

REFERENCES

1. G. Miller, et al., Phys. Rev. D6 3011 (1972).
2. D. H. Perkins, in Proceedings of the XVI International Conference on High Energy Physics, V. 4, 189 (1972). For more recent results see the Proceedings of the 4th International Conference on Neutrino Physics and Astrophysics; Downingtown, Pennsylvania (1974).
3. R. P. Feynman, Photon Hadron Interactions, (W.A. Benjamin, Reading, Massachusetts, 1972), J. D. Bjorken and E. Paschos, Phys. Rev. 185 1975 (1969).
4. D. Gross and F. Wilczek, Phys. Rev. Letters 30 1343 (1973);  
H. D. Politzer, Phys. Rev. Letters 30, 1346 (1973);  
G. 't Hooft (unpublished).
5. D. Gross and F. Wilczek, Phys. Rev. D9 980 (1974).
6. H. Georgi and H. D. Politzer, Phys. Rev. D9 416 (1974).
7. D. Bailin, A. Love and D. V. Nanopoulos, Lettere al Nuovo Cimento 9 501 (1974).
8. Some of these results have already been briefly presented. See  
A. Zee; Proceedings of the 4th International Conference of Neutrino Physics and Astrophysics, Downingtown, Pennsylvania (1974).
9. C. Callan and D. Gross, Phys. Rev. Letters 22 156 (1969).
10. We thank R. Feynman for a discussion on this point. Calculations by M. Calvo (Princeton University thesis, unpublished) suggests that  $\gamma_n(\bar{g}) = a_1 \log n + a_2 \bar{g}^2 \log^2 n + \dots$ , so the  $n \rightarrow \infty$  limit is highly nonuniform.
11. K. Wilson, Phys. Rev. 179 1499 (1969);  
G. Mack, Phys. Rev. Letters 25 400 (1970);  
S. Cicciarello, R. Gatto, G. Sartori and M. Tonin, Phys. Letters 30B 546 (1969).