

Direct Observations of Particle Dynamics in Magnetized Collisionless Shock Precursors in Laser-Produced Plasmas

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We present the first laboratory observations of time-resolved electron and ion velocity distributions in magnetized collisionless shock precursors. Thomson scattering of a probe laser beam was used to observe the interaction of a laser-driven, supersonic piston plasma expanding through an ambient plasma in an external magnetic field. From the Thomson-scattered spectra we measure time-resolved profiles of electron density, temperature, and ion flow speed, as well as spatially resolved magnetic fields from proton radiography. We observe direct evidence of the coupling between piston and ambient plasmas, including the acceleration of ambient ions driven by magnetic and pressure gradient electric fields, and deformation of the piston ion flow, key steps in the formation of magnetized collisionless shocks. Even before a shock has fully formed, we observe strong density compressions and electron heating associated with the pileup of piston ions. The results demonstrate that laboratory experiments can probe particle velocity distributions relevant to collisionless shocks, and can complement, and in some cases overcome, the limitations of similar measurements undertaken by spacecraft missions.

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Collisionless shocks are commonly found in systems in which strongly driven flows interact with preexisting magnetic fields, including planetary bow shocks in the heliosphere [1–3] and astrophysical shocks in supernova remnants [4–6]. In collisionless plasmas, these shocks form on spatial scales much smaller than the collisional mean free path due to dissipation mediated by electromagnetic fields. For most observed shocks, the fast inflow of particles can only be managed through the magnetic reflection of some particles back upstream, resulting in complex interactions between populations of inflowing, reflected, and shocked ions and electrons that are not fully understood. Consequently, fundamental questions, such as how energy is partitioned between electrons and ions across a collisionless shock [7–9], remain unanswered.

A key method for addressing these questions is the direct probing of particle velocity distributions, which has primarily been undertaken through *in situ* measurements by spacecraft. These missions have yielded a wealth of information on shock physics [10], and have recently begun to address the question of energy partitioning [11] as improved diagnostics have allowed high-resolution sampling of velocity distributions. Even so, spacecraft remain fundamentally limited, as they rely on the inherently noisy process of sampling shock crossings through multiple orbits and have difficulty gauging large-scale, 3D effects

due to undersampling [12,13]. Laboratory experiments, with reproducible and controllable plasma conditions, can complement and overcome some of these limitations to help address fundamental questions [14], and have recently extended the regimes of magnetized shock formation to strongly driven laser plasmas [15,16]. Moreover, velocity distributions can be similarly probed in the laboratory by measuring the Thomson scattering of light off plasma waves [17,18]. Early experiments [19,20] pioneered the use of Thomson scattering to study magnetized shocks, but were limited to a sparse sampling of the electron velocity distribution. Recent experiments have used this diagnostic to study velocity distributions in collisional shocks [21,22] and in unmagnetized collisionless counterstreaming flows [23,24].

In this Letter, we present the first laboratory observations of temporally resolved electron and ion velocity distributions in magnetized collisionless shock precursors. The distributions were acquired through Thomson scattering of a probe laser that diagnosed the interaction of a laser-driven, supersonic piston plasma expanding through an ambient plasma in an external magnetic field. Spatially resolved 2D proton radiography images of the magnetic field were also acquired. We directly observe the coupling between the piston and ambient plasmas, including the acceleration of ambient ions by magnetic and pressure

77 gradient electric fields and the pileup of piston ions behind
 78 the resulting compressed magnetic field. These effects
 79 constitute a shock precursor [25], a key step in piston-
 80 driven shock formation, and are found to depend critically
 81 on the presence of the ambient plasma and background
 82 magnetic field. The results build on an experimental plat-
 83 form that has studied high-Mach-number magnetized
 84 collisionless shocks [16,26], laser-driven magnetic recon-
 85 nection [27], and Weibel-mediated shocks [28].

