 628
 the top surface, the “trapping” effect encourages higher 629
 absorption for reflected rays (196 W for case 1 and 157 W 630
 for case 2). The reflected rays significantly heat the rear 631
 keyhole wall (primarily in the lower portion) to above 632
 the boiling point. Local evaporation and recoil pressure 633
 is induced, pushing the rear keyhole wall backward and 634
 enhancing the keyhole width. With the keyhole wall area 635
 much larger in case 1 (approximately 7×10^{-8} m²) than 636
 in case 2 (approximately 4.5×10^{-8} m²), the reflected ray 637
 absorption intensity density should be generally lower in 638
 case 1 (averaged at approximately 2.8×10^9 W/m²) than 639
 in case 2 (averaged at approximately 3.5×10^9 W/m²). 640
 As a result, the temperature on the rear keyhole wall in 641
 case 1 averages approximately 3300 K with the maximum 642
 value above 3600 K, while in case 2 the average tempera- 643
 ture is above 3400 K with the maximum value above 3700 644
 K. Because of the slightly higher temperature in case 2, a 645
 higher recoil pressure is available on the rear wall to push 646
 it farther away from the front wall. Therefore, the keyhole 647
 width is larger in case 2 (120 μ m) than in case 1 (73 μ m), 648
 as given in Fig. 3.

649 In case 3, the keyhole width is rather large near the sur- 650
 face of the keyhole, which offers a large exit for the rays 651
 to escape. According to our simulation, all the reflected 652
 rays will only be absorbed once by the rear keyhole wall, 653
 and their reflections will be able to leave the keyhole. 654
 Therefore, the reflected ray absorption power is drastically 655
 reduced to 56 W and it causes no evaporation on the rear 656
 keyhole wall to modify the general shape of the keyhole.

657 2. Dynamic fluctuation of rear keyhole wall

658 Strong keyhole fluctuation is found in cases 1 and 2, and 659
 it is primarily caused by the random movement of the rear 660
 keyhole wall away from its general shape. Such random 661
 movement of the rear keyhole wall results from the ran- 662
 domly varying distribution of reflected rays absorbed on 663
 the rear keyhole wall, which is initiated by the wavy front 664
 from the protrusions on the front keyhole wall. While the 665
 fresh rays are presumably parallel to each other before they 666
 land on the protrusions, the reflected rays follow different 667
 directions and illuminate the rear keyhole wall at differ- 668
 ent locations. The directions of the reflected rays become 669
 essentially random beyond the first reflection. Therefore, 670
 no fixed pattern of absorption intensity distribution can be

TABLE II. Order of magnitude of the major forces driving the molten pool flow and keyhole evolution. These forces are evaluated at the front keyhole wall, rear keyhole wall, and top surface of the molten pool in the three selected cases studied in this paper.

	Case 1 $P = 416 \text{ W}$ $V = 300 \text{ mm/s}$	Case 2 $P = 416 \text{ W}$ $V = 400 \text{ mm/s}$	Case 3 $P = 416 \text{ W}$ $V = 800 \text{ mm/s}$
Front keyhole wall			
Maximum temperature (K)	~ 4200	~ 4200	~ 4000
Maximum recoil pressure (Pa)	10^6	10^6	10^5
Maximum curvature (1/m)	-10^4	-10^4	-10^4
Maximum capillary force (Pa)	-10^4	-10^4	-10^4
Maximum temperature gradient (K/s)	10^7	10^7	10^7
Maximum thermocapillary force (Pa)	10^4	10^4	10^4
Rear keyhole wall			
Maximum temperature (K)	~ 3600	~ 3700	~ 2800
Maximum recoil pressure (Pa)	10^5	10^5	0
Maximum curvature (1/m)	-10^4	-10^4	-10^4
Maximum capillary force (Pa)	-10^4	-10^4	-10^4
Maximum temperature gradient (K/s)	10^6	10^6	10^6
Maximum thermocapillary force (Pa)	10^3	10^3	10^3
Top surface of molten pool			
Maximum temperature (K)	< 3000	< 3000	< 3000
Maximum recoil pressure (Pa)	0	0	0
Maximum curvature (1/m)	10^3	10^3	-10^3
Maximum capillary force (Pa)	10^3	10^3	-10^3
Maximum temperature gradient (K/s)	10^6	10^6	10^6
Maximum thermocapillary force (Pa)	10^3	10^3	10^3

671 found for reflected rays, and hot spots with high absorp-
 672 tion intensities can appear at random locations. At each
 673 hot spot, local evaporation is induced and recoil pressure
 674 is generated that pushes the rear keyhole wall backward,
 675 creating a local dent on the rear keyhole wall. As the hot
 676 spots wander, the dents move accordingly. There is a hot
 677 spot near the bottom in Fig. 5(f) and also around the middle
 678 depth in Fig. 5(h).