86 *Setup.*—The experiments were carried out on the
 87 OMEGA laser facility [29] and are shown schematically
 88 in Fig. 1. The experiment utilizes two planar plastic (CH)
 89 2 targets and a set of copper coils to generate a magnetic
 90 field. The “piston” target is attached to the coils 3 mm
 91 from the target chamber center (TCC) and defines the exper-
 92 imental coordinate system, with \hat{x} along the target normal, \hat{y}
 93 parallel to the long edge, and \hat{z} parallel to the short edge. A
 94 second “ambient” target is centered at TCC along \hat{x} and
 95 offset 5 mm diagonally at a 45° angle. A background
 96 magnetic field is generated by the coils [30]. The initial
 97 field B_y has a peak strength of 10 T near the piston target
 98 and falls off like $1/x$ along \hat{x} , while it is nearly uniform
 99 across the target surface due to the elongated and stacked
 100 coil structure. A beam (351 nm, 100 J, 1 ns) incident on the
 101 ambient target creates an ambient plasma that expands
 102 through the background field. Twelve ns later, at time t_0 ,
 103 two drive beams (351 nm, 350 J, 2 ns) irradiate the piston
 104 target to generate a supersonic piston plasma, which
 105 expands through and sweeps up the ambient plasma and
 106 magnetic field. Note that while the ambient plasma also
 107 sweeps out magnetic flux, observations presented here and
 108 elsewhere [27] indicate that the ambient plasma is at least
 109 partially magnetized at time t_0 . This implies anomalously
 110 fast magnetic diffusion, which is not unprecedented and has
 111 been experimentally observed in other laser plasma experi-
 112 ments [31,32] where the magnetic cavity collapses on
 113 timescales much faster than dictated by Spitzer or Bohm
 114 diffusion.

115 The primary diagnostic was temporally resolved Thomson
 116 scattering using a 2ω probe beam (527 nm, 30-50 J, 2 ns)
 117 [33]. Scattered light from the probe beam was collected from
 118 a localized volume ($50 \times 50 \times 70 \mu\text{m}^3$) such that the probed
 119 wave vector $\mathbf{k} = \mathbf{k}_i - \mathbf{k}_s$ was directed along the piston
 120 expansion direction (i.e., along \hat{x}), where \mathbf{k}_i is the incident
 121 wave vector and \mathbf{k}_s is the scattered wave vector [Fig. 1(b)].
 122 The scattering angle was 63°, yielding a scattering parameter
 123 $\alpha = 1/k\lambda_{de} \approx 1.5$ for typical plasma parameters (i.e., the
 124 collective regime). The collected light was split along two
 125 beam paths. One path measured light scattered from electron
 126 plasma waves (EPW), which can provide information on the
 127 electron density and temperature. The other path measured
 128 light scattered from ion acoustic waves (IAW), which can
 129 also diagnose the electron temperature, as well as the ion
 130 temperature and flow speed. The EPW and IAW signals were
 131 passed through spectrometers with wavelength resolutions

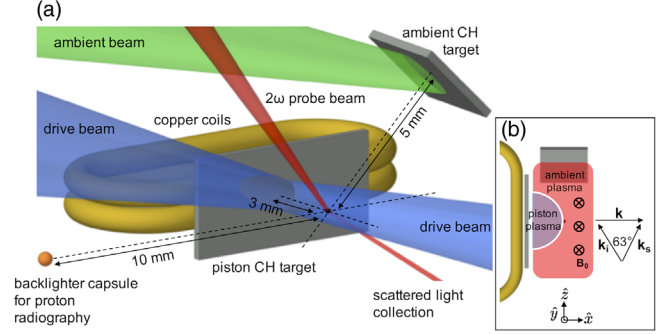


FIG. 1. (a) Experimental setup. A background magnetic field primarily directed along \hat{y} is preimposed using current-carrying copper wires. A laser ablates a CH target to create an ambient plasma. Two drive beams then generate a CH piston plasma that expands through the ambient plasma to drive a shock. Temperature, density, and velocity are diagnosed in the \hat{x} direction using Thomson scattering with a 2ω probe beam. Twenty beams (not shown) compress a DHe3 backlighter capsule to generate monoenergetic protons that probe the magnetic field structure in the x - y plane. (b) Top-down schematic view of the setup and Thomson scattering geometry.

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of 0.5 and 0.05 nm, respectively, and imaged onto streak cameras with a temporal resolution of 50 ps. The location of the probed plasma ranged from 3 to 4 mm from the piston target along \hat{x} . The scattered signal was streaked for 2 ns starting 3 to 4.5 ns after t_0 .

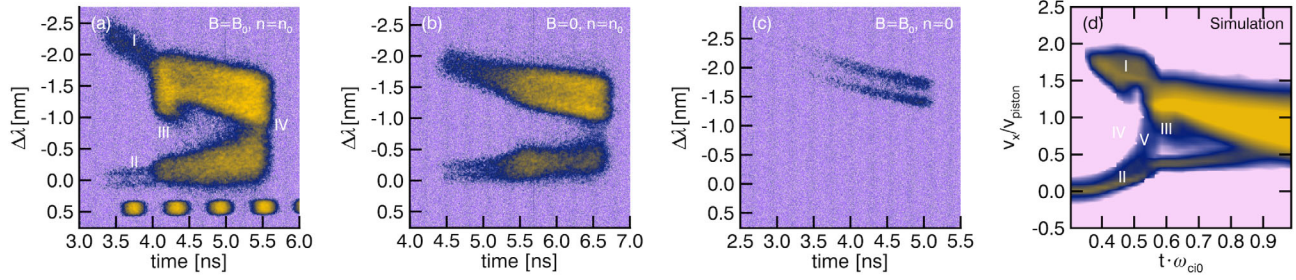
The magnetic field structure was measured using proton radiography [34]. A 420 μm diameter glass capsule filled with DHe3 was placed 10 mm from TCC along \hat{z} and irradiated by 20 beams at $t_0 + 3$ ns. The resulting implosion produced 3 and 14.7 MeV protons as fusion byproducts, which passed through the plasma and were collected on CR-39 plates placed 154 mm from TCC (geometric magnification $M = 16.4$). The protons leave tracks in the CR-39 that correspond to a 2D map of proton deflections in the x - y plane, which can be converted to path-integrated magnetic field amplitudes.

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Results.—Figure 2 shows streaked IAW spectra taken under three experimental configurations: (a) a magnetized piston-ambient interaction, (b) an unmagnetized piston-ambient interaction, and (c) a magnetized piston expansion. The EPW spectrum corresponding to Fig. 2(a) is shown in Fig. 3(a), and a proton radiograph taken under the same conditions is shown in Fig. 3(c). The ambient plasma was measured at TCC using Thomson scattering in the absence of a piston plasma over the same time intervals as in Fig. 2. The measurements yielded a time-averaged mean electron density $n_{e0} = 0.9 \pm 0.2 \times 10^{18} \text{ cm}^{-3}$ and temperature $T_{e0} = 40 \pm 10 \text{ eV}$ [35].

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The spectra show qualitative signatures of a magnetized collisionless shock precursor, and can be divided into four distinct regions in the IAW spectra, labeled I–IV in Fig. 2(a). Region I consists of piston ions that are streaming through the ambient plasma (region II) but largely



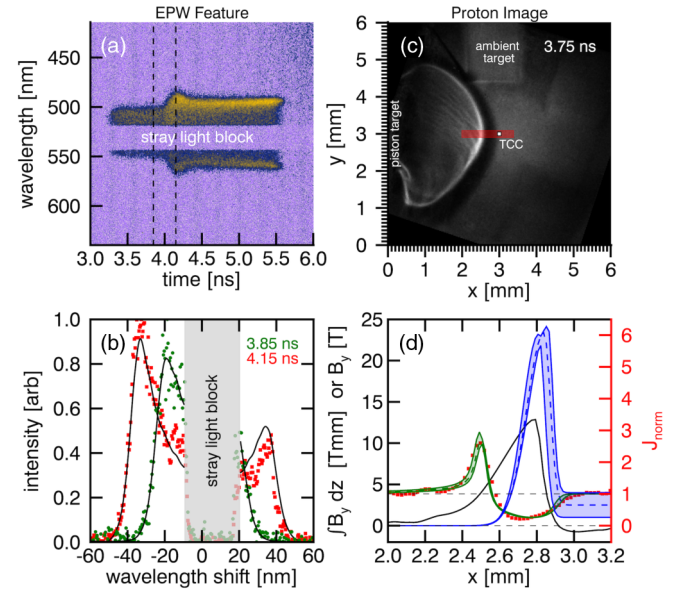
F2:1 FIG. 2. IAW spectra of piston-ambient interactions under three experimental conditions: (a) magnetized ambient plasma,
 F2:2 (b) unmagnetized ambient plasma, and (c) no ambient plasma. Data in (a) and (c) were taken at $x = 3$ mm (TCC), while in (b) were
 F2:3 taken at $x = 4$ mm. The wavelength shift $\Delta\lambda$ is relative to the probe beam, and the marks at the bottom of (a) are timing fiducials.
 F2:4 (d) Simulated ion velocity space in conditions similar to (a), with velocity relative to the piston speed and time relative to the upstream
 F2:5 gyrofrequency. Regions of interest are labeled with Roman numerals and discussed in the text.

165 unaffected by the magnetic field. A key step in piston-
 166 driven shock formation is the sweeping up of ambient
 167 plasma [36] and the resulting compression of the magnetic
 168 field.