679 The fluctuation is significantly reduced in case 3 for two
 680 reasons. First, the front keyhole wall is nearly flat, with
 681 a relatively uniform distribution of laser absorption and
 682 recoil pressure. This helps the molten metal on the front
 683 keyhole wall to develop a steady flow going around the
 684 keyhole to build a stable rear keyhole wall. Second, the
 685 reflected ray absorption is insignificant on the rear keyhole
 686 wall. No evaporation is induced and thus no meaningful
 687 disturbances on the rear keyhole wall.

688 D. Effect of major forces on the keyhole dynamics

689 After the discussion of the laser absorption and the
 690 dynamic phenomena on the front and rear keyhole wall,
 691 here we present a direct comparison of various driving
 692 forces for both the molten pool flow and keyhole evolution.
 693 The major driving forces to be investigated are the recoil
 694 pressure, capillary force, and thermocapillary force. The
 695 orders of magnitude of these forces at different locations
 696 of the keyhole are calculated from the simulation results

(formula given in Appendix B) and are listed in Table II.
 Note that the recoil pressure and capillary force are applied
 along the normal direction of the molten pool surface and
 keyhole wall, while the thermocapillary force is applied
 along the tangential direction.

On the front keyhole wall, the temperature is above 4000
 K because of the intensive heating from the fresh laser
 rays, which induces strong evaporation to generate a strong
 recoil pressure on the order of $10^5 - 10^6$ Pa. This pressure
 is applied along the normal direction of the keyhole
 wall toward the liquid region and acts as the major force
 to maintain the inclined front keyhole wall while pushing
 it forward. Since the front keyhole wall is a concave inter-
 face, the curvature [K in Eq. (B8)] is negative on the order
 of -10^4 m^{-1} . With the surface tension coefficient [γ in
 Eq. (B8)] being 1.6 for the liquid metal, the capillary force
 is on the order of -10^4 Pa. While this force acts along
 the opposite direction with respect to the recoil pressure,
 it is at least 2 orders of magnitude smaller and therefore is
 overruled by the recoil pressure. With the maximum sur-
 face temperature gradient [$\nabla_s T$ in Eq. (B3)] on the order
 of 10^7 K/m and the surface tension coefficient [$\partial\gamma/\partial T$ in
 Eq. (B3)] being 0.0043 for the liquid metal, the thermo-
 capillary force is on the order of 10^4 Pa. Although this
 force is also 2 orders of magnitude smaller than the recoil
 pressure, it acts along a different direction from that of the
 recoil pressure and hence is not totally overruled. It can

724 still effectively drive the liquid metal on the interface to
725 flow from hot regions to cold regions (i.e., the Marangoni
726 flow).

727 On the rear keyhole wall, the temperature is still above
728 the boiling point in cases 1 and 2. The recoil pressure is on
729 the order of 10^5 Pa and is still dominant over the capillary
730 force, which is on the order of -10^4 Pa. However, the tem-
731 perature of the rear keyhole wall drops below the boiling
732 point in case 3, where no evaporation or recoil pressure is
733 available. The capillary force becomes dominant along the
734 normal direction of the rear keyhole wall. Along the tan-
735 gential direction, the thermocapillary force is on the order
736 of 10^3 Pa in all three cases and still effectively drives a
737 Marangoni flow on the rear keyhole wall.

738 On the top surface of the molten pool, the temperature is
739 usually below the boiling point and, hence, no recoil pres-
740 sure is available. This area is primarily controlled by the
741 capillary force and thermocapillary force, both of which
742 are on the order of magnitude of 10^3 Pa. The capillary
743 force tends to suppress any unevenness on the top sur-
744 face. The thermocapillary force drives the Marangoni flow
745 toward the tail and the side of the molten pool and is
746 primarily responsible for the development of vortex A in
747 Figs. 10(a)–10(c).

748 IV. CONCLUSIONS

749 In this paper, we present a combination of an
750 advanced multiphysics numerical model with state-of-the-
751 art dynamic x-ray radiography experiments to understand
752 the dynamic phenomena of laser-induced keyholing in
753 stainless steel 304. On the experimental side, *in situ* DXR
754 experiments provide observations of the keyhole shape
755 and fluctuation as a function of laser power and scanning
756 speed. As the laser scanning speed is increased at a con-
757 stant power, the keyhole depth decreases, the front wall
758 inclination and keyhole width increase, and the magnitude
759 of keyhole fluctuation decreases. On the modeling side,
760 our multiphysics model successfully reproduces the typ-
761 ical shapes observed by DXR under different conditions.
762 Moreover, the models provide information regarding the
763 distribution of laser absorption on the keyhole walls and
764 its effects on evaporation, thermofluidic flow, and keyhole
765 shape and fluctuation, all of which are inaccessible via
766 experiment.