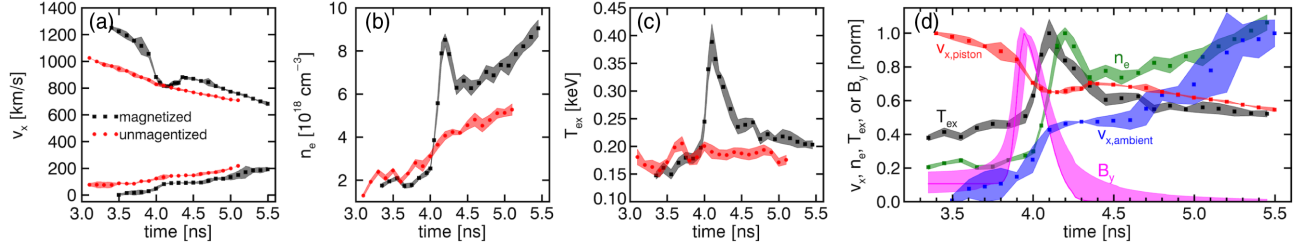
169 The increased field then causes a pileup of piston plasma
 170 and deformation of the piston flow. Both the ambient ion
 171 acceleration and piston deformation are seen in region III,
 172 which also corresponds to the peak in the EPW spectra
 173 in Fig. 3(a). Eventually, most of the ambient ions not
 174 participating in shock formation are swept up by the piston,
 175 which results in the merging of the piston and ambient
 176 plasmas in region IV. Without a background magnetic field
 177 [Fig. 2(b)], no ion pileup or flow deformation is observed,
 178 though the ambient ions are still eventually swept up.
 179 Likewise, Fig. 2(c) shows that with only a magnetized
 180 piston plasma, no shock precursor forms. These last two
 181 cases indicate that the presence of both the ambient plasma
 182 and background field is critical to shock formation. Lastly,
 183 Fig. 2(d) is the x component of the ion velocity distribution
 184 in the Thomson-scattering volume as a function of time
 185 from a 1D multispecies (CH) PSC [37,38] particle-in-cell
 186 simulation under conditions similar to Fig. 2(a). The four
 187 regions of Fig. 2(a) are clearly visible in the simulation and
 188 show that there is strong correspondence between the
 189 velocity distributions and the Thomson-scattered spectra.
 190 The simulation also shows additional features due to the
 191 H plasma (region V), but calculations indicate that the
 192 H ion acoustic waves would be heavily Landau damped
 193 relative to the C waves and so would not be observable.

194 Figure 3(c) shows a 14.7 MeV proton image taken at
 195 $t_0 + 3.75$ ns under the same conditions as Fig. 2(a). White,
 196 high-fluence and dark, low-fluence shells that result from
 197 the deflection of protons by the B_y magnetic field can be
 198 clearly seen, and represent large gradients in the path-
 199 integrated magnetic fields. Since there is no evidence at this
 200 late time of large-scale proton deflections due the expansion
 201 of the ambient plasma, the observed features primarily
 202 represent the 2D projection of the 3D magnetic cavity
 203 created by the expanding piston plasma. In Fig. 3(d), along
 204 a 1D profile through TCC [red box in Fig. 3(c)], we invert

205 the proton fluence profile (red squares) [39] to reconstruct
 206 the path-integrated magnetic field $\int B_y dz$ (black line).
 207 The result indicates that magnetic flux has been swept
 208 up into a thin shell, consistent with an expanding piston.
 209 To unfold the original field, we assume a simple 3D model
 210 $B_y(x, y, z)$ for the magnetic cavity [40] and iteratively fit
 211 the model parameters to find the best match between the
 212 model-generated synthetic fluence and measured fluence.



F3:1 FIG. 3. (a) Streaked Thomson-scattered spectrum of the EPW
 F3:2 feature taken at TCC, corresponding to Fig. 2(a). (b) Two
 F3:3 example profiles at time $t_0 + 3.85$ ns (green) and $t_0 + 4.15$ ns
 F3:4 (red), along with best fits (black). (c) Proton radiography image
 F3:5 taken at time $t_0 + 3.75$ ns using 14.7 MeV protons. (d) Proton
 F3:6 intensity (red squares) taken from the red region in (c), normalized
 F3:7 to the mean intensity, and the associated reconstructed
 F3:8 path-integrated magnetic field $\int B_y dz$ (black). Also shown is the
 F3:9 normalized proton intensity (green dashed) forward-modeled
 F3:10 from a 2D synthetic magnetic field $B_y(x, z)$, which has the dashed
 F3:11 blue profile at $z = 0$. The model uncertainties are shown as
 F3:12 shaded regions.



F4:1 FIG. 4. Thomson scattering results. Measured (a) piston and ambient ion flow speed v_x , (b) electron density n_e , and (c) electron
 F4:2 temperature T_{ex} for a piston plasma expanding through a magnetized (black squares) and unmagnetized (red circles) ambient plasma.
 F4:3 (d) Electron temperature (black), density (green), magnetic field (magenta), and piston (red) and ambient (blue) ion flow speed for the
 F4:4 magnetized case. The magnetic field profile was constructed from the model in Fig. 3(d). Error bars are shown as shaded regions. The
 F4:5 magnetized plasma were probed at $x = 3$ mm (TCC), while the unmagnetized plasma was probed at $x = 4$ mm. The unmagnetized data
 F4:6 have been shifted forward in time by 1.5 ns for ease of comparison.

213 The best fit synthetic fluence (green line) is shown in
 214 Fig. 3(d), along with the corresponding model field $B_y(x)$
 215 (blue line) that passes through the location of the Thomson
 216 scattering measurements. Here, the model field has a peak
 217 value $B_{y,peak} = 23 \pm 2$ T at $x_{peak} = 2.85 \pm 0.05$ mm, though
 218 the local upstream value $B_{y0} = 3 \pm 2$ T is not well con-
 219 strained. Comparison of the 3 and 14.7 MeV protons
 220 images also indicates that the protons are not significantly
 221 deflected by electric fields [40].

222 We can further quantify the Thomson-scattered spectra
 223 in Figs. 2 and 3 by iteratively fitting the data with a spectral
 224 model of the scattered power [33] to extract the time-
 225 resolved electron density n_e and x component of the
 226 electron temperature T_{ex} and ion flow speed v_x . An
 227 example EPW spectrum and fit is shown in Fig. 3(b). To
 228 perform error analysis, we employ a Monte Carlo approach
 229 in which the extracted plasma parameters represent the
 230 mean value over 50 fits, with error bars corresponding to
 231 the standard deviation. In all cases, the EPW spectral fits
 232 assumed Maxwellian velocity distributions. In contrast, the
 233 IAW spectra involve multiple ion species (C and H) and
 234 multiple flows from potentially non-Maxwellian ion dis-
 235 tributions. Extracting parameters from these spectra is
 236 beyond the scope of this Letter and will be reported
 237 separately. Instead, we only determine the ion flow speed
 238 from the Doppler shifts of the spectra, which can be
 239 accurately resolved without knowing the exact form of
 240 the scattered power [18]. Based on the results of these
 241 fits (see Fig. 4), we can justify the use of Maxwellian
 242 distributions by estimating the electron τ_{ee} and electron-ion
 243 τ_{ei} collision times relative to the fastest gradient timescales
 244 $\tau_s \sim 200$ ps and the electron plasma frequency ω_{pe} . We find
 245 for the electrons that $\tau_{ee} < \tau_s$, indicating that the electrons
 246 are well thermalized, and that $\tau_{pe} \ll \tau_{ei} < \tau_{ee}$, so that
 247 collisions do not significantly affect the EPW spectra.
 248 Furthermore, the spatial scales are dominantly determined
 249 by the piston-ambient ion (and eventually ambient-ambient
 250 ion) interaction, which is highly collisionless ($\tau_{pa}/\tau_s \gg 1$)
 251 due the large flow velocities in these experiments.

A summary of the Thomson scattering results is shown
 in Fig. 4 for magnetized (black) and unmagnetized (red)
 piston-ambient interactions. Figure 4(a) shows two sets of
 flow speeds v_x extracted from the IAW spectra that
 correspond to the faster (piston) and slower (ambient)
 moving populations. For the magnetized case, the piston
 ions exhibit a rapid deceleration around $t_0 + 4.0$ ns,
 coincident with the onset of the region of ion pileup in
 Fig. 2(a). Over the same time the ambient ions are
 accelerated, and then plateau for several hundred ps
 before being accelerated again as they begin to merge with
 the piston plasma. In the unmagnetized case, the piston
 ions show no deceleration and are consistent with a free-
 streaming expansion ($v \propto 1/t$). Figures 4(b) and 4(c) show
 electron density and temperature extracted from the EPW
 spectra. In the region of ion pileup, the magnetized case
 exhibits a strong density compression $n_e/n_{e0} \approx 10$,
 steep density ramp $\tau_n \sim 200$ ps, and electron heating
 $T_{ex}/T_{e0} \approx 10$, where n_{e0} and T_{e0} are the density and
 temperature in the stand-alone ambient plasma. The shock
 precursor has a time-of-flight speed of $v_{sh} \approx 750$ km/s,
 corresponding to an acoustic Mach number $M_s =$
 $v_{sh}/C_{s,C} \approx 15$, where $C_{s,C}$ is the upstream C ion sound
 speed evaluated at T_{e0} . No density compression or electron
 heating is observed in the unmagnetized case.

Figure 4(d) combines temperature (black), density
 (green), magnetic field (purple), and piston (red) and
 ambient (blue) ion flow results for the magnetized case.
 The field is plotted assuming that it is slowly changing on
 the timescales of interest, so that the spatial profile can be
 converted to a temporal profile using the time-of-flight
 speed $v_{field} = 760 \pm 20$ km/s. The combined profiles
 directly demonstrate the piston-ambient ion coupling that
 is critical to shock formation. First, the magnetic field acts
 as an interface between the highly magnetized ambient and
 piston electrons: swept-over ambient electrons compress
 the field at the leading edge while piston electrons expel
 the field [41]. Consequently, the piston electrons (and ions)
 will necessarily pile up behind the magnetic compression,
 as observed. This results in a localized electron density
 peak that then transitions into the smooth ablation profile of

293 the piston plume. The temperature in turn rises adiabatically
 294 ($T_e \propto n_e^{2/3}$) with the density, consistent with collisional
 295 electrons.