767 On the front keyhole wall, the phenomena are domi-
768 nated by the absorption of fresh rays. In the case of a
769 lower scanning speed, a higher recoil pressure and a longer
770 fresh ray dwell time are available to accelerate drilling on
771 the front wall, leading to a higher drilling velocity and a
772 smaller inclination angle of the front wall. The incident
773 angle of fresh rays on the front wall is closer to 90° , and the
774 fresh ray absorption becomes less stable. Protrusions form
775 on the top of the front wall and flow downward, causing

unstable wavy fluctuation of the front wall. At higher scan- 776
777 ning speeds, the front wall is more inclined and the fresh
778 ray absorption becomes more stable, generating a flat and
779 stable front wall.

780 On the rear keyhole wall, the general shape is primar-
781 ily determined by the fluid dynamics initiated on the front
782 keyhole wall, and the reflected ray absorption is mainly
783 responsible for the fluctuation of the rear wall from its
784 nominal shape. At lower scanning speeds, the melt flow
785 initiated on the front keyhole wall circulates around the
786 keyhole with a lower speed along the horizontal direc-
787 tion. This slow horizontal flow immediately fills the empty
788 space behind the keyhole, resulting in a smaller keyhole
789 width. The smaller keyhole width significantly enhances
790 multiple reflection of the laser light, resulting in higher
791 absorbed power from reflected ray absorption. The distri-
792 bution of the reflected ray absorption varies significantly
793 over the rear keyhole wall (especially in the lower portion)
794 and causes the rear keyhole wall to fluctuate. At higher
795 scanning speeds, the melt flow goes around the keyhole at
796 a much higher speed along the horizontal direction. The
797 fast horizontal flow does not immediately fill the empty
798 space behind the keyhole, resulting in a larger keyhole
799 width. The reflected ray absorption is considerably lower
800 and does not cause any noticeable fluctuation of the rear
801 keyhole wall.

802 Multiple driving forces (namely, recoil pressure, capil-
803 lary force, and thermocapillary force) exist on the molten
804 pool surface and keyhole wall, and they are of different
805 significance to the molten pool flow and keyhole behavior.
806 Along the normal direction of the interface, the recoil pres-
807 sure pushes the molten metal away, thus maintaining the
808 opening of the keyhole, while the capillary force tends to
809 suppress any unevenness of the interface. The recoil pres-
810 sure from evaporation (when present) is usually orders of
811 magnitude larger than the capillary force on both the front
812 and rear keyhole wall, and the capillary force is only sig-
813 nificant when the surface temperature is below the boiling
814 point. The thermocapillary force is active along the tangent
815 direction of the interface and drives Marangoni flow. The
816 thermocapillary force is also orders of magnitude smaller
817 than the recoil pressure but its effect is always present.

818 The combination of DXR observation and multiphysics
819 modeling provides a valuable framework for understand-
820 ing the complex phenomena in laser-induced keyholes.
821 The framework is expected to contribute to understanding
822 the mechanisms of pore/spatter formation in laser-based
823 additive manufacturing and welding processes.

824 ACKNOWLEDGMENTS

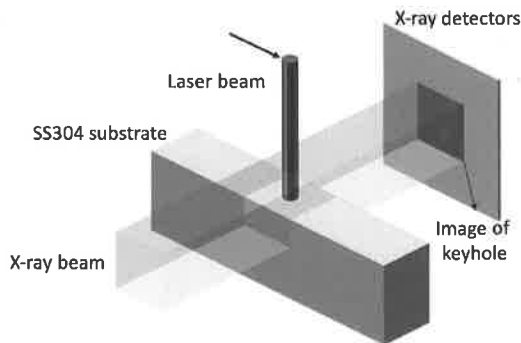
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843 APPENDIX A: METHODOLOGY FOR IN SITU 844 OBSERVATION OF LASER-INDUCED KEYHOLE 845 WITH DYNAMIC X-RAY RADIOGRAPHY

846 High-speed dynamic x-ray radiography experiments are
 847 performed at the 32-ID-B beamline at the APS. Figure 11
 848 shows a schematic of the experimental setup. In each
 849 case, a SS304 substrate about 400 μm thick is processed
 850 with a specific combination of laser power and scanning
 851 speed. The metal specimen is placed inside a SS cham-
 852 ber filled with argon gas to create an inert environment.
 853 The experiments are conducted at room temperature with-
 854 out preheating of the metal specimens. A laser beam of
 855 approximately 100 μm diameter, shown in Fig. 11, is used
 856 to scan along the centerline of the metal plate, orthogonal
 857 to the x-ray beam. The high-speed imaging camera cap-
 858 tures a side view of the laser keyhole shape that is exposed
 859 on the x-ray detectors and records full field x-ray images
 860 with a rate of 50 kHz and a spatial resolution of about 2 μm
 861 per pixel.