296 While at this stage in formation the density profile
 297 primarily reflects piston dynamics, it also crucially leads
 298 to the sweeping up of ambient ions through the pressure
 299 gradient electric field $E_P = -\nabla P_e/en_e$, where $P_e = n_e T_e$,
 300 and the magnetic gradient field $E_B = -\nabla B^2/2\mu_0 en_e$. As
 301 seen in Fig. 4(d), the change in ambient ion speed between
 302 3.75 and 4.15 ns is $\Delta v \sim 66$ km/s. Assuming that gradients
 303 of the temporally or spatially varying functions f can be
 304 related through $df/dx \approx (1/v_{\text{field}})df/dt$, we can estimate
 305 the effect of these fields on the ambient ions from our
 306 measured data. Over the same time range, the pressure
 307 gradient field accelerates the ions by $\Delta v_{EP} = \int (Z_C e/m_C)$
 308 $E_P dt \approx 38$ km/s, while the magnetic gradient field accel-
 309 erates the ions by $\Delta v_{EB} \approx 19$ km/s, giving a combined
 310 change $\Delta v_E \approx 57$ km/s that agrees well with the measure-
 311 ments. The relative contributions are also consistent with
 312 the $\beta_e = 2\mu_0 P_e/B^2 > 1$ conditions in these experiments,
 313 since $E_P/E_B \propto \beta_e$. While the magnitude of these electric
 314 fields has been measured [42] or inferred from simulations
 315 [36] in previous $\beta_e < 1$ piston-ambient coupling experi-
 316 ments, this is the first time that they have been estimated
 317 directly from localized measurements when $\beta_e > 1$. The
 318 piston ion flow is correspondingly affected by the electric
 319 fields, decelerating slightly behind the density compres-
 320 sion, and then strongly accelerating at the leading edge. The
 321 deformation of the ion flows is therefore a key signature of
 322 the onset of piston-driven shock formation.

323 In summary, we have measured for the first time through
 324 Thomson scattering the evolution of electron and ion
 325 velocity distributions in magnetized collisionless shock
 326 precursors. We have extracted time-resolved profiles of
 327 electron temperature, density, and ion flow speed, which
 328 indicate the development of strong density compressions
 329 and electron heating associated with the pileup of piston
 330 ions and acceleration of ambient ions by magnetic and
 331 pressure gradient electric fields. Proton radiography images
 332 confirm that there is an associated strong magnetic comp-
 333 ression in the same region. This acceleration of ambient
 334 ions and subsequent deformation of the piston ion flow is
 335 a key component of magnetized shock formation, and is
 336 not observed without both a background magnetic field
 337 and ambient plasma. Since the distributions can in principle
 338 be probed along any direction, these results will enable
 339 future experiments to study multidimensional distribution
 340 functions in a manner analogous to spacecraft, allowing
 341 direct comparisons between studies of space and laboratory
 342 collisionless shocks.

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[1] E. J. Smith, L. Davis, D. E. Jones, P. J. Coleman, D. S. 362
 Colburn, P. Dyal, and C. P. Sonett, *Science* **188**, 451 (1975). 363
 [2] E. J. Smith, L. Davis, D. E. Jones, P. J. Coleman, D. S. 364
 Colburn, P. Dyal, and C. P. Sonett, *Science* **207**, 407 (1980). 365
 [3] A. H. Sulaiman, A. Masters, M. K. Dougherty, D. Burgess, 366
 M. Fujimoto, and G. B. Hospodarsky, *Phys. Rev. Lett.* **115**, 367
 125001 (2015). 368
 [4] P. J. Cargill and K. Papadopoulos, *Astrophys. J. Lett.* **329**, 369
 L29 (1988). 370
 [5] D. S. Spicer, S. P. Maran, and R. W. Clark, *Astrophys. J.* 371
356, 549 (1990). 372
 [6] A. Bamba, R. Yamazaki, M. Ueno, and K. Koyama, 373
Astrophys. J. **589**, 827 (2003). 374
 [7] M. Balikhin, M. Gedalin, and A. Petrukovich, *Phys. Rev.* 375
Lett. **70**, 1259 (1993). 376
 [8] B. Lembège, P. Savoini, M. Balikhin, S. Walker, and V. 377
 Krasnoselskikh, *J. Geophys. Res.* **108**, 1256 (2003). 378
 [9] S. J. Schwartz, E. Henley, J. Mitchell, and V. Krasnoselskikh, 379
Phys. Rev. Lett. **107**, 215002 (2011). 380
 [10] D. Burgess and M. Scholer, *Collisionless Shocks in Space* 381
Plasmas: Structure and Accelerated Particles, Cambridge 382
 Atmospheric and Space Science Series (Cambridge 383
 University Press, Cambridge, England, 2015). 384
 [11] L.-J. Chen *et al.*, *Phys. Rev. Lett.* **120**, 225101 (2018). 385
 [12] V. V. Lobzin, V. V. Krasnoselskikh, J.-M. Bosqued, J.-L. 386
 Pinçon, S. J. Schwartz, and M. Dunlop, *Geophys. Res. Lett.* 387
34, L05107 (2007). 388
 [13] A. Johlander *et al.*, *Phys. Rev. Lett.* **117**, 165101 (2016). 389
 [14] G. G. Howes, *Phys. Plasmas* **25**, 055501 (2018). 390
 [15] C. Niemann, W. Gekelman, C. G. Constantin, E. T. Everson, 391
 D. B. Schaeffer, A. S. Bondarenko, S. E. Clark, D. Winske, 392
 S. Vincena, B. Van?Compernelle, and P. Pribyl, *Geophys.* 393
Res. Lett. **41**, 7413 (2014). 394
 [16] D. B. Schaeffer, W. Fox, D. Haberberger, G. Fiksel, 395
 A. Bhattacharjee, D. H. Barnak, S. X. Hu, and K. 396
 Germaschewski, *Phys. Rev. Lett.* **119**, 025001 (2017). 397
 [17] D. H. Froula, J. S. Ross, L. Divol, and S. H. Glenzer, *Rev.* 398
Sci. Instrum. **77**, 10E522 (2006). 399
 [18] J. Sheffield, D. H. Froula, S. H. Glenzer, and N. Luhmann, 400
Plasma Scattering of Electromagnetic Radiation, 2nd ed. 401
 (Academic Press, New York, 2011). 402
 [19] J. W. M. Paul, G. C. Goldenbaum, A. Iiyoshi, L. S. Holmes, 403
 and R. A. Hardcastle, *Nature (London)* **216**, 363 (1967). 404
 [20] A. W. DeSilva and J. A. Stamper, *Phys. Rev. Lett.* **19**, 1027 405
 (1967). 406