862 Experiments are performed in the parameter window
 863 of interest to metal AM. The laser power P is varied
 864 between 208 and 520 W with increments of 104 W and
 865 the laser scanning speed V between 300 and 1200 mm/s



with increments of 100 or 200 mm/s. For the purpose of
 this study, 14 combinations of P and V are tested using the
 DXR experiment to explore typical cases of the keyhole
 formation.

870 APPENDIX B: FORMULATION OF THE 871 MULTIPHYSICS MODELING OF 872 LASER-INDUCED KEYHOLE

873 A multiphysics model is employed to account for
 874 keyhole wall movement, laser-matter interaction, and mul-
 875 tiphase and multispecies thermofluidic flow in the laser-
 876 induced keyhole and molten pool. Model inputs include
 877 the laser parameters (laser power, spot size, power distri-
 878 bution, scanning speed); substrate geometry; and material
 879 properties. The model then predicts the 3D keyhole shape
 880 and the 3D distributions of temperature, pressure, and
 881 velocity in solid, liquid, and gas phases. The primary
 882 formulations of the model are briefly described in what
 883 follows, with a more detailed description of the model
 884 provided in Refs. [32] and [33].

885 In the current model, the level-set (LS) method [45] is
 886 used to track the interface between the condensed phases
 887 (solid and liquid metal) and noncondensed phases (ambient
 888 gas and metallic vapor). The LS equation is written as

$$\frac{\partial \varphi}{\partial t} + \left(\vec{V}_c - \frac{m_{\text{evap}}}{\rho_c} \cdot \vec{N} \right) \cdot \nabla \varphi = 0. \quad (\text{B1}) \quad 889$$

890 In Eq. (B1), φ is the LS function, which quantifies the
 891 distance from the centroid of a control volume to the molten
 892 pool surface. The molten pool surface (with a portion of
 893 it being the keyhole wall) is captured by the zero iso-LS
 894 surface, and the LS function is defined to be negative in
 895 the condensed phases and positive in the noncondensed
 896 phases. The two terms in the bracket of the left-hand side
 897 (lhs) of Eq. (B1) are two components that drive the inter-
 898 face motion. The first term, \vec{V}_c , is the convection velocity
 899 of the condensed phase at the interface. The second term
 900 is the surface recession velocity due to evaporation, with
 901 m_{evap} as the interface evaporation rate, ρ_c as the density
 902 of the condensed phase, and \vec{N} as the unit normal vector
 903 at the interface pointing to the noncondensed phases. We
 904 adopt the calculation of m_{evap} used by a numerical model
 905 [27] that was originally derived in Ref. [46].

906 The transport phenomena in all phases are governed
 907 by the conservation of mass, momentum, energy, and
 908 chemical species, which can be written, respectively, as

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{V}) = 0, \quad (\text{B2}) \quad 909$$

F11:1 FIG. 11. Sketch of dynamic x-ray radiography experiments.
 F11:2 The laser scanning direction is shown by the black arrow.

$$\begin{aligned} \frac{\partial \rho \vec{V}}{\partial t} + \nabla \cdot (\rho \vec{V} \vec{V}) &= -\nabla p + \nabla \cdot (\mu \nabla \vec{V}) - \frac{\mu}{\kappa} \vec{V} \\ &+ \rho_r \vec{g} \beta_T (T - T_r) \\ &+ \nabla_s T \frac{\partial \gamma}{\partial T} \delta_m(\varphi), \end{aligned} \quad (\text{B3})$$

$$\frac{\partial \rho h}{\partial t} + \nabla \cdot (\rho \vec{V} h) = \nabla \cdot (k \nabla T), \quad (\text{B4})$$

$$\frac{\partial \rho Y}{\partial t} + \nabla \cdot (\rho \vec{V} Y) = \nabla \cdot (\rho D \nabla Y). \quad (\text{B5})$$