- 407 [21] L. G. Suttle, J. D. Hare, S. V. Lebedev, G. F. Swadling, G. C. 446
 408 Burdiak, A. Ciardi, J.P. Chittenden, N.F. Loureiro, 447
 409 N. Niasse, F. Suzuki-Vidal, J. Wu, Q. Yang, T. Clayson, 448
 410 A. Frank, T. S. Robinson, R. A. Smith, and N. Stuart, *Phys.* 449
 411 *Rev. Lett.* **116**, 225001 (2016). 450
- 412 [22] H. G. Rinderknecht, H.-S. Park, J. S. Ross, P. A. Amendt, 451
 413 D. P. Higginson, S. C. Wilks, D. Haberberger, J. Katz, D. H. 452
 414 Froula, N. M. Hoffman, G. Kagan, B. D. Keenan, and E. L. 453
 415 Vold, *Phys. Rev. Lett.* **120**, 095001 (2018). 454
- 416 [23] J. S. Ross, H.-S. Park, R. Berger, L. Divol, N. L. Kugland, 455
 417 W. Rozmus, D. Ryutov, and S. H. Glenzer, *Phys. Rev. Lett.* 456
 418 **110**, 145005 (2013). 457
- 419 [24] T. Morita, Y. Sakawa, K. Tomita, T. Ide, Y. Kuramitsu, K. 458
 420 Nishio, K. Nakayama, K. Inoue, T. Moritaka, H. Ide, M. 459
 421 Kuwada, K. Tsubouchi, K. Uchino, and H. Takabe, *Phys.* 460
 422 *Plasmas* **20**, 092115 (2013). 461
- 423 [25] D. B. Schaeffer, E. T. Everson, D. Winske, C. G. Constantin, 462
 424 A. S. Bondarenko, L. A. Morton, K. A. Flippo, D. S. 463
 425 Montgomery, S. A. Gaillard, and C. Niemann, *Phys.* 464
 426 *Plasmas* **19**, 070702 (2012). 465
- 427 [26] D. B. Schaeffer, W. Fox, D. Haberberger, G. Fiksel, A. 466
 428 Bhattacharjee, D. H. Barnak, S. X. Hu, K. Germaschewski, 467
 429 and R. K. Follett, *Phys. Plasmas* **24**, 122702 (2017). 468
- 430 [27] G. Fiksel, W. Fox, A. Bhattacharjee, D. H. Barnak, P.-Y. 469
 431 Chang, K. Germaschewski, S. X. Hu, and P. M. Nilson, 470
 432 *Phys. Rev. Lett.* **113**, 105003 (2014). 471
- 433 [28] W. Fox, G. Fiksel, A. Bhattacharjee, P.-Y. Chang, K. 472
 434 Germaschewski, S. X. Hu, and P. M. Nilson, *Phys. Rev.* 473
 435 *Lett.* **111**, 225002 (2013). 474
- 436 [29] T. R. Boehly, R. S. Craxton, T. H. Hinterman, J. H. Kelly, 475
 437 T. J. Kessler, S. A. Kumpan, S. A. Letzring, R. L. McCrory, 476
 438 S. F. B. Morse, W. Seka, S. Skupsky, J. M. Soures, and C. P. 477
 439 Verdon, *Rev. Sci. Instrum.* **66**, 508 (1995). 478
- 440 [30] G. Fiksel, A. Agliata, D. Barnak, G. Brent, P.-Y. Chang, L. 479
 441 Folsbee, G. Gates, D. Hasset, D. Lonobile, J. Magoon, D. 480
 442 Mastro Simone, M. J. Shoup, and R. Betti, *Rev. Sci. Instrum.* 481
 443 **86**, 016105 (2015). 482
- 444 [31] C. Niemann, W. Gekelman, C. G. Constantin, E. T. Everson, 483
 445 D. B. Schaeffer, S. E. Clark, D. Winske, A. B. Zylstra, 484
 P. Pribyl, S. K. P. Tripathi, D. Larson, S. H. Glenzer, 446
 and A. S. Bondarenko, *Phys. Plasmas* **20**, 012108 447
 (2013). 448
- [32] D. B. Schaeffer, L. R. Hofer, E. N. Knall, P. V. Heuer, C. G. 449
 Constantin, and C. Niemann, *High Power Laser Sci. Eng.* **6**, 450
 e17 (2018). 451
- [33] R. K. Follett, J. A. Delettrez, D. H. Edgell, R. J. Henchen, J. 452
 Katz, J. F. Myatt, and D. H. Froula, *Rev. Sci. Instrum.* **87**, 453
 11E401 (2016). 454
- [34] R. D. Petrasso, C. K. Li, F. H. Seguin, J. R. Rygg, J. A. 455
 Frenje, R. Betti, J. P. Knauer, D. D. Meyerhofer, 456
 P. A. Amendt, D. H. Froula, O. L. Landen, P. K. Patel, 457
 J. S. Ross, and R. P. J. Town, *Phys. Rev. Lett.* **103**, 458
 085001 (2009). 459
- [35] Because of heating of the electrons by the probe beam for 460
 temperatures less than 100 eV (see for example Ref. [18]), 461
 these measurements are most likely an overestimate of the 462
 true electron temperature. 463
- [36] A. S. Bondarenko, D. B. Schaeffer, E. T. Everson, S. E. 464
 Clark, B. R. Lee, C. G. Constantin, S. Vincena, B. Van 465
 Comperolle, S. K. P. Tripathi, D. Winske, and C. Niemann, 466
Nat. Phys. **13**, 573 (2017). 467
- [37] K. Germaschewski, W. Fox, S. Abbott, N. Ahmadi, K. 468
 Maynard, L. Wang, H. Ruhl, and A. Bhattacharjee, 469
J. Comput. Phys. **318**, 305 (2016). 470
- [38] W. Fox, J. Matteucci, C. Moissard, D. B. Schaeffer, A. 471
 Bhattacharjee, K. Germaschewski, and S. X. Hu, *Phys.* 472
Plasmas **25**, 102106 (2018). 473
- [39] A. F. A. Bott, C. Graziani, P. Tzeferacos, T. G. White, D. Q. 474
 Lamb, G. Gregori, and A. A. Schekochihin, *J. Plasma Phys.* 475
83, 905830614 (2017). 476
- [40] See Supplemental Material at [http://link.aps.org/](http://link.aps.org/supplemental/10.1103/PhysRevLett.000.000000) 477
[supplemental/10.1103/PhysRevLett.000.000000](http://link.aps.org/supplemental/10.1103/PhysRevLett.000.000000) for more 478
 information on the analysis of the proton radiography 479
 images. 480
- [41] T. P. Wright, *Phys. Fluids* **14**, 1905 (1971). 481
- [42] J. Bonde, S. Vincena, and W. Gekelman, *Phys. Plasmas* **25**, 482
 042110 (2018). 483