In Eqs. (B2)–(B5), ρ is the density, \vec{V} is the velocity, p is the pressure, μ is the viscosity, κ is the isotropic permeability in the Kozeny-Carman equation, ρ_r is a reference density at the reference temperature T_r , \vec{g} is the gravitational force vector, β_T is the thermal expansion coefficient, T is the temperature, $\nabla_s T$ is the surface gradient of temperature, γ is the surface tension, $\partial \gamma / \partial T$ is the surface tension coefficient, $\delta_m(\varphi)$ is the modified delta function, h is the enthalpy, k is the thermal conductivity, Y is the mass fraction of iron vapor in the ambient gas, and D is the mass diffusivity of iron vapor. Note that the third through fifth terms on the right-hand side (rhs) of Eq. (B3) represent the dragging force in the mushy region, the thermal-buoyancy force, and the thermal-capillary force, respectively.

At the interface between the condensed and noncondensed phases, a set of jump conditions is enforced as boundary conditions of Eqs. (B2)–(B5). We use the notation $[X]$ to denote the jump between the condensed and noncondensed phases of a physical quantity X . First, a jump in temperature is enforced according to the theory of the Knudsen layer [47]:

$$[X] = T_c - T_{nc}, \quad (\text{B6})$$

where T_c and T_{nc} are the temperature of the condensed and noncondensed phase at the interface and are calculated according to Ref. [47]. The other jump conditions are obtained from the conservation of mass, momentum, and energy across the interface [48]. The mass conservation is written as

$$[V_N] = m_{\text{evap}} \left[\frac{1}{\rho} \right], \quad (\text{B7})$$

where V_N is the velocity normal to the interface. Equation (B7) enforces a jump of the velocity normal across the interface that results from evaporation. The momentum conservation is written as

$$[p] = \gamma K + p_{\text{recoil}}(T_c), \quad (\text{B8})$$

where K is the curvature of molten pool surface and p_{recoil} is the recoil pressure generated by the metal vapor. The term p_{recoil} is a function of T_c and is given in Refs. [32]

and [33]. Equation (B8) states that the pressure jump on the lhs is balanced by the summation of capillary force and recoil pressure on the rhs. Finally, the energy conservation is written as

$$[k \nabla T] = Q_{\text{laser}} - \sigma \epsilon (T^4 - T_\infty^4) - m_{\text{evap}} L_{\text{evap}}, \quad (\text{B9})$$

where Q_{laser} is the power intensity of laser absorption, σ is the Stefan-Boltzmann constant, ϵ is the emissivity, T_∞ is the ambient temperature, and L_{evap} is the latent heat of evaporation. The lhs of Eq. (B9) represents the jump of the conductive heat flux. The three terms on the rhs are the heat fluxes due to laser absorption, radiation, and evaporation, respectively.

The term Q_{laser} is predicted by a ray-tracing model [32]. The laser is divided into a finite number of rays, each with specific size, position, direction, and power. Upon each incidence of the ray on the keyhole wall, the power of the ray is partially absorbed by the surface according to the local laser absorptivity A , and the remainder of the power goes to the reflected rays, with the reflected direction determined by the law of reflection. The keyhole wall is divided into a finite number of segments in the model, and Q_{laser} for each segment is calculated by dividing the power absorption on each segment by the area of the segment. An example of predicted distribution of Q_{laser} over the entire keyhole is given in Fig. 4(b). Note that A is calculated as a function of the laser incident angle, χ , according to the Fresnel equation

$$\begin{aligned} A(\chi) &= 1 - \frac{1}{2} \\ &\times \left[\frac{1 + (1 - \epsilon_c \cos \chi)^2}{1 + (1 + \epsilon_c \cos \chi)^2} + \frac{\epsilon_c^2 - 2\epsilon_c \cos \chi + 2 \cos^2 \chi}{\epsilon_c^2 + 2\epsilon_c \cos \chi + 2 \cos^2 \chi} \right]. \end{aligned} \quad (\text{B10})$$

Here, ϵ_c is a material constant related to its electrical conductance. The laser is assumed unpolarized in this work, so the absorptivity is the average of the p -polarized and s -polarized lasers. The calculated absorptivity is shown by the purple curve in Fig. 8(b).

All the governing equations are solved using an in-house numerical solver based on the preconditioning method [49]. All the boundary conditions are treated with a “sharp interface” method [50]. For all the simulations in this work, the laser parameters and substrate geometry are set to be identical to those in the corresponding subset of DXR experiments, and the physical properties of SS304 are obtained from Refs. [32,33]. The three forces of interest in Table II, namely, the recoil pressure, capillary force, and thermocapillary force, are calculated using the terms p_{recoil} , γK , and $\nabla_s T \cdot \partial \gamma / \partial T$, respectively, in Eqs. (B3) and (B8).

